

# Detector technologies

## A brief overview

Many thanks to Erik Butz, [Simon Spannagel](#),  
[Freya Blekman](#), [Peter Schleper](#), Erika Garutti

[jory.sonneveld@cern.ch](mailto:jory.sonneveld@cern.ch)

Nikhef



UNIVERSITY  
OF AMSTERDAM

# Ik ben onderzoeker in natuurkunde

Ik werk aan de universiteit van Amsterdam en ontwikkel detectoren in de Nikhef detector R&D groep.

Ik heb ook detectoren bediend en data geanalyseerd: sta ik voor een deeltjesdetector waaraan gewerkt heb op CERN.

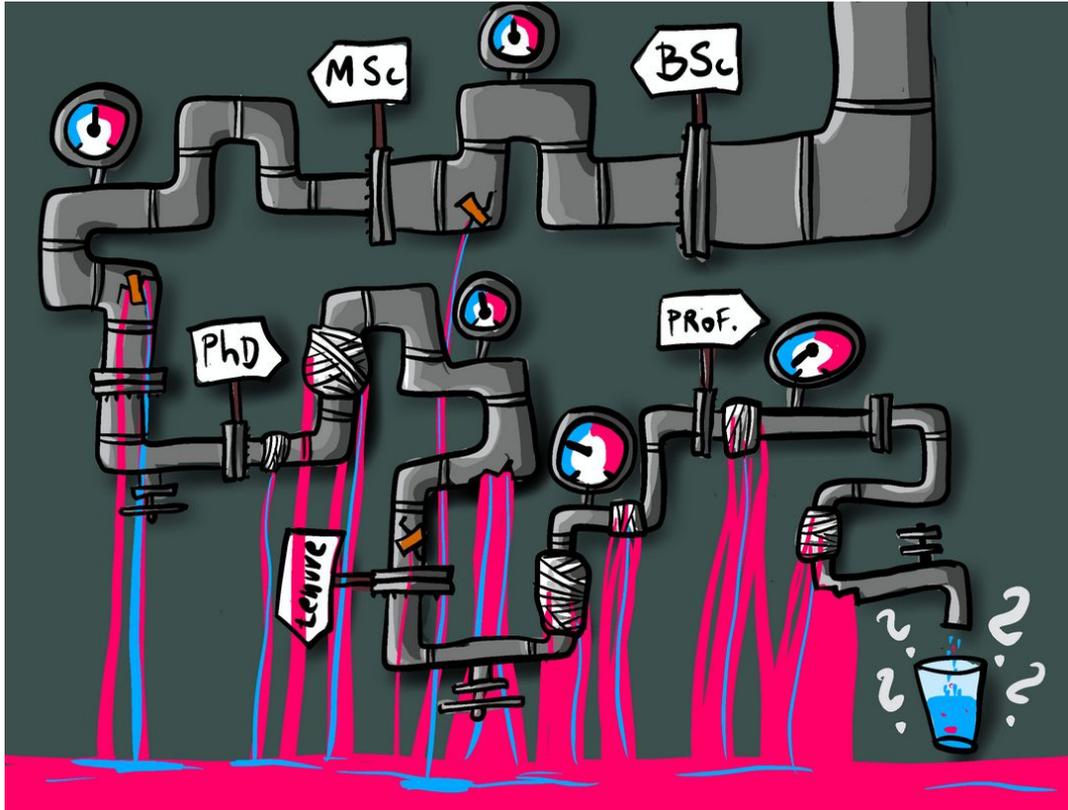
Door de buis gaan hele kleine deeltjes met bijna de lichtsnelheid!



# Diversiteit

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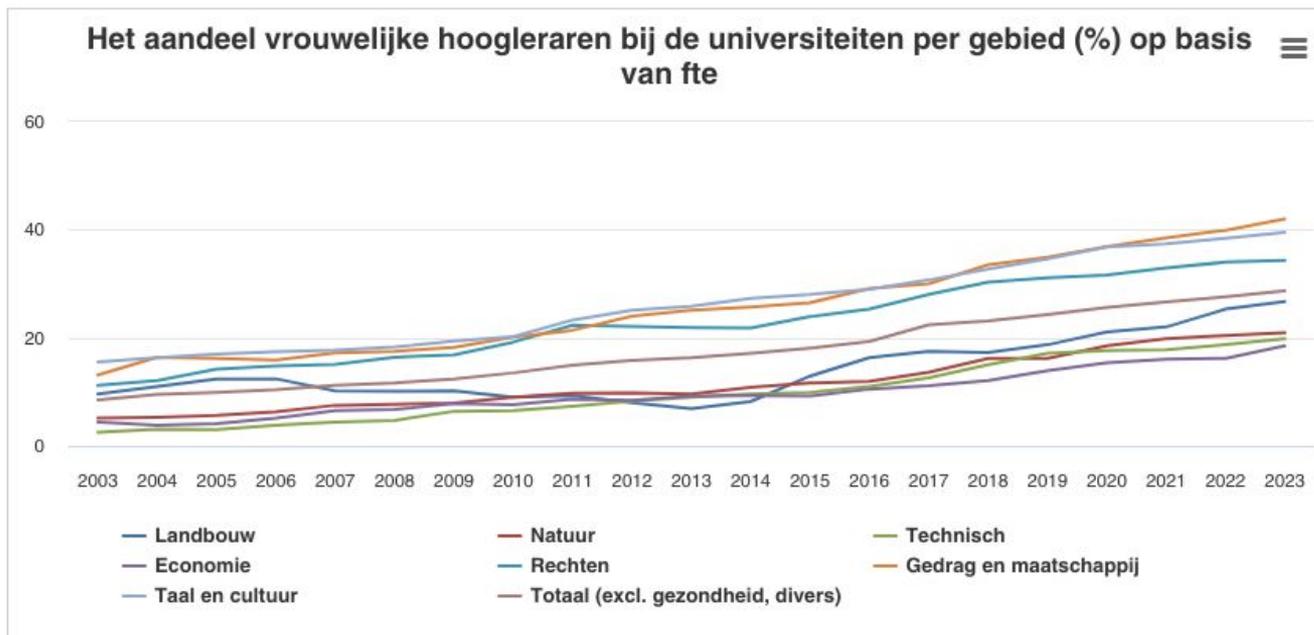
# Onderweg naar de carrière gaat veel talent verloren



Onze  
bachelorinstroom  
is maar 25%  
vrouw

<https://iop.uva.nl/people/d-i-council/posters/leaky-pipeline.html>

# Voorbeelden zijn er weinig



Van het [Rathenau instituut](#)

**Gegevens:** [Download als CSV bestand](#)

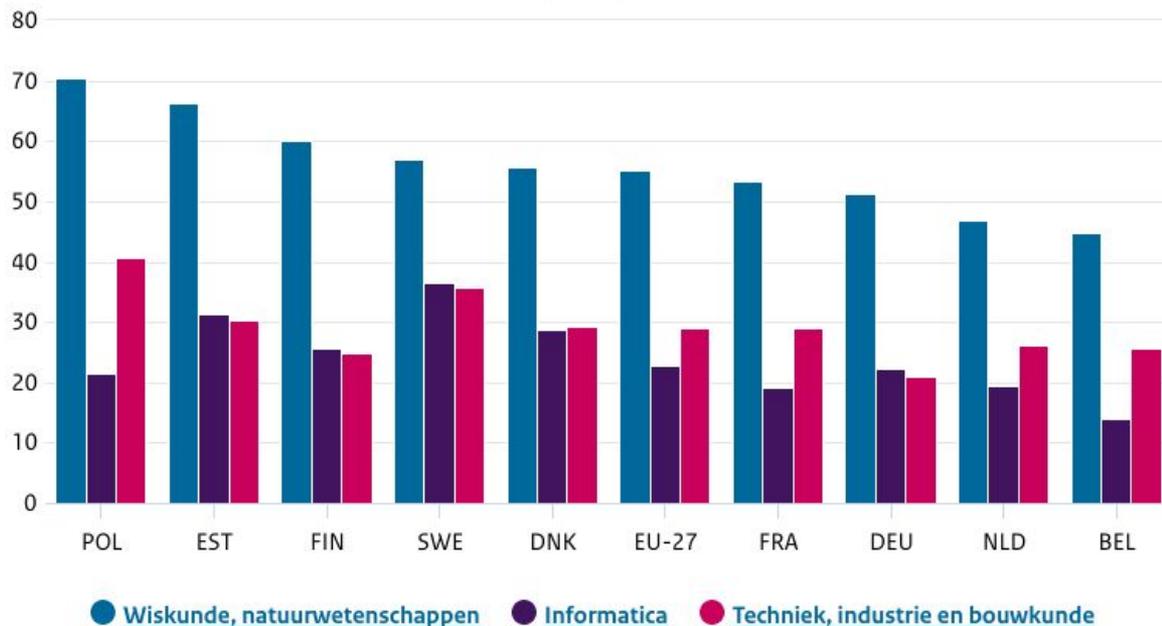
**Bron:** UNL/ WOPI

**Notities:** Exclusief Gezondheid en hoogleraren bij de universitair medische centra. Het percentage vrouwelijke hoogleraren aan de universitair medische centra is toegenomen van 17,3% in 2012 tot 25,7% in 2020. (Zie 'Extra toelichting WOPI').

