

Status of the FAMU experiment



Cecilia Pizzolotto
for the FAMU Collaboration



PSAS'2024 - International Conference on Precision Physics of Simple Atomic Systems
Zurich, June 2024

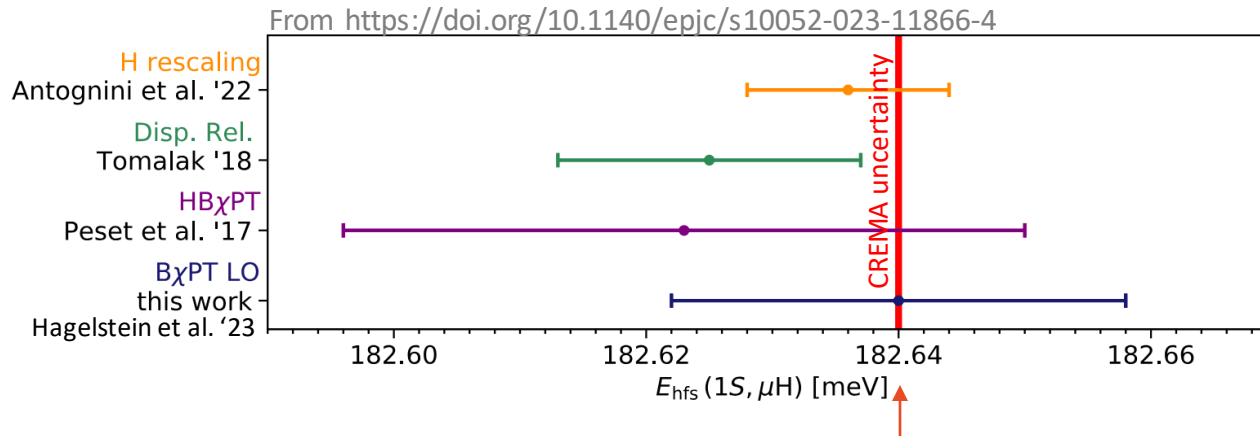


FAMU aims

Measure the Hyperfine splitting (HFS) of muonic hydrogen ground level

→ the Zemach radius of the proton

Theoretical prediction of the HFS energy for muonic hydrogen



FAMU expected uncertainty in the same order of magnitude

FAMU started the frequency scan in 2023 and will continue in 2024 and 2025

FAMU = Fisica degli Atomu MUonici



FAMU experimental hall

≈40 researcher
13 institutions



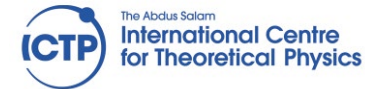
RIKEN Nishina Center
“The RIKEN-RAL Muon Facility”
(International Research Collaboration between RIKEN and SFTC (Science and Technology Facility Council) in the UK)



ALMA MATER STUDIORUM A.D. 1088
UNIVERSITÀ DI BOLOGNA



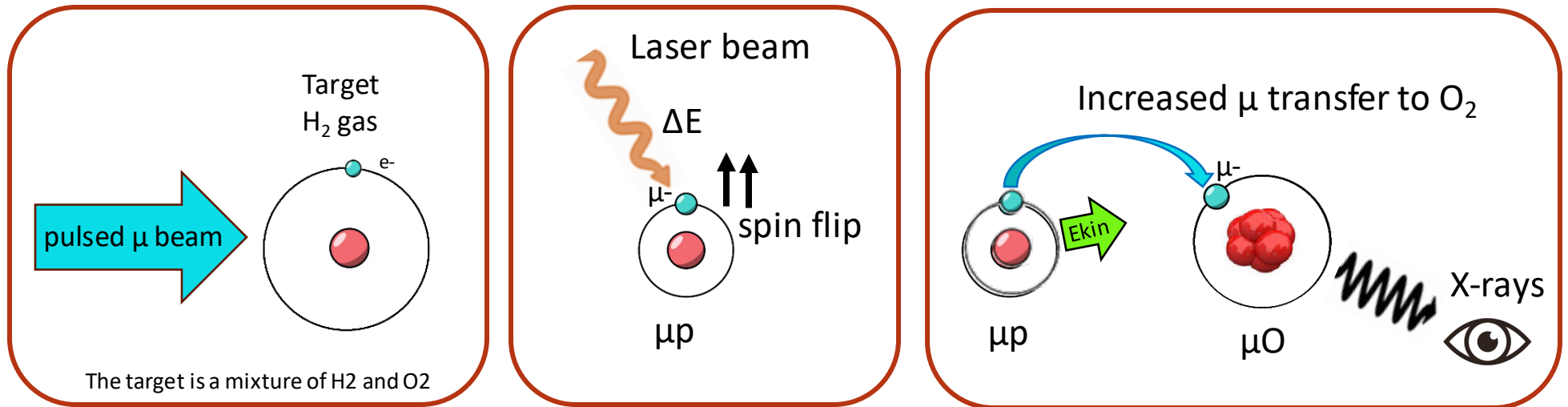
UNIVERSITÀ
DEGLI STUDI
DI UDINE
hic sunt futura



The Henryk Niewodniczański
INSTITUTE OF NUCLEAR PHYSICS
POLISH ACADEMY OF SCIENCES



FAMU method and workflow

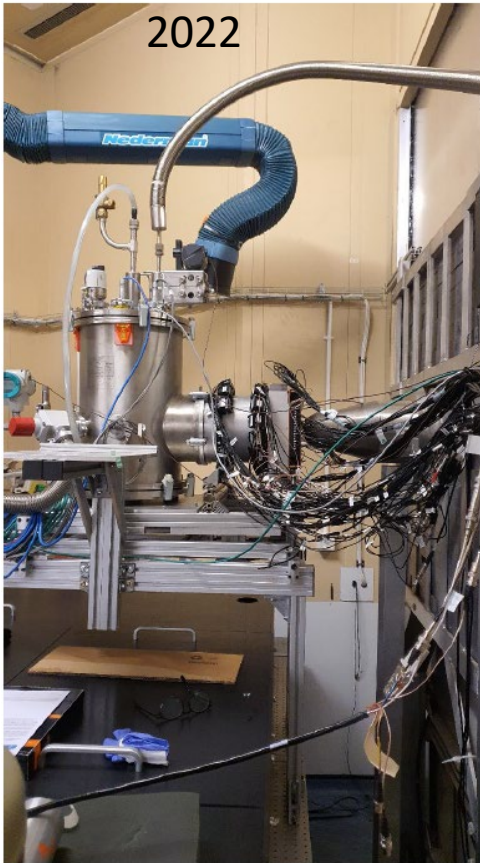


- Create muonic hydrogen and wait for its thermalization;
- Shoot laser at the hyperfine splitting energy ($\lambda_0 \sim 6.8 \mu\text{m}$) and change spin state of μp from 1^1S_0 to 1^3S_1 , spin is flipped: $\mu p(\uparrow\downarrow) \rightarrow \mu p(\uparrow\uparrow)$;
- De-excitation and acceleration of μp (~ 120 meV)
- If μp are accelerated, the μ transfer to Oxygen increases (O₂ has an energy-dependent rate);
- The hyperfine splitting energy is determined by varying the wavelength of the laser beam and search the maximum number of oxygen X-rays

Temperature dependence of the μ transfer to Oxygen
<https://doi.org/10.1016/j.physleta.2021.127401>
<https://doi.org/10.1016/j.physleta.2020.126667>

