

# The neutron lifetime experiment $\tau$ SPECT

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# Ultracold Neutrons - UCN

Subatomic Particles at Human Velocities

$$E_{\text{kin}} \lesssim 335 \text{ neV}$$



$$v < 8 \text{ m s}^{-1} \simeq 30 \text{ km h}^{-1} \simeq 18.6 \text{ mph}$$

# UCN Interactions

- Strong Interaction
  - Neutron Optical Potential (Fermi Potential):
  - $V_F \propto \rho b_{\text{coh}}$
  - $^{58}\text{Ni}$ :  $\sim 335$  neV, Stainless steel:  $\sim 190$  neV, Al:  $\sim 54$  neV

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# UCN Interactions

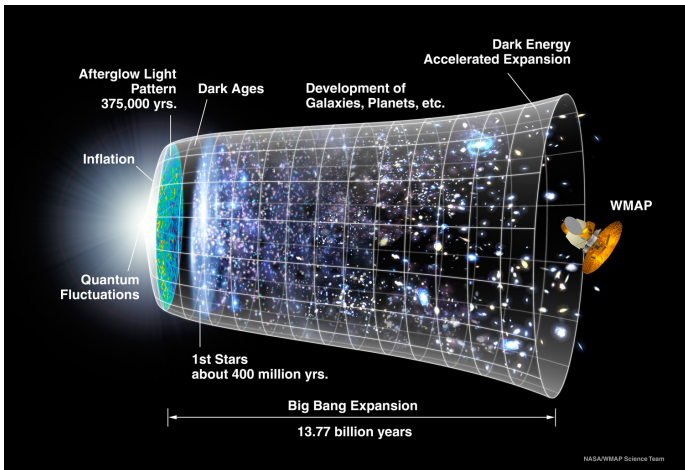
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- Gravity
  - $102.5 \text{ neV m}^{-1}$
- Magnetism
  - Spin polarization with strong magnetic fields.
  - $\mu_n = -60.3 \text{ neV T}^{-1}$

# Why Neutron Lifetime?

## a) Big Bang Nucleosynthesis (He abundance)

[Cyburt et al., doi:10.1103/RevModPhys.88, 2016]

# Big Bang Nucleosynthesis



@t = 2 min:  $n/p \approx 1/6$

@t = 4 min:  $n/p \approx 1/7$

# Neutron Lifetime

## Why n-lifetime?

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### b) CKM Unitarity ( $V_{ud}$ )

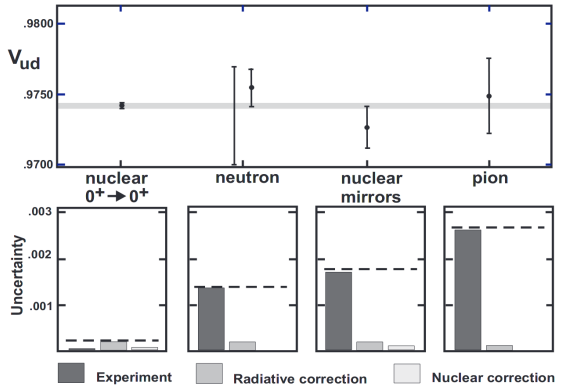
[Czarnecki, Marciano, Sirlin, doi:10.1103/PhysRevD.100.073008, 2019]

# Cabibbo–Kobayashi–Maskawa matrix

$$\begin{bmatrix} |V_{ud}| & |V_{us}| & |V_{ub}| \\ |V_{cd}| & |V_{cs}| & |V_{cb}| \\ |V_{td}| & |V_{ts}| & |V_{tb}| \end{bmatrix}$$