# Weinig vrouwen in techniek in Nederland

Percentage gediplomeerde vrouwen in de bètatechniek per studierichting

2020-2021



Van [ocwincijfers.nl](https://ocwincijfers.nl)

# Diversiteit

Image: [kcl.ac.uk](https://www.kcl.ac.uk)



**Young girls become interested in STEM subjects around the age of 11 and then quickly lose interest when they are 15**

*Study from 2017 commissioned by Microsoft.*

## **Causes**

Social pressure

Lack of mentors

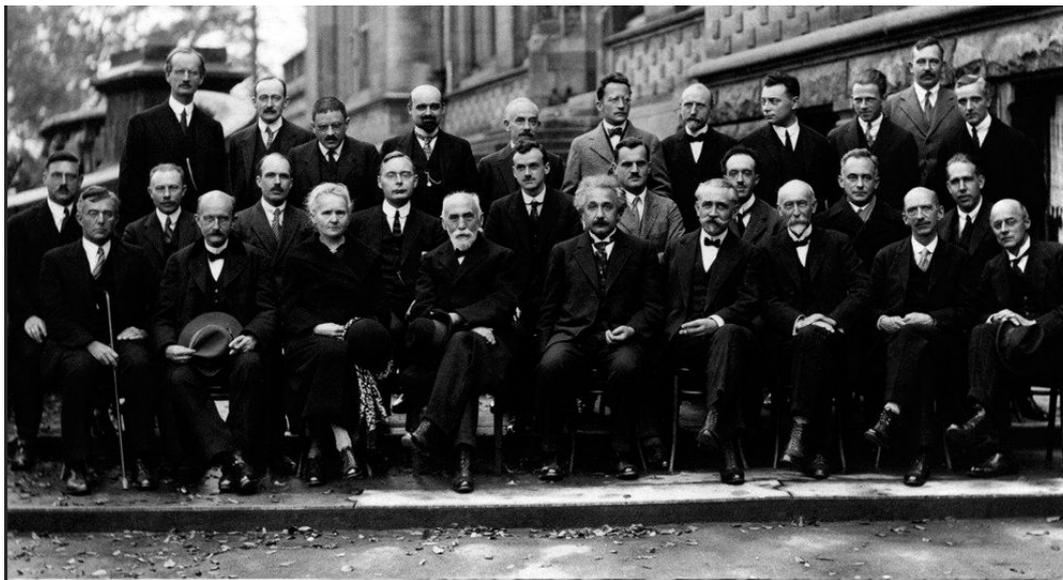
Lack of access to hands-on-learning

**Girls' confidence declines dramatically during puberty, even when they outperform boys in school.**

*A study published in 2018 asked 1,300 pupils:*

- Similar confidence up to age 12
- 30% lower confidence in girls after

# Diversiteit: waar is iedereen?



Niet-Nederlandse komaf kiest:

- economische opleiding
- universitaire rechtenstudie
- Niet 'Cultuur en maatschappij': te "soft"
- Niet 'natuur en techniek': te zwaar (!)

Van [wikipedia](#)

Vaak hebben allochtone studenten ook een grote 'drive' om te willen slagen binnen hun schoolloopbaan, omdat ze het **gevoel hebben dat ze zich meer moeten bewijzen dan autochtone studenten.**

# You can make a difference



Affirm girls' achievements and identities, especially between ages 8 and 14:

**“The praise I received from my favourite teachers when I was in middle school has stayed with me throughout college. [...] Sometimes all it takes is one teacher affirming you—especially when your cultural identities are not being affirmed anywhere else.”**

Van Tigdankay Saccoh in de [Economist](#)

## **What can you do?**

Emphasize importance of trial-and-error

Affirm achievements

Deeltjes: waar komen ze vandaan?

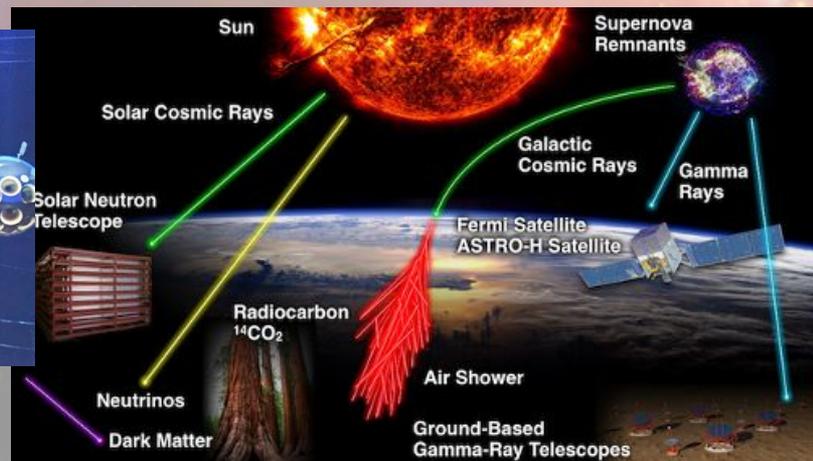
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# Deeltjes vanuit de ruimte

10000 keer per seconde gaan er deeltjes uit kosmische stralen door jou heen

Wat zijn die deeltjes en hoe gedragen ze zich?  
Waarvan zijn wij en het universum gemaakt?

From <https://www.km3net.org/>



**KM3net: cubic kilometer neutrino telescope**  
in Mediterranean Sea, can detect photons  
kan gamma stralen detecteren

Hoe kunnen we zo'n deeltje  
detecteren?

---

$\gamma$  stralen: fotonen

### Cherenkov licht:

- Licht is 0.25% langzamer in water
- Ultra-hoge-energie deeltjes kunnen sneller dan licht in water
- Een blauwe flash van "Cherenkov licht" ontstaat
- Net als een sonic boom van een vliegtuig dat door de geluidsbarriere gaat

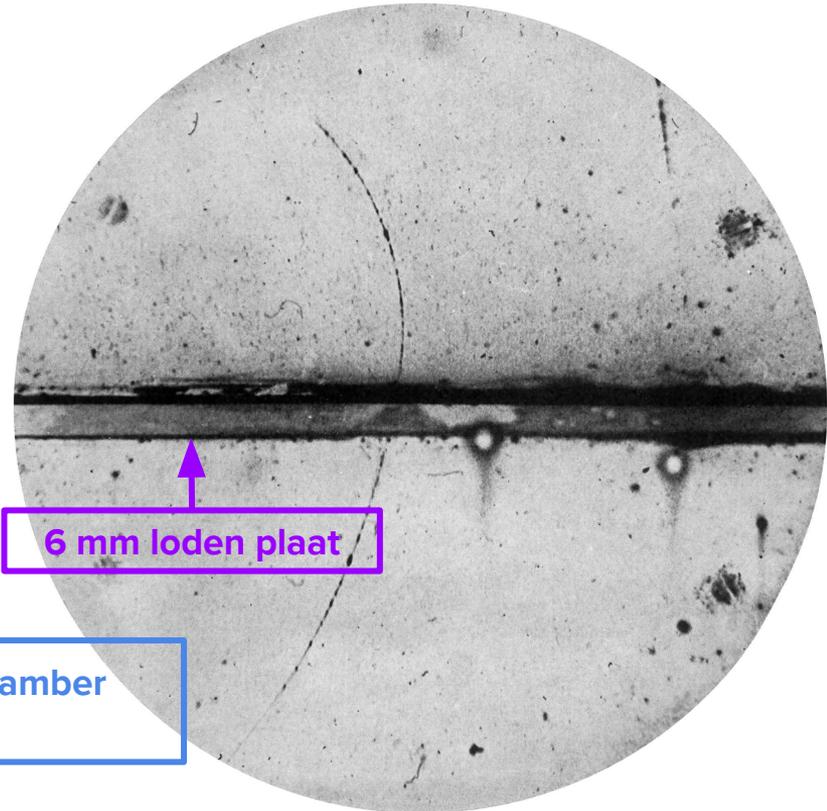
# Ontdekking van antimaterie

<https://upload.wikimedia.org/wikipedia/commons/6/69/PositronDiscovery.jpg>

C.D. Anderson <https://journals.aps.org/pr/pdf/10.1103/PhysRev.43.491>

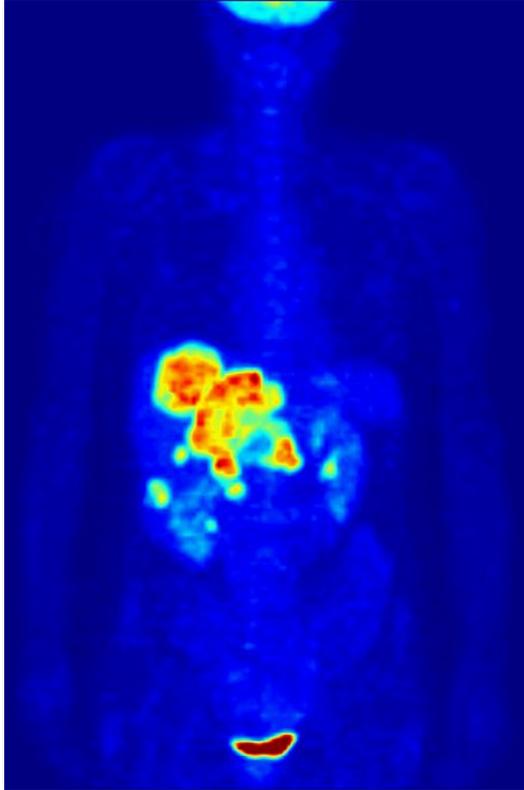
Het eerste positron ooit geobserveerd!

**Nevelkamer:** gasmengsel van superverzadigd water of alcohol. Een energetisch deeltje ioniseert het gas en ionen vormen condensatiecentra die zichtbaar worden als wolk.