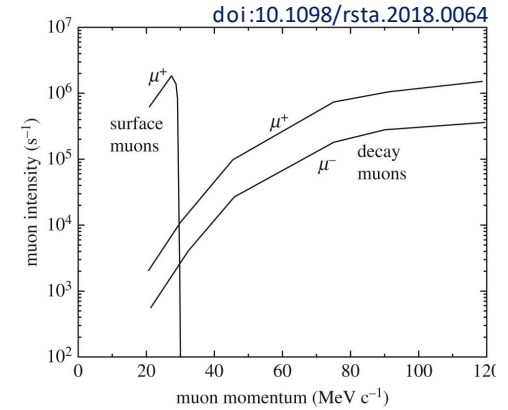
FAMU Characteristics

- high intensity pulsed muon beam
- proper gas mixture H+O
- optical cavity
- high energy and fine-tunable laser
- fast and accurate X-rays detectors

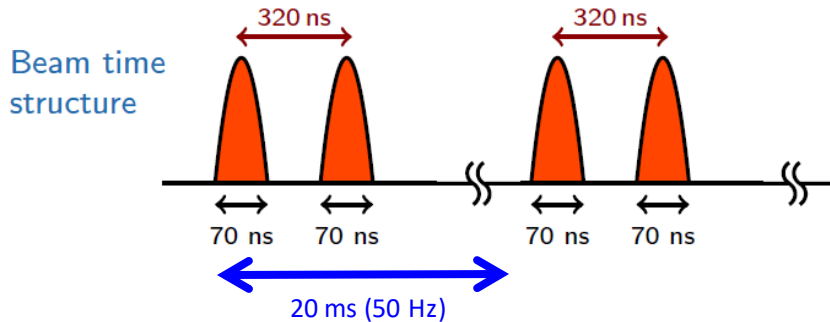
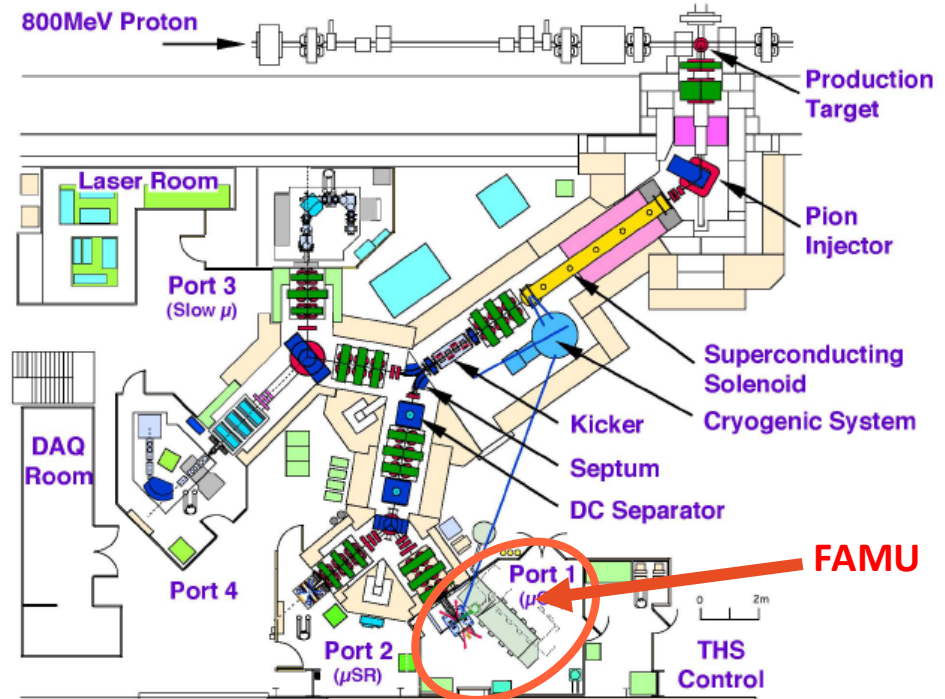


Where: Rutherford Appleton Laboratory - UK

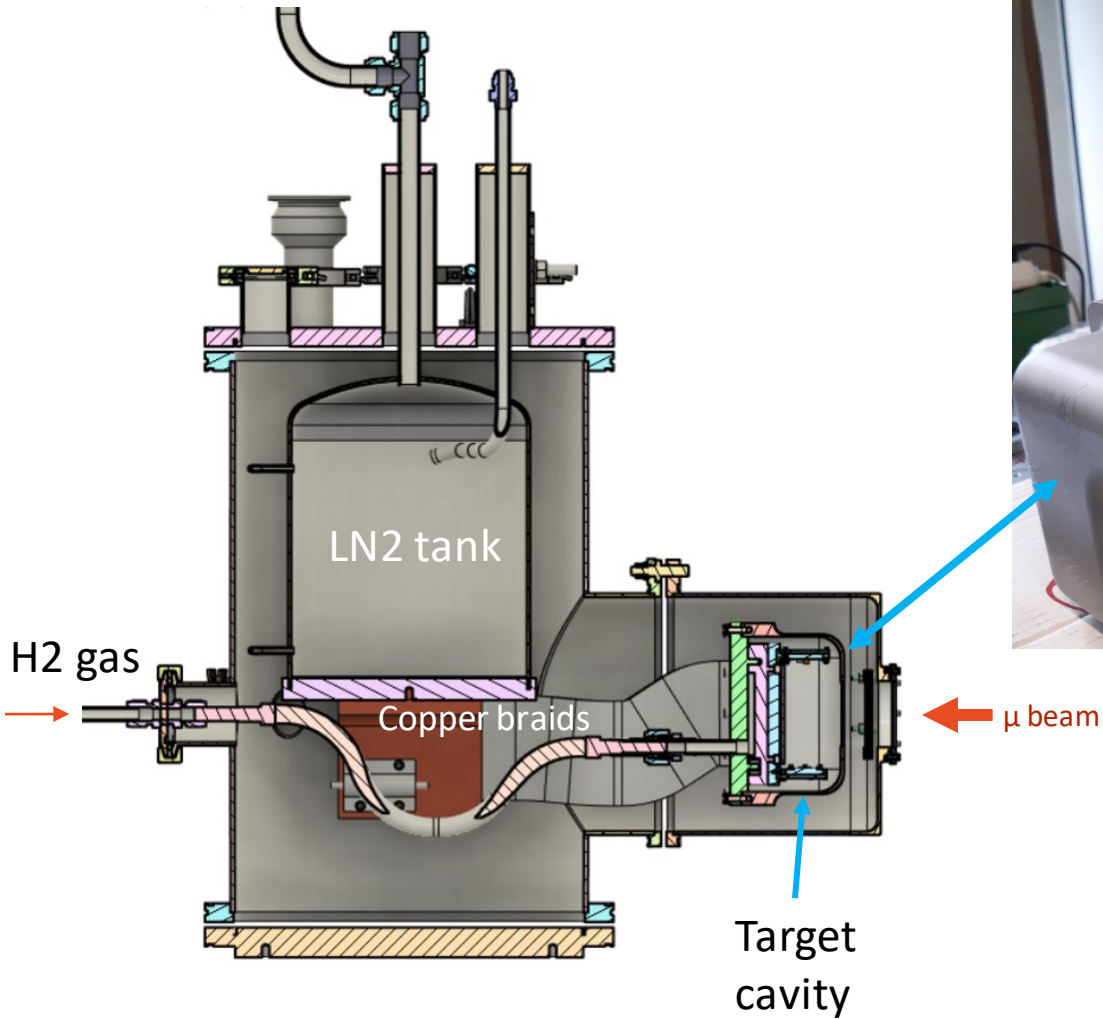
The brightest pulsed muon beam facility in the world!



