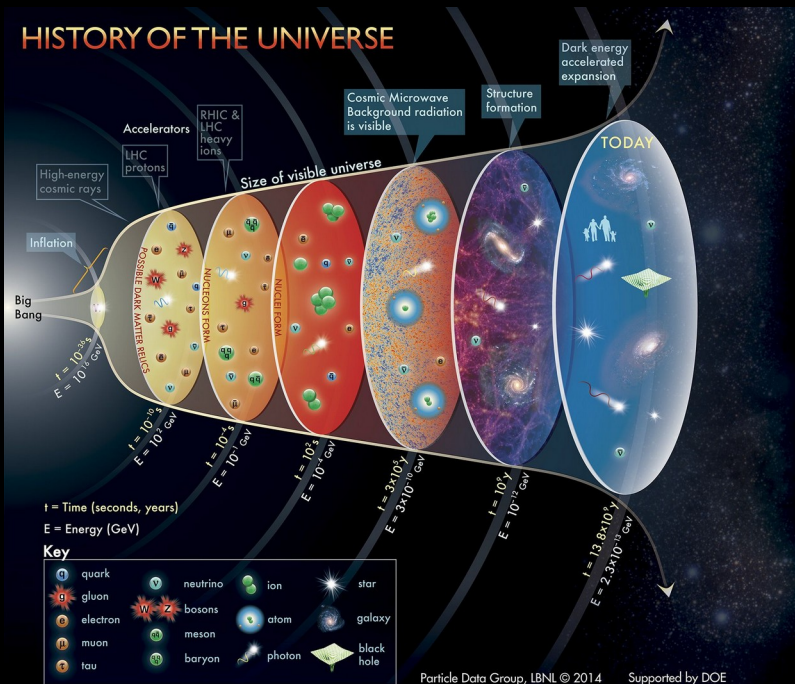




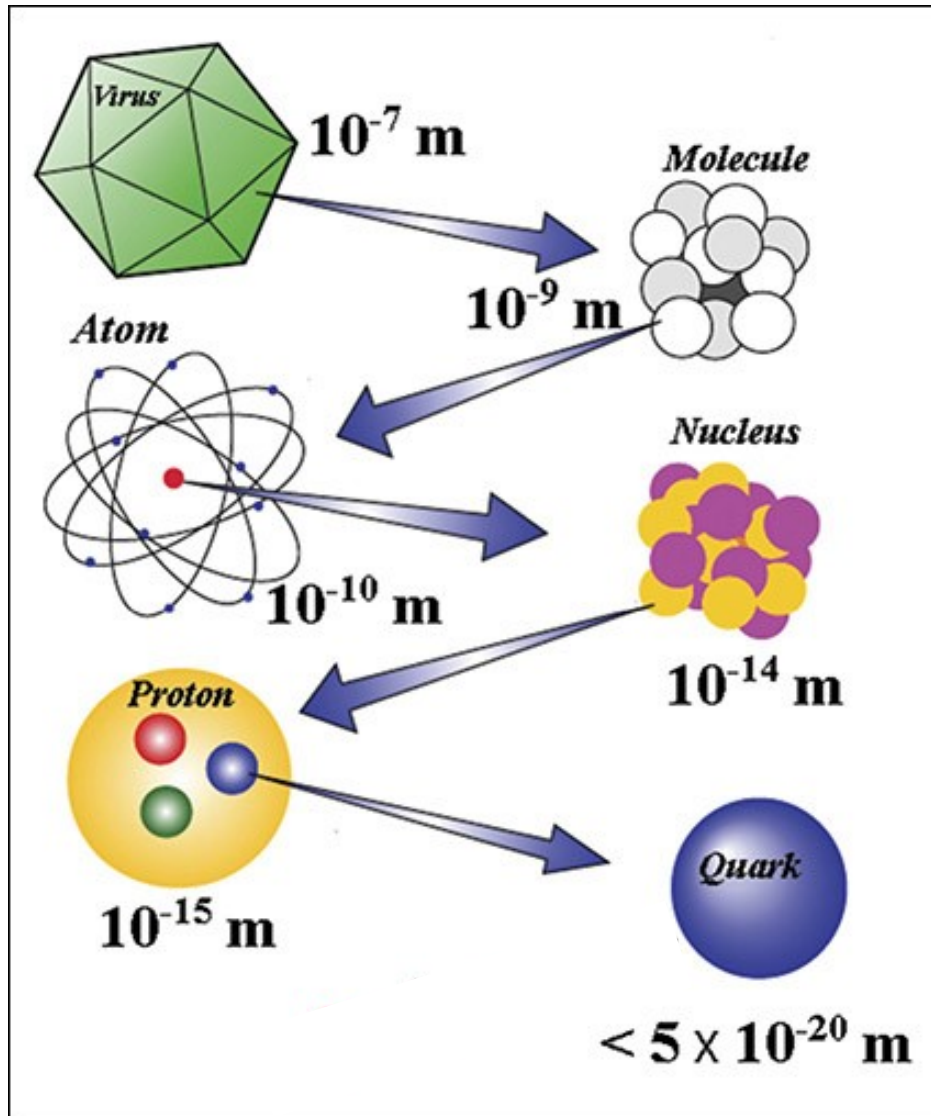
# The early universe

## as a particle physics laboratory

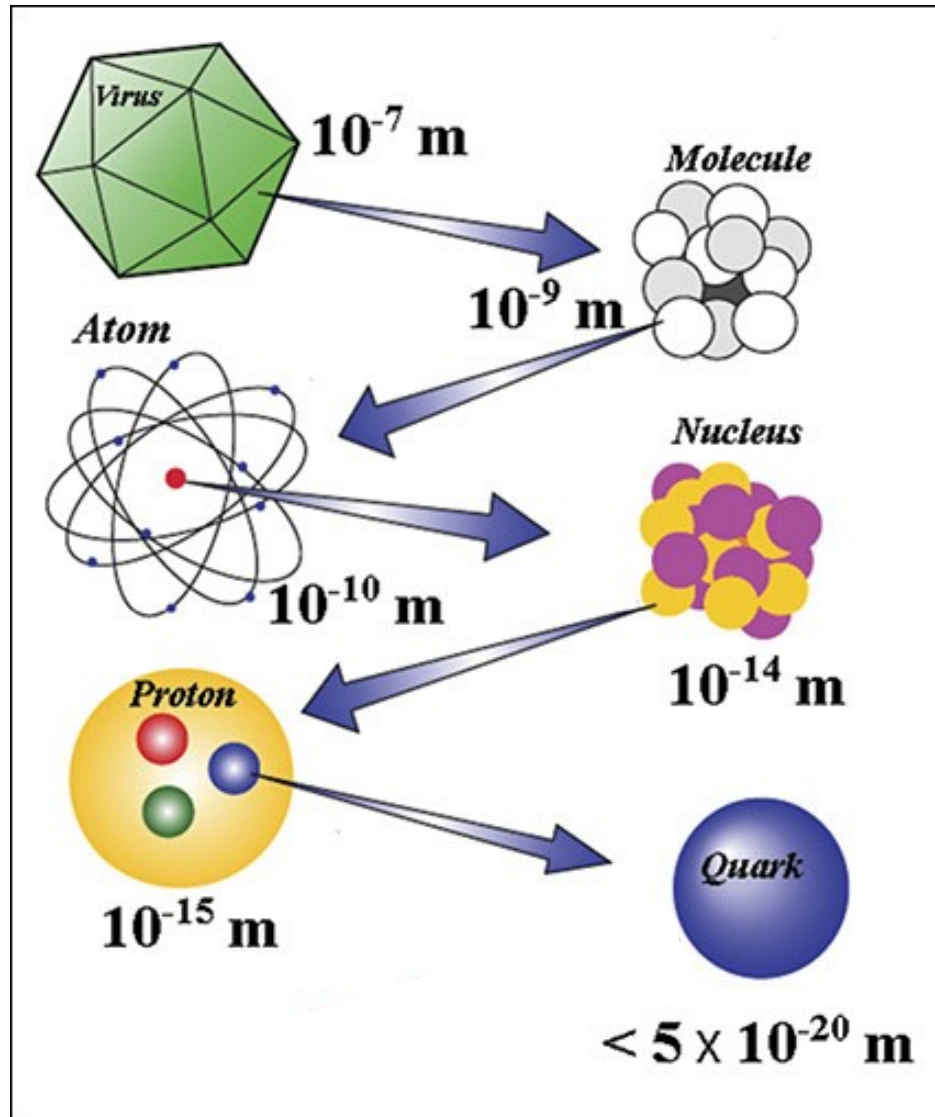


**Valerie Domcke**  
CERN TH

# From the smallest distances...



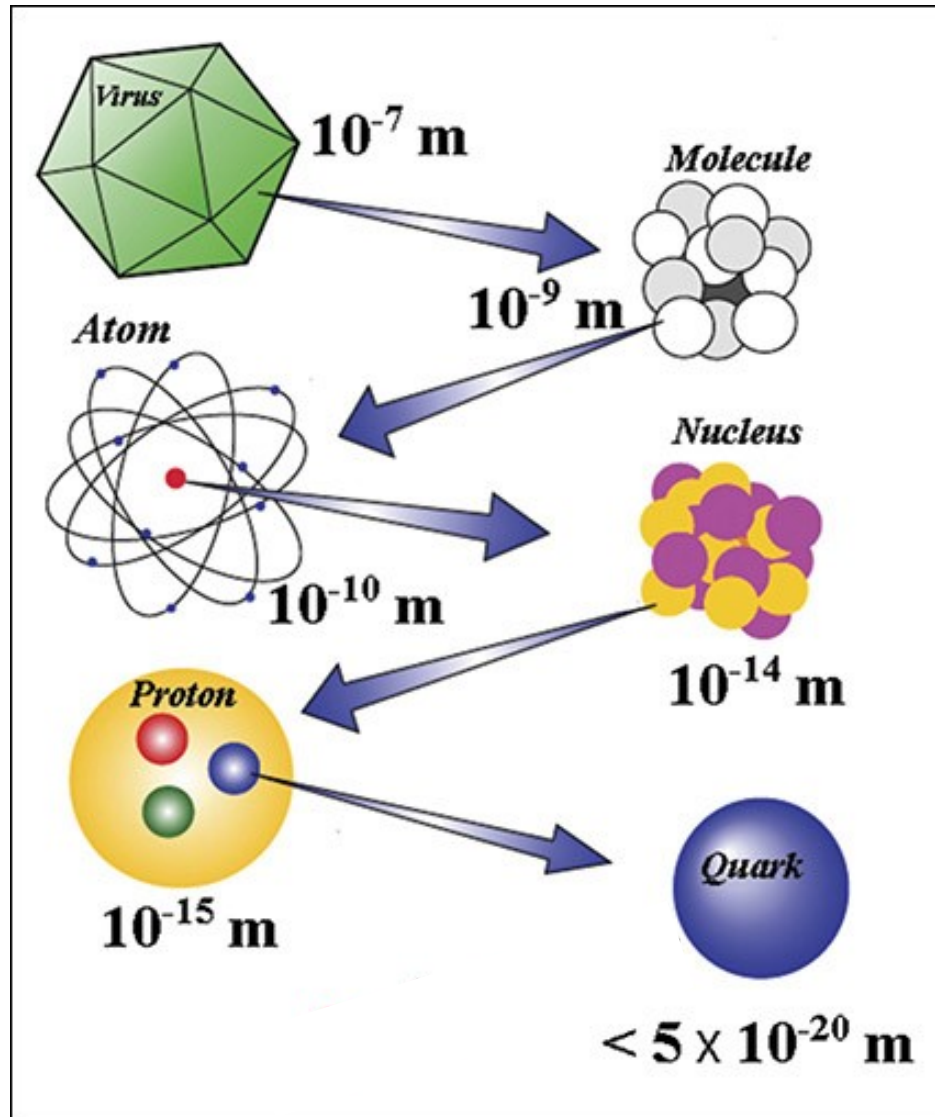
# From the smallest distances...



A periodic table of elements, showing the arrangement of chemical elements. The table includes element symbols, atomic numbers, and names. The elements are arranged in rows and columns, with the first row starting with Hydrogen (H) and Helium (He).



# From the smallest distances...

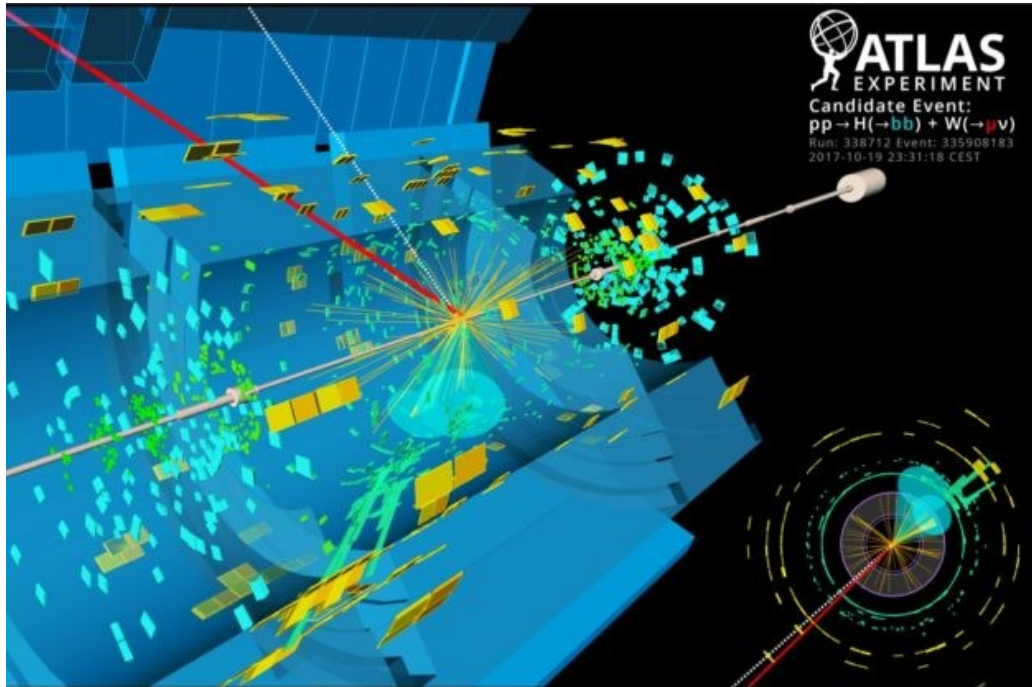


A standard periodic table of elements, showing the arrangement of chemical elements based on their atomic number and chemical properties. The table includes element symbols, names, and atomic numbers.