15000 Gauss = 1.5T magnetisch veld Wilson cloud chamber (nevelkamer) voor detectie van kosmische straling

# Meer deeltjesdetectoren



Magnetic resonance imaging

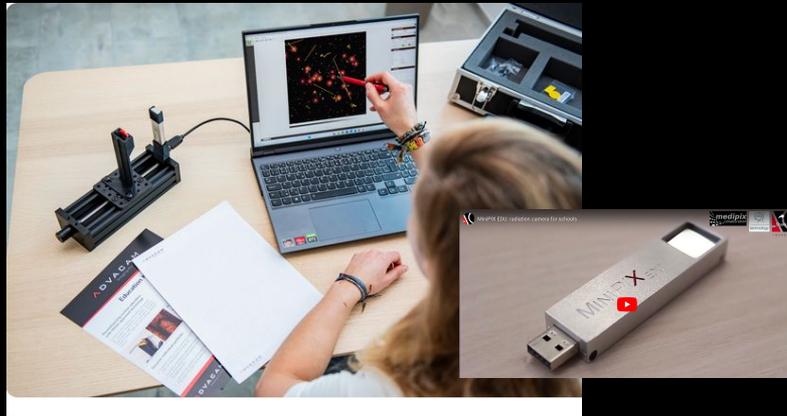
(b)

Positron emission tomography



# Detectie van deeltjes

Deeltjes laten sporen achter!

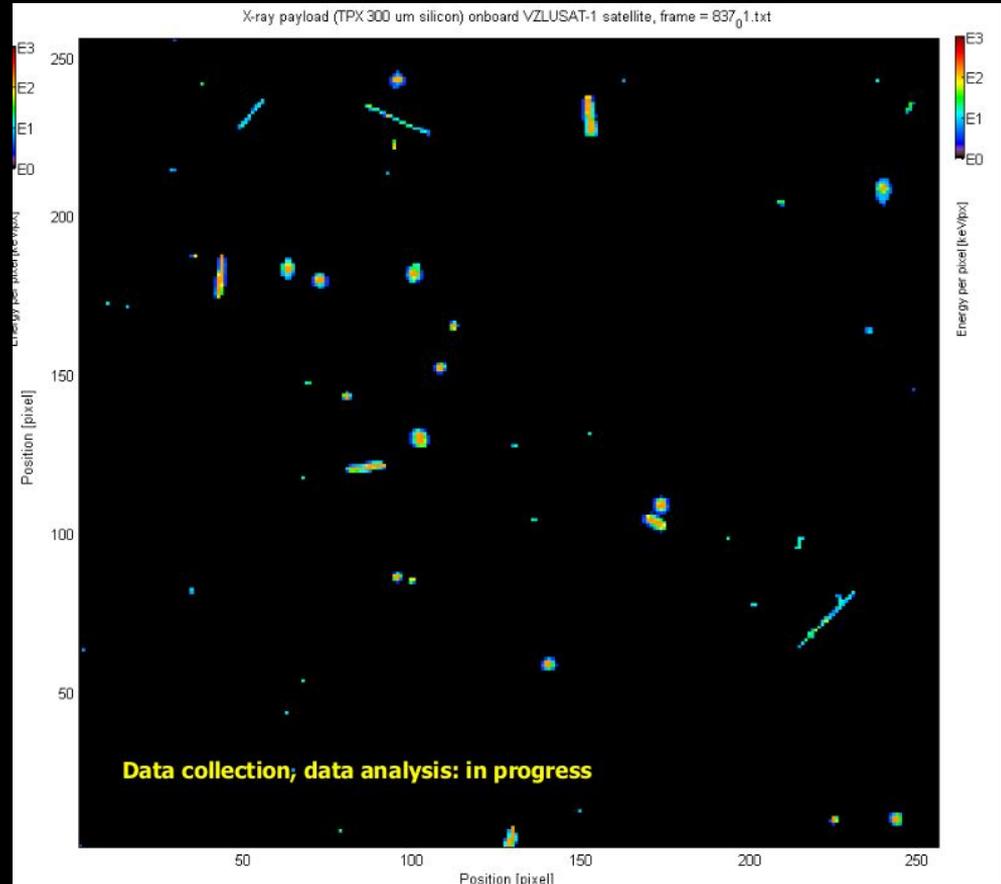


See also

<https://advacam.com/camera/edu-kit/>

or no need for the whole package:

<https://advacam.com/camera/minipix-edu/>



# Build your own muon detector

Silicon photomultiplier



Alle producten Fabrikanten Services

Start > Semiconductors > ICs > IC Sensors > Silicon Photomultiplier

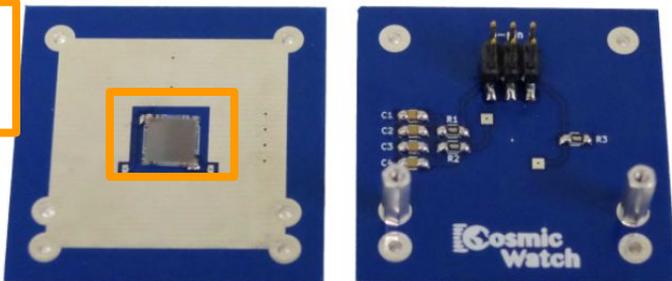
**MICROFC-60035-SMT-TR1**

Silicon Photomultiplier (SiPM), C-Series, 6mmx6mm

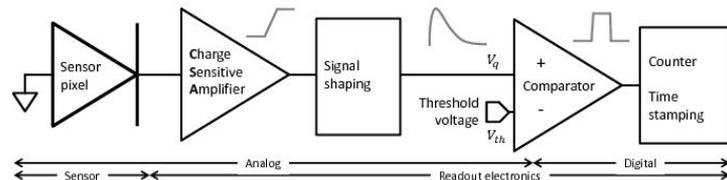
Download Code



[Cosmic Watch Detector](#)



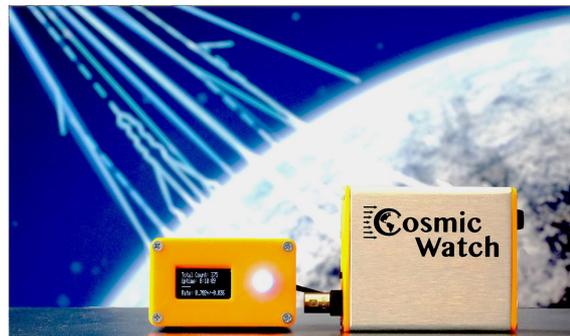
With a scintillator (plastic) and a silicon photomultiplier you can build your own muon detector!



Readout: arduino



scintillator



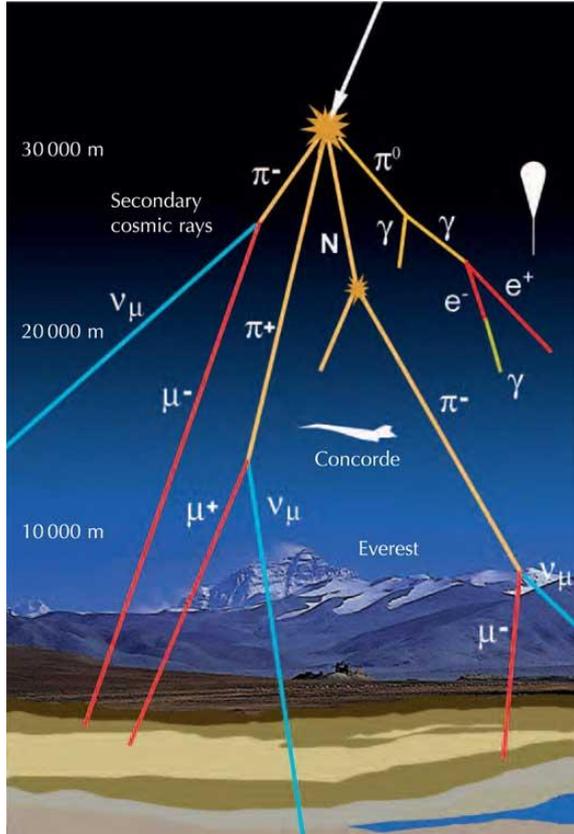
Figures from [here](#) and [here](#)

Wat voor deeltjes?

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# Vele verschillende deeltjes!

Waar is het proton?



## Standard Model of Elementary Particles

	three generations of matter (fermions)			interactions / force carriers (bosons)	
	I	II	III		
mass	$\approx 2.2 \text{ MeV}/c^2$	$\approx 1.28 \text{ GeV}/c^2$	$\approx 173.1 \text{ GeV}/c^2$	0	$\approx 124.97 \text{ GeV}/c^2$
charge	$\frac{2}{3}$	$\frac{2}{3}$	$\frac{2}{3}$	0	0
spin	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1	0
	<b>u</b> up	<b>c</b> charm	<b>t</b> top	<b>g</b> gluon	<b>H</b> higgs
	<b>d</b> down	<b>s</b> strange	<b>b</b> bottom	<b><math>\gamma</math></b> photon	
	<b>e</b> electron	<b><math>\mu</math></b> muon	<b><math>\tau</math></b> tau	<b>Z</b> Z boson	
	<b><math>\nu_e</math></b> electron neutrino	<b><math>\nu_\mu</math></b> muon neutrino	<b><math>\nu_\tau</math></b> tau neutrino	<b>W</b> W boson	

**QUARKS** (left side)

**LEPTONS** (left side)

**GAUGE BOSONS VECTOR BOSONS** (bottom)

**SCALAR BOSONS** (right side)

Ieder deeltje gedraagt zich anders, afhankelijk van massa, lading. Op aarde zien we veel **muonen**.



