

Last time

DIS (cont'd)

$$\frac{d\sigma}{d^3k} = \frac{d_{EM}^2}{Q^4 EE'} l_{\mu\nu} W^{\mu\nu}$$

We found the leptonic tensor

$$l_{\mu\nu} = 2 (k_{\mu} k'_{\nu} + k_{\nu} k'_{\mu} - g_{\mu\nu} k_0 k'_0)$$

and hadronic tensor was written as

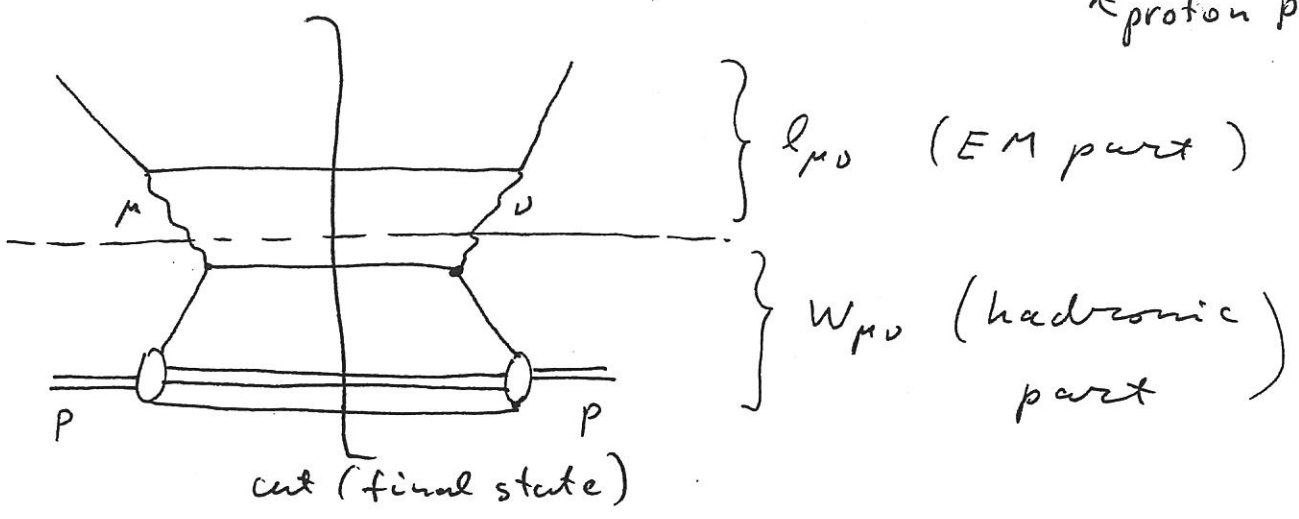
$$W_{\mu\nu} = \frac{1}{4\pi m_p} \int d^4x e^{iq \cdot x} \langle p | j_{\mu}(x) j_{\nu}(0) | p \rangle$$

with  $j_{\mu}(x)$  the quark electro-magnetic current.



$$W_{\mu\nu} = \frac{1}{4\pi m_p} \int d^4x e^{iq \cdot x} \langle p | j_\mu(x) j_\nu(0) | p \rangle$$

