

# Nal Quenching factor measurement



Particle physics day 2023 - Jyväskylä







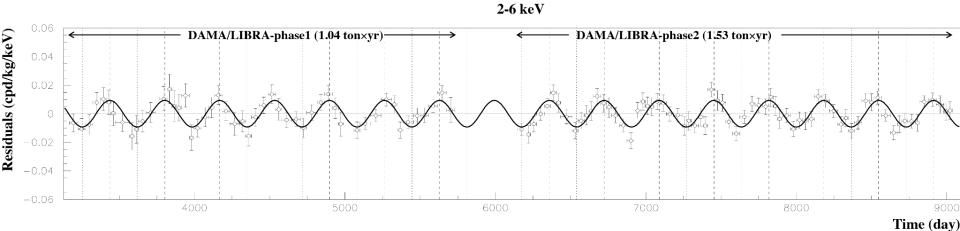


#### **Outline**

- Background: Direct detection & DAMA
- Setup
- Analysis
- Results

#### **Background: Direct detection & DAMA**

- Assume WIMPs have some interaction with ordinary matter
- Need (very) low background detectors
- Event rate above background is a dark matter signal
  - Modulation searches just takes all events and looks for a modulation
- DAMA: 13.7σ signal [1]



**Earth** 

December

June

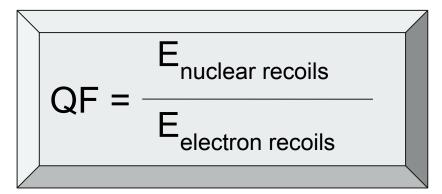
Sun

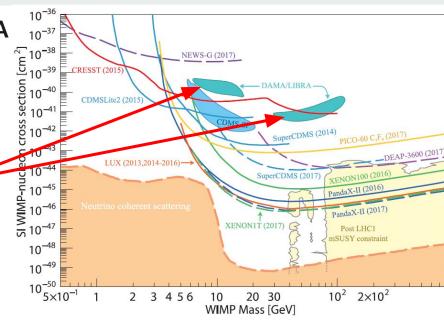
**WIMP** 

wind

#### **Background: Direct detection & DAMA**

- DAMA is located in Gran Sasso, Italy
- Ultrapure Nal crystals w. Tl dopant
- See a significant statistical excess in nuclear recoils
- Only see scintillation light
- Through light yield and Quenching factor
  we get the recoil energy

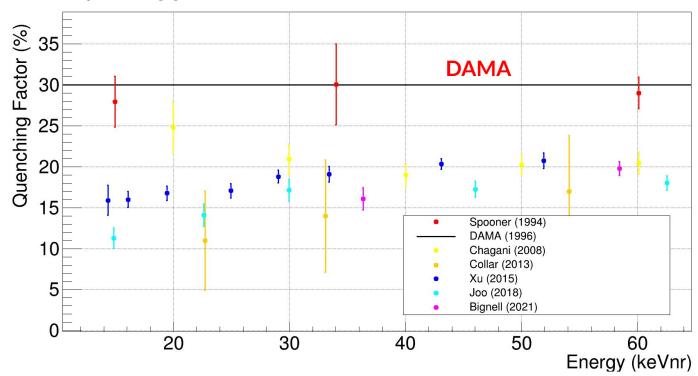




Exclusion plot of WIMP type dark matter [3]

## Background

- Different QF results [2]



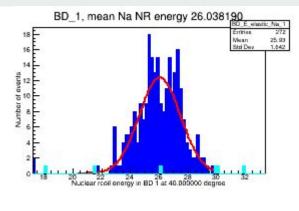
#### **Setup - Facility**

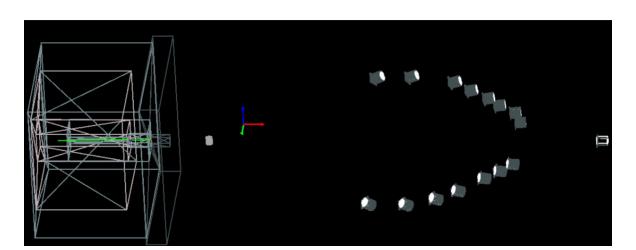
- Measurement done at Triangle Universities Nuclear Laboratory (TUNL)
- 5 Nal crystals with varying levels of Thallium dopant
- Proton accelerator -> neutrons via <sup>7</sup>Li(p,n)<sup>7</sup>Be reaction
  - Mean E ≃ 1.28MeV
- Nal w. PMT placed in front of collimator
- 15+1 Backing detectors (Liquid scintillators coupled to PMTs)
  - Angles from 7 to 40 degrees (E<sub>Na</sub> ≈ 0.8 26keV) and 0-degree detector
- Trigger scheme:
  - If BD triggers -> save cached Nal pulse



#### **Setup - Simulations**

- Simulate the recoil spectra in Geant4
- 100 Million events
- Obtain Recoil spectra for each scattering angle
- Beam profile from previous simulations
  - Working on improvements with data from 0-degree backing detectors