Hoe detecteer je zo'n deeltje op  
CERN?

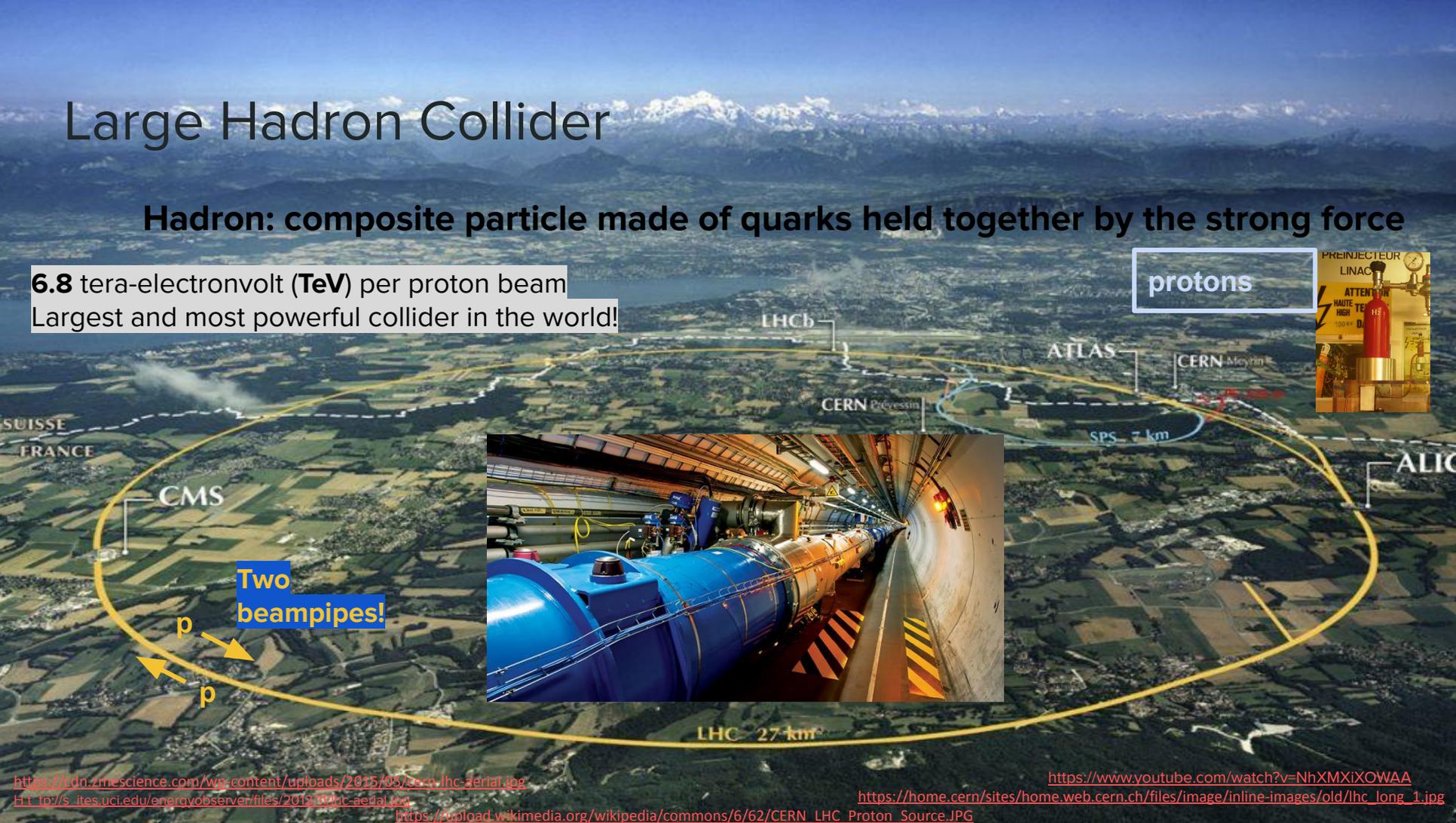
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# Large Hadron Collider

**Hadron: composite particle made of quarks held together by the strong force**

**6.8** tera-electronvolt (**TeV**) per proton beam  
Largest and most powerful collider in the world!

protons



Two beampipes!



LHC 27 km

[https://cdn.zmescience.com/wp-content/uploads/2015/05/cern\\_lhc-aerial.jpg](https://cdn.zmescience.com/wp-content/uploads/2015/05/cern_lhc-aerial.jpg)

<http://sites.uci.edu/energyobserver/files/2012/10/lhc-aerial.jpg>

[https://upload.wikimedia.org/wikipedia/commons/6/62/CERN\\_LHC\\_Proton\\_Source.JPG](https://upload.wikimedia.org/wikipedia/commons/6/62/CERN_LHC_Proton_Source.JPG)

<https://www.youtube.com/watch?v=NhXMXiXOWAA>

[https://home.cern/sites/home.web.cern.ch/files/image/inline-images/old/lhc\\_long\\_1.jpg](https://home.cern/sites/home.web.cern.ch/files/image/inline-images/old/lhc_long_1.jpg)

# Large Hadron Collider

6.8 tera-electron volt (TeV) per proton beam

2x2800 bunches of protons 25 ns apart

eV = energy to move an electron through 1 V:  $1 \text{ eV} = 1.602 \times 10^{-19} \text{ J}$

MeV =  $10^6 \text{ eV}$

GeV =  $10^9 \text{ eV}$

TeV =  $10^{12} \text{ eV}$

PeV =  $10^{15} \text{ eV}$

protons

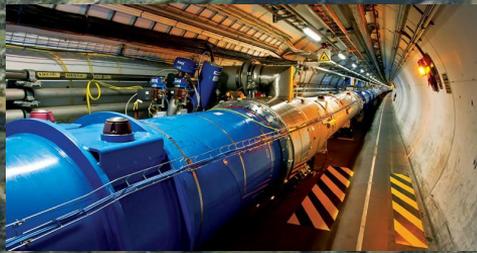
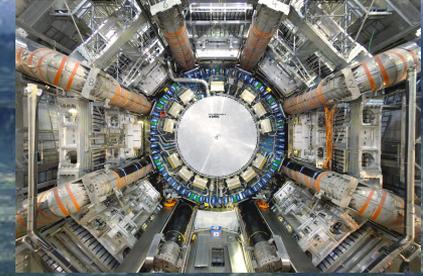
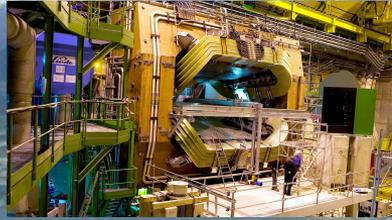
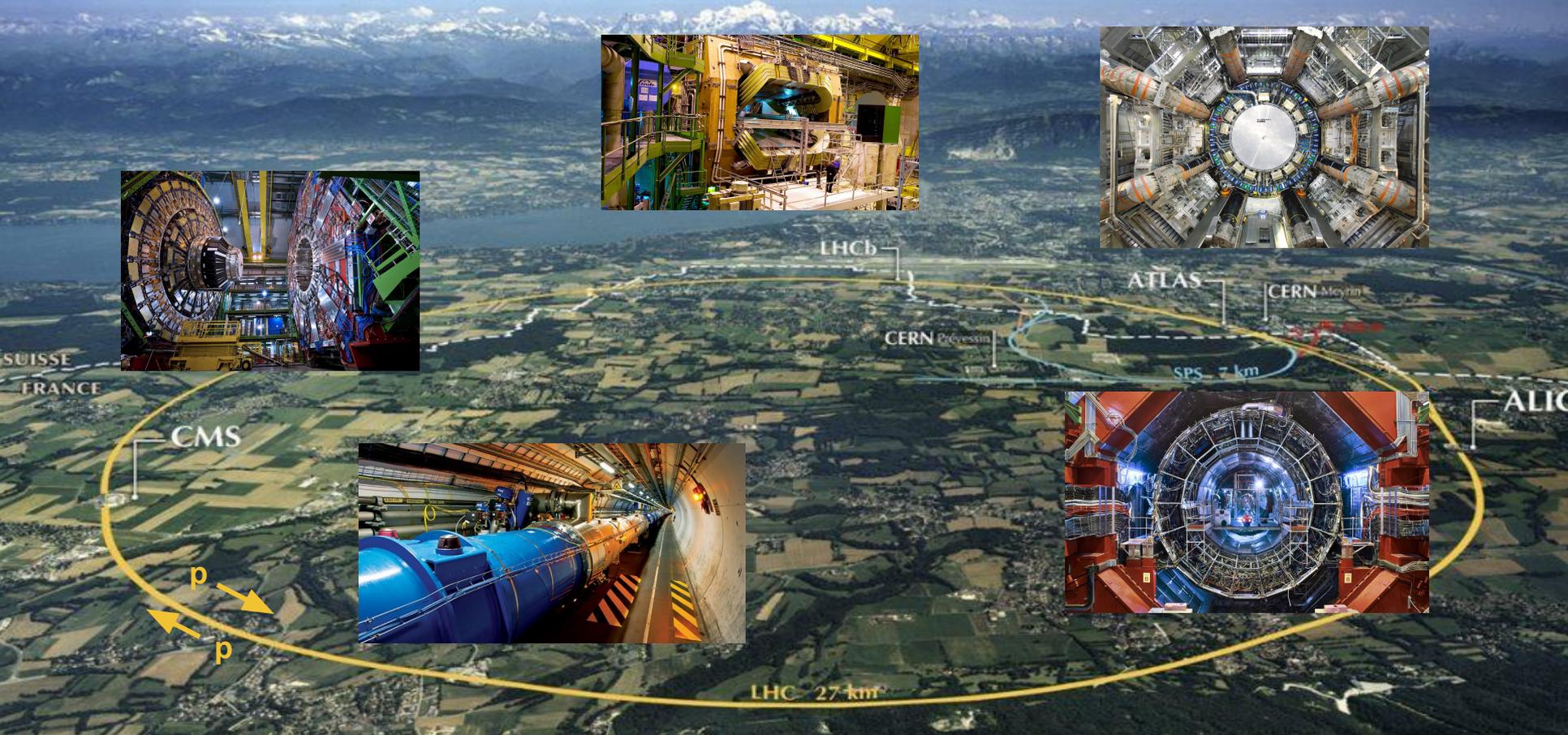


