



ACAT 2024



大阪公立大学
Osaka Metropolitan University



東京都市大学
TOKYO CITY UNIVERSITY



Computing the Wave

Where the Gravitational Wave Community
benefits from High-Energy Physics, and where it differs ?

22nd International Workshop on Advanced Computing and Analysis Techniques in Physics Research
March 13th, 2024 - Stony Brook University

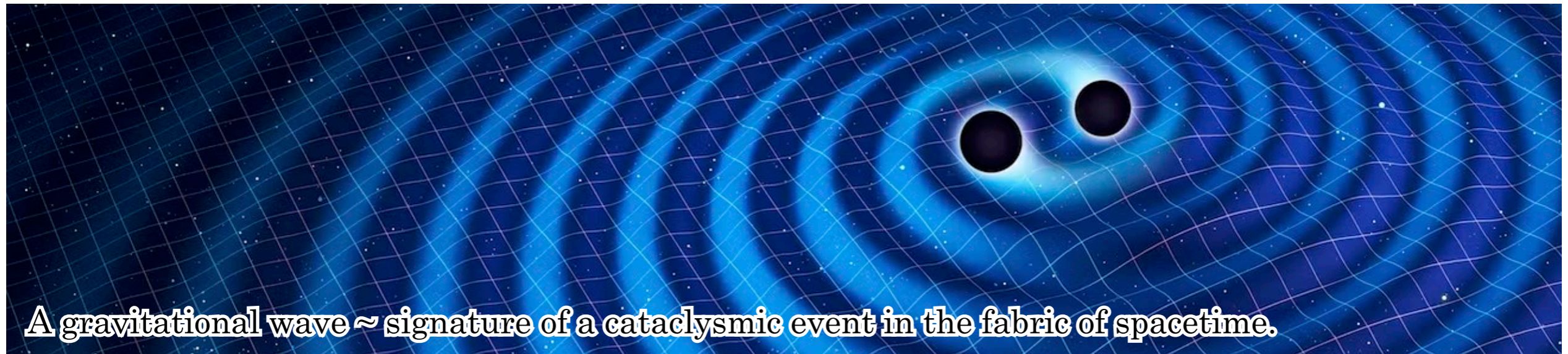
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Hirotaka Takahashi², Ken-Ichi Oohara⁴, Kazuki Sakai⁵

On Behalf of KAGRA Collaboration

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³University of Illinois at Urbana-Champaign, ⁴Niigata University, ⁵NIT Nagaoka College



GRAVITATIONAL WAVES (GW) IN A NUTSHELL



[e.g. two black holes (BH) or neutron stars (NS) colliding in space]

★ Key milestones in the History of GW Research:

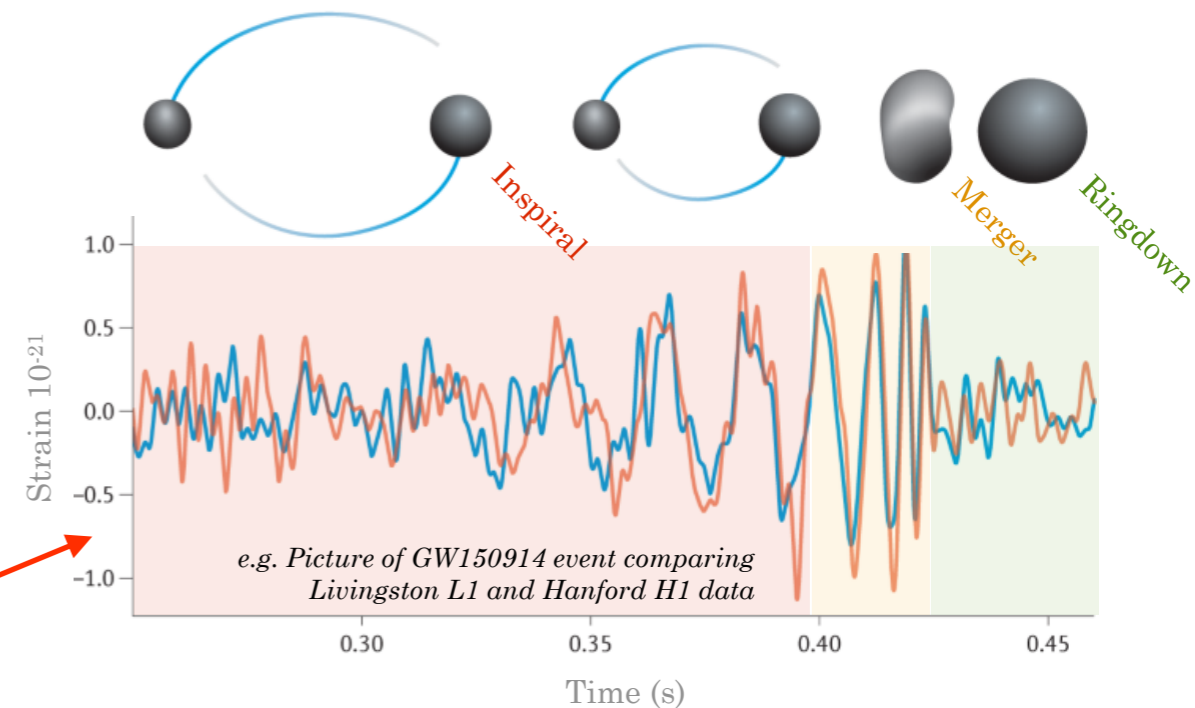
1916 — Prediction of existence by A. Einstein

1974 — First indirect evidence: Hulse-Taylor Pulsar

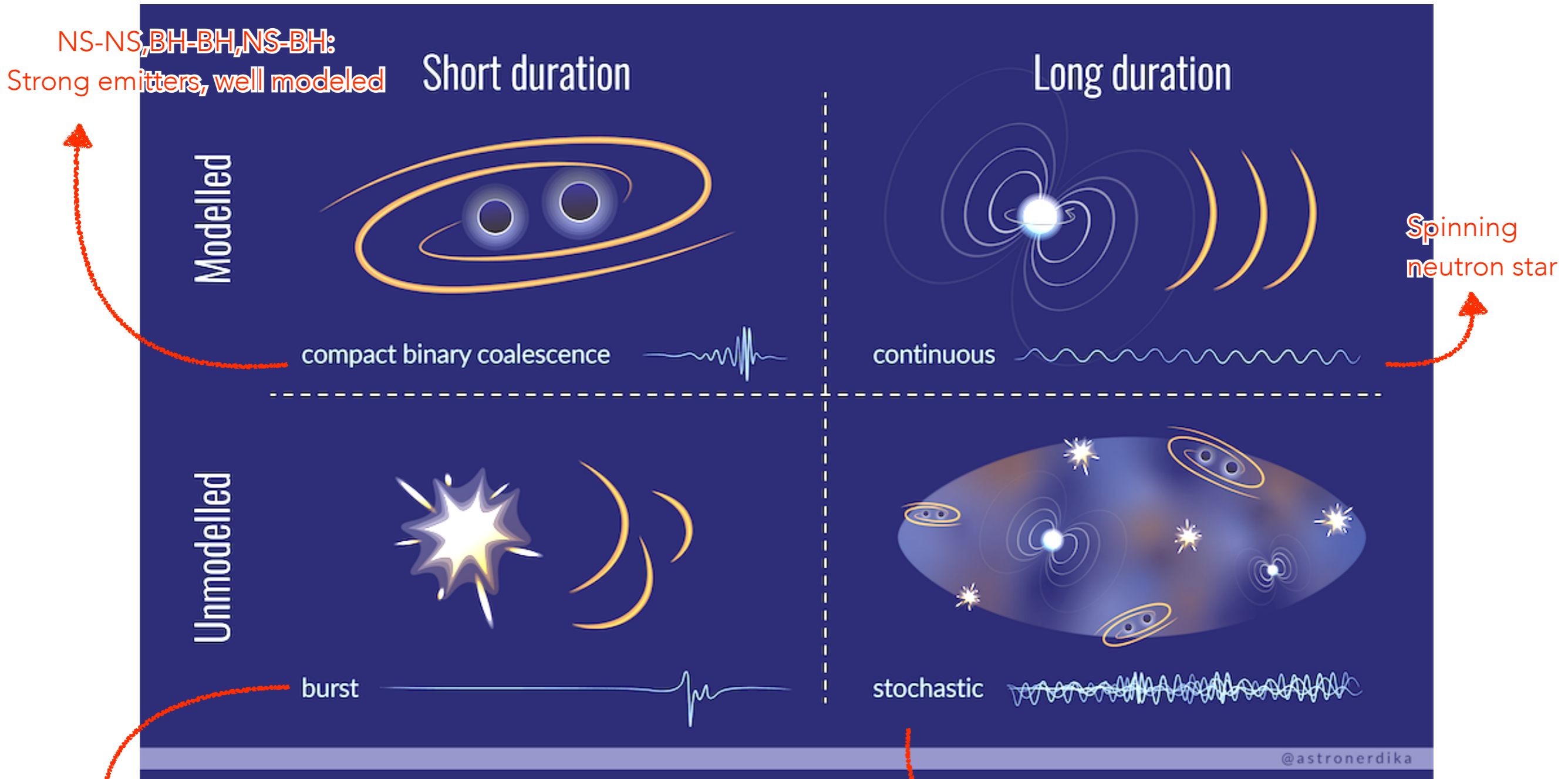
90-00's — LIGO construction (Livingston & Hanford, US)

2015 — First direct BH-BH merger detection: [GW150914](#)

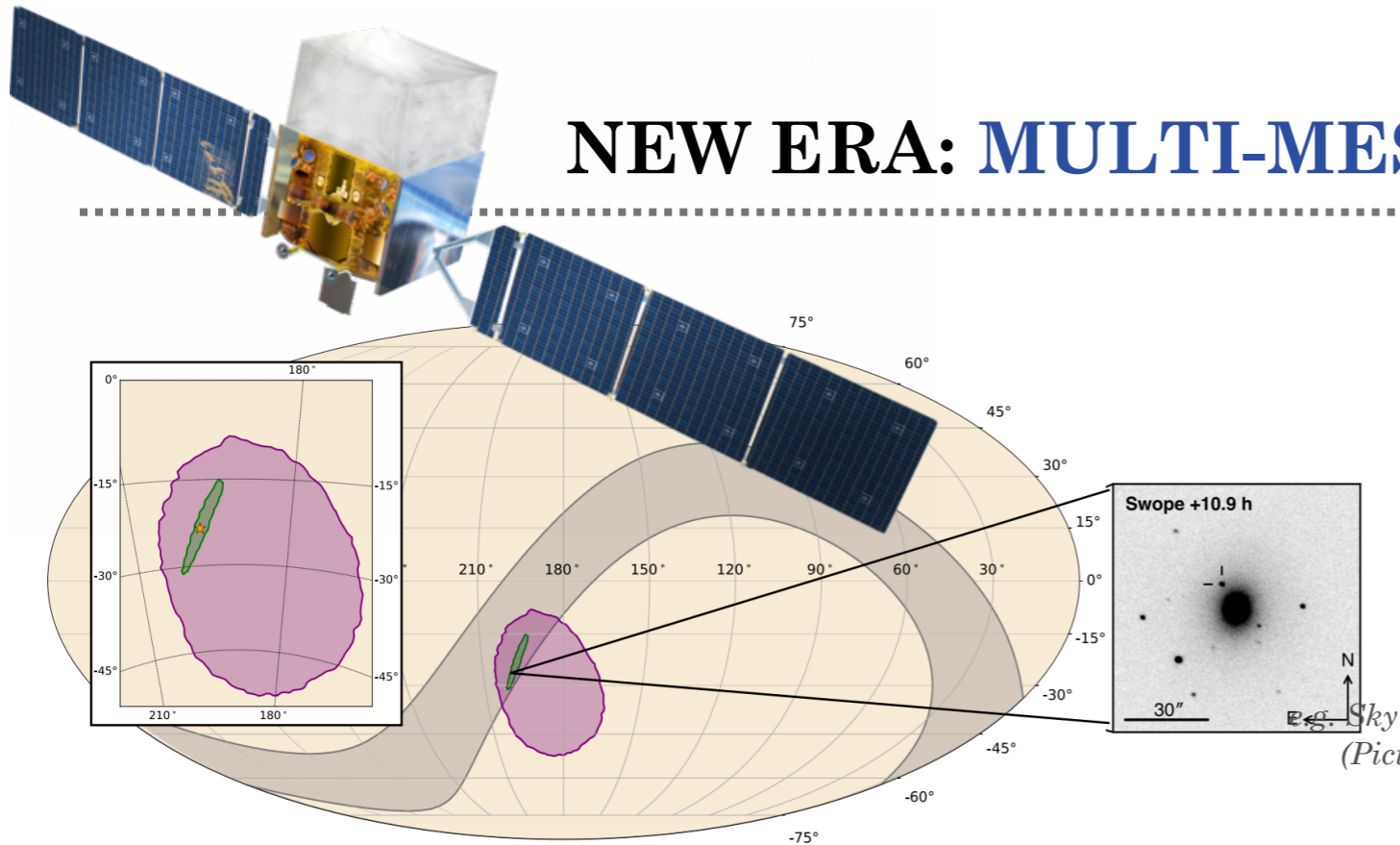
2017 — First NS-NS detection merger with Electromagnetic Counterpart: [GW170817](#) / [GRB170817A](#) (New Era)