Riken RAL Muon Facility:



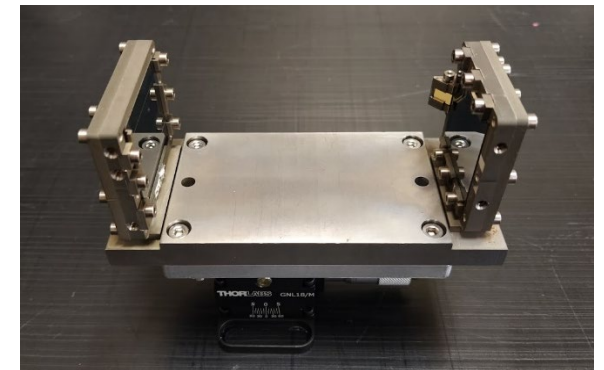
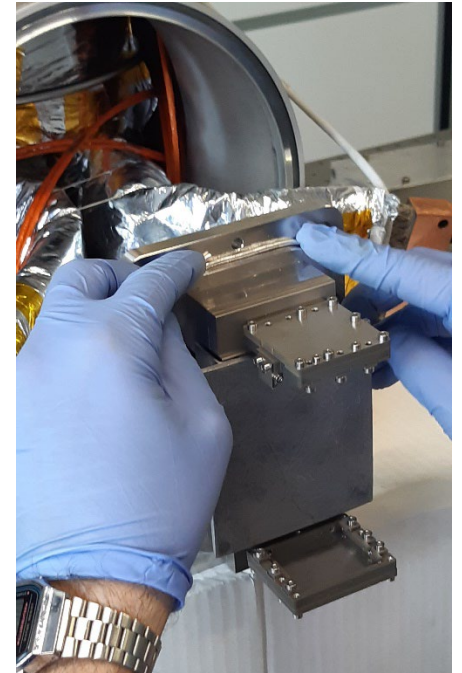
Cyogenic target and optical cavity



Target

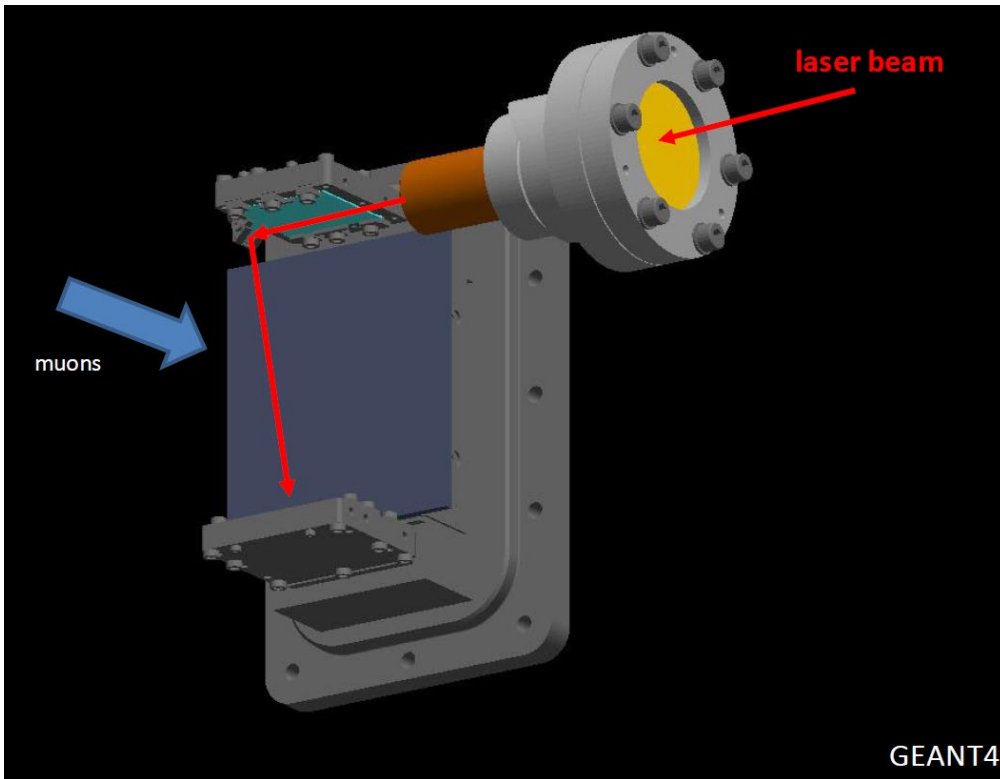


Cavity before inserting the mirrors



Gas mixture of Hydrogen and 1.5% Oxygen at ~ 80 K

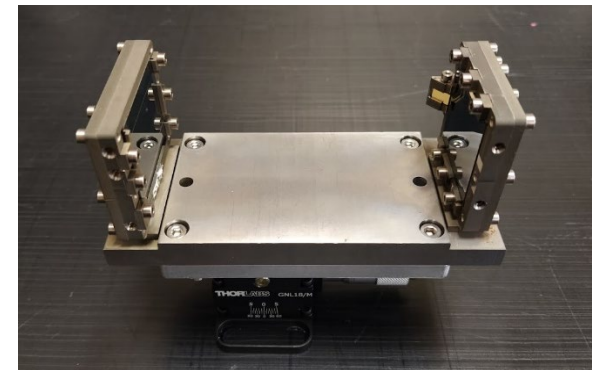
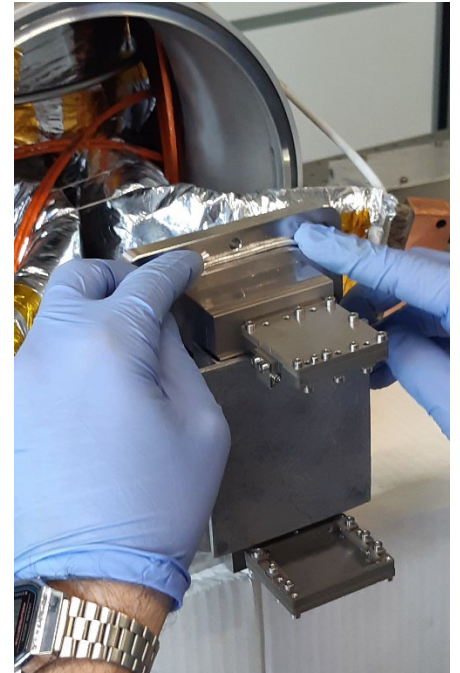
Cyogenic target and optical cavity