## Standard-Modell der Elementarteilchen

Drei Generationen der Materie (Fermionen)			Wechselwirkungen (Bosonen)	
I	II	III		
Masse Ladung Spin	$\approx 2.2 \text{ MeV}/c^2$ +2/3 1/2 <b>u</b> Up	$\approx 1.28 \text{ GeV}/c^2$ +2/3 1/2 <b>c</b> Charm	$\approx 173.1 \text{ GeV}/c^2$ +2/3 1/2 <b>t</b> Top	0 0 1 <b>g</b> Gluon
	$\approx 4.7 \text{ MeV}/c^2$ -1/3 1/2 <b>d</b> Down	$\approx 96 \text{ MeV}/c^2$ -1/3 1/2 <b>s</b> Strange	$\approx 4.18 \text{ GeV}/c^2$ -1/3 1/2 <b>b</b> Bottom	0 0 1 <b><math>\gamma</math></b> Photon
<b>QUARKS</b>	$\approx 0.511 \text{ MeV}/c^2$ -1 1/2 <b>e</b> Elektron	$\approx 105.66 \text{ MeV}/c^2$ -1 1/2 <b><math>\mu</math></b> Muon	$\approx 1.7768 \text{ GeV}/c^2$ -1 1/2 <b><math>\tau</math></b> Tau	0 0 1 <b>Z</b> Z-Boson
<b>LEPTONEN</b>	$< 1.0 \text{ eV}/c^2$ 0 1/2 <b><math>\nu_e</math></b> Elektron-Neutrino	$< 0.17 \text{ MeV}/c^2$ 0 1/2 <b><math>\nu_\mu</math></b> Muon-Neutrino	$< 18.2 \text{ MeV}/c^2$ 0 1/2 <b><math>\nu_\tau</math></b> Tau-Neutrino	$\approx 80.39 \text{ GeV}/c^2$ $\pm 1$ 1 <b>W</b> W-Boson
				<b>H</b> Higgs
				<b>SKALARBOSONEN</b>
				<b>EICHBOSONEN VEKTORBOSONEN</b>

# ... to the highest energies ...



- At „normal“ temperatures, quarks are confined into protons and neutrons, protons are stable
- At high-energy collisions, free quarks become „free“ for a short time period
- More massive virtual (short-lived) elementary particles can be formed

$$10^{-20} \text{ m} = \hbar c / 10 \text{ TeV} \quad \rightarrow \quad \frac{L}{10^{-20} \text{ m}} \hat{=} \frac{10 \text{ TeV}}{E}$$

# ... to the earliest times



universe today

- cold:  $-270\text{ }^{\circ}\text{C}$  (2.7 K)
- largely empty
- inhomogeneous
- matter consists of atoms, molecules, ...
- expanding

# ... to the earliest times

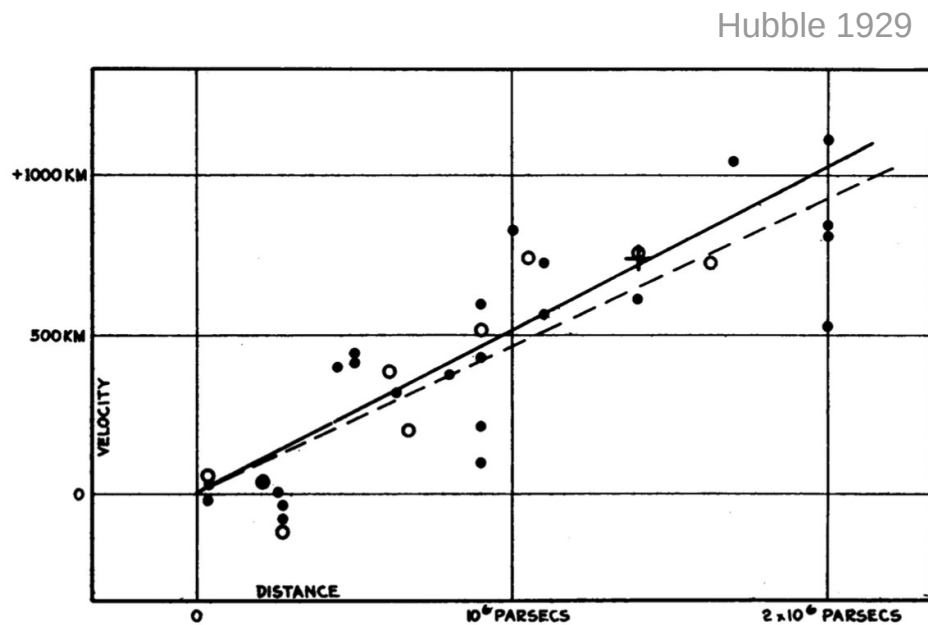


FIGURE 1

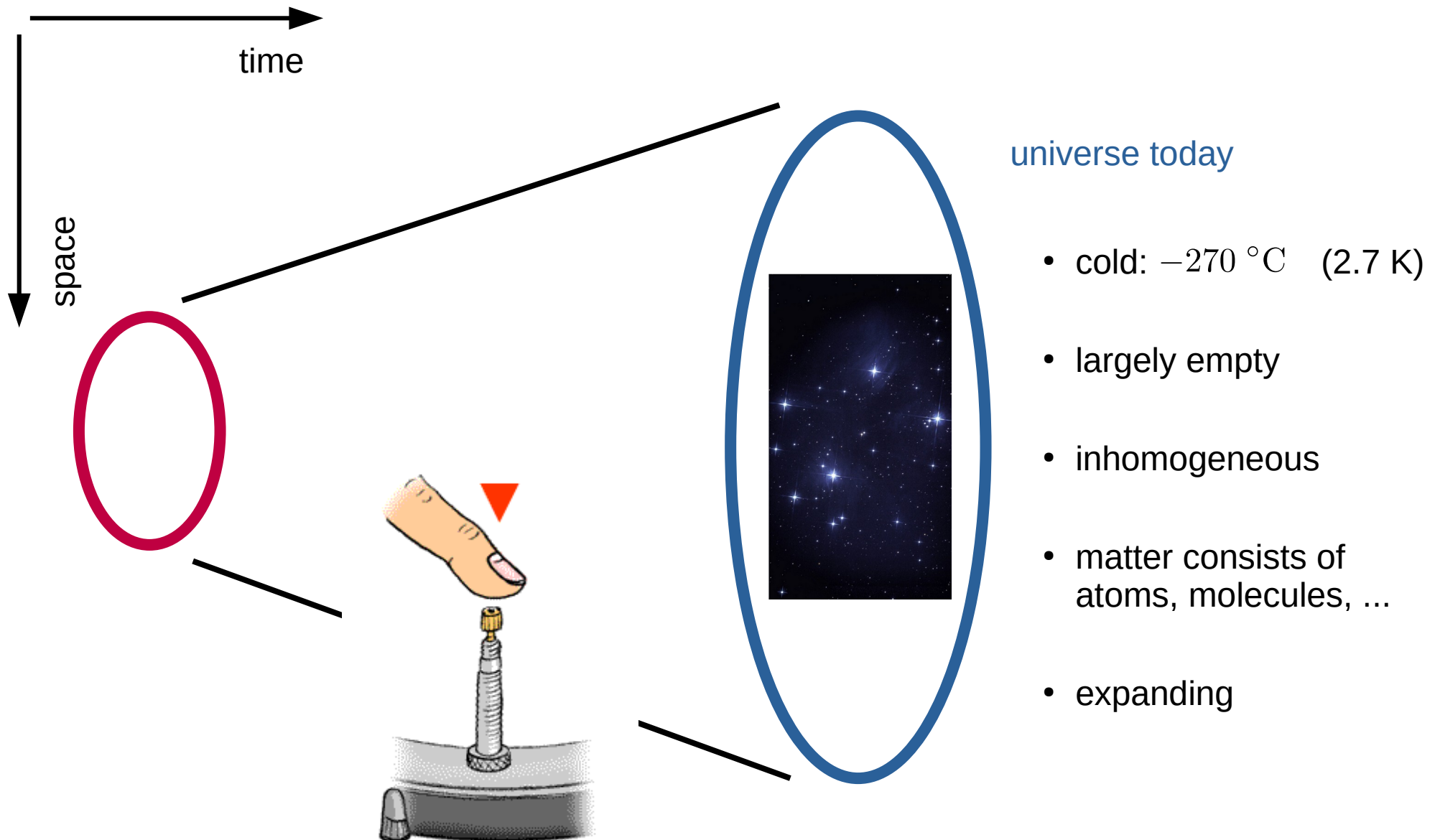
Velocity-Distance Relation among Extra-Galactic Nebulae.



universe today

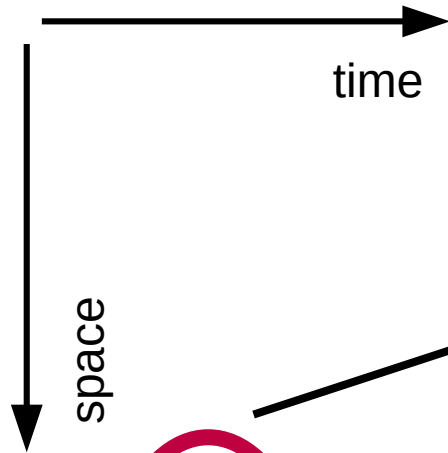
- cold:  $-270\text{ }^{\circ}\text{C}$  (2.7 K)
- largely empty
- inhomogeneous
- matter consists of atoms, molecules, ...
- expanding

# ... to the earliest times





# ... to the earliest times



13 billion years ago

- hot:  $\sim 10^{20} \text{ }^\circ\text{C}$
- homogeneous plasma

elementary building blocks of matter exist as free particles

**Standard-Modell der Elementarteilchen**

	Drei Generationen der Materie (Fermionen)			Wechselwirkungen (Bosonen)	
	I	II	III		
Masse	$\sim 2.2 \text{ MeV}/c^2$	$\sim 1.28 \text{ GeV}/c^2$	$\sim 173.1 \text{ GeV}/c^2$	0	$\sim 124.97 \text{ GeV}/c^2$
Ladung	$\frac{2}{3}$	$\frac{2}{3}$	$\frac{2}{3}$	0	0
Spin	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	0	0
<b>QUARKS</b>	<b>u</b> Up	<b>c</b> Charm	<b>t</b> Top	<b>g</b> Gluon	<b>H</b> Higgs
	$\frac{1}{3}$	$\frac{1}{3}$	$\frac{1}{3}$	0	0
	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$	0	0
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	1
	<b>d</b> Down	<b>s</b> Strange	<b>b</b> Bottom	<b><math>\gamma</math></b> Photon	
<b>LEPTONEN</b>	$\frac{1}{3}$	$\frac{1}{3}$	$\frac{1}{3}$	0	0
	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$	0	0
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	1
	<b>e</b> Elektron	<b><math>\mu</math></b> Muon	<b><math>\tau</math></b> Tau	<b>Z</b> Z-Boson	
	$\frac{1}{3}$	$\frac{1}{3}$	$\frac{1}{3}$	0	0
	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$	0	0
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	1
	<b><math>\nu_e</math></b> Elektron-Neutrino	<b><math>\nu_\mu</math></b> Muon-Neutrino	<b><math>\nu_\tau</math></b> Tau-Neutrino	<b>W</b> W-Boson	
	$\frac{1}{3}$	$\frac{1}{3}$	$\frac{1}{3}$	0	0
	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$	0	0
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	1
	<b><math>\nu_e</math></b> Elektron-Neutrino	<b><math>\nu_\mu</math></b> Muon-Neutrino	<b><math>\nu_\tau</math></b> Tau-Neutrino	<b>W</b> W-Boson	

**EICHBOSONEN VEKTORBOSONEN** (Z, W)

**SKALARBOSONEN** (H)

universe today



- cold:  $-270 \text{ }^\circ\text{C}$  (2.7 K)
- largely empty
- inhomogeneous
- matter consists of atoms, molecules, ...
- expanding

# Standard Model (SM) or Particle Physics

**Standard-Modell der Elementarteilchen**

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	I	II	III		
Masse	$\approx 2.2 \text{ MeV}/c^2$	$\approx 1.28 \text{ GeV}/c^2$	$\approx 173.1 \text{ GeV}/c^2$	0	$\approx 124.97 \text{ GeV}/c^2$
Ladung	$\frac{2}{3}$	$\frac{2}{3}$	$\frac{2}{3}$	0	0
Spin	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	0
	<b>u</b> Up	<b>c</b> Charm	<b>t</b> Top	<b>g</b> Gluon	<b>H</b> Higgs
<b>QUARKS</b>	$\approx 4.7 \text{ MeV}/c^2$	$\approx 96 \text{ MeV}/c^2$	$\approx 4.18 \text{ GeV}/c^2$	0	
	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$	0	
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	
	<b>d</b> Down	<b>s</b> Strange	<b>b</b> Bottom	<b><math>\gamma</math></b> Photon	
<b>LEPTONEN</b>	$\approx 0.511 \text{ MeV}/c^2$	$\approx 105.66 \text{ MeV}/c^2$	$\approx 1.7768 \text{ GeV}/c^2$	$\approx 91.19 \text{ GeV}/c^2$	
	-1	-1	-1	0	
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	
	<b>e</b> Elektron	<b><math>\mu</math></b> Muon	<b><math>\tau</math></b> Tau	<b>Z</b> Z-Boson	
	$< 1.0 \text{ eV}/c^2$	$< 0.17 \text{ MeV}/c^2$	$< 18.2 \text{ MeV}/c^2$	$\approx 80.39 \text{ GeV}/c^2$	
	0	0	0	$\pm 1$	
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	
	<b><math>\nu_e</math></b> Elektron-Neutrino	<b><math>\nu_\mu</math></b> Muon-Neutrino	<b><math>\nu_\tau</math></b> Tau-Neutrino	<b>W</b> W-Boson	

**SKALARBOSONEN** (Higgs, Photon)  
**EICHBOSONEN VEKTORBOSONEN** (Gluon, Z, W)

$$\begin{aligned}
 \mathcal{L} = & -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\
 & + i\bar{\psi} \not{D} \psi + h.c. \\
 & + \sum_i y_{ij} \bar{\psi}_i \phi \psi_j + h.c. \\
 & + |D_\mu \phi|^2 - V(\phi)
 \end{aligned}$$

Elementary „building blocks“ in the framework of quantum field theory

# Standard Model (SM) or Particle Physics

**Standard-Modell der Elementarteilchen**

Drei Generationen der Materie (Fermionen)			Wechselwirkungen (Bosonen)		
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Ladung	$\frac{2}{3}$	$\frac{2}{3}$	$\frac{2}{3}$	0	0
Spin	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	0
	<b>u</b> Up	<b>c</b> Charm	<b>t</b> Top	<b>g</b> Gluon	<b>H</b> Higgs
<b>QUARKS</b>	$\approx 4.7 \text{ MeV}/c^2$	$\approx 96 \text{ MeV}/c^2$	$\approx 4.18 \text{ GeV}/c^2$	0	
	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$	0	
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	
	<b>d</b> Down	<b>s</b> Strange	<b>b</b> Bottom	<b><math>\gamma</math></b> Photon	
<b>LEPTONEN</b>	$\approx 0.511 \text{ MeV}/c^2$	$\approx 105.66 \text{ MeV}/c^2$	$\approx 1.7768 \text{ GeV}/c^2$	$\approx 91.19 \text{ GeV}/c^2$	
	-1	-1	-1	0	
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	
	<b>e</b> Elektron	<b><math>\mu</math></b> Muon	<b><math>\tau</math></b> Tau	<b>Z</b> Z-Boson	
	$< 1.0 \text{ eV}/c^2$	$< 0.17 \text{ MeV}/c^2$	$< 18.2 \text{ MeV}/c^2$	$\approx 80.39 \text{ GeV}/c^2$	
	0	0	0	$\pm 1$	
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	
	<b><math>\nu_e</math></b> Elektron-Neutrino	<b><math>\nu_\mu</math></b> Muon-Neutrino	<b><math>\nu_\tau</math></b> Tau-Neutrino	<b>W</b> W-Boson	

**SKALARBOSONEN** (Higgs)  
**EICHBOSONEN VEKTORBOSONEN** (Photon, Z, W)

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + i\bar{\psi} \not{D} \psi + h.c. + \chi_i Y_{ij} \chi_j \phi + h.c. + |D_\mu \phi|^2 - V(\phi)$$

Elementary „building blocks“ in the framework of quantum field theory

High Energy Frontier:

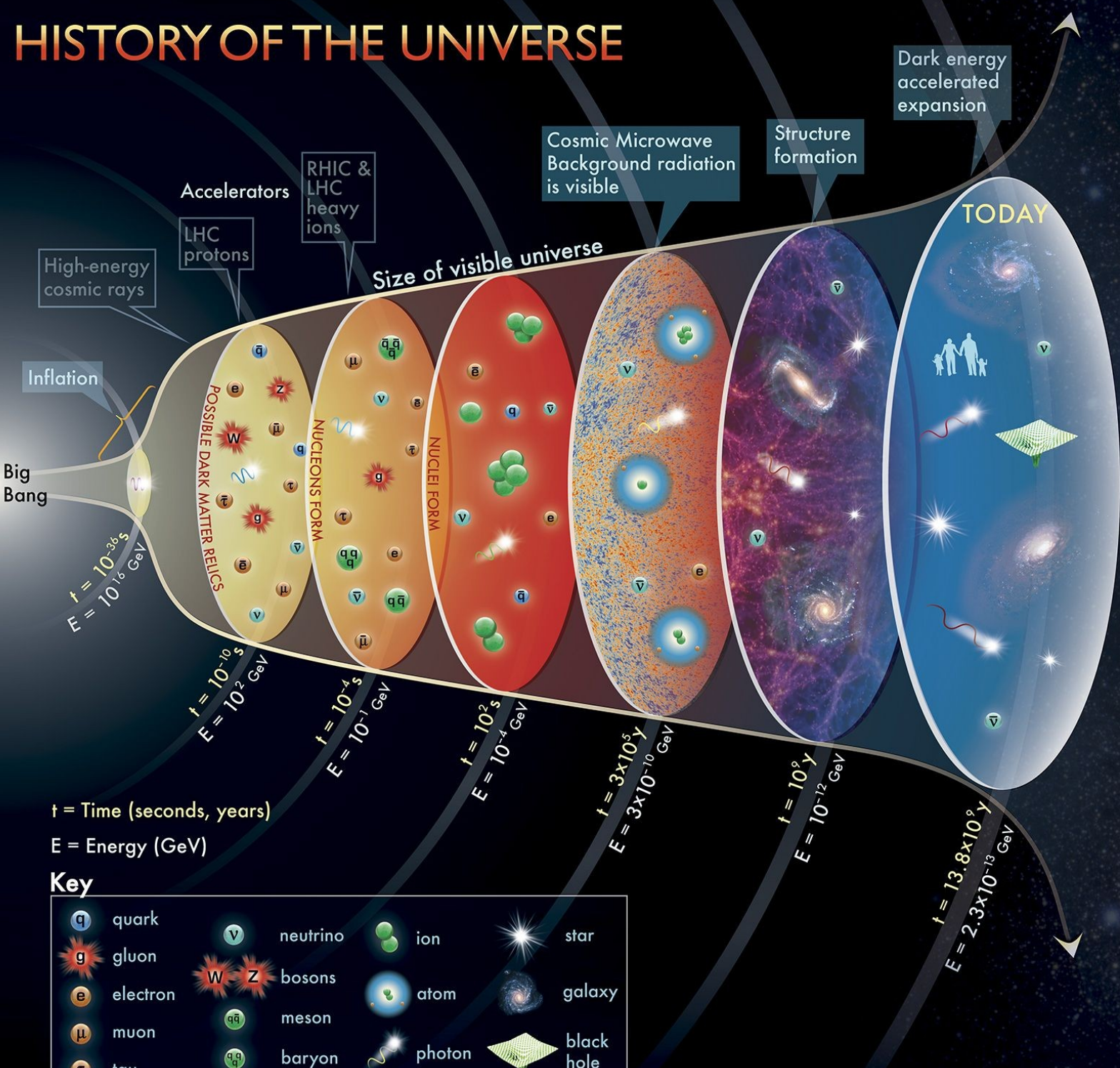
- Other elementary particles?
- Are the SM particles truly elementary?

Early Universe Frontier:

- Can the SM explain all observations to date?
- Possible relics from earlier times/higher energies?



# HISTORY OF THE UNIVERSE



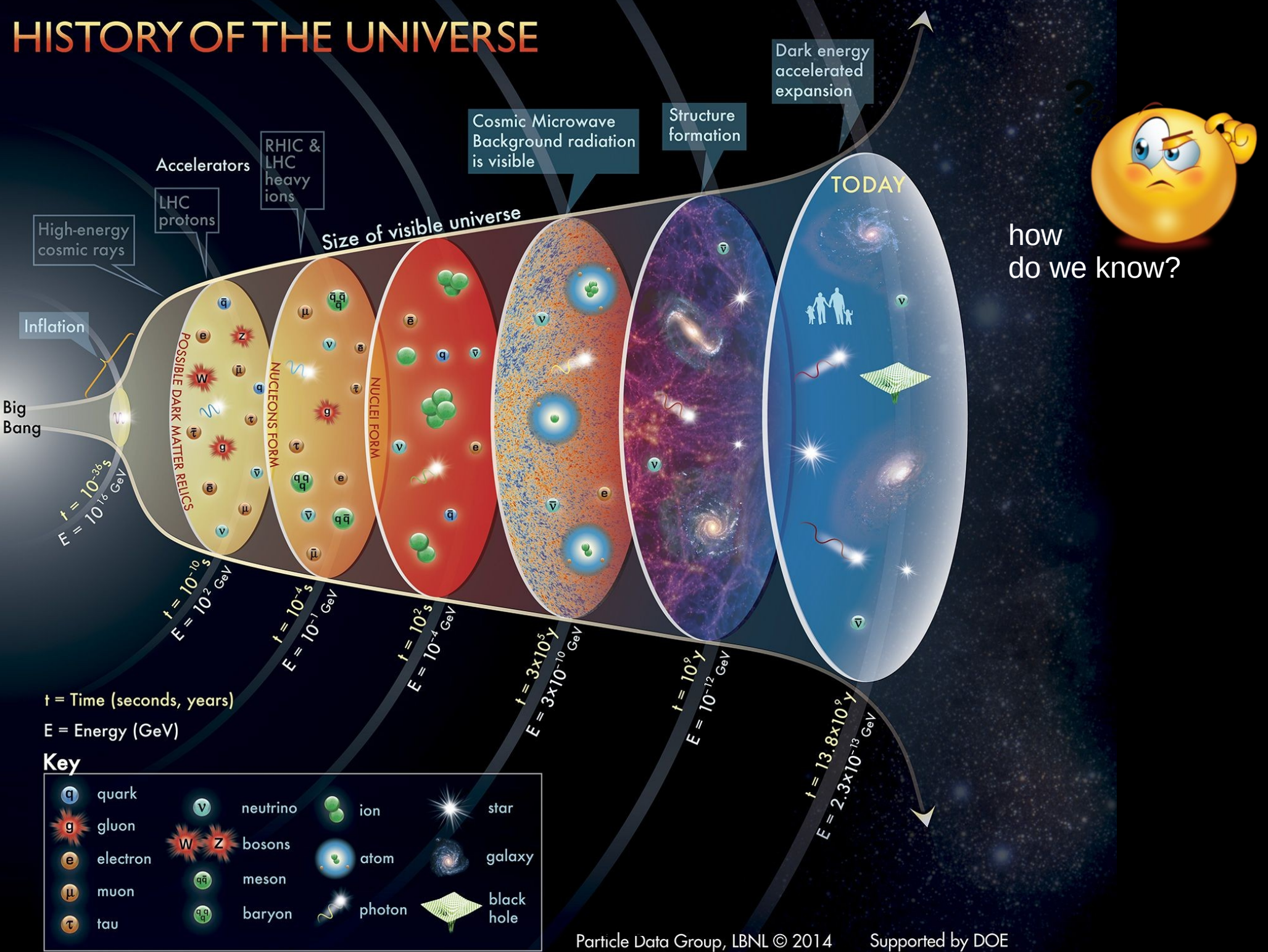
t = Time (seconds, years)  
E = Energy (GeV)

**Key**

quark	neutrino	ion	star
gluon	bosons	atom	galaxy
electron	meson	photon	black hole
muon	baryon		
tau			

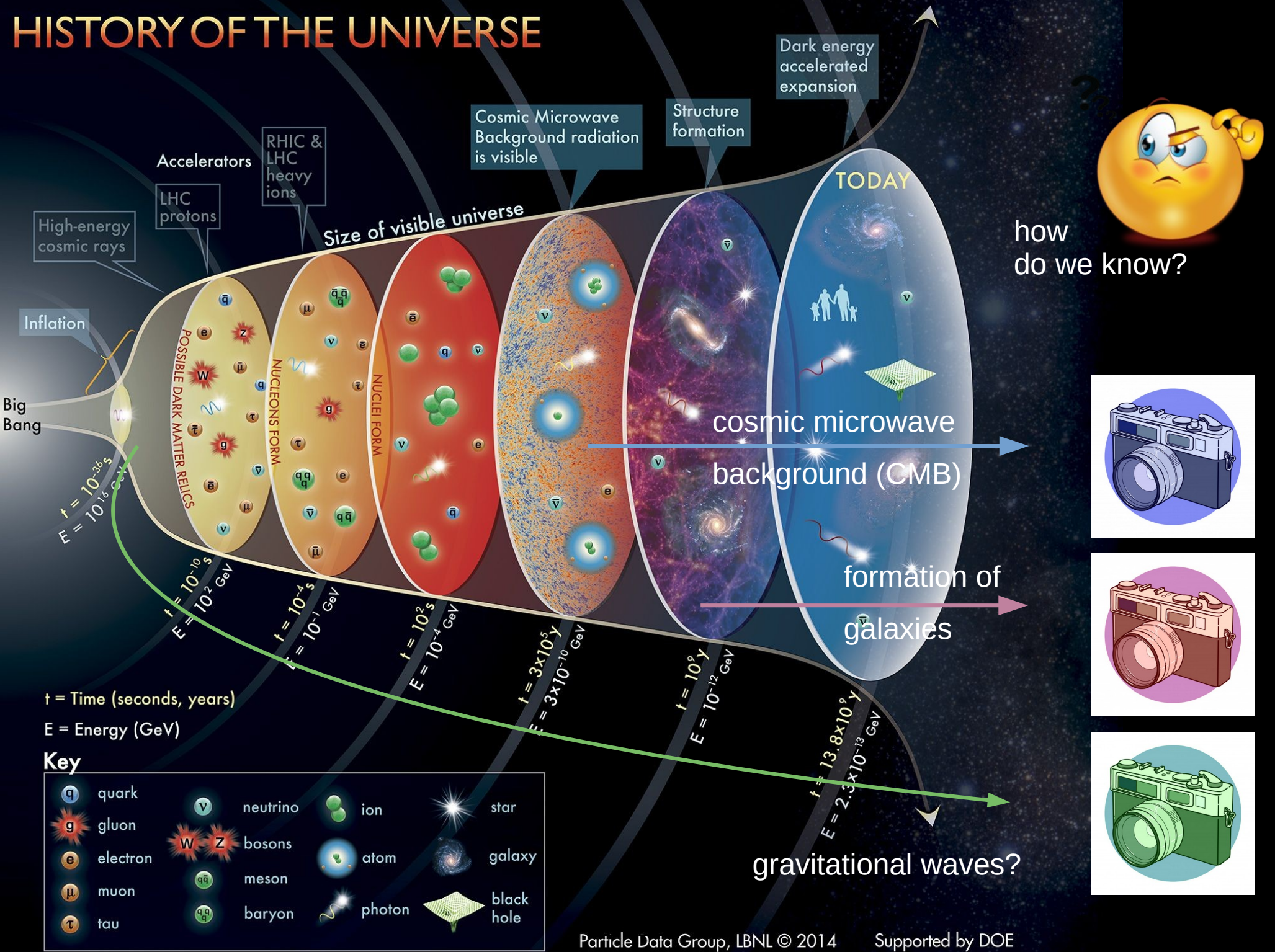


# HISTORY OF THE UNIVERSE





# HISTORY OF THE UNIVERSE



$t = 10^{-36} s$   
 $E = 10^{16} GeV$

$t = 10^{-10} s$   
 $E = 10^2 GeV$

$t = 10^{-4} s$   
 $E = 10^{-1} GeV$

$t = 10^2 s$   
 $E = 10^{-4} GeV$

$t = 3 \times 10^5 y$   
 $E = 3 \times 10^{-10} GeV$

$t = 10^9 y$   
 $E = 10^{-12} GeV$

$t = 13.8 \times 10^9 y$   
 $E = 2.3 \times 10^{-13} GeV$

$t$  = Time (seconds, years)  
 $E$  = Energy (GeV)

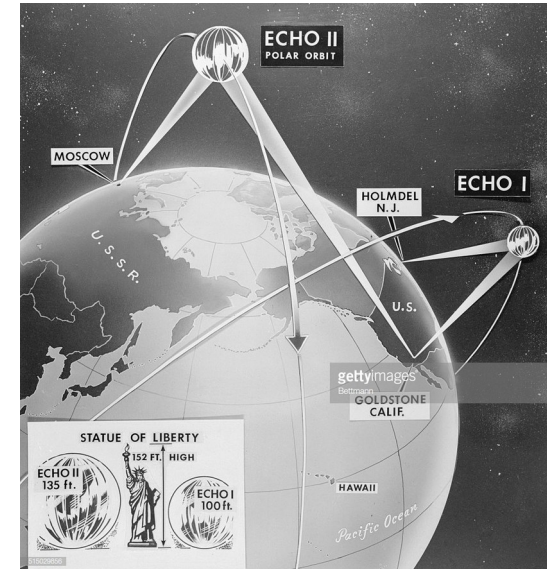
**Key**

quark	neutrino	ion	star
gluon	bosons	atom	galaxy
electron	meson	photon	black hole
muon	baryon		
tau			

# US East Coast, 1960s ...



Arno Penzias, Robert Wilson 1964



Project Echo, 1960

- Bell Lab's Horn Antenna: a 6m radio telescope promising unprecedented sensitivity
- But a background noise is disrupting the measurements ...



# The search for the culprit begins...

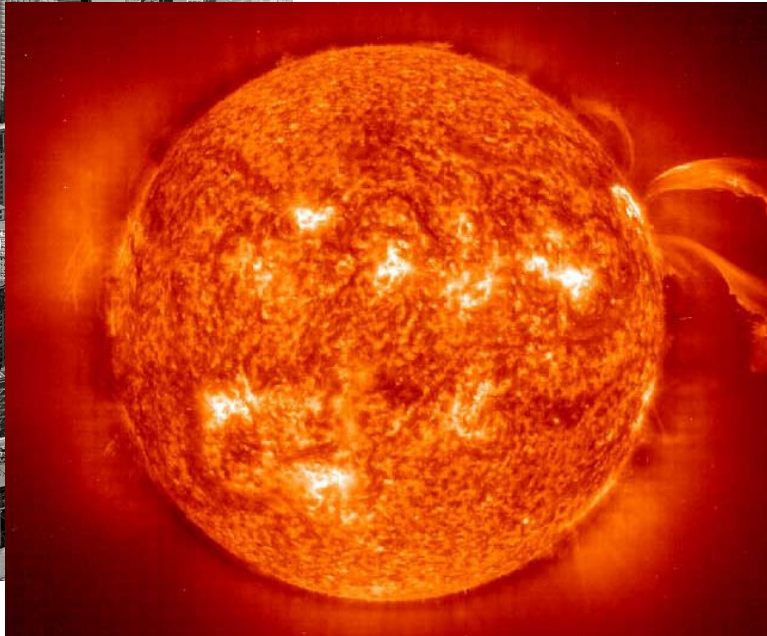


New York?