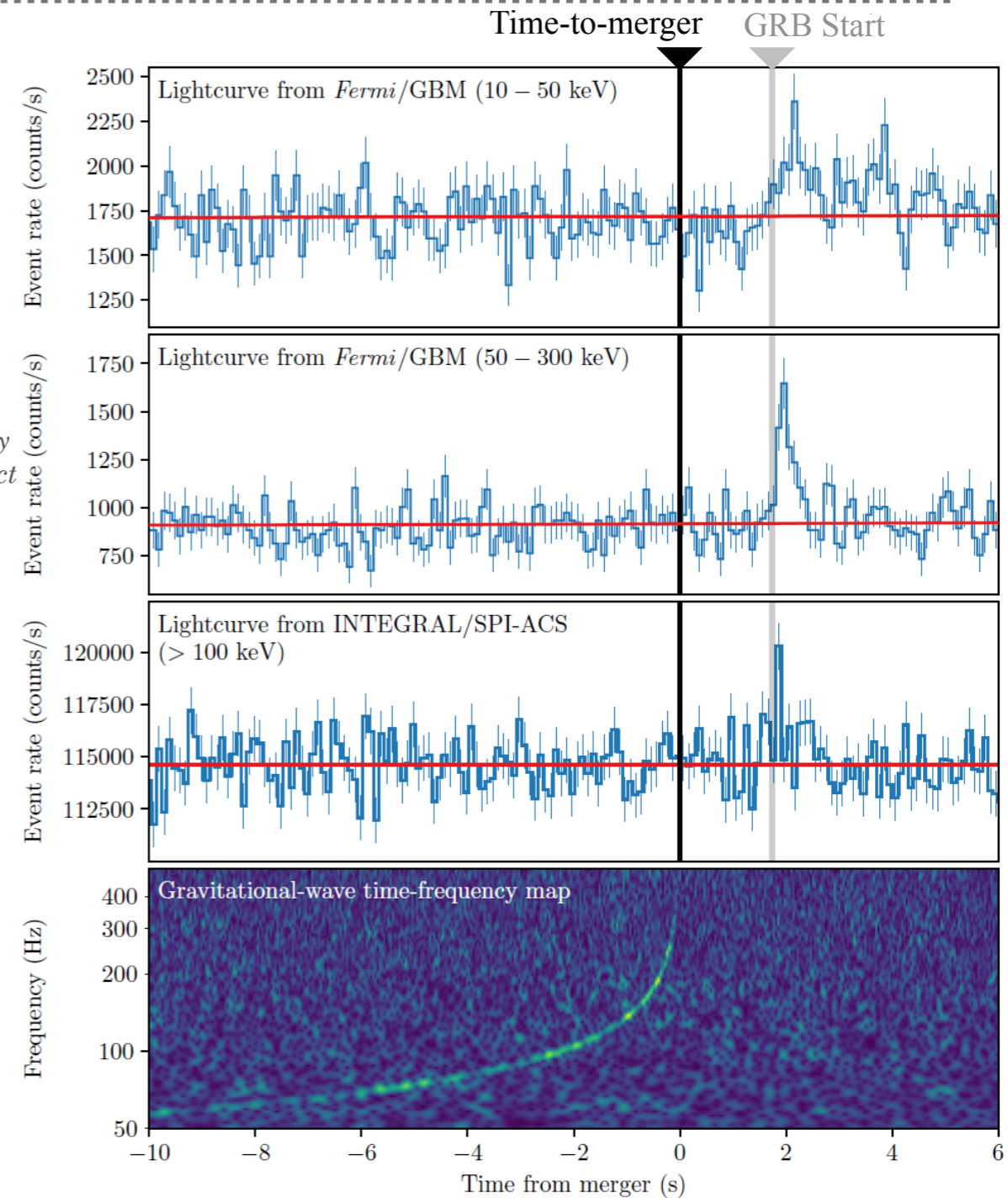
TOPOLOGY OF GRAVITATIONAL WAVES



NEW ERA: MULTI-MESSENGER ASTRONOMY

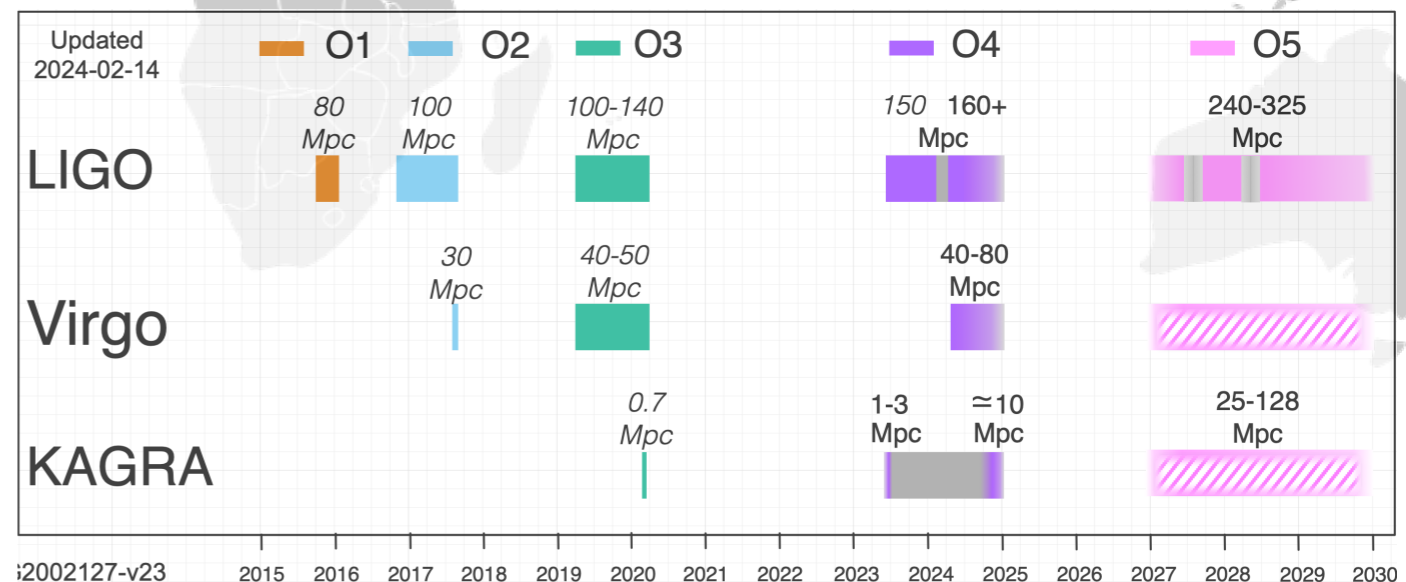
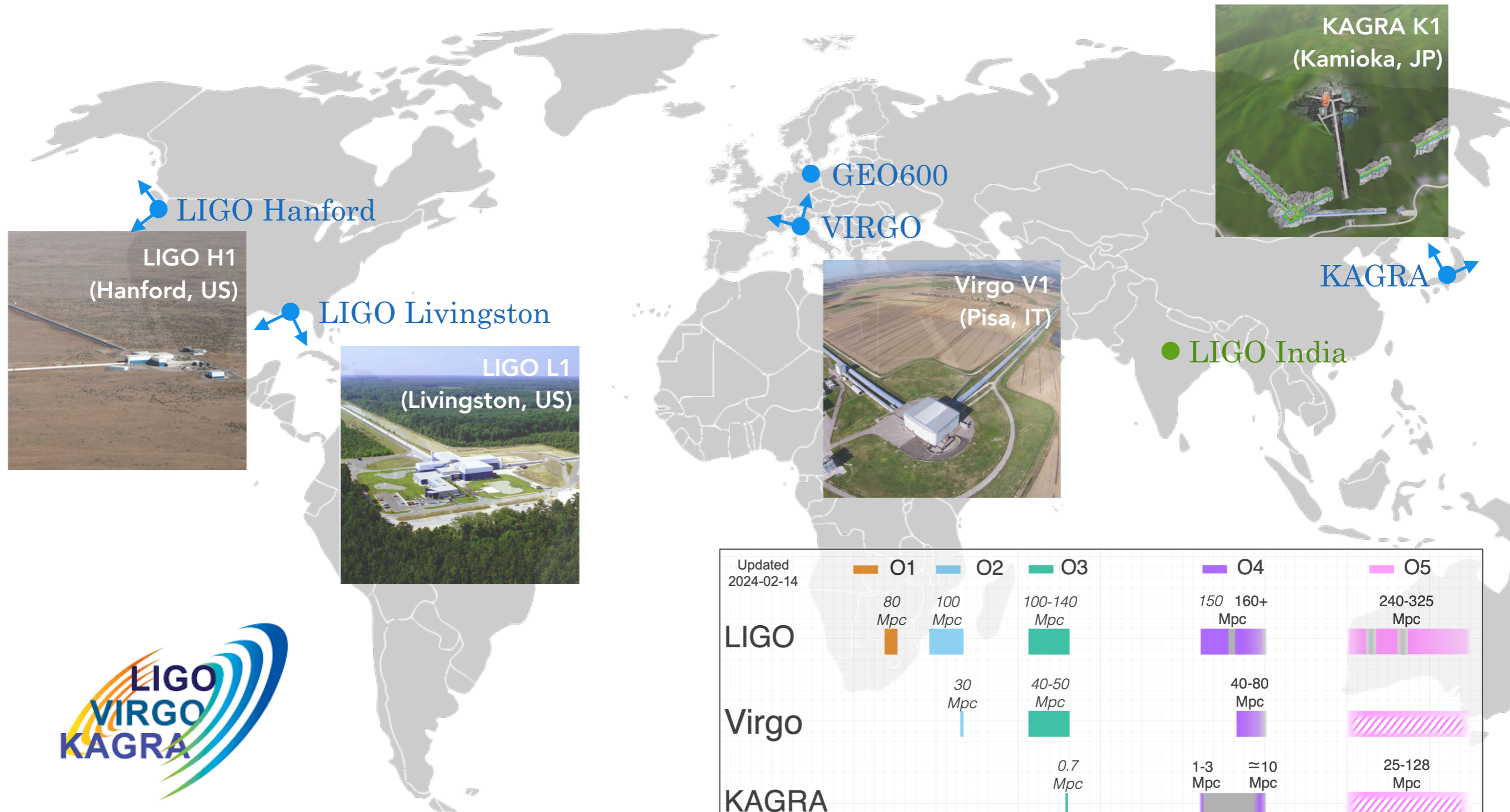