Two  
beampipes!

p  
p

LHC 27 km

# Large Hadron Collider

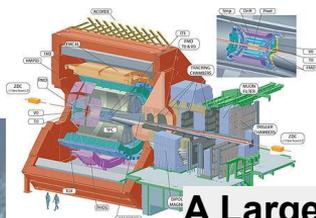
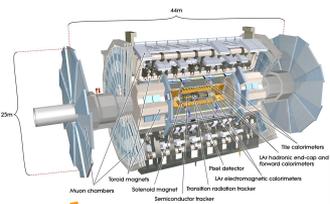


# LHC magneten: Kouder dan het universum!

Naar 1.8 K zodat de magneten (in het blauwe omhulsel) supergeleidend worden

Het universum is warmer: 2.73 kelvin!

# Detectors at the LHC

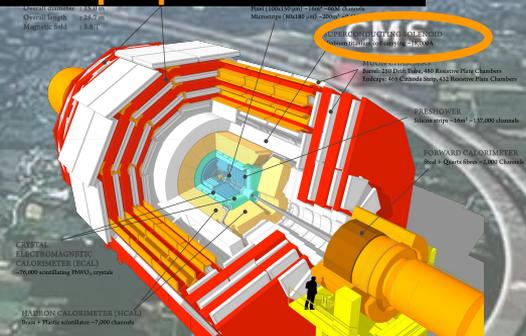


**A Large Ion Collider Experiment:** specialized in heavy ion collisions and quark-gluon plasma: fraction of second after big bang!

## Compact Muon Solenoid

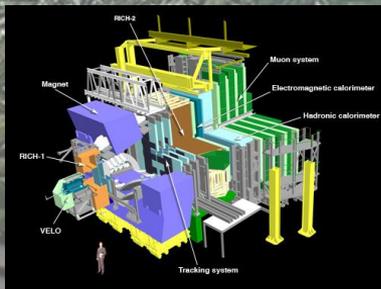
14000 tons: 1.5\* Eiffel tower weight, half the size of ATLAS: 15 m x 15 m x 21 m very compact!  
Largest superconducting solenoid magnet ever made

## Why a second multipurpose detector?



## A Toroidal LHC Apparatus:

25 m x 25 m x 46m  
The inner detector has 3 air core toroidal magnets and one solenoidal magnet.  
**Multipurpose detector.**



## LHC beauty:

A single-arm forward spectrometer designed for the study of particles containing b or c quarks.

**Other detectors:** MoEDAL, TOTEM, LHCforward, Faser, SND

Wat detecteren we?

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# Niet alle bekende elementaire deeltjes!

Directly detect:

Vervalsproducten

jets

Indirectly detect:

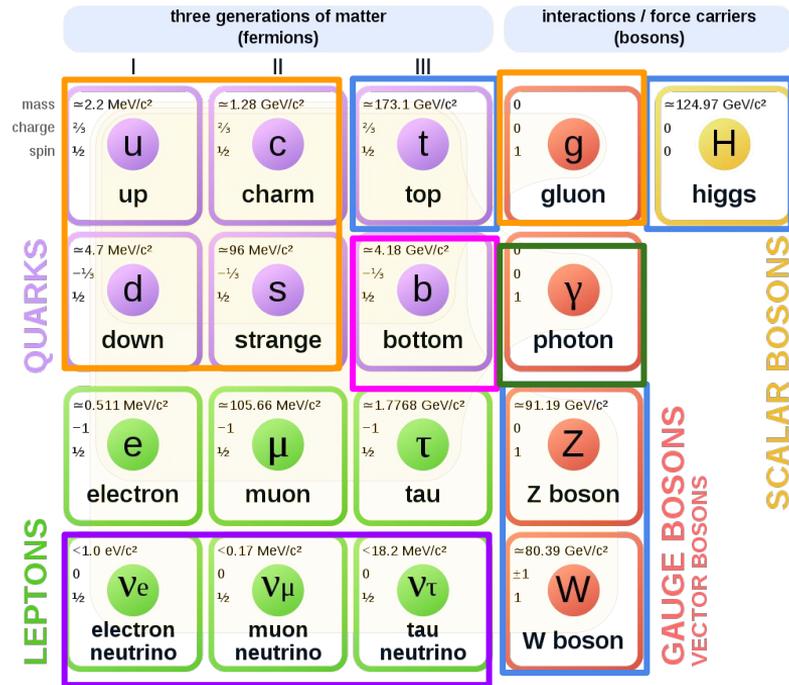
Missing energy

Secondary vertex + jets

Neutral particles

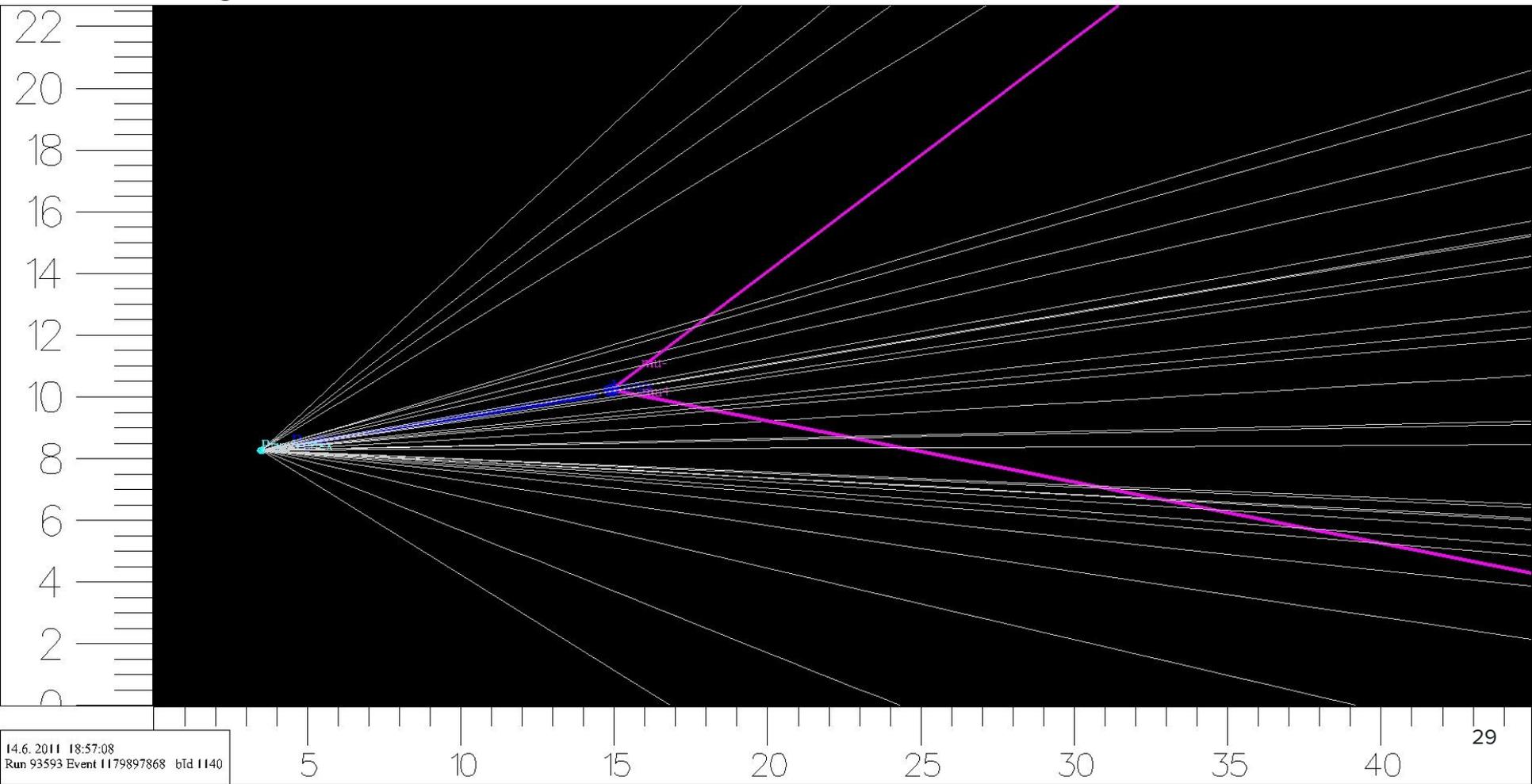
Should be able to detect and identify:  
 $e^\pm, \mu^\pm, \gamma, \pi^\pm, K^\pm, p^\pm, K^0, n$   
 using mass, charge, interaction