# The search for the culprit begins...

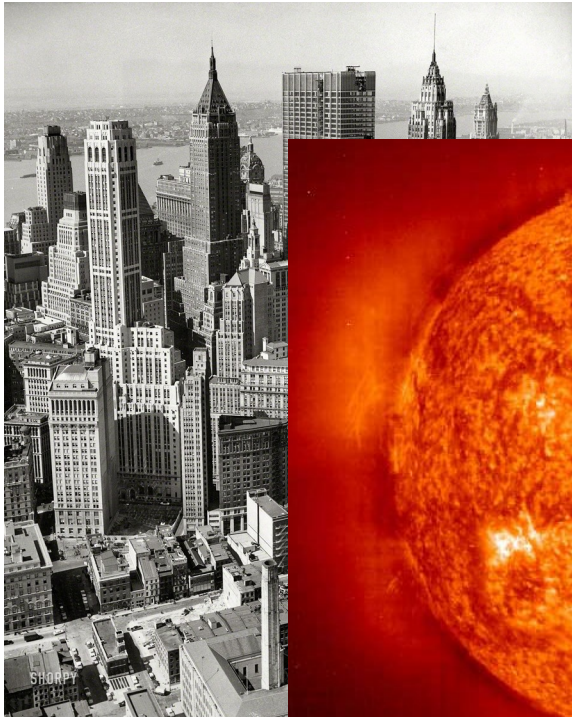


New York?



The sun?

# The search for the culprit begins...



New York?



The sun?



The galaxy?



# The search for the culprit begins...



New York?



The sun?



The galaxy?



pigeons?

# A bold theory

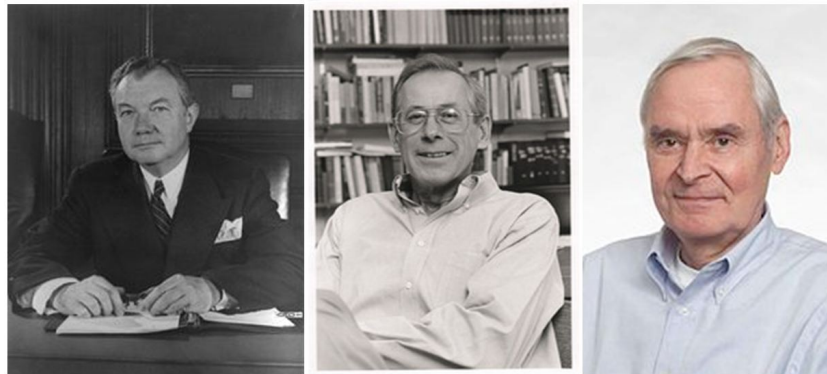
At the same time in Princeton, 60 km away

- Theoretical physicists are discussing the very nature of the universe:

”Steady State” or ”Big Bang” ?

- Robert Dicke, Jim Peebles and David Wilkinson’

Big Bang Theory → cosmic background radiation as relic of the primordial universe



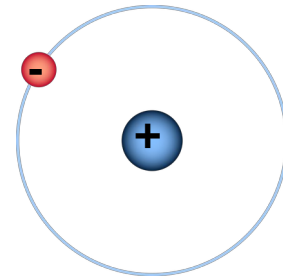
# the cosmic microwave background

binding energy of hydrogen atom:  $T \sim 3000^\circ\text{C}$

$T > 3000^\circ\text{C}$

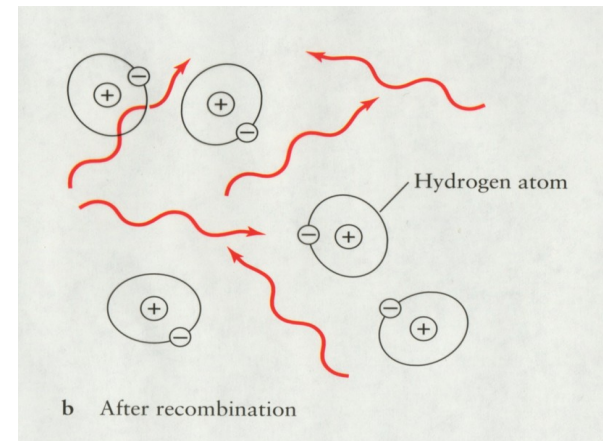
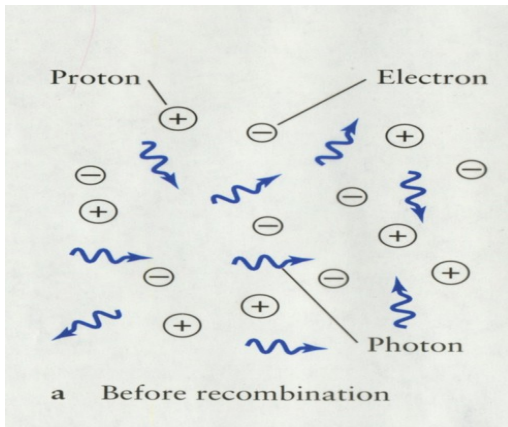


$T < 3000^\circ\text{C}$



- many free charged particles (electrons & protons)
- photons scatter multiple times, universe not transparent

- electrons and protons from electrically neutral hydrogen atoms
- universe becomes transparent



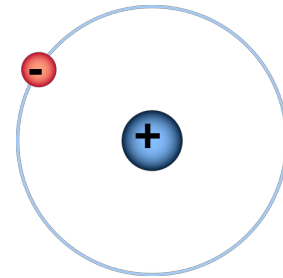
# the cosmic microwave background

binding energy of hydrogen atom:  $T \sim 3000^\circ\text{C}$

$T > 3000^\circ\text{C}$

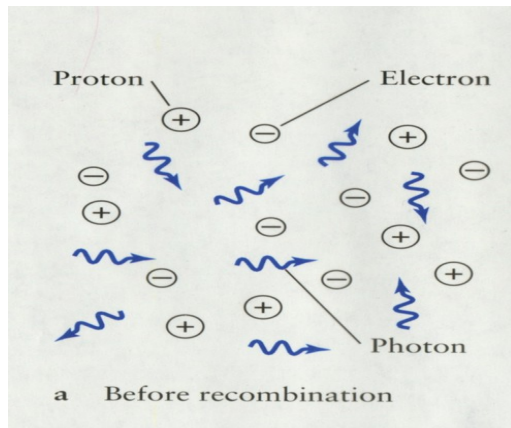


$T < 3000^\circ\text{C}$

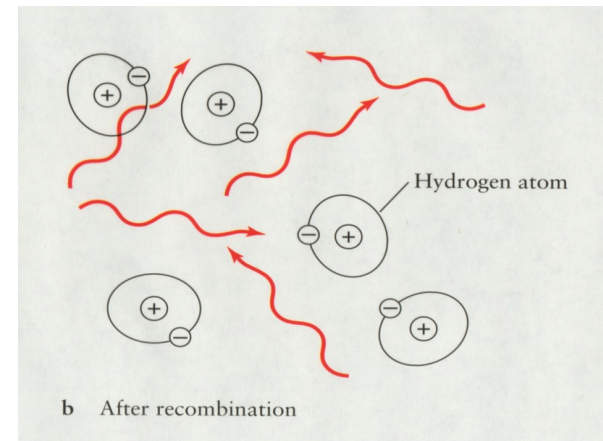


- many free charged particles (electrons & protons)
- photons scatter multiple times, universe not transparent

- electrons and protons from electrically neutral hydrogen atoms
- universe becomes transparent



- thermal radiation with  $T \sim 3000\text{ C}$  as cosmic background radiation
- cools in expanding universe to  $T \ll 3000\text{ C}$



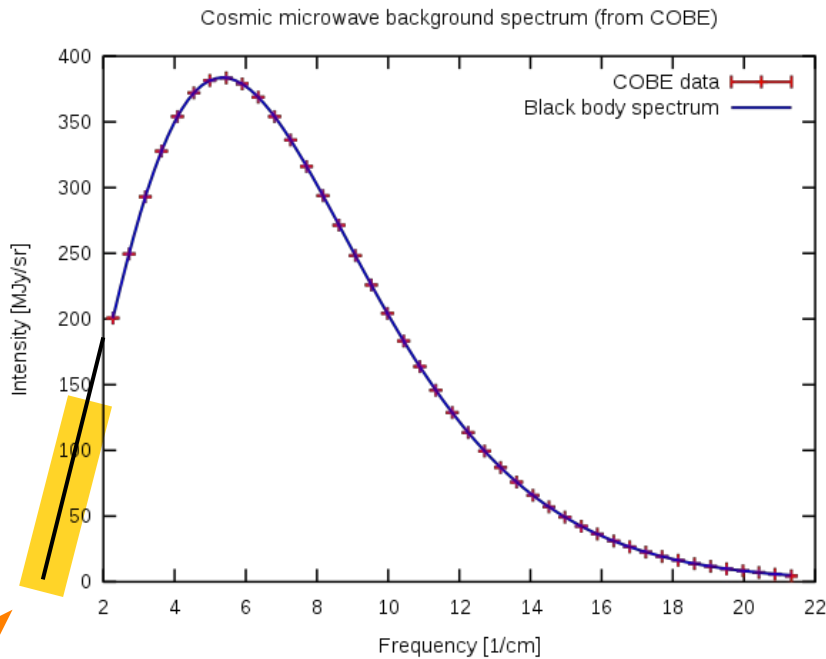


# CMB black body radiation



COBE satellite,  
1989-93

- cosmic microwave background well measured today
- black body radiation with  $T = 2.7 \text{ K}$  ( $-270 \text{ C}$ ) (microwaves)



confirms key prediction  
of `big bang' theory



2019 nobel prize Peebles for his  
contributions to theoretical cosmology

Penzias, Wilson (nobel prize 1978)

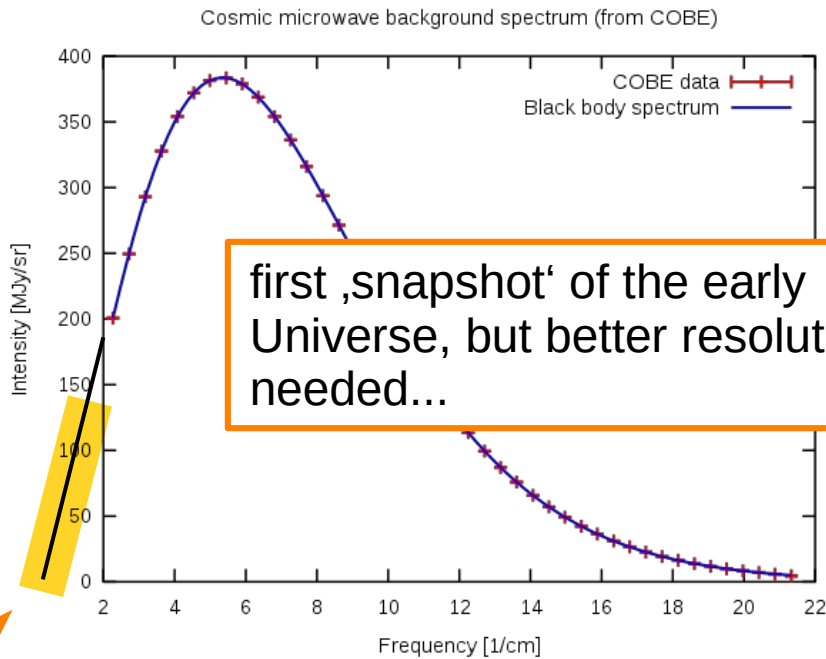


# CMB black body radiation



COBE satellite,  
1989-93

- cosmic microwave background well measured today
- black body radiation with  $T = 2.7 \text{ K}$  ( $-270 \text{ C}$ ) (microwaves)



first 'snapshot' of the early Universe, but better resolution needed...

confirms key prediction of 'big bang' theory



2019 nobel prize Peebles for his contributions to theoretical cosmology

Penzias, Wilson (nobel prize 1978)

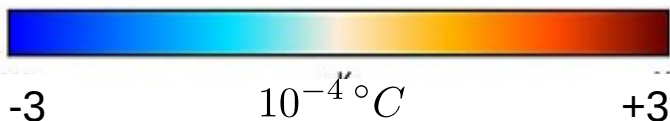
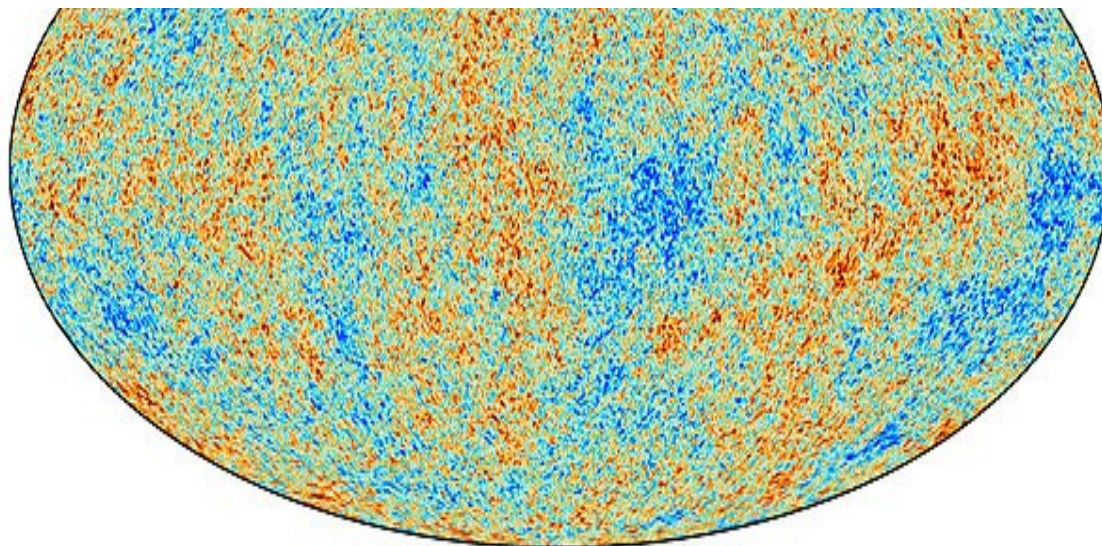
# anisotropies in the CMB

completely homogeneous plasma → homogeneous universe after cooling

- small perturbations needed as seeds for galaxies to form through gravitational collapse
- anisotropies in the CMB, deviation from black body radiation  $1:10^4$