- ▶ Multi-messenger Astronomy (MMA) Era (started in 2017)
 - IGWN: Ground-based coordination for low-latency alerts
 - NASA: Coordination between space and astronomers.
 - This challenge is about speed (< 60s) not size (~1Mb/s)
- ▶ Stereoscopic sky localization to be provided
 - The more data, the better the localization is.
 - Goal: GW signal with negative latency
(meaning.. detection prior time-to-merger)
- ▶ Interest: connecting GW and HEP fields.
 - Gravitational wave observatories
 - Electromagnetic telescopes (e.g. *Fermi*)

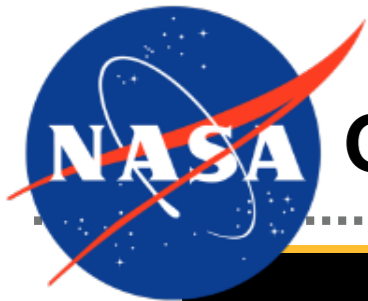


First NS-NS Detection with Electromagnetic Counterpart
(Landmark event: GW170817 / GRB170817A)

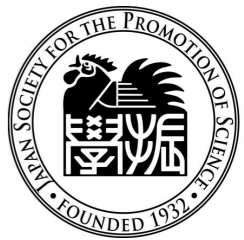
INTL. GRAVITATIONAL WAVE NETWORK (IGWN)



e.g. Future projects: Einstein telescope, LISA interferometer, etc.



GENERAL COORDINATES NETWORK (GCN)

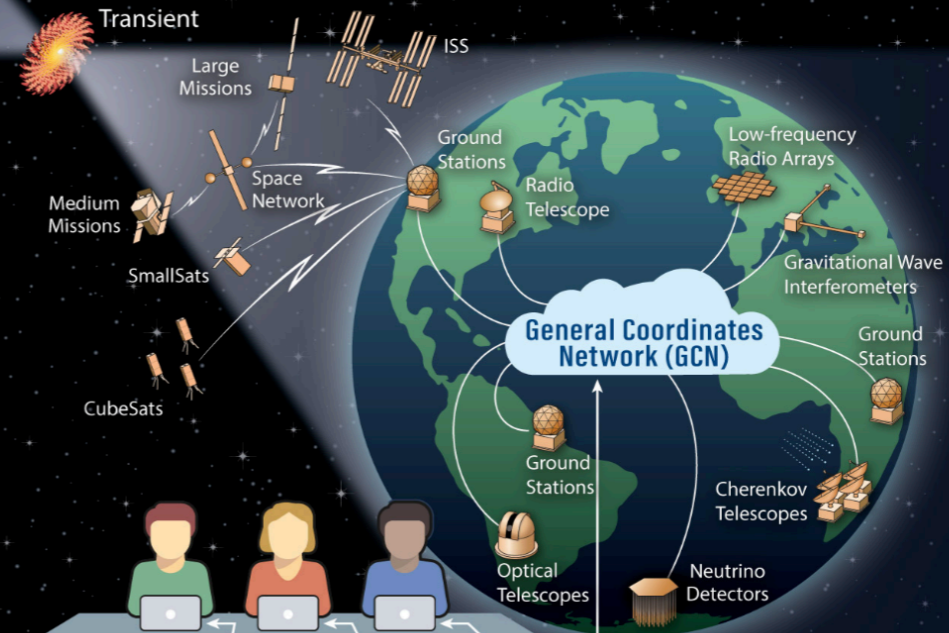


GCN: NASA's Time-Domain and Multimessenger Alert System

GCN distributes alerts between space- and ground-based observatories, physics experiments, and thousands of astronomers around the world.

[Start streaming GCN Notices](#)

[Post a GCN Circular](#)



LIGO/Virgo/KAGRA Public Alerts .. March 1st, 2024

O4 Significant Detection Candidates: **81** (92 Total - 11 Retracted)

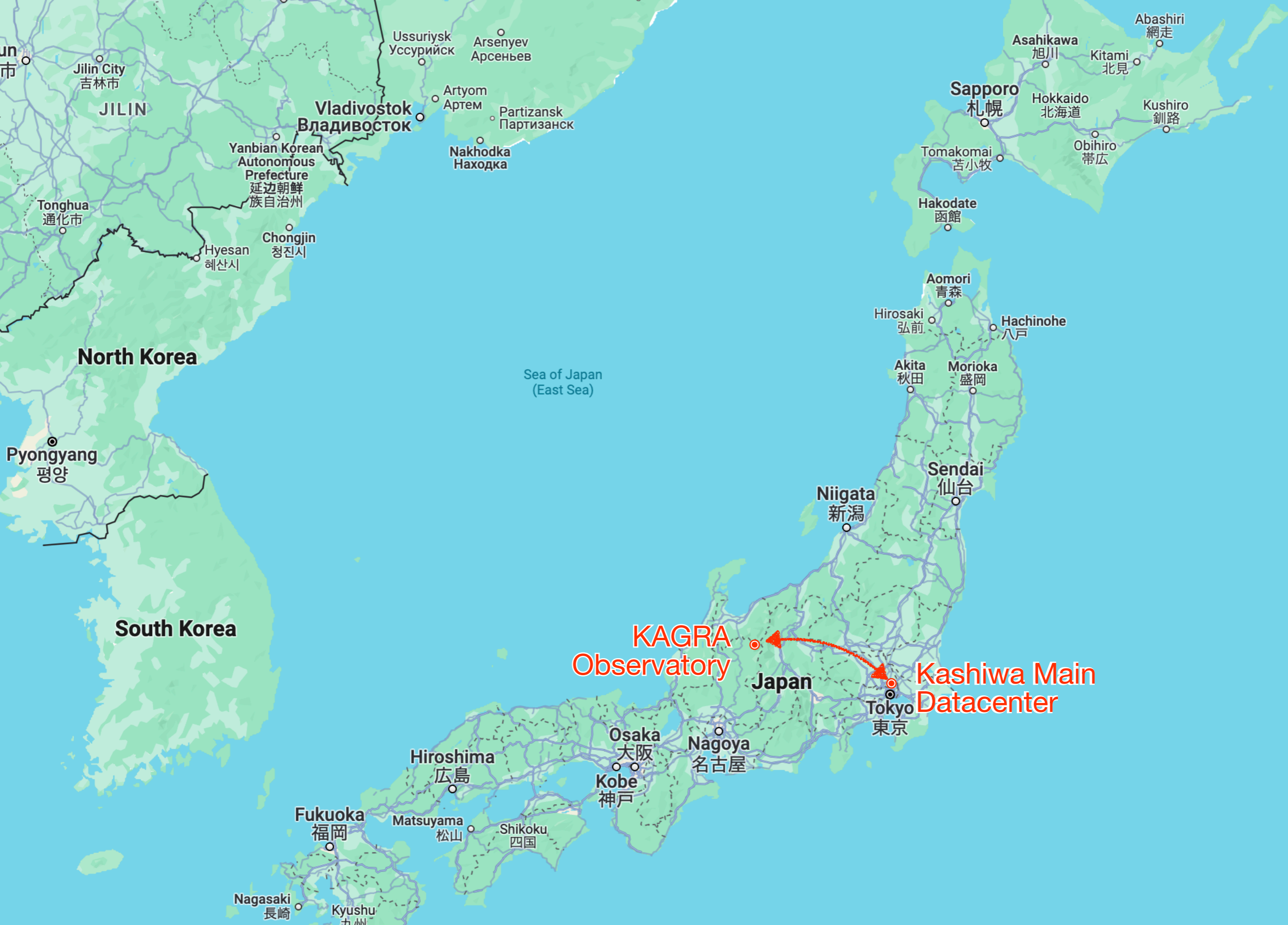
O4 Low Significance Detection Candidates: **1610** (Total)

Gravitational-Wave Candidate Event Database (GraceDB)

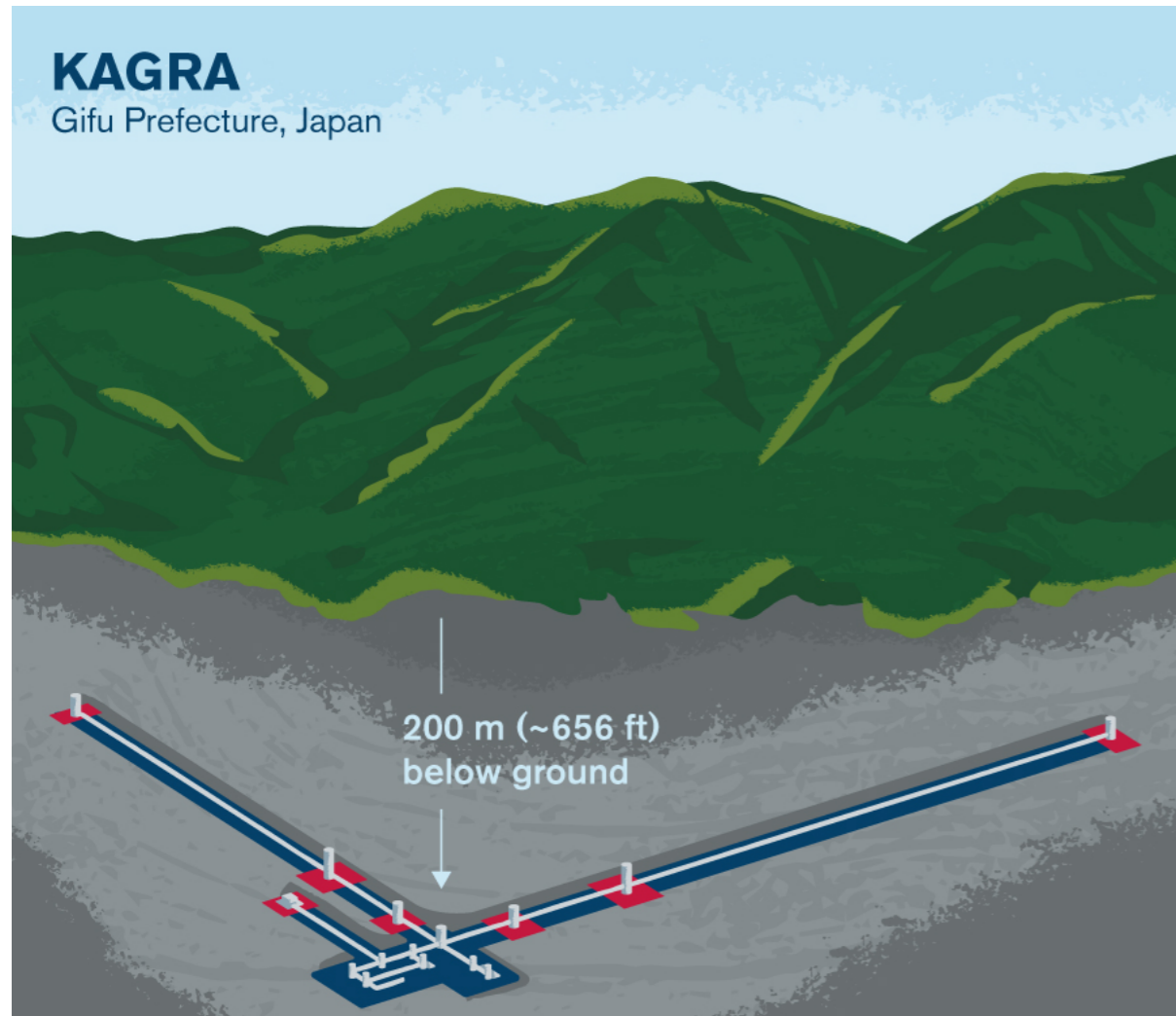
Event ID	Possible Source	Significant	UTC	GCN	Location	FAR
S240109a	BBH (99%)	Yes	Jan. 9, 2024 05:04:31 UTC	GCN Circular Query Notices VOE		1 per 4.3136 years
S240107b	BBH (97%) Terrestrial (3%)	Yes	Jan. 7, 2024 01:32:15 UTC	GCN Circular Query Notices VOE		1.8411 per year
S240104bl	BBH (>99%)	Yes	Jan. 4, 2024 16:49:32 UTC	GCN Circular Query Notices VOE		1 per 8.9137e+08 years
S231231ag	BBH (>99%)	Yes	Dec. 31, 2023 15:40:16 UTC	GCN Circular Query Notices VOE		1 per 3.7932e+06 years

- **GraceDB** — Centralized API for public alerts and sending GCN Circular Query
- **Celery message broker** — Full orchestration performed with GWCelery package
- **Early public warning alerts** — for rapid comms. (<https://gcn.gsfc.nasa.gov/>)

...