Target

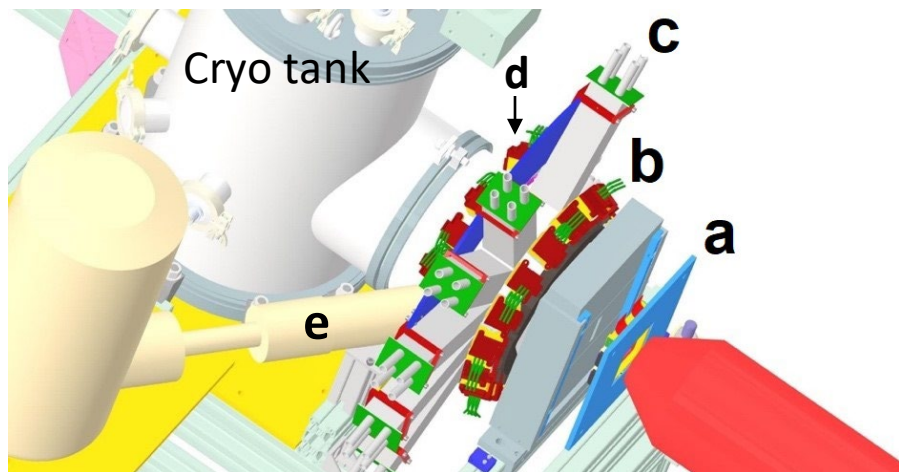


Cavity before inserting the mirrors



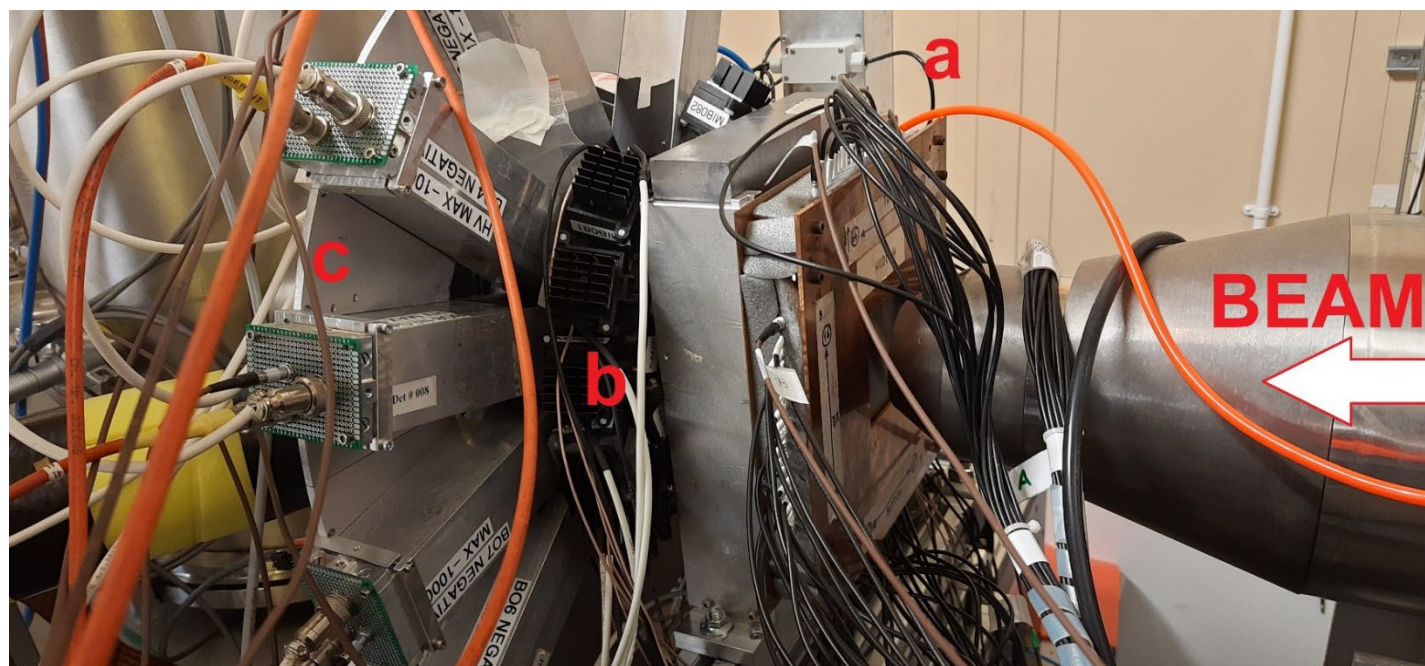
Gas mixture of Hydrogen and 1.5% Oxygen at ~ 80 K

Detectors



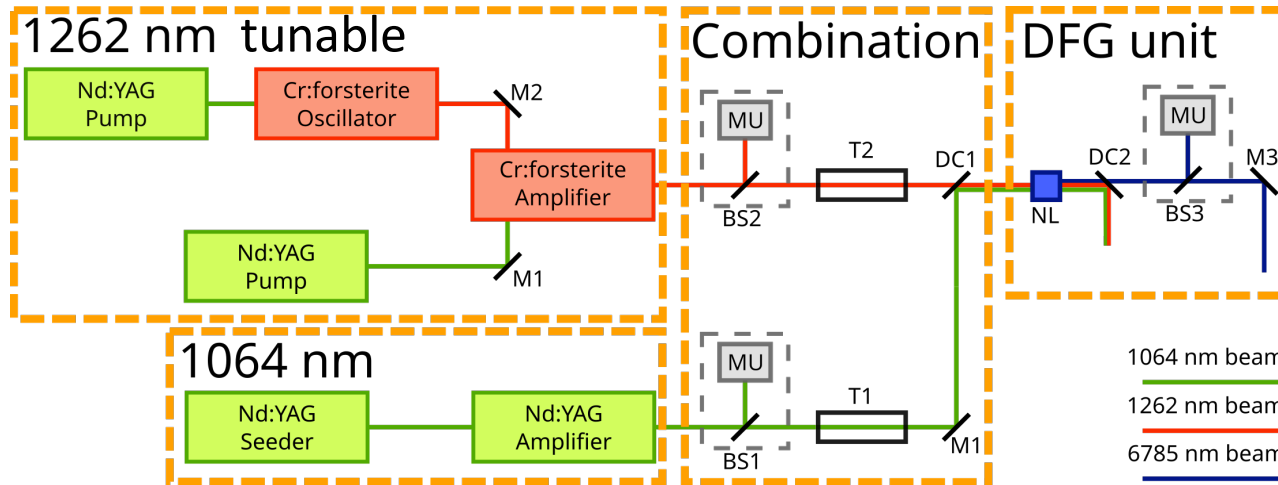
2023 DETECTOR SETUP

- (a) the **beam hodoscope**, 1mm fibers
- (b) the upstream ring of 1 inch **LaBr₃(Ce)** crystals+ SiPM
- (c) the central ring with 6 PMT readout prototypes and SiPM
- (d) (not visible here) backstream ring with ½ inch LaBr₃(Ce) + SiPM
- (e) One HPGe detector for cross calibration



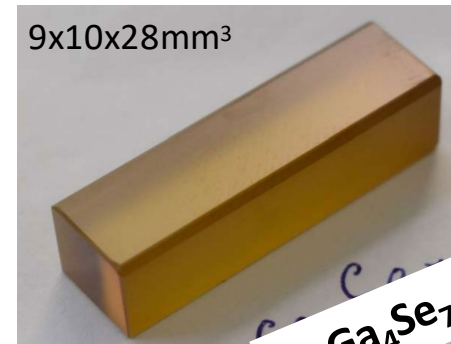
FAMU Laser

Characteristics:	Wavelength range	6800 ± 50 nm	} We reached better values than our goal
	Energy output	> 1 mJ	
	Linewidth	< 30 pm	
	Tunability steps	~9 pm	
	Pulses duration	10 ns	
	Repetition rate	25 Hz	