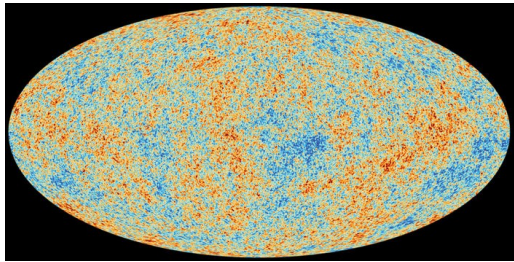


PLANCK satellite,  
2009 - 2013

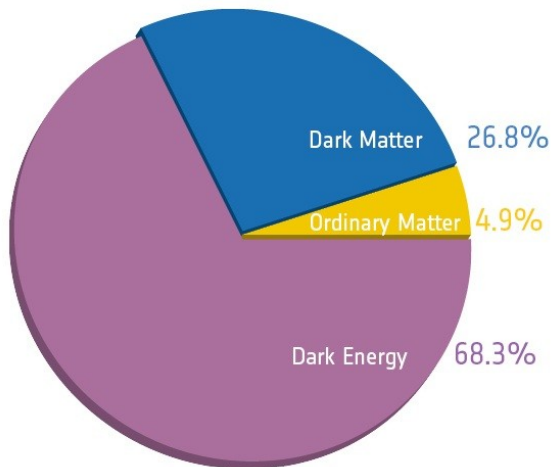
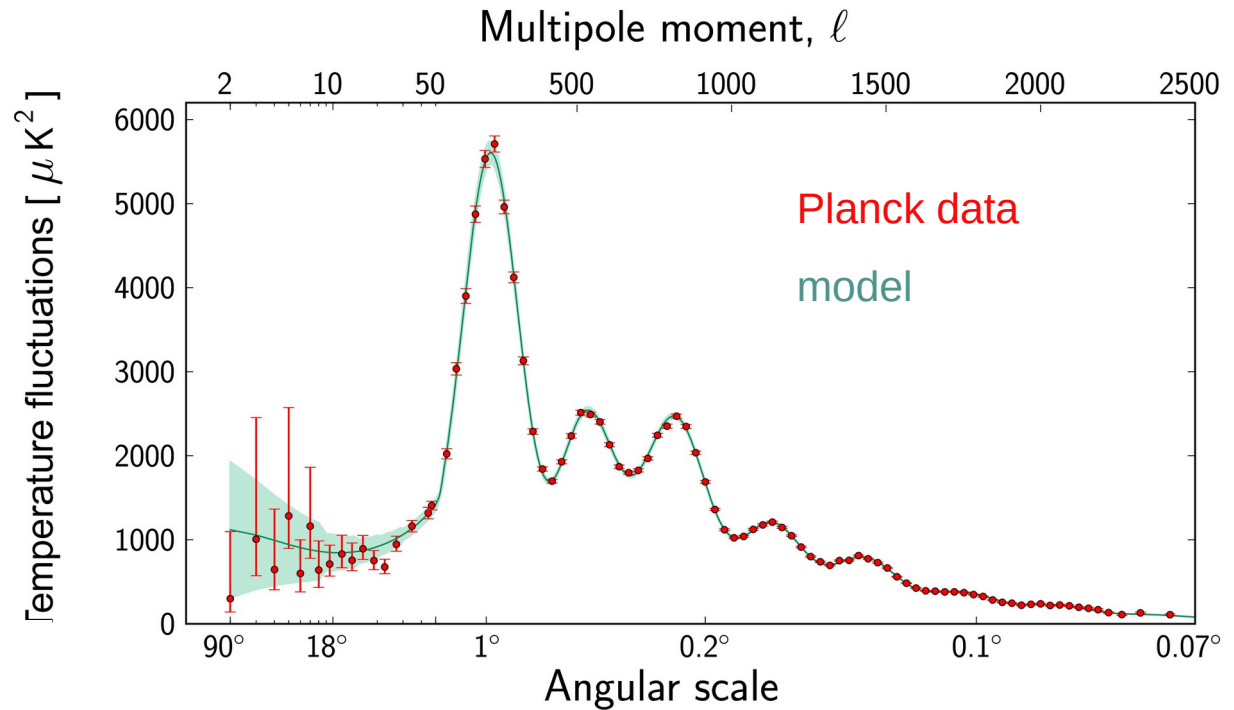


PLANCK 2018 data release

# anisotropies in the CMB



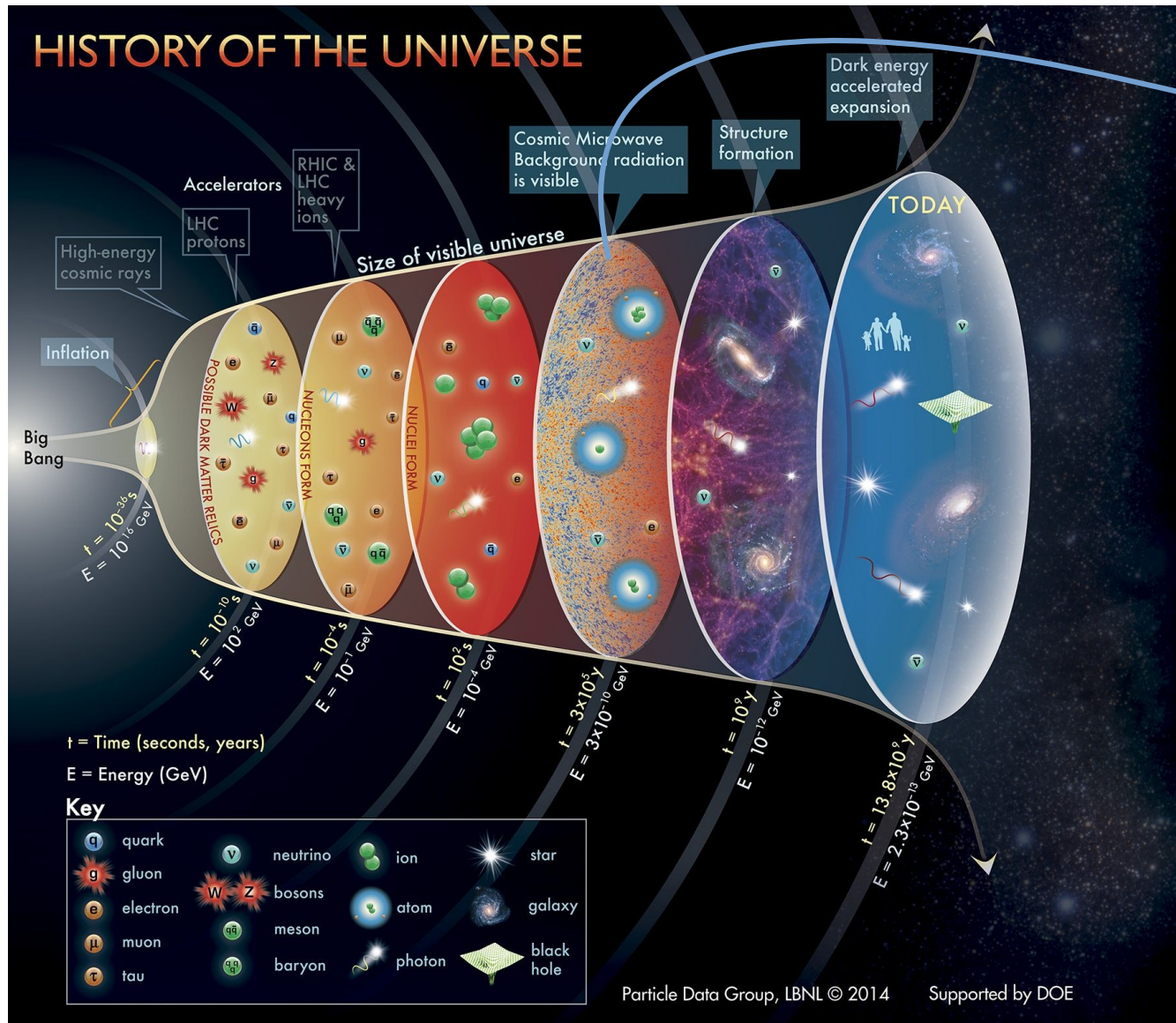
statistical analysis



standard model  
of cosmology ( $\Lambda$ CDM)

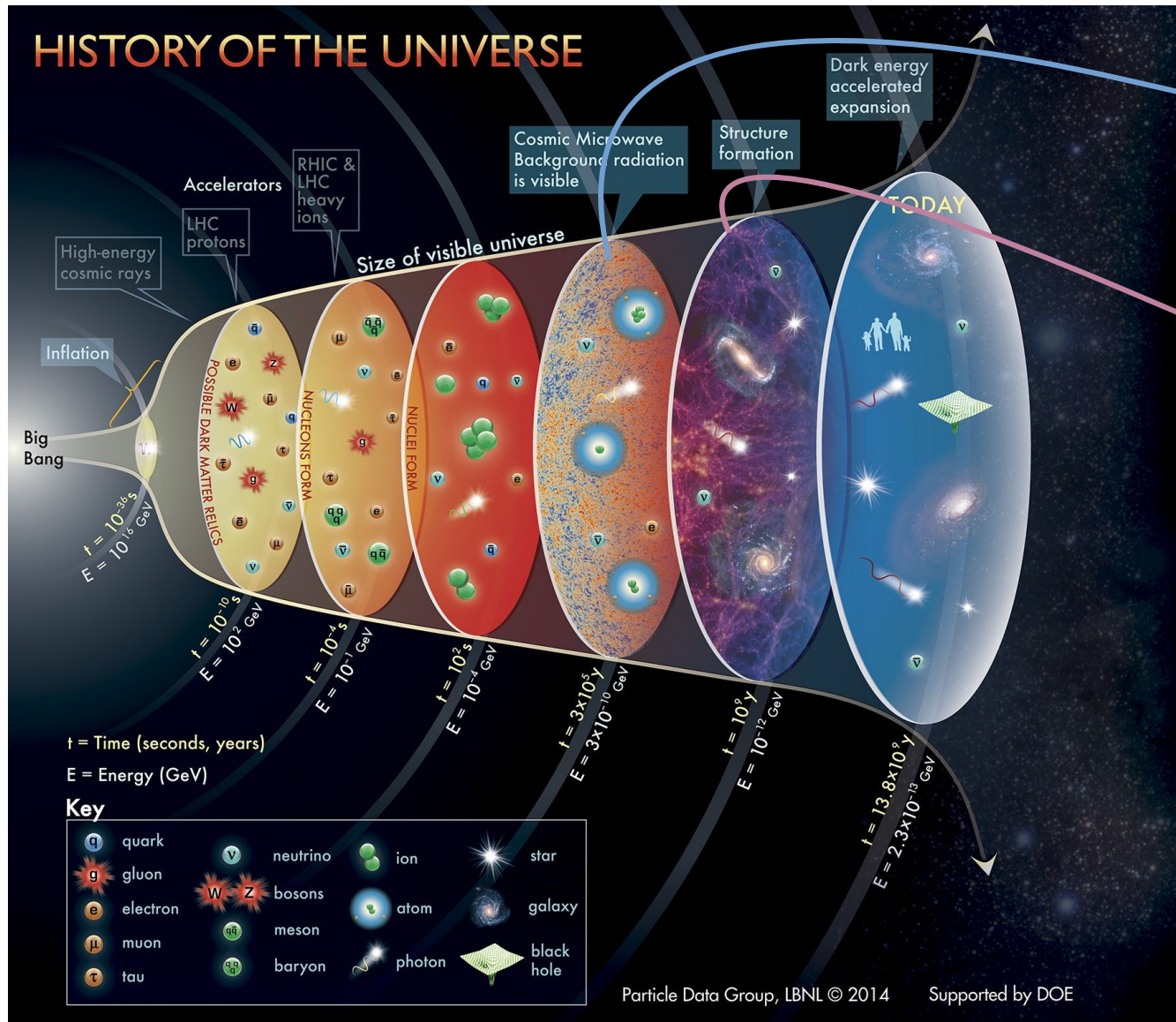


# snapshots of our universe



CMB as relic thermal radiation from the early universe, decoupled in neutral universe

# snapshots of our universe

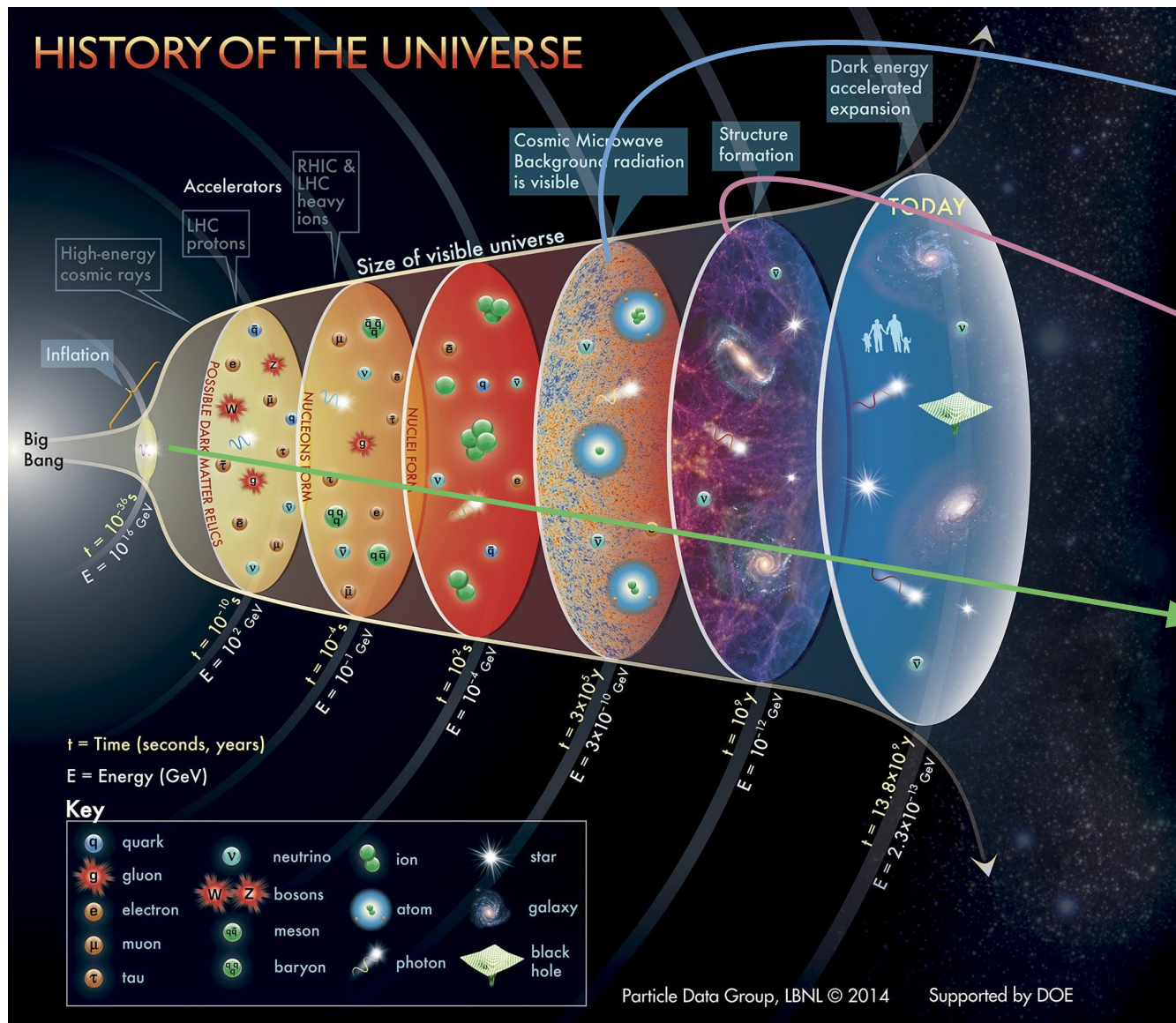


CMB as relic thermal radiation from the early universe, decoupled in neutral universe