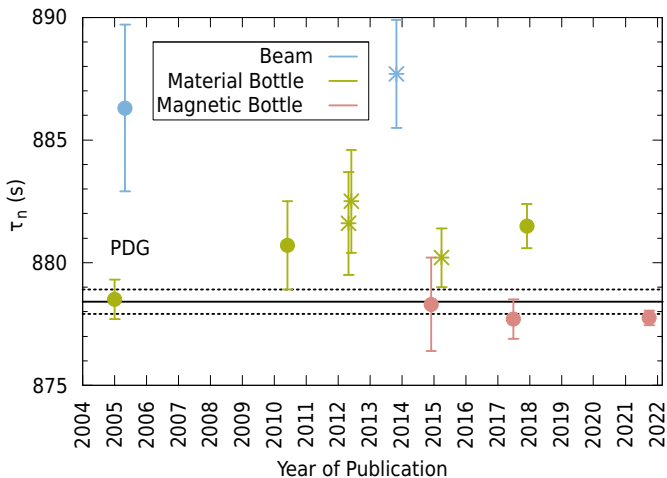
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[Hardy and Towner, doi:10.48550/arXiv.1807.01146, 2018]

# The Lifetime Puzzle



# Neutron Lifetime

## Why n-lifetime?

### a) Big Bang Nucleosynthesis (He abundance)

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### b) CKM Unitarity ( $V_{ud}$ )

[Marciano and Sirlin, doi:10.1103/PhysRevLett.96.032002, 2006]

### c) "It's 2024. We cannot agree on $\tau_n$ to better than 10s?!"

$$\tau_{n,\text{beam}} = 887.7 \pm 1.2 \pm 1.9\text{s}$$

$\neq$

$$\tau_{n,\text{stored}} = 877.75 \pm 0.28 \pm 0.22\text{s}$$



# $\tau$ SPECT

## Concept:

- 3-D magnetic storage
  - Two solenoids + Octupole

# $\tau$ SPECT

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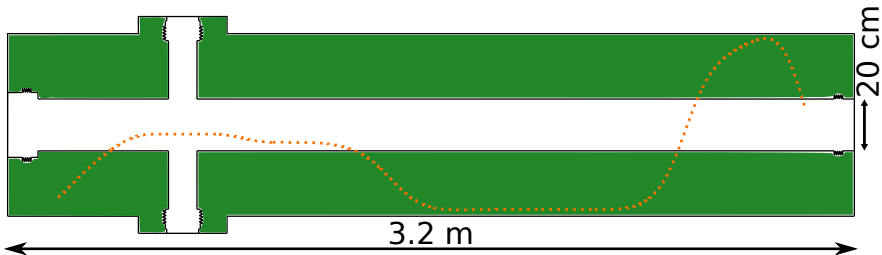
- 3-D magnetic storage
  - Two solenoids + Octupole
- Spinflip-loading
  - Holding field polarizes neutrons
  - Fast adiabatic spinflip as loading mechanism

# $\tau$ SPECT

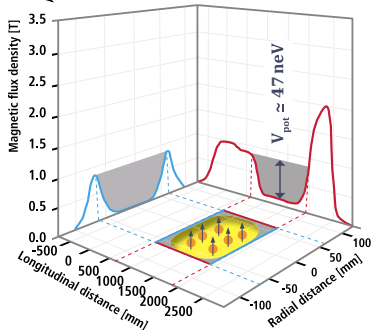
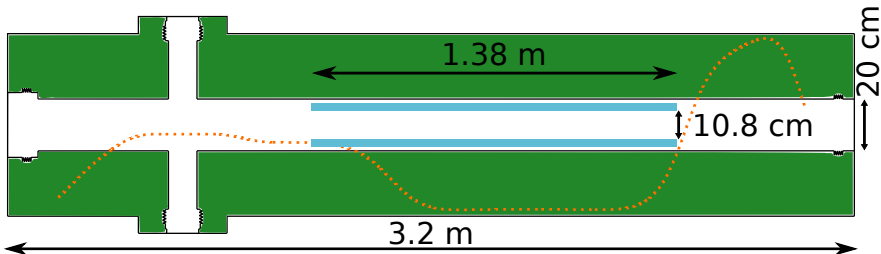
## Concept:

- 3-D magnetic storage
  - Two solenoids + Octupole
- Spinflip-loading
  - Holding field polarizes neutrons
  - Fast adiabatic spinflip as loading mechanism
- In-situ UCN detection
  - Minimizes extraction losses
  - High detector requirements wrt temp. & B-field

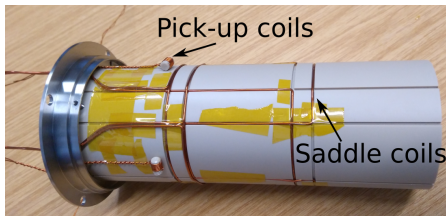
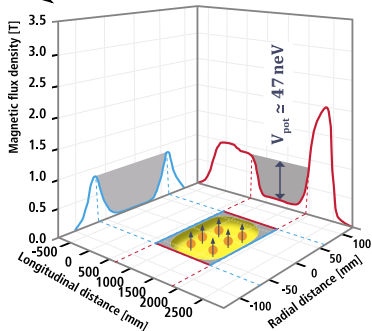
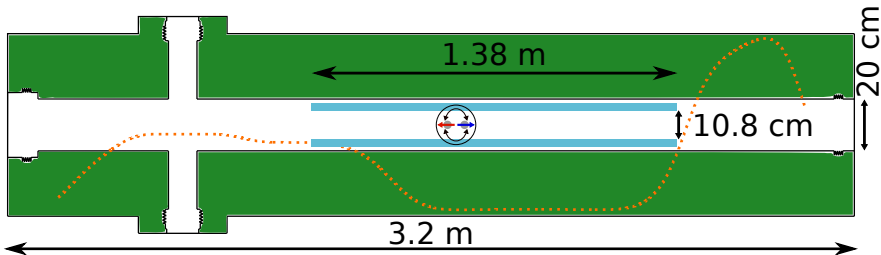
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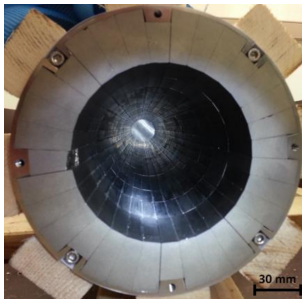
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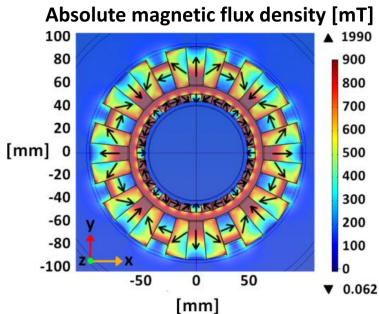
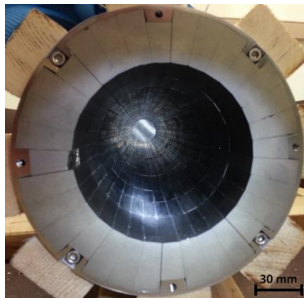


# Gradient compensation in SF region



work of K.U. Ross

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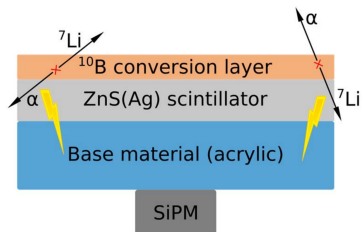


# UCN Detection

Slow neutrons are fundamentally hard to detect  
( = to generate an electric signal )

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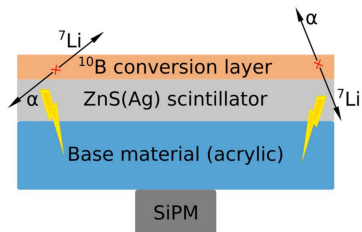
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- Neutron capture on  $^{10}\text{B}$
- Subsequent decay into  $\alpha + ^7\text{Li}$  back-to-back
- Charged particle generates light in scintillator
- Detect light in Silicon Photomultiplier (SiPM)

