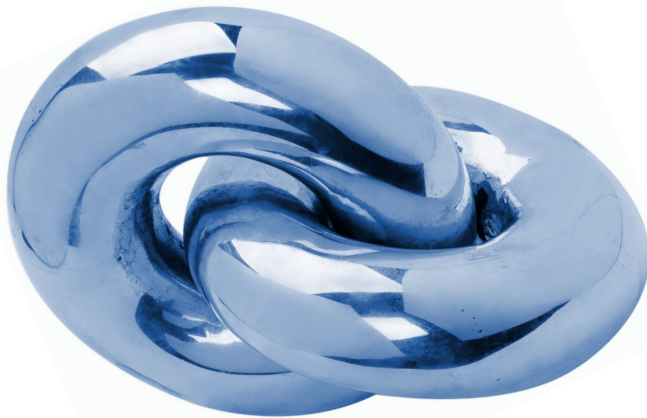


# (Finite) modular symmetries and the strong CP problem

in collaboration with S.T. Petcov [[2404.08032](#)]

João Penedo (INFN, Roma Tre)

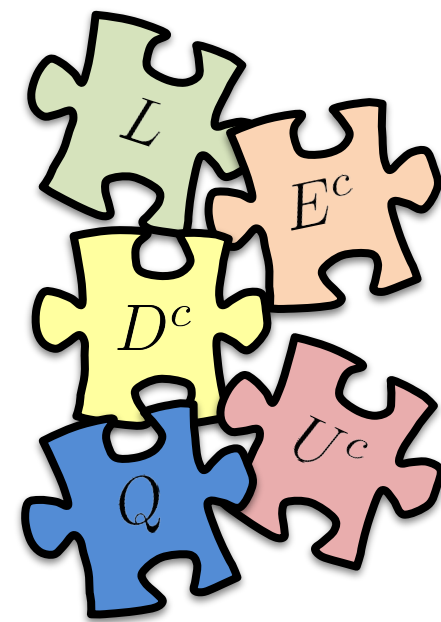
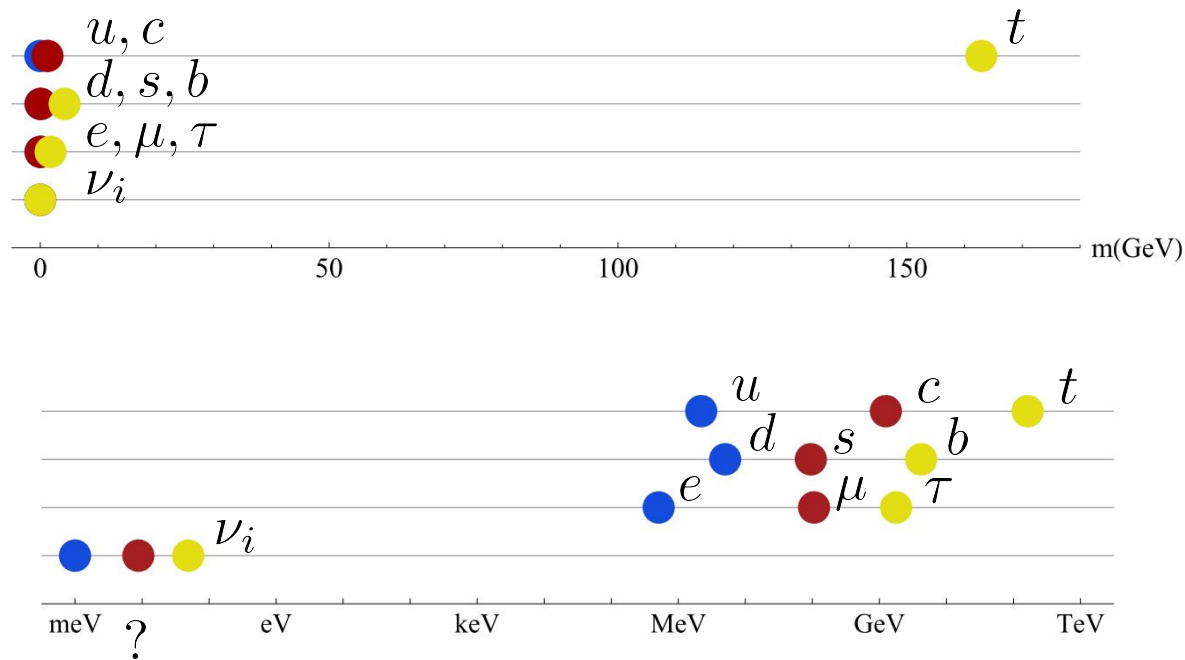


7 June 2024

PLANCK @ Lisbon



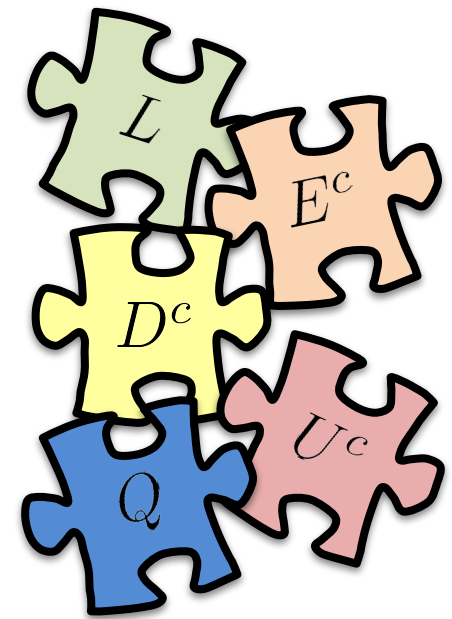
# The flavour puzzle



adapted from R. Toorop's PhD thesis

# The flavour puzzle

$$U_{\text{PMNS}} \sim \begin{array}{c} e \\ \mu \\ \tau \end{array} \begin{array}{c} \nu_1 \quad \nu_2 \quad \nu_3 \\ \left[ \begin{array}{ccc} \blacksquare & \blacksquare & \cdot \\ \blacksquare & \blacksquare & \blacksquare \\ \blacksquare & \blacksquare & \blacksquare \end{array} \right] \end{array}$$



adapted from P. Novichkov's slides at PASCOS 2021

# 3ν flavour paradigm



## Masses: ordering

$$\frac{\Delta m_{\odot}^2}{|\Delta m_A^2|} \sim \frac{1}{30}$$

Normal ordering (NO)

$$m_1 < m_2 < m_3$$

$$\text{————— } m_3$$

$$\text{————— } m_2$$

$$\text{————— } m_1$$

?

Inverted ordering (IO)

$$m_3 < m_1 < m_2$$

$$\text{————— } m_2$$

$$\text{————— } m_1$$

$$\text{————— } m_3$$

?

VS.

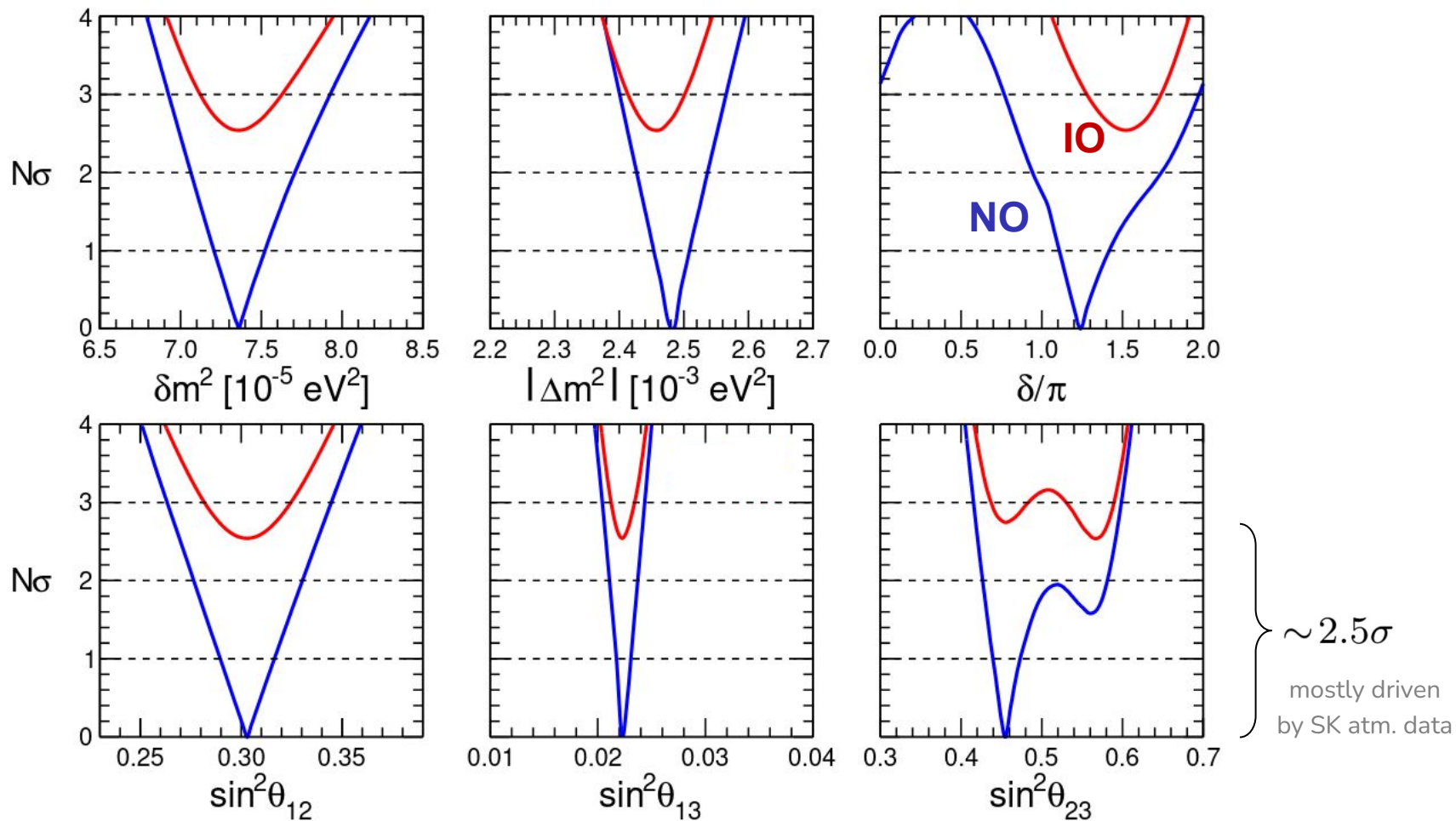
## Mixing matrix parameterisation

$$c_{ij} \equiv \cos \theta_{ij}, \quad s_{ij} \equiv \sin \theta_{ij}$$

$$U_{\text{PMNS}} = \begin{pmatrix} 1 & & \\ & c_{23} & s_{23} \\ & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & & s_{13}e^{-i\delta} \\ & 1 & \\ -s_{13}e^{i\delta} & & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & \\ -s_{12} & c_{12} & \\ & & 1 \end{pmatrix} \begin{pmatrix} 1 & & \\ & e^{i\alpha_{21}/2} & \\ & & e^{i\alpha_{31}/2} \end{pmatrix}$$

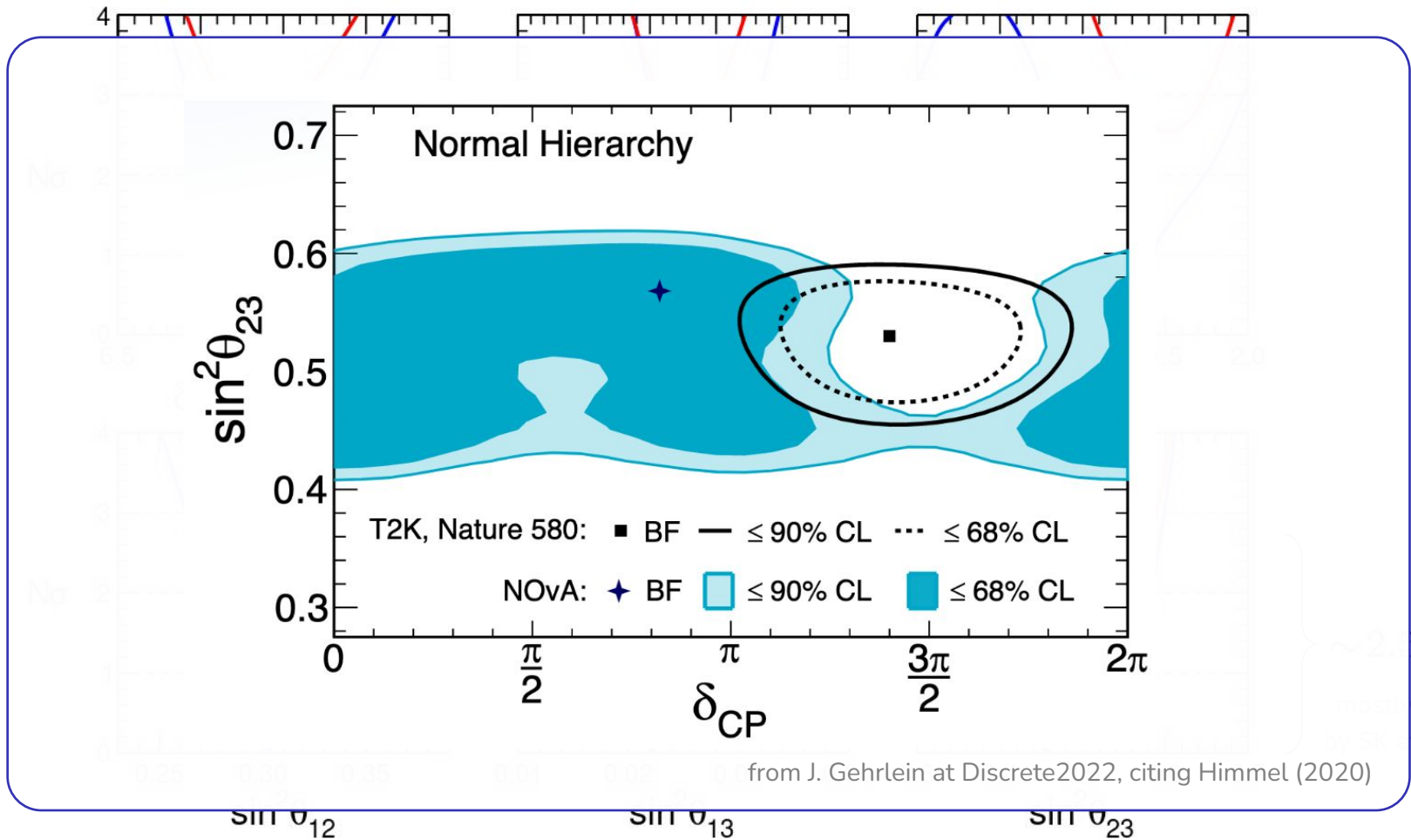
# 3ν flavour paradigm

from Capozzi et al. [2107.00532],  
 see also València [2006.11237], NuFIT [2007.14792]



# 3ν flavour paradigm

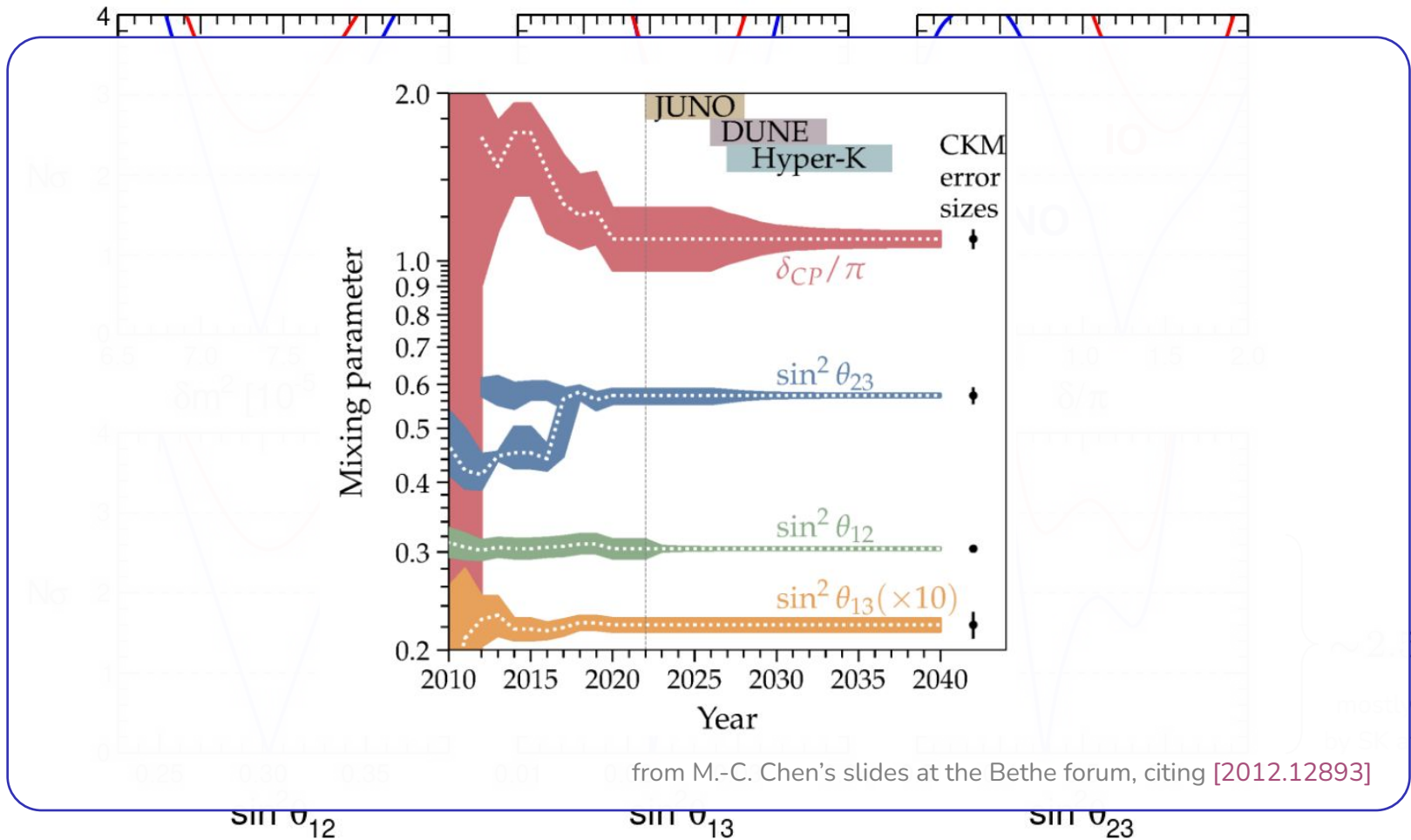
from Capozzi et al. [2107.00532],  
 see also València [2006.11237], NuFIT [2007.14792]



$\sim 2.5\sigma$   
 driven  
 tm. data

# 3ν flavour paradigm

from Capozzi et al. [2107.00532],  
see also València [2006.11237], NuFIT [2007.14792]



from M.-C. Chen's slides at the Bethe forum, citing [2012.12893]

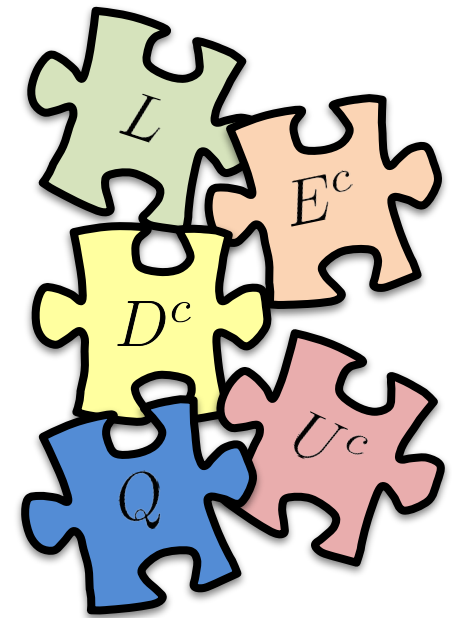
$\sim 2.5\sigma$   
most driven  
tm. data

# The flavour puzzle

$$U_{\text{PMNS}} \sim \begin{array}{c} e \\ \mu \\ \tau \end{array} \begin{array}{c} \nu_1 \\ \nu_2 \\ \nu_3 \end{array} \begin{bmatrix} \blacksquare & \blacksquare & \cdot \\ \blacksquare & \blacksquare & \blacksquare \\ \blacksquare & \blacksquare & \blacksquare \end{bmatrix}$$

$$V_{\text{CKM}} \sim \begin{array}{c} u \\ c \\ t \end{array} \begin{array}{c} d \\ s \\ b \end{array} \begin{bmatrix} \blacksquare & \cdot & \cdot \\ \cdot & \blacksquare & \cdot \\ \cdot & \cdot & \blacksquare \end{bmatrix}$$

Recall talk by G. Martinelli



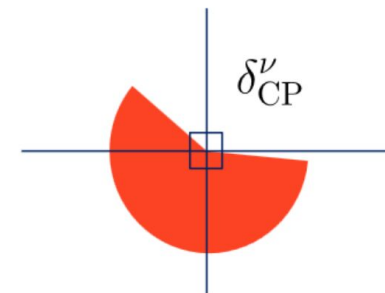
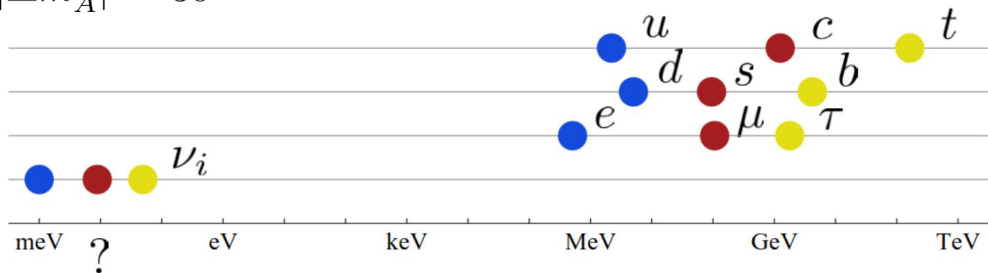
adapted from P. Novichkov's slides at PASCOS 2021



# Motivation

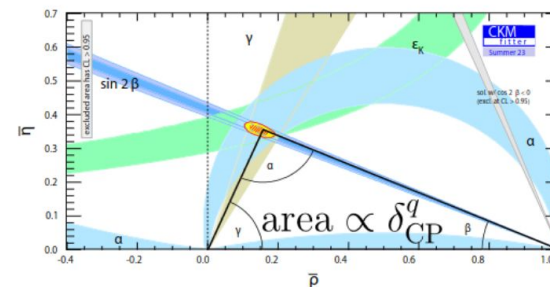
In search of an organising principle...

$$\frac{\Delta m_{\odot}^2}{|\Delta m_A^2|} \sim \frac{1}{30}$$



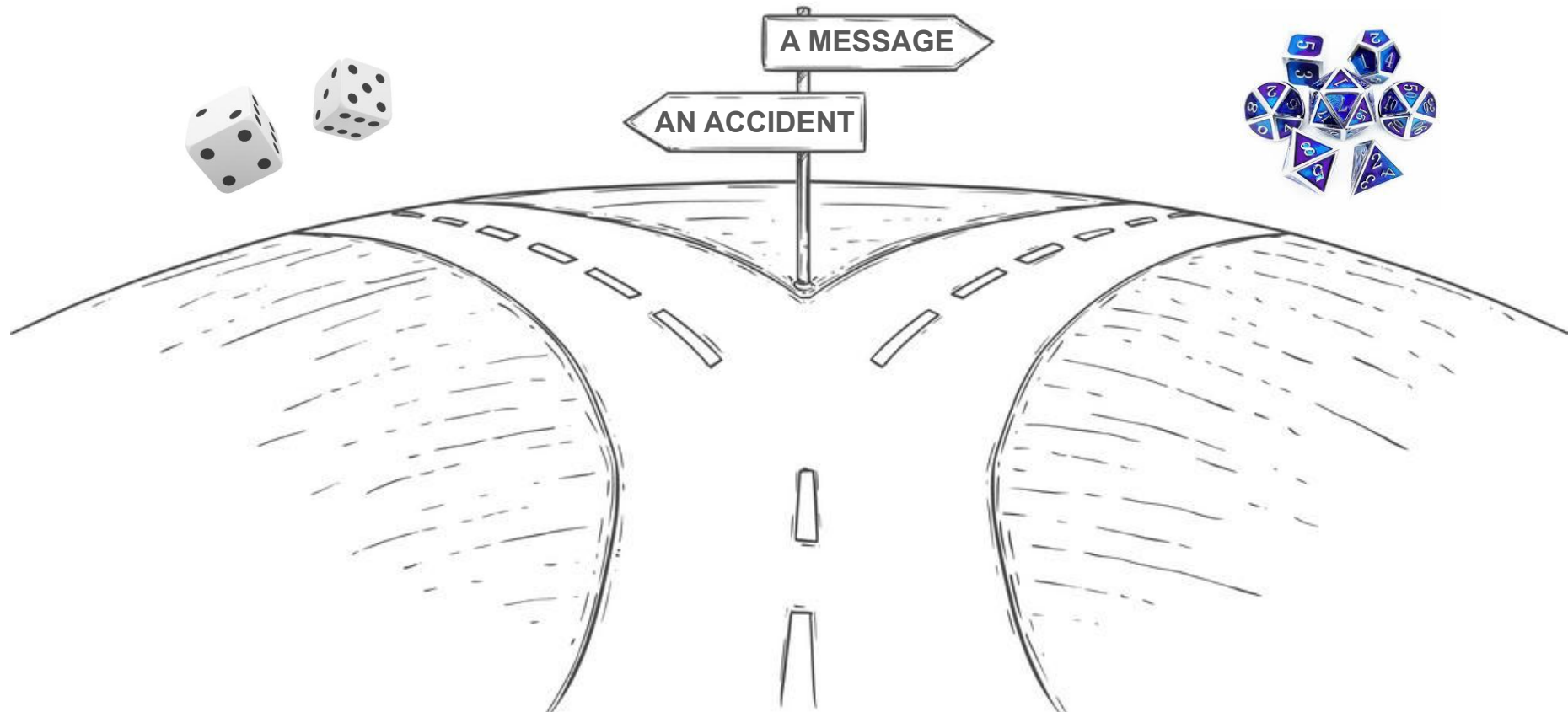
$$U_{PMNS} \sim \begin{bmatrix} \nu_1 & \nu_2 & \nu_3 \\ e & \blacksquare & \blacksquare & \blacksquare \\ \mu & \blacksquare & \blacksquare & \blacksquare \\ \tau & \blacksquare & \blacksquare & \blacksquare \end{bmatrix}$$

$$U_{CKM} \sim \begin{bmatrix} d & s & b \\ u & \blacksquare & \blacksquare & \blacksquare \\ c & \blacksquare & \blacksquare & \blacksquare \\ t & \blacksquare & \blacksquare & \blacksquare \end{bmatrix}$$

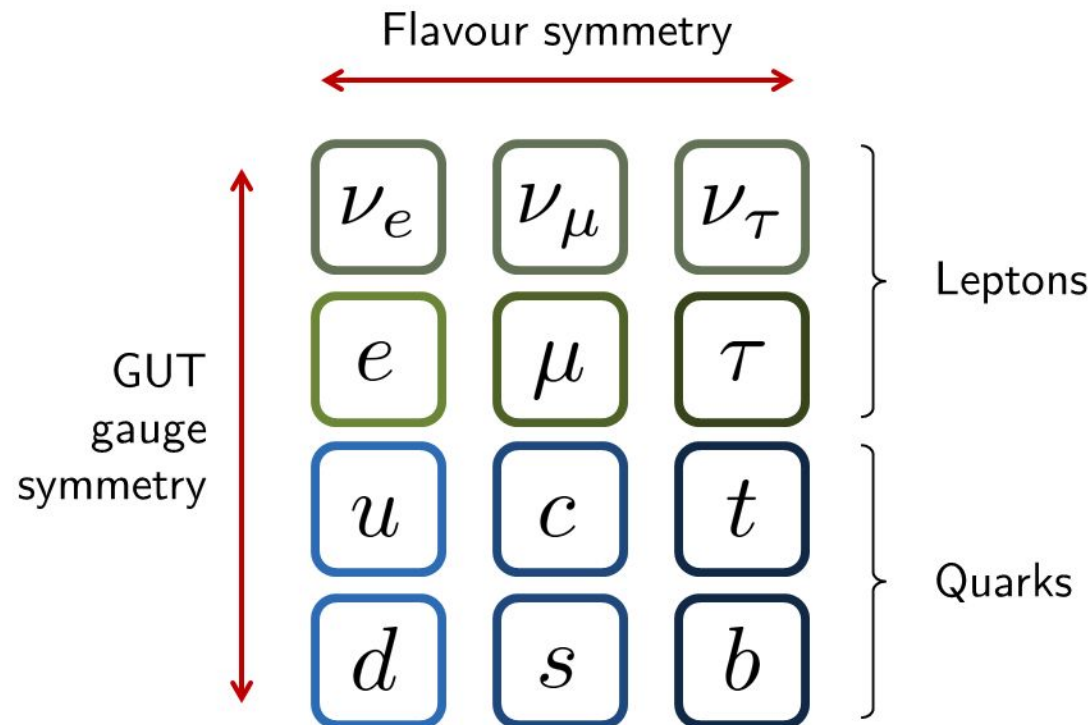


# Motivation

Is there an organising principle?