CMB anisotropies as sees for galaxy formation



# snapshots of our universe

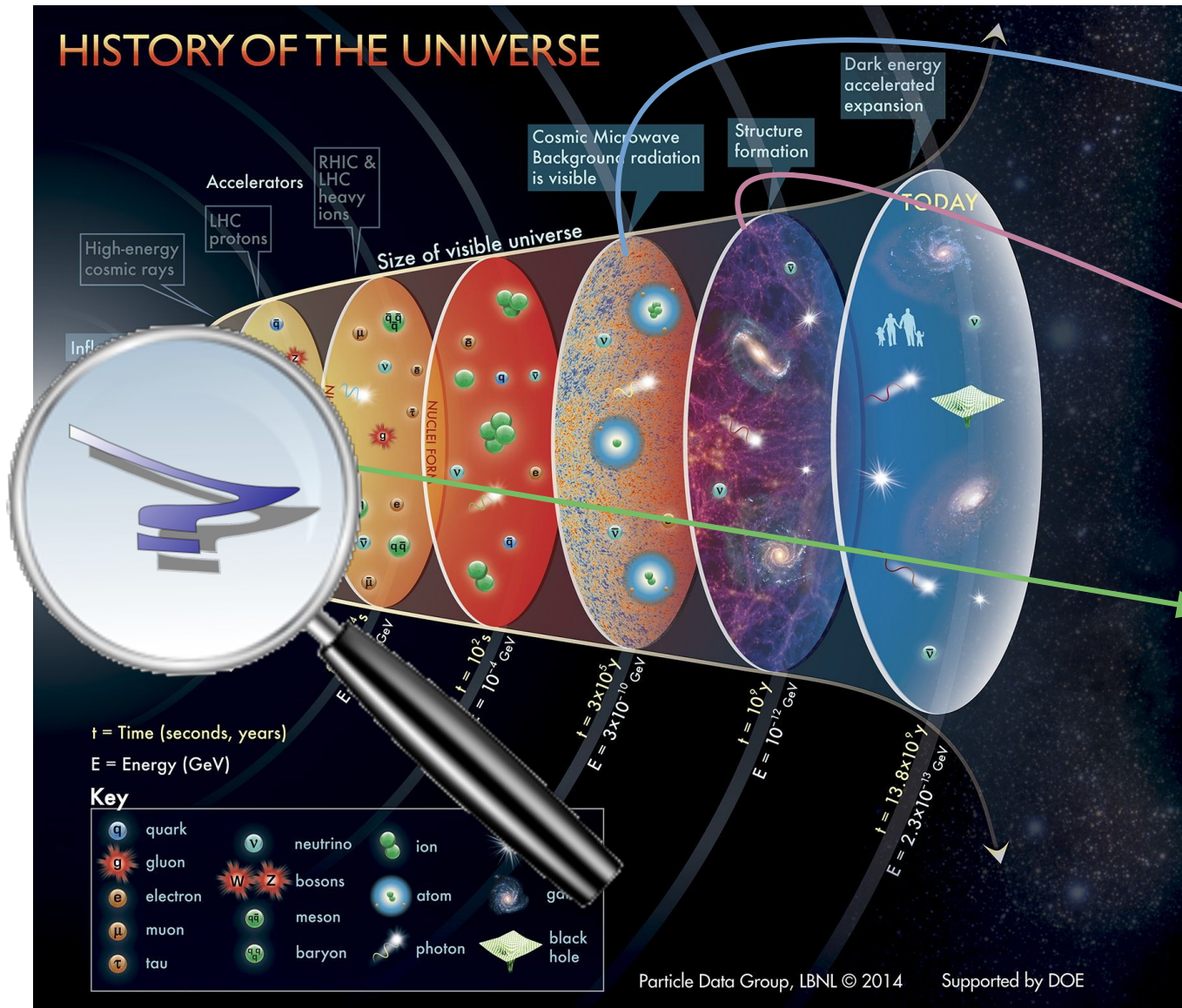


CMB as relic thermal radiation from the early universe, decoupled in neutral universe

CMB anisotropies as sees for galaxy formation

gravitational waves as new window to the early universe

# snapshots of our universe



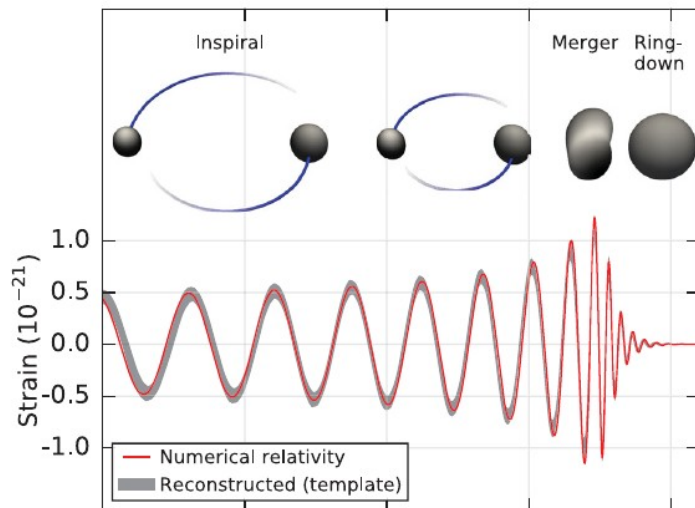
CMB as relic thermal radiation from the early universe, decoupled in neutral universe

CMB anisotropies as sees for galaxy formation

gravitational waves as new window to the early universe



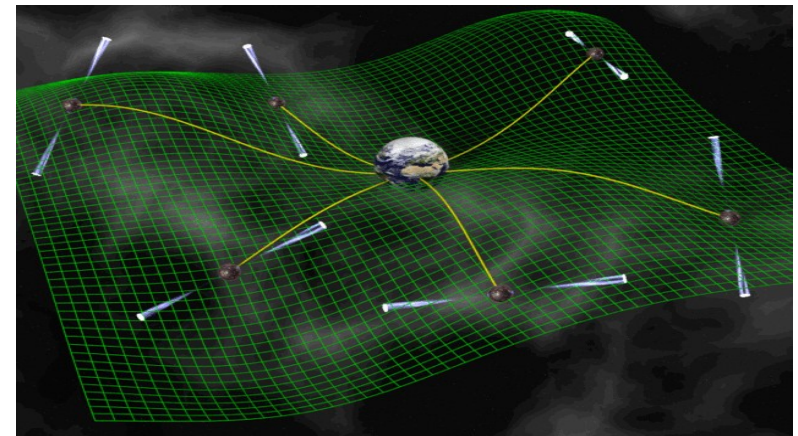
# gravitational waves



LIGO Livingston, USA

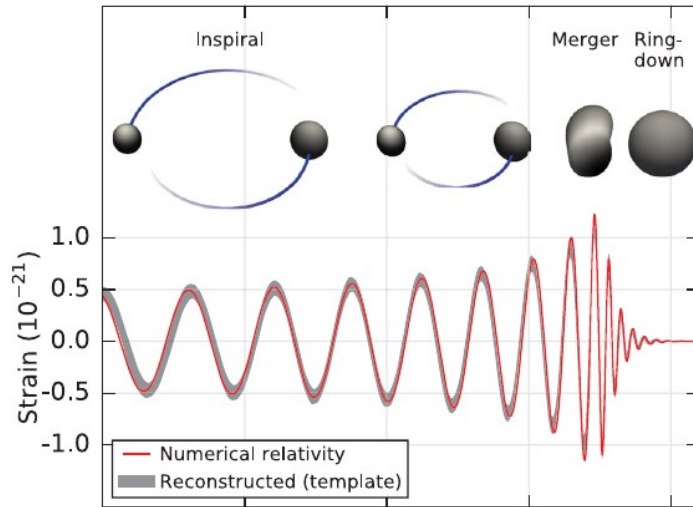


2015: first direct observation of GWs, collision of two black holes a billion years ago





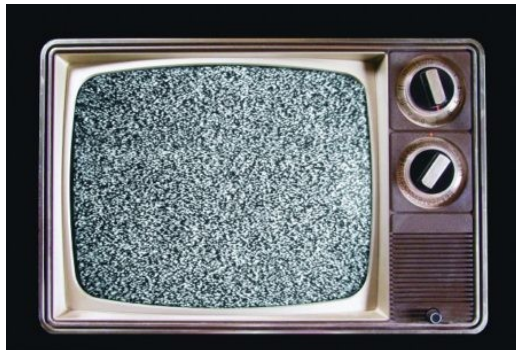
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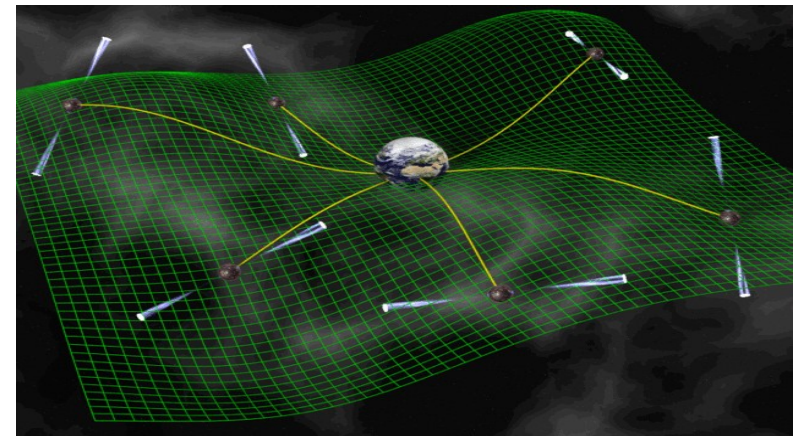


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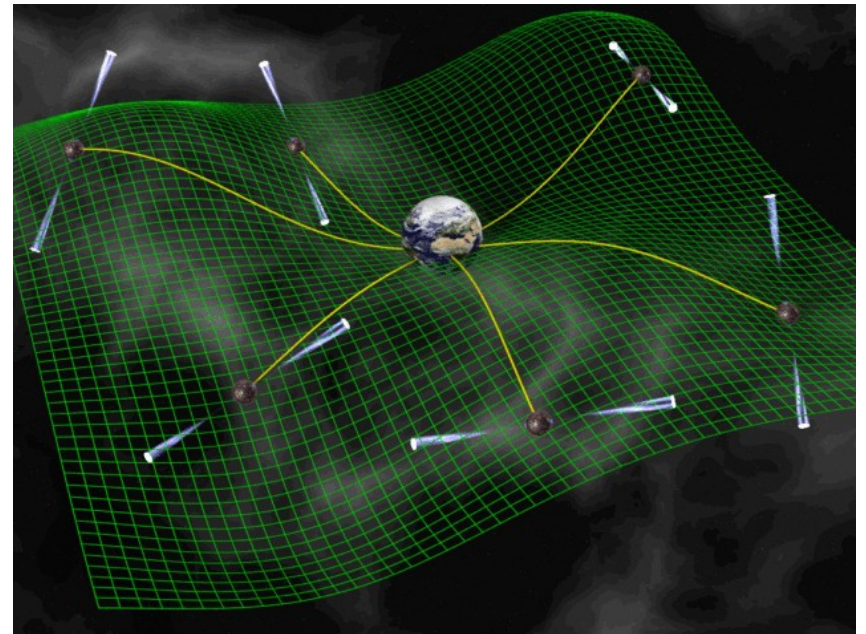
next challenge:  
stochastic gravitational  
wave background

pulsar timing arrays



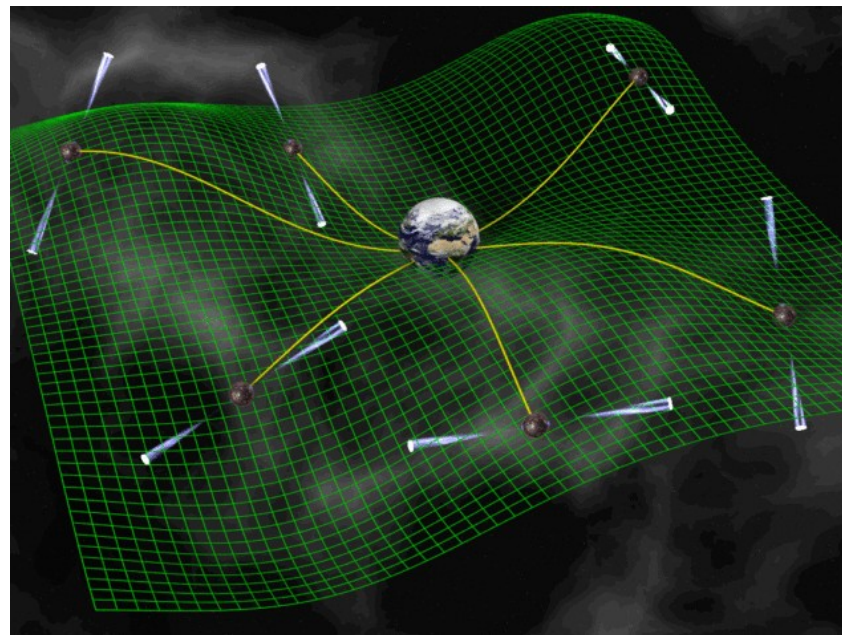
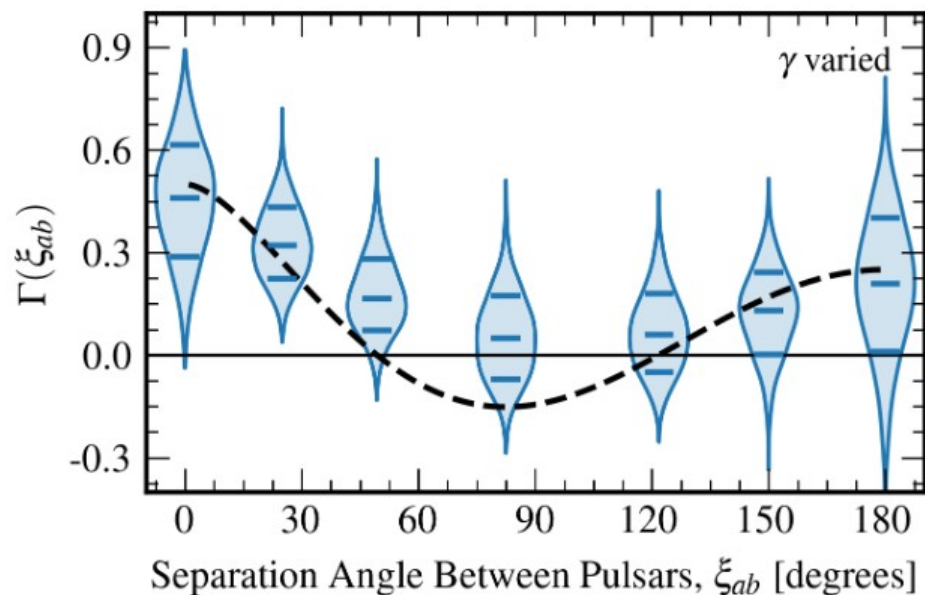
# pulsar timing arrays

- search for delays in pulse arrivals
- 2020: evidence for common stochastic noise component across all pulsars
- 2023: evidence for Hellings-Down correlation (i.e. gravitational waves)



