



ACAT 2024



大阪公立大学  
Osaka Metropolitan University



東京都市大学  
TOKYO CITY UNIVERSITY



# Computing the Wave

Where the Gravitational Wave Community  
benefits from High-Energy Physics, and where it differs ?

22<sup>nd</sup> International Workshop on Advanced Computing and Analysis Techniques in Physics Research  
March 13th, 2024 - Stony Brook University

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Hirotaka Takahashi<sup>2</sup>, Ken-Ichi Oohara<sup>4</sup>, Kazuki Sakai<sup>5</sup>

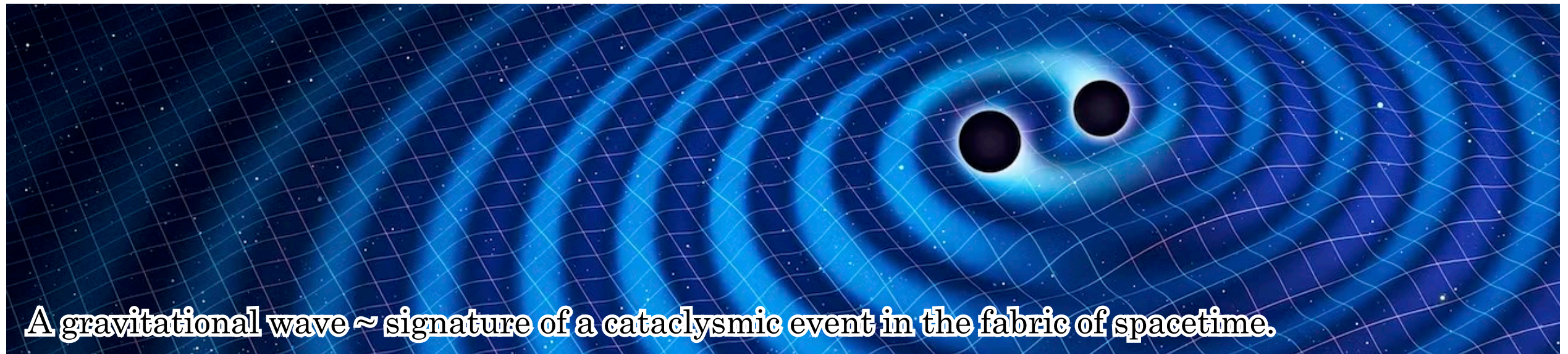
*On Behalf of KAGRA Collaboration*

<sup>1</sup>Osaka Metropolitan University, <sup>2</sup>Tokyo City University,  
<sup>3</sup>University of Illinois at Urbana-Champaign, <sup>4</sup>Niigata University, <sup>5</sup>NIT Nagaoka College





# GRAVITATIONAL WAVES (GW) IN A NUTSHELL



[ e.g. two black holes (BH) or neutron stars (NS) colliding in space ]

## ★ Key milestones in the History of GW Research:

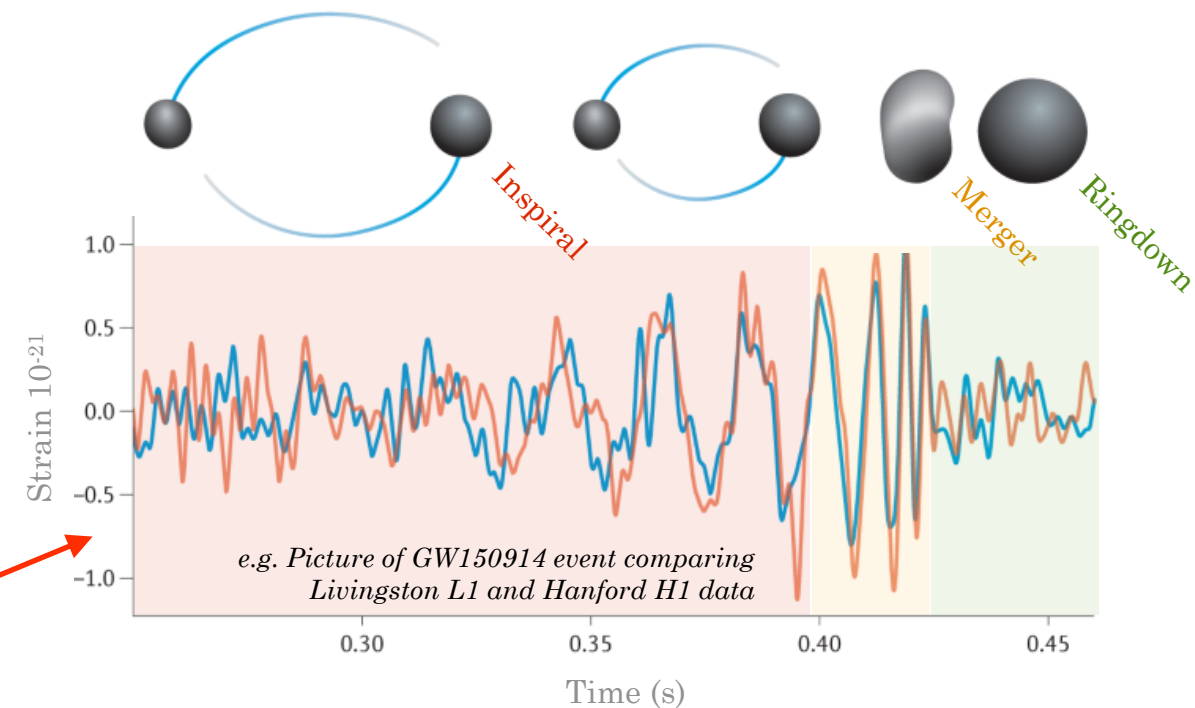
1916 — Prediction of existence by A. Einstein

1974 — First indirect evidence: Hulse-Taylor Pulsar

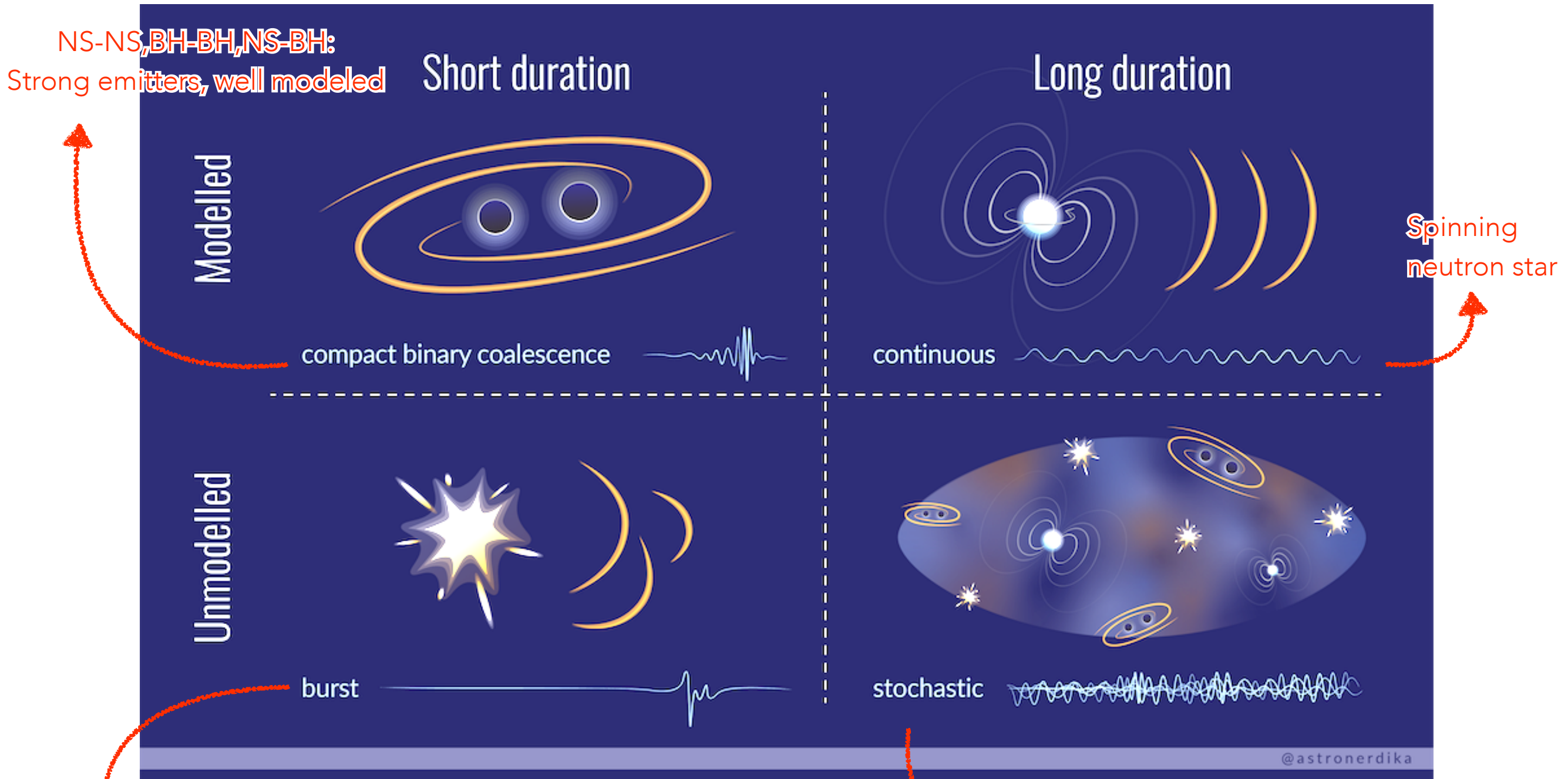
90-00's — LIGO construction (Livingston & Hanford, US)

2015 — First direct BH-BH merger detection: [GW150914](#)

2017 — First NS-NS detection merger with Electromagnetic Counterpart: [GW170817](#) / [GRB170817A](#) (New Era)