# Flavour symmetries

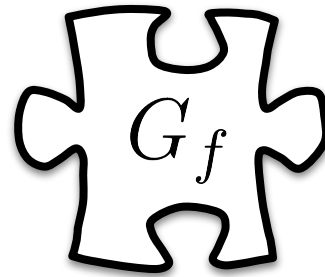


For reviews, see: Altarelli and Feruglio (2010), Ishimori et al. (2010), King and Luhn (2013), Petcov (2017), Feruglio and Romanino (2019), Ding and Valle (2024)

# Flavour symmetries

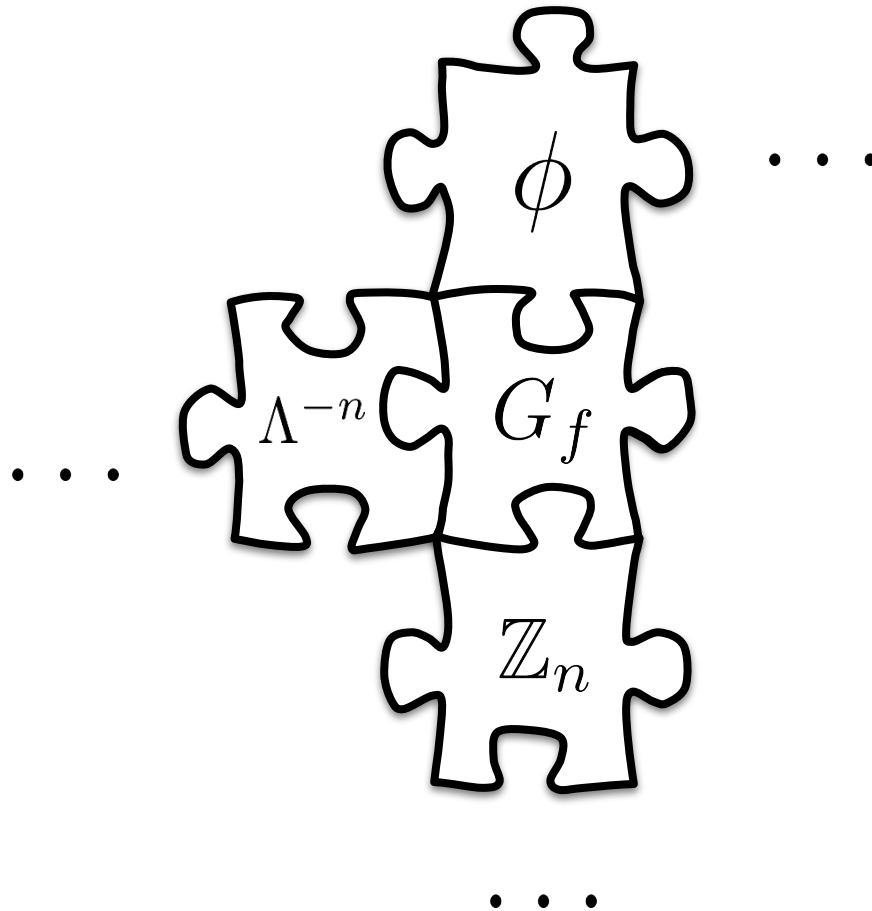


## Non-Abelian discrete flavour symmetries

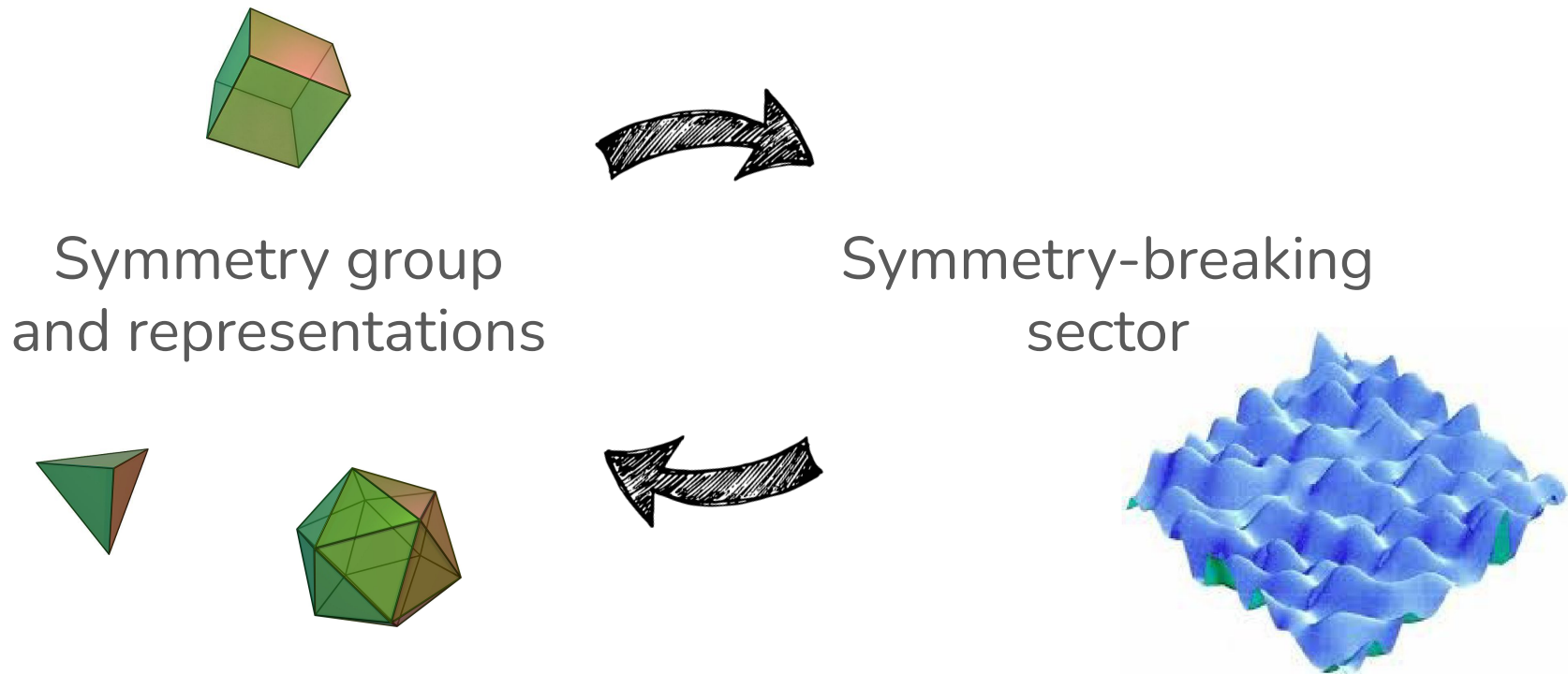


model-independent approaches relying on residual symmetries  
constrain mixing and the Dirac phase

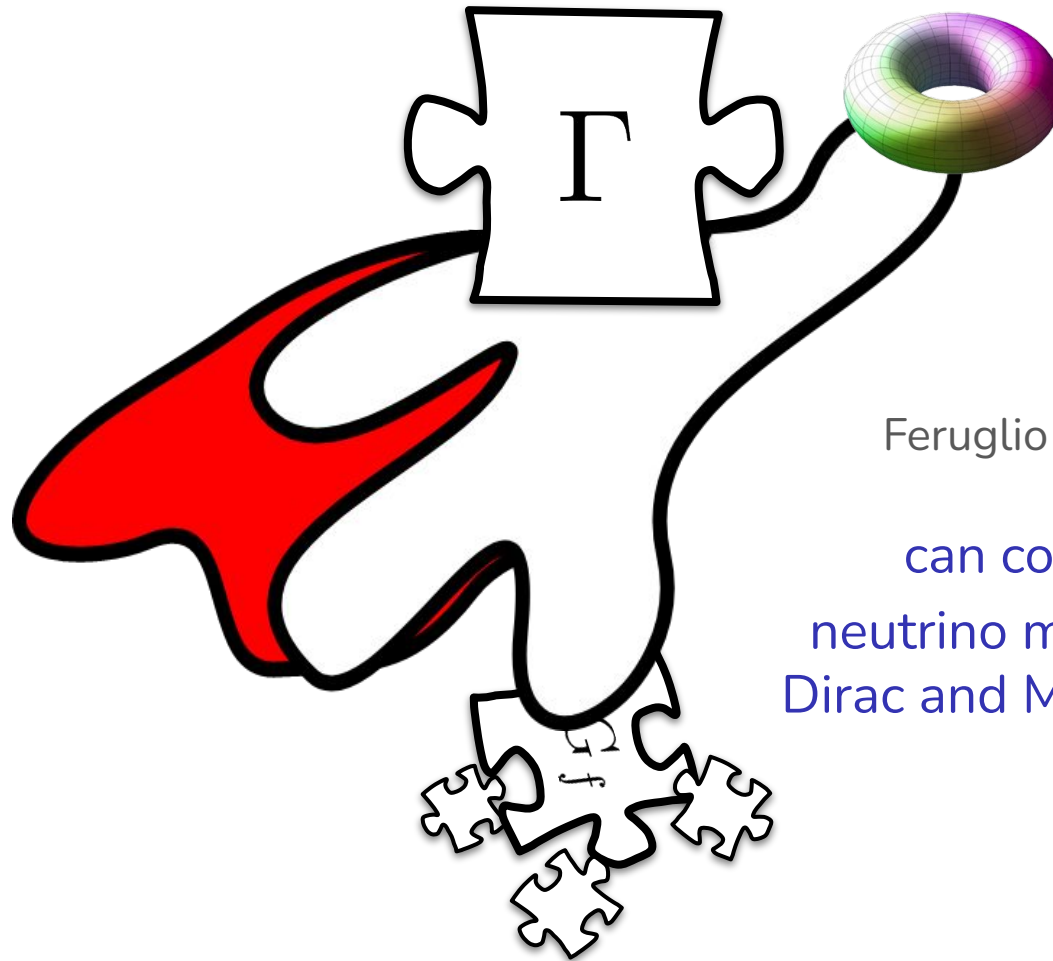
# Problems with the usual approach



# A reversal of the usual logic



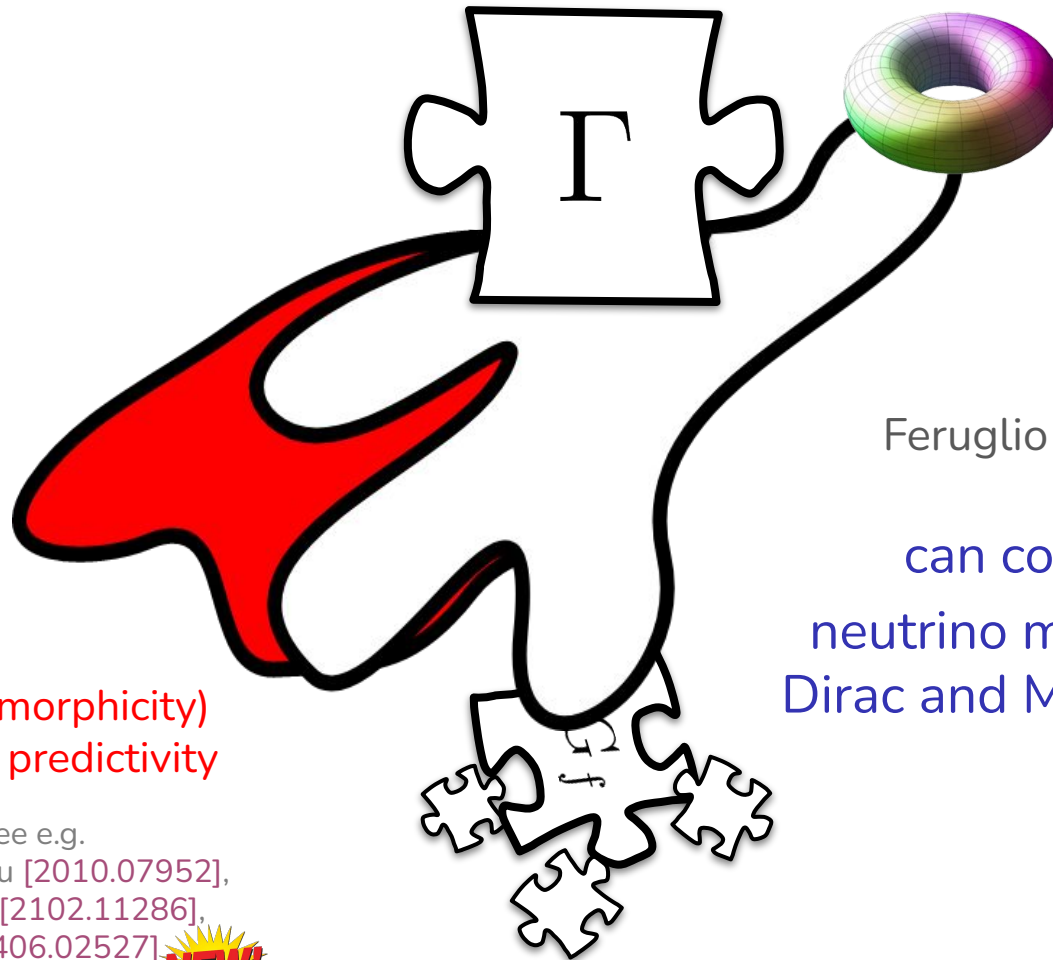
# Modular symmetry to the rescue!



Feruglio [1706.08749]

can constrain all:  
neutrino masses, mixing,  
Dirac and Majorana phases

# Modular symmetry to the rescue!



**SUSY** (holomorphicity)  
required for predictivity

...but see e.g.  
Ding, Feruglio, Liu [2010.07952],  
Almumin et al. [2102.11286],  
Qu, Ding [2406.02527]

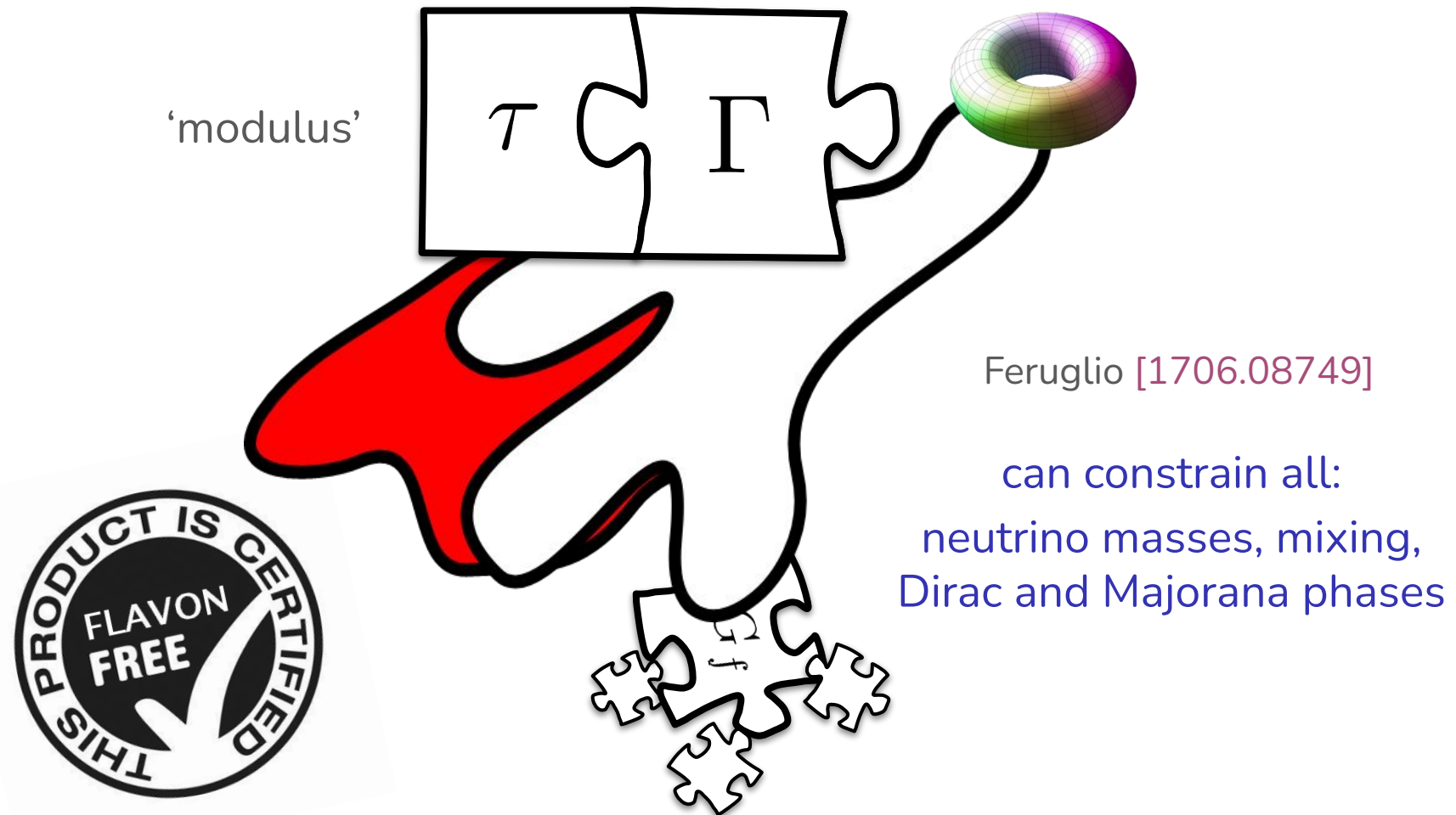
**NEW!**

Feruglio [1706.08749]

can constrain all:  
neutrino masses, mixing,  
Dirac and Majorana phases



# Modular symmetry to the rescue!



# Modular symmetry to the rescue!

'modulus'  $\tau$   $\Gamma$

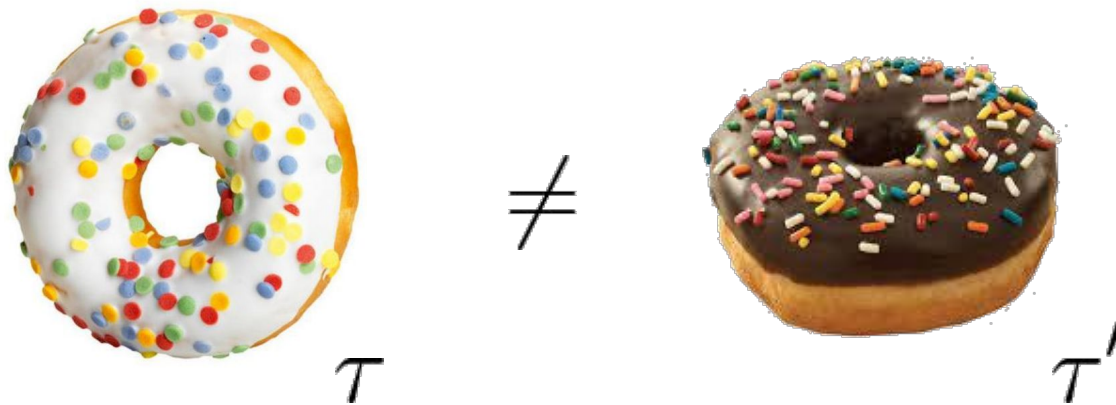
*the dream*

naturally correct:  
fermion masses, mixing,  
Dirac and Majorana phases



**The (bottom-up)  
framework**

# The modulus



$\tau$  may describe a torus compactification

(we assume only 1 unfrozen modulus)

In the **bottom-up** modular approach  $\tau$  is a dimensionless **spurion**

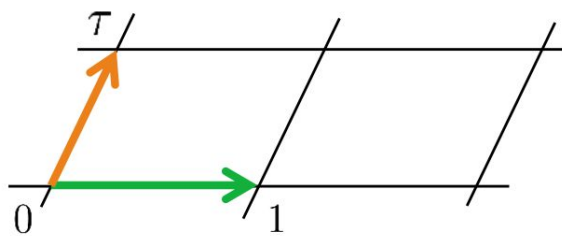
# The modulus



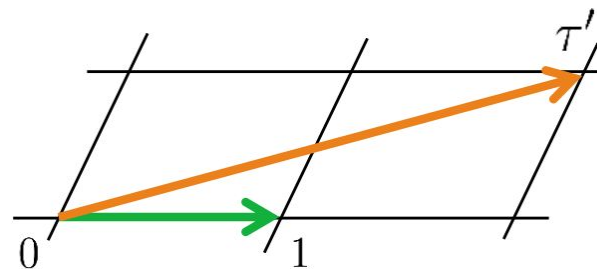
=



$$\tau' = \frac{a\tau + b}{c\tau + d}$$

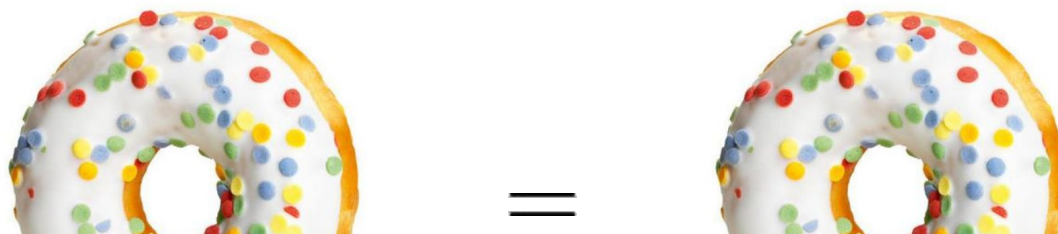


$$\text{Im } \tau > 0$$



$$ad - bc = 1 \quad a, b, c, d \in \mathbb{Z}$$

# The modulus



$$\gamma = \begin{pmatrix} a & b \\ c & d \end{pmatrix}$$

$$\tau' = \frac{a\tau + b}{c\tau + d}$$

$$\det \gamma = 1$$

$$a, b, c, d \in \mathbb{Z}$$

*the modular group*

$$\Gamma \equiv SL(2, \mathbb{Z}) = \{\gamma\}$$

# The modular group

$$\tau \rightarrow \frac{a\tau + b}{c\tau + d}$$

$$\Gamma \equiv SL(2, \mathbb{Z}) = \left\{ \gamma = \begin{pmatrix} a & b \\ c & d \end{pmatrix} \mid a, b, c, d \in \mathbb{Z}, \det \gamma = 1 \right\}$$

**Presentation in terms of generators S, T, R:**

$$S^2 = R, \quad (ST)^3 = R^2 = \mathbb{1}, \quad RT = TR$$

# The modular group

$$\tau \rightarrow \frac{a\tau + b}{c\tau + d}$$

$$\Gamma \equiv SL(2, \mathbb{Z}) = \left\{ \gamma = \begin{pmatrix} a & b \\ c & d \end{pmatrix} \mid a, b, c, d \in \mathbb{Z}, \det \gamma = 1 \right\}$$

$$S = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} :$$

$$\tau \rightarrow -1/\tau$$

**inverSion**

$$T = \begin{pmatrix} 1 & 1 \\ 0 & 1 \end{pmatrix} :$$

$$\tau \rightarrow \tau + 1$$

**Translation**

$$R = \begin{pmatrix} -1 & 0 \\ 0 & -1 \end{pmatrix} :$$

$$\tau \rightarrow \tau$$

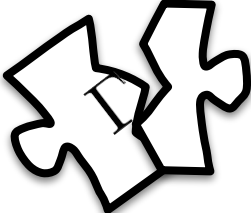
**Redundant**

but can affect fields...



# The modular group

$$\langle \tau \rangle \mapsto \frac{a\tau + b}{c\tau + d}$$



$$\equiv SL(2, \mathbb{Z}) = \left\{ \gamma = \begin{pmatrix} a & b \\ c & d \end{pmatrix} \mid a, b, c, d \in \mathbb{Z}, \det \gamma = 1 \right\}$$

$$S = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} :$$

$$\tau \rightarrow -1/\tau$$

**inversion**

$$T = \begin{pmatrix} 1 & 1 \\ 0 & 1 \end{pmatrix} :$$

$$\tau \rightarrow \tau + 1$$

**Translation**

$$R = \begin{pmatrix} -1 & 0 \\ 0 & -1 \end{pmatrix} :$$

$$\tau \rightarrow \tau$$

**Redundant**

but can affect fields...

## The field transformations

$$\psi \rightarrow (c\tau + d)^{-k} \rho(\gamma) \psi$$

# The field transformations

$$\psi \rightarrow \boxed{(c\tau + d)^{-k}} \rho(\gamma) \psi$$

Weight  $k \in \mathbb{Z}$

*NEW!* *automorphy factor*

# The field transformations

\* not necessarily:  
rare from top-down!

$$\psi \rightarrow \boxed{(c\tau + d)^{-k}} \rho(\gamma) \psi$$

Weight  $k \in \mathbb{Z}^*$

*automorphy factor*

**NEW!**

# The field transformations

\* not necessarily:  
rare from top-down!

$$\psi \rightarrow \overset{\text{NEW!}}{\underbrace{(c\tau + d)^{-k}}_{\text{Weight } k \in \mathbb{Z}^*}} \overset{\text{automorphy factor}}{\underbrace{\rho(\gamma)}} \psi$$

Weight  $k \in \mathbb{Z}^*$

“Almost trivial”  
representation of  
the modular group

$$\rho(\Gamma(N)) = \mathbb{1}$$

$$\rho(T\Gamma(N)) = \rho(T)$$

$$\rho(S\Gamma(N)) = \rho(S)$$

... Feruglio [1706.08749]

\* not necessarily:  
rare from top-down!

# The field transformations

$$\psi \rightarrow \overset{\text{NEW!}}{\text{automorphy factor}} \left( c\tau + d \right)^{-k} \rho(\gamma) \psi$$

Weight  $k \in \mathbb{Z}^*$

“Almost trivial”  
representation of  
the modular group

$$\Gamma(N) \subset \Gamma$$

Principal congruence subgroup of level  $N$

$$\Gamma(N) \equiv \left\{ \begin{pmatrix} a & b \\ c & d \end{pmatrix} \equiv \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \pmod{N} \right\}$$

$$\rho(\Gamma(N)) = \mathbb{1}$$

$$\rho(T\Gamma(N)) = \rho(T)$$

$$\rho(S\Gamma(N)) = \rho(S)$$

... Feruglio [1706.08749]

$$\rho(\gamma) \text{ is effectively a representation of } \Gamma'_N \equiv \Gamma/\Gamma(N)$$

other choices are possible: in general, *vector-valued modular forms*, see e.g. [2112.14761, 2311.10136]

# The finite modular groups

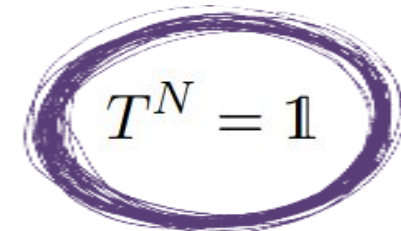
$\Gamma'_N \equiv \Gamma/\Gamma(N)$  *behave like flavour groups*

$N$	2	3	4	5
$\Gamma_N$	$S_3$	$A_4$	$S_4$	$A_5$
$\Gamma'_N$	$S_3$	$A'_4 \equiv T'$	$S'_4 \equiv SL(2, \mathbb{Z}_4)$	$A'_5 \equiv SL(2, \mathbb{Z}_5)$

← drop the **R**  
generator  
(in general there in TD)

Presentation in terms of generators **S, T, R**:

$$S^2 = R, \quad (ST)^3 = R^2 = \mathbb{1}, \quad RT = TR, \quad T^N = \mathbb{1}$$



# The finite modular groups

$$\Gamma'_N \equiv \Gamma/\Gamma(N) \text{ behave like flavour groups}$$

$N$	2	3	4	5
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← drop the  $\mathbf{R}$   
generator

(in general there in TD)

$$\Gamma_2 \simeq S_3$$

Kobayashi et al. [1803.10391]  
Meloni, Parriciatu [2306.09028]

$$\Gamma_3 \simeq A_4$$

Feruglio [1706.08749]

$$\Gamma_4 \simeq S_4$$

JP, Petcov [1806.11040]

$$\Gamma_5 \simeq A_5$$

Novichkov et al. [1812.02158]

summary in Appendices of  
Novichkov, JP, Petcov, Titov  
[1905.11970]

$$\Gamma'_3 \simeq A'_4$$

Liu, Ding [1907.01488]

$$\Gamma'_4 \simeq S'_4$$

Novichkov, JP, Petcov [2006.03058]

$$\Gamma'_5 \simeq A'_5$$

Wang, Yu, Zhou [2010.10159]

For early top-down, see e.g.:

Kobayashi et al. [1804.06644];  
Kobayashi, Tamba [1811.11384];  
de Anda et al. [1812.05620];  
Baur et al. [1901.03251, 1908.00805];  
Kariyazono et al. [1904.07546];  
Nilles et al. [2001.01736, 2004.05200,  
2006.03059]; Kobayashi, Otsuka  
[2001.07972, 2004.04518];  
Abe et al. [2003.03512];  
Ohki et al. [2003.04174];  
Kikuchi et al. [2005.12642]

by now, O(200) papers...