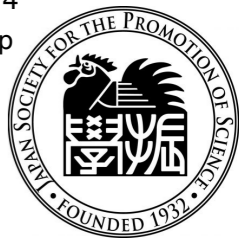


FOCUS ON KAGRA EXPERIMENT

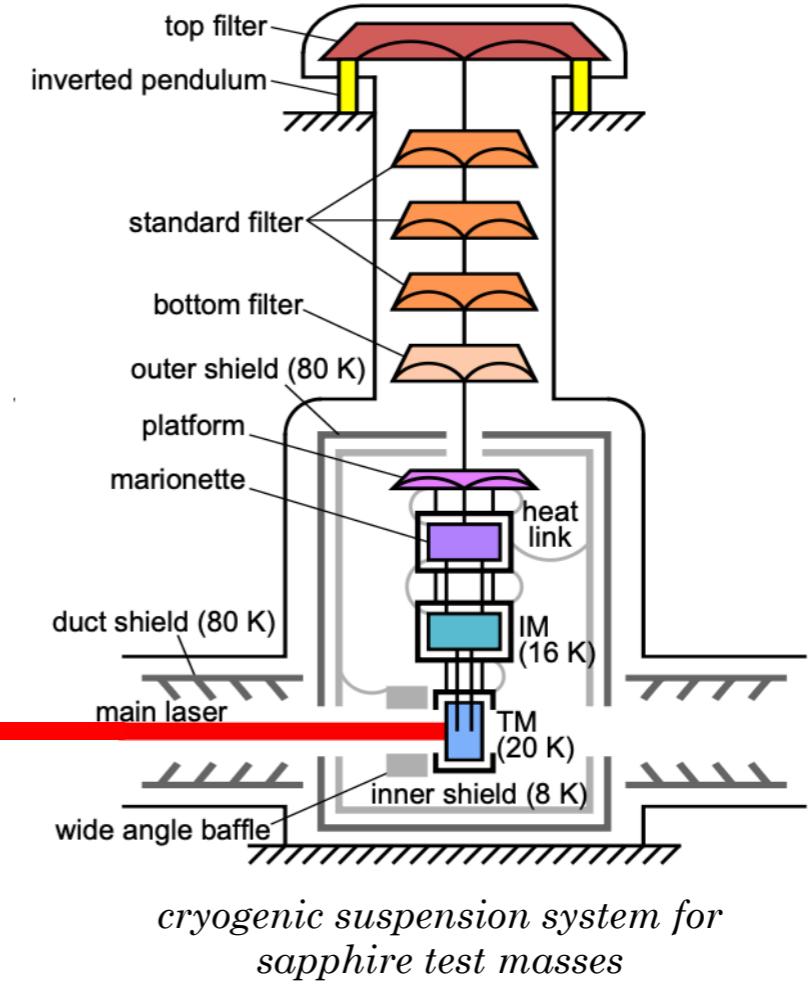
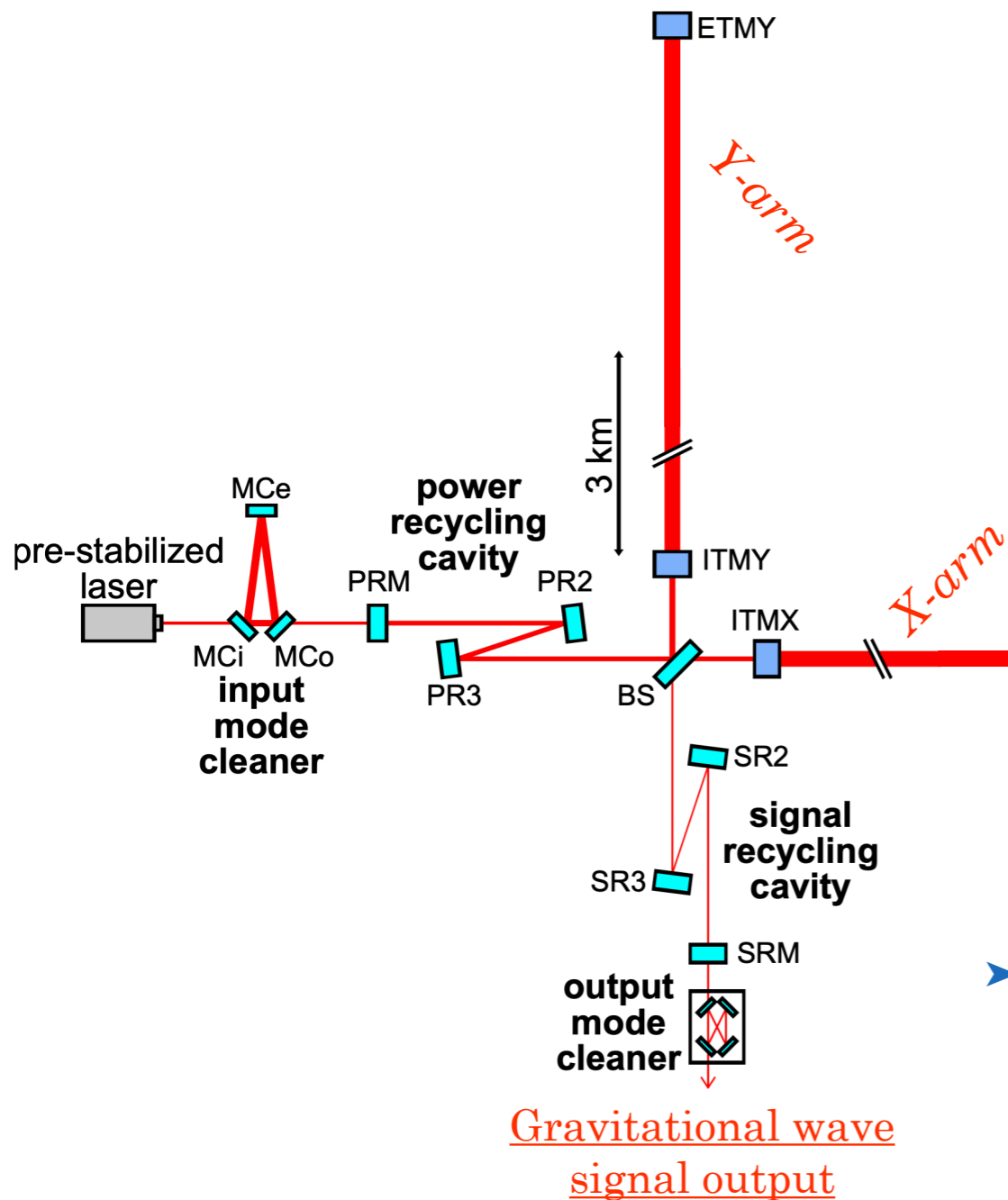


KAGRA Observatory has been established in 2016 in Kamioka, Gifu prefecture
(*same place as Kamiokande neutrino experiment*)

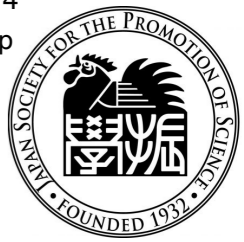
The telescope is a 3 km Fabry-Perot interferometer
with **20K Cryogenic Sapphire Mirrors** in Mt. Ikenoyama.



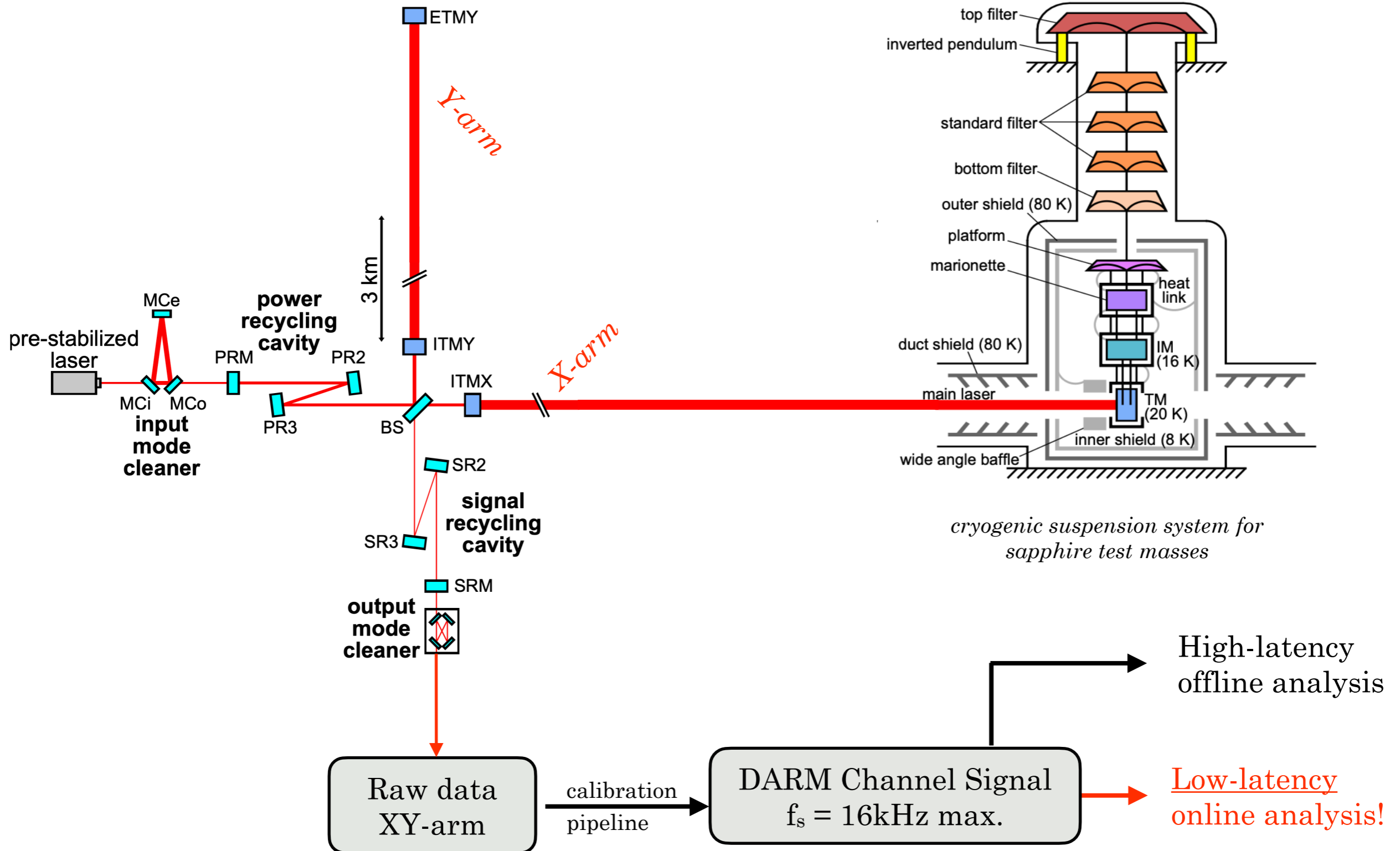
ON-SITE DATA ACQUISITION (AT KAGRA)