# TOPOLOGY OF GRAVITATIONAL WAVES

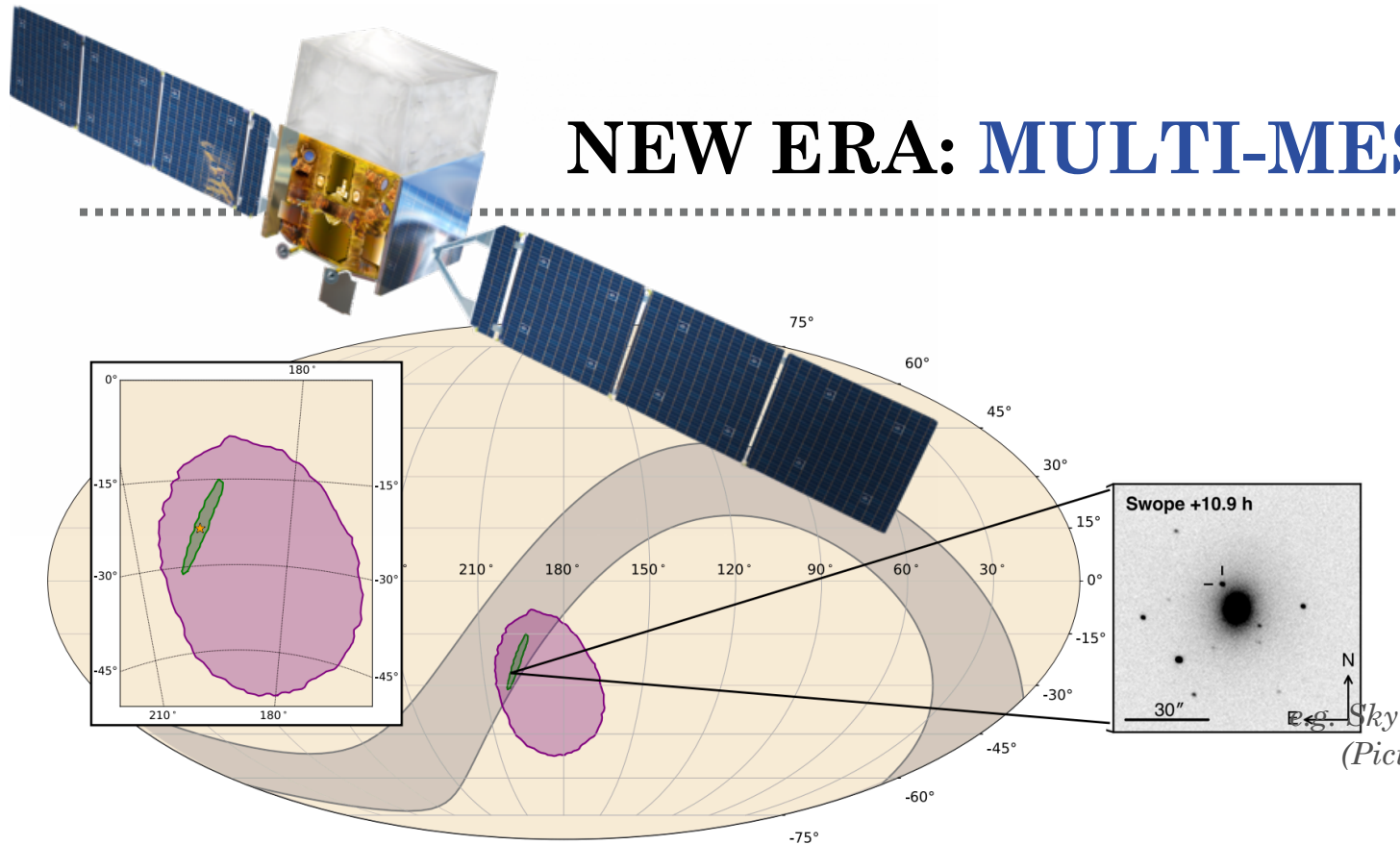


Not well modeled, transients, weak signal, soft gamma-repeater, ..

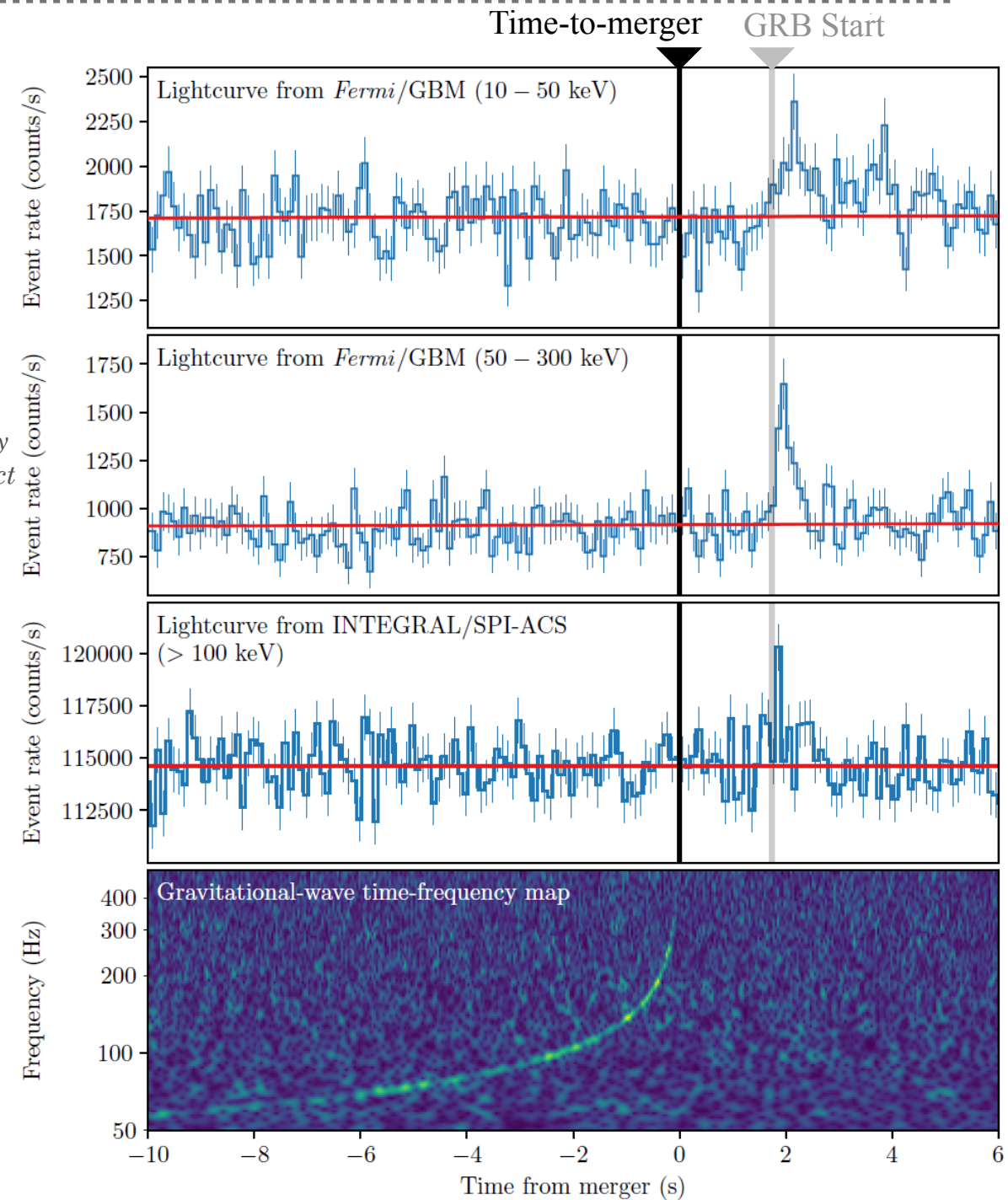
Residual of Big-Bang, cosmic GW background, ..



# NEW ERA: MULTI-MESSENGER ASTRONOMY



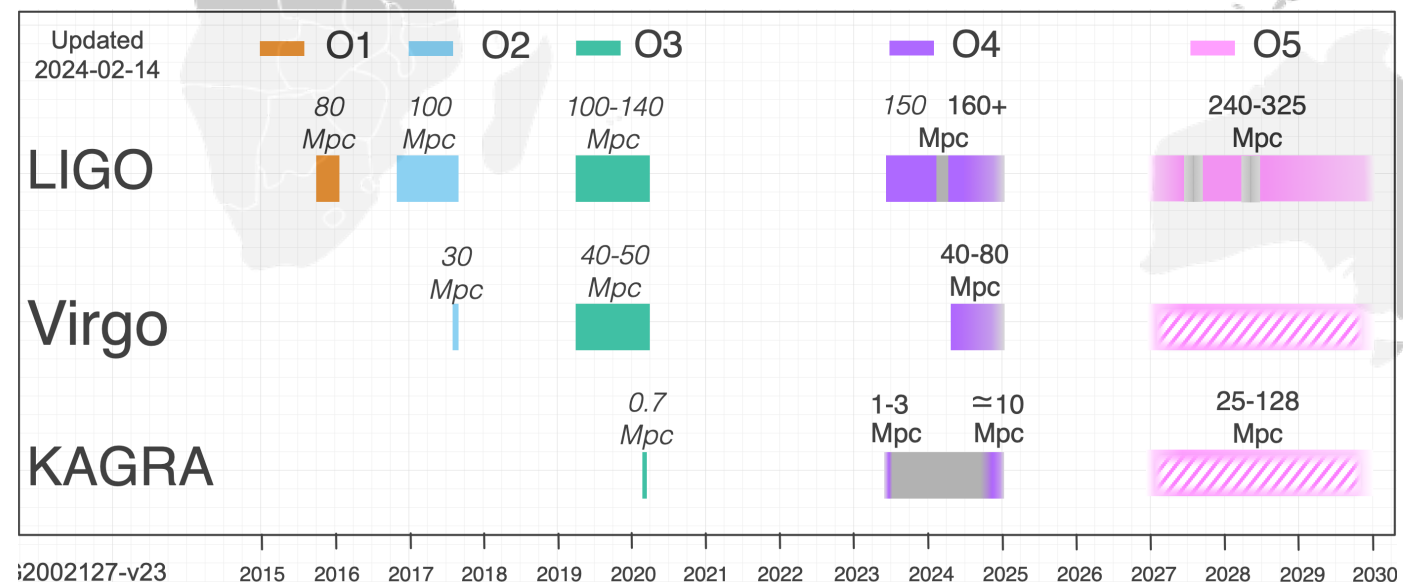
- ▶ **Multi-messenger Astronomy (MMA) Era (started in 2017)**
  - IGWN: Ground-based coordination for low-latency alerts
  - NASA: Coordination between space and astronomers.
  - **This challenge is about speed (< 60s) not size (~1Mb/s)**
- ▶ **Stereoscopic sky localization to be provided**
  - The more data, the better the localization is.
  - Goal: GW signal with negative latency  
(meaning.. detection prior time-to-merger)
- ▶ **Interest: connecting GW and HEP fields.**
  - Gravitational wave observatories
  - Electromagnetic telescopes (*e.g. Fermi*)



First NS-NS Detection with Electromagnetic Counterpart  
(Landmark event: GW170817 / GRB170817A)

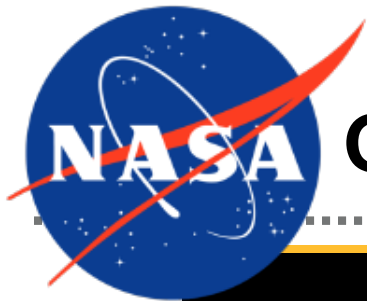


# INTL. GRAVITATIONAL WAVE NETWORK (IGWN)



e.g. Future projects: Einstein telescope, LISA interferometer, etc.