## Standard Model of Elementary Particles



# LHCb $B_s \rightarrow \mu^+ \mu^-$

<http://lhcb-public.web.cern.ch/lhcb-public/>



# What do we measure and how?

Observable	Measurable quantity
Momentum (p)	Bending radius in magnetic field
Speed (v)	Time of flight, Cherenkov radiation
Charge (Q)	Bending in magnetic field
Lifetime ( $\tau$ )	Distance traveled before decay
Energy (E)	Absorption in calorimeters
Mass (m)	Indirectly from momentum
Spin	<u><a href="#">Angular distributions</a></u>

- $d = c\tau\gamma$
- $\gamma = 1/\sqrt{1-\beta^2}$
- $\beta = v/c$
- $E^2 = m^2c^4 + p^2c^2$
- $p = \gamma mv = mv/\sqrt{1-v^2/c^2}$

For some examples of measuring spin see

<https://arxiv.org/pdf/1202.6660.pdf> and  
<http://moriond.in2p3.fr/QCD/2013/proceedings/Muehleitner.pdf>

**Need 1) a magnetic field and 2) interaction with material**

# Hoe detecteren we?

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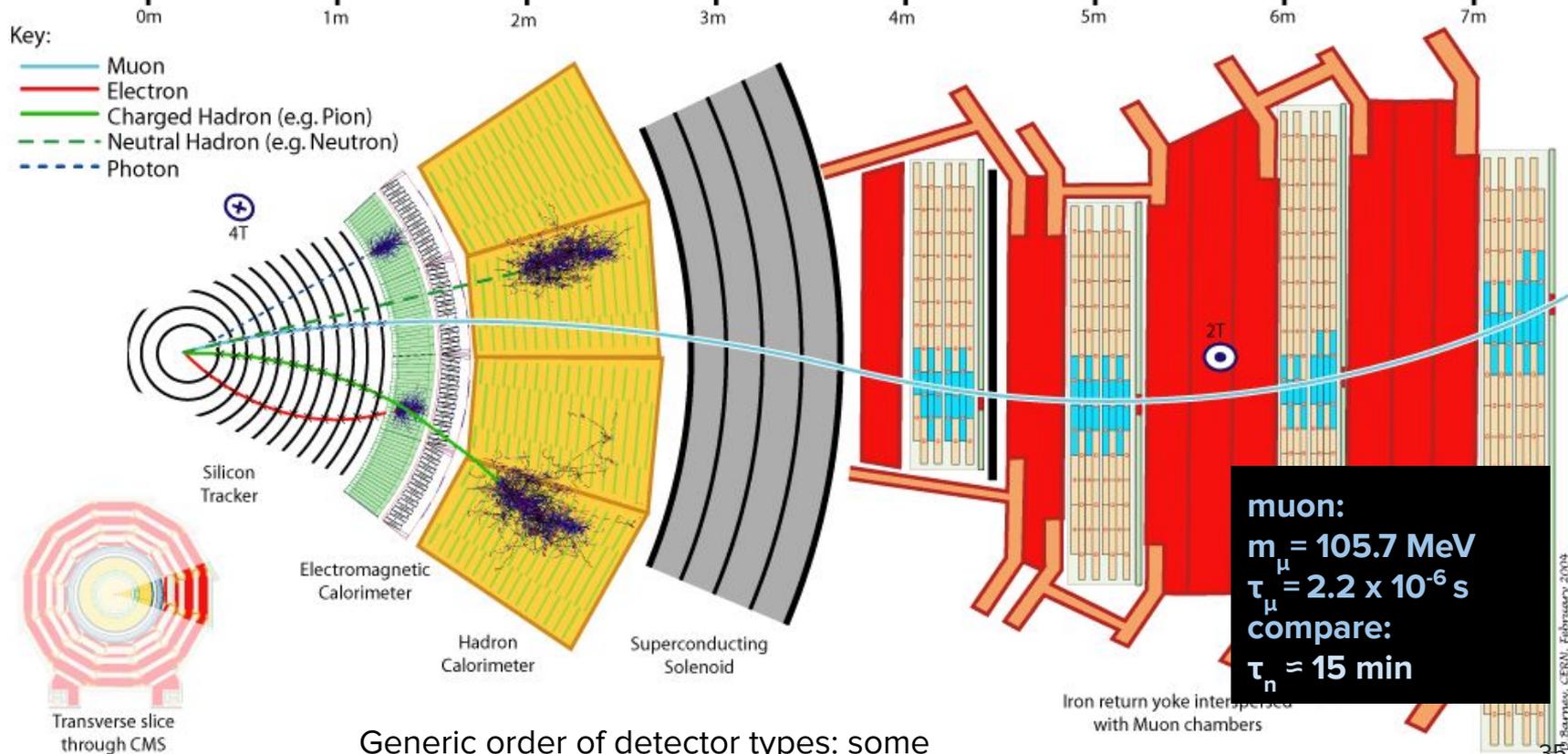
## Compact Muon Solenoid

100 meter onder de grond is een holte voor experimentatie, van de grootte van een kathedraal, die 1 van de twee experimenten herbergt die het Higgs boson heeft gevonden: CMS, een detector van 14000 ton !



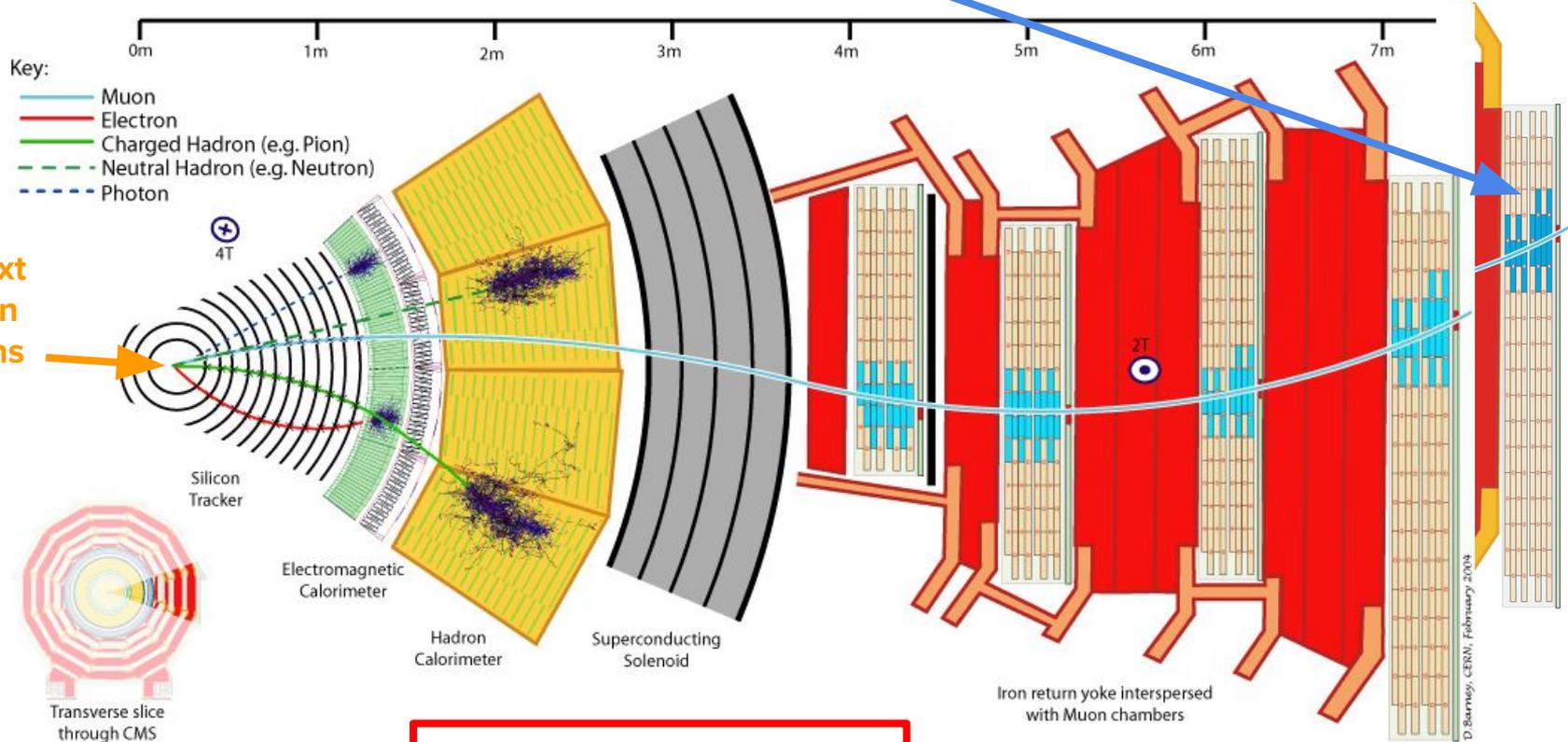


# Detectors at the large hadron collider: onion-like



Generic order of detector types: some measurements destructive!