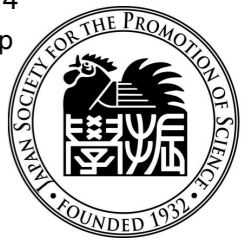


- **Fabry-Perot Interferometer (power recycling)**
 - Laser 180W continuous source at $\lambda = 1064 \text{ nm}$
 - XY-arms, 20K-cryogenic cooled suspended mirrors
- **On-site infrastructure actions**
 - Data acquisition, pre-processing, environmental monitor.
 - Feedback control to actuators on ETMX,ETMY

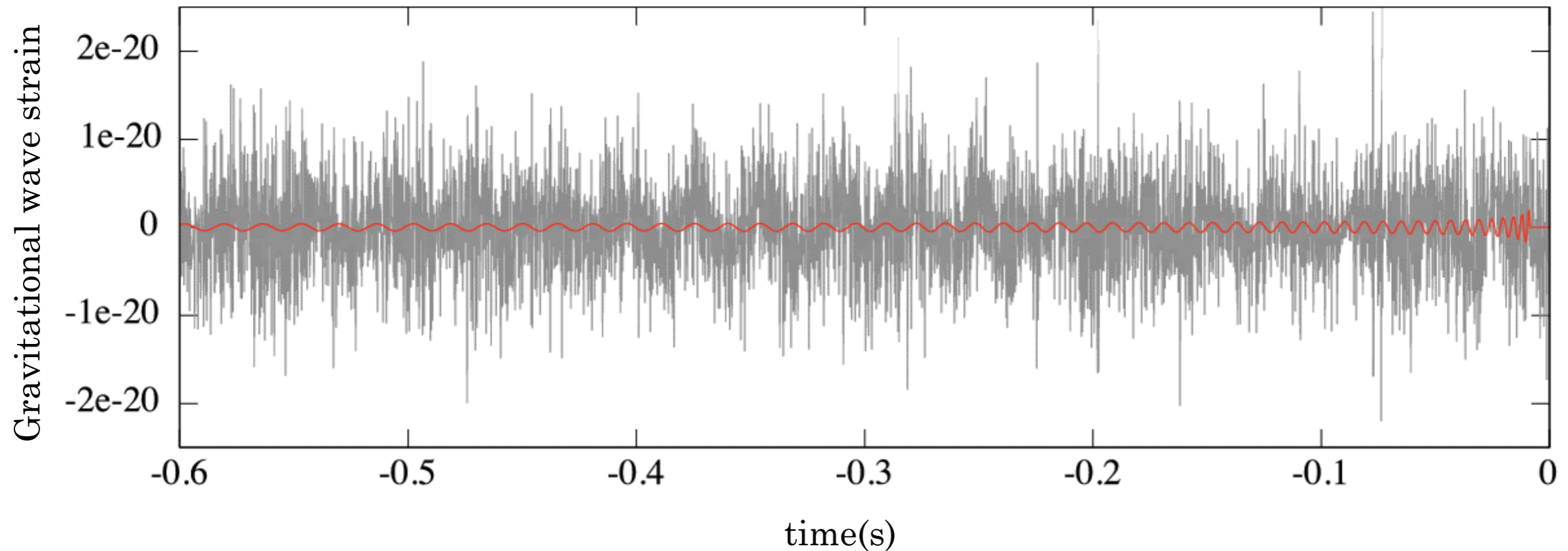


ON-SITE DATA ACQUISITION (AT KAGRA)





ON-SITE DATA ACQUISITION (AT KAGRA)



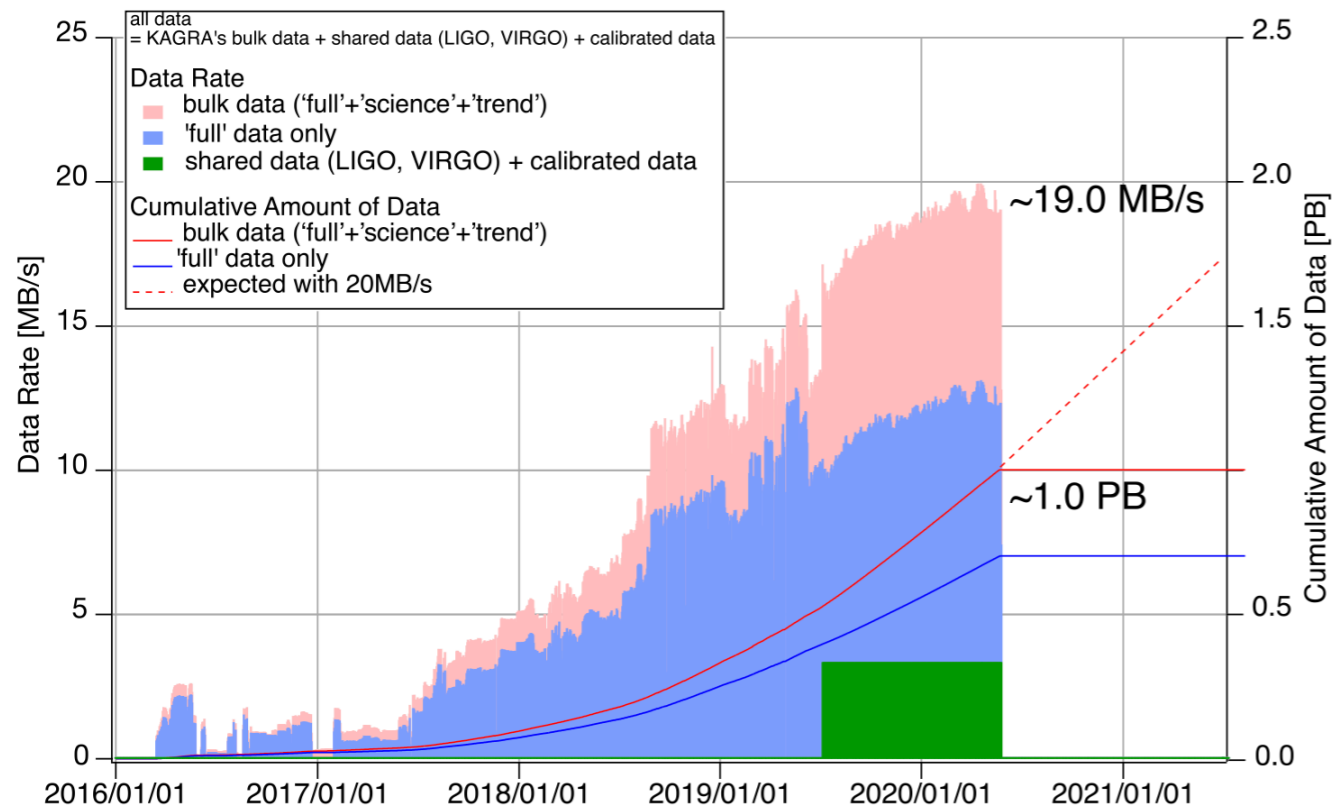
- Gravitation wave signal (here is LIGO experiment); typical signal injection in red
- Data transfer via Apache Kafka (message broker): low-latency reliable sustained data-stream
- Typical latencies at Kashiwa Data center for “1 second” files:

LIGO to Kashiwa	6-8s latency
VIRGO to Kashiwa	10s latency
KAGRA to Kashiwa	2.5s latency

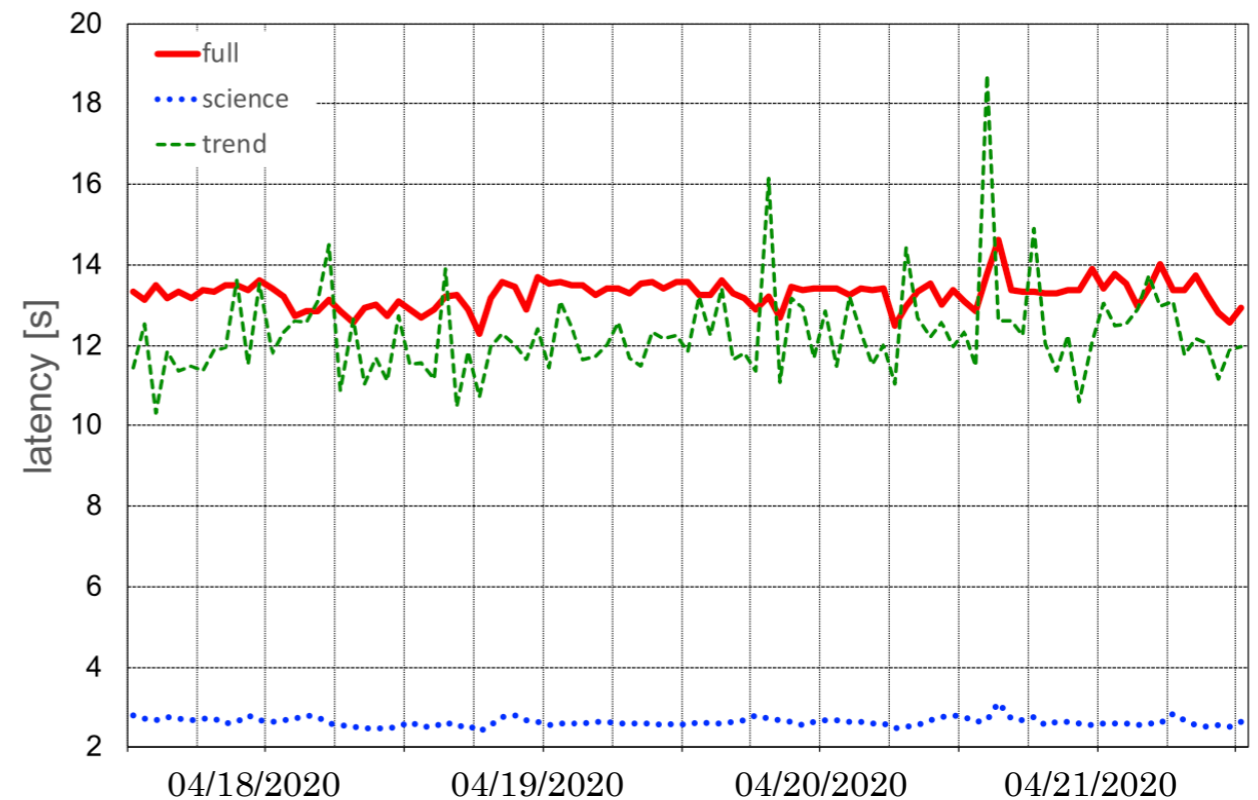


(Latency = transfer from ADC to on-site servers + FIR filtering
+ transfer from On-Site (L1,H1,V1,K1) to Kashiwa)

FOCUS ON: KAGRA (KAMIOKA) TO KASHIWA MAIN CENTER (TOKYO)



(Data rate and quantity since beginning of KAGRA and projection for O4)