# A lot of model building...



- models based on finite modular groups of higher  $N$
- modular models of unification (also without GUTs)
- modular models of leptogenesis
- models with multiple moduli

based on symplectic modular invariance (**Siegel modular group**)  
and automorphic forms

$$\tau \rightarrow \begin{pmatrix} \tau_1 & \tau_3 \\ \tau_3 & \tau_2 \end{pmatrix}$$

- models relating modular flavour symmetries and inflation
- models exploring the interplay of modular and gCP symmetries

## ...and a vast literature...



- models based on finite modular groups of higher  $N$

[2004.12662, 2108.02181, 2307.01419]

- modular models of unification (also without GUTs)

[1906.10341, 2012.01397, 2101.02266, 2101.12724, 2103.02633, 2103.16311, 2108.09655, 2206.14675, 2312.09255]

- modular models of leptogenesis

[1909.06520, 2007.00545, 2103.07207, 2201.10429, 2204.08338, 2205.08269, 2206.14675, 2402.18547, 2405.09363]

- models with multiple moduli

[1811.04933, 1812.11289, 1906.02208, 1908.02770, 2304.05958]

based on symplectic modular invariance (**Siegel modular group**)

and automorphic forms: Ding, Feruglio, Liu [2010.07952, 2402.14915]

From TD, see e.g.: Nilles et al. [2105.08078], Baur et al. [2012.09586], Kikuchi et al. [2305.16709]

$$\tau \rightarrow \begin{pmatrix} \tau_1 & \tau_3 \\ \tau_3 & \tau_2 \end{pmatrix}$$

- models relating modular flavour symmetries and inflation Recall talk by X. Wang

[2208.10086, 2303.02947, 2405.06497, 2405.08924]

- models exploring the interplay of modular and gCP symmetries

[1901.03251, 1905.11970, 1910.11553, 2006.03058, 2012.01688, 2012.13390, 2102.06716, 2106.11659]

But how does it work?

$$\psi \sim (\mathbf{r}, k)$$

$$W \sim g(\psi_1 \dots \psi_n) \mathbf{1}$$

$$\psi \rightarrow (c\tau + d)^{-k} \rho_{\mathbf{r}}(\gamma) \psi$$

Need modular forms

$$\psi \sim (\mathbf{r}, k)$$

$$W \sim g(Y(\tau) \psi_1 \dots \psi_n) \mathbf{1}$$

$$\psi \rightarrow (c\tau + d)^{-k} \rho_{\mathbf{r}}(\gamma) \psi$$

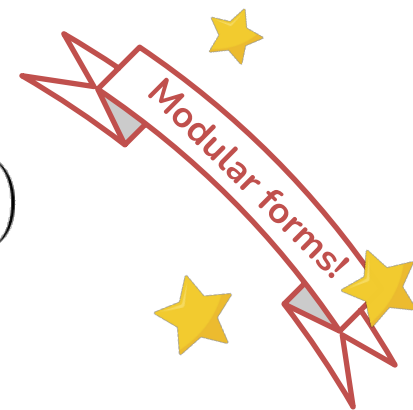
Need modular forms

$$\psi \sim (\mathbf{r}, k)$$

$$W \sim g(Y(\tau) \psi_1 \dots \psi_n) \mathbf{1}$$

$$\psi \rightarrow (c\tau + d)^{-k} \rho_{\mathbf{r}}(\gamma) \psi$$

$$Y(\tau) \rightarrow (c\tau + d)^{k_Y} \rho_Y(\gamma) Y(\tau)$$



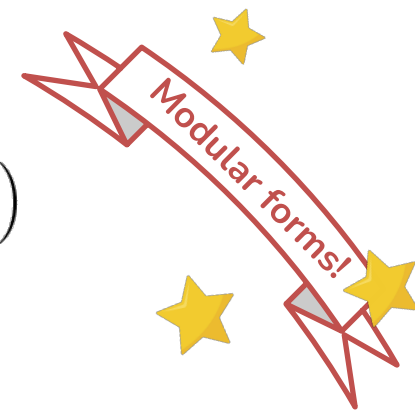
Need modular forms

$$\psi \sim (\mathbf{r}, k)$$

$$W \sim g(Y(\tau) \psi_1 \dots \psi_n) \mathbf{1}$$

$$\begin{aligned} \psi &\rightarrow (c\tau + d)^{-k} \rho_{\mathbf{r}}(\gamma) \psi \\ Y(\tau) &\rightarrow (c\tau + d)^{k_Y} \rho_Y(\gamma) Y(\tau) \end{aligned}$$

$$\begin{cases} k_Y = k_1 + \dots + k_n \\ \rho_Y \otimes \rho_1 \otimes \dots \otimes \rho_n \supset \mathbf{1} \end{cases}$$



Need modular forms

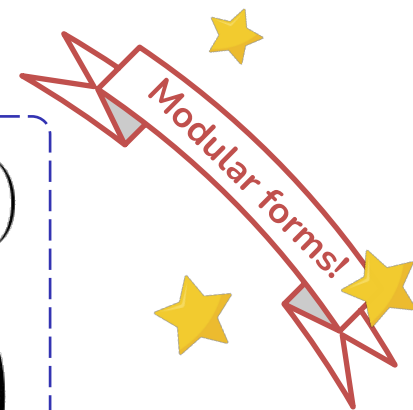
$$\psi \sim (\mathbf{r}, k)$$

$$W \sim g(Y(\tau) \psi_1 \dots \psi_n) \mathbf{1}$$

$$\psi \rightarrow (c\tau + d)^{-k} \rho_{\mathbf{r}}(\gamma) \psi$$

$$Y(\tau) \rightarrow (c\tau + d)^{k_Y} \rho_Y(\gamma) Y(\tau)$$

$$= Y\left(\frac{a\tau + b}{c\tau + d}\right)$$



# The modular forms

$N$	2	3	4	5
$\Gamma_N$	$S_3$	$A_4$	$S_4$	$A_5$
$\Gamma'_N$	$S_3$	$A'_4 \equiv T'$	$S'_4 \equiv SL(2, \mathbb{Z}_4)$	$A'_5 \equiv SL(2, \mathbb{Z}_5)$
$\dim \mathcal{M}_k(\Gamma(N))$	$k/2 + 1$	$k + 1$	$2k + 1$	$5k + 1$

*Not so many available!*

A **finite set** of  
functions for each  $k$



# The modular forms

$N$	2	3	4	5
$\Gamma_N$	$S_3$	$A_4$	$S_4$	$A_5$
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*Not so many available!*

**A finite set of functions for each  $k$**

**Lowest-weight  $k$  modular forms for each group:**

$$\Gamma_N^{(\prime)} \quad Y_{\mathbf{r}}^{(k)}$$

$$\Gamma_2 \simeq S_3 \quad Y_2^{(2)}$$

$$\Gamma'_3 \simeq A'_4 \quad Y_{\hat{2}}^{(1)}$$

$$\Gamma_3 \simeq A_4 \quad Y_3^{(2)}$$

$$\Gamma'_4 \simeq S'_4 \quad Y_{\hat{3}}^{(1)}$$

$$\Gamma_4 \simeq S_4 \quad Y_2^{(2)}, Y_{3'}^{(2)}$$

$$\Gamma'_5 \simeq A'_5 \quad Y_{\hat{6}}^{(1)}$$

$$\Gamma_5 \simeq A_5 \quad Y_3^{(2)}, Y_{3'}^{(2)}, Y_5^{(2)}$$

**non-singular**, unlike modular functions. Can still have an interpretation, see Feruglio, Strumia, Titov [2305.08908]

**can generalize** modular group to e.g. the larger metaplectic group and get half-integer weight forms, see Liu et al. [2007.13706]

# Example

Let's build a modular-invariant term!

$$W \supset NN$$

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$$\Gamma_3 \simeq A_4$$

$$N \sim (\mathbf{3}, 1)$$

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$$W \supset NN$$

Let's build a modular-invariant term!

$$\Gamma_3 \simeq A_4$$

$$N \sim (\mathbf{3}, 1)$$

$$W \supset \Lambda \left( N \otimes N \otimes Y_{\mathbf{3}}^{(2)} \right)_{\mathbf{1}}$$

# Example

$$W \supset NN$$

Let's build a modular-invariant term!

$$\Gamma_3 \simeq A_4$$

$$N \sim (\mathbf{3}, 1)$$

$$W \supset \Lambda \left( N \otimes N \otimes Y_{\mathbf{3}}^{(2)} \right)_{\mathbf{1}}$$



$$Y_{\mathbf{3}}^{(2)}(\tau) = \begin{pmatrix} Y_1(\tau) \\ Y_2(\tau) \\ Y_3(\tau) \end{pmatrix}$$

$$M_N = \Lambda \begin{pmatrix} 2Y_1(\tau) & -Y_3(\tau) & -Y_2(\tau) \\ -Y_3(\tau) & 2Y_2(\tau) & -Y_1(\tau) \\ -Y_2(\tau) & -Y_1(\tau) & 2Y_3(\tau) \end{pmatrix}$$

# Example

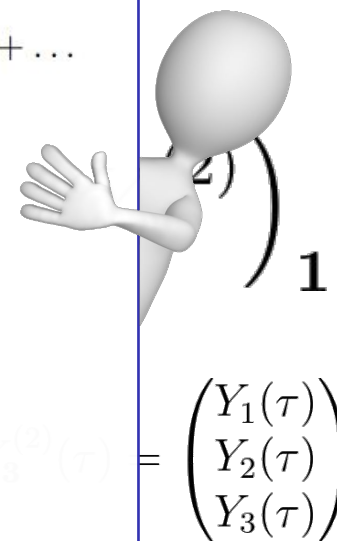
$$W \supset NN$$

let's build a modular-invariant term!

$$Y_1(\tau) = \frac{i}{2\pi} \left[ \frac{\eta'(\frac{\tau}{3})}{\eta(\frac{\tau}{3})} + \frac{\eta'(\frac{\tau+1}{3})}{\eta(\frac{\tau+1}{3})} + \frac{\eta'(\frac{\tau+2}{3})}{\eta(\frac{\tau+2}{3})} - \frac{27\eta'(3\tau)}{\eta(3\tau)} \right] = 1 + 12q + 36q^2 + 12q^3 + \dots$$

$$Y_2(\tau) = \frac{-i}{\pi} \left[ \frac{\eta'(\frac{\tau}{3})}{\eta(\frac{\tau}{3})} + \omega^2 \frac{\eta'(\frac{\tau+1}{3})}{\eta(\frac{\tau+1}{3})} + \omega \frac{\eta'(\frac{\tau+2}{3})}{\eta(\frac{\tau+2}{3})} \right] = -6q^{1/3}(1 + 7q + 8q^2 + \dots)$$

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with

$$\omega = e^{2\pi i/3} \quad \eta(\tau) = q^{1/24} \prod_{n=1}^{\infty} (1 - q^n) \quad q \equiv e^{i2\pi\tau}$$

obey  $Y_2^2 + 2Y_1Y_3 = 0$

Feruglio [1706.08749]

$$MN = \Lambda \begin{pmatrix} -Y_3(\tau) & 2Y_2(\tau) & -Y_1(\tau) \\ -Y_2(\tau) & -Y_1(\tau) & 2Y_3(\tau) \end{pmatrix}$$

# Example

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so now we can build models...

# Example: an $S_4$ lepton model

Novichkov, JP, Petcov, Titov [1811.04933]

**Ingredients:** Choose group, field content

$$\psi \sim (\mathbf{r}, k)$$



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$$N^c \sim (\mathbf{3}', 0), \quad L \sim (\mathbf{3}, 2)$$

$$E^c \sim (\mathbf{1}', 0) \oplus (\mathbf{1}, 2) \oplus (\mathbf{1}', 2)$$

**Ingredients:** Choose group, field content

$$\psi \sim (\mathbf{r}, k)$$

$$W = \alpha \left( E_1^c L Y_{\mathbf{3}'}^{(2)} \right)_1 H_d + \beta \left( E_2^c L Y_{\mathbf{3}}^{(4)} \right)_1 H_d + \gamma \left( E_3^c L Y_{\mathbf{3}'}^{(4)} \right)_1 H_d$$

$$+ g \left( N^c L Y_{\mathbf{2}}^{(2)} \right)_1 H_u + \underbrace{g'}_{\in \mathbb{C} \text{ only physical phase}} \left( N^c L Y_{\mathbf{3}'}^{(2)} \right)_1 H_u + \Lambda (N^c N^c)_1,$$

**Procedure:** Fit couplings and  $\tau$       $\min \chi^2(\tau, g'/g, g^2/\Lambda, \alpha, \beta, \gamma)$

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**Procedure:** Fit couplings and  $\tau$   $\min \chi^2(\tau, g'/g, g^2/\Lambda, \alpha, \beta, \gamma)$

$$g\text{CP} \Rightarrow g' \in \mathbb{R}$$

Novichkov, JP, Petcov, Titov [1905.11970]

$\tau$  can be the only source of CPV

# Example: an $S_4$ lepton model (results)

Novichkov, JP, Petcov, Titov  
[1811.04933, 1905.11970]

$$\sin^2 \theta_{23} \sim 0.49$$

$$\delta \sim 1.6\pi$$

$$\alpha_{21} \sim 0.3\pi$$

$$\alpha_{31} \sim 1.3\pi$$

$$|\langle m \rangle|_{\beta\beta} \sim 12 \text{ meV}$$

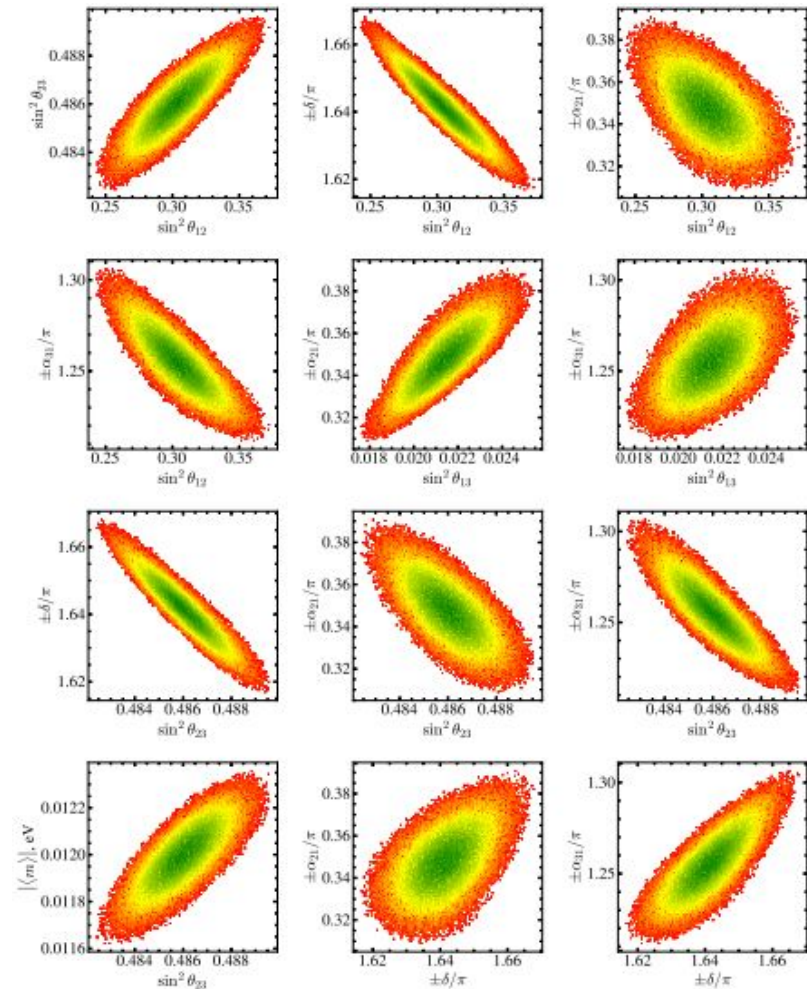
$$\sum_i m_i \sim 0.08 \text{ eV}$$

**7 (4) parameters**

**vs.**

**12 (9) observables**

Minimal model found with one less parameter:  
Ding, Liu, Yao [2211.04546]



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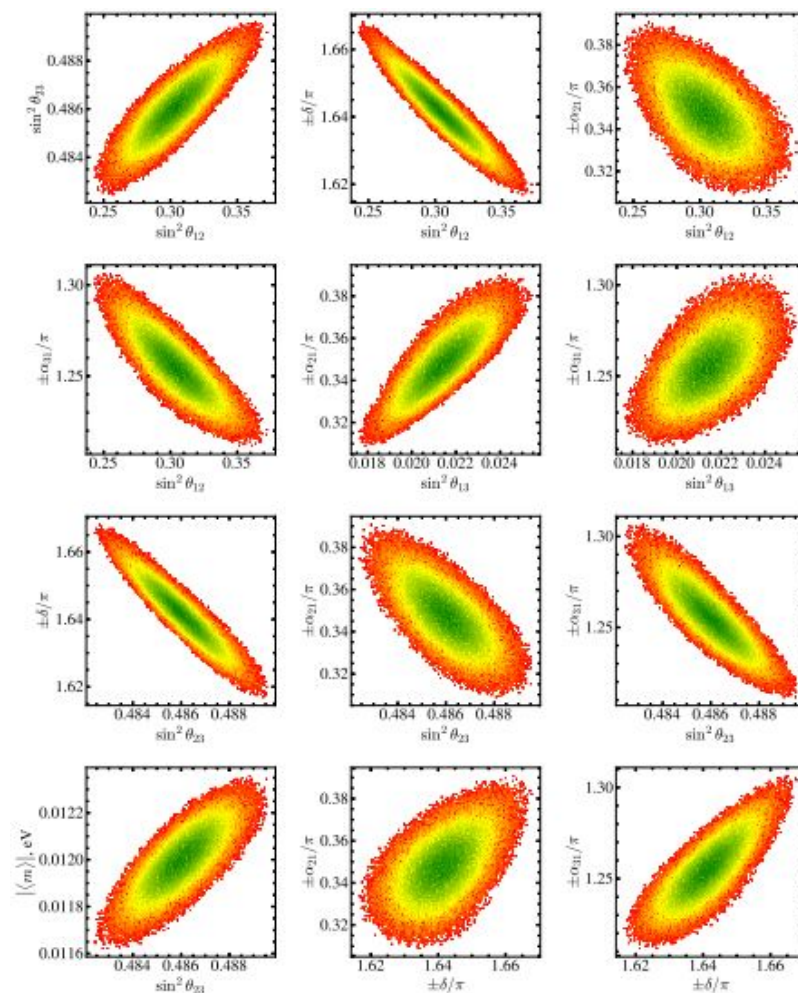
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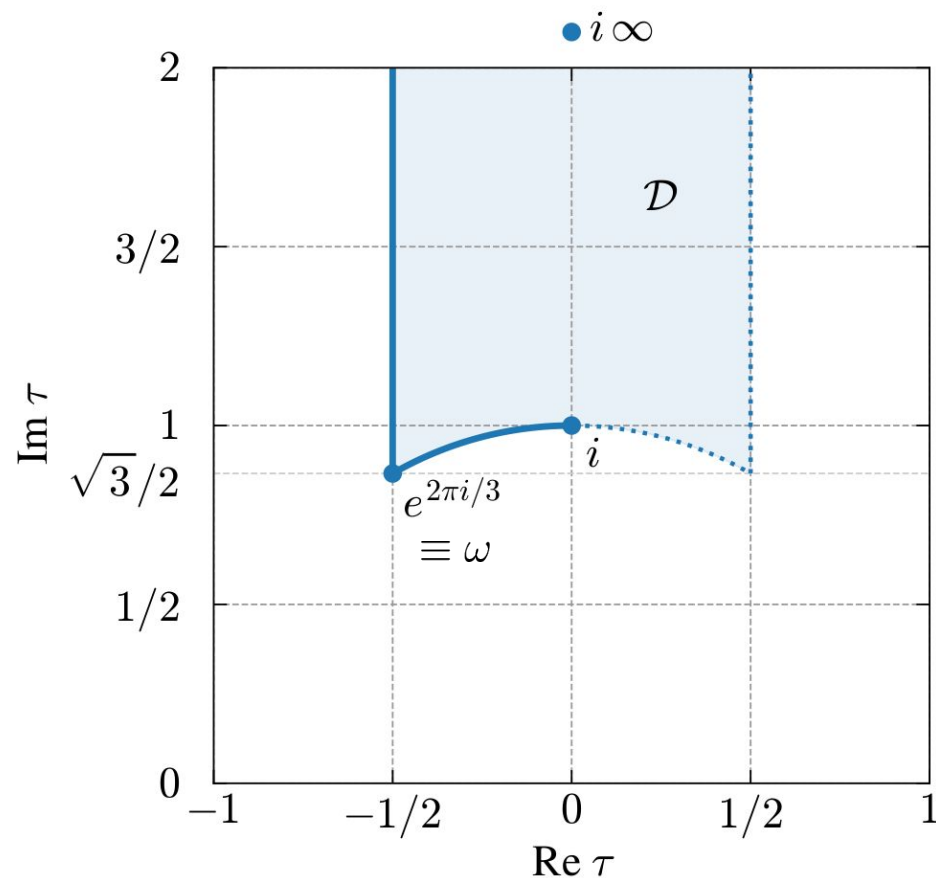
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Minimal model found with one less parameter:  
Ding, Liu, Yao [2211.04546]



...and one does not need to consider the whole  $\frac{1}{2}$  plane

# The fundamental domain



- **Any  $\tau$**  breaks the full modular symmetry
- To **fit** a model which is invariant under the full modular group, it is enough to scan  $\tau$  in the **fundamental domain**

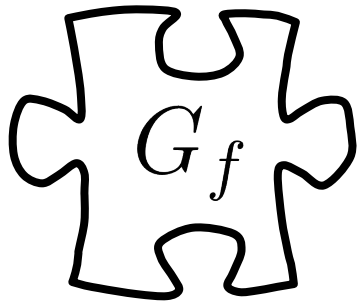
In some cases, can avoid fit by looking at invariants  
see Chen et al. [\[2211.04546\]](#)

...and one does not need to consider the whole  $\frac{1}{2}$  plane

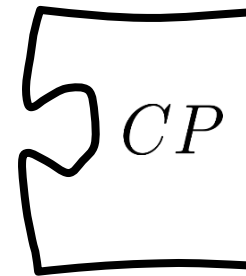


**CP violation**

# Flavour symmetries + gCP (generalized CP)



$$\psi(x) \rightarrow \rho_{\mathbf{r}}(g) \psi(x)$$



$$\psi(x) \rightarrow X_{\mathbf{r}}^{\text{CP}} \overline{\psi}(x_{\text{P}})$$

Branco, Lavoura, Rebelo (1986)

Harrison, Scott (2002)

Grimus, Lavoura (2003)

Farzan, Smirnov (2006)

Ferreira et al. (2012)

...

# Flavour symmetries + gCP (generalized CP)

$$\psi(x) \rightarrow \rho_{\mathbf{r}}(g) \psi(x) \quad \begin{array}{c} \text{Puzzle Piece} \\ G_f \quad CP \end{array} \quad \psi(x) \rightarrow X_{\mathbf{r}}^{\text{CP}} \bar{\psi}(x_P)$$

can constrain mixing, the Dirac **and the Majorana phases**

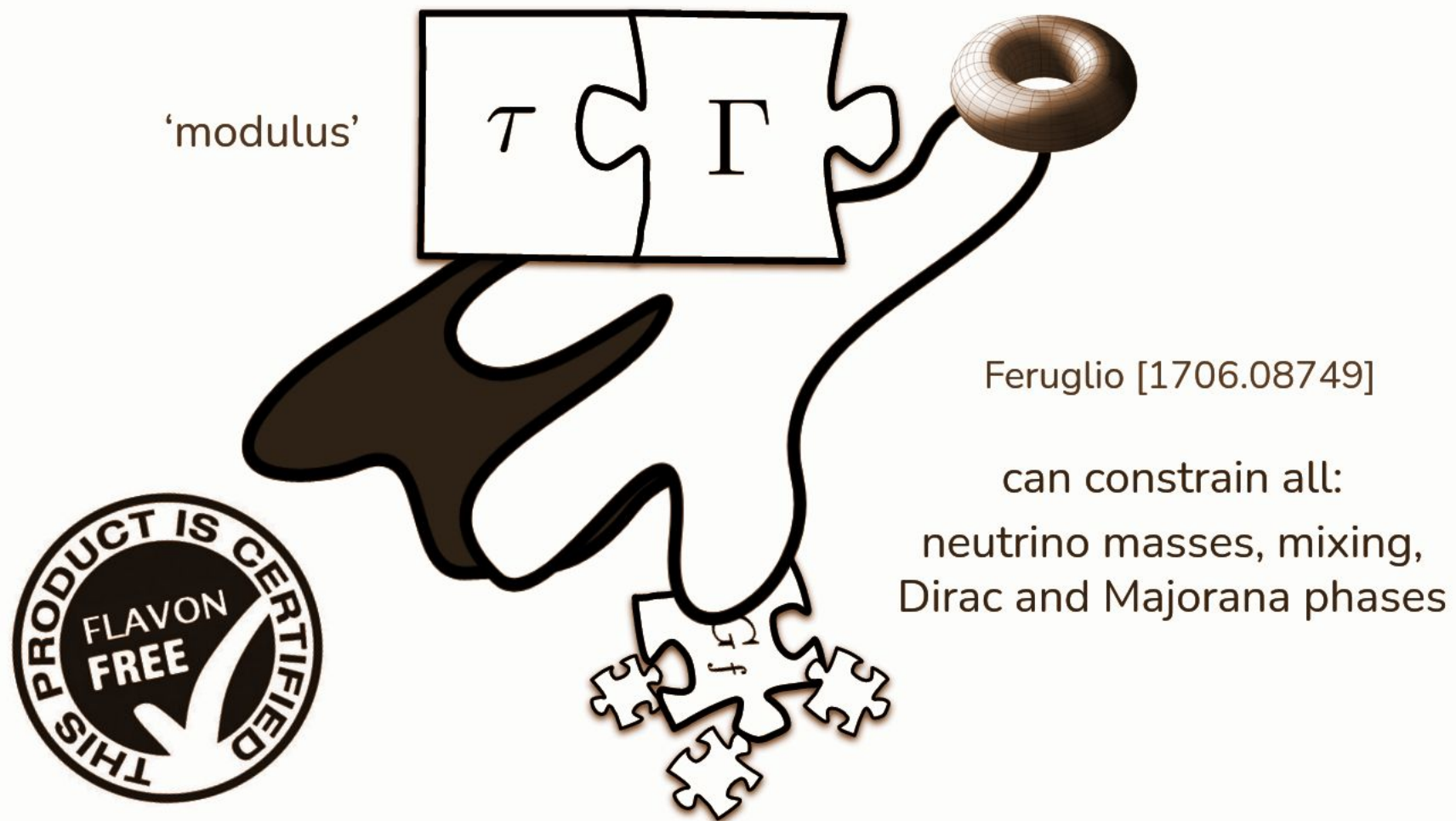
Consistency condition [Feruglio et al. (2012), Holthausen et al. (2012)]

$$X_{\mathbf{r}}^{\text{CP}} \rho_{\mathbf{r}}^*(g) \left(X_{\mathbf{r}}^{\text{CP}}\right)^{-1} = \rho_{\mathbf{r}}(u(g))$$

$u$  is a class-inverting outer automorphism [Chen et al. (2014)]

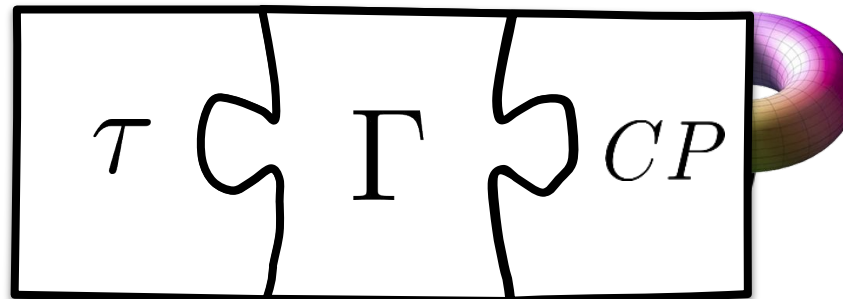


# Modular symmetry to the rescue!



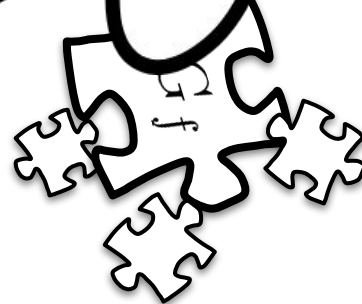
# Modular symmetry + gCP

'modulus'