# Note when the muon arrives here



The next collision happens here:

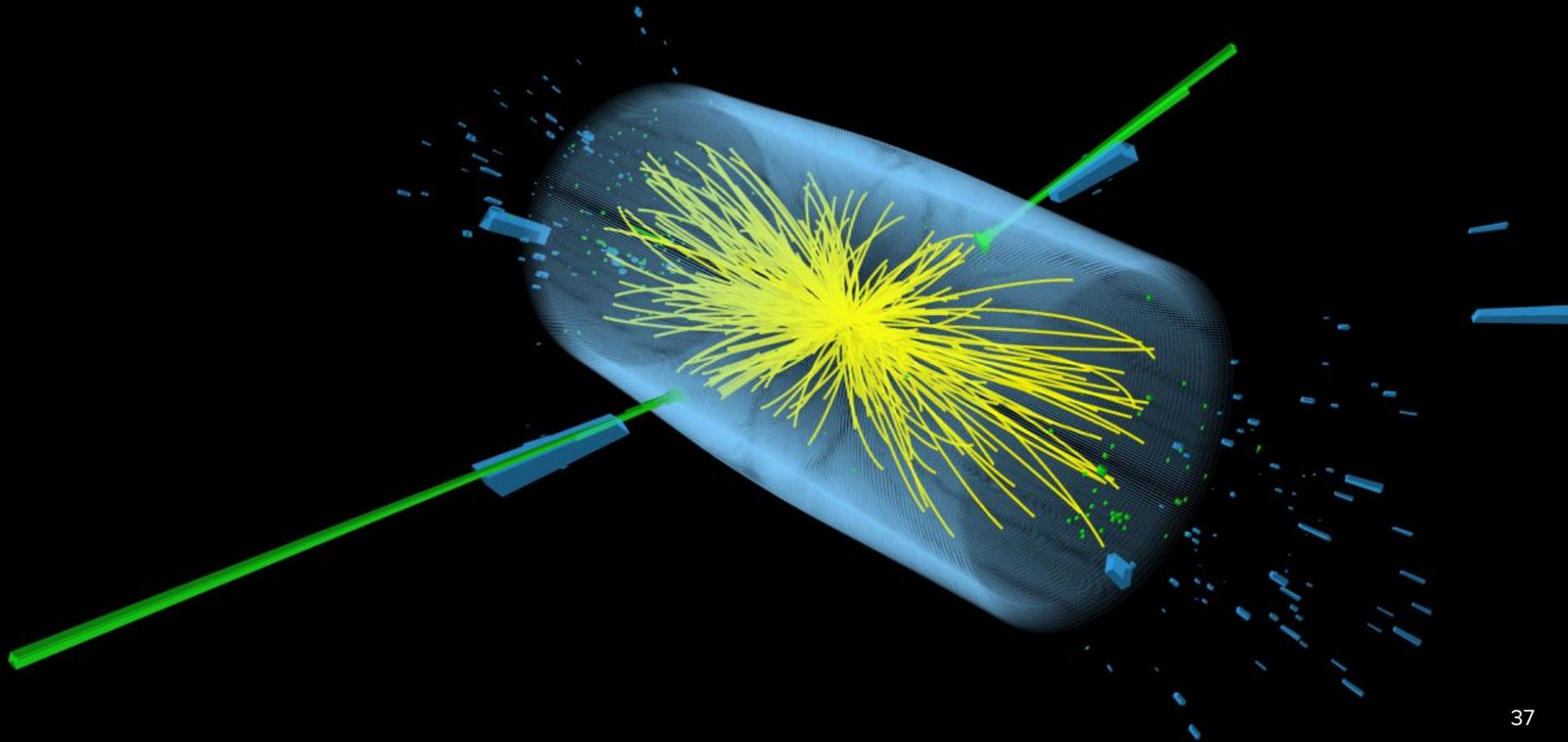
$$25 \text{ ns} \cdot c \approx 7.5 \text{ m}$$



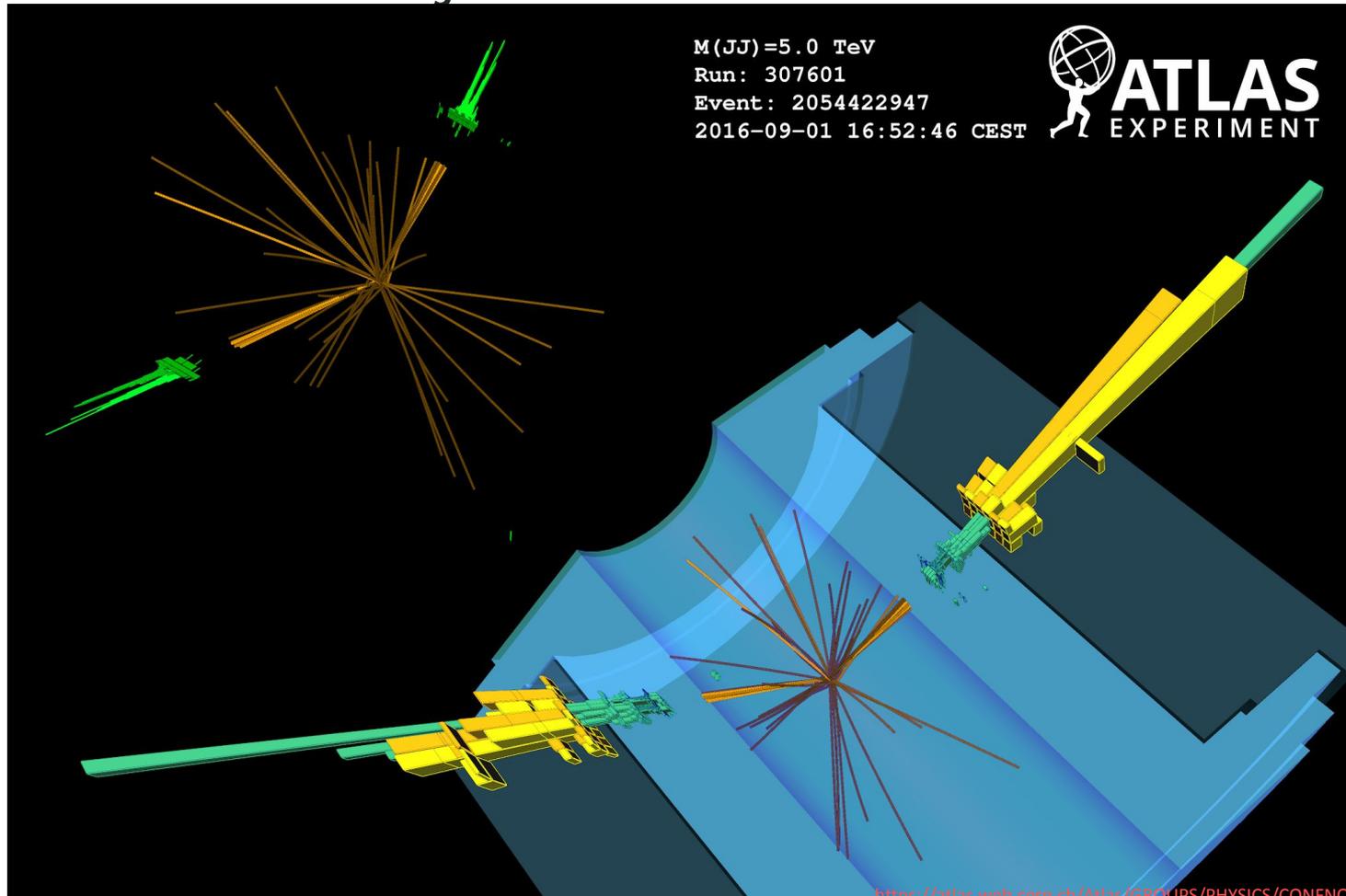
CMS Experiment at the LHC, CERN

Data recorded: 2016-May-11 21:40:47.974592 GMT

Run / Event / LS: 273158 / 238962455 / 150



# Diboson event: jets in the ATLAS detector



# Zware ionen

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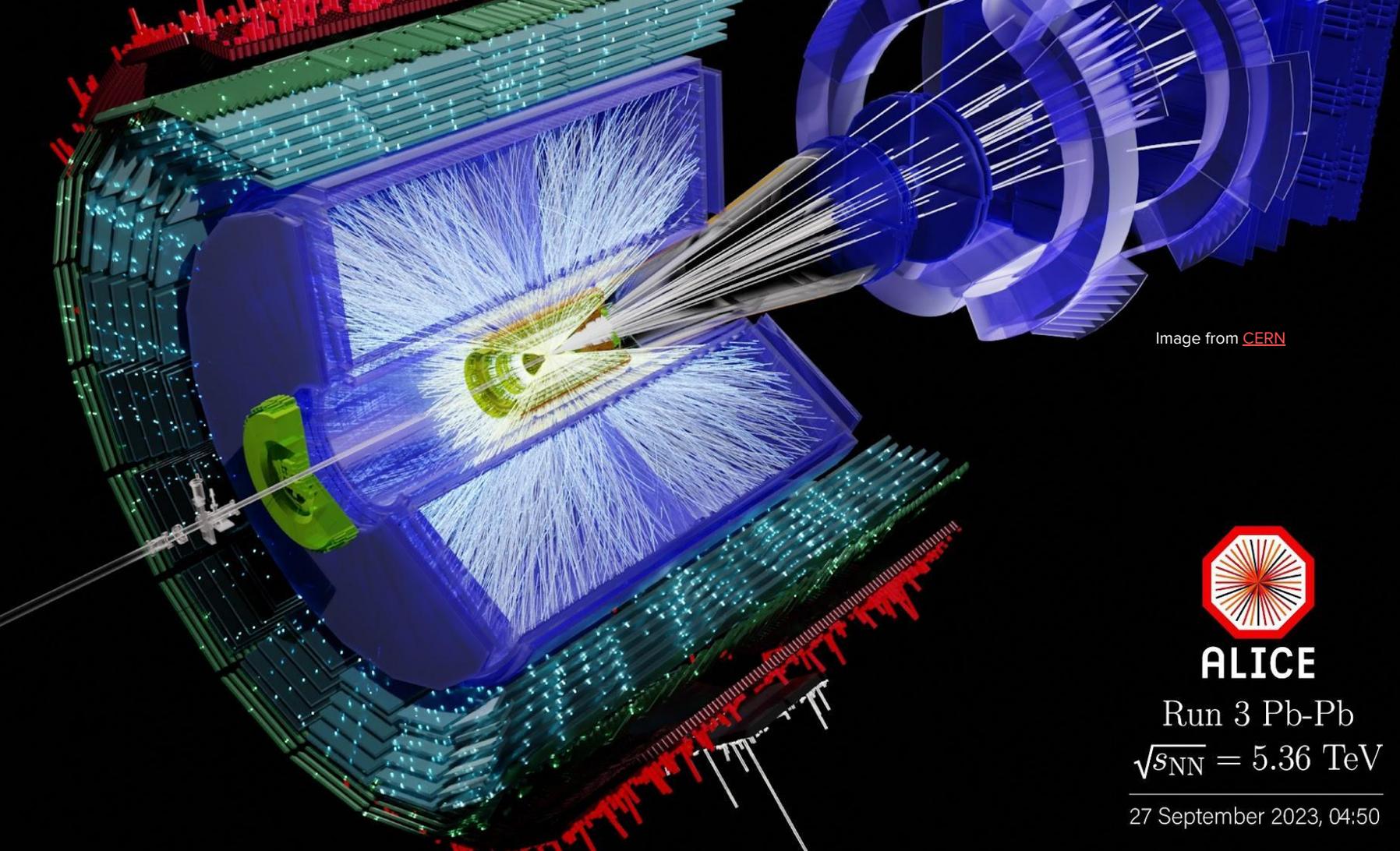