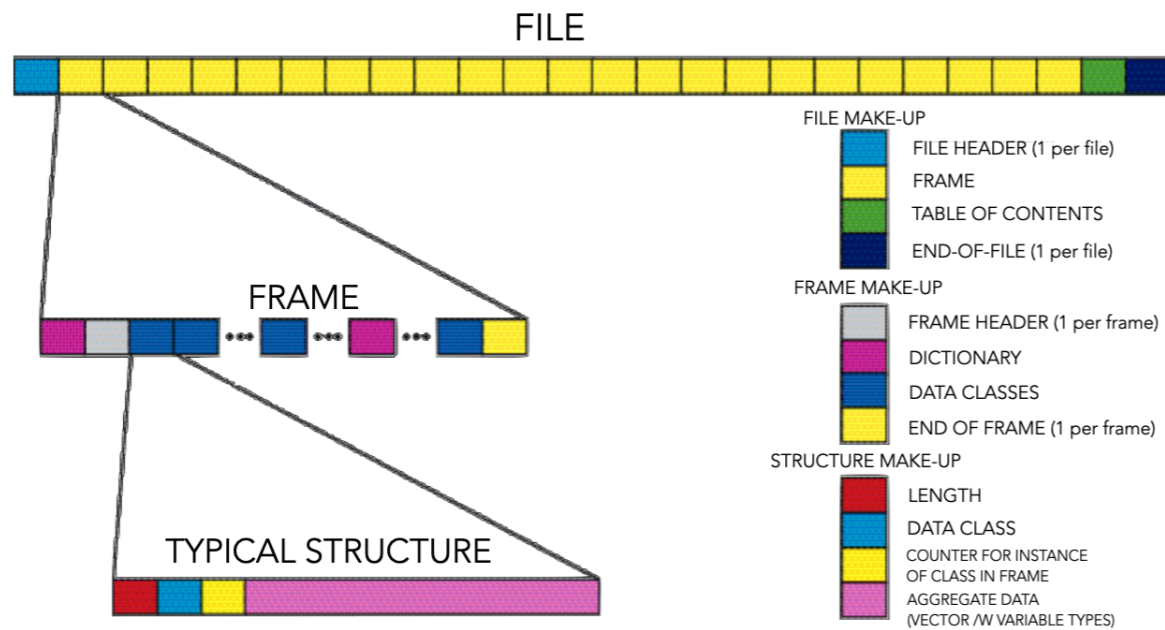


(Latency transfer of larger 32-seconds data files from KAGRA to Kashiwa)

- **Low-latency Online Analysis O4a ~ O(1min)**
 - Data transfer of 1s-files via “Apache Kafka”
 - Run O4a ~ 1.5PB/y of raw data
 - Research pipeline and rapid sky localization
 - Goal: Alert generation and flagging

- **High-latency Offline Analysis**
 - Data management /w “Rucio” (GridFTP/CERN)
 - Offline event search, detector characterization
 - Make use of for sharing data:
 - * CVMFS (CERN) for “data” and “software”
 - * OSDF Data origin (OSG) for “data”

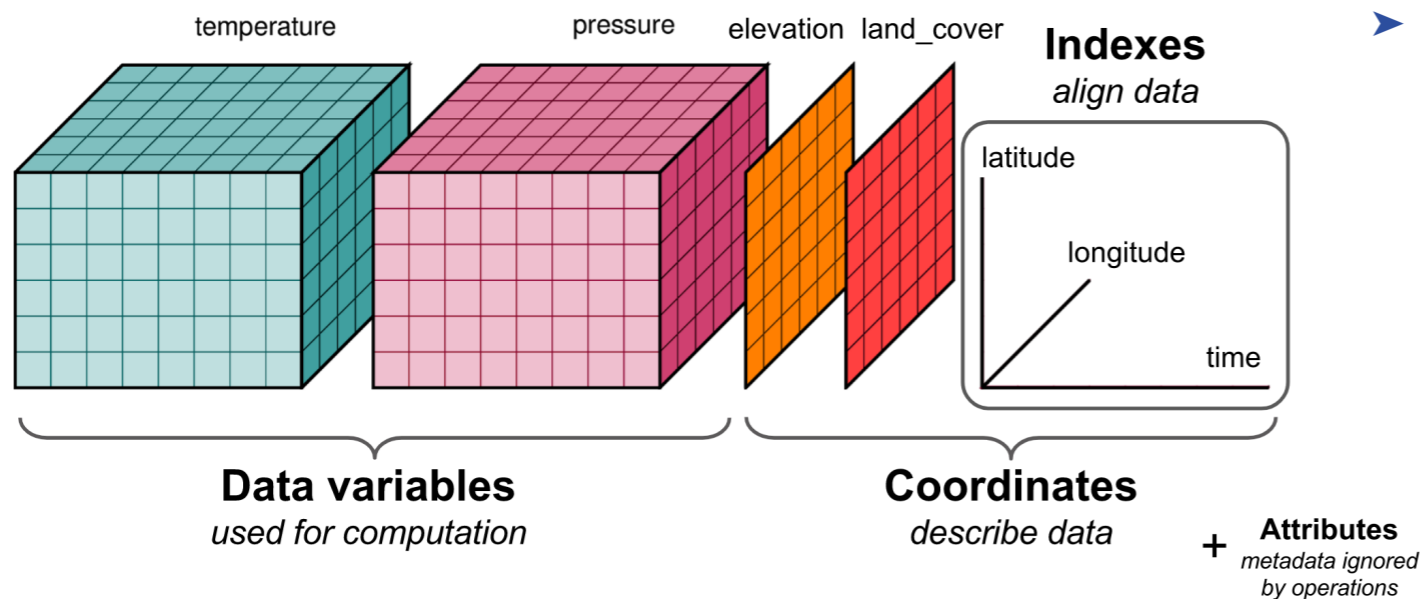
GRAVITATIONAL WAVE DATA FORMATS



— Picture of the Frame Format data structure for low latency —

- **Low-latency data format:** `*.gwf`
 - “Frame Format” standard by LVK collaboration
 - Since 2019, strain data $h(t)$ are 1s frame size.
 - Optimized format to fast reading/writing

Data Format specific to LVK collaboration



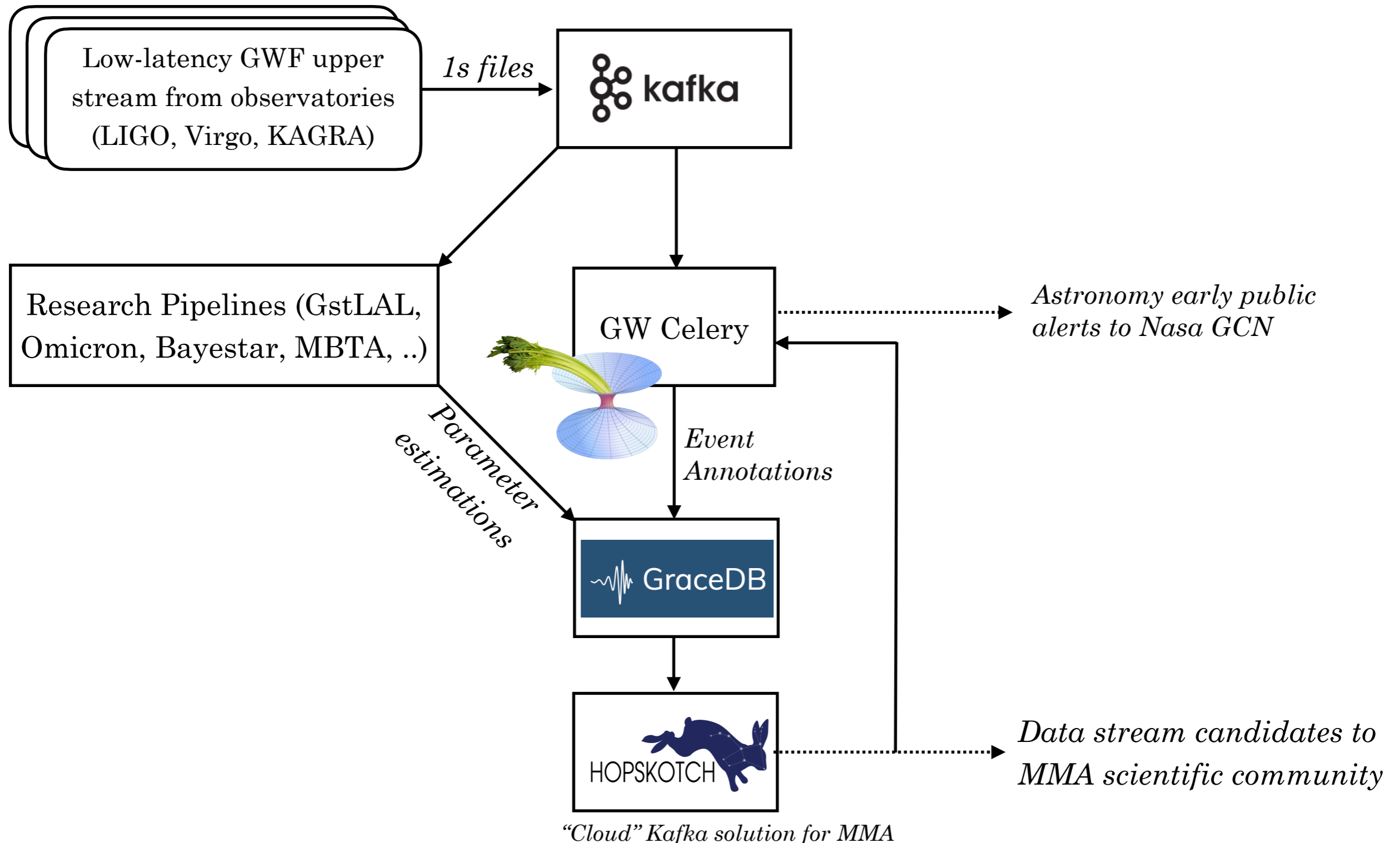
— Diagram of the HDF5 data format for high latency —

- **High-latency data formats:** `*.gwf` `*.hdf5`
 - Usually HDF5 is used for Open Science Data
 - More flexible when using with Python



(<https://gitlab.cern.ch/escalade/standalone/hdf2root>)

LOW-LATENCY ORCHESTRATION OVERVIEW





IGWN O3/O4 COMPUTING RESOURCES

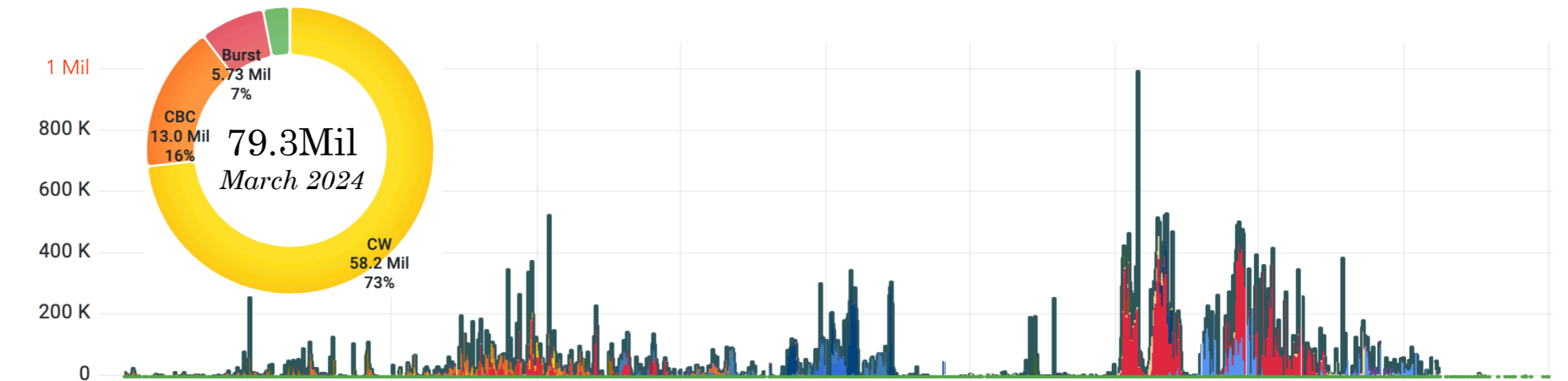
► Observing run periods:

- O3a: April 1, 2019 to October 1, 2019
- O3b: November 1, 2019 to March 27, 2020
- O4a: May 24, 2023 to January 16, 2024
- O4b: April 3, 2024 to the end of year

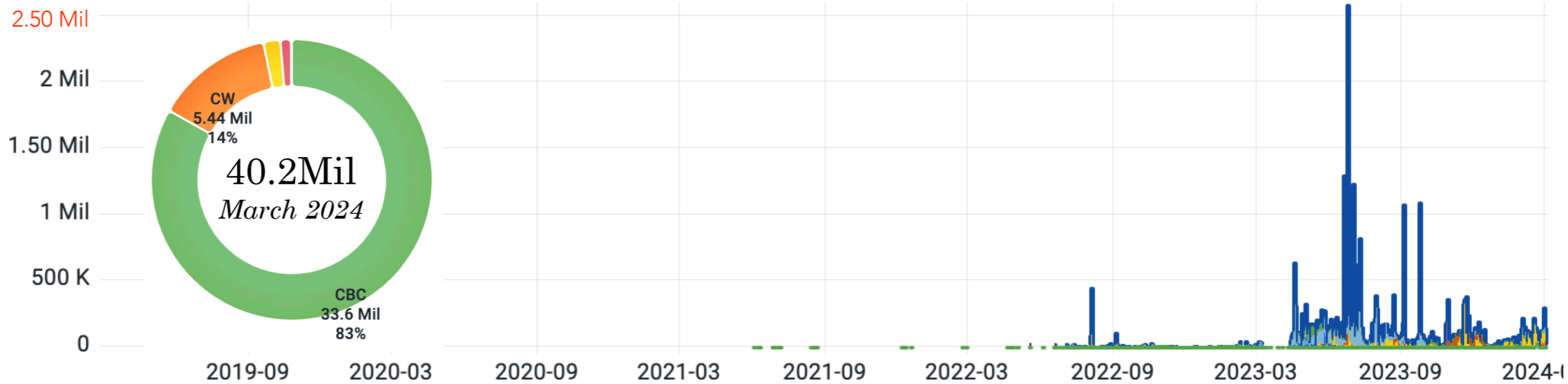
► IGWN Computing Resources

- Computing policy and common coordination
- Code management: Gitlab, Conda & Docker
- Resource monitoring: Grafana
- No common framework for development.

(O3)



(O4)



COMPUTING PIPELINES AND COMMON FRAMEWORK

- ▶ **IGWN Collaboration Goal: to reduce the amount of tools**
 - Interferometer sensitivity \nearrow ; Detection range \nearrow ; Smaller noise \searrow ; Signal-to-Noise-Ratio (SNR) \nearrow
 - Raw data is not increasing much over years (main computing resources is about template generation)
 - New technics developed in addition to the classic technics. (Matched Filtering, Wavelet, etc.)
- ▶ **Computing pipelines are heterogeneous and complex (no common framework)**
 - Many languages: Matlab, C/C++, Python, ..
 - Main scientific package: LVK Algorithm Library Suite (LALSuite; C/C++)
[Noise simulation, template generation (GR modeling for matching data)]
 - Machine learning increasing usage: First machine learning pipeline MLy is online.



*ROOT has already been used in GW community
(mainly people in Europe, originating from HEP).*

- ▶ **Requirements to make use of ROOT in GW community**
 - (1) Handling typical I/O files: GWF files and HDF5 files
 - (2) Advanced digital signal processing and (3) complex digital filtering
- ▶ **Personal interests in using CERN modern/high-level tools** (RNTuple, RDataFrame and SOFIE)
 - investigation and development planned for RDataFrame to include TimeSeries capability (1 event = 1s Δt)

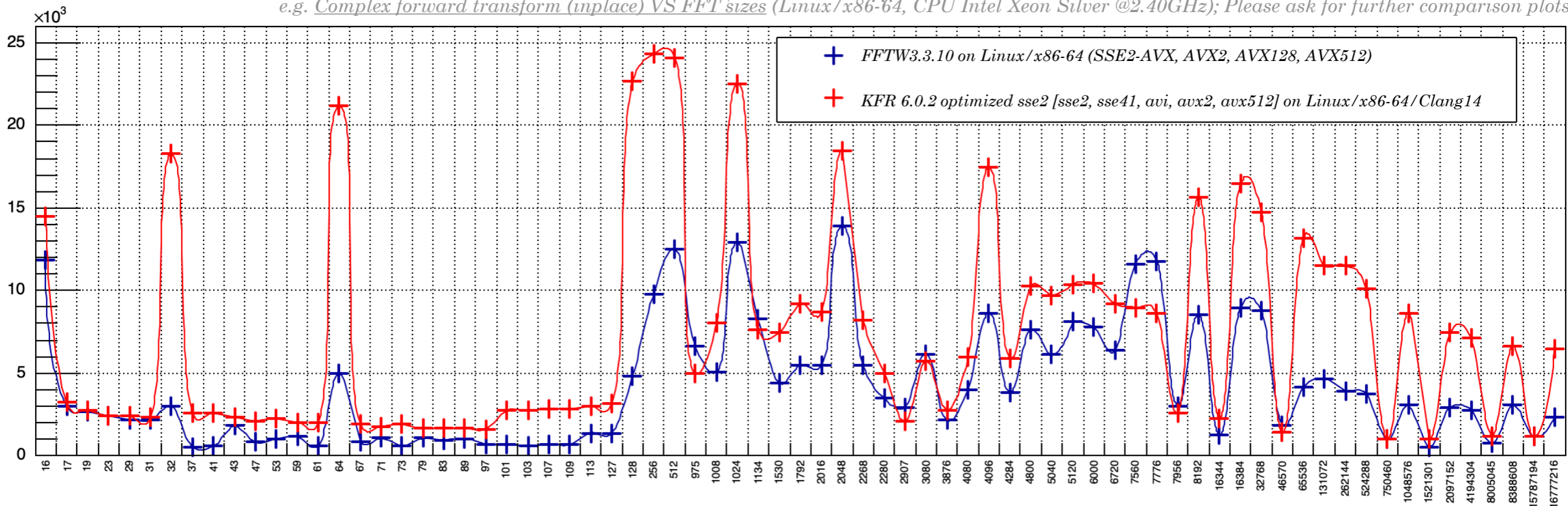
FAST, MODERN C++ DSP FRAMEWORK

KFR is built for developers who need a powerful toolkit for audio and dsp applications

- KFR library under GPL licensing.
 - DFT & FFT at N-dimensions available
 - Advanced signal processing (IIR/FIR filtering)
- Highly optimized modern C++ library
 - optimal on x86-64 architectures using Clang
 - optimized perf. up to 40% in best cases vs. FFTW3
- Compatible with both ARM64 and x86-64 archs.
 - SIMD SSE, SSE2, AVX instructions, for x86_64
 - SIMD NEON64 instructions for ARM64.
- Benchmark tests performed with last version.
 - Comparison between KFR6.0.2 and FFTW3.3.10
 - <https://git.ligo.org/kagra/containers/fft-benchmark>

e.g. Complex forward transform (inplace) VS FFT sizes (Linux/x86-64, CPU Intel Xeon Silver @2.40GHz); Please ask for further comparison plots

MFlops (The Higher, The Better!)



— One of the benchmark test result (Mflops) comparing KFR 6.0.2 with FFTW3.3.10 / **The Higher the Better!** —

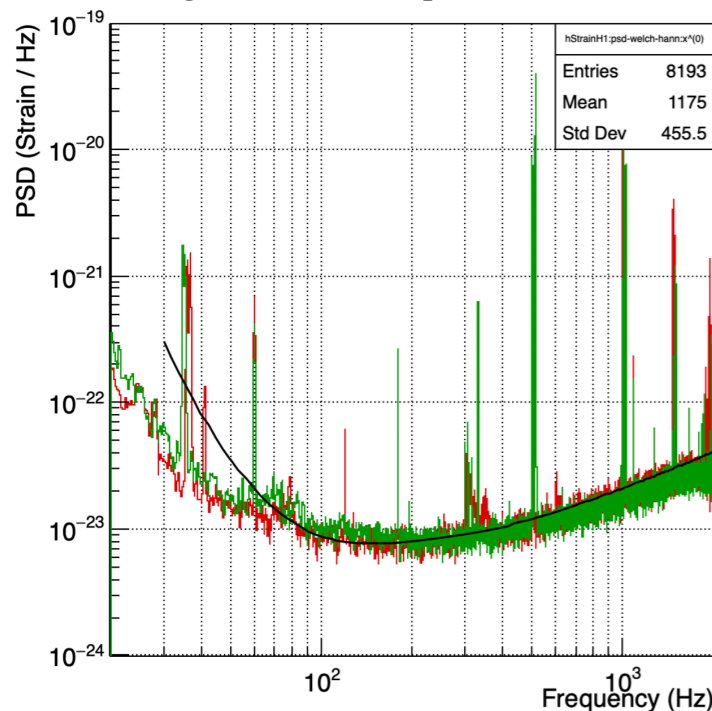
FFT Size



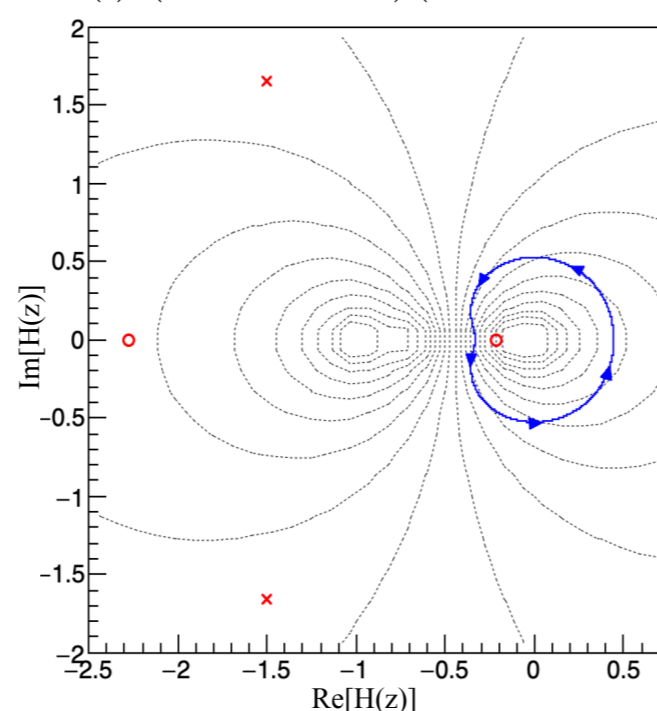
ROOT SIGNAL LIBRARY PROTOTYPE

- **Advanced DFT features using KFR library.**
 - ASD/PSD calculation, SNR calculation, Matched filtering, Convolution, Window, resampling, etc..
- **Analog and digital filter $H = B/A$ implementations. (ZPK, BA, SOS, SS representations)**
 - Finite/Infinite Impulse Response filtering
 - Chebychev, Butterworth, Elliptical filter implemented, [..]
- **TKFR Prototype class implemented with ROOT / Make use of the TVirtualFFT interface too**
 - ✓ **ASD, PSD, Hilbert transform, Elliptical filters, Wiener & Kalman filter**
 - FIR/IRR, Windowing, Decimation, Waveform generation (sine, impulse, steps, ramps, etc.), etc.**