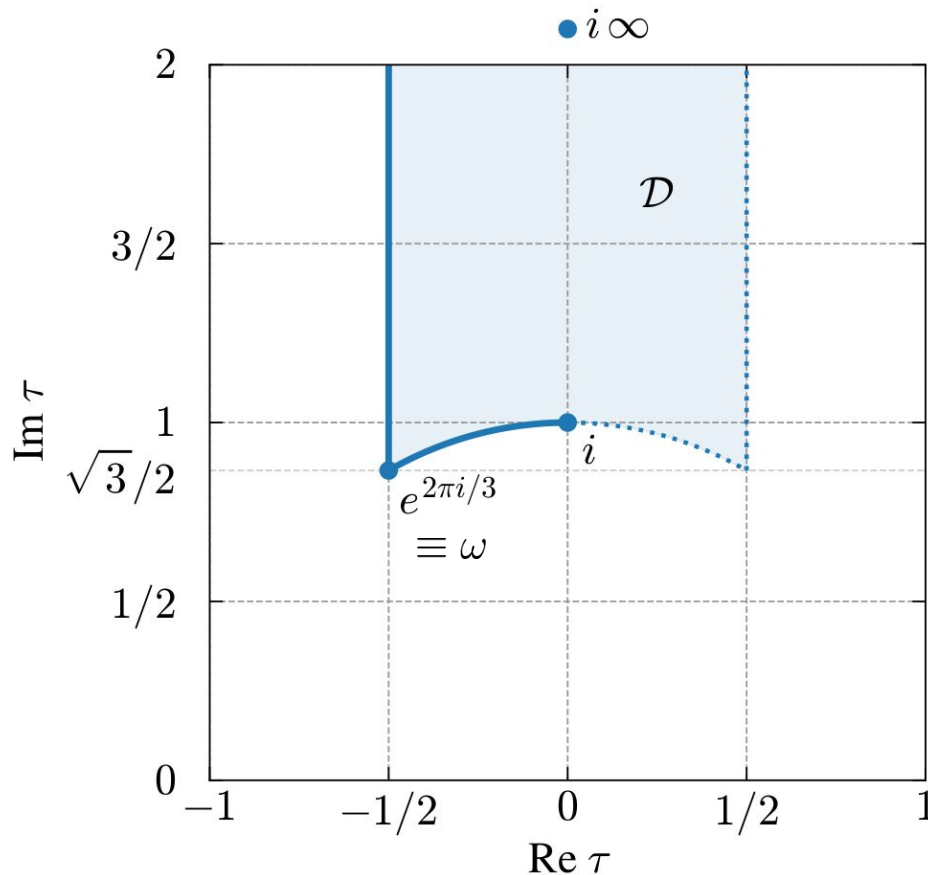


Feruglio [1706.08749]

can constrain all:  
neutrino masses, mixing,  
Dirac and Majorana phases



# Modular symmetry + gCP (back to the fundamental domain)



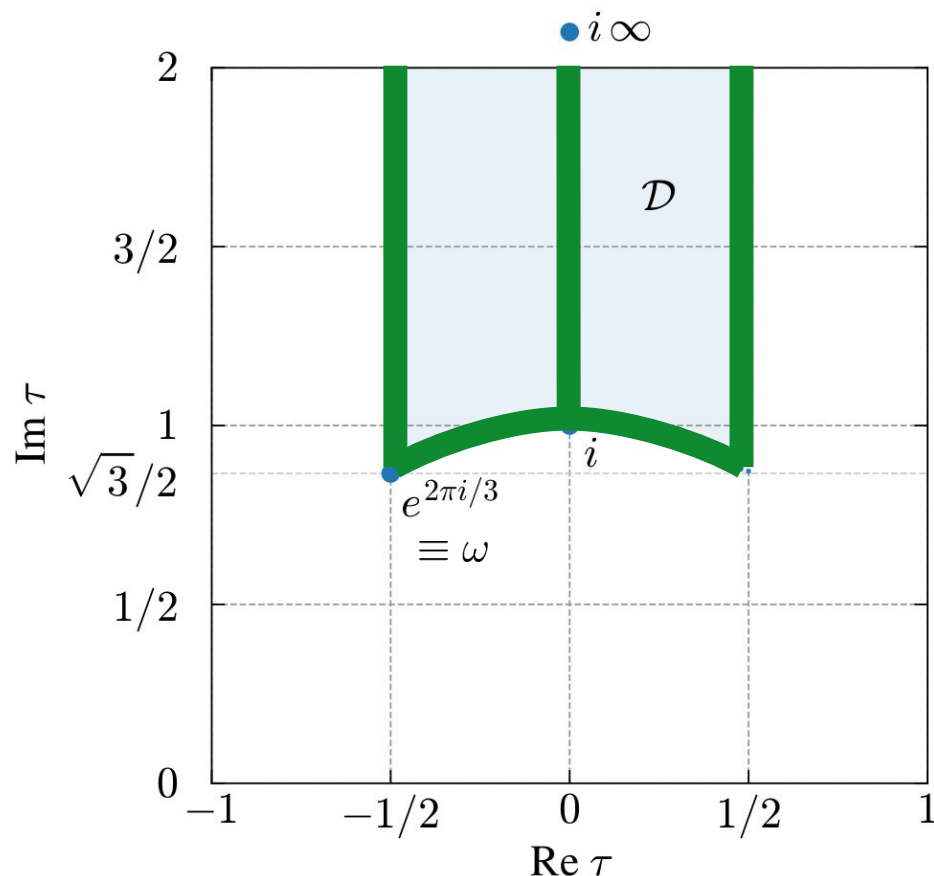
- **Any**  $\tau$  breaks the full modular symmetry
- Special values of  $\tau$  preserve the CP symmetry
- The modulus can be the **only source of CP violation!**  
(recall the  $S_4$  model of slide 50...)

- CP is violated by the modulus unless

$$-\tau^* = \gamma\tau$$

special regions of the fundamental domain

# Modular symmetry + gCP (back to the fundamental domain)



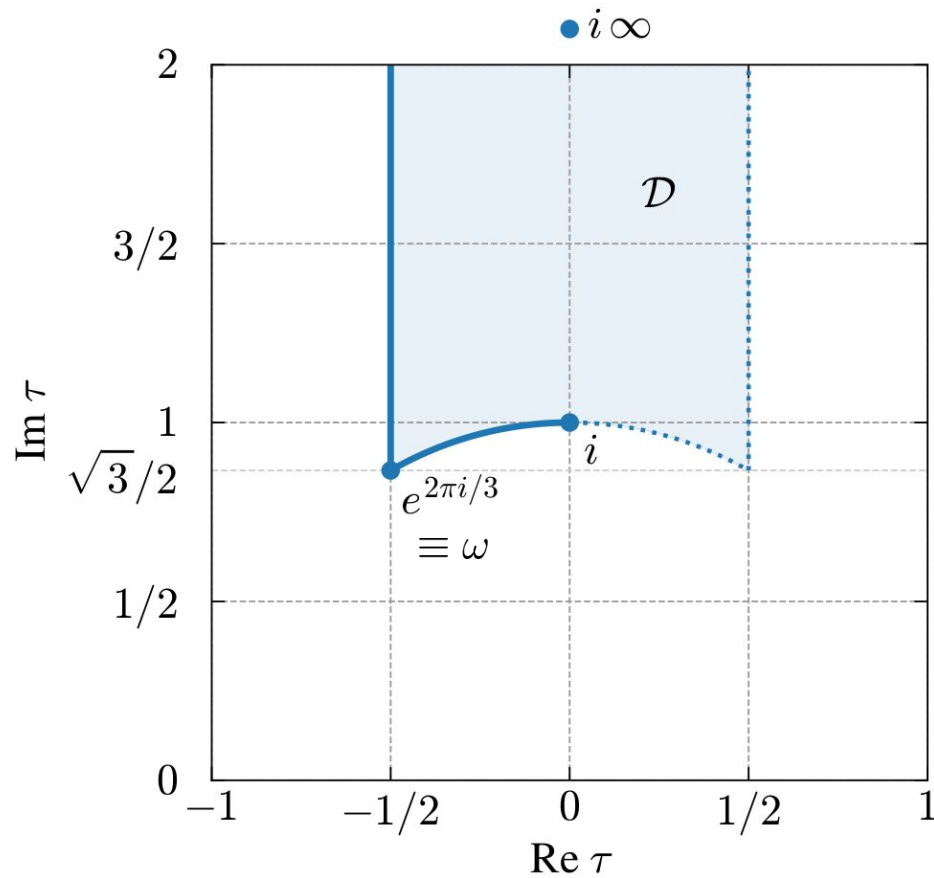
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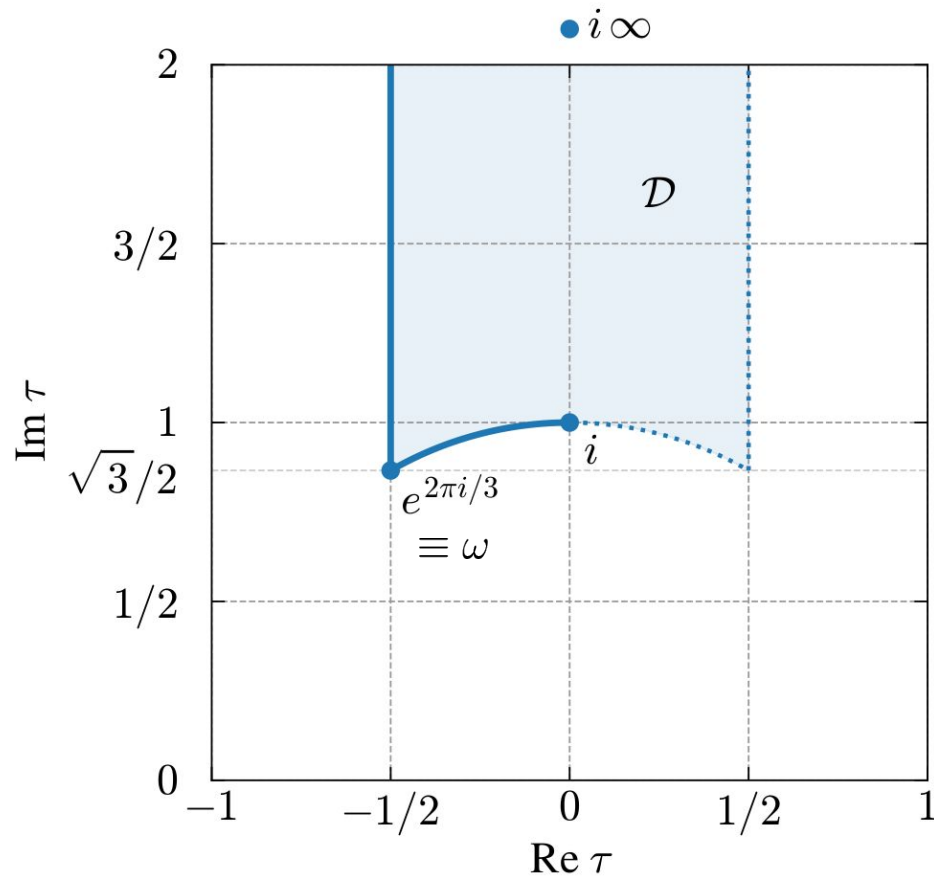
special regions of the fundamental domain

# Intermezzo: Residual modular symmetries



Recall talks by M. Tanimoto, M. Levy

# Intermezzo: Residual modular symmetries



- At special values of  $\tau$ , some residual symmetry remains

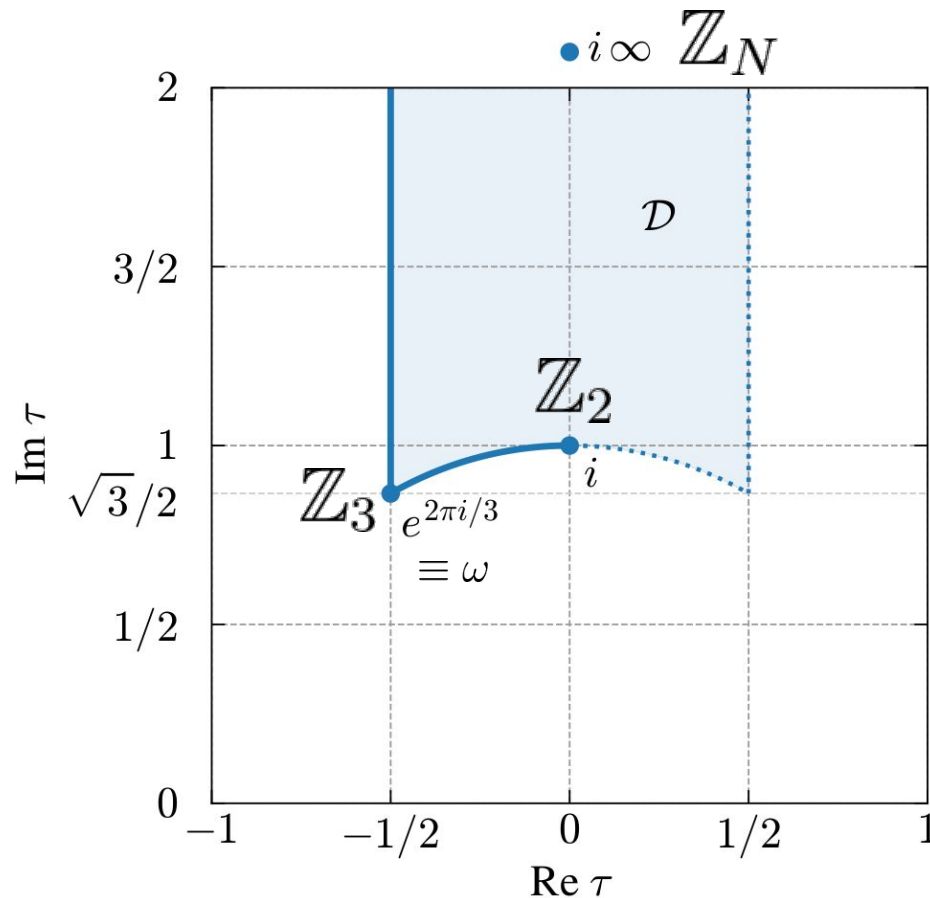
see e.g. Novichkov et al. [1811.04933];  
Novichkov et al. [1812.11289]

- Near them, these symmetries are linearly realized

see e.g. Feruglio [2302.11580]

Recall talks by M. Tanimoto, M. Levy

# Intermezzo: Residual modular symmetries



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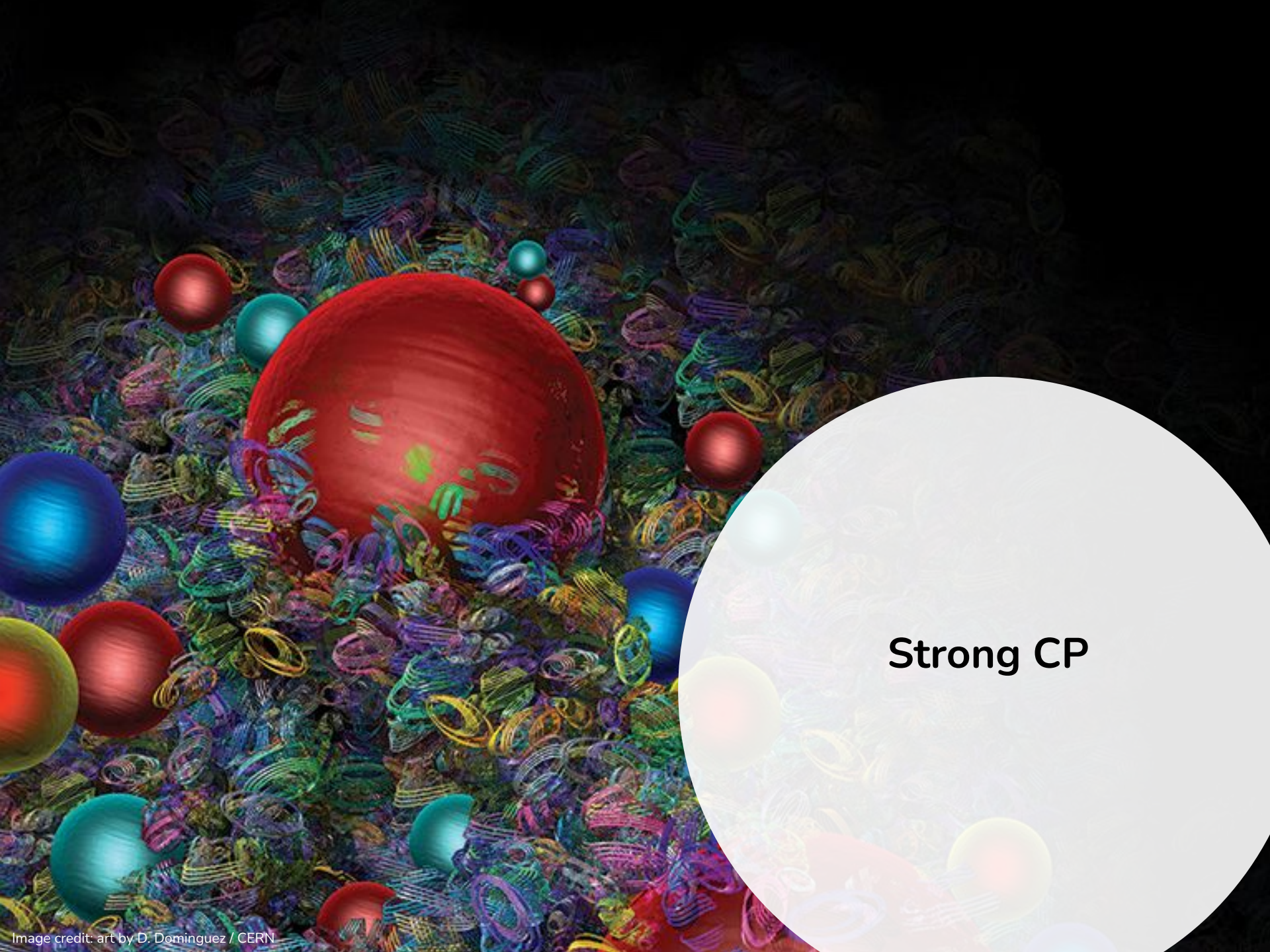
Key idea for hierarchies:

some couplings vanish as we approach a symmetric point

Novichkov, JP, Petcov [2102.07488]

Recall talks by M. Tanimoto, M. Levy

if the base symmetry is smaller, more stabilizers arise,  
see e.g. Varzielas, Levy, Zhou [2008.05329]



# Strong CP

Image credit: art by D. Dominguez / CERN



# The problem

$$\mathcal{L}_{\text{QCD}} = \bar{q}(i\not{D} - M_q)q - \frac{1}{4}G^{a\mu\nu}G_{\mu\nu}^a + \theta\frac{g_s^2}{32\pi^2}G^{a\mu\nu}\tilde{G}_{\mu\nu}^a$$

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The physical quantity:

$$\bar{\theta} = \theta_{\text{QCD}} + \arg \det M_u M_d$$

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$O(1)$  a priori

$$|d_n| < 1.8 \times 10^{-26} \text{ e}\cdot\text{cm} \quad (90\% \text{ C.L.})$$

[2001.11966]



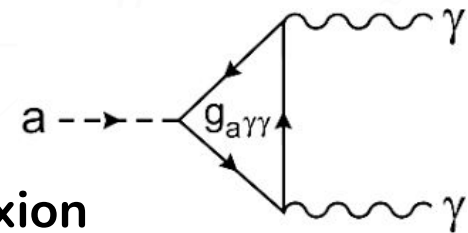
$$|\bar{\theta}| \lesssim 10^{-10}$$

[hep-ph/9908508]



# The problem...

## ...and some solutions



Promotion to a dynamical field: the (invisible) **axion**

Peccei and Quinn [1977], Wilczek, Weinberg [1978]

The physical quantity:

$$\theta = \theta_{\text{QCD}} + \arg \det M_u M_d$$

Solutions with **spontaneously broken CP symmetry**

- Nelson-Barr [1984] (minimal e.g.: Bento-Branco-Parada [1991])
- NHDM with Abelian symmetries — see talk by C. Manzari
- the Feruglio-Strumia-Titov modular solution [2305.08908]

[2001.11000]

[hep-ph/9908500]

# The problem

...and some solutions

- We want to impose CP in the UV
- We need to break it to get the CKM phase
- We need to ensure that after breaking  $\arg \det(Y_u Y_d) = 0$

• NHDM with Abelian symmetries — see talk by C. Manzari  
from C. Manzari slides, on Tuesday

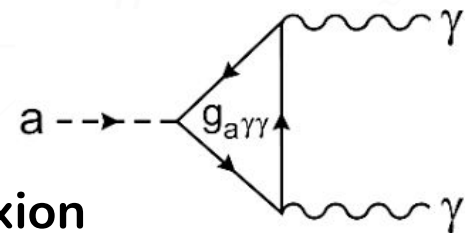
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[hep-ph/9908500]

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[hep-ph/9908500]

# The modular idea

 by Feruglio, Strumia, Titov [2305.08908]

- No axions! see instead [2002.06931, 2402.02071] for a modular origin of the axion
- Need to **produce quark CPV phase** in the CKM mixing matrix
- Need to **suppress**:

$$\bar{\theta} = \theta_{\text{QCD}} + \arg \det M_u M_d$$

↑  
vanishes due to  
imposed gCP

↑  
vanishes due to  
special structure

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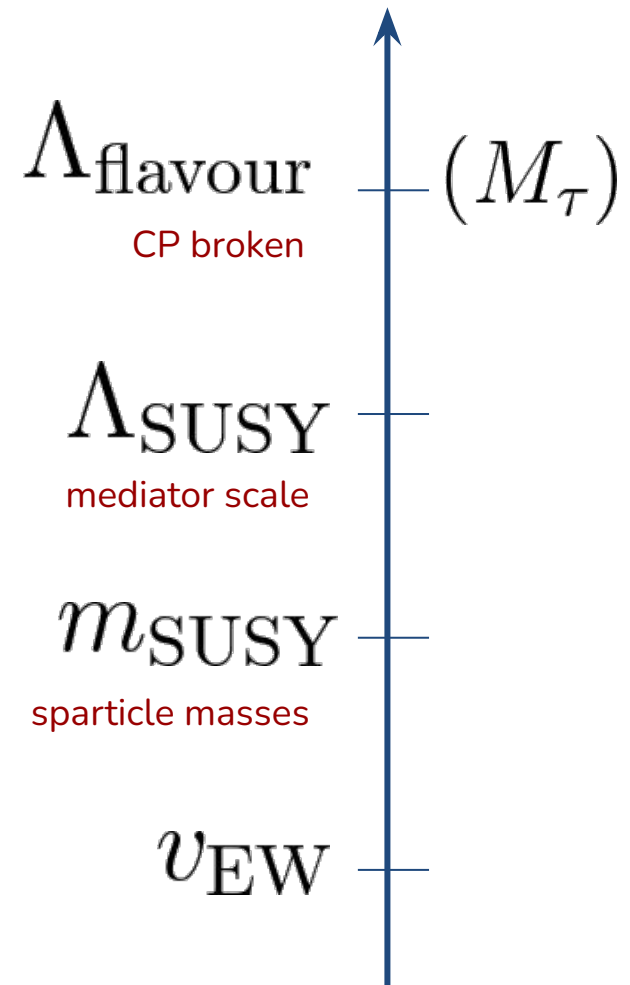
vanishes due to  
special structure

- It turns out to be holomorphic → insensitive to Kähler! 🤔
- Relies on the fact that mass matrix **determinants are modular forms**

already noted e.g. in Ding, Liu, Yao [2211.04546]  
derivation in app. A of JP, Petcov [2404.08032]



...its quality...



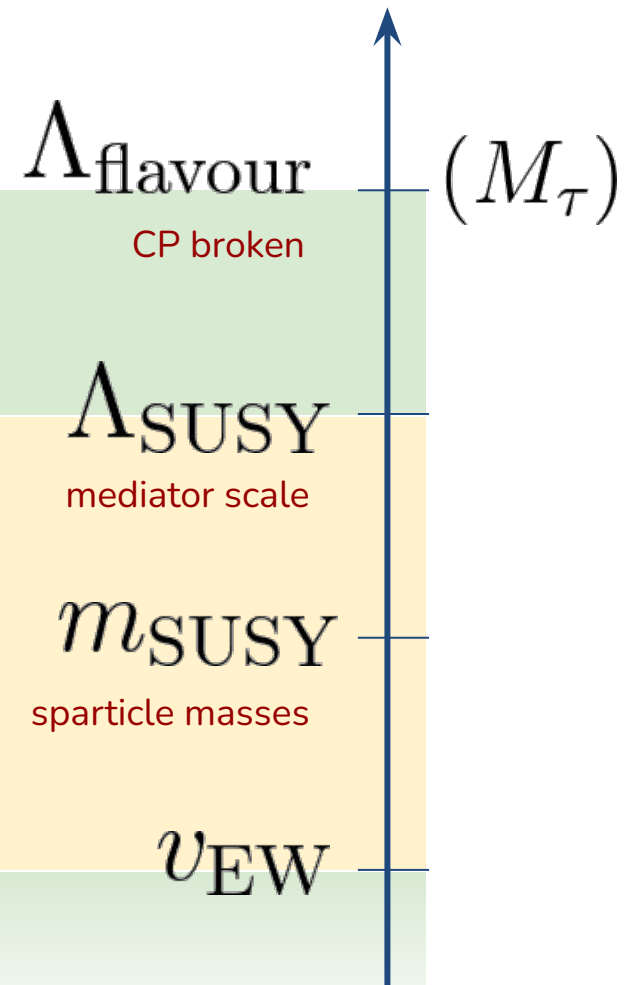
**Separation of scales** also needed to trust model predictions

see Feruglio and Criado [1807.01125]

see also A. Titov slides @ Cosmic WISPer 2024

## ...its quality...

- **No more operators** in the limit of unbroken SUSY
- SUSY **non-renormalization** theorems
- **gluino mass real** if SUSY-breaking sector is CP conserving w/ zero modular charges
- assume gauge or anomaly mediation:  
controls dangerous **threshold corrections**
- consistency will automatically give **real  $\mu$  term**
- SM contribution is **negligible** (4 loops)



**Separation of scales** also needed to trust model predictions

see Feruglio and Criado [1807.01125]

see also A. Titov slides @ Cosmic WISPers 2024

## ...& the consequences

- The determinants have weights

$$k_{\text{det}}^q = 3k_{H_q} + \sum_i k_i + k_i^c$$

$$Q_i \xrightarrow{\gamma} (c\tau + d)^{-k_i} \rho_{ij}(\gamma) Q_j$$

$$q_i^c \xrightarrow{\gamma} (c\tau + d)^{-k_i^c} \rho_{ij}^c(\gamma) q_j^c$$

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$$k_{\text{det}}^q = 3k_{H_q} + \sum_i k_i + k_i^c$$

- To avoid massless quarks, must have  $k_{\text{det}}^q \geq 0$
- Then, to make  $\arg \det M_u M_d$  vanish we require

$$k_{\text{det}}^u = k_{\text{det}}^d = 0$$

so that both determinants are  $\tau$ -independent **constants** (weight 0) and **real**, due to the imposed gCP!