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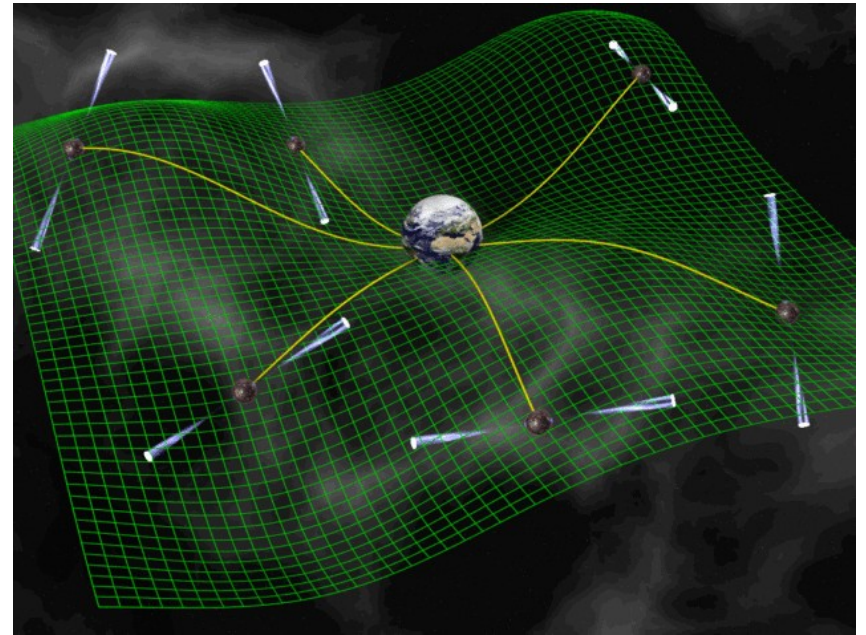
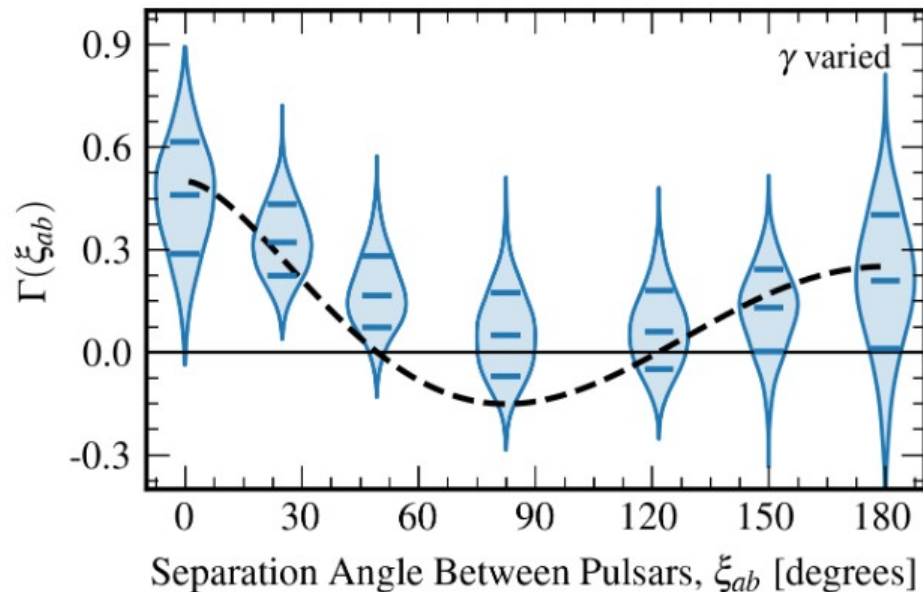
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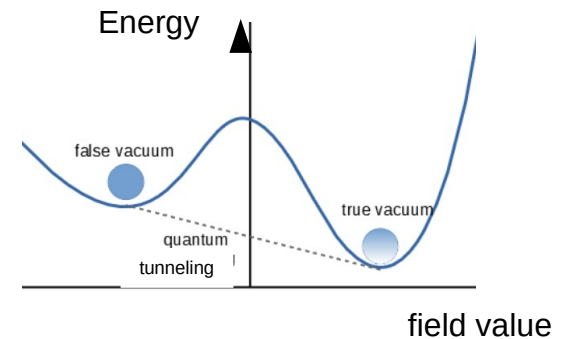


- likely origin: supermassive BH binaries
- SGWB or individual source?  
→ frequency dependence, anisotropy
- cosmological or astrophysical?  
→ anisotropy

# example : first order phase transition

Electroweak symmetry breaking: Cross-over in the SM,  
new physics in the Higgs sector can make it 1<sup>st</sup> order

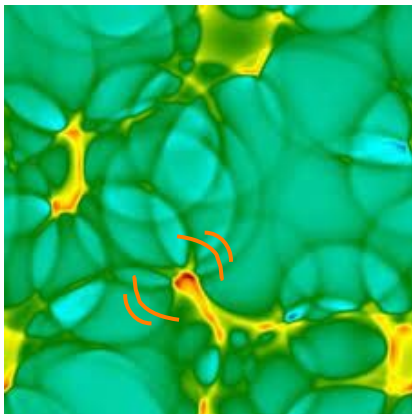
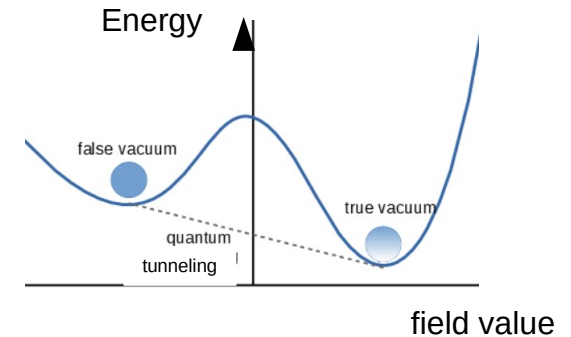
.. and beyond: extended symmetry groups (eg GUTs)  
spontaneously broken in cooling Universe



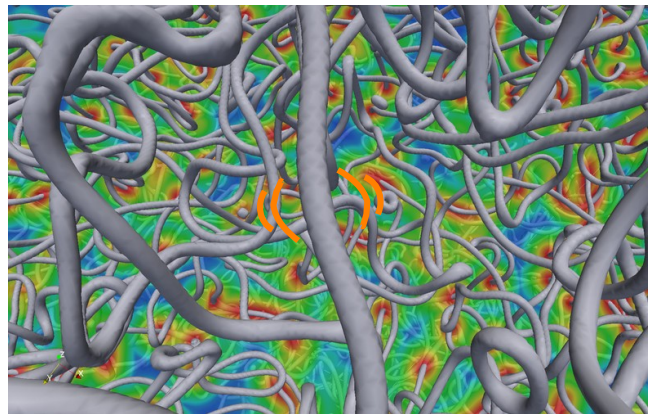
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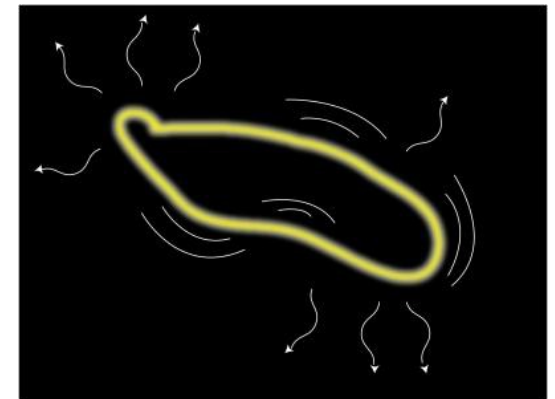
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1st order PT sources GWs



topological defects formed during PT radiate GWs

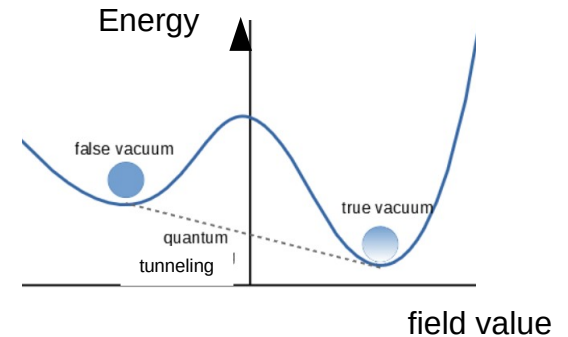




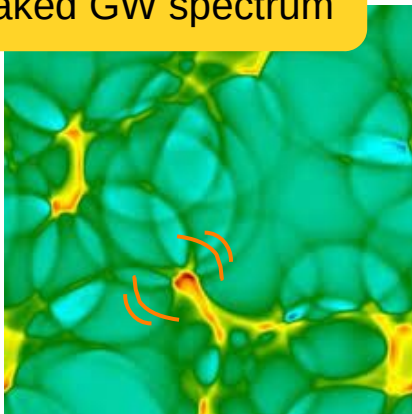
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transient event →  
peaked GW spectrum

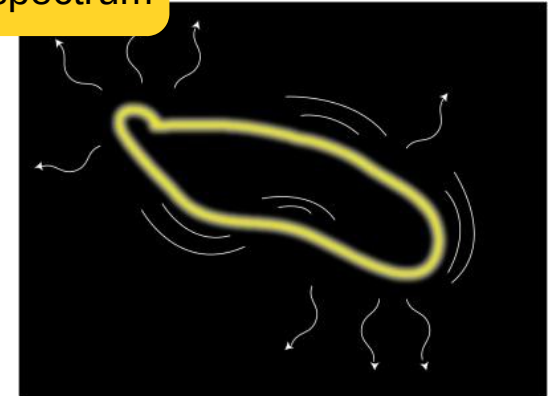


1st order PT sources GWs

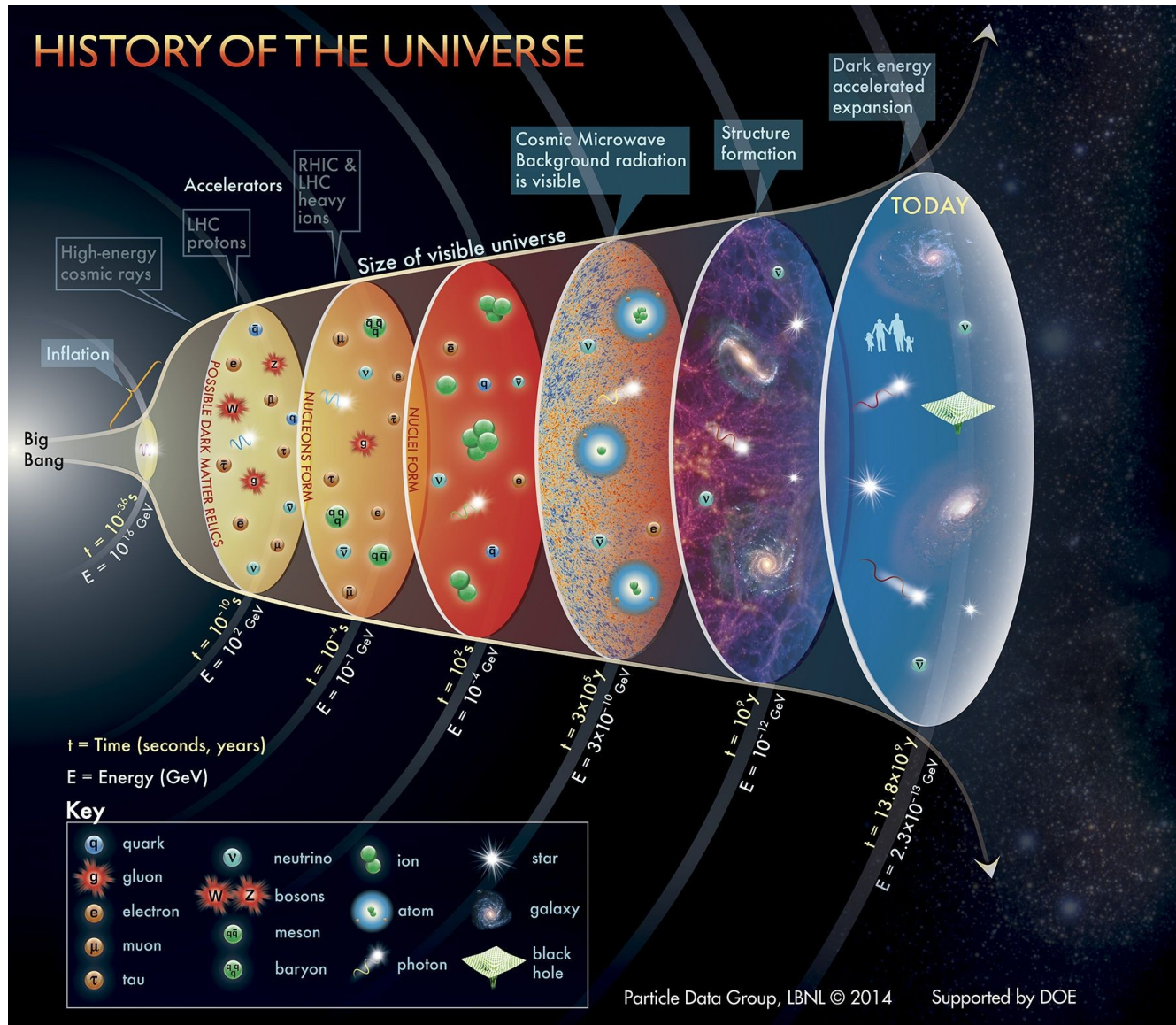
persistent source →  
extended GW spectrum



topological defects formed during PT radiate GWs



# conclusions and outlook

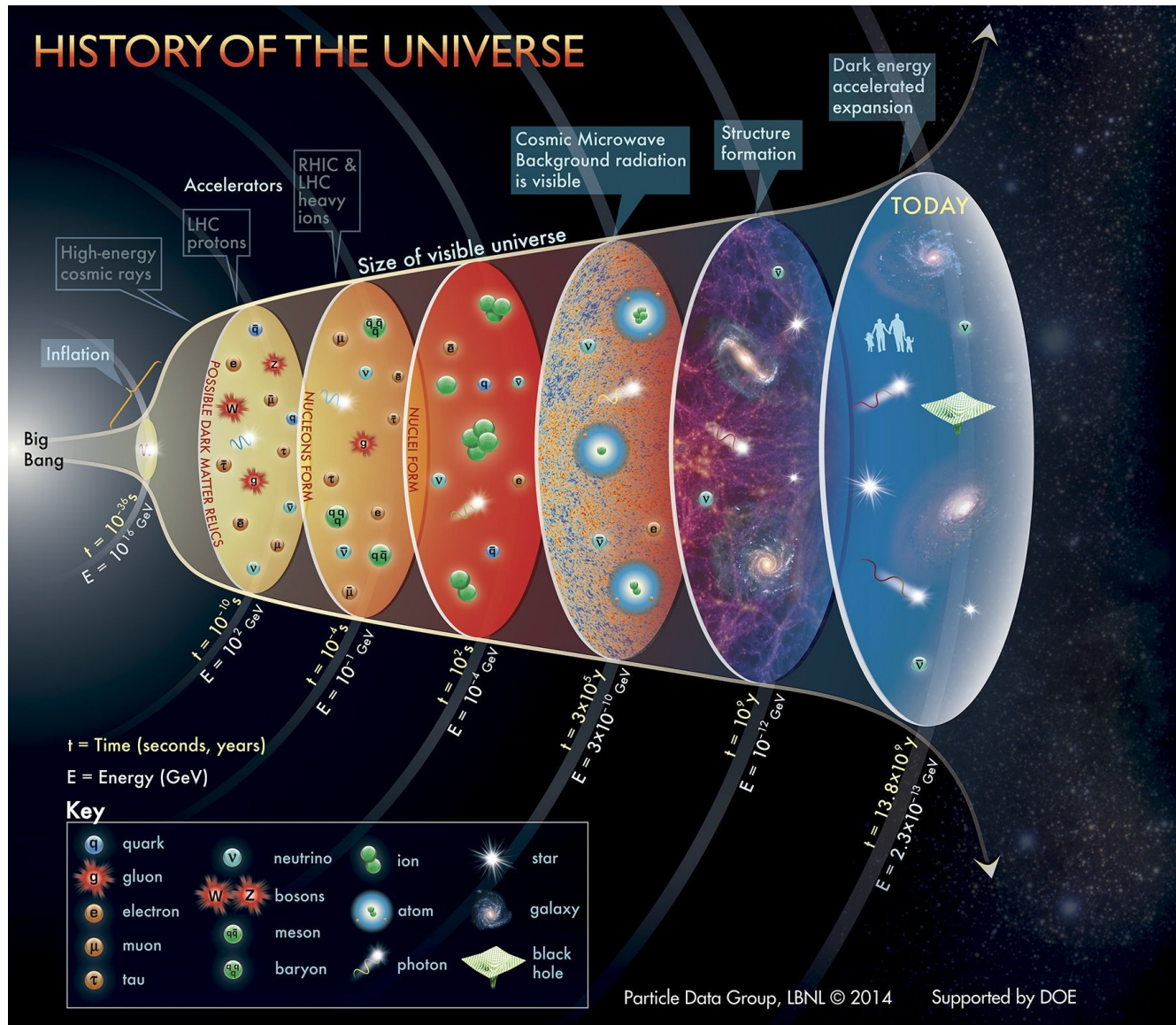


the discovery of the CMB revolutionarized our understanding of the universe

what surprises do gravitational waves reserve for us?