non linear crystal

9x10x28mm³



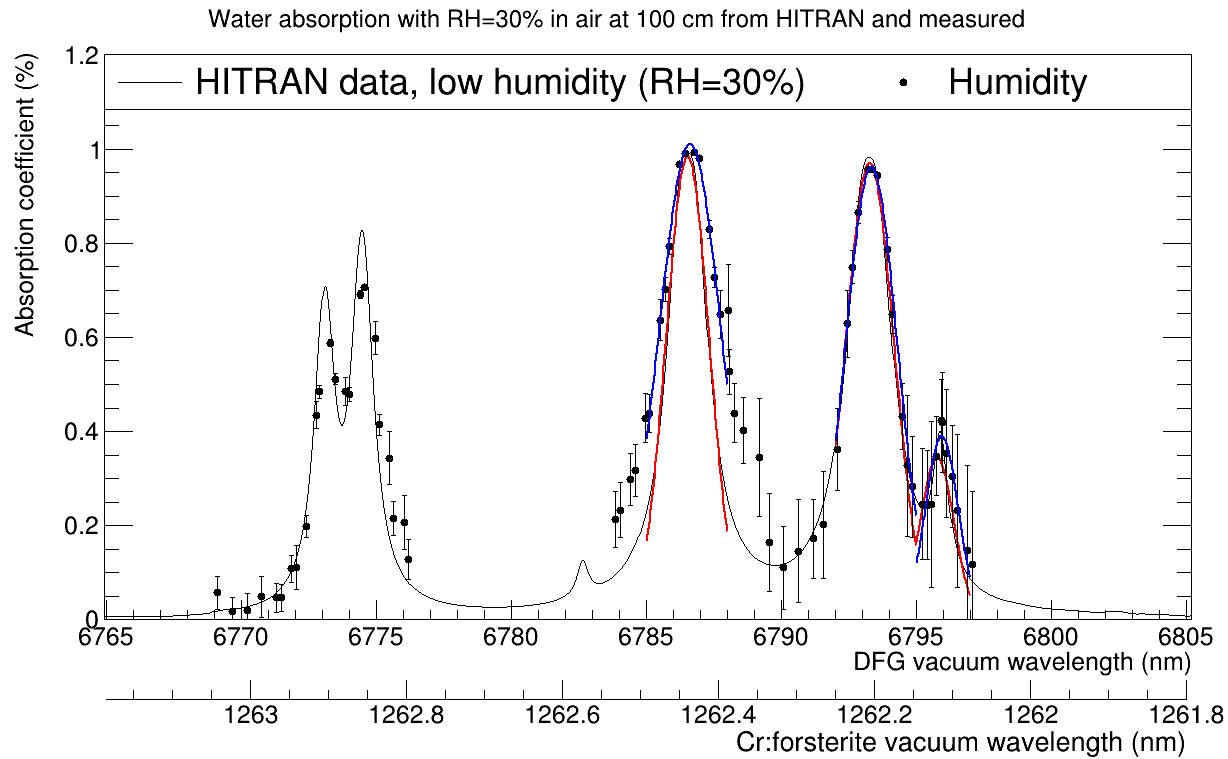
BaGa₄Se₇

$$\lambda_{DFG}^{-1} = \lambda_1^{-1} - \lambda_2^{-1}$$

M1 - Mirror HR 1064 nm, M2 - Mirror HR 1262 nm, M3 - Mirror HR 1064&1262&6785 nm, M4 - Mirror HR 6785 nm, T1 and T2 - telescopes, BS1 - beamsplitter/beamsampler 1064 nm, BS2 - beamsplitter/beamsampler 1262 nm, BS3 - beamsplitter/beamsampler 6785 nm, DC1 - dichroic mirror (reflecting 1064 nm, transmitting 1262 nm), DC2 - dichroic mirror (reflecting 1064 nm and 1262 nm, transmitting 6785 nm), NL - nonlinear crystal, MU - measuring units (wavelength meter, energy meter, dimensions)

Laser absolute calibration

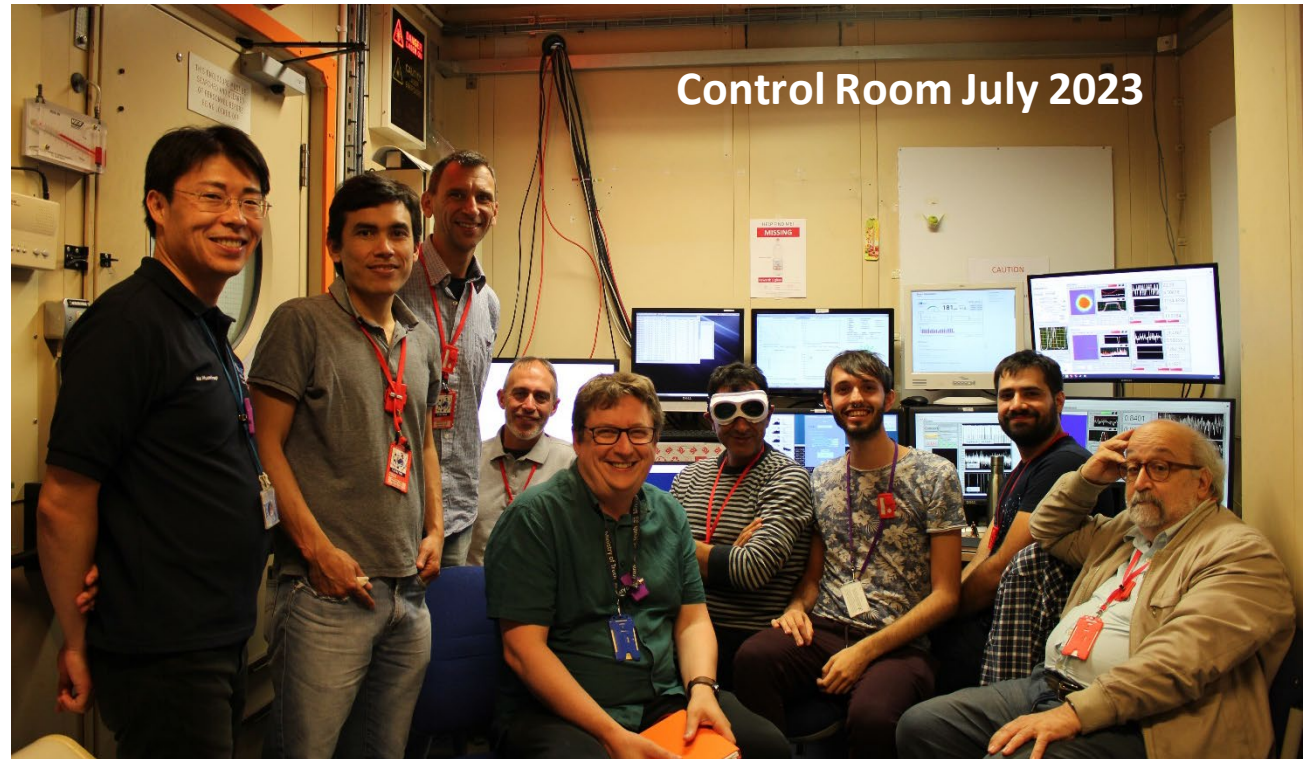
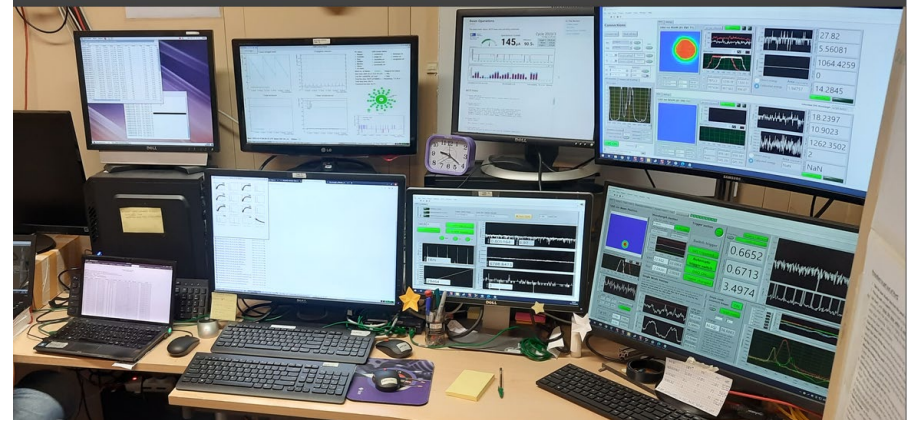
Water absorption measured in 100 cm of air RH 40-45%
for the absolute energy calibration