# GENERAL COORDINATES NETWORK (GCN)

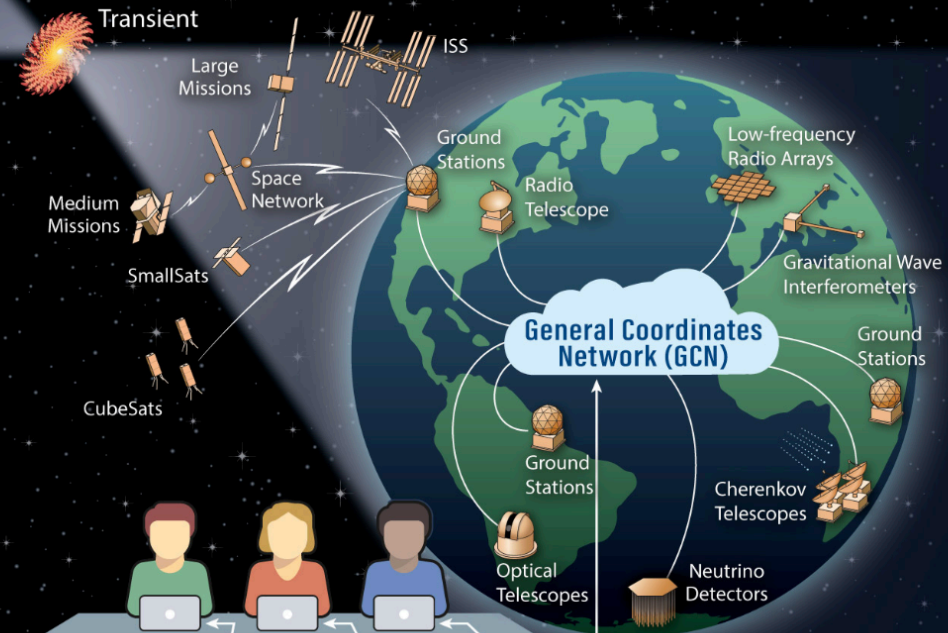


## GCN: NASA's Time-Domain and Multimessenger Alert System

GCN distributes alerts between space- and ground-based observatories, physics experiments, and thousands of astronomers around the world.

[Start streaming GCN Notices](#)

[Post a GCN Circular](#)



LIGO/Virgo/KAGRA Public Alerts .. March 1<sup>st</sup>, 2024

O4 Significant Detection Candidates: **81** (92 Total - 11 Retracted)

O4 Low Significance Detection Candidates: **1610** (Total)

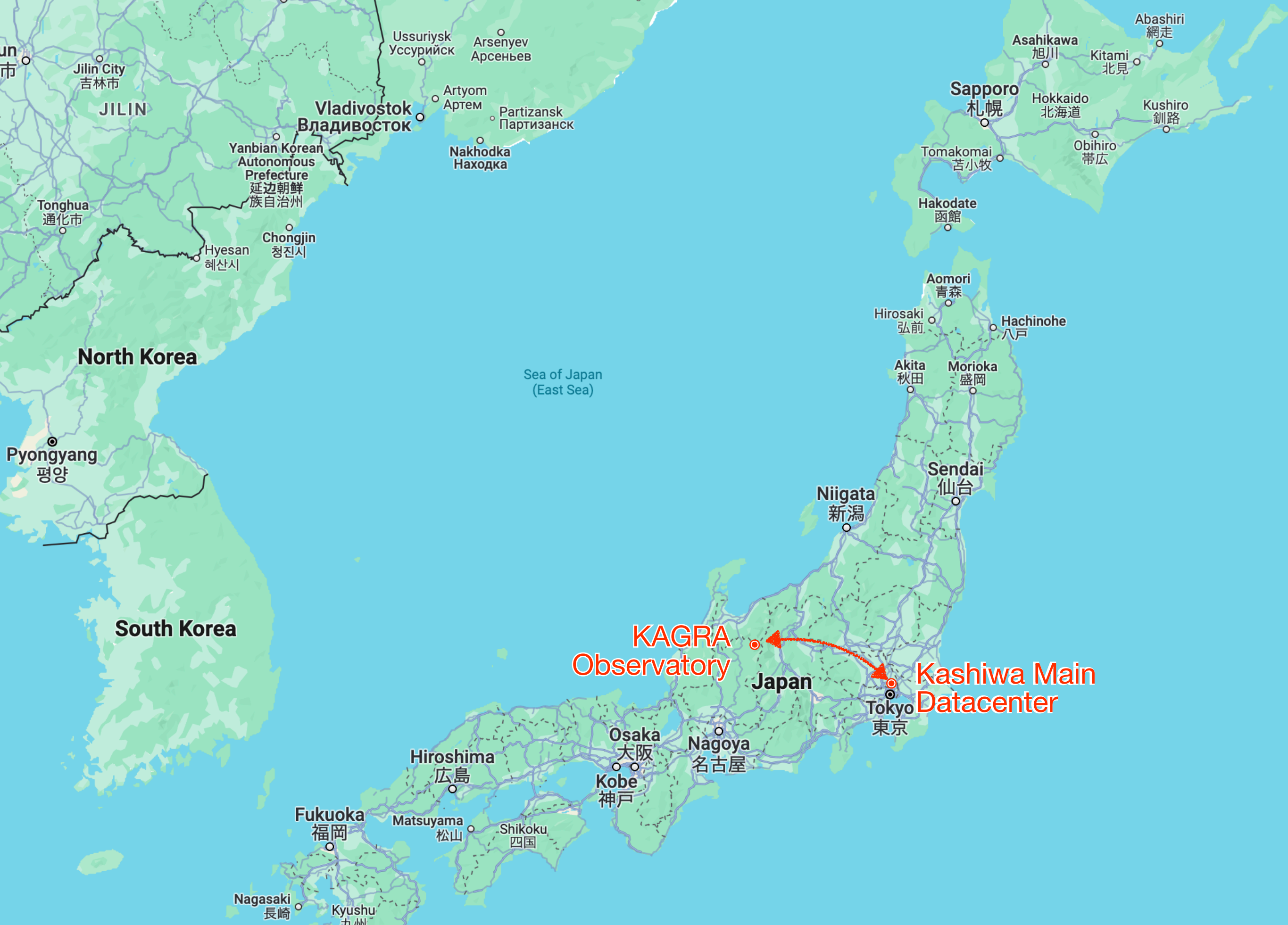
## Gravitational-Wave Candidate Event Database (GraceDB)

Event ID	Possible Source	Significant	UTC	GCN	Location	FAR
S240109a	BBH (99%)	Yes	Jan. 9, 2024 05:04:31 UTC	<a href="#">GCN Circular Query</a> <a href="#">Notices   VOE</a>		1 per 4.3136 years
S240107b	BBH (97%) Terrestrial (3%)	Yes	Jan. 7, 2024 01:32:15 UTC	<a href="#">GCN Circular Query</a> <a href="#">Notices   VOE</a>		1.8411 per year
S240104bl	BBH (>99%)	Yes	Jan. 4, 2024 16:49:32 UTC	<a href="#">GCN Circular Query</a> <a href="#">Notices   VOE</a>		1 per 8.9137e+08 years
S231231ag	BBH (>99%)	Yes	Dec. 31, 2023 15:40:16 UTC	<a href="#">GCN Circular Query</a> <a href="#">Notices   VOE</a>		1 per 3.7932e+06 years

- **GraceDB** — Centralized API for public alerts and sending GCN Circular Query
- **Celery message broker** — Full orchestration performed with GWCelery package
- **Early public warning alerts** — for rapid comms. (<https://gcn.gsfc.nasa.gov/>)

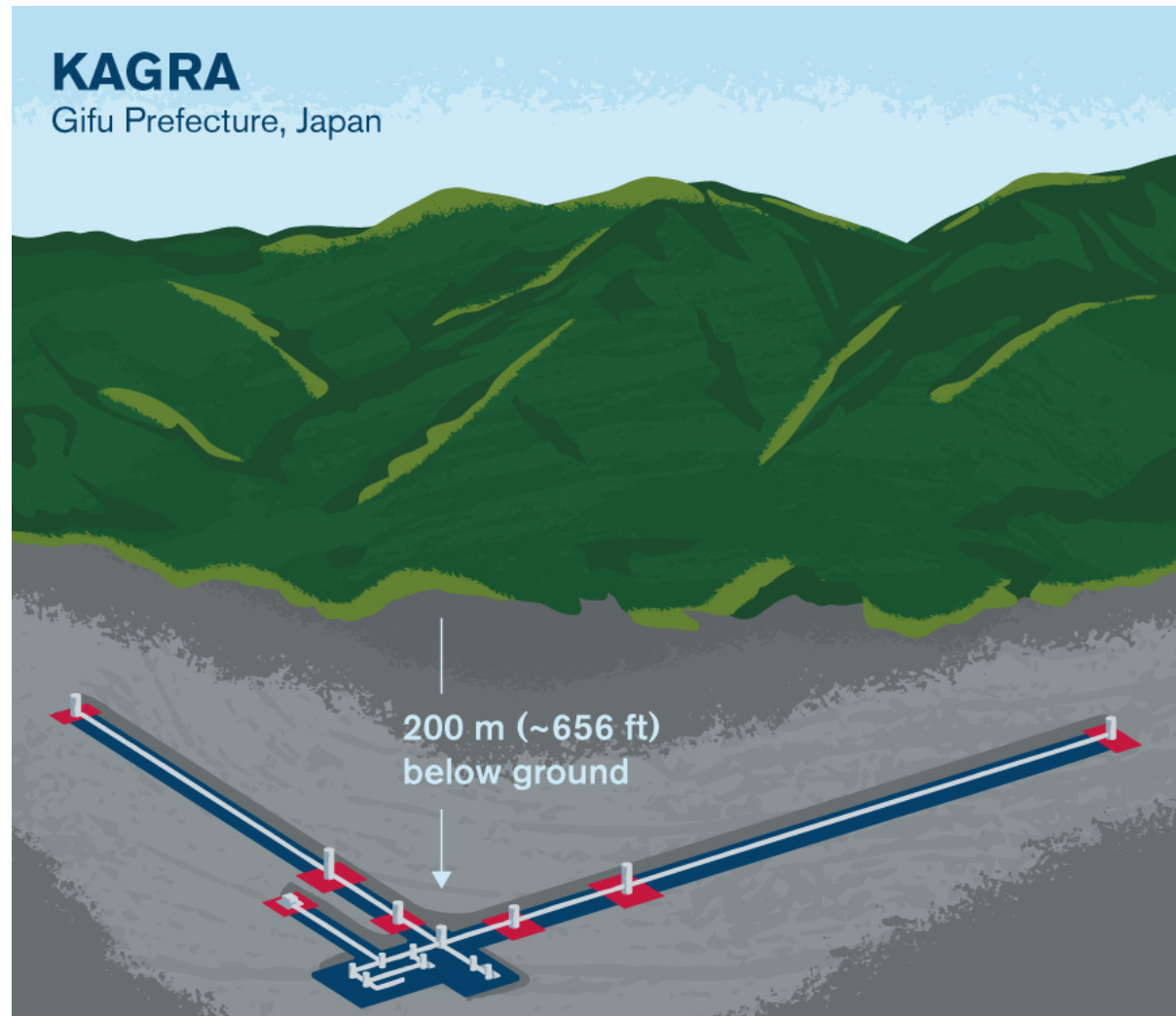
...