(averaged over  $\sigma$ )  
 $\uparrow$  proton polarization



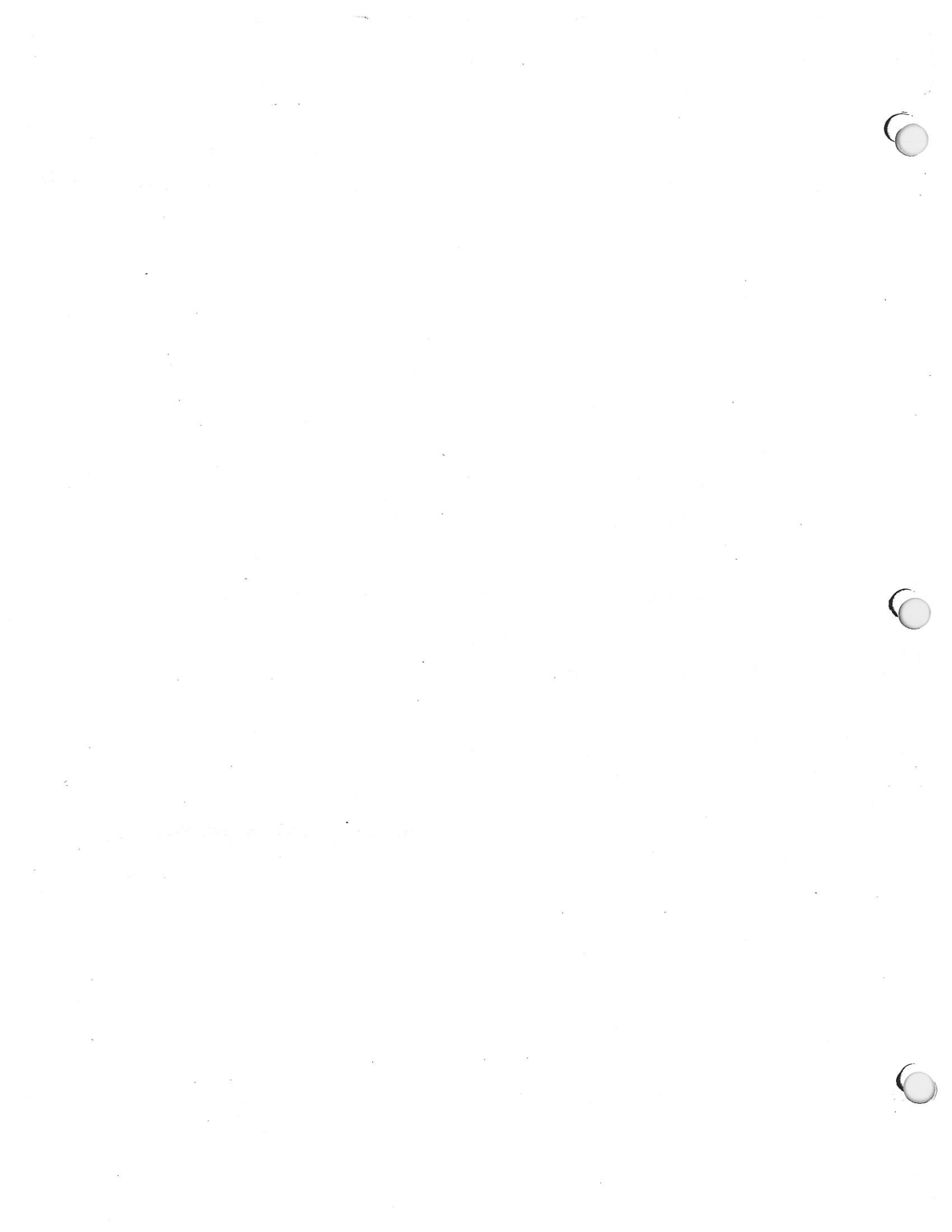
$$W_{\mu\nu}(p, q) = A(x, Q^2) p_\mu p_\nu + B(x, Q^2) g_\mu g_\nu + C(x, Q^2) g_{\mu\nu} + D(x, Q^2) (p_\mu q_\nu + p_\nu q_\mu) + E(x, Q^2) (p_\mu q_\nu - p_\nu q_\mu) + F(x, Q^2) \epsilon_{\mu\nu\sigma\rho} p^\sigma q^\rho$$

$F = 0$  in  $\gamma^* p, \gamma^* A$  (F comes from  $\gamma_5$ 's, appears in DIS or for polarized protons.)

(1)  $q_\mu W^{\mu\nu} = 0$  (current conservation)  
 $q_\mu \rightarrow \partial_\mu \Rightarrow \partial_\mu j^\mu(x) = 0$

$$A p_\nu (p \cdot q) + B q_\nu \cdot q^2 + C q_\nu + D (p \cdot q q_\nu + q^2 p_\nu) + E (p \cdot q q_\nu - q^2 p_\nu) = 0$$

(2)  $q_\nu W^{\mu\nu} = 0 = A p_\mu (p \cdot q) + B q^2 g_\mu + C q_\mu + D (p \cdot q g_\mu + q^2 p_\mu) + E (p_\mu q^2 - p \cdot q g_\mu) = 0$



$$(1) - (2) = 0 \Rightarrow (E = 0.)$$

(51)

$p_\mu$  and  $q_\mu$  are independent  $\Rightarrow$

$$0 = A p \cdot q + D q^2$$

$$0 = B q^2 + C + D p \cdot q$$

$$D = -A \frac{p \cdot q}{q^2}$$

$$B = -\frac{1}{q^2} C + A \left( \frac{p \cdot q}{q^2} \right)^2$$

$$W_{\mu\nu} = A \left[ p_\mu p_\nu - \frac{p \cdot q}{q^2} (p_\mu q_\nu + p_\nu q_\mu) + \left( \frac{p \cdot q}{q^2} \right)^2 q_\mu q_\nu \right] + C \left[ g_{\mu\nu} - \frac{q_\mu q_\nu}{q^2} \right]$$

Usually one writes

$$W_{\mu\nu} = -W_1(x, Q^2) \left[ g_{\mu\nu} - \frac{q_\mu q_\nu}{q^2} \right] + \frac{W_2(x, Q^2)}{m_p^2} \cdot$$

$$\cdot \left[ p_\mu p_\nu - \frac{p \cdot q}{q^2} (p_\mu q_\nu + p_\nu q_\mu) + \left( \frac{p \cdot q}{q^2} \right)^2 q_\mu q_\nu \right]$$

$W_1$  &  $W_2$  are structure functions (Def.)