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- The determinants have weights

$$k^q = 2k_{H_u} + \sum k_{H_d} + k^c$$

- Consistent w/ cancellation of the QCD **modular anomaly**, which further implies

$$k_{H_u} + k_{H_d} = 0$$

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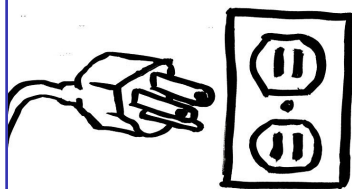
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- The determinants have weights

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**incompatible** w/ modular hierarchy mechanism  
of Novichkov, JP, Petcov [2102.07488] :(

Recall talk by M. Tanimoto

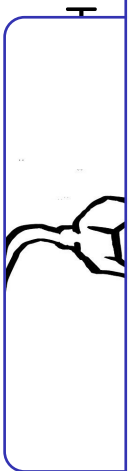
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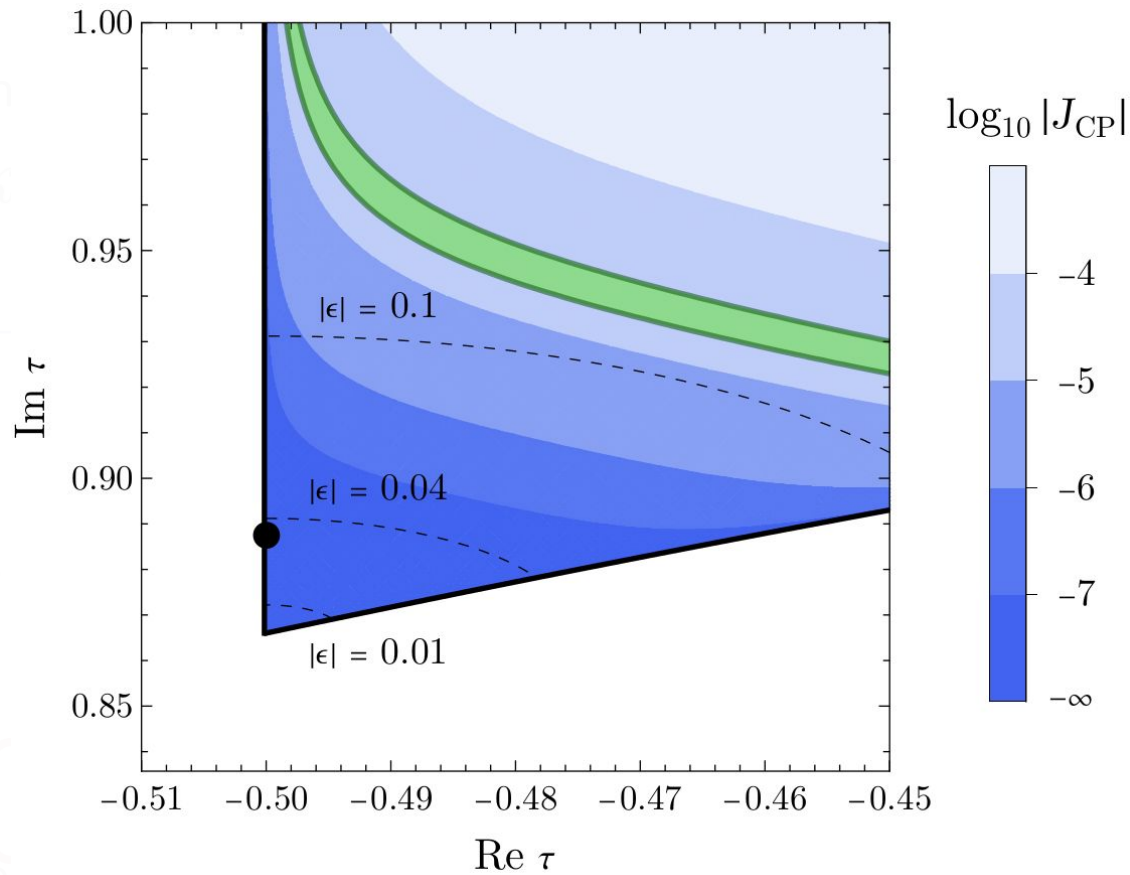
NB.: it was **already difficult** to marry CKM CPV with hierarchies...

$Q_j$   
 $q_j^c$

• The  
 $k$



so  
an



Recall talk by M. Levy

Varzielas, Levy, JP, Petcov [2307.14410]



# Which matrices work?

(weight structures)

$$M : \begin{pmatrix} k_{11} & k_{12} & k_{13} \\ k_{21} & k_{22} & k_{23} \\ k_{31} & k_{32} & k_{33} \end{pmatrix} \quad k_{ij} = k_i + k_j^c$$

$$k_{\text{det}} = k_{11} + k_{22} + k_{33} = k_{12} + k_{23} + k_{31} = \dots \stackrel{!}{=} 0$$

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(weight structures)

$$M : \begin{pmatrix} k + k' & k' & 0 \\ k & 0 & -k' \\ 0 & -k & -k - k' \end{pmatrix}$$

at least one of  $k, k'$  non-zero

JP, Petcov [2404.08032]

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e.g.  $k, k' < 0$

$$M = v_q \begin{pmatrix} 0 & 0 & \alpha_{13} \\ 0 & \alpha_{22} & \alpha_{23} \mathcal{Y}_1^{(|k'|)} \\ \alpha_{31} & \alpha_{32} \mathcal{Y}_2^{(|k|)} & \alpha_{33} \mathcal{Y}_3^{(|k+k'|)} \end{pmatrix}$$

at least one of  $k, k'$  non-zero

$$\det M = -v_q^3 \alpha_{13} \alpha_{22} \alpha_{31}$$

# Which matrices work?

(weight structures in both sectors)

$$M_u : \begin{pmatrix} k + k' & k' & 0 \\ k & 0 & -k' \\ 0 & -k & -k - k' \end{pmatrix}$$

$$M_d : \begin{pmatrix} k + k' & k' & 0 \\ k & 0 & -k' \\ 0 & -k & -k - k' \end{pmatrix}$$

at least one of  $k, k'$  non-zero

up to **simultaneous** permutations from the left  
and independent permutations from the right (weak basis transformations)

# Which irreps work?

- Irreps beyond 1D imply **extra relations** between weights
- **Cannot have triplet** irreps
- Potentially viable non-singlet case: **2+1**

$$Q \sim (\mathbf{2}_Q, k_2) \oplus (\mathbf{1}_Q, k_1),$$

$$u^c \sim (\overline{\mathbf{2}}_Q, -k_2) \oplus (\overline{\mathbf{1}}_Q, -k_1),$$

$$d^c \sim (\overline{\mathbf{2}}_Q, -k_2) \oplus (\overline{\mathbf{1}}_Q, -k_1),$$

$$M_q \propto \begin{pmatrix} |\alpha_1^q/\beta_q| & 0 & \cos \theta_q e^{i\phi_1^q} \\ 0 & |\alpha_1^q/\beta_q| & \sin \theta_q e^{i\phi_2^q} \\ 0 & 0 & |\alpha_2^q/\beta_q| \end{pmatrix}^{(T)}$$

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 \quad
 M_q \propto \begin{pmatrix} |\alpha_1^q/\beta_q| & 0 & \cos \theta_q e^{i\phi_1^q} \\ 0 & |\alpha_1^q/\beta_q| & \sin \theta_q e^{i\phi_2^q} \\ 0 & 0 & |\alpha_2^q/\beta_q| \end{pmatrix}^{(T)}$$

too much strain on the model, does not work → **quarks must furnish 1D irreps**

$$k_2 > k_1$$

in the interval  $[0, 1]$ , implying the upper bound  $|V_{us}| \lesssim m_u/m_c \simeq 0.002$ . This is two orders of magnitude smaller than the value  $|V_{us}| \simeq 0.225$  [45] required by quark data. By a similar procedure, one can also show that

$$\left| |V_{ub}| - \frac{m_d}{m_s} \right| \lesssim \frac{m_u}{m_c}, \quad (41)$$

i.e. that  $|V_{ub}| \in [0.048, 0.053]$ , again in contradiction with the data, which requires the much smaller  $|V_{ub}| \simeq 0.003$ .

$$k_2 < k_1$$

once again leading to an upper bound on the magnitude of a mixing matrix element. In this case one has  $|V_{cb}| \lesssim m_s/m_b \simeq 0.014$ , while data requires  $|V_{cb}| \simeq 0.036$ , a value more than twice as large.

# What can we do with 1D?

$$\begin{aligned} Q &\sim (\mathbf{1}_{Q_1}, k_1) \oplus (\mathbf{1}_{Q_2}, k_2) \oplus (\mathbf{1}_{Q_3}, k_3), \\ u^c &\sim (\overline{\mathbf{1}_{Q_1}}, -k_1) \oplus (\overline{\mathbf{1}_{Q_2}}, -k_2) \oplus (\overline{\mathbf{1}_{Q_3}}, -k_3) \\ d^c &\sim (\overline{\mathbf{1}_{Q_1}}, -k_1) \oplus (\overline{\mathbf{1}_{Q_2}}, -k_2) \oplus (\overline{\mathbf{1}_{Q_3}}, -k_3) \end{aligned}$$

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$$\text{I : } \begin{pmatrix} \alpha_1^q & 0 & \tilde{\alpha}_{13}^q \mathcal{Y}_{q,13}^{(k+k')} \\ 0 & \alpha_2^q & \tilde{\alpha}_{23}^q \mathcal{Y}_{q,23}^{(k)} \\ 0 & 0 & \alpha_3^q \end{pmatrix}, \quad k > 0, k' \geq 0,$$

$$\text{II : } \begin{pmatrix} \alpha_1^q & \tilde{\alpha}_{12}^q \mathcal{Y}_{q,12}^{(k')} & \tilde{\alpha}_{13}^q \mathcal{Y}_{q,13}^{(k+k')} \\ 0 & \alpha_2^q & 0 \\ 0 & 0 & \alpha_3^q \end{pmatrix}, \quad k \geq 0, k' > 0,$$

$$\text{III : } \begin{pmatrix} \alpha_1^q & \tilde{\alpha}_{12}^q \mathcal{Y}_{q,12}^{(k')} & \tilde{\alpha}_{13}^q \mathcal{Y}_{q,13}^{(k+k')} \\ 0 & \alpha_2^q & \tilde{\alpha}_{23}^q \mathcal{Y}_{q,23}^{(k)} \\ 0 & 0 & \alpha_3^q \end{pmatrix}, \quad k, k' > 0,$$

$$\text{IV : } \begin{pmatrix} \alpha_{11}^q & \alpha_{12}^q & \tilde{\alpha}_{13}^q \mathcal{Y}_{q,13}^{(k)} \\ \alpha_{21}^q & \alpha_{22}^q & \tilde{\alpha}_{23}^q \mathcal{Y}_{q,23}^{(k)} \\ 0 & 0 & \alpha_3^q \end{pmatrix}, \quad k > 0,$$

$$\text{V : } \begin{pmatrix} \alpha_1^q & \tilde{\alpha}_{12}^q \mathcal{Y}_{q,12}^{(k')} & \tilde{\alpha}_{13}^q \mathcal{Y}_{q,13}^{(k')} \\ 0 & \alpha_{22}^q & \alpha_{23}^q \\ 0 & \alpha_{32}^q & \alpha_{33}^q \end{pmatrix}, \quad k' > 0,$$

- **New irreps** and **forms** beyond  $\Gamma$
- Zeros from **non-trivial 1D** or **missing form**
- Several forms may contribute to each entry

$$\mathcal{Y}_{q,ij}^{(w)} \equiv \sum_{n=1} g_n^q Y_{\mathbf{1}_{ij},n}^{(w)}$$

- Canonical rescaling may help with hierarchies

$$\tilde{\alpha}_{ij}^q \equiv \alpha_{ij}^q (2 \operatorname{Im} \tau)^{k_{ij}/2}$$

- Forms are related, so getting CPV is tricky!

# How “minimal” can we be?

$$\begin{aligned}
 Q &\sim (\mathbf{1}_{Q_1}, k_1) \oplus (\mathbf{1}_{Q_2}, k_2) \oplus (\mathbf{1}_{Q_3}, k_3), \\
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$$k \equiv k_2 - k_3, \quad k' \equiv k_1 - k_2$$

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 \geq 6 \text{ params.}$$

$$\text{II : } \begin{pmatrix} \alpha_1^q & \tilde{\alpha}_{12}^q \mathcal{Y}_{q,12}^{(k')} & \tilde{\alpha}_{13}^q \mathcal{Y}_{q,13}^{(k+k')} \\ 0 & \alpha_2^q & 0 \\ 0 & 0 & \alpha_3^q \end{pmatrix}, \quad k \geq 0, k' > 0, \\
 \geq 6 \text{ params.}$$

$$\text{III : } \begin{pmatrix} \alpha_1^q & \tilde{\alpha}_{12}^q \mathcal{Y}_{q,12}^{(k')} & \tilde{\alpha}_{13}^q \mathcal{Y}_{q,13}^{(k+k')} \\ 0 & \alpha_2^q & \tilde{\alpha}_{23}^q \mathcal{Y}_{q,23}^{(k)} \\ 0 & 0 & \alpha_3^q \end{pmatrix}, \quad k, k' > 0, \\
 \geq 7 \text{ params.}$$

$$\text{IV : } \begin{pmatrix} \alpha_{11}^q & \alpha_{12}^q & \tilde{\alpha}_{13}^q \mathcal{Y}_{q,13}^{(k)} \\ \alpha_{21}^q & \alpha_{22}^q & \tilde{\alpha}_{23}^q \mathcal{Y}_{q,23}^{(k)} \\ 0 & 0 & \alpha_3^q \end{pmatrix}, \quad k > 0, \\
 \geq 8 \text{ params.}$$

$$\text{V : } \begin{pmatrix} \alpha_1^q & \tilde{\alpha}_{12}^q \mathcal{Y}_{q,12}^{(k')} & \tilde{\alpha}_{13}^q \mathcal{Y}_{q,13}^{(k')} \\ 0 & \alpha_{22}^q & \alpha_{23}^q \\ 0 & \alpha_{32}^q & \alpha_{33}^q \end{pmatrix}, \quad k' > 0, \\
 \geq 8 \text{ params.}$$

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- Canonical rescaling may help with hierarchies

$$\tilde{\alpha}_{ij}^q \equiv \alpha_{ij}^q (2 \text{Im } \tau)^{k_{ij}/2} \in \mathbb{R}$$

- Forms are related, so getting CPV is **tricky!**  
*need more params. than naively expected*

# The minimal & next-to-minimal landscapes

# params. = 6+6+2 = 14

# params. = 7+7+2 = 16

$(\mathbf{k}, \mathbf{k}')$	Minimal models (I and II)	Next-to-minimal models (I and II)	Next-to-minimal models (III)
All $\Gamma'_N$	(10, 12), (12, 14), (14, 16)	(16, 18), (18, 20), (20, 22)	(4, 8), (4, 14), (6, 6), (6, 10), (6, 14), (8, 8), (8, 10), (8, 14), (10, 10)
$S_3$ only	(10', 12), (10, 18'), (10', 18), (12, 12'), (12, 14'), (12, 16'), (12', 16), (12, 20'), (12', 20), (14', 16), (14, 18'), (14', 18), (14, 22'), (14', 22), (16, 16'), (16', 18), (16', 18'), (16, 20'), (16', 20), (18, 20'), (18', 20'), (20, 20'), (20', 22), (20', 22')	(16, 18'), (16, 24'), (16', 24), (18, 18'), (18', 20), (18, 22'), (18', 22), (18, 26'), (18', 26), (20, 22'), (20, 24'), (20', 24), (20, 28'), (20', 28), (22, 22'), (22, 24'), (22', 24'), (22, 26'), (22', 26), (24', 26), (24', 26'), (26, 26'), (26, 28'), (26', 28')	(4, 14'), (4, 20'), (6', 6'), (6', 10'), (6, 12'), (6', 12'), (6', 14'), (6, 16'), (6', 16'), (6, 20'), (6', 20'), (8, 10'), (8, 14'), (8, 16'), (8, 20'), (10', 10'), (10, 12'), (10', 12'), (10, 14'), (10', 14), (10, 16'), (10', 16'), (12', 14), (12', 14'), (14, 14')
$A'_4$ only	(8', 12), (8', 18), (10', 12), (10, 16'), (10', 16), (10, 20''), (10', 20), (12, 12'), (12, 12''), (12, 14'), (12, 14''), (12, 16''), (12', 16), (12'', 16'), (12, 18'), (12, 18''), (12', 18), (12'', 18), (12, 22''), (12', 22), (12'', 22'), (14, 16'), (14', 16), (14', 16'), (14'', 16), (14'', 16'), (14', 18), (14, 20'), (14, 20''), (14', 20), (14', 20''), (14'', 20), (14'', 20'), (14, 24''), (14'', 24'), (16, 16''), (16', 16''), (16, 18'), (16, 18''), (16', 18'), (16', 18''), (16'', 18), (16'', 20'), (16, 22''), (16', 22''), (16'', 22), (16'', 22'), (16'', 26'), (18, 18'), (18, 18''), (18', 20), (18', 20'), (18', 20''), (18'', 20), (18'', 20'), (18'', 20''), (18, 22''), (18', 22), (18'', 22'), (18', 24''), (18'', 24'), (20, 22''), (20', 22''), (20'', 22''), (22, 22''), (22', 22''), (22'', 24'), (22'', 24''), (22'', 26'), (22'', 26''), (22'', 28''), (22'', 28''), (22'', 28), (22'', 28'), (22'', 32'), (24', 24''), (24', 26), (24', 26'), (24', 26''), (24'', 26), (24'', 26'), (24'', 26''), (24'', 30''), (24'', 30')	(14', 24), (16, 16'), (16', 18), (16, 20''), (16', 20), (16, 22'), (16', 22), (16, 26''), (16', 26), (18, 20'), (18, 20''), (18, 24'), (18, 24''), (18', 24), (18'', 24), (18, 28''), (18', 28), (18'', 28'), (20, 20'), (20, 20''), (20', 20''), (20, 22'), (20', 22), (20', 22'), (20'', 22), (20'', 22'), (20, 24''), (20'', 24'), (20, 26'), (20, 26''), (20', 26), (20', 26''), (20'', 26), (20'', 26'), (20, 30''), (20'', 30'), (22, 22'), (22, 24'), (22, 24''), (22', 24'), (22', 24''), (22'', 24'), (22'', 24''), (22, 26''), (22', 26), (22, 28''), (22'', 28''), (22'', 28), (22'', 28'), (22'', 32'), (24', 24''), (24', 26), (24', 26'), (24', 26''), (24'', 26), (24'', 26'), (24'', 26''), (24'', 30''), (24'', 30')	(4', 8''), (4, 12'), (4', 12''), (4', 14''), (4, 16''), (4', 16''), (4, 18'), (4', 18''), (4, 22''), (4', 22''), (6, 10'), (6, 14'), (6, 14''), (6, 18'), (6, 18''), (6, 22''), (8, 8'), (8', 8''), (8'', 8''), (8'', 10'), (8, 12'), (8, 12''), (8', 12'), (8', 12''), (8'', 12'), (8'', 12''), (8, 14'), (8', 14), (8', 14''), (8'', 14'), (8'', 14''), (8, 16''), (8'', 16''), (8, 18'), (8, 18''), (8', 18'), (8', 18''), (8'', 18'), (8'', 18''), (8, 22''), (8'', 22''), (10, 10'), (10', 10'), (10, 12'), (10', 12''), (10, 14'), (10, 14''), (10', 14), (10', 14''), (10, 16''), (10', 16''), (10, 18''), (10', 18'), (12', 12'), (12'', 12''), (12', 14), (12', 14'), (12', 14''), (12'', 14), (12'', 14'), (12'', 14''), (12', 18'), (12'', 18''), (14, 14''), (14', 14'), (14, 16''), (14'', 16'')

• • •

listed all 462 one can get with  $\Gamma'_N$

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$S_3$ only	(10', 12), (10, 18'), (10', 18), (12, 12'), (12, 14'), (12, 16'), (12', 16), (12, 20'), (12', 20), (14', 16), (14, 18'), (14', 18), (14, 22'), (14', 22), (16, 16'), (16', 18), (16', 18'), (16, 20'), (16', 20), (18, 20'), (18', 20'), (20, 20'), (20', 22), (20', 22')	(16, 18'), (16, 24'), (16', 24), (18, 18'), (18', 20), (18, 22'), (18', 22), (18, 26'), (18', 26), (20, 22'), (20, 24'), (20', 24), (20, 28'), (20', 28), (22, 22'), (22, 24'), (22', 24'), (22, 26'), (22', 26), (24', 26), (24', 26'), (26, 26'), (26, 28'), (26', 28')	(4, 8), (4, 14), (6, 6), (6, 10), (6, 14), (8, 8), Feruglio, Strumia, Titov [2305.08908] (8, 20'), (10', 10'), (10, 12'), (10', 12'), (10, 14'), (10', 14), (10, 16'), (10', 16'), (12', 14), (12', 14'), (14, 14')
$A'_4$ only	(8', 12), (8', 18), (10', 12), (10, 16'), (10', 16), (10, 20''), (10', 20), (12, 12'), (12, 12''), (12, 14'), (12, 14''), (12, 16''), (12', 16), (12'', 16'), (12, 18'), (12, 18''), (12', 18), (12'', 18), (12, 22''), (12', 22), (12'', 22'), (14, 16'), (14', 16), (14', 16'), (14'', 16), (14'', 16'), (14', 18), (14, 20'), (14, 20''), (14', 20), (14', 20''), (14'', 20'), (14, 24''), (14'', 24'), (16, 16''), (16', 16''), (16, 18'), (16, 18''), (16', 18'), (16', 18''), (16'', 18), (16'', 20'), (16, 22''), (16', 22''), (16'', 22), (16'', 22'), (16'', 26'), (18, 18'), (18, 18''), (18', 20), (18', 20'), (18', 20''), (18'', 20), (18'', 20'), (18'', 20''), (18, 22''), (18', 22), (18'', 22'), (18', 24''), (18'', 24'), (20, 22''), (20', 22''), (20'', 22''), (22, 22''), (22', 22''), (22'', 24'), (22'', 24''), (22'', 26'), (22'', 26''), (22'', 28''), (22'', 28''), (22'', 28), (22'', 28'), (22'', 32'), (24', 24''), (24', 26), (24', 26'), (24', 26''), (24'', 26), (24'', 26'), (24'', 26''), (24', 30''), (24'', 30')	(14', 24), (16, 16'), (16', 18), (16, 20''), (16', 20), (16, 22'), (16', 22), (16, 26''), (16', 26), (18, 20'), (18, 20''), (18, 24'), (18, 24''), (18', 24), (18'', 24), (18, 28''), (18', 28), (18'', 28'), (20, 20'), (20, 20''), (20', 20''), (20, 22'), (20', 22), (20', 22'), (20'', 22), (20'', 22'), (20, 24''), (20'', 24'), (20, 26'), (20, 26''), (20', 26), (20', 26''), (20'', 26), (20'', 26'), (20, 30''), (20'', 30'), (22, 22'), (22, 24'), (22, 24''), (22', 24'), (22', 24''), (22'', 24'), (22'', 24''), (22'', 26'), (22'', 26''), (22'', 28'), (22'', 28''), (22'', 28), (22'', 28'), (22'', 32'), (24', 24''), (24', 26), (24', 26'), (24', 26''), (24'', 26), (24'', 26'), (24'', 26''), (24', 30''), (24'', 30')	(4', 8''), (4, 12'), (4, 16''), (4, 16'), (4, 22''), (4', 22''), (6, 14''), (6, 18'), (6, 18''), (8', 8''), (8'', 8''), (8'', 10'), (8, 12'), (8, 12''), (8', 12'), (8'', 12'), (8'', 12''), (8, 14'), (8', 14), (8', 14''), (8'', 14'), (8'', 14''), (8, 16''), (8'', 16''), (8, 18'), (8, 18''), (8', 18'), (8', 18''), (8'', 18'), (8'', 18''), (8, 22''), (8'', 22''), (10, 10'), (10', 10'), (10, 12'), (10', 12''), (10, 14'), (10, 14''), (10', 14), (10', 14''), (10, 16''), (10', 16''), (10, 18''), (10', 18'), (12', 12'), (12'', 12''), (12', 14), (12', 14'), (12', 14''), (12'', 14), (12'', 14'), (12'', 14''), (12', 18'), (12'', 18''), (14, 14''), (14', 14'), (14, 16''), (14'', 16'')
			Petcov, Tanimoto [2404.00858]



listed all 462 one can get with  $\Gamma'_N$

# The minimal & next-to-minimal landscapes



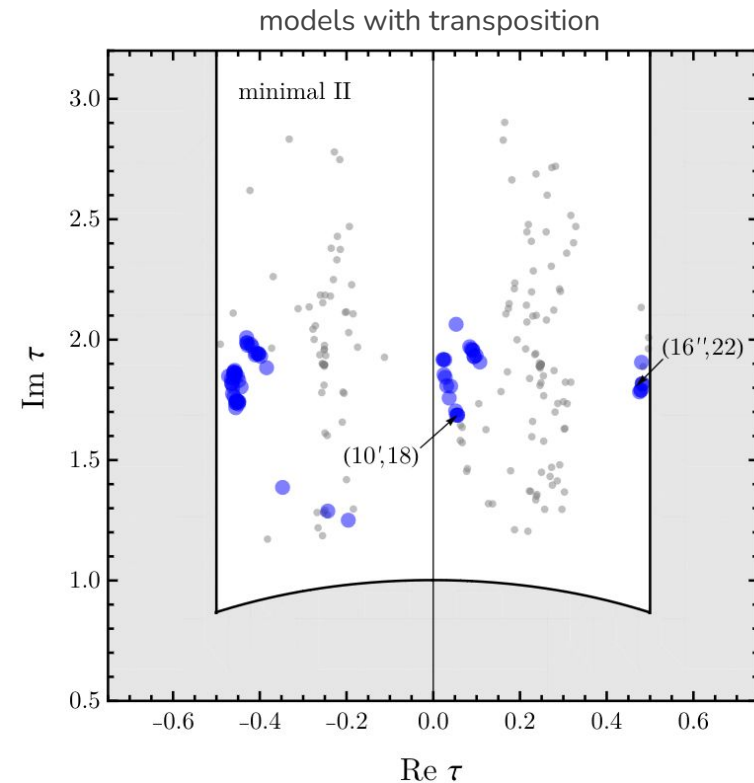
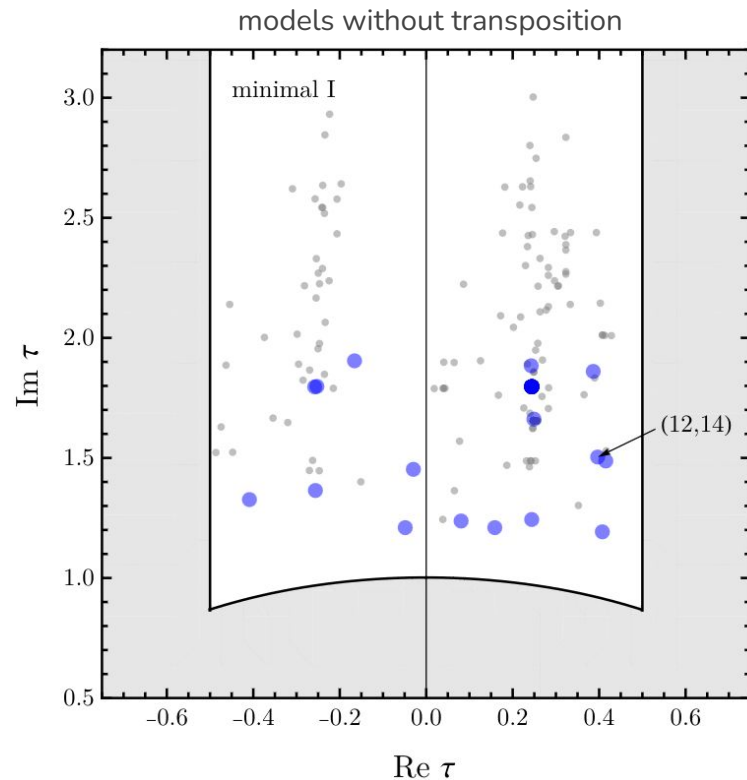
$S_4$   
only