# FOCUS ON KAGRA EXPERIMENT

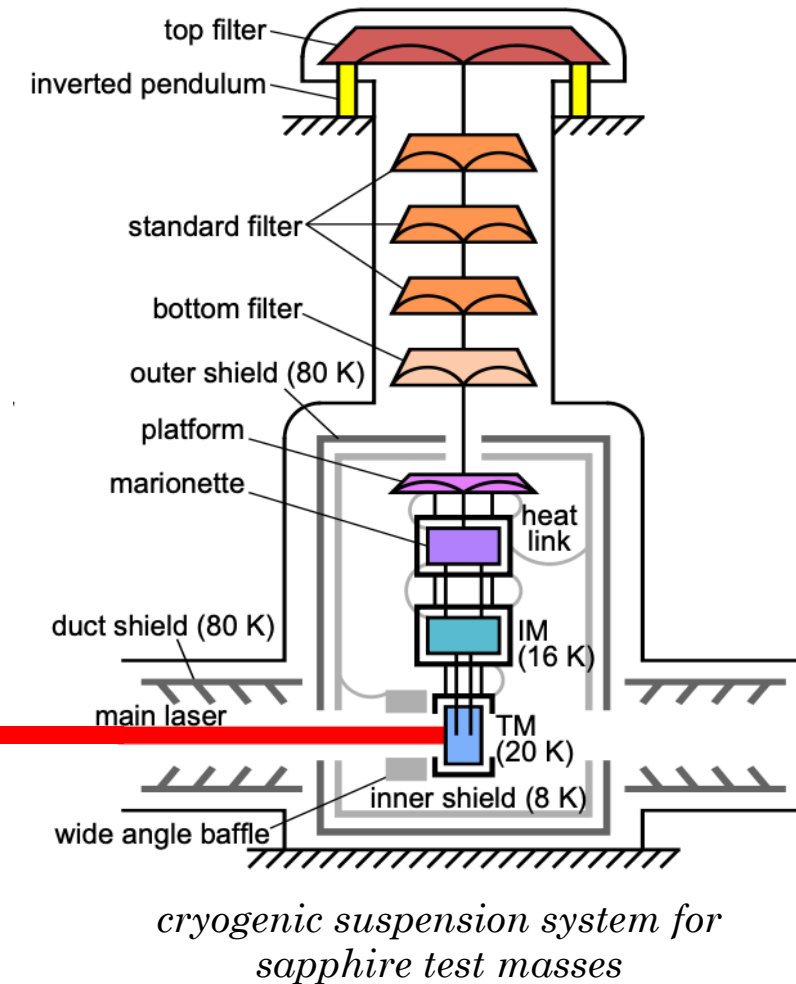
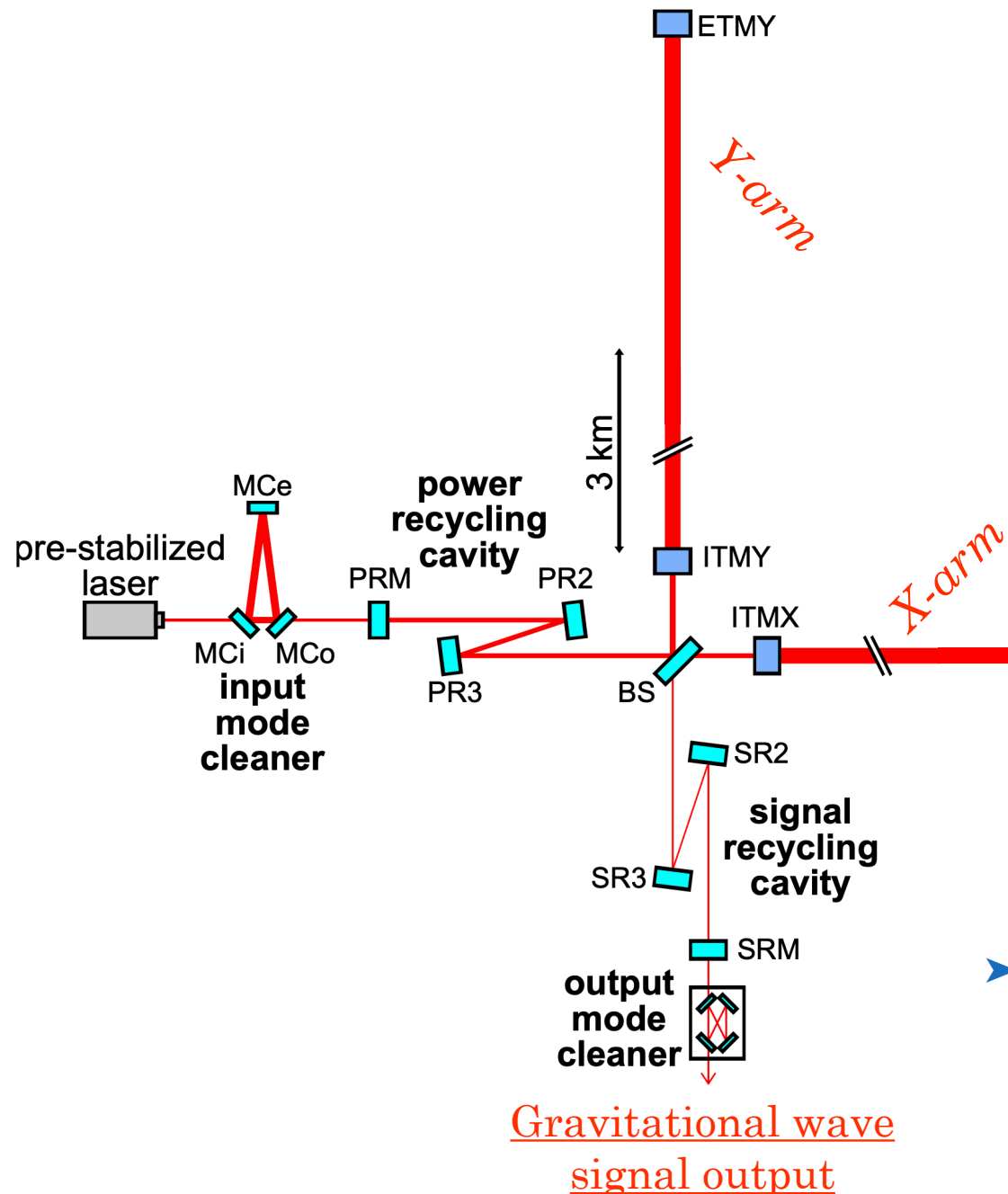


KAGRA Observatory has been established in 2016 in Kamioka, Gifu prefecture  
(*same place as Kamiokande neutrino experiment*)

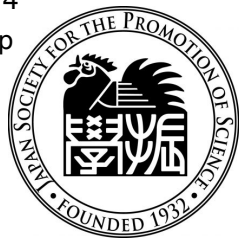
The telescope is a 3 km Fabry-Perot interferometer  
with **20K Cryogenic Sapphire Mirrors** in Mt. Ikenoyama.



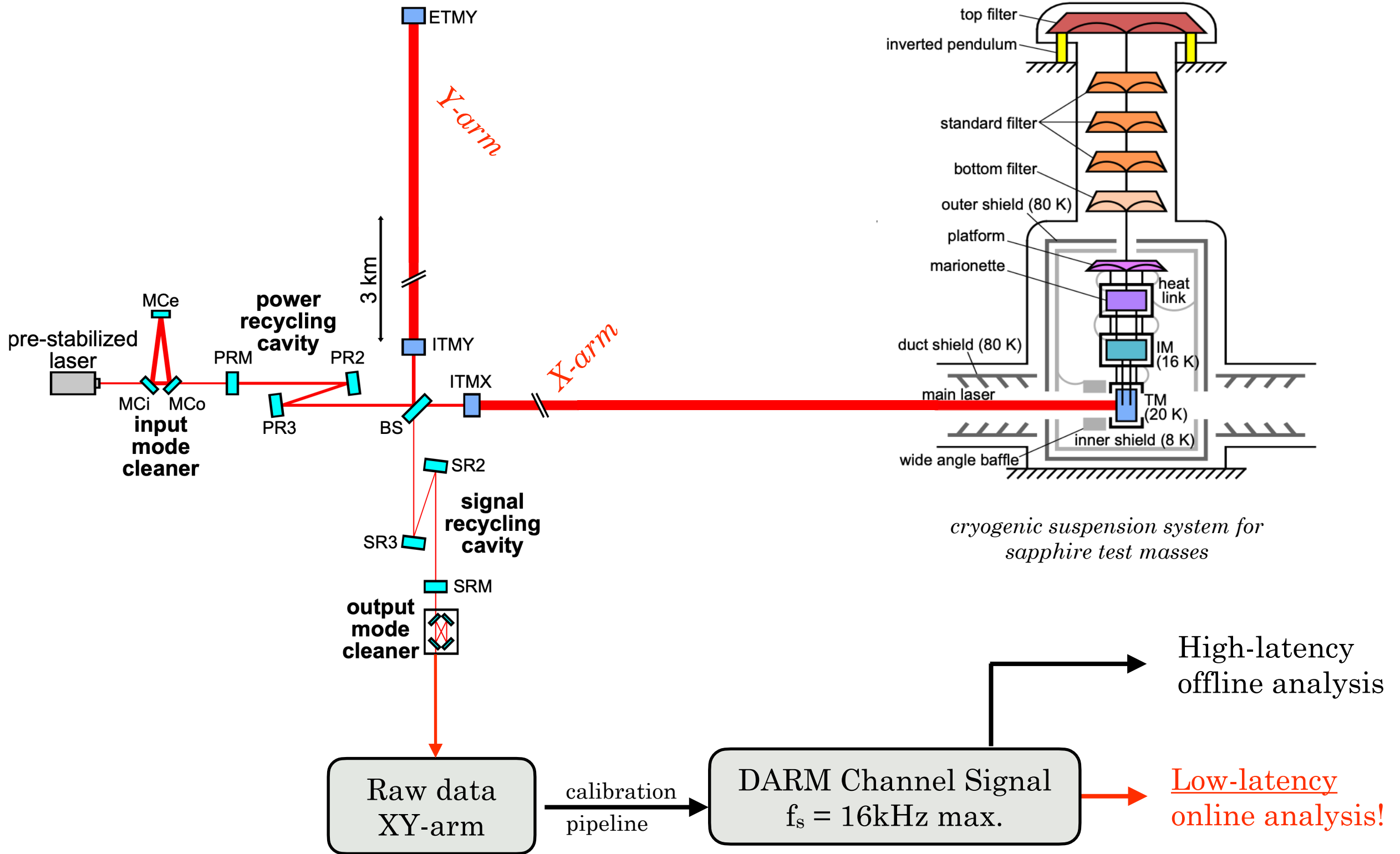
# ON-SITE DATA ACQUISITION (AT KAGRA)



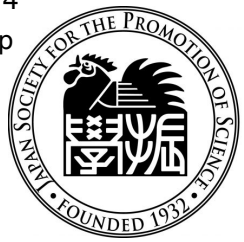
- **Fabry-Perot Interferometer (power recycling)**
  - Laser 180W continuous source at  $\lambda = 1064 \text{ nm}$
  - XY-arms, 20K-cryogenic cooled suspended mirrors
- **On-site infrastructure actions**
  - Data acquisition, pre-processing, environmental monitor.
  - Feedback control to actuators on ETMX, ETMY