Using  $q_\mu \ell^{\mu\nu} = q_\nu \ell^{\mu\nu} = 0$  yields

$$\ell_{\mu\nu} W^{\mu\nu} = -W_1 (-4 k \cdot k') + \frac{2W_2}{m_p^2} \left[ 2 p_0 k_0 p_0 k'_0 - m^2 k \cdot k' \right]$$

$$2 \varepsilon \varepsilon' \sin^2 \frac{\theta}{2}$$

$$2m^2 \varepsilon \varepsilon' - 2m^2 \varepsilon \varepsilon' \sin^2 \frac{\theta}{2} =$$

$$= 2m^2 \varepsilon \varepsilon' \cos^2 \frac{\theta}{2}$$

$$g_{\mu\nu} l^{\mu\nu} = (k-k')_{\mu} 2(k^{\mu} k'^{\nu} + k^{\nu} k'^{\mu} - g^{\mu\nu} k \cdot k') = \textcircled{52}$$

$$= 2 \left( k^2 k'^{\nu} + \cancel{k^{\nu} k^{\mu} k'^{\mu}} - \cancel{k^{\nu} k \cdot k'} - \cancel{k' \cdot k k'^{\nu}} - k^{\nu} k'^{\nu} \right) + \cancel{k^{\mu} k^{\nu} k'^{\mu}} = 2(k^2 k'^{\nu} - k'^2 k^{\nu}) \approx 0 \text{ as } k^2 \approx k'^2 \approx 0$$

(neglect electron's mass),  $g_{\nu\lambda} l^{\mu\nu} = 0$  (similar)

$$\Rightarrow l_{\mu\nu} W^{\mu\nu} = l^{\mu\nu} \left[ -W_1 \left( g_{\mu\nu} - \frac{g_{\mu}^{\alpha} g_{\nu}^{\beta}}{g^2} \right) + \frac{W_2}{m_p^2} \right]$$

$$\left( p_{\mu} p_{\nu} - \frac{p \cdot g}{g^2} \left( p_{\mu} g_{\nu}^{\alpha} + p_{\nu} g_{\mu}^{\alpha} \right) + \left( \frac{p \cdot g}{g^2} \right)^2 g_{\mu}^{\alpha} g_{\nu}^{\beta} \right)$$

$$= -l^{\mu}_{\mu} W_1 + \frac{W_2}{m_p^2} p_{\mu} p_{\nu} l^{\mu\nu} = \left[ \text{as } l^{\mu\nu} = 2(k^{\mu} k'^{\nu} + k^{\nu} k'^{\mu} - g^{\mu\nu} k \cdot k') \right]$$

$$= -2(2k \cdot k' - 4k \cdot k') W_1 + \frac{W_2}{m_p^2} 2(2p \cdot k p \cdot k' - p^2 k \cdot k')$$

$$= 4k \cdot k' W_1 + 2 \frac{W_2}{m_p^2} (2p \cdot k p \cdot k' - m_p^2 k \cdot k')$$

remember:  $k^{\mu} = (\epsilon, 0, 0, \epsilon)$ ,  $k'^{\mu} = (\epsilon', \epsilon' \sin \theta, 0, \epsilon' \cos \theta)$   
 $p^{\mu} = (m_p, \vec{0})$

$$\Rightarrow k \cdot k' = 2\epsilon\epsilon' \sin^2(\theta/2); \quad p \cdot k = m_p \epsilon, \quad p \cdot k' = \epsilon' m_p$$

$$L_{\mu\nu} W^{\mu\nu} = 4\epsilon\epsilon' \left[ 2W_1 \sin^2 \frac{\theta}{2} + W_2 \cos^2 \frac{\theta}{2} \right]$$

$$\frac{d\sigma}{d^3k'} = \frac{4d_{EM}^2}{Q^4} \left[ 2W_1 \sin^2 \frac{\theta}{2} + W_2 \cos^2 \frac{\theta}{2} \right]$$

By varying the angle  $\theta$  can separate  $W_1$  &  $W_2$  contributions in experiments.  $\Rightarrow$  Rosenbluth separation

Usually one defines  $F_1(x, Q^2) = m_p W_1(x, Q^2)$ ,  $F_2(x, Q^2) = \nu W_2(x, Q^2)$

The Parton Model.