- $(\widehat{7}', 12), (\widehat{7}', 18), (\widehat{9}', 12), (\widehat{9}', 16), (\widehat{9}', 20),$   
 $(10', 12), (10, \widehat{15}'), (10', \widehat{15}'), (10, 18'),$   
 $(10', 18), (10, \widehat{21}'), (10', \widehat{21}'), (\widehat{11}', 12),$   
 $(\widehat{11}', 16), (\widehat{11}', 18), (\widehat{11}', 22), (12, 12'),$   
 $(12, \widehat{13}), (12, \widehat{13}'), (12, 14'), (12, 15),$   
 $(12', \widehat{15}'), (12, 16'), (12', 16), (12, \widehat{17}'),$   
 $(12, \widehat{17}'), (12, \widehat{19}), (12', \widehat{19}'), (12, 20'),$   
 $(12', 20), (12, \widehat{23}), (12', \widehat{23}'), (\widehat{13}, \widehat{15}'),$   
 $(\widehat{13}', \widehat{15}'), (\widehat{13}', 16), (\widehat{13}, 18), (\widehat{13}, 18'),$   
 $(\widehat{13}', 18), (\widehat{13}', 18'), (\widehat{13}', 20), (\widehat{13}, \widehat{21}'),$   
 $(\widehat{13}', \widehat{21}'), (\widehat{13}, 24'), (14, \widehat{15}'), (14', \widehat{15}'),$   
 $(14', 16), (14, 18'), (14', 18), (14, \widehat{19}'),$   
 $(14', \widehat{19}'), (14, \widehat{21}'), (14', \widehat{21}'), (14, 22'),$   
 $(14', 22), (14, \widehat{25}), (14', \widehat{25}'), (\widehat{15}, \widehat{15}'),$   
 $(\widehat{15}, 16), (\widehat{15}', 16'), (\widehat{15}', \widehat{17}'), (\widehat{15}', \widehat{17}'),$   
 $(\widehat{15}, 18'), (\widehat{15}, 19'), (\widehat{15}', 19), (\widehat{15}, 20),$   
 $(\widehat{15}', 20'), (\widehat{15}, 22'), (\widehat{15}, \widehat{23}'), (\widehat{15}', \widehat{23}'),$   
 $(\widehat{15}, 26'), (16, 16'), (16, \widehat{17}'), (16, \widehat{17}'),$   
 $(16', 18), (16', 18'), (16, \widehat{19}), (16', \widehat{19}'),$   
 $(16, 20'), (16', 20), (16', \widehat{21}'), (16', \widehat{21}'),$   
 $(16, \widehat{23}), (16', \widehat{23}'), (\widehat{17}, 18), (\widehat{17}, 18'),$   
 $(\widehat{17}', 18), (\widehat{17}', 18'), (\widehat{17}, \widehat{19}'), (\widehat{17}', \widehat{19}'),$   
 $(\widehat{17}', 20), (\widehat{17}, \widehat{21}'), (\widehat{17}', \widehat{21}'), (\widehat{17}, 22),$   
 $(\widehat{17}, 22'), (\widehat{17}', 22), (\widehat{17}', 22'), (\widehat{17}, 24'),$   
 $(\widehat{17}, \widehat{25}'), (\widehat{17}', \widehat{25}'), (\widehat{17}, 28'), (18, \widehat{19}),$   
 $(18', \widehat{19}), (18, 20'), (18', 20'), (18, \widehat{23}),$   
 $(18', \widehat{23}), (\widehat{19}, \widehat{19}'), (\widehat{19}, 20), (\widehat{19}', 20'),$   
 $(\widehat{19}, \widehat{21}'), (\widehat{19}, \widehat{21}'), (\widehat{19}, 22'), (\widehat{19}, \widehat{23}'),$   
 $(\widehat{19}', \widehat{23}'), (\widehat{19}, 24'), (\widehat{19}, 26'), (20, 20'),$   
 $(20', \widehat{21}'), (20', 22), (20', 22'),$   
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 $(\widehat{21}, \widehat{23}), (\widehat{21}', \widehat{23}), (22, \widehat{23}), (22', \widehat{23}),$   
 $(\widehat{23}, \widehat{23}'), (\widehat{23}, 24'), (\widehat{23}, 25'), (\widehat{23}, 25'),$   
 $(\widehat{23}, 26'), (\widehat{23}, 28')$
- $(\widehat{13}', 24), (\widehat{15}', 16), (\widehat{15}', 18), (\widehat{15}', 20),$   
 $(\widehat{15}', 22), (\widehat{15}', 26), (16, 18'), (16, \widehat{21}'),$   
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 $(16', \widehat{27}'), (\widehat{17}', 24), (\widehat{17}', 28), (18, 18'),$   
 $(18, \widehat{19}'), (18', \widehat{19}'), (18', 20), (18, \widehat{21}'),$   
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 $(18', \widehat{23}'), (18, \widehat{25}), (18', \widehat{25}'), (18, 26'),$   
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 $(20, 24'), (20', 24), (20, \widehat{25}), (20, \widehat{25}'),$   
 $(20, \widehat{27}'), (20', \widehat{27}'), (20, 28'), (20', 28),$   
 $(20, \widehat{31}), (20', \widehat{31}'), (\widehat{21}, \widehat{21}'), (\widehat{21}, 22),$   
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 $(\widehat{21}', \widehat{23}'), (\widehat{21}, 24'), (\widehat{21}, \widehat{25}'), (\widehat{21}', \widehat{25}'),$   
 $(\widehat{21}, 26), (\widehat{21}, 26'), (\widehat{21}', 26), (\widehat{21}', 26'),$   
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 $(\widehat{27}, 28'), (\widehat{27}, 29), (\widehat{27}, \widehat{29}'), (\widehat{27}, 32'),$   
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 $(29', \widehat{31}'), (29, 32'), (31, 32')$
- $(\widehat{3}', 9), (\widehat{3}', \widehat{13}), (\widehat{3}', \widehat{15}), (\widehat{3}', \widehat{17}), (\widehat{3}', \widehat{19}),$   
 $(\widehat{3}', \widehat{23}), (4, \widehat{11}'), (4, 14'), (4, \widehat{17}'), (4, \widehat{17}'),$   
 $(4, 20'), (4, \widehat{23}), (6', 6'), (6, \widehat{9}'), (6', \widehat{9}),$   
 $(6', 10'), (6, 12'), (6', 12'), (6, \widehat{13}'),$   
 $(6', \widehat{13}), (6', 14'), (6, \widehat{15}), (6', \widehat{15}), (6, 16'),$   
 $(6', 16'), (6, \widehat{17}'), (6', \widehat{17}'), (6, \widehat{19}), (6', \widehat{19}),$   
 $(6, 20'), (6', 20'), (6, \widehat{23}), (6', \widehat{23}), (\widehat{7}', 8),$   
 $(\widehat{7}', 9), (\widehat{7}', \widehat{11}'), (\widehat{7}', \widehat{13}), (\widehat{7}', 14), (\widehat{7}', 14'),$   
 $(\widehat{7}', 15), (\widehat{7}', \widehat{17}'), (\widehat{7}', \widehat{19}), (\widehat{7}', 20'),$   
 $(8, 10'), (8, \widehat{11}'), (8, \widehat{13}), (8, \widehat{13}'), (8, 14'),$   
 $(8, 16'), (8, \widehat{17}'), (8, \widehat{17}'), (8, \widehat{19}), (8, 20'),$   
 $(8, \widehat{23}), (\widehat{9}, 9), (\widehat{9}, 9'), (\widehat{9}', 9'), (\widehat{9}, 10'),$   
 $(\widehat{9}', 10), (\widehat{9}, \widehat{11}'), (\widehat{9}, 12'), (\widehat{9}', 12'), (\widehat{9}, \widehat{13}),$   
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 $(\widehat{9}', 17'), (\widehat{9}', 17'), (\widehat{9}, 19), (\widehat{9}, 20'), (\widehat{9}', 20'),$   
 $(\widehat{9}, 23), (10', 10'), (10, \widehat{11}'), (10', \widehat{11}'),$   
 $(10, 12'), (10', 12'), (10, \widehat{13}'), (10', \widehat{13}),$   
 $(10, 14'), (10', 14), (10, \widehat{15}), (10', \widehat{15}),$   
 $(10, 16'), (10', 16'), (10, \widehat{17}'), (10', \widehat{17}'),$   
 $(10, \widehat{19}), (10', \widehat{19}), (\widehat{11}, \widehat{11}'), (\widehat{11}', \widehat{13}'),$   
 $(\widehat{11}', 14), (\widehat{11}', 14'), (\widehat{11}', \widehat{15}), (\widehat{11}', 16'),$   
 $(\widehat{11}', \widehat{17}'), (\widehat{11}', 20'), (12', \widehat{13}), (12', \widehat{13}'),$   
 $(12', 14), (12', 14'), (12', \widehat{17}'), (12', \widehat{17}'),$   
 $(\widehat{13}, \widehat{13}), (\widehat{13}, \widehat{13}'), (\widehat{13}', \widehat{13}'), (\widehat{13}, 14),$   
 $(\widehat{13}', 14'), (\widehat{13}, \widehat{15}), (\widehat{13}, 16'), (\widehat{13}', 16'),$   
 $(\widehat{13}, 19), (14, 14'), (14, \widehat{15}), (14', \widehat{15}),$   
 $(14, \widehat{17}'), (14', \widehat{17}'), (\widehat{15}, \widehat{17})$

fin



# A peek into the minimal model landscape



**Moduli selection** from the minimization of **parameter hierarchies**  
(log spread), using the normalization proposal in Petcov [\[2311.04185\]](#)

# An $S_4$ ' benchmark

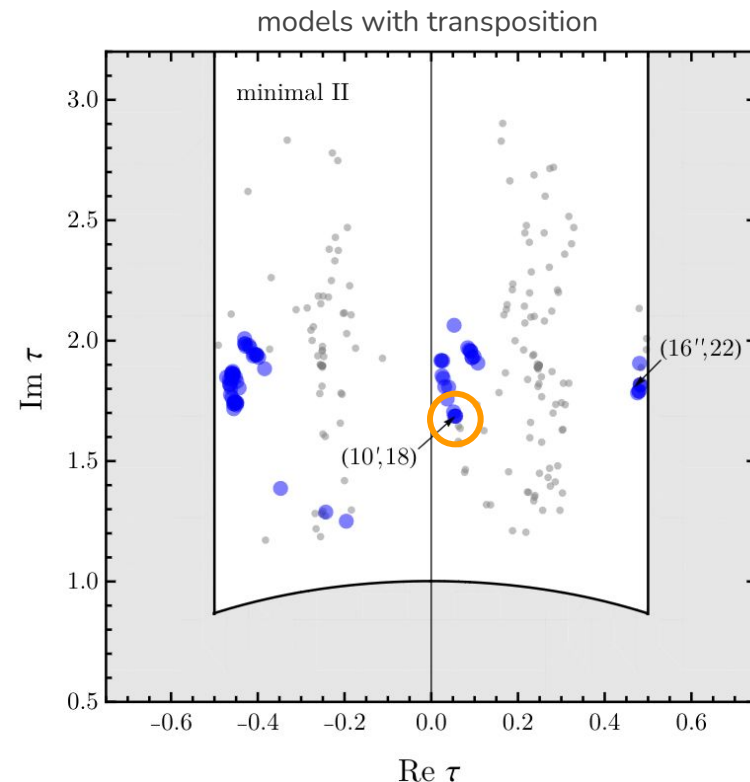
- Quark mass hierarchies and mixing angles can be reproduced by  **$O(1)$  parameters**

Also in FST [2305.08908]

- Overall tendency: parametric hierarchies mostly in downs

$$\alpha_2^u \simeq 1.00 \alpha_1^u, \quad \alpha_3^u \simeq 1.65 \alpha_1^u, \quad \alpha_{12}^u \simeq 1.27 \alpha_1^u, \\ \alpha_{13,1}^u \simeq 1.47 \alpha_1^u, \quad \alpha_{13,2}^u \simeq 1.61 \alpha_1^u,$$

$$\alpha_2^d \simeq 0.48 \alpha_1^d, \quad \alpha_3^d \simeq 0.05 \alpha_1^d, \quad \alpha_{12}^d \simeq 7.92 \alpha_1^d, \\ \alpha_{13,1}^d \simeq -7.90 \alpha_1^d, \quad \alpha_{13,2}^d \simeq 0.05 \alpha_1^d,$$



$$\tau \simeq 0.06 + 1.69 i$$

# Extra! Extra! Read all about it!

Feruglio, Parriciatu, Strumia, Titov [2406.01689] **NEW!**



- can use **lower weights** if we **forego minimality**
- can get **non-Abelian irreps** if we **add vector-like quarks**
- consider models with level  $N > 1$ , either using  $\Gamma_N$  or  $\Gamma(N)$  instead of  $\Gamma$

$\Gamma_2$ :

	SM quarks			Extra vector-like quarks			
	$Q$	$D^c$	$U^c$	$D^{c'}$	$D'$	$U^{c'}$	$U'$
$SU(2)_L \otimes U(1)_Y$	$2_{1/6}$	$1_{1/3}$	$1_{-2/3}$	$1_{1/3}$	$1_{-1/3}$	$1_{-2/3}$	$1_{2/3}$
Flavour symmetry $\Gamma_2$	$2 \oplus 1_0$	$2 \oplus 1_1$	$2 \oplus 1_0$	$2 \oplus 1_0$	$2 \oplus 1_1$	$2 \oplus 1_0$	$2 \oplus 1_0$
Modular weights $k_\Phi$	-2	-2	-2	+2	+2	+2	+2

$\Gamma_3$ :

	SM quarks			Extra vector-like quarks			
	$Q$	$D^c$	$U^c$	$D^{c'}$	$D'$	$U^{c'}$	$U'$
$SU(2)_L \otimes U(1)_Y$	$2_{1/6}$	$1_{1/3}$	$1_{-2/3}$	$1_{1/3}$	$1_{-1/3}$	$1_{-2/3}$	$1_{2/3}$
Flavour symmetry $\Gamma_3$	$3$	$3$	$3$	$3$	$3$	$3$	$3$
Modular weights $k_\Phi$	-1	$\pm 1$	$\pm 1$	+1	$\mp 1$	+1	$\mp 1$

+ modifications to the minimal Kähler

To close, I hope I have convinced you that...

Modular symmetries can...

...offer a **predictive framework** for flavour

...provide an origin for **CP violation** (CPV)

...help solve the **strong CP** problem

Modular symmetries can...

...offer a **predictive framework** for flavour

...provide an origin for **CP violation** (CPV)

...help solve the **strong CP** problem

they can arise from stringy constructions...

# Parting words (i.e. what next?)



- modular symmetry breaking as the only source of CPV?
- natural origin of mass hierarchies?
- hints of universality?
- use TD to fix Kähler, irreps, weights?
- pheno beyond masses and mixing?

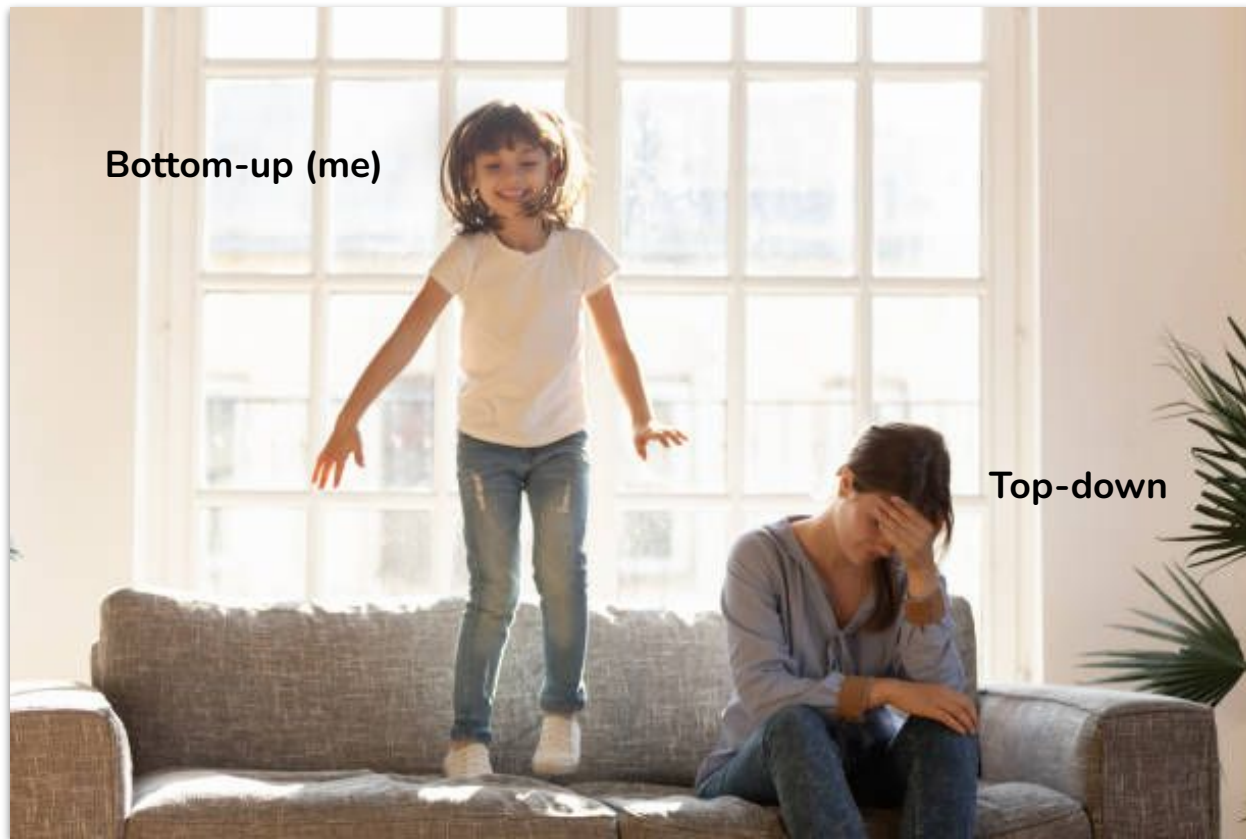
# Parting words (i.e. what next?)



- modular symmetry breaking as the only source of CPV?
- natural origin of mass hierarchies?
- hints of universality?
- use TD to fix Kähler, irreps, weights!
- pheno beyond masses and mixing?



# State of the art



See next talk, by H. P. Nilles

Obrigado!



Backup slides

# Modular-invariant SUSY actions

Ferrara et al, '89

$$W(\psi; \tau) = \sum_n \sum_{\{i_1, \dots, i_n\}} \sum_s g_{i_1 \dots i_n, s} (Y_{i_1 \dots i_n, s}(\tau) \psi_{i_1} \dots \psi_{i_n})_{\mathbf{1}, s}$$

$$\mathcal{S} = \int d^4x d^2\theta d^2\bar{\theta} K(\psi, \bar{\psi}; \tau, \bar{\tau}) + \int d^4x d^2\theta W(\psi; \tau) + \text{h.c.}$$

$\tau$  is a dimensionless spurion: once its value is fixed, it **parameterizes all** modular sym. breaking

One may argue that  $Y$ 's play the role of flavons, but structures are **completely fixed** given the modulus VEV

# SUSY breaking effects?



- **RGEs & threshold corrections** need to be considered, depend on  $\tan \beta$  and unknown SUSY spectrum
- **SUSY-breaking** corrections can be made negligible via separation of scales (power counting argument)
- Under reasonable conditions, predictions may be unaffected

Feruglio and Criado [1807.01125]

# What about... the Kähler?

- **Not holomorphic:** unconstrained by the symmetry!
- Under a modular transformation, invariant up to:

$$K(\chi_i, \bar{\chi}_i; \tau, \bar{\tau}) \rightarrow K(\chi_i, \bar{\chi}_i; \tau, \bar{\tau}) + f(\chi_i; \tau) + f(\bar{\chi}_i; \bar{\tau})$$

- Minimal choice:

$$K(\chi_i, \bar{\chi}_i; \tau, \bar{\tau}) = -h \Lambda_0^2 \log(-i(\tau - \bar{\tau})) + \sum_i \frac{|\chi_i|^2}{(-i(\tau - \bar{\tau}))^{k_i}}$$

impacts pheno → should be justified from the top-down

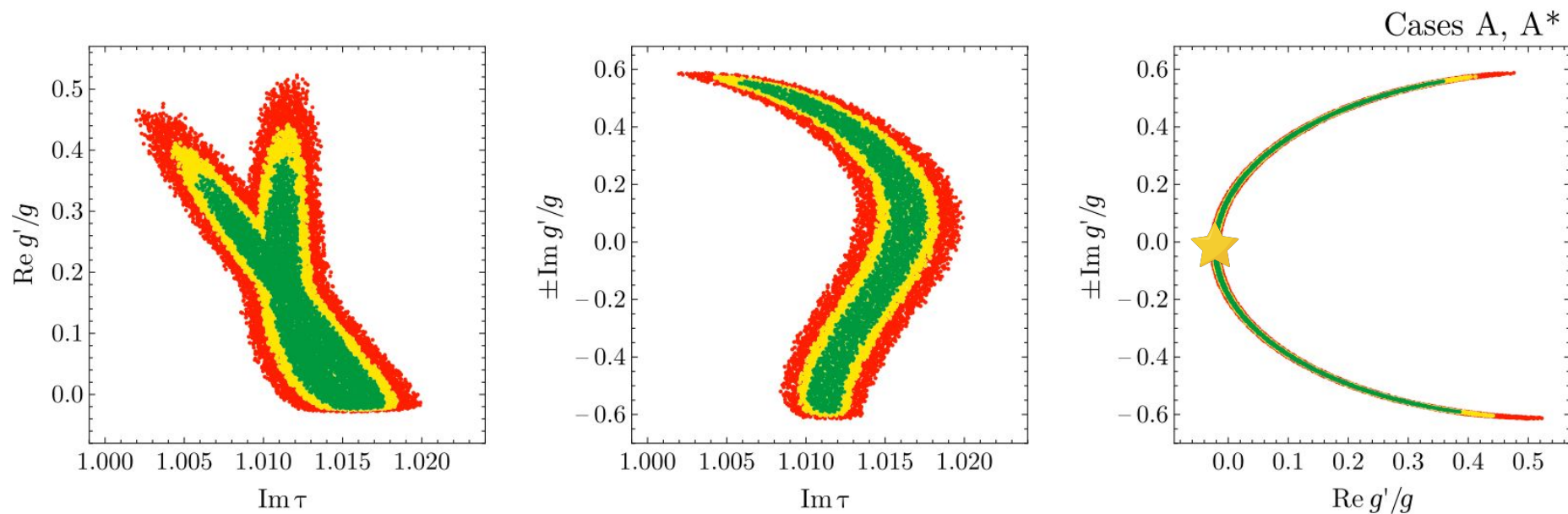
Chen, Ramos-Sánchez and Ratz [\[1909.06910\]](#)

- Further constraints may arise from the (unavoidable...) combination of modular + traditional flavour symmetries

Nilles, Ramos-Sanchez, Vaudrevange [\[2004.05200\]](#)



# Correlations between parameters in the first $S_4$ example model



Novichkov, JP, Petcov, Titov [1811.04933]

# The QCD angle is holomorphic

Furthermore, extra non-minimal kinetic terms are possible, because the  $3 \times 3$  kinetic matrices  $Z_f(\tau, \tau^\dagger)$  of fermions  $f = \{u_R, d_R, Q\}$  are not holomorphic in  $\tau$ , and modular invariance allows them to depend on the CP-violating parameters  $\tau, \tau^\dagger$  in new ways. These non-minimal kinetic terms reduce the predictive power of flavour models based on modular symmetries [28, 41–43] and are often assumed to be negligible.

Such extra complex terms are not a problem for our proposed interpretation of the QCD problem,  $\bar{\theta} = 0$ . Indeed each kinetic matrix  $Z_f$  can be brought to canonical form via a general linear transformation of the three generations of  $f_{1,2,3}$  quarks: a linear transformation affects both  $\arg \det M_q$  and  $\theta_{\text{QCD}}$  (via the anomaly) but leaves the physical combination  $\bar{\theta}$  invariant. Furthermore, these linear transformations can be chosen in ways that leave  $\arg \det M_q$  and  $\theta_{\text{QCD}}$  separately invariant, by decomposing each kinetic matrix  $Z_f$  either as  $Z_f = H_f^\dagger H_f$  (where  $H_f$  is an hermitian matrix, see e.g. [44]) or as  $Z_f = V_f^\dagger \Delta_f^2 V_f$  (where  $\Delta_f$  is a diagonal matrix with real positive entries and  $V_f$  is a product of 3 complex rotations with unit determinant). The consequent linear transformation of quark fields affects their masses and mixings (including the CKM phase) without affecting  $\arg \det M_q$ .

This discussion shows that, unlike fermion masses and mixing angles, the physical  $\bar{\theta}$  angle is a holomorphic quantity completely insensitive to the Kähler potential and can be effectively constrained by modular invariance alone, at least in the limit of unbroken supersymmetry.



# Modular anomalies

## Canonical normalisation

$$K \supset \frac{\Phi^\dagger \Phi}{(-i\tau + i\tau^\dagger)^{k_\Phi}} = \Phi_{\text{can}}^\dagger \Phi_{\text{can}} \quad \Phi_{\text{can}} = \{\phi_{\text{can}}, \psi_{\text{can}}\}$$

$$\psi_{\text{can}} \rightarrow \left( \frac{c\tau + d}{c\tau^\dagger + d} \right)^{-\frac{k_\Phi}{2}} \psi_{\text{can}} = e^{-ik_\Phi \alpha(\tau)} \psi_{\text{can}} \quad \alpha(\tau) = \arg(c\tau + d)$$

Modular symmetry acts on canonically normalised fields as a  **$\tau$ -dependent phase rotation** (with  $\tau = \tau(x)$ )

## Conditions for modular-gauge anomaly cancellation

$$\text{SU}(3)_C : \quad A \equiv \sum_{i=1}^3 \left( 2k_{Q_i} + k_{u_{Ri}} + k_{d_{Ri}} \right) = 0$$

$$\text{SU}(2)_L : \quad \sum_{i=1}^3 \left( 3k_{Q_i} + k_{L_i} \right) + k_{H_u} + k_{H_d} = 0$$

$$\text{U}(1)_Y : \quad \sum_{i=1}^3 \left( k_{Q_i} + 8k_{u_{Ri}} + 2k_{d_{Ri}} + 3k_{L_i} + 6k_{e_{Ri}} \right) + 3 \left( k_{H_u} + k_{H_d} \right) = 0$$

# tests of modulus couplings

G-J. Ding, FF,  
2003.13448

non standard neutrino interactions

$$\mathcal{L} = i \sum_{f=e,e^c,\nu} \bar{f} \bar{\sigma}^\mu \partial_\mu f + \frac{1}{2} \partial_\mu \varphi_\alpha \partial^\mu \varphi_\alpha - \frac{1}{2} M_\alpha^2 \varphi_\alpha^2 - (m_e + Z_\alpha^e \varphi_\alpha) e^c e - \frac{1}{2} \nu (m_\nu + Z_\alpha^\nu \varphi_\alpha) \nu + h.c. + \dots$$

$$\tau = \langle \tau \rangle + \frac{\varphi_u + i \varphi_v}{\sqrt{2}}$$

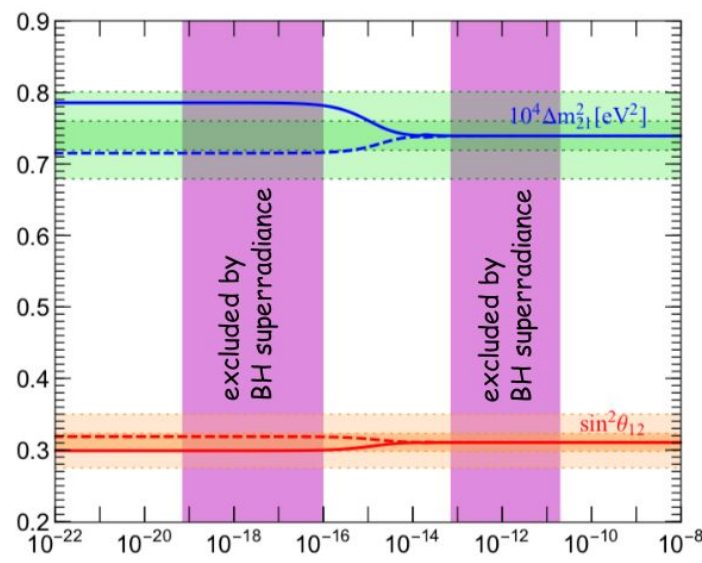


in medium with non-zero electron number density

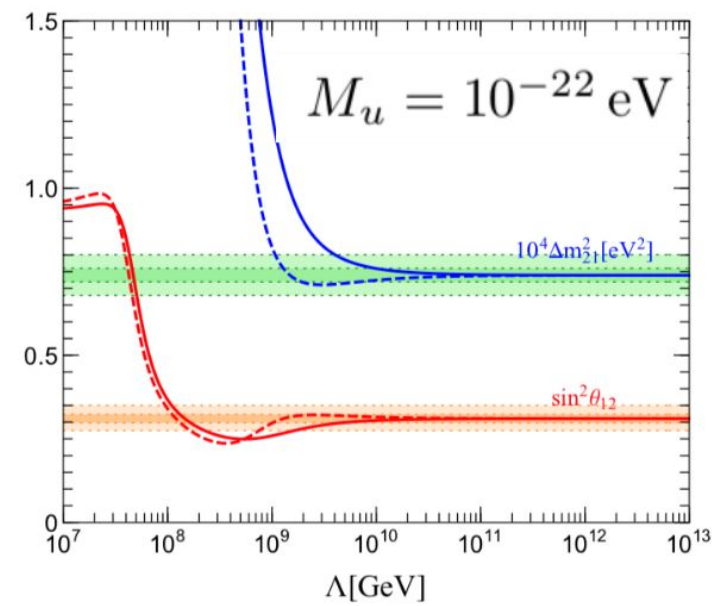
small, unless the modulus is very light

$$\delta m_\nu(0) = -n_e^0 \frac{\text{Re}(Z^e) Z^\nu}{M^2(R)},$$

in the sun:



$$\Lambda = 5 \times 10^9 \text{ GeV} \quad \begin{matrix} M_u [\text{eV}] \\ [\text{modulus VEV}] \end{matrix}$$



from Feruglio's slides at Bethe Workshop, citing [2003.13448]

# But what if I normalize *this* way?

A comment on normalizations

$$W \supset \sum_i \left( \psi \alpha_i Y_i^{(K)} \psi^c \right)_{\mathbf{1}} H_q$$

mod. form of weight  $K$

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A comment on normalizations

$$W \supset \sum_i \left( \psi \alpha_i Y_i^{(K)} \psi^c \right)_{\mathbf{1}} H_q$$

mod. form of weight  $K$

<b>1</b>	<b>100</b>
<b>10</b>	<b>10</b>
<b>100</b>	<b>1</b>
$ \alpha $	$\ Y\ _{\mathbb{E}}$

**Same model predictions!**

*how can we discuss natural  $\alpha$ 's?*

*how do we interpret a hierarchy between  $\alpha$ 's?*

*how can we claim modular symmetries are responsible for hierarchies, not  $\alpha$ 's?*

(norms of  $Y$ 's are not fixed by group theory)