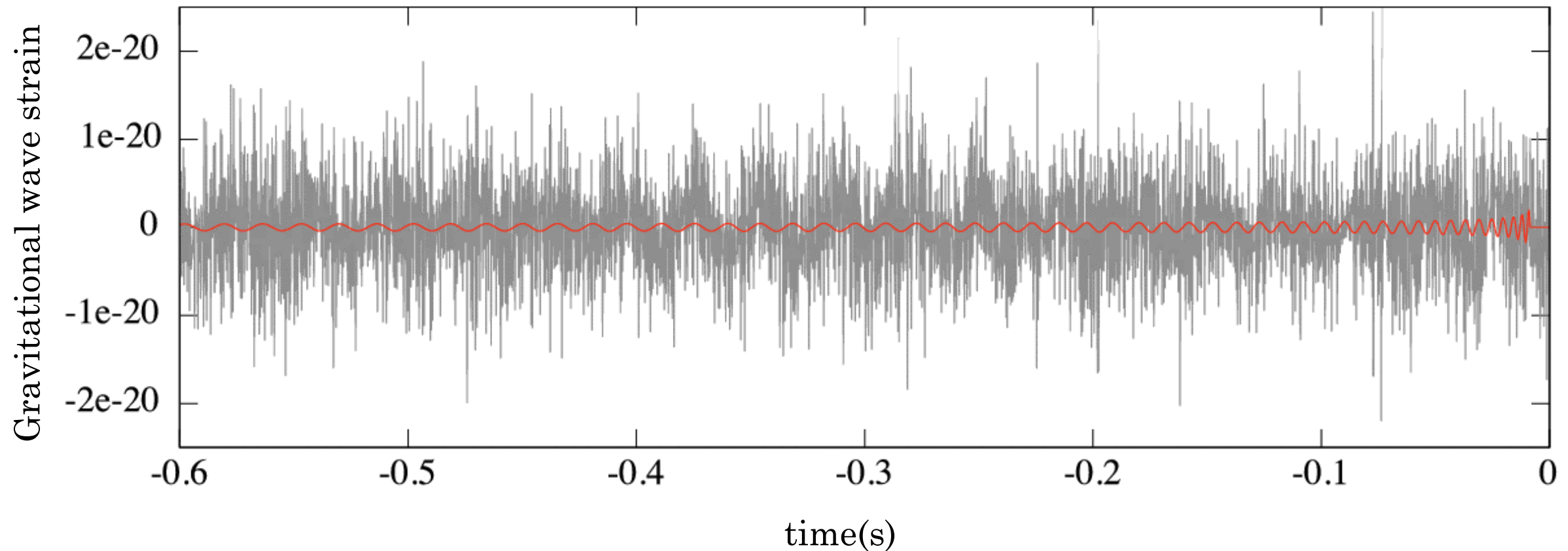
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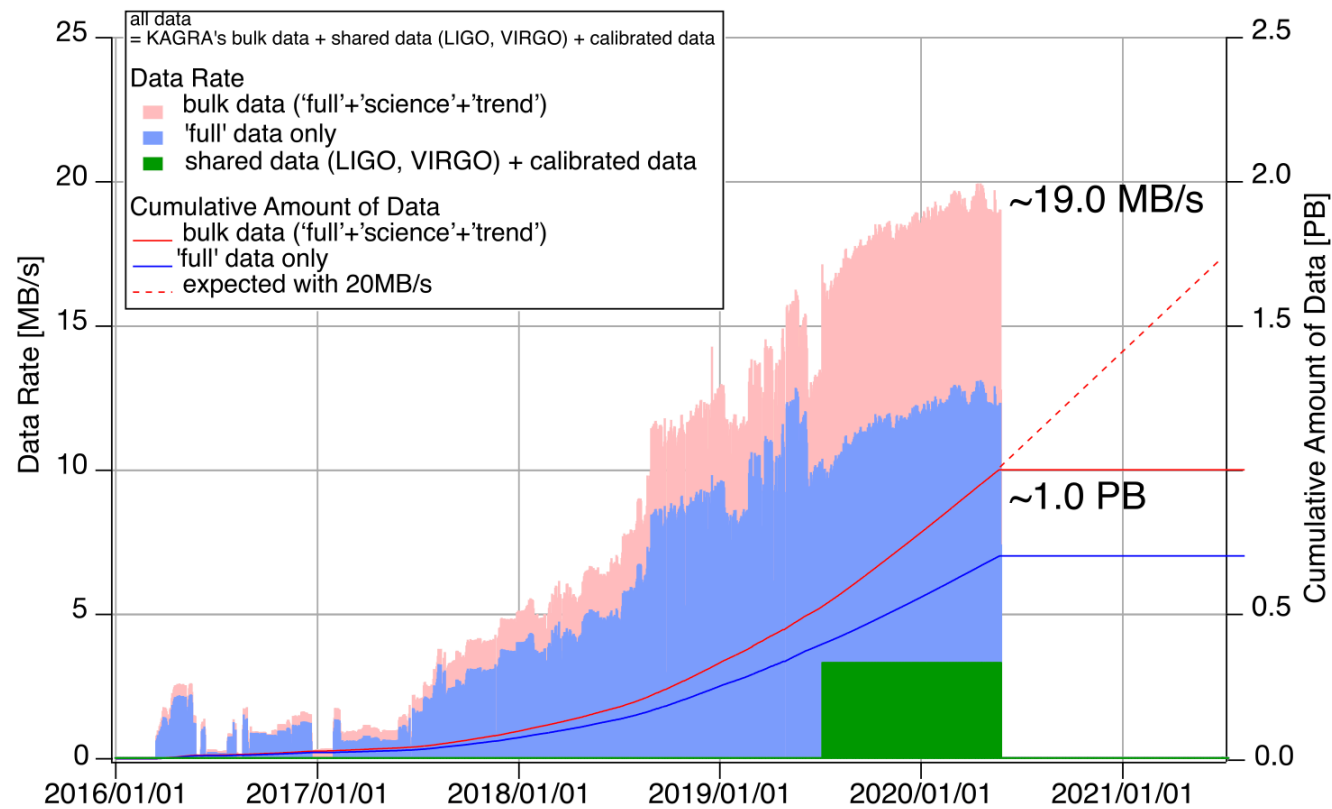
- Gravitation wave signal (here is LIGO experiment); typical signal injection in red
- Data transfer via Apache Kafka (message broker): low-latency reliable sustained data-stream
- Typical latencies at Kashiwa Data center for “1 second” files:

LIGO to Kashiwa	6-8s latency
VIRGO to Kashiwa	10s latency
KAGRA to Kashiwa	2.5s latency

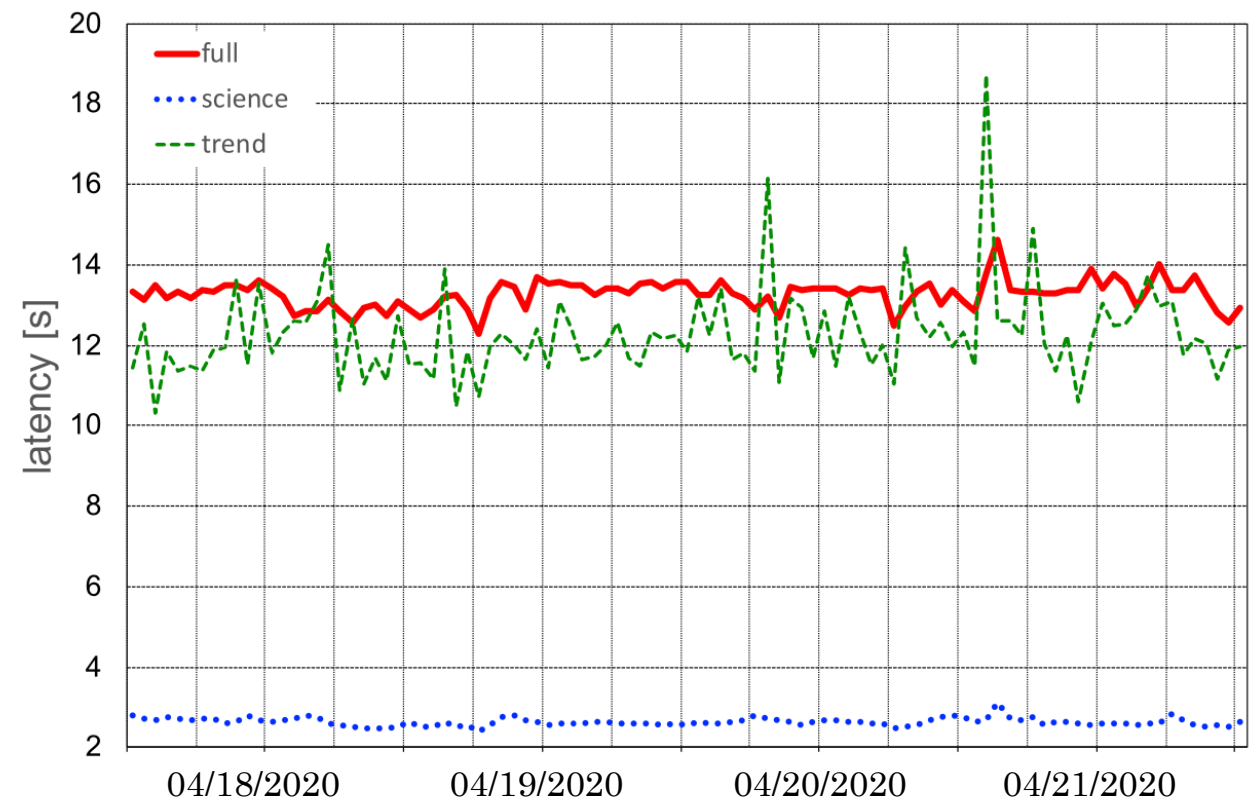


(Latency = transfer from ADC to on-site servers + FIR filtering  
+ transfer from On-Site (L1,H1,V1,K1) to Kashiwa)

# FOCUS ON: KAGRA (KAMIOKA) TO KASHIWA MAIN CENTER (TOKYO)



(Data rate and quantity since beginning of KAGRA and projection for O4)



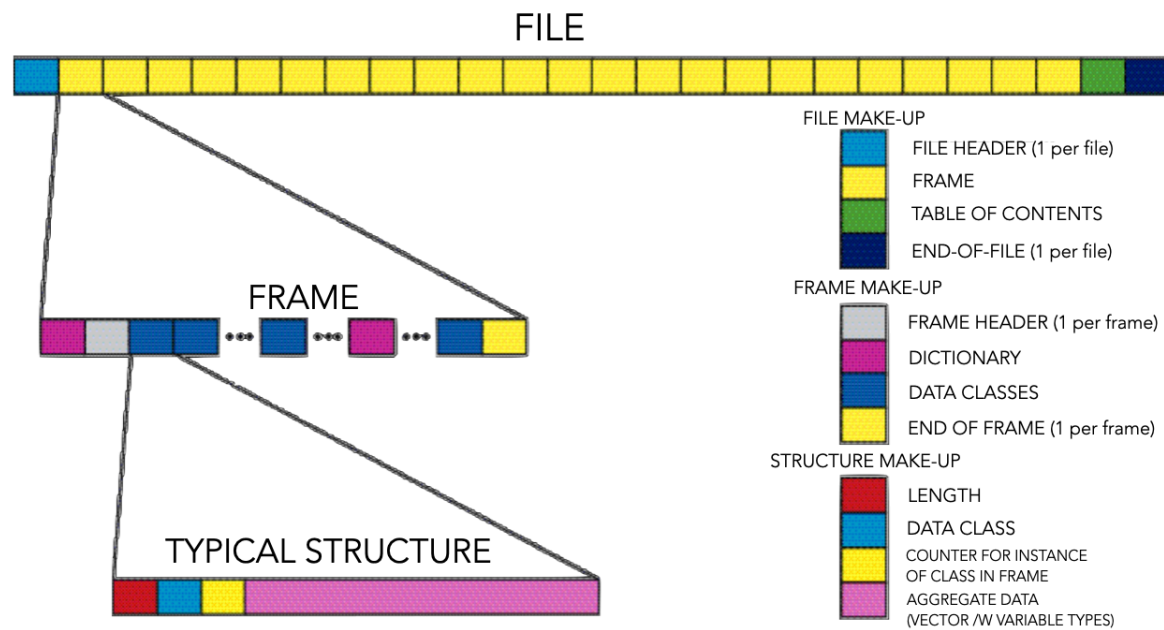
(Latency transfer of larger 32-seconds data files from KAGRA to Kashiwa)