Data taking in 2023

July 2023: commissioning

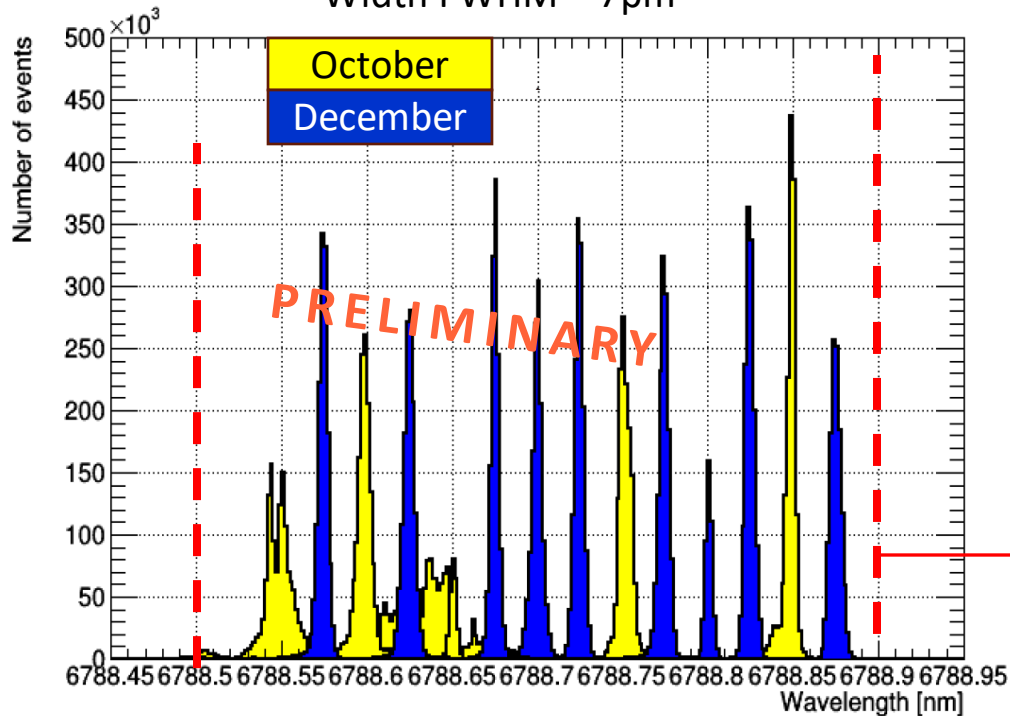
Physics data taking:
October and December 2023
A total of 21 days



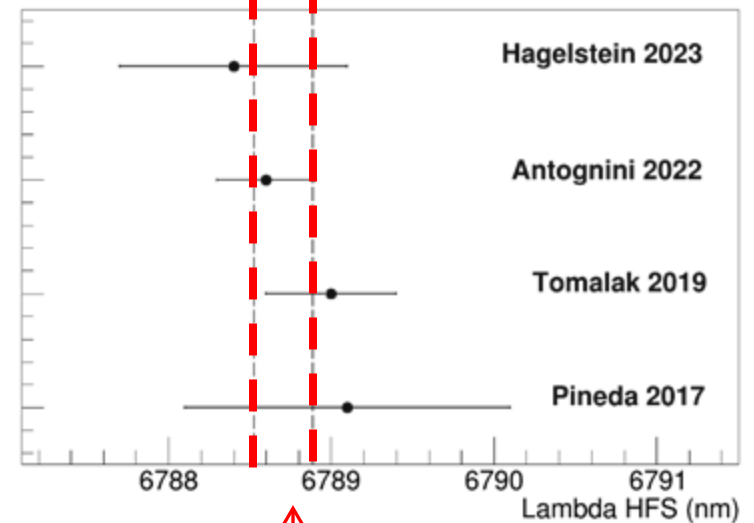
FAMU investigated regions in 2023

Laser wavelengths investigated in 2023:

14 frequencies investigated
in steps of 25 pm
~ 24 h for each frequency
Width FWHM ~7pm

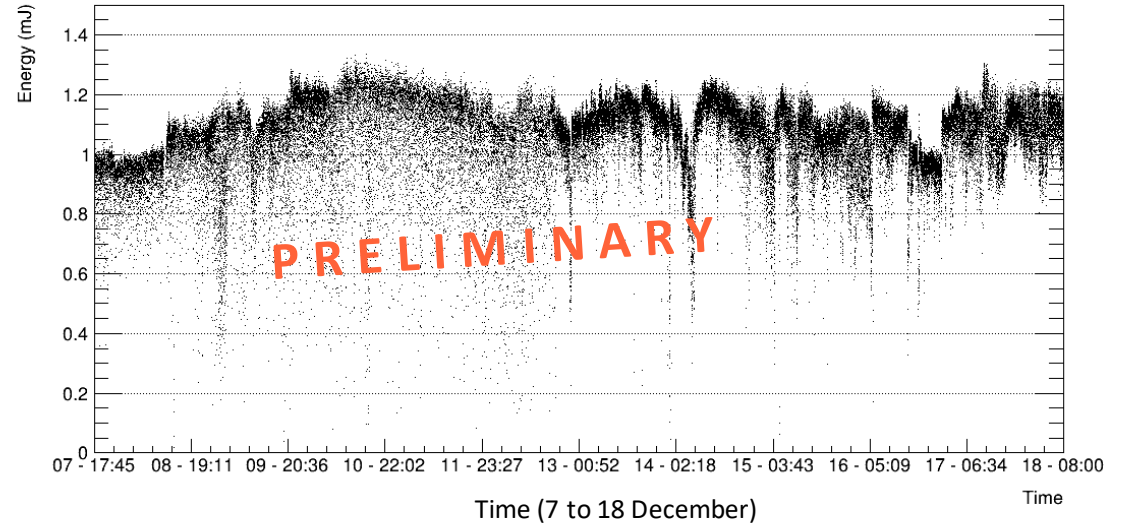
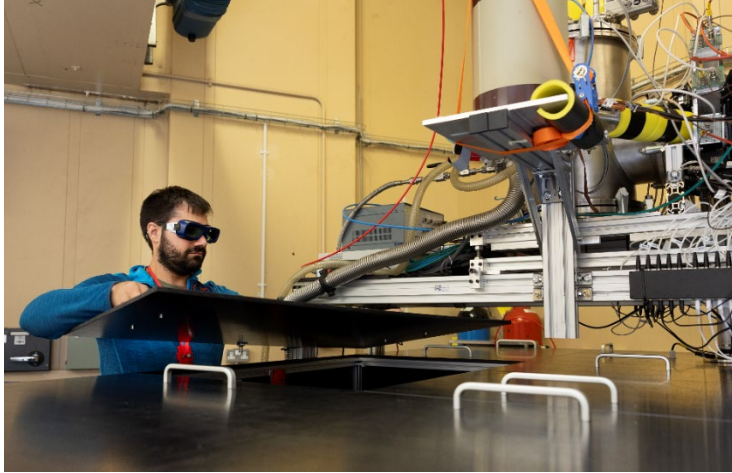


Theoretical prediction of the HFS



Laser operations during beam time

@7um Mean energy 1.05 ± 0.12 mJ



	October	December
Continuous laser operations	9 days	12 days
Number of laser shots	20M	26M
Mean daily uptime (deadtime: WL change, laser checks and alignment)	22h (~90%)	23h (~96%)
Number of measured wavelengths	5	9
Time required to change wavelength	1 – 2.5 hours	30-60 min