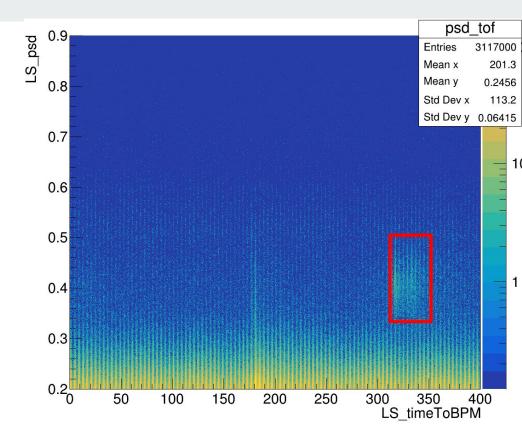




## **Analysis**

#### Part 1:

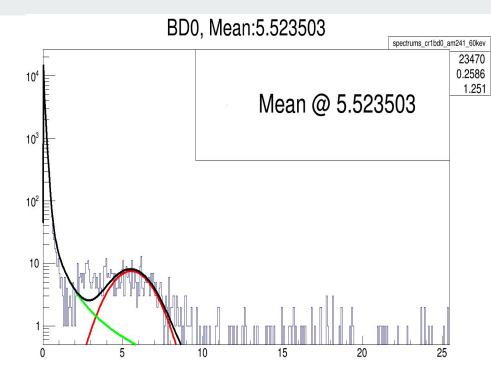
- Calibrate using different gamma sources
- Separate out neutron recoils from gamma events
- Output -> signal and background histograms



## **Analysis**

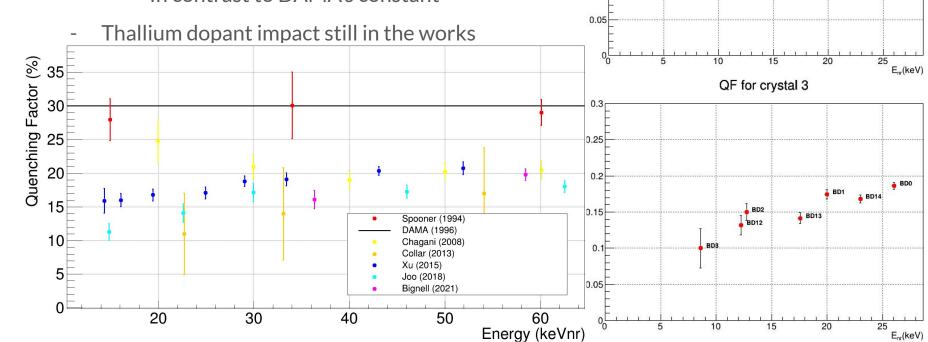
#### Part 2:

- Find the nuclear recoil peak
- Fit Gaussian + Background to signal histogram
  - Monte Carlo minimizing each potential point
  - Background is a function fit to histogram
  - Gaussian width is fixed from calibration resolution
- Compare to simulations



## Results & Thank you

- Energy dependent QFs
  - In contrast to DAMA's constant



**Quenching Factor** 

0.15

QF for crystal 1

BD2

BD12

BD11

ĕ BD13

BD0

ĕ BD14

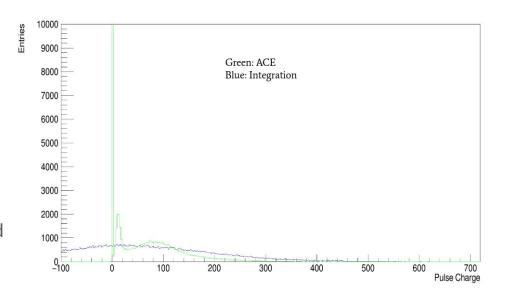
#### References

- [1] R Bernabei et al 2023 J. Phys.: Conf. Ser. 2586 012096
- [2] D. Cintas et al 2021 J. Phys.: Conf. Ser. 2156 012065
- [3] S. Hamdan, arXiv:2108.07752

## Backup/Extra slides

## **Energy integration**

- Tried two methods:
  - Integration
  - Adopted Charge Estimate
- ACE = sum everything above threshold
- ACE has better low energy resolution

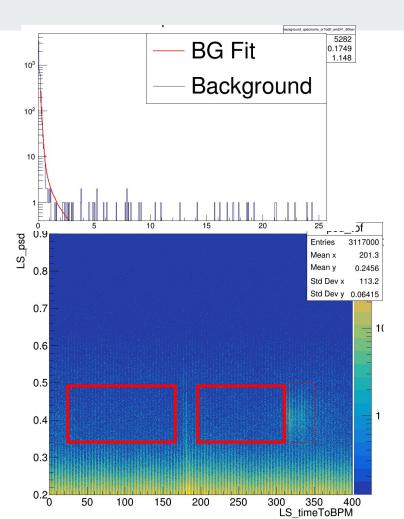


## Background

- "inverse" the cut criteria from neutron events
- Calibrate with same source and function as signal
- Save histogram
- Spectrum fitter fits function:

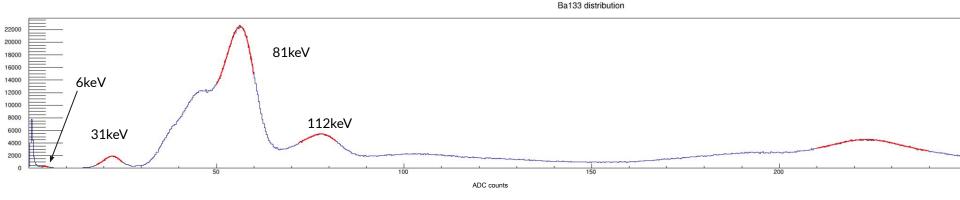
- 
$$A^*e^{-b^*x} + C^*x^d$$

- Function is then used in peak finding



#### **Calibration**

- Sources (Am241, Ba133, Cs137) placed on top of crystal
- Record PMT with no trigger scheme
- Fit Peak with Gaussian + Linear function
  - Extract Gaussian mean and std. dev.



#### **Resolution function**

- Get the std. dev of the Ba133 calibration peaks
- Plot as a function of mean energy
- Add simulation "resolution" in quadrature