- ▶ **Low-latency Online Analysis O4a ~ O(1min)**
  - Data transfer of 1s-files via “Apache Kafka”
  - Run O4a ~ 1.5PB/y of raw data
  - Research pipeline and rapid sky localization
  - Goal: Alert generation and flagging

- ▶ **High-latency Offline Analysis**
  - Data management /w “Rucio” (GridFTP/CERN)
  - Offline event search, detector characterization
  - Make use of for sharing data:
    - \* CVMFS (CERN) for “data” and “software”
    - \* OSDF Data origin (OSG) for “data”



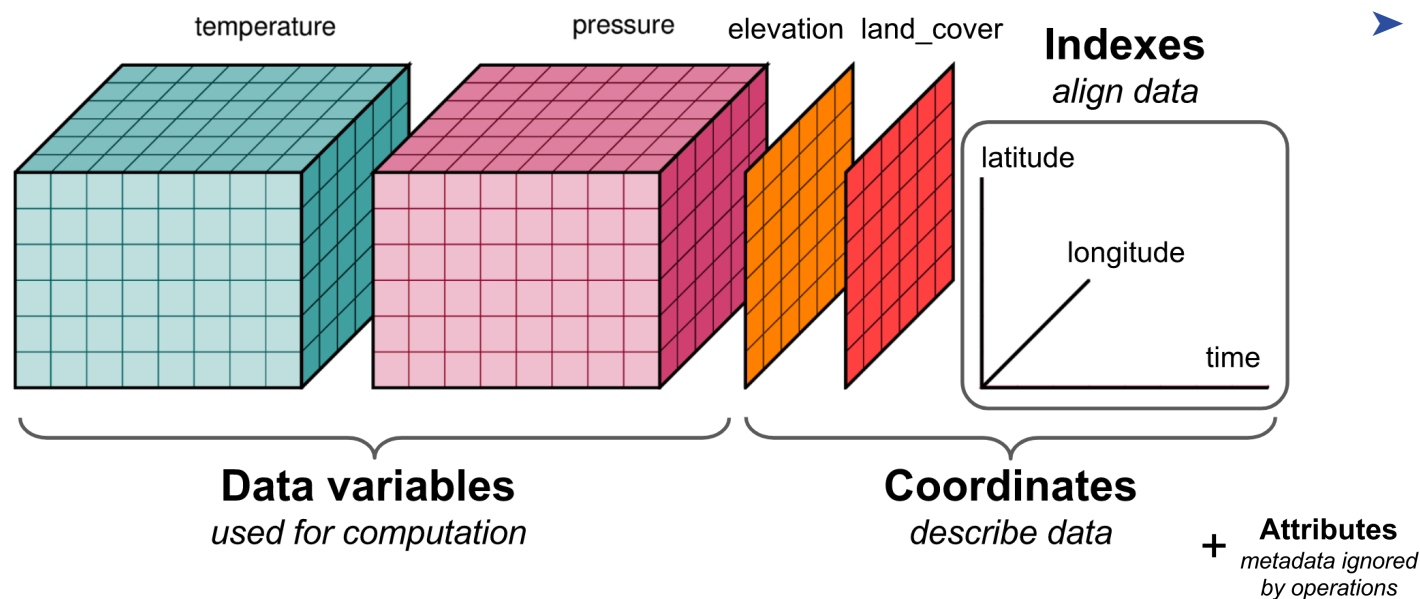
# GRAVITATIONAL WAVE DATA FORMATS



— Picture of the Frame Format data structure for low latency —

- Low-latency data format: `*.gwf`
  - “Frame Format” standard by LVK collaboration
  - Since 2019, strain data  $h(t)$  are 1s frame size.
  - Optimized format to fast reading/writing

Data Format specific to LVK collaboration



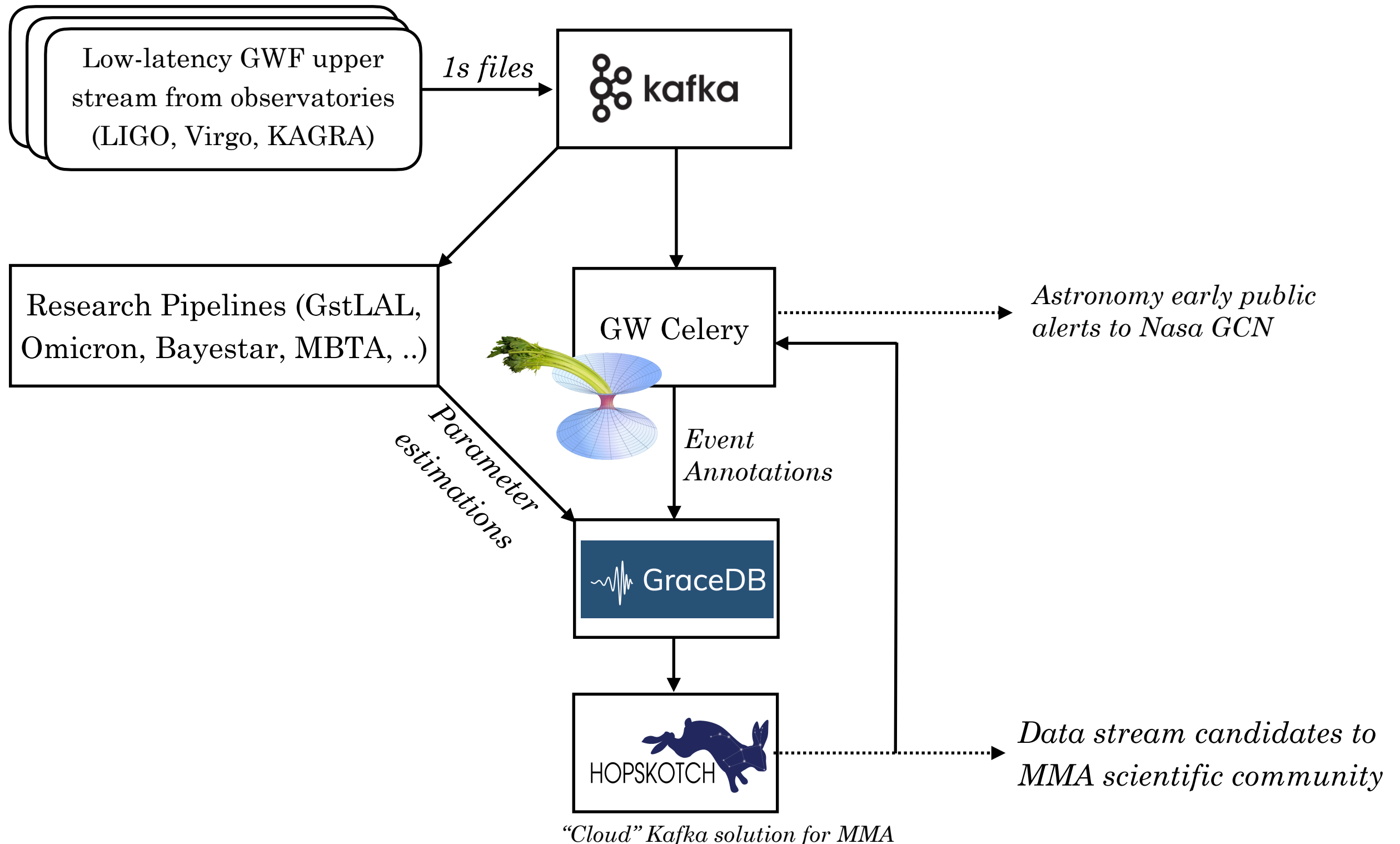
— Diagram of the HDF5 data format for high latency —

- High-latency data formats: `*.gwf` `*.hdf5`
  - Usually HDF5 is used for Open Science Data
  - More flexible when using with Python



(<https://gitlab.cern.ch/escalade/standalone/hdf2root>)

# LOW-LATENCY ORCHESTRATION OVERVIEW







# IGWN O3/O4 COMPUTING RESOURCES

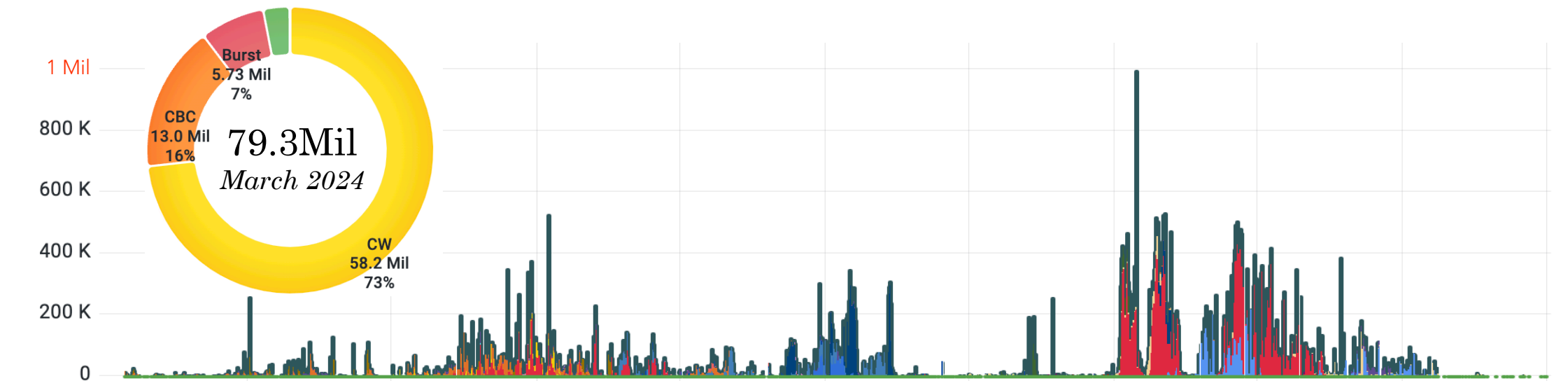
► Observing run periods:

- O3a: April 1, 2019 to October 1, 2019
- O3b: November 1, 2019 to March 27, 2020
- O4a: May 24, 2023 to January 16, 2024
- O4b: April 3, 2024 to the end of year

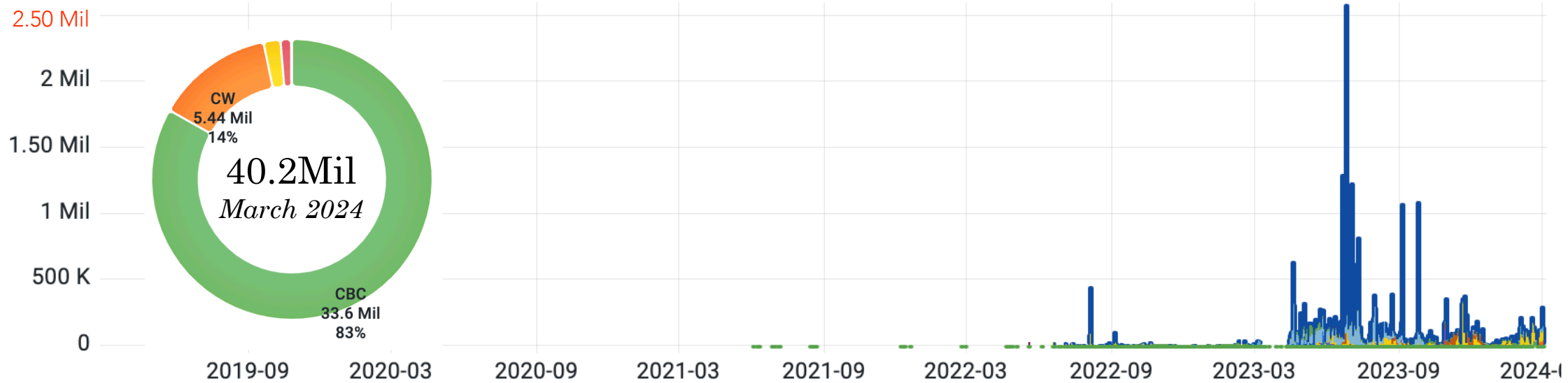
► IGWN Computing Resources

- Computing policy and common coordination
- Code management: Gitlab, Conda & Docker
- Resource monitoring: Grafana
- No common framework for development.