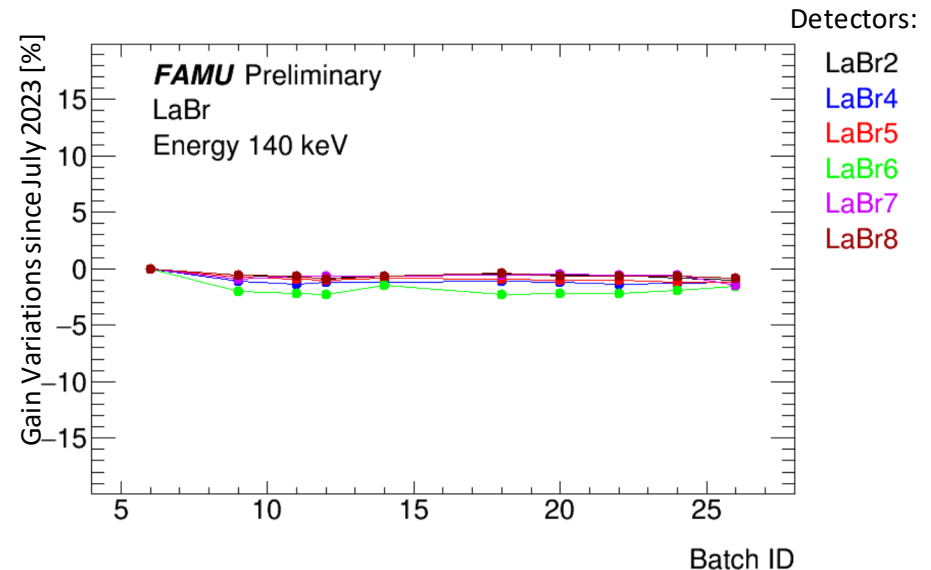
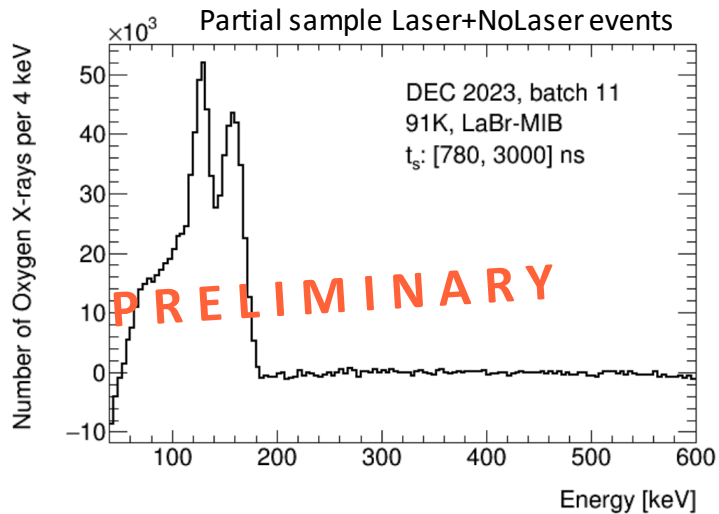


Ongoing data analysis

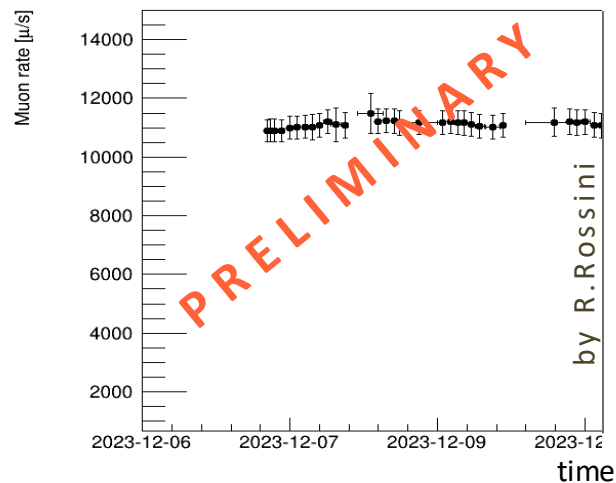
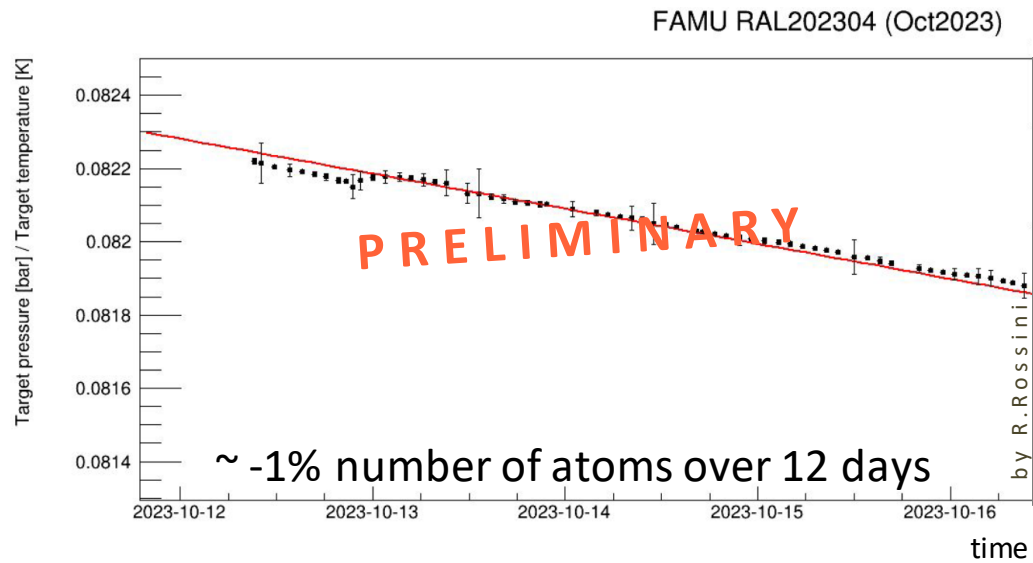
Presently focused on time effects and systematics.

For example:

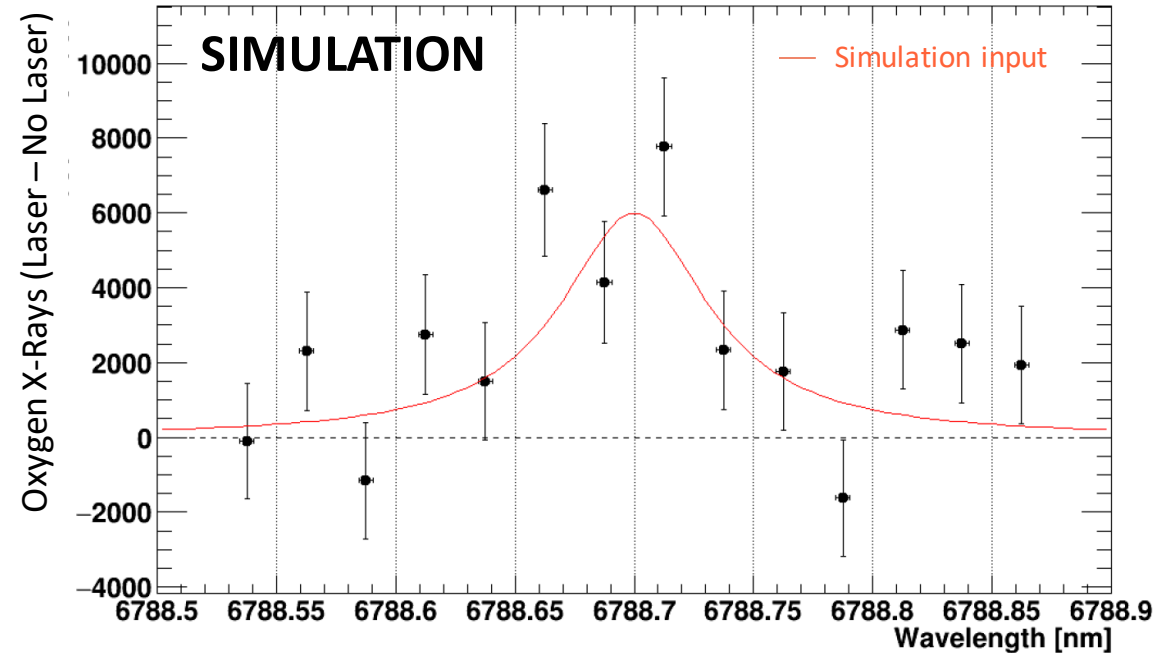
- Detector performances, stability in terms of gain and efficiency
- Monitor of the muon beam flux variation with the hodoscope
- Target condition (T, P) and other time dependencies
- Stability and characteristics of the laser, resolution, etc
- Better understanding of errors with MC