# But what if I normalize *this* way?

A comment on normalizations

can trust these:

$$Y_3 \sim \begin{pmatrix} 1 \\ \epsilon \\ \epsilon^2 \end{pmatrix}$$

cannot trust these:

$$M \sim \left( \begin{array}{cc|c} 1 & \epsilon & 1 \\ \epsilon & \epsilon^2 & \epsilon \\ \hline \epsilon & \epsilon^2 & \epsilon \end{array} \right)$$

*how can we claim modular symmetries are responsible for hierarchies, not  $\alpha$ 's?*

discussion in Varzielas, Levy, JP, Petcov [2307.14410]

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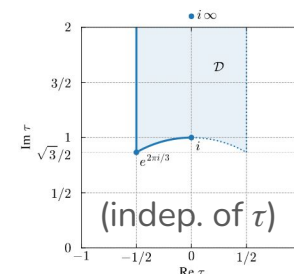
# A normalization proposal / choice

Petcov [2311.04185], D. Zagier (1981)

A “global” normalisation based on the Petersson inner product

$$N \left[ Y^{(K)} \right]^2 \equiv \iint_{\mathcal{D}} \sum_i \left| Y_i^{(K)}(x + iy) \right|^2 (2y)^K \frac{dx dy}{y^2} \stackrel{!}{=} 1$$

different prescription for non-cusp forms (yet another if  $K=1$ )



- Is there a general top-down recipe?
- Basis ambiguity if there are several forms of the same weight and irrep

# Lessons from eclectic flavor symmetries

see [1901.03251, 1908.00805, 2001.01736, 2004.05200, 2006.03059, 2112.06940]  
 and Nilles, Ramos-Sánchez [2404.16933] for a recent summary

- There is no possible scheme with just modular flavor symmetries
- Also discrete R-symmetries seem unavoidable
- A limited type of groups appear (e.g.  $T'$ )

nature of symmetry		outer automorphism of Narain space group	flavor groups				
eclectic	modular	rotation $S \in \text{SL}(2, \mathbb{Z})_T$ rotation $T \in \text{SL}(2, \mathbb{Z})_T$	$\mathbb{Z}_4$ $\mathbb{Z}_3$	$T'$			$\Omega(2)$
	traditional flavor	translation A translation B	$\mathbb{Z}_3$ $\mathbb{Z}_3$	$\Delta(27)$	$\Delta(54)$	$\Delta'(54, 2, 1)$	
		rotation $C = S^2 \in \text{SL}(2, \mathbb{Z})_T$	$\mathbb{Z}_2^R$				
		rotation $R = \gamma_{(3)} \in \text{SL}(2, \mathbb{Z})_U$	$\mathbb{Z}_9^R$				

Nilles, Ramos-Sanchez, Vaudrevange [2006.03059]



# Lessons from eclectic flavor symmetries

see [1901.03251, 1908.00805, 2001.01736, 2004.05200, 2006.03059, 2112.06940]

and Nilles, Ramos-Sánchez [2404.16933] for a recent summary

- There is no possible scheme with just modular flavor symmetries
- Also discrete R-symmetries seem unavoidable
- A limited type of groups appear (e.g.  $T'$ )
- Only some weights and irreps available at low energy
- Weights (typically fractional) correlated with irreps

sector	matter fields $\Phi_n$	eclectic flavor group $\Omega(2)$								$\mathbb{Z}_9^R$ $R$
		modular $T'$ subgroup				traditional $\Delta(54)$ subgroup				
		irrep $\mathbf{s}$	$\rho_{\mathbf{s}}(\text{S})$	$\rho_{\mathbf{s}}(\text{T})$	$n$	irrep $\mathbf{r}$	$\rho_{\mathbf{r}}(\text{A})$	$\rho_{\mathbf{r}}(\text{B})$	$\rho_{\mathbf{r}}(\text{C})$	
bulk	$\Phi_0$	$\mathbf{1}$	1	1	0	$\mathbf{1}$	1	1	+1	0
	$\Phi_{-1}$	$\mathbf{1}$	1	1	-1	$\mathbf{1}'$	1	1	-1	3
$\theta$	$\Phi_{-2/3}$	$\mathbf{2}' \oplus \mathbf{1}$	$\rho(\text{S})$	$\rho(\text{T})$	$-2/3$	$\mathbf{3}_2$	$\rho(\text{A})$	$\rho(\text{B})$	$+\rho(\text{C})$	1
	$\Phi_{-5/3}$	$\mathbf{2}' \oplus \mathbf{1}$	$\rho(\text{S})$	$\rho(\text{T})$	$-5/3$	$\mathbf{3}_1$	$\rho(\text{A})$	$\rho(\text{B})$	$-\rho(\text{C})$	-2
$\theta^2$	$\Phi_{-1/3}$	$\mathbf{2}'' \oplus \mathbf{1}$	$(\rho(\text{S}))^*$	$(\rho(\text{T}))^*$	$-1/3$	$\bar{\mathbf{3}}_1$	$\rho(\text{A})$	$(\rho(\text{B}))^*$	$-\rho(\text{C})$	2
	$\Phi_{+2/3}$	$\mathbf{2}'' \oplus \mathbf{1}$	$(\rho(\text{S}))^*$	$(\rho(\text{T}))^*$	$+2/3$	$\bar{\mathbf{3}}_2$	$\rho(\text{A})$	$(\rho(\text{B}))^*$	$+\rho(\text{C})$	5
super-potential	$\mathcal{W}$	$\mathbf{1}$	1	1	-1	$\mathbf{1}'$	1	1	-1	3

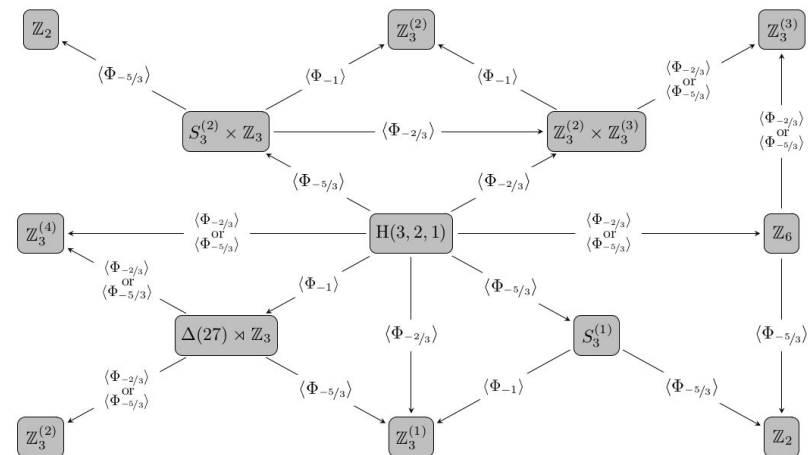
# Lessons from eclectic flavor symmetries

see [1901.03251, 1908.00805, 2001.01736, 2004.05200, 2006.03059, 2112.06940]  
and Nilles, Ramos-Sánchez [2404.16933] for a recent summary

- There is no possible scheme with just modular flavor symmetries
- Also discrete R-symmetries seem unavoidable
- A limited type of groups appear (e.g.  $T$ )
- Only some weights and irreps available at low energy
- Weights (typically fractional) correlated with irreps
- Breaking reintroduces flavons :(



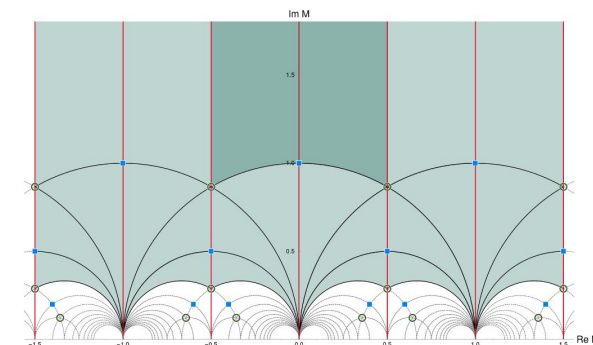
Baur et al. [2112.06940]



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- Both Kähler and superpotential play a crucial role
- Larger fundamental domains ( $\Gamma(N)$  instead of  $\Gamma?$ )
- Top-down and bottom-up do **not yet** meet



but there are a few BU attempts:  
Chen et al. [2108.02240]; Ding et al. [2303.02071];  
Li, Ding [2308.16901]; Li, Lu, Ding [2405.13460]

# Moduli stabilization



early attempts: [1909.05139, 1910.11553]

# Simplest modular-invariant potentials?

- Studied by Cvetič, Font, Ibáñez, Lüst and Quevedo (1991)  
 $\mathcal{N} = 1$  SUGRA

$$V(\tau, \bar{\tau}) = \frac{\Lambda_V^4}{8(\text{Im } \tau)^3 |\eta|^{12}} \left[ \frac{4}{3} \left| iH' + \frac{3}{2\pi} H \hat{G}_2 \right|^2 (\text{Im } \tau)^2 - 3|H|^2 \right]$$

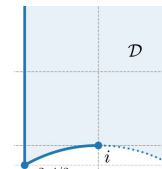
$$H(\tau) = (j(\tau) - 1728)^{m/2} j(\tau)^{n/3}$$

$$W(\tau) = \Lambda_W^3 \frac{H(\tau)}{\eta(\tau)^6}$$

$$m, n = 0, 1, 2, \dots$$

- This potential is **modular-** and **CP-invariant**
- Simplified model, independent of the level  $N$

# $q$ - and $u$ -expansions of $\eta$



$$|q| \leq e^{-\sqrt{3}\pi} \simeq 0.004$$

$$\eta = q^{1/24} \sum_{n=-\infty}^{\infty} (-1)^n q^{\frac{3n^2-n}{2}} = q^{1/24} (1 - q - q^2 + q^5 + q^7 - q^{12} - q^{15} + \mathcal{O}(q^{22}))$$

$$u \equiv \frac{\tau - \omega}{\tau - \omega^2}$$

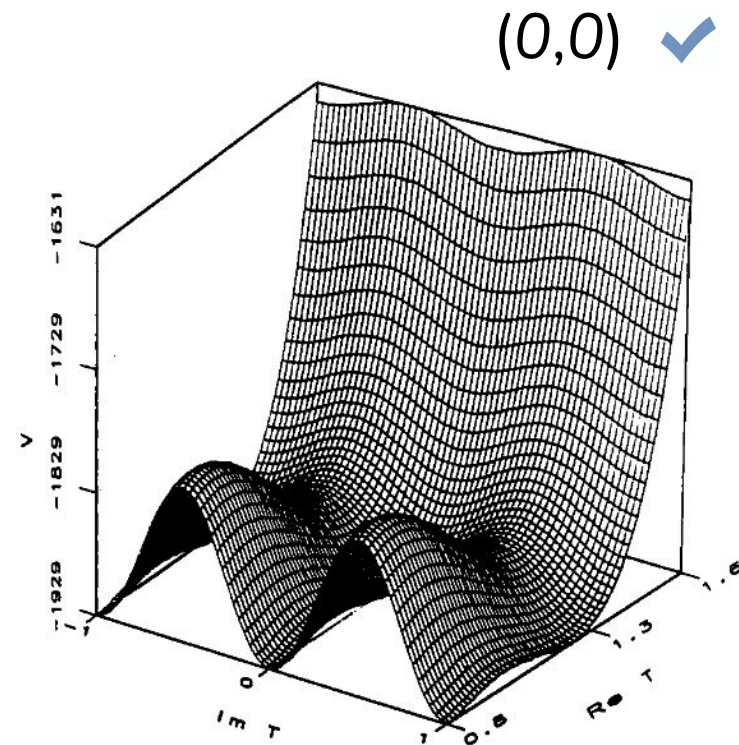
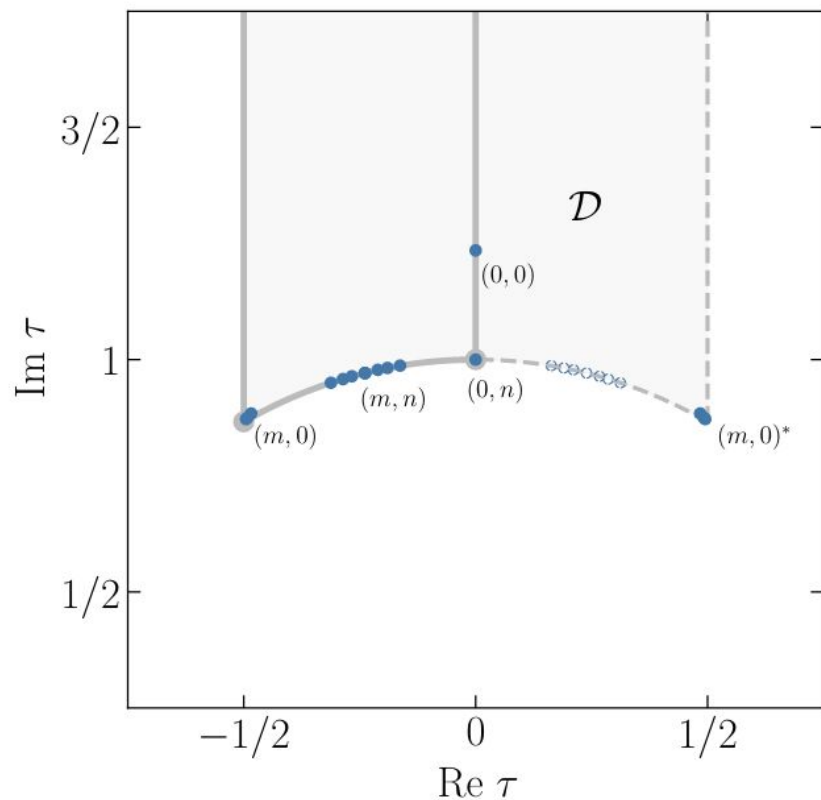
$$\tilde{\eta}(u) \equiv \frac{\eta(u)}{\sqrt{1-u}}$$

$$u \xrightarrow{ST} \omega^2 u$$

$$\tilde{\eta}(u) \xrightarrow{ST} \tilde{\eta}(u)$$

$$\begin{aligned} \tilde{\eta}(u) &\simeq e^{-i\pi/24} (0.800579 - 0.573569u^3 - 0.780766u^6 - 0.150007u^9) + \mathcal{O}(u^{12}) \\ &\equiv e^{-i\pi/24} (\tilde{\eta}_0 + \tilde{\eta}_3 u^3 + \tilde{\eta}_6 u^6 + \tilde{\eta}_9 u^9) + \mathcal{O}(u^{12}), \end{aligned}$$

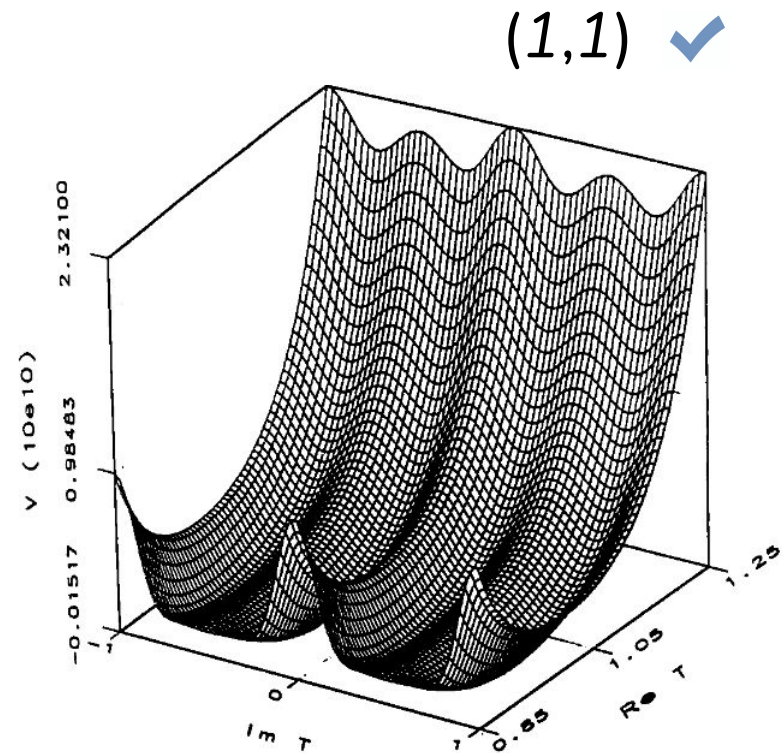
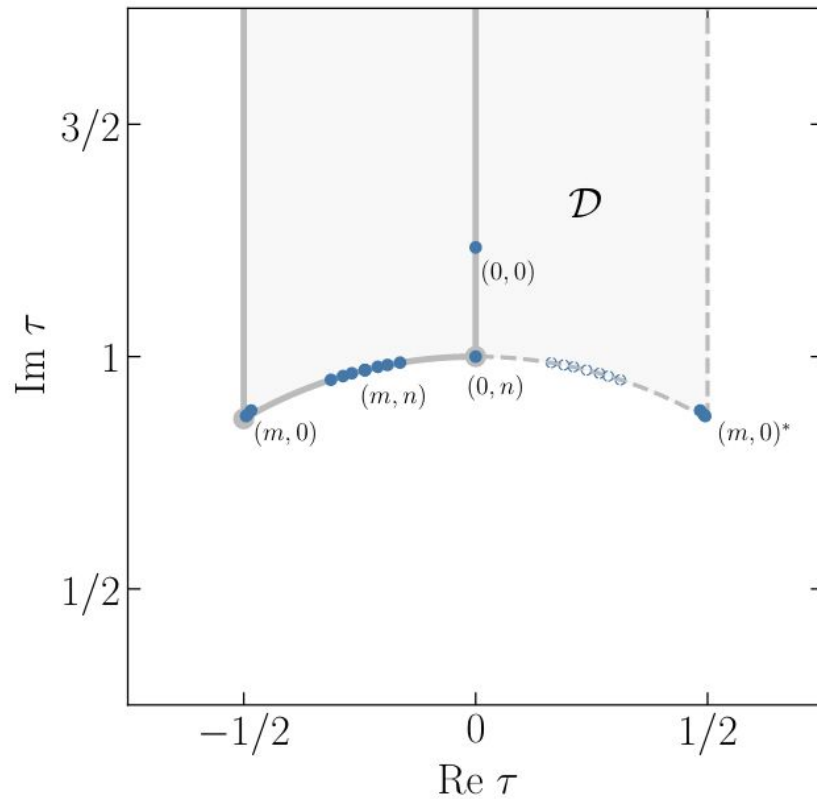
# Global minima for $(m,n)$ -potentials



“(...) we conjecture that all extrema of  $V$  entirely lie on [the boundary].” — Cvetič et al.

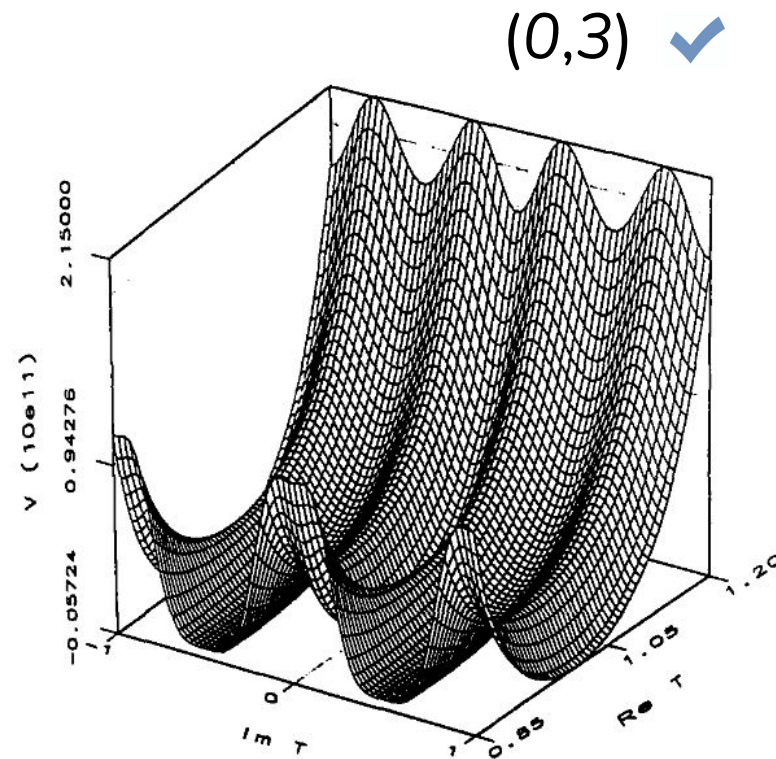
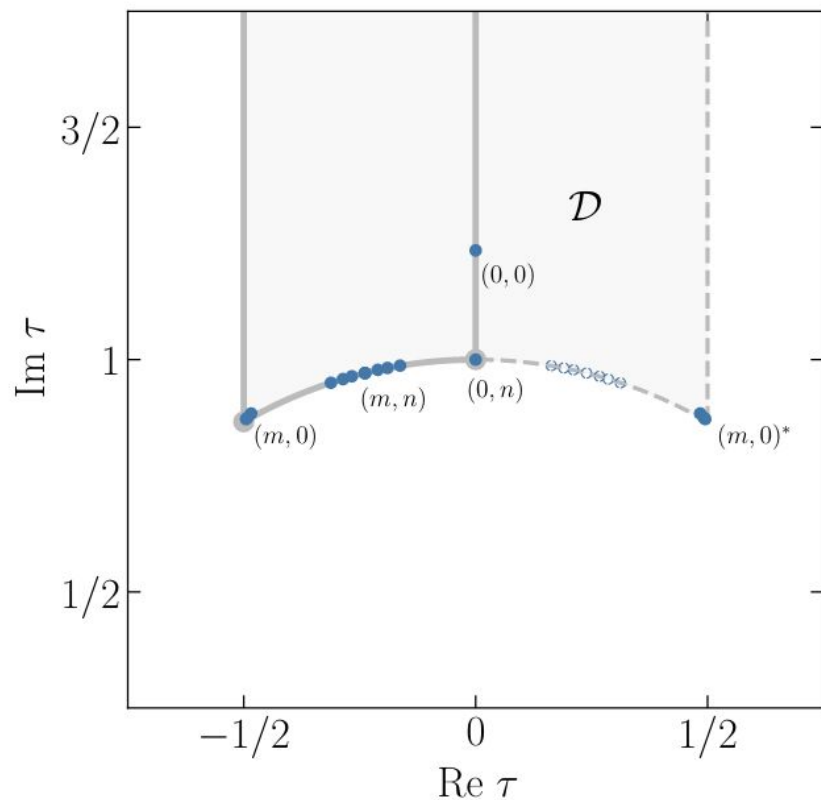


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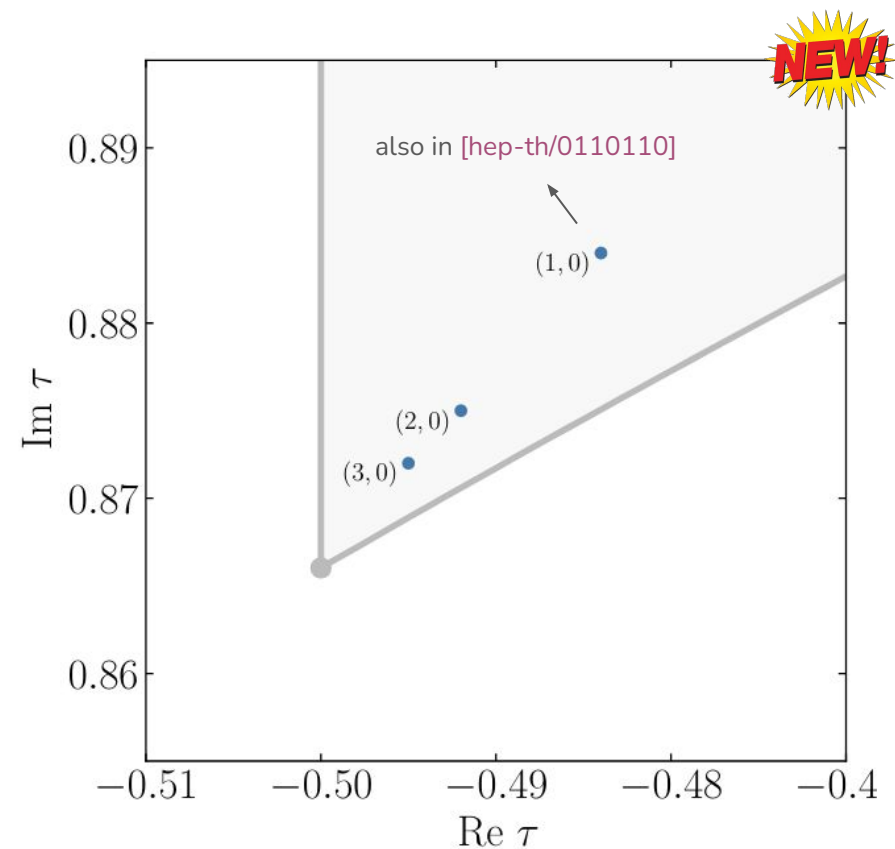
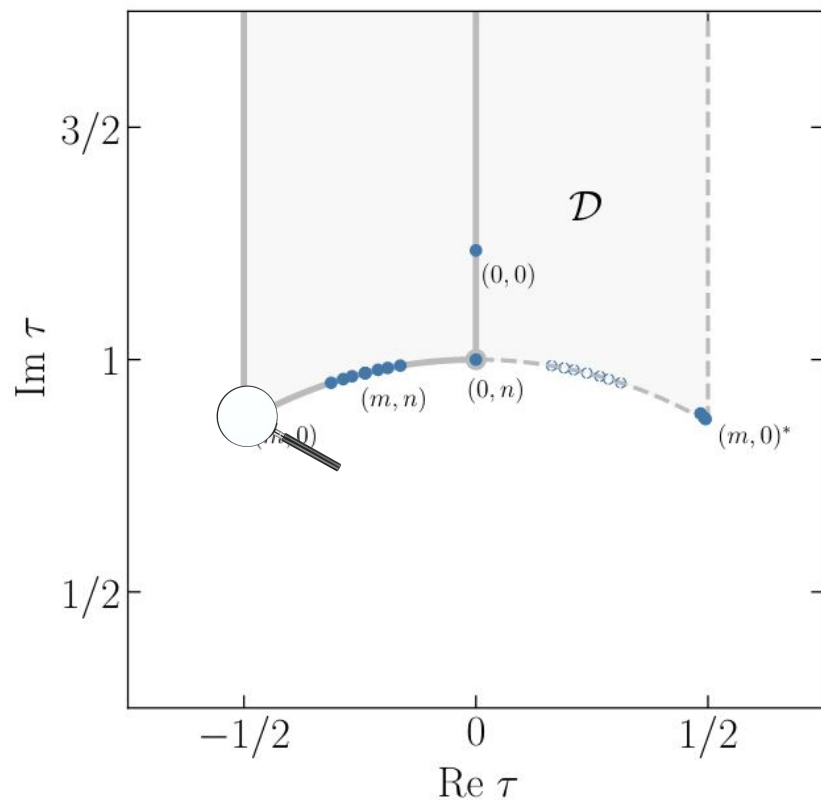
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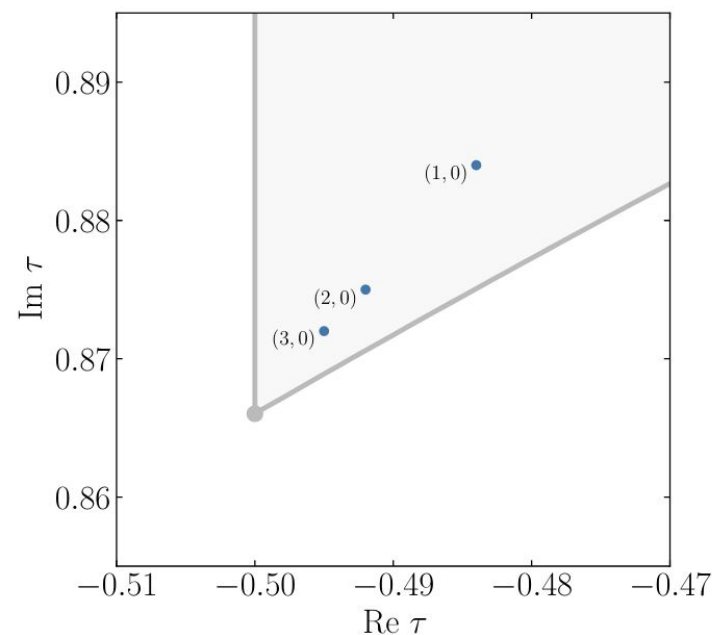
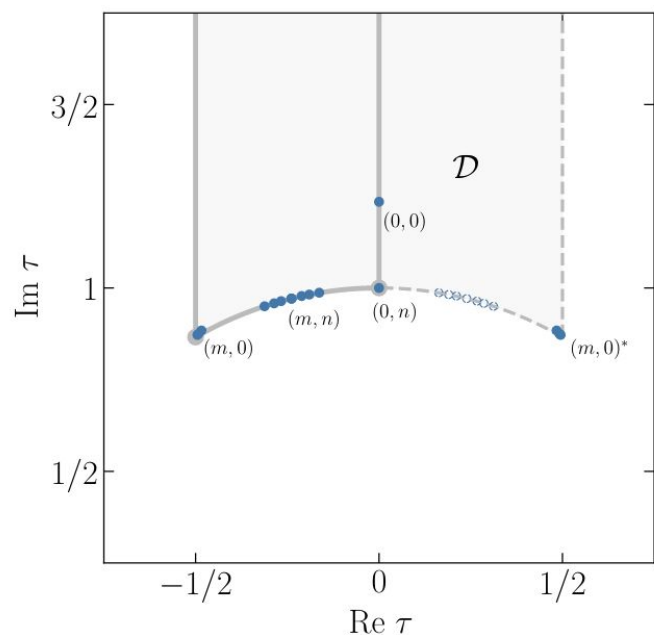
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# Global minima for $(m,n)$ -potentials



“(...) we conjecture that all extrema of  $V$  entirely lie on [the boundary].” — Cvetič et al.

these results later confirmed by Leedom, Righi, Westphal [2212.03876]



$(\mathbf{0}, \mathbf{0})$  is a single minimum at  $\tau \simeq 1.2i$  on the imaginary axis, corresponding to the case  $m = n = 0$ ;

$(\mathbf{0}, n)$  is a single minimum at the symmetric point  $\tau = i$  attained when  $m = 0, n \neq 0$ ;

$(m, \mathbf{0})$  and  $(m, \mathbf{0})^*$  are a pair of degenerate minima for each  $m \neq 0$  and  $n = 0$ :  $(m, 0)$  is located in the vicinity of the left cusp  $\tau = \omega$ , approaching this symmetric point as  $m$  increases, while  $(m, 0)^*$  is its CP-conjugate;

$(m, n)$  is a series of minima on the unit arc, corresponding to  $m \neq 0, n \neq 0$ ; these minima shift towards  $\tau = \omega$  ( $\tau = i$ ) along the arc as  $m$  ( $n$ ) grows.

