

# Searches for Axions in the LUX-ZEPLIN Dark Matter Detection Experiment

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## Aim

The LUX-ZEPLIN (LZ) experiment is an extremely sensitive dark matter direct detection experiment that uses xenon as the target to detect signals in a liquid-gas dual-phase time projection chamber. Produced scintillation light (S1) in the liquid chamber and ionization light (S2) in the gas chamber could be the hints of dark matter candidates: axions or axion-like particles. In details, the main topics about dark matter candidates in the LZ are:

- Limit solar axions masses less than  $1 \text{ keV}/c^2$ .
- Limit axion-like particles' mass in the range of  $1\text{-}11 \text{ keV}/c^2$ .
- Explore properties of other dark matter candidates: weakly interacting massive particle, neutrinos, and hidden photon.

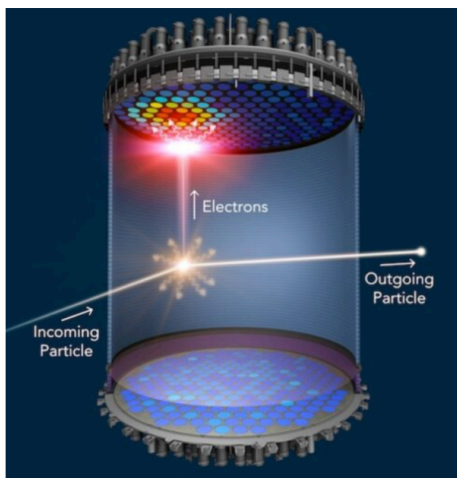


Figure 1. Dark matter detection mechanism in the LZ: in a cylinder of 7 tons of liquid xenon, S1 light is produced by prompt scintillation in liquid xenon, S2 light is generated through electroluminescence when electrons are drifted by a vertical field across the liquid surface into the thin gas phase.

## Reference

- (1) Aalbers, J.; Akerib, D.; Al Musalhi, A.; Alder, F.; Amarasinghe, C.; Ames, A.; Anderson, T.; Angelides, N.; Araújo, H.; Armstrong, J., et al. Search for new physics in low-energy electron recoils from the first LZ exposure, *Physical Review D* **2023**, *108*, 072006.

## Method

### Axion Production Mechanisms

Pseudo-scalar Nambu-Goldstone boson axions have theoretical couplings to leptons, hadrons, and photons. As a result, axion production is possible within star like our sun via nuclear and thermal processes:

Coupling	Formular	Process
Axion-electron	$g_{ae} = \frac{C_e m_e}{f_a}$	Atomic, Bremsstrahlung, Compton (ABC)
Axion-nucleon	$g_{aN} = \frac{C_N m_N}{f_a}$	$^{57}\text{Fe}$ de-excitation
Axion-photon	$g_{a\gamma} = \frac{C_\gamma}{f_a}$	Primakoff effect

Table 1. Axions and axion-like particles' couplings and mechanisms.  $C_i$  is the specific parameter for each coupling.  $m_e$  and  $m_N$  are the mass of electrons and nuclei.  $f_a$  is the relevant decay constant.

In our previous work, only the ABC process was calculated because Primakoff effect is expected to be small in comparison [1]; future work aims to consider both mechanisms.

### Axion Detection Mechanisms

Axion detection mechanisms in LZ now are mostly axio-electric effect [1], allowing atomic (xenon) ionization by absorption of axions. The Inverse Primakoff effect is another possible detection mechanism which will be considered in the future.

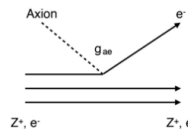


Figure 2. The diagram of axio-electric effect

### The Axion's Coupling Constant and Mass

The strength of interactions between axions and normal matter is given by its coupling constants:  $g_{ae}$  or  $g_{a\gamma}$ . Through fitting axions event rate, axion couplings constant  $g_{ae}$  or  $g_{a\gamma}$  can be derived. After deducing the value of decay constant  $f_a$ , the mass of axions will be given by the equation:

$$m_A \simeq \frac{6 \times 10^6 \text{ GeV}}{f_a} eV/c^2 \quad (1)$$

Axion-like particles have similarly production and detection mechanisms in LZ. The mass of axion-like particles and its couplings constant  $g_{ae}$  can also be determined through fitting the event rate.

## Result

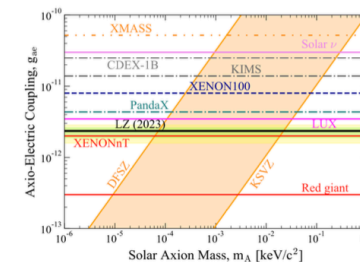


Figure 3. The 90% C.L. upper limit (black line) on solar axion  $g_{ae}$  coupling constant  $g_{ae} = 2.35 \times 10^{-12}$ . The green and yellow bands represent the  $1\sigma$  and  $2\sigma$  sensitivity range, respectively, for repeated background-only experiments [1].

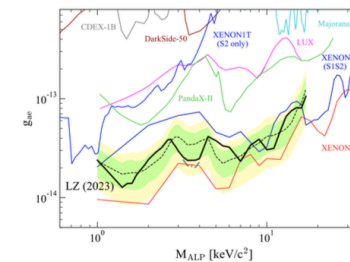


Figure 4. The 90% C.L. upper limit (solid black line) on solar axion-like particles  $g_{ae}$  coupling constant. And the dashed black line is the median sensitivity. The green and yellow bands represent the  $1\sigma$  and  $2\sigma$  sensitivity range, respectively, for repeated background-only experiments [1].

## Discussion

LZ is collecting 1000 days data after a period of calibrations and detector state optimizations. Subsequent science data will benefit from further decay of both  $^{127}\text{Xe}$  and  $^{37}\text{Ar}$ , which will especially improve the solar axion sensitivity [1].

To limit the space of  $g_{ae}$ , in next step, LZ will improve the axion analysis by studying the Primakoff and inverse Primakoff effect in axion production and detection mechanisms.