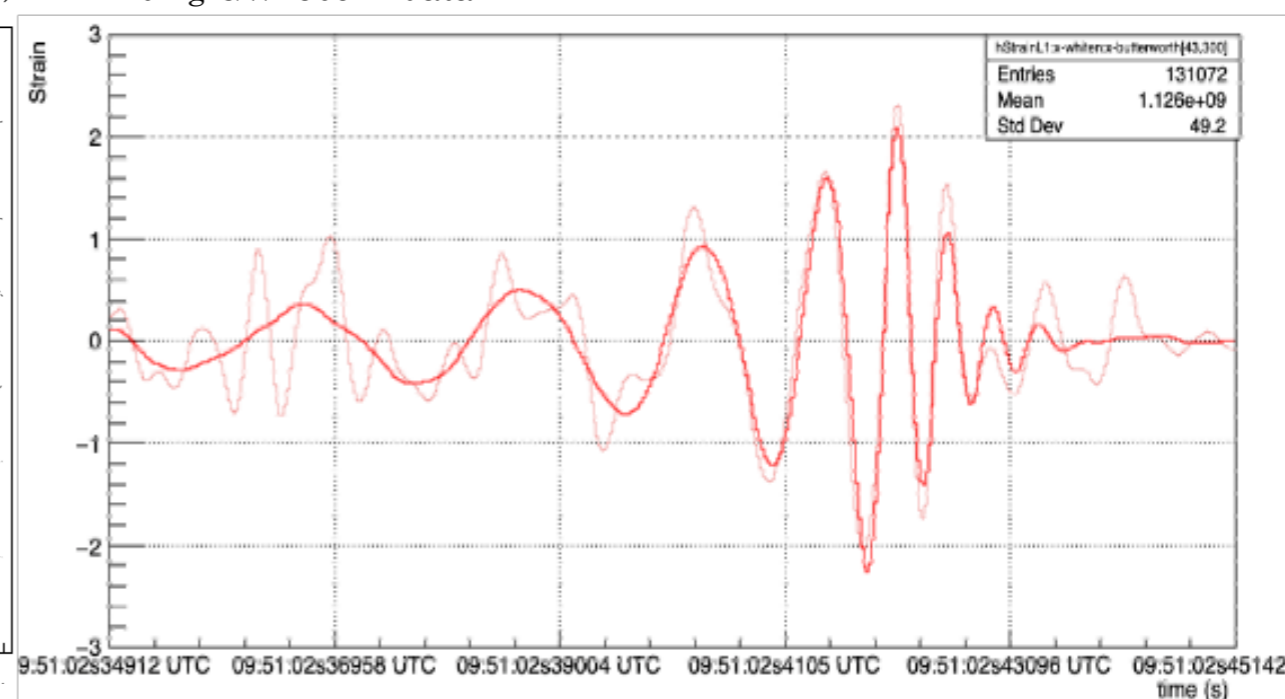
Power Spectrum Density calculation using GW150914 OpenScience data



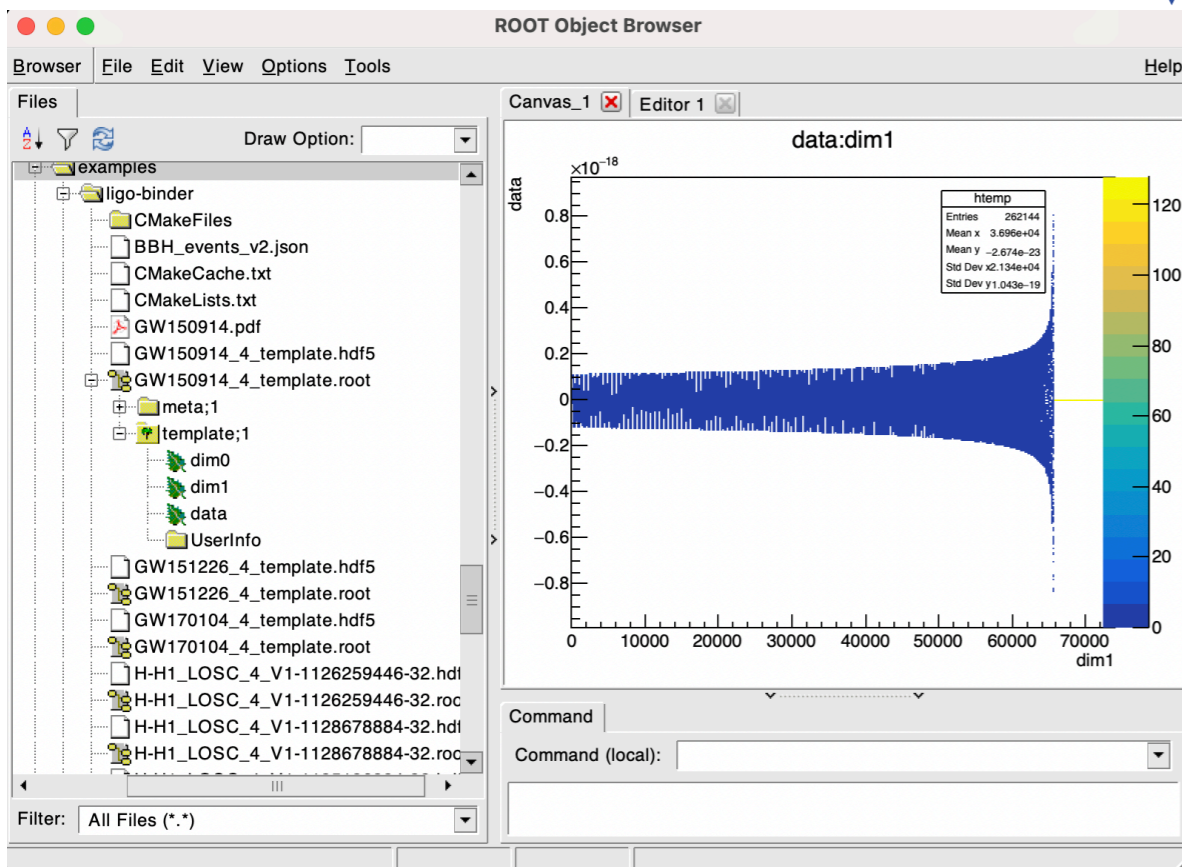
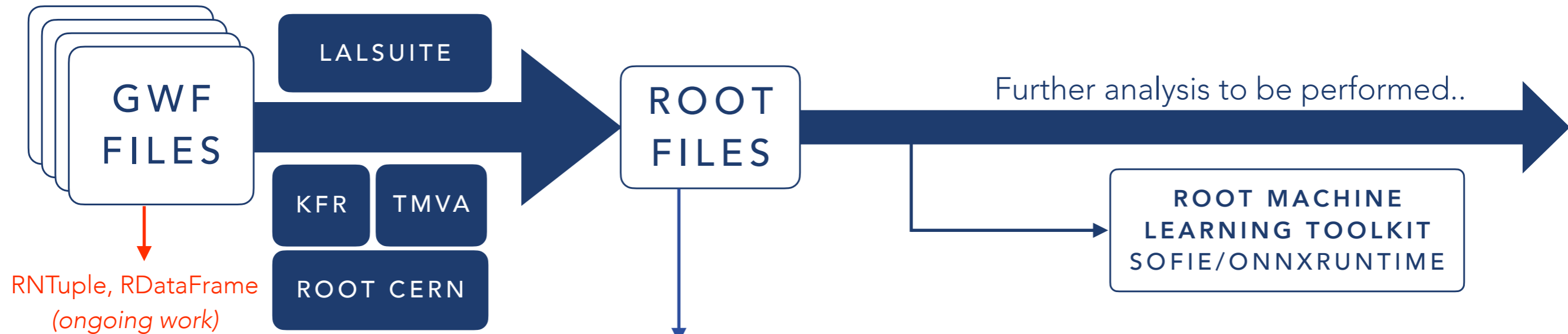
Example of digital filter implementation $H(z)=(5.0+3.0z+1.0z^2)/(1.0+5.0z+2.0z^2)$



Matched filtering techniques with “best template” using GW150914 data



SUMMARY: PROPOSED GW WORKFLOW



➤ ROOT is very much compatible with GW analysis:

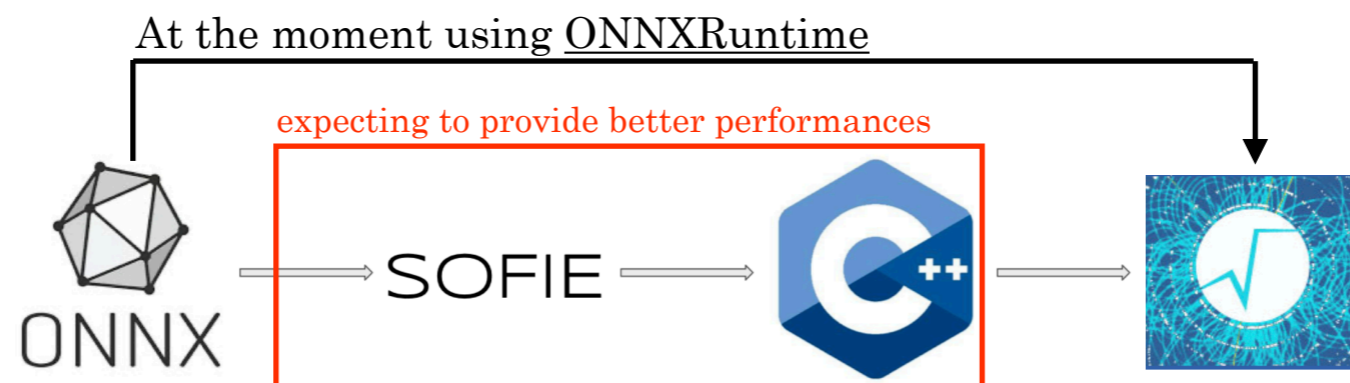
- ✓ KFR Library, TTimeSeries/TFrequencySeries prototype
- ✓ TKafka consumer/broker client prototype
- ✓ TConfigParser (YAML,JSON,TOML,XML,.. run-time files) prototype
- ✓ TCurlFile / TArchive prototype (download/cache files and archive handling)
- ✓ THDF5 parser prototype

➤ LALSuite Package linkable for direct use.

- ✓ TStrain, TStrainNoise, TStrainTemplate
(for direct strain, noise and template generation)
- ✓ TLigoFormat for configuration file parsing
- ✓ TFrameChain GWF proprietary format. (MT)

ULTIMATE GOAL: MACHINE LEARNING INFERENCE WITH ROOT

- ▶ **Starformer project**, a transformer-based model for early detection of CBC mergers
 - Collaboration with ISLab (*experts in transformer architecture development, Osaka University*)
 - *Explainable AI is of highest importance, flexibility for production models, parallel performances.*
- ▶ **Machine learning universal format for models**: “Open Neural Network Exchange” (ONNX) files.
(Training: pyTorch / Inference: ROOT)
- ▶ **Inference using ROOT + ONNX Runtime**:
 - ONNX Runtime: developed by Microsoft and Facebook. (Please visit <https://huggingface.co/> too!)
Interoperability in many languages. (Python, C++, ..)
 - Performance enhanced expected using SOFIE library in the future (maybe)



CONCLUSION

- **Gravitational wave community is building a new computing paradigm:**
 - Computing power requirement and increasing detector sensitivity.
 - Ongoing work (since many years), now resulting in the IGWN computing collaboration.
- **Computing power doesn't grow that much in GW compared to LHC experiments**
 - Still very challenging to communicate quickly with MMA institutes and astronomers.
 - Kafka integration prototype in an external ROOT library.
- **Goal: provide mature tools as a bedrock for analysis and pipeline development.**
 - Common framework, performant tools + keep flexibility towards new IT technologies
 - Machine learning models is (since many years) intensively used
- **Use of ROOT as main data framework in the future**
 - provide tools to run optimized DSP analysis
 - communicate with cloud computing services.
- **Development of Starformer project** (transformer-based ML model)
 - collaboration with ISLab (Osaka University) for innovative architecture developments.