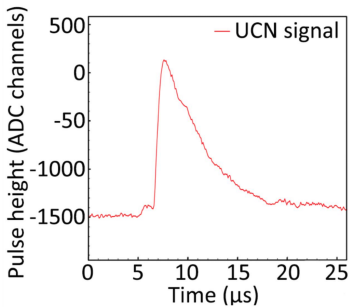
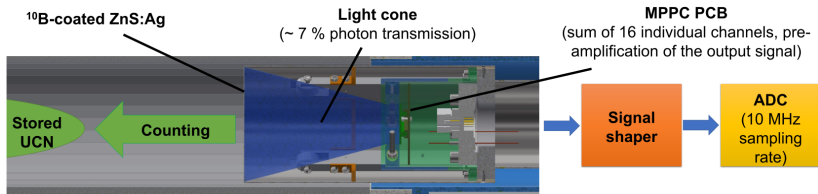
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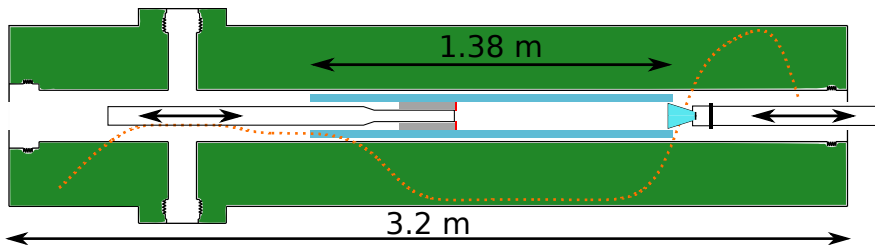
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# $\tau$ SPECT Detector



PhD work of J. Kahlenberg

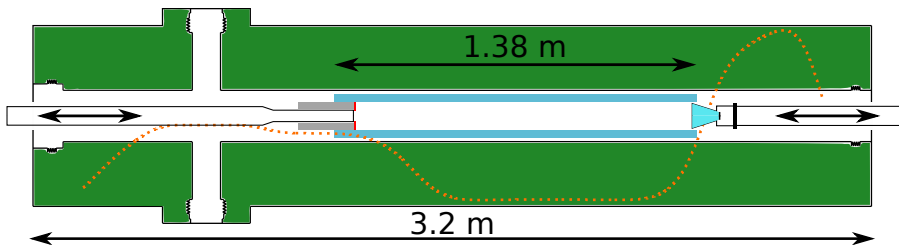
# Measurement Procedure



## 1. Fill UCN into $\tau$ SPECT Magnet from the left

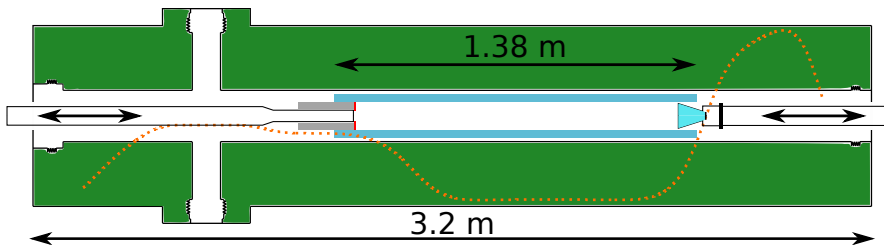
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- Simultaneously: Intensity Monitoring (non-trappable UCN)