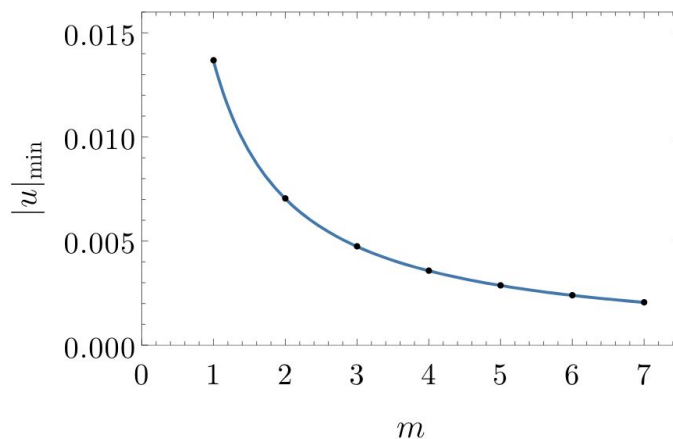
# The $(m,0)$ family of potentials

- $u$ -expand  $(m,0)$  potentials to analyse them near the left cusp

$$V_{m,0} = \Lambda_V^4 \frac{1728^m}{\sqrt{3} \tilde{\eta}_0^{12}} \left\{ -1 - 2|u|^2 + (A_m^2 - 3)|u|^4 \right\} + \mathcal{O}(|u|^6)$$

- “Mexican”-hat potential  
(cusp is a maximum!)

$$A_m \equiv \frac{864 |\tilde{\eta}_3|^3}{\pi^6 \tilde{\eta}_0^{27}} m + \frac{6 |\tilde{\eta}_3|}{\tilde{\eta}_0} \\ \simeq 68.78 m + 4.30$$



$$|u|_{\min} \simeq (A_m^2 - 3)^{-1/2} \\ \simeq A_m^{-1} = \frac{0.0145}{\boxed{m} + 0.0625}$$

# The $(m,0)$ family of potentials

(phase dependence)

$$u = |u|e^{i\phi}$$

- $u$ -expanding to higher order shows dependence on  $\phi \in [-\pi/3, 0]$

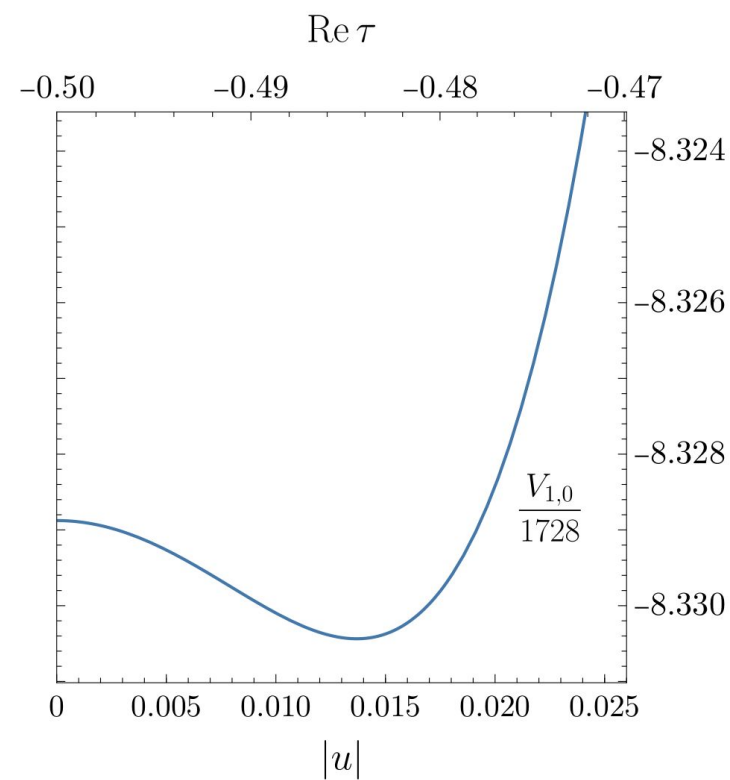
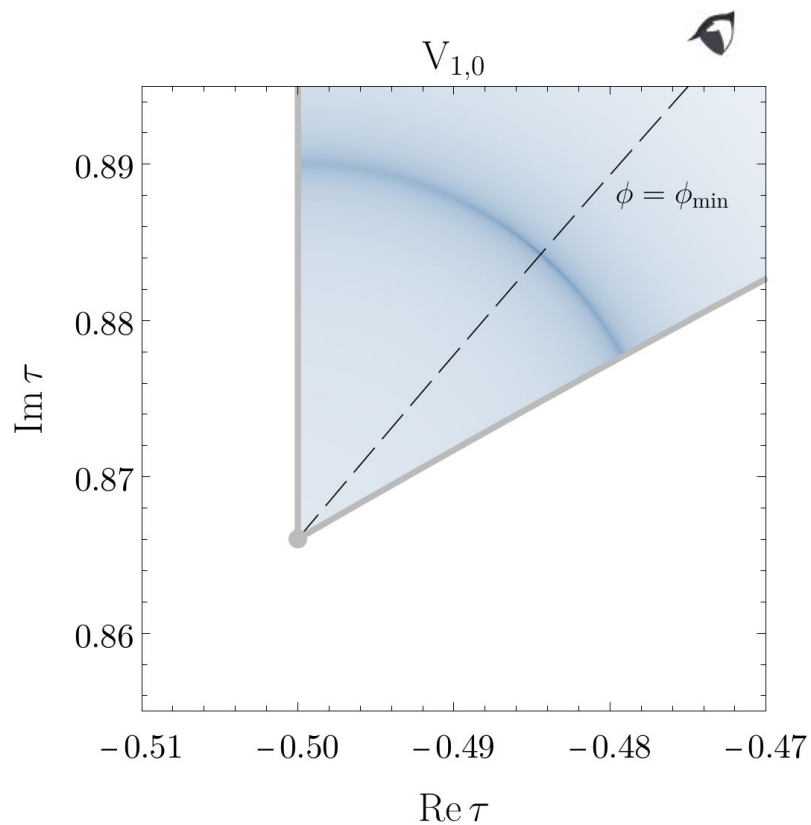
$$V_{m,0} \propto -1 - 2|u|^2 + (A_m^2 - 3)|u|^4 + (-4 + 2A_m^2 + B_m^2 \cos 6\phi)|u|^6 \\ + 2A_mB_m^2 \cos 3\phi |u|^7 + (-5 + 3A_m^2 + 2B_m^2 \cos 6\phi)|u|^8 + \mathcal{O}(|u|^9)$$

$$B_m^2 \equiv \frac{864 |\tilde{\eta}_3|^3}{\pi^6 \tilde{\eta}_0^{27}} m \left[ \frac{864 |\tilde{\eta}_3|^3}{\pi^6 \tilde{\eta}_0^{27}} (m-2) + \frac{3(31\tilde{\eta}_3^2 - 10\tilde{\eta}_0\tilde{\eta}_6)}{\tilde{\eta}_0|\tilde{\eta}_3|} \right] + \frac{6(7\tilde{\eta}_3^2 - 2\tilde{\eta}_0\tilde{\eta}_6)}{\tilde{\eta}_0^2} \\ \simeq 4730.60 m^2 - 2069.73 m + 33.26.$$

- Phase of  $u$  mostly determined by  $|u|^6$  and  $|u|^7$  terms

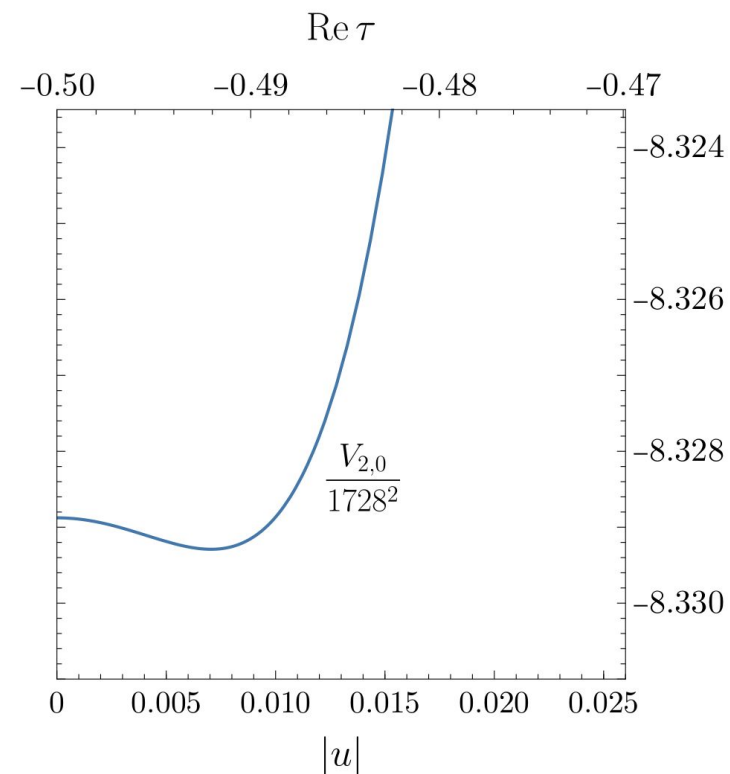
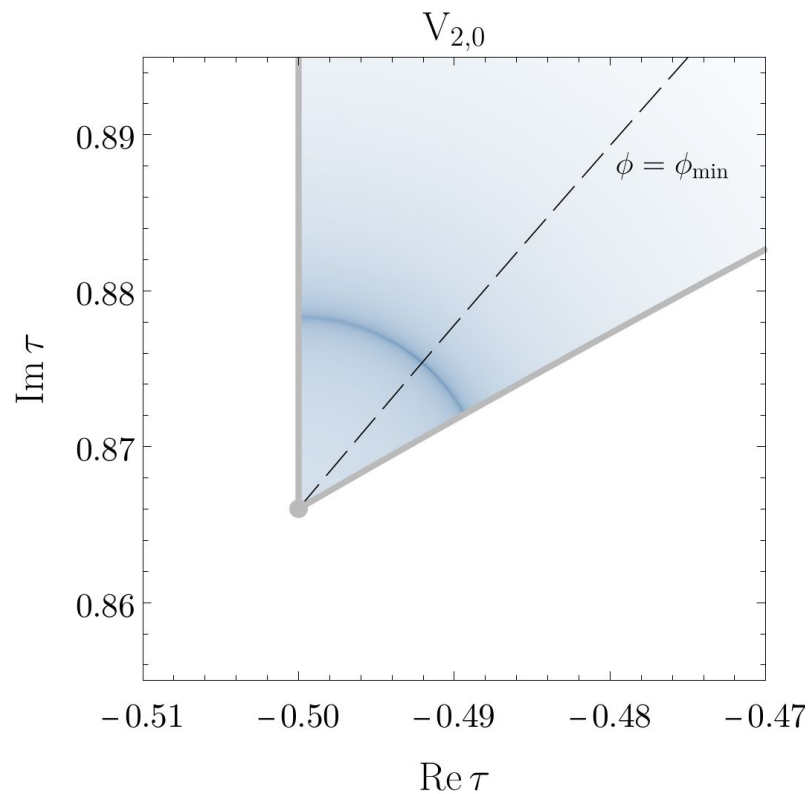
$$\phi_{\min} \simeq -\frac{2\pi}{9} = -40^\circ$$

# The $(m,0)$ family of potentials ( $m = 1$ )



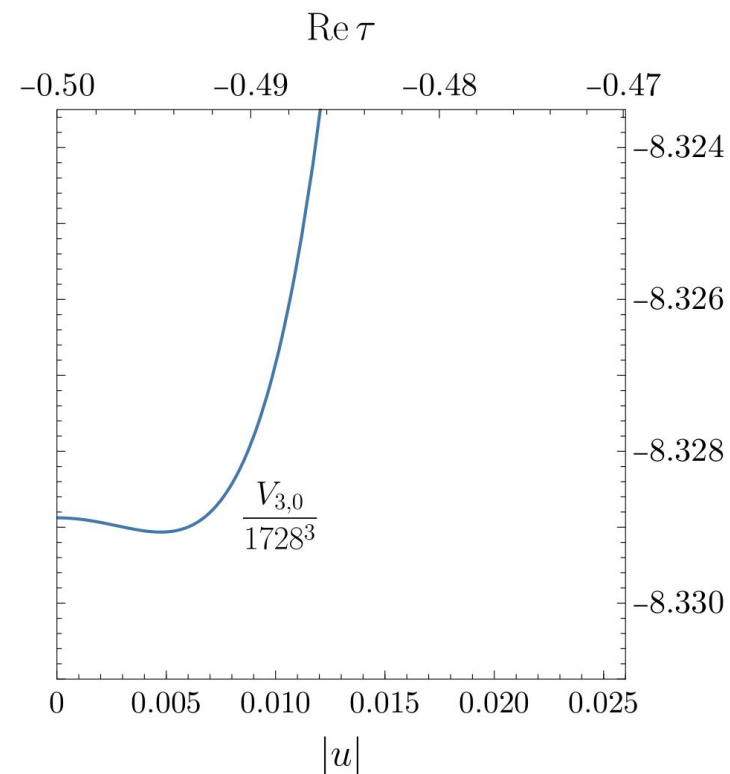
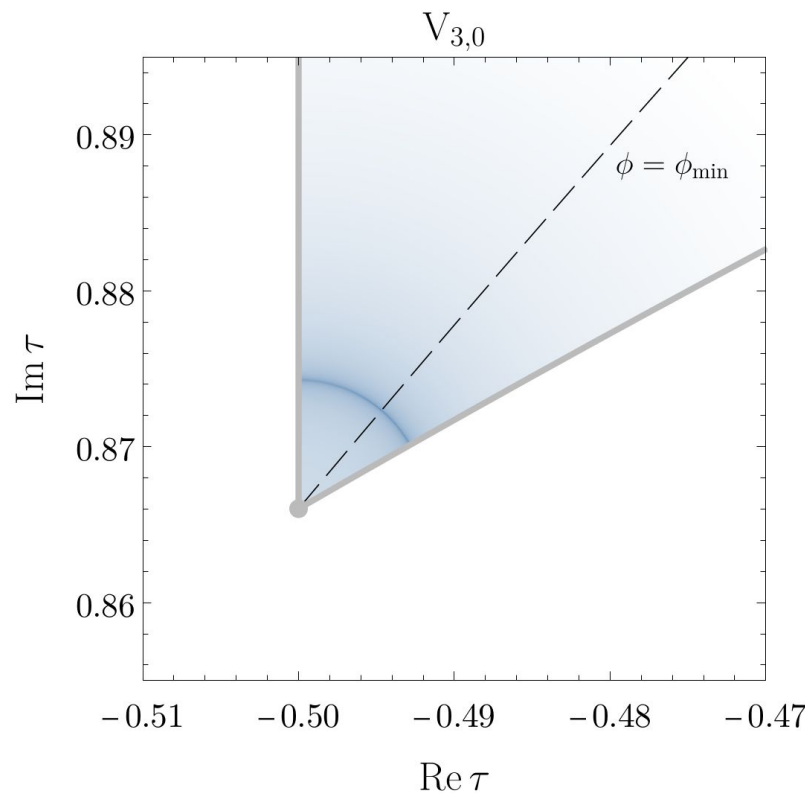
"Mexican"-hat potential

# The $(m,0)$ family of potentials ( $m = 2$ )

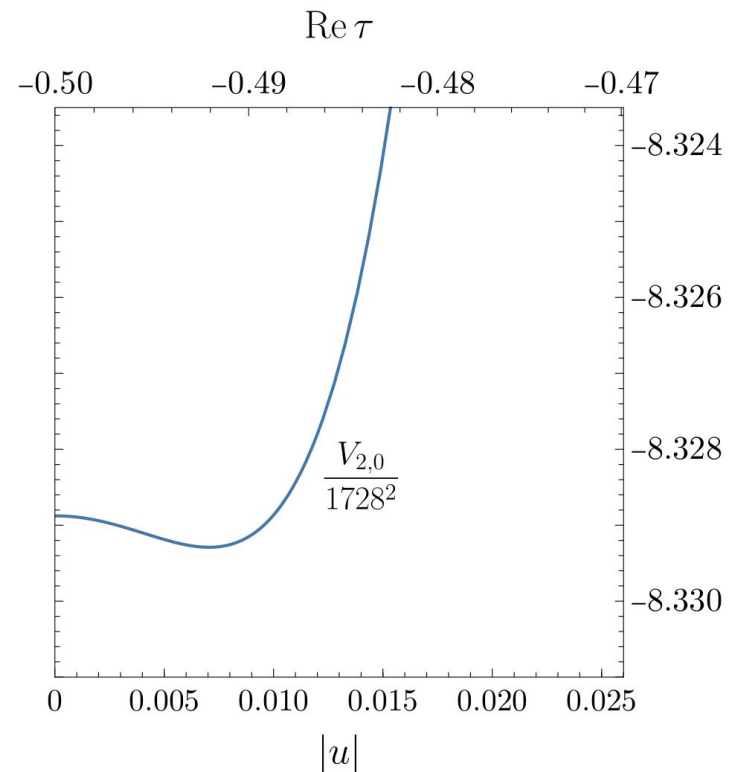
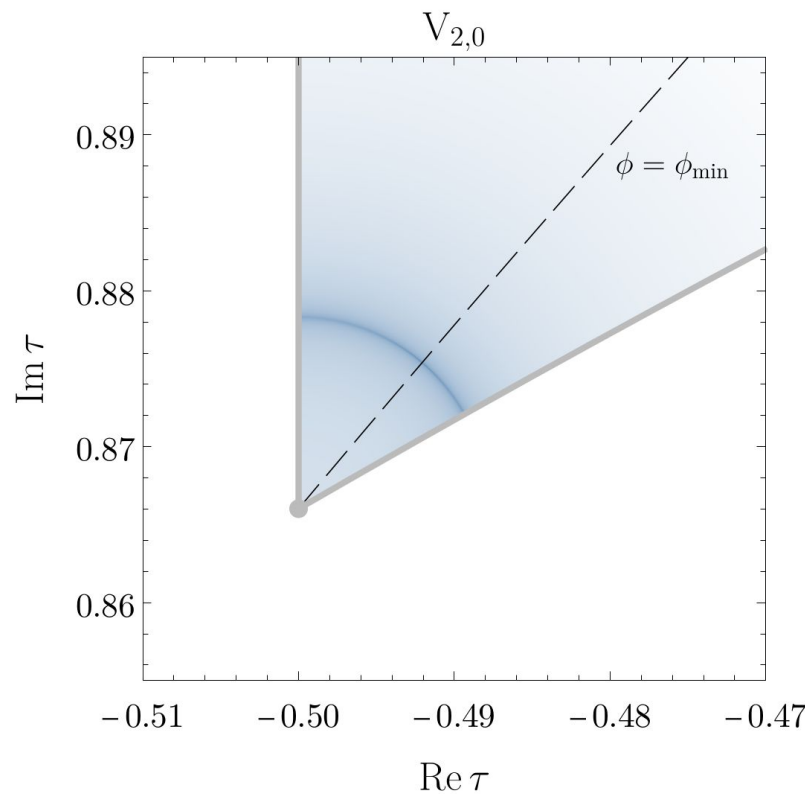




# The $(m,0)$ family of potentials ( $m = 3$ )

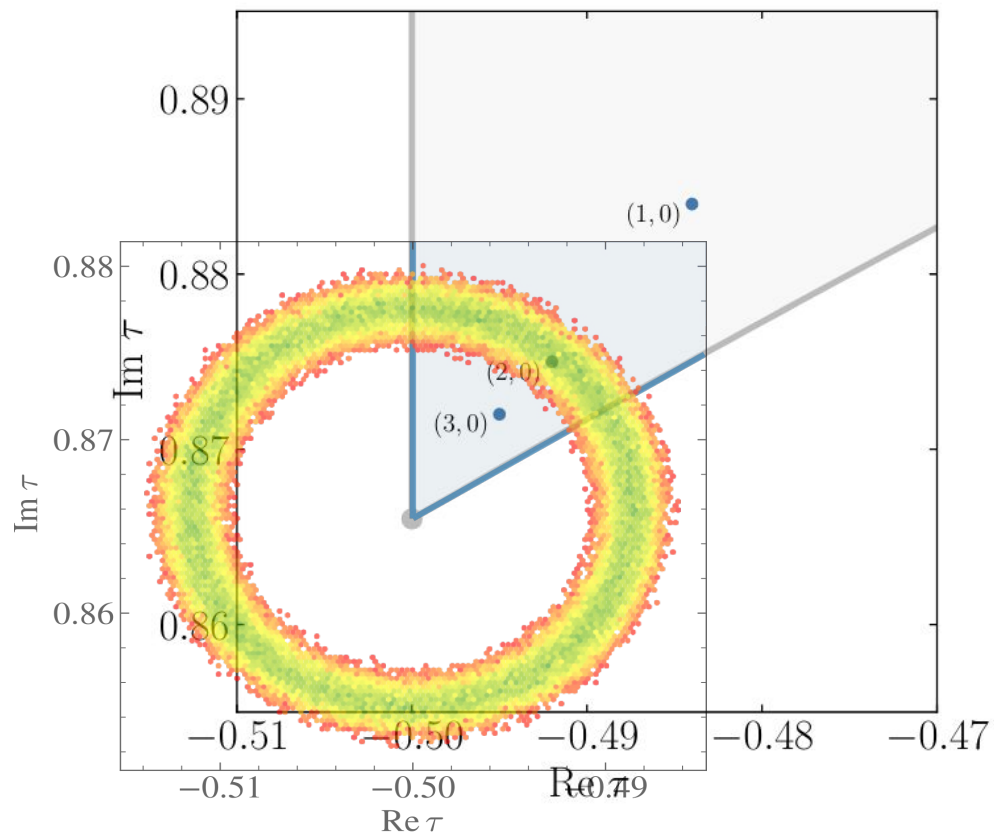


# The $(m,0)$ family of potentials ( $m = 2$ )



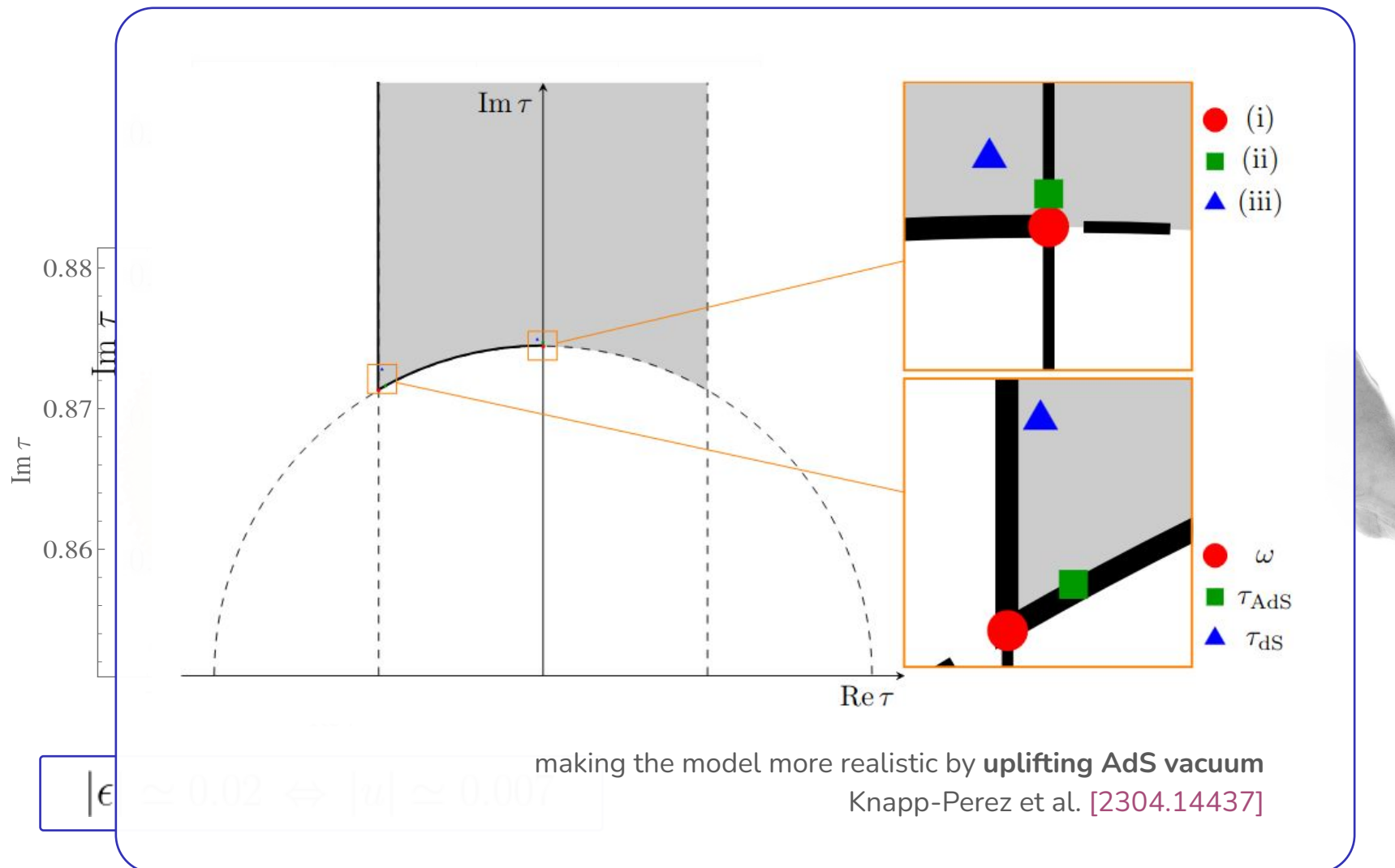
$$|\epsilon| \simeq 0.02 \Leftrightarrow |u| \simeq 0.007$$

# Matching puzzle pieces?



$$|\epsilon| \simeq 0.02 \Leftrightarrow |u| \simeq 0.007$$

# Matching puzzle pieces?



No, there is no tuning in choosing this form of the superpotential (arguably)

$$H(\tau) \propto (J(\tau) - 1)^{m/2}$$

Subset of all possible  $H(\tau)$  which vanish only at the symmetric point  $\tau=i$  (itself distinguished by modular symmetry)

$$J(\tau) \equiv j(\tau)/1728$$

# The global SUSY limit (a comment)

$$\mathfrak{n} = \kappa^2 \Lambda_K^2 \rightarrow 0$$

$$K(\tau, \bar{\tau}) = -\Lambda_K^2 \log(2 \operatorname{Im} \tau)$$

$$\kappa^2 = 8\pi/M_P^2$$

$$W(\tau) = \Lambda_W^3 H(\tau)$$

$$H(\tau) = (j(\tau) - 1728)^{m/2} j(\tau)^{n/3} \mathcal{P}(j(\tau))$$

$$V(\tau, \bar{\tau}) = \frac{4\Lambda_W^6}{\Lambda_K^2} (\operatorname{Im} \tau)^2 |H'(\tau)|^2$$

- Global minima are zeros of  $H'$
- non-trivial  $\mathcal{P}(j)$  can be engineered to produce minima at arbitrary points in the fundamental domain