(O3)



(O4)



# COMPUTING PIPELINES AND COMMON FRAMEWORK

- ▶ **IGWN Collaboration Goal: to reduce the amount of tools**
  - Interferometer sensitivity  $\nearrow$  ; Detection range  $\nearrow$  ; Smaller noise  $\searrow$  ; Signal-to-Noise-Ratio (SNR)  $\nearrow$
  - Raw data is not increasing much over years (main computing resources is about template generation)
  - New technics developed in addition to the classic technics. (Matched Filtering, Wavelet, etc.)
- ▶ **Computing pipelines are heterogeneous and complex (no common framework)**
  - Many languages: Matlab, C/C++, Python, ..
  - Main scientific package: LVK Algorithm Library Suite (LALSuite; C/C++)  
*[Noise simulation, template generation (GR modeling for matching data)]*
  - Machine learning increasing usage: First machine learning pipeline M<sub>Ly</sub> is online.



*ROOT has already been used in GW community  
(mainly people in Europe, originating from HEP).*

- ▶ **Requirements to make use of ROOT in GW community**
  - (1) Handling typical I/O files: GWF files and HDF5 files
  - (2) Advanced digital signal processing and (3) complex digital filtering
- ▶ **Personal interests in using CERN modern/high-level tools** (RNTuple, RDataFrame and SOFIE)
  - investigation and development planned for RDataFrame to include TimeSeries capability (1 event = 1s  $\Delta t$ )



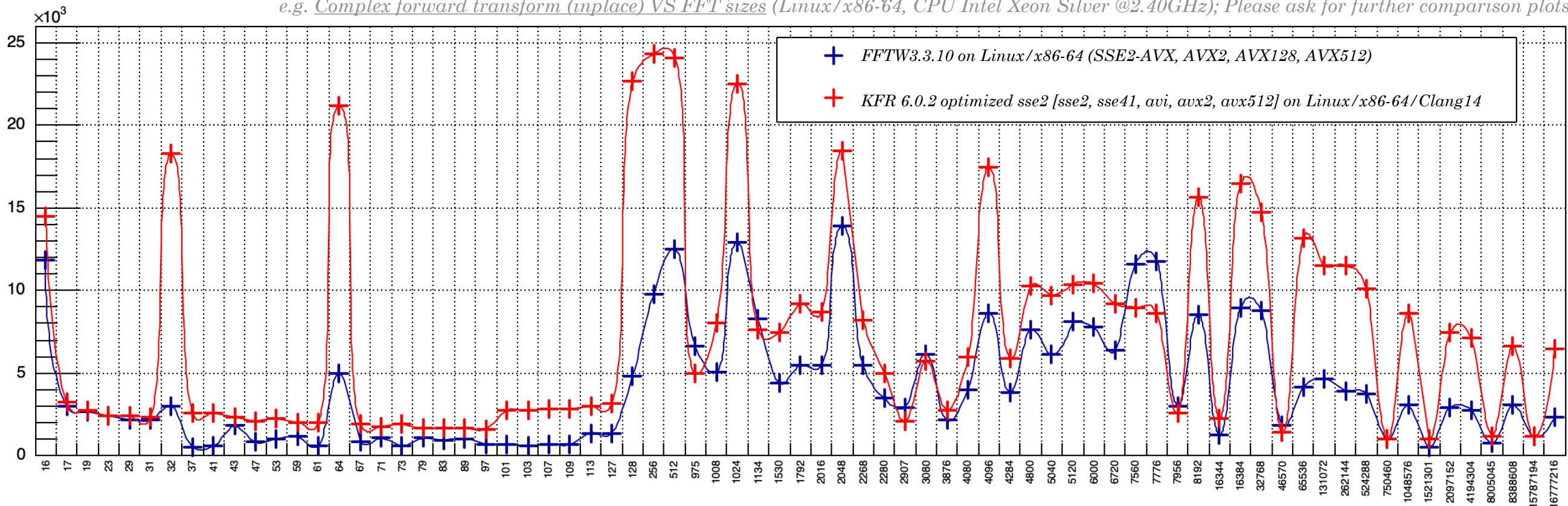
# FAST, MODERN C++ DSP FRAMEWORK

KFR is built for developers who need a powerful toolkit for audio and dsp applications

- KFR library under GPL licensing.
  - DFT & FFT at N-dimensions available
  - Advanced signal processing (IIR/FIR filtering)
- Highly optimized modern C++ library
  - optimal on x86-64 architectures using Clang
  - optimized perf. up to 40% in best cases vs. FFTW3
- Compatible with both ARM64 and x86-64 archs.
  - SIMD SSE, SSE2, AVX instructions, for x86\_64
  - SIMD NEON64 instructions for ARM64.
- Benchmark tests performed with last version.
  - Comparison between KFR6.0.2 and FFTW3.3.10
  - <https://git.ligo.org/kagra/containers/fft-benchmark>

e.g. *Complex forward transform (inplace) VS FFT sizes (Linux/x86-64, CPU Intel Xeon Silver @2.40GHz); Please ask for further comparison plots*

MFlops (The Higher, The Better!)



— One of the benchmark test result (Mflops) comparing KFR 6.0.2 with FFTW3.3.10 / **The Higher the Better!** —

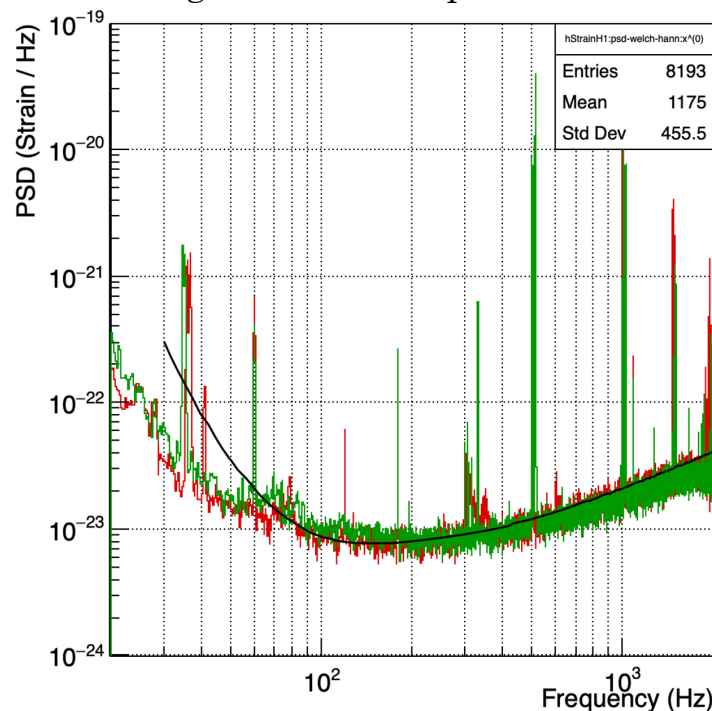
FFT Size



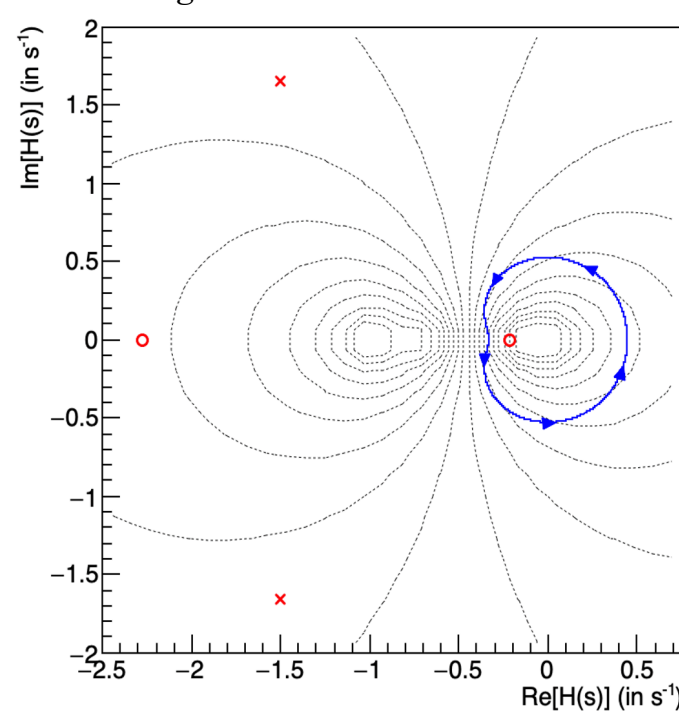
# ROOT SIGNAL LIBRARY PROTOTYPE

- **Advanced DFT features using KFR library.**
  - ASD/PSD calculation, SNR calculation, Matched filtering, Convolution, Window, resampling, etc..
- **Analog and digital filter  $H = B/A$  implementations. (ZPK, BA, SOS, SS representations)**
  - Finite/Infinite Impulse Response filtering
  - Chebychev, Butterworth, Elliptical filter implemented, [..]
- **TKFR Prototype class implemented with ROOT / Make use of the TVirtualFFT interface too**
  - ✓ **ASD, PSD, Hilbert transform, Elliptical filters, Wiener & Kalman filter**
  - FIR/IRR, Windowing, Decimation, Waveform generation (sine, impulse, steps, ramps, etc.), etc.**