Some studies of stability



Expectations from Montecarlo



Simulation inputs:

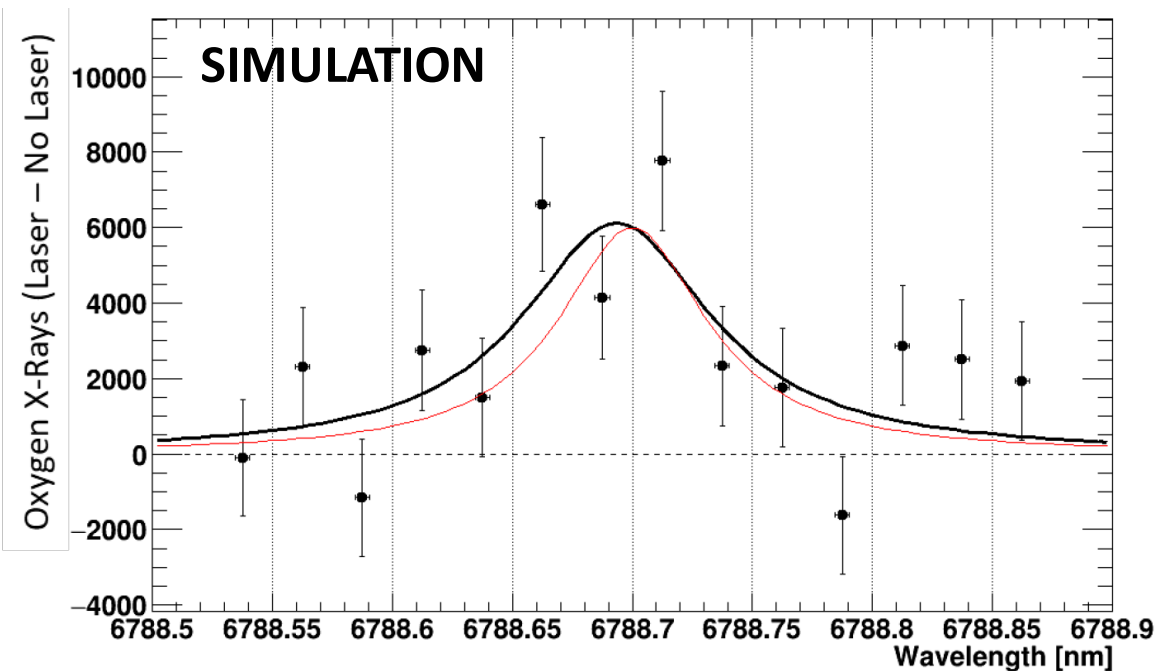
Resonance center 6788.700 nm

Resonance width 75 pm

Effect 3%

Simulation includes our experimental conditions

Expectations from Montecarlo



Simulation inputs:

Resonance center 6788.700nm

Resonance width 75 pm

Effect 3%

Simulation includes our experimental conditions

The fit reconstructs

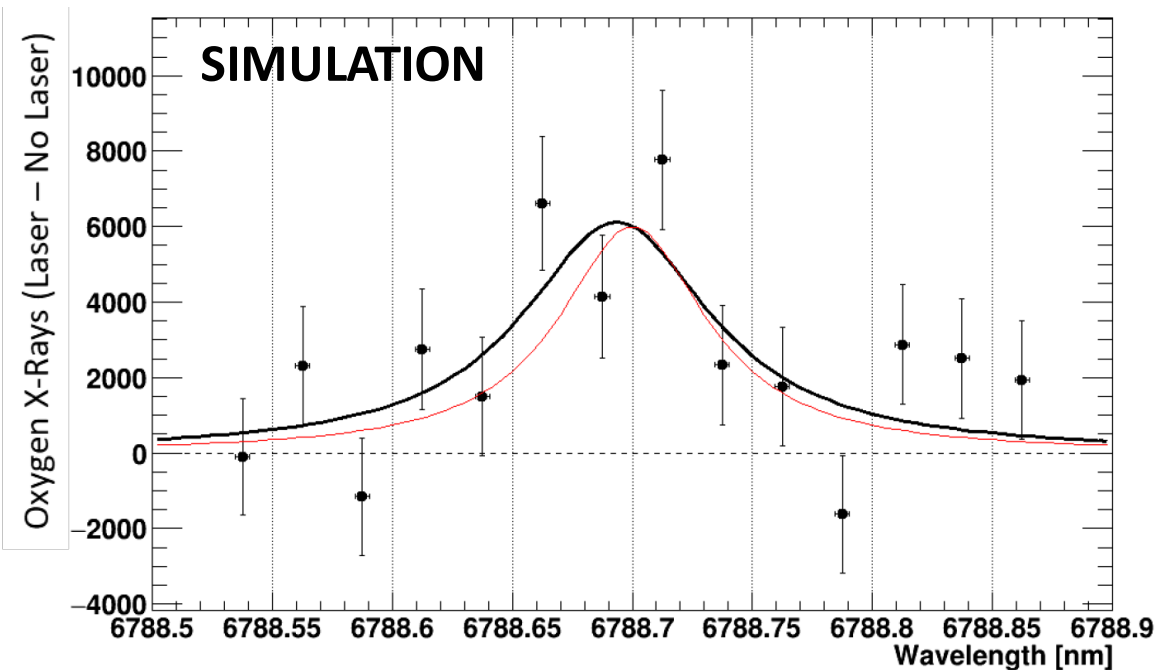
the position of the center with an

uncertainty on the 2nd digit

≡ less than 2 part per million

Significance 4.7 sigma

Expectations from Montecarlo



Simulation inputs:

Resonance center 6788.700nm

Resonance width 75 pm

Effect 3%

Simulation includes our experimental conditions

The fit reconstructs
the position of the center with an
uncertainty on the 2nd digit
≡ less than 2 part per million
Significance 4.7 sigma

What we can improve:

- Reduce the wavelength step between measurements
- Increase the statistics
- Increase the laser energy

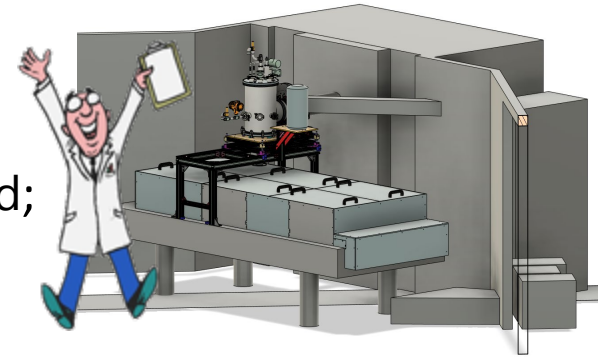
Present laser characteristics are not the main source of uncertainty

Although the target was designed to suppress the signals for materials different by the gas
what affect our precision: X-rays background

Summary and outlook

- Successful physics data taking in 2023 ;
 - target, detectors, cavity, and laser are performing as expected;
 - 14 frequencies investigated
- The analysis of the 2023 dataset is ongoing

- New physics data taking planned in July and till 2025
 - Improvement: larger coverage from X-rays detectors
 - Enlarge the investigated WL region by 2/3



Future:

Working on an improvement of the laser scheme in terms of energy and stability



Thank you for your
attention