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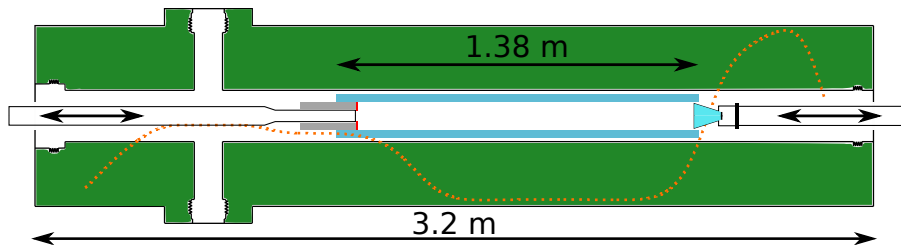
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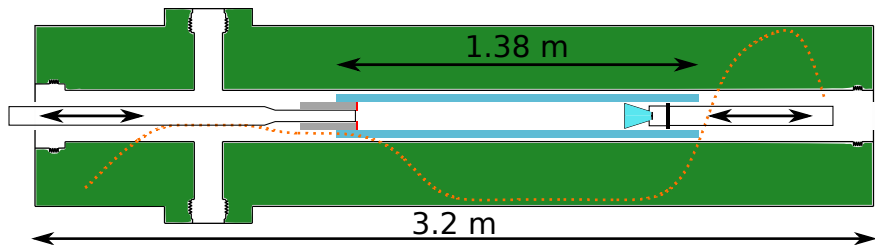
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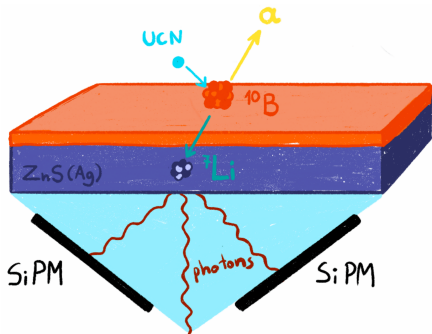
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  - Polarization due to high Magnetic Field, SF on
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4. Wait ...
5. Count UCN

# UCN Intensity Monitor

In order to know  $N(t=0)$ , we need an in-beamline monitor.

# UCN Intensity Monitor

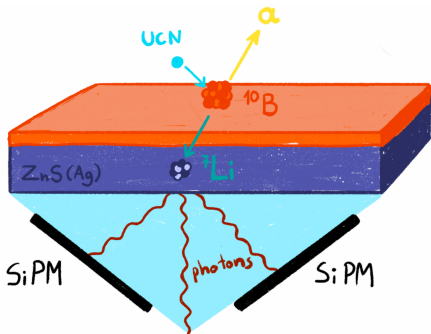
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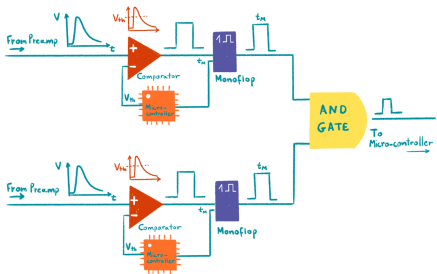
Illustrations: D. Kanta

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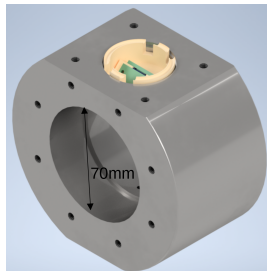
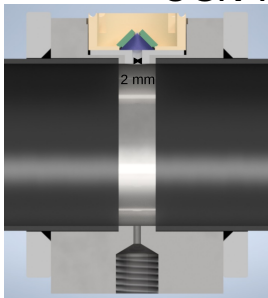
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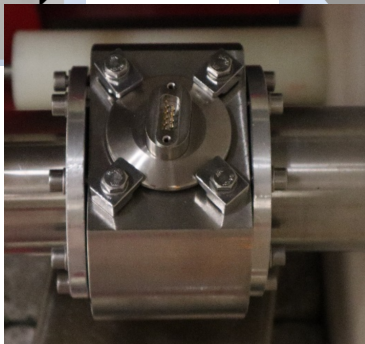
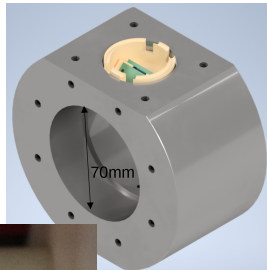
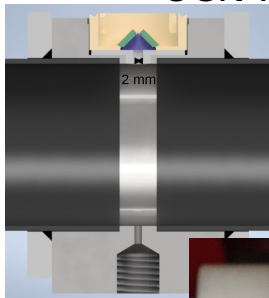
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# UCN Intensity Monitor