Sterman ch 14, Peskin 17.5  
Y.K. & Levin, ch. 2.2

Go to Infinite Momentum Frame:

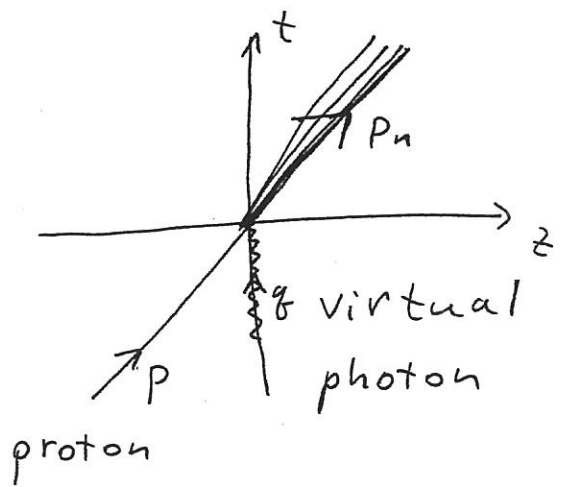
$$p^\mu \approx \left( p + \frac{m^2}{2p}, 0, 0, p \right),$$

$\vec{q} = (q^1, q^2) \sim 2d$  vector in transverse plane

$$q^\mu = \left( q_0, \vec{q}, 0 \right),$$

$Q^2$  and  $x$  are 2 invariants  
 $\leftarrow$  large,  $Q \gg \Lambda_{QCD}$

$$p \cdot q = m\nu = q_0 \cdot p$$



$\Rightarrow q_0 = \frac{m\nu}{p} \sim$  small as  $p$  goes large,  $p \gg Q$

$$\Rightarrow \text{as } \nu = \frac{Q^2}{2m_p x} \Rightarrow q_0 = \frac{Q^2}{2xp}$$

$$\Rightarrow Q^2 = -\vec{q}^2 = \vec{q}^2$$

$\Rightarrow q_0 \ll Q$  since  $xp \gg Q$ .

$$F_1(x, Q^2) = m_p W_1(x, Q^2)$$

$$F_2(x, Q^2) = v W_2(x, Q^2) = \frac{Q^2}{2m_p x} W_2(x, Q^2)$$

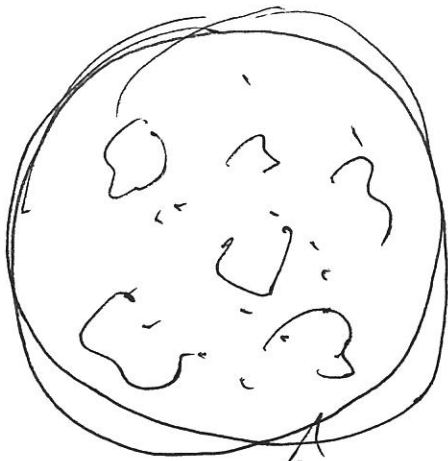
$$\frac{d\sigma}{d^3k'} = \frac{4d_{EM}^2}{Q^4} \left[ 2 \frac{1}{m_p} F_1(x, Q^2) \sin^2 \frac{\theta}{2} + \frac{2m_p x}{Q^2} F_2(x, Q^2) \cdot \cos^2(\theta/2) \right]$$

$F_1, F_2$  are dimensionless

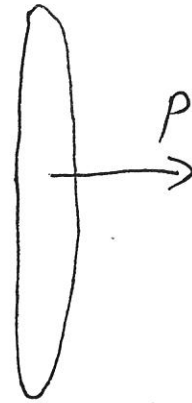


proton at rest

infinite momentum frame (IMF)



ultra-boost  
→



quantum fluctuations

$$t_{\text{fluct.}} = \delta \tau_{\text{fluct}} = \delta \frac{1}{\Lambda_{\text{QCD}}}$$

time scale

$$\delta = \frac{P}{m_p}$$

$$\tau_{\text{fluct.}} \approx \frac{1}{\Lambda_{\text{QCD}}} \leftarrow \text{only scale in the problem}$$

$$\Rightarrow t_{\text{fluct}} = \frac{P}{m_p} \frac{1}{\Lambda_{\text{QCD}}}$$

$$\frac{P}{m_p} \approx 7000 \text{ at LHC} \quad (\sqrt{s} = 14 \text{ TeV})$$

⇒ fluctuations live much longer in IMF ⇒  
⇒ can probe them

⇒ IMF "freezes" time for quantum fluctuations, allowing us to measure them.