# conclusions and outlook



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**Thank you for your attention !**

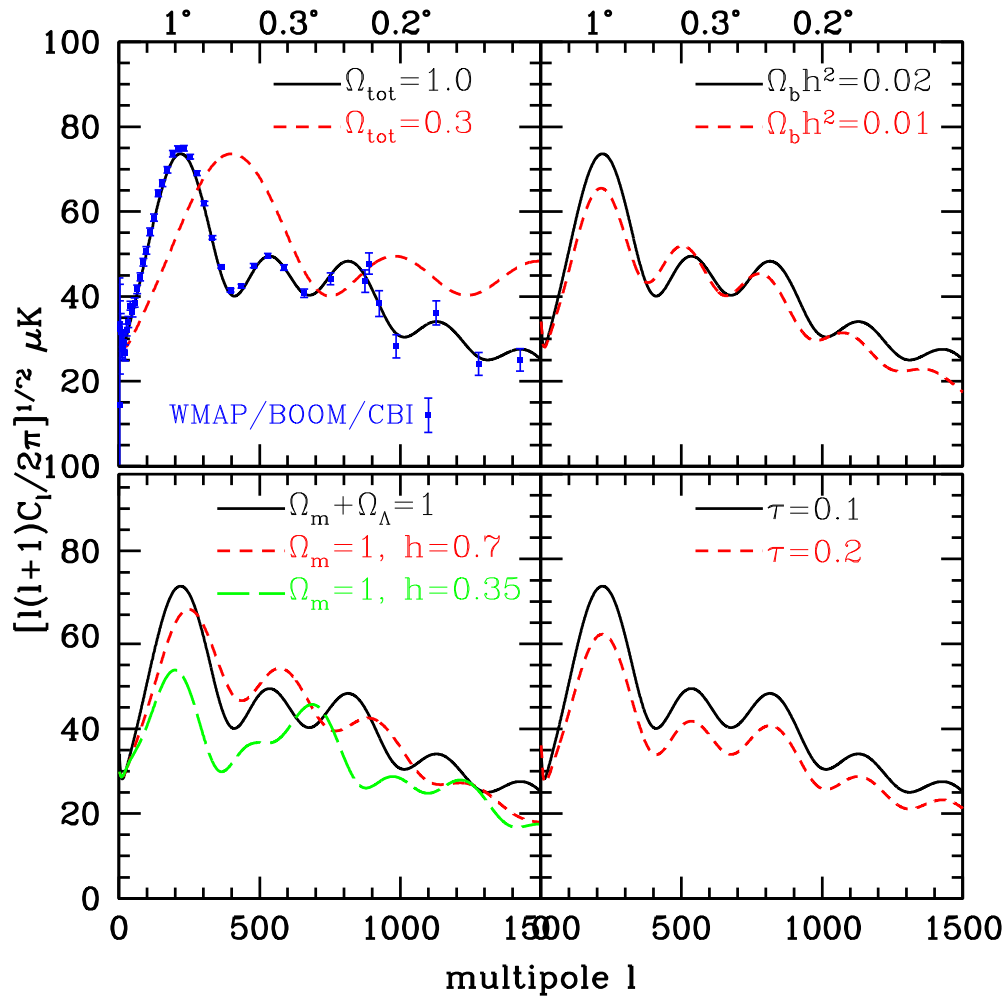


## Backup Slides

# Anisotropien im CMB

Gesamtenergie

Anteil der dunklen Energie



Anteil der Baryonen

Zeitpunkt der Re-ionisation

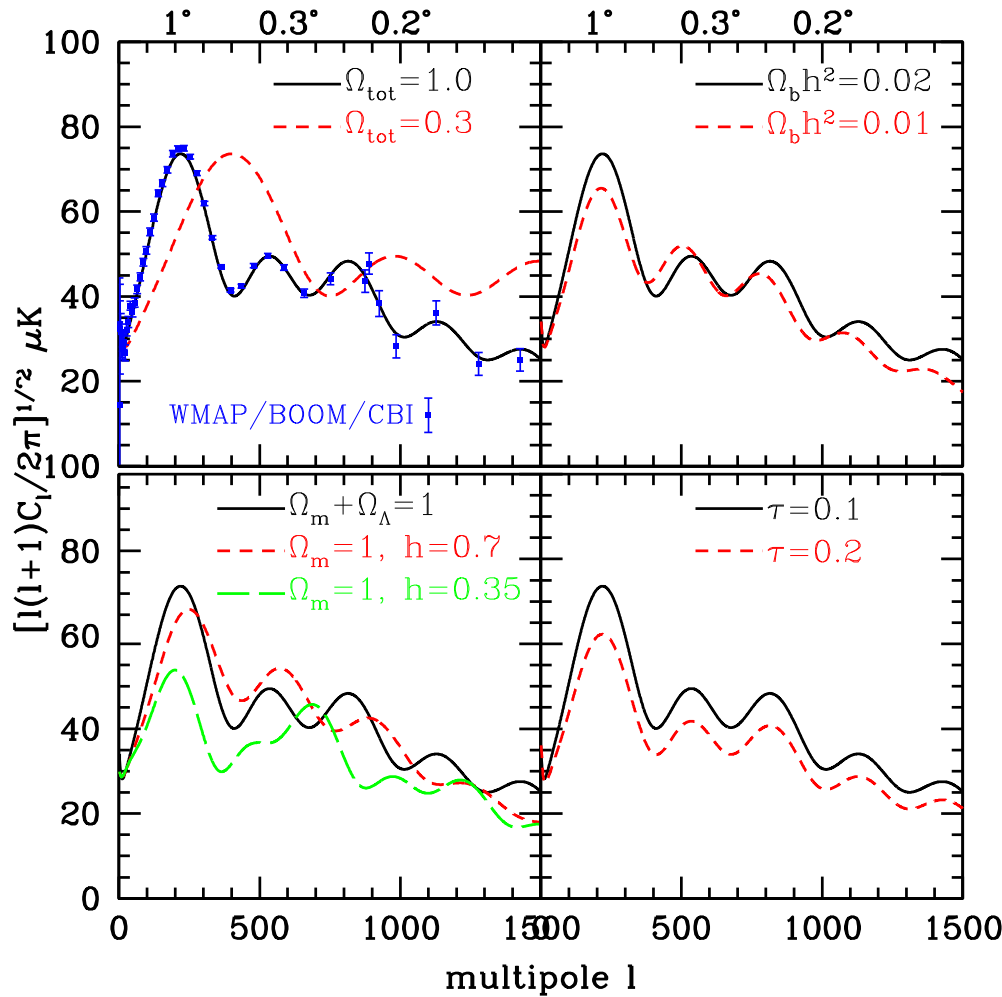
Ermittlung der 6 Parameter durch Fit an die Daten

Kamionkowski '07

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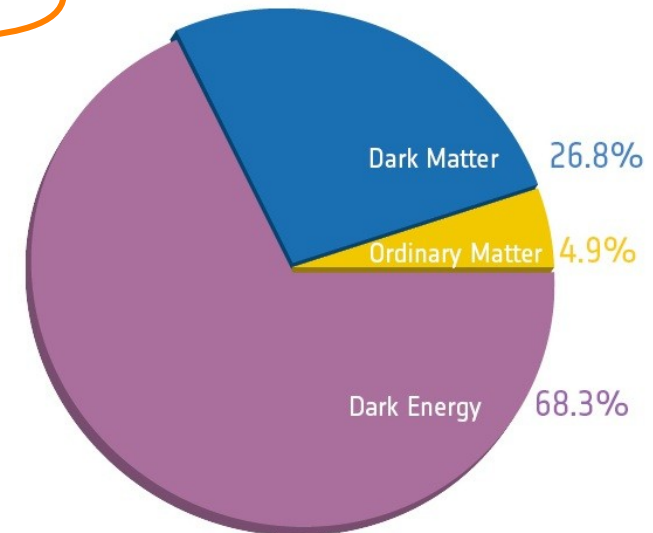


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Standardmodell der Kosmologie



Kamionkowski '07

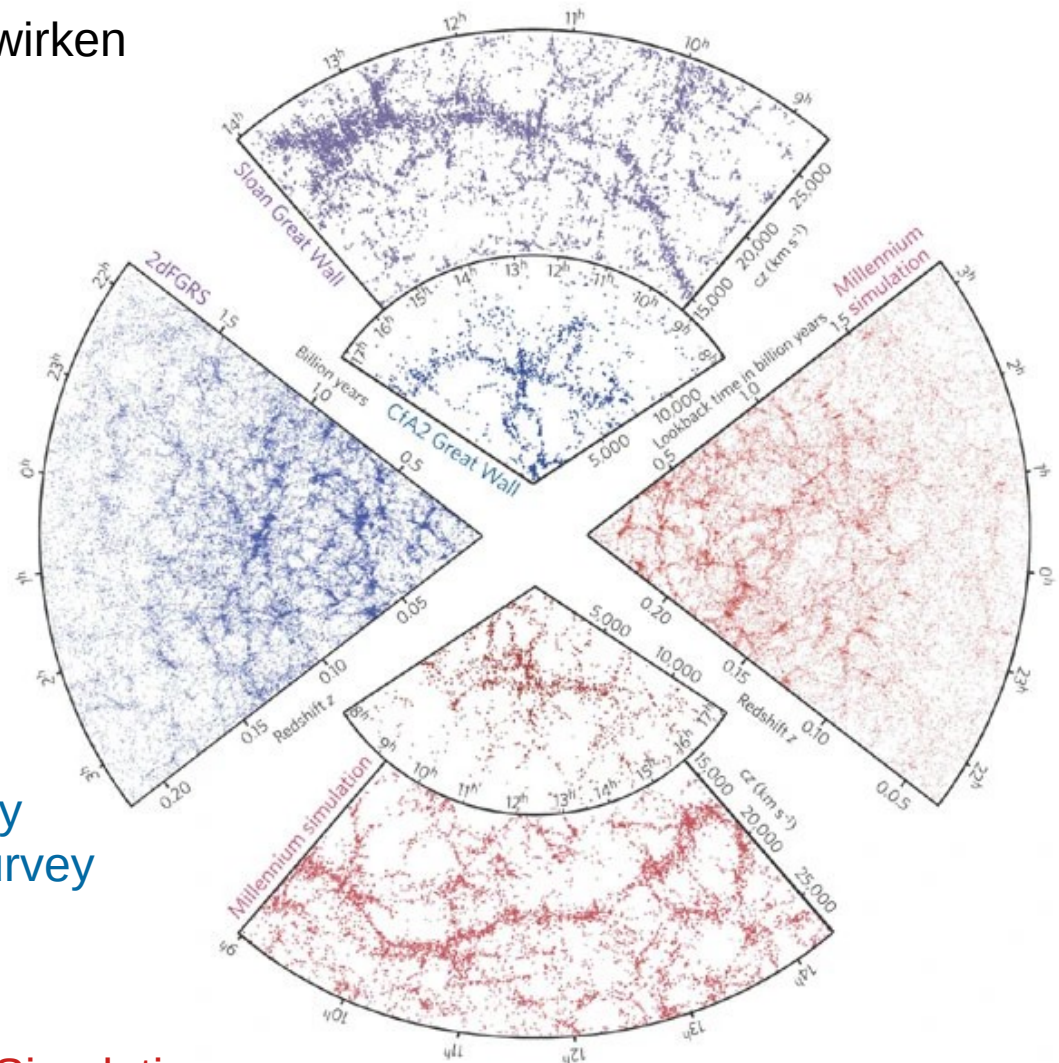
Momentaufnahme mit einigen Überraschungen !



# Entstehung von Galaxien

- Anisotropien im primordialen Plasma wirken als Kristallaktionskerne, Regionen mit höherer Dichte ziehen mehr Materie an (Gravitationskraft)
- Bildung von Sternen, Galaxien und Filamenten
- Statistische Eigenschaften der CMB Anisotropien als Erklärung für die Verteilung von Galaxien

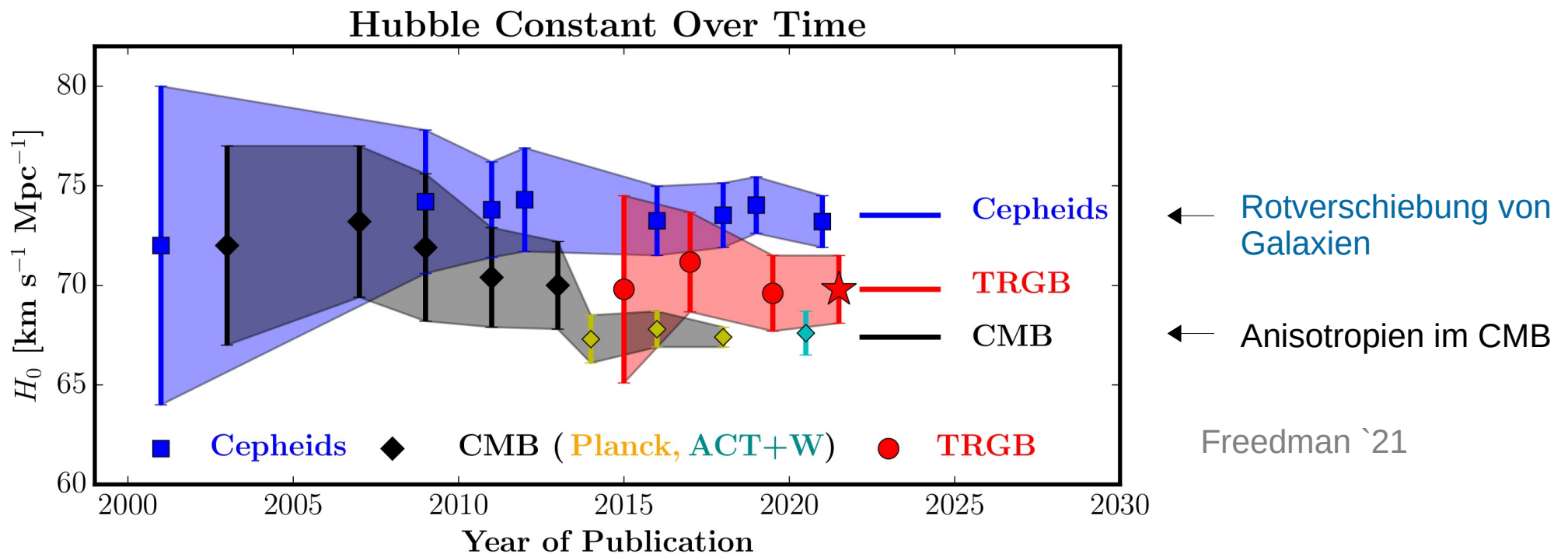
Sloan Digital Sky Survey  
2dF Galaxy Redshift Survey



Millennium Simulation

# Ein neues Rätsel

genauere Messungen weisen eine Diskrepanz bei der Bestimmung der Expansionsrate des Universums  $H_0$  auf:



Unbekannte, systematische Ungenauigkeiten in den Messmethoden?  
Oder ein Hinweis auf eine nötige Erweiterung der Standardmodels der Kosmologie?