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# Systematics

- Gaps:
- Wall losses:

# Systematics

- Gaps:  $\rightarrow 0$  ✓
- Wall losses:  $\rightarrow 0$  ✓
- Depolarisation:



# Systematics

- Gaps:  $\rightarrow 0$  ✓
- Wall losses:  $\rightarrow 0$  ✓
- Depolarisation:  $\ll 0.1$  s ✓
- Rest gas interactions:

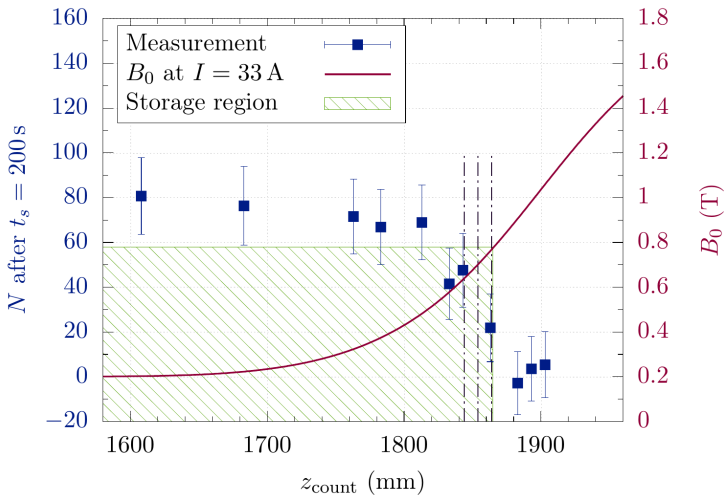
# Systematics

- Gaps:  $\rightarrow 0$  ✓
- Wall losses:  $\rightarrow 0$  ✓
- Depolarisation:  $\ll 0.1$  s ✓
- Rest gas interactions:  $\lesssim 0.1$  s ✓
- Microphonic heating:
- Marginally trapped neutrons:

# Systematics

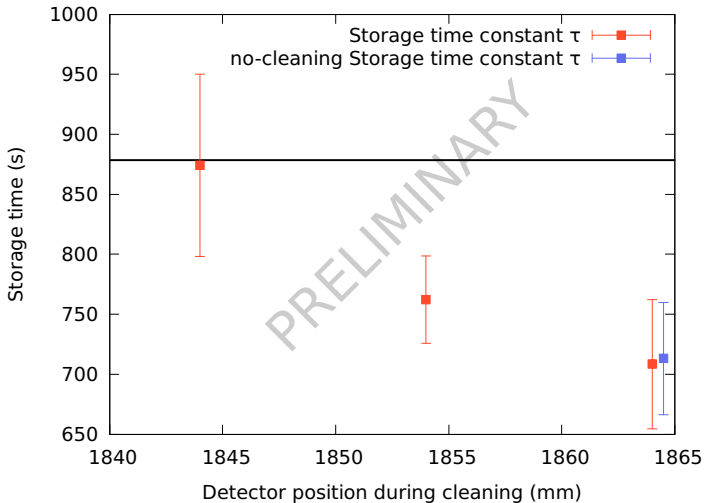
- Gaps:  $\rightarrow 0$  ✓
- Wall losses:  $\rightarrow 0$  ✓
- Depolarisation:  $\ll 0.1$  s ✓
- Rest gas interactions:  $\lesssim 0.1$  s ✓
- Microphonic heating: Has not been observed, measure. ✓
- Marginally trapped neutrons: Spectrum cleaning necessary! ✓