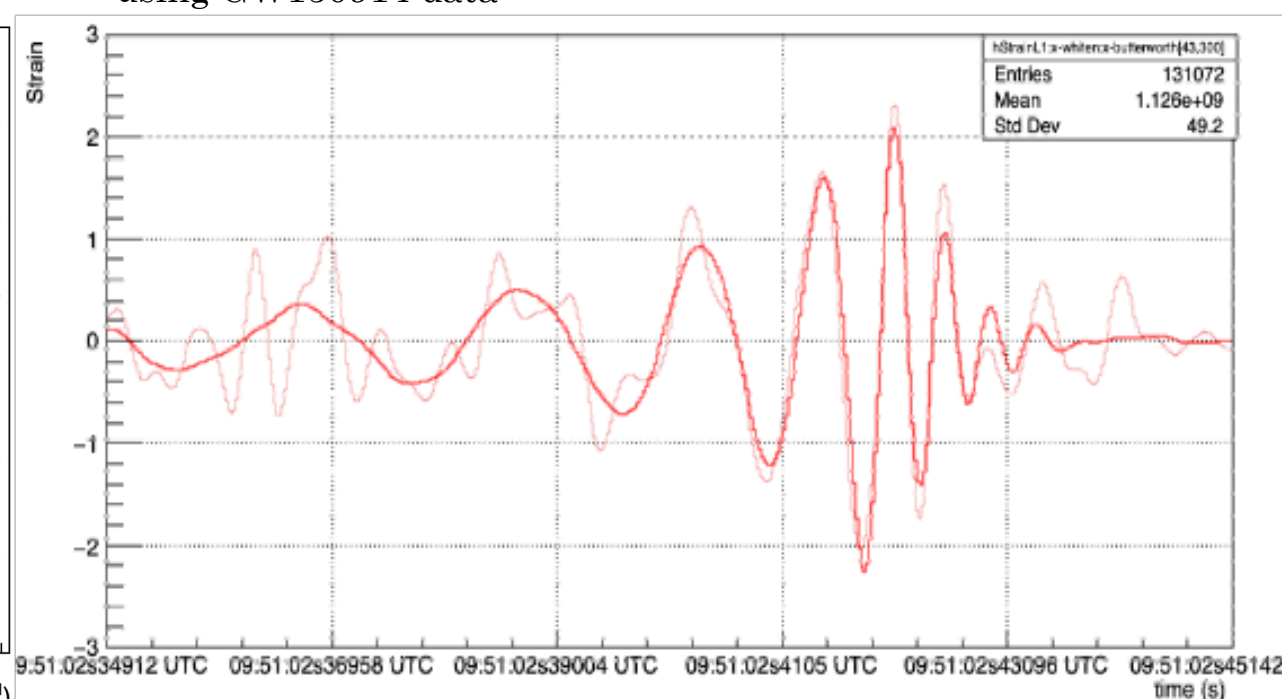
Power Spectrum Density calculation using GW150914 OpenScience data



Example of analog filter  $H(s)$  using KFR



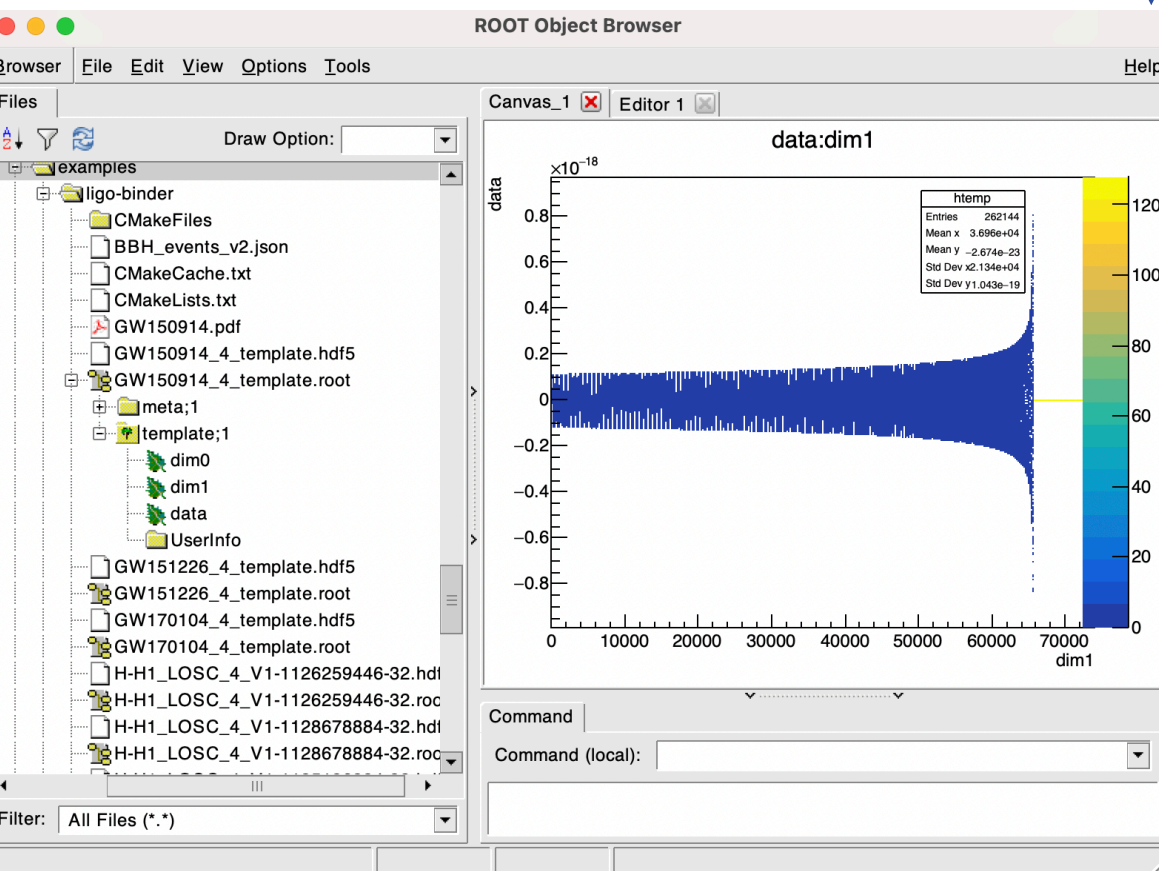
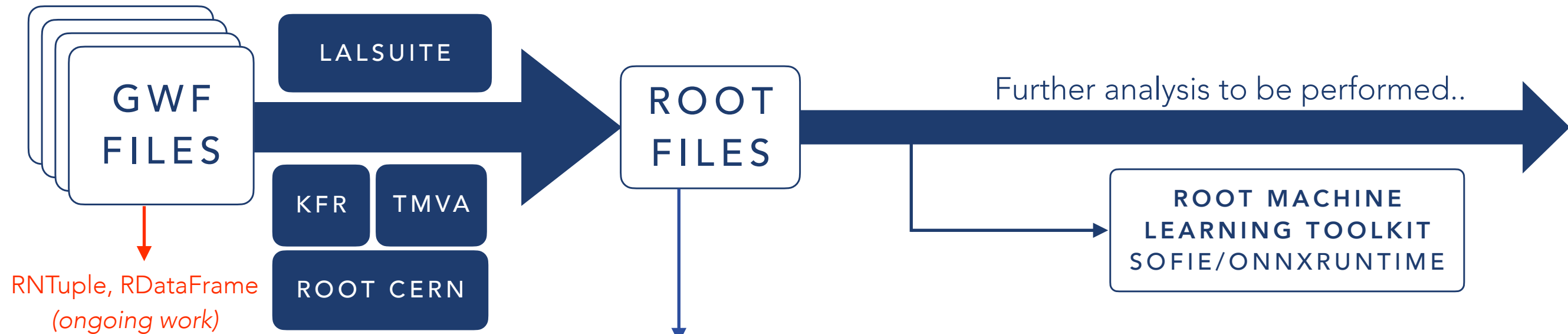
Matched filtering technics with “best template” using GW150914 data







# SUMMARY: PROPOSED GW WORKFLOW



➤ ROOT is very much compatible with GW analysis:

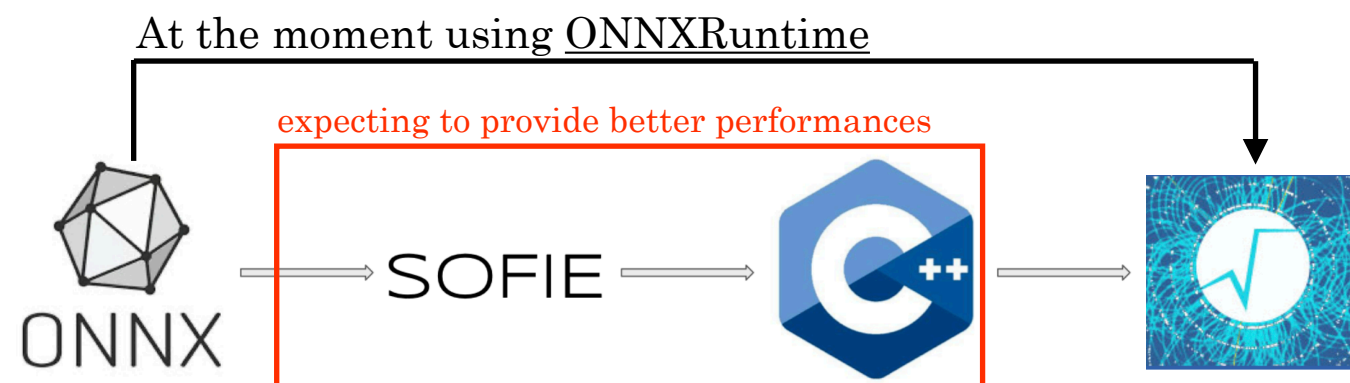
- ✓ KFR Library, TTimeSeries/TFrequencySeries prototype
- ✓ TKafka consumer/broker client prototype
- ✓ TConfigParser (YAML,JSON,TOML,XML,.. run-time files) prototype
- ✓ TCurlFile / TArchive prototype (download/cache files and archive handling)
- ✓ THDF5 parser prototype

➤ LALSuite Package linkable for direct use.

- ✓ TStrain, TStrainNoise, TStrainTemplate  
(for direct strain, noise and template generation)
- ✓ TLigoFormat for configuration file parsing
- ✓ TFrameChain GWF proprietary format. (MT)

## ULTIMATE GOAL: MACHINE LEARNING INFERENCE WITH ROOT

- ▶ **Starformer project**, a transformer-based model for early detection of CBC mergers
  - Collaboration with ISLab (*experts in transformer architecture development, Osaka University*)
  - *Explainable AI is of highest importance, flexibility for production models, parallel performances.*
- ▶ **Machine learning universal format for models**: “Open Neural Network Exchange” (ONNX) files.  
(Training: pyTorch / Inference: ROOT)
- ▶ **Inference using ROOT + ONNX Runtime**:
  - ONNX Runtime: developed by Microsoft and Facebook. (Please visit <https://huggingface.co/> too!)  
Interoperability in many languages. (Python, C++, ..)
  - Performance enhanced expected using SOFIE library in the future (maybe)





# CONCLUSION

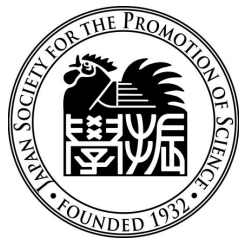
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- **Gravitational wave community is building a new computing paradigm:**
  - Computing power requirement and increasing detector sensitivity.
  - Ongoing work (since many years), now resulting in the IGWN computing collaboration.
- **Computing power doesn't grow that much in GW compared to LHC experiments**
  - Still very challenging to communicate quickly with MMA institutes and astronomers.
  - Kafka integration prototype in an external ROOT library.
- **Goal: provide mature tools as a bedrock for analysis and pipeline development.**
  - Common framework, performant tools + keep flexibility towards new IT technologies
  - Machine learning models is (since many years) intensively used
- **Use of ROOT as main data framework in the future**
  - provide tools to run optimized DSP analysis
  - communicate with cloud computing services.
- **Development of Starformer project** (transformer-based ML model)
  - collaboration with ISLab (Osaka University) for innovative architecture developments.

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# BACKUP SLIDES





# HIGH-LATENCY ORCHESTRATION OVERVIEW