# Countermeasures



K. Ross

# Countermeasures



# Systematics Control

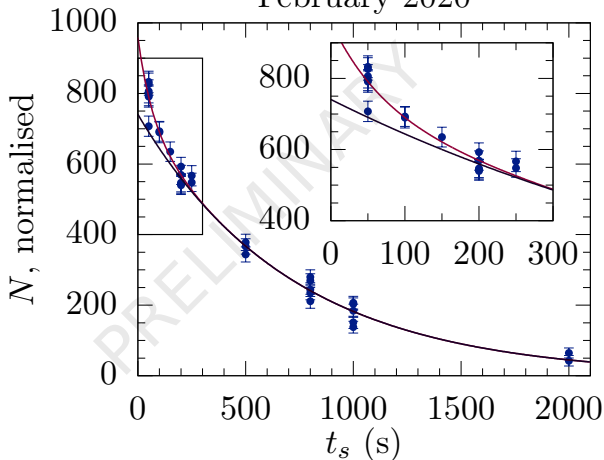
- Marginally trapped neutrons:
  - Clean spectrum with active detector before  $t = 0$
  - Demonstrated to work
  - 2 parameters: position and duration
  - Too aggressive cleaning  $\rightarrow$  lower statistics
  - Introduce asymmetry:  $\tau$ SPECT at a small tilt angle

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  - Too aggressive cleaning  $\rightarrow$  lower statistics
  - Introduce asymmetry:  $\tau$ SPECT at a small tilt angle
- Microphonic heating:
  - Microphonic heating has not been observed
  - Can be measured via change in arrival time
  - Can be measured by "cleaning" again!
  - New  $\tau$ SPECT timing controller can do that.

# Without Energy Spectrum Cleaning

February 2020



Decay times:

Fast:

$$\tau = 64.5 \text{ s}$$

Slow:

$$\tau = 740(47) \text{ s}$$

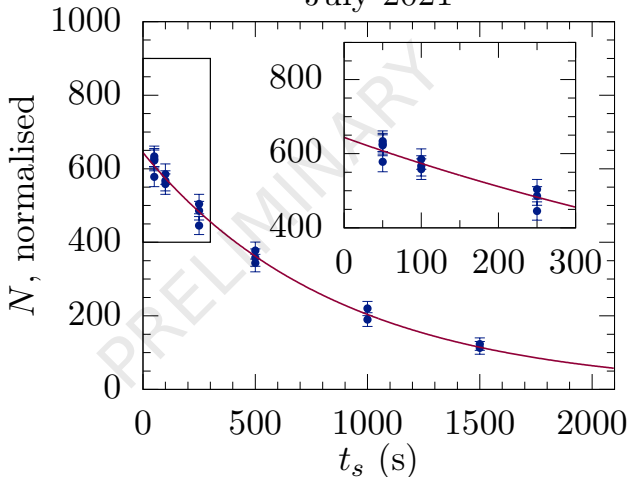
$$\chi^2 = 1.6$$

K. U. Roß



# With Energy Spectrum Cleaning

July 2021



Decay times:

$$\tau = 869(29) \text{ s}$$

$$\chi^2 = 0.6$$

K. U. Roß

# $\tau$ SPECT at PSI



# $\tau$ SPECT at PSI



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Goal: Show statistical reach and systematics control for a physics run aiming for a precision of 0.3 s in the next years.



# Future

- $\tau$ SPECT still statistically limited at PSI.
- UCN source energy spectrum and guides not optimized
- next generation in preparation:
  - Larger volume 10 L to 60 L
  - Stronger magnets 1 T to 3 T
  - almost- $4\pi$  loss detector on magnet inside
  - 0.1 s and below

# Team



+ W. Heil & P. Blümler & B. Lauss & S. Vanneste

## Team



J. Auler<sup>1</sup>, P. Blümler<sup>1</sup>, M. Engler<sup>2</sup>, M. Fertl<sup>1</sup>, K. Franz<sup>2</sup>, W. Heil<sup>1</sup>,  
S. Kaufmann<sup>2</sup>, B. Lauss<sup>3</sup>, N. Pfeifer<sup>1</sup>, D. Ries<sup>3</sup>, S. Vanneste<sup>1</sup>,  
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<sup>2</sup> Institute of Nuclear Chemistry, Johannes Gutenberg University Mainz, Germany

<sup>3</sup> Paul Scherrer Institute, Villigen, Switzerland



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Thank you for your attention!

# Backup