Image from [CERN](#)



**ALICE**

Run 3 Pb-Pb

$\sqrt{s_{NN}} = 5.36 \text{ TeV}$

27 September 2023, 04:50

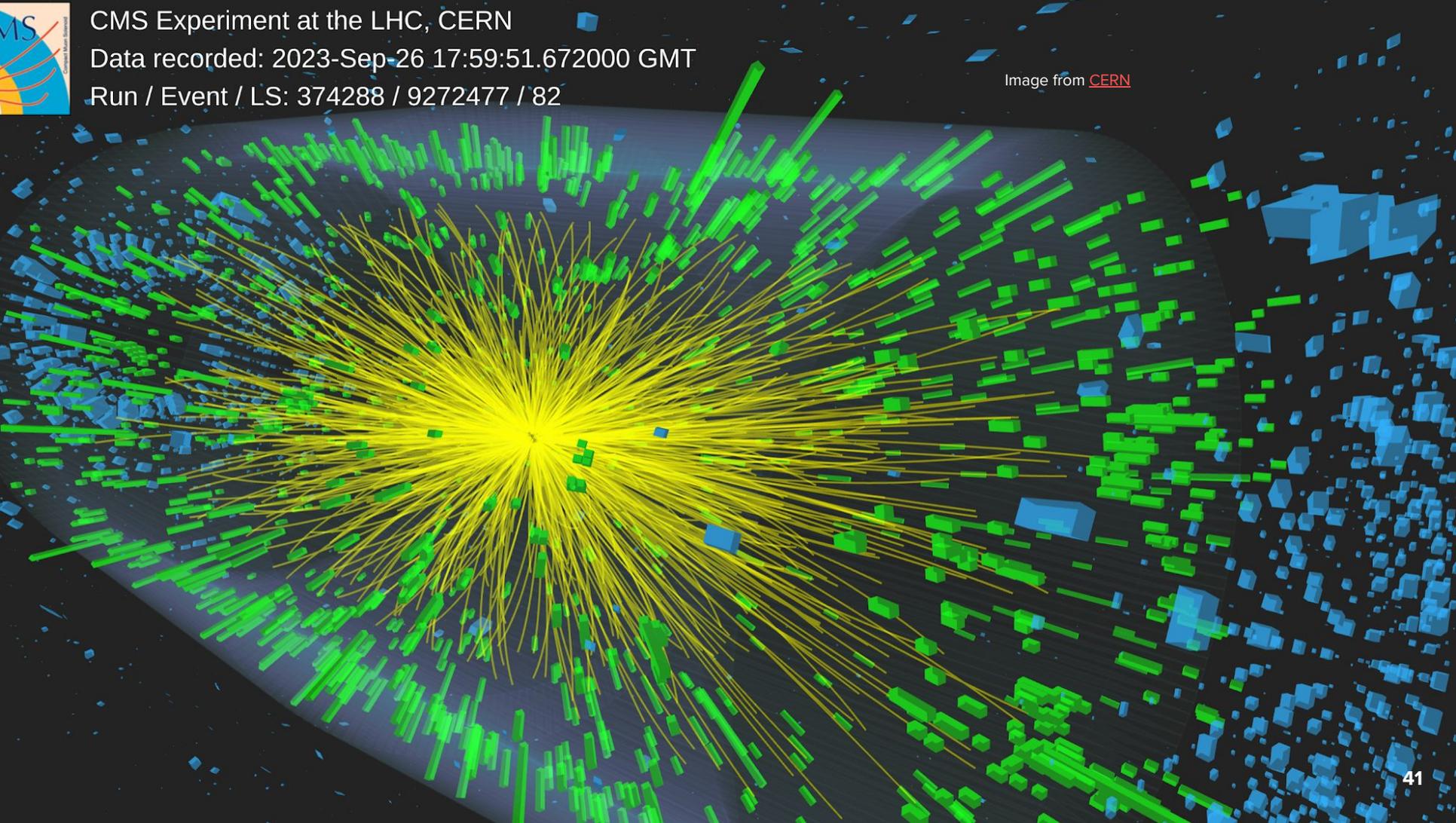


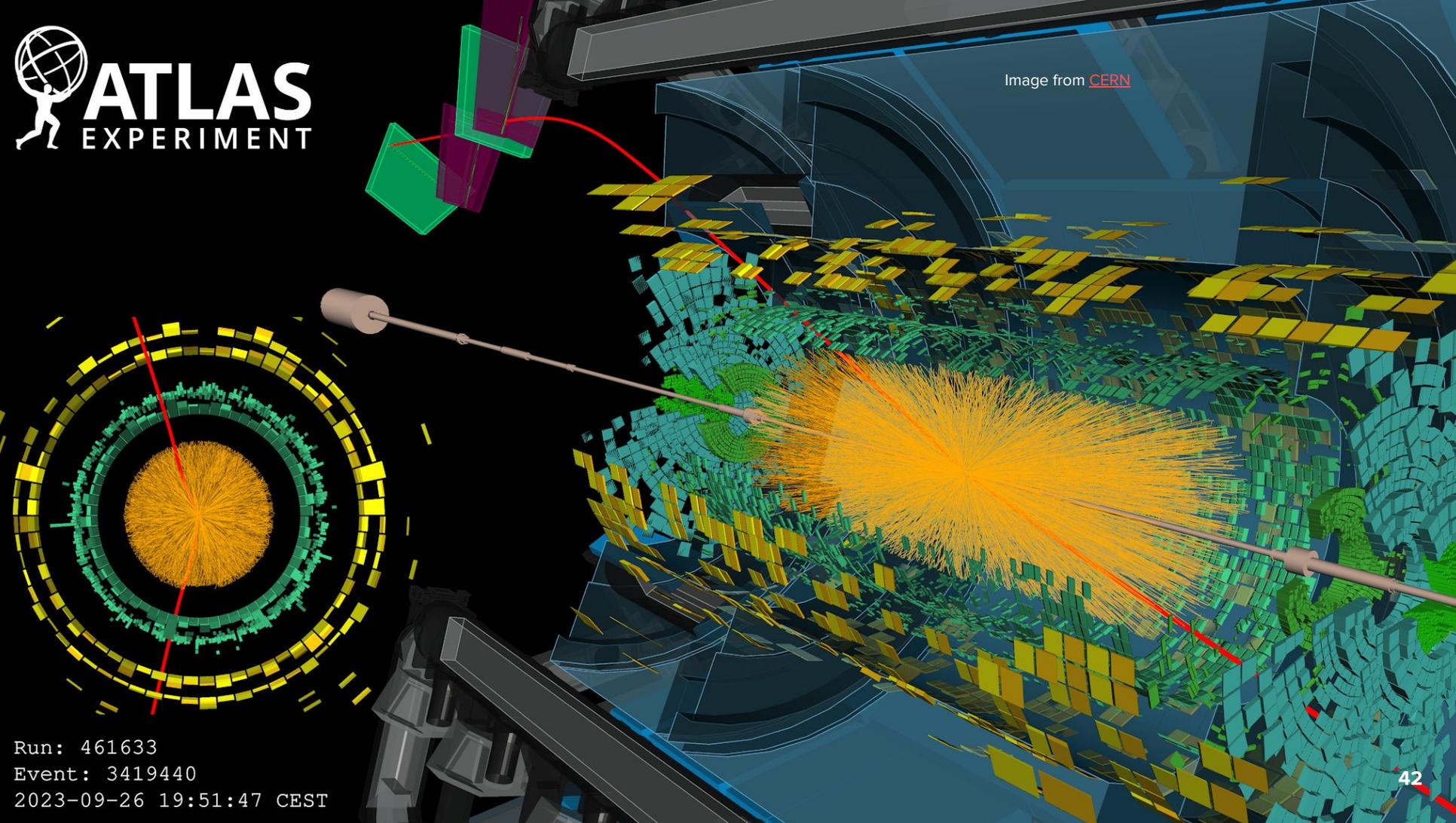
CMS Experiment at the LHC, CERN

Data recorded: 2023-Sep-26 17:59:51.672000 GMT

Run / Event / LS: 374288 / 9272477 / 82

Image from [CERN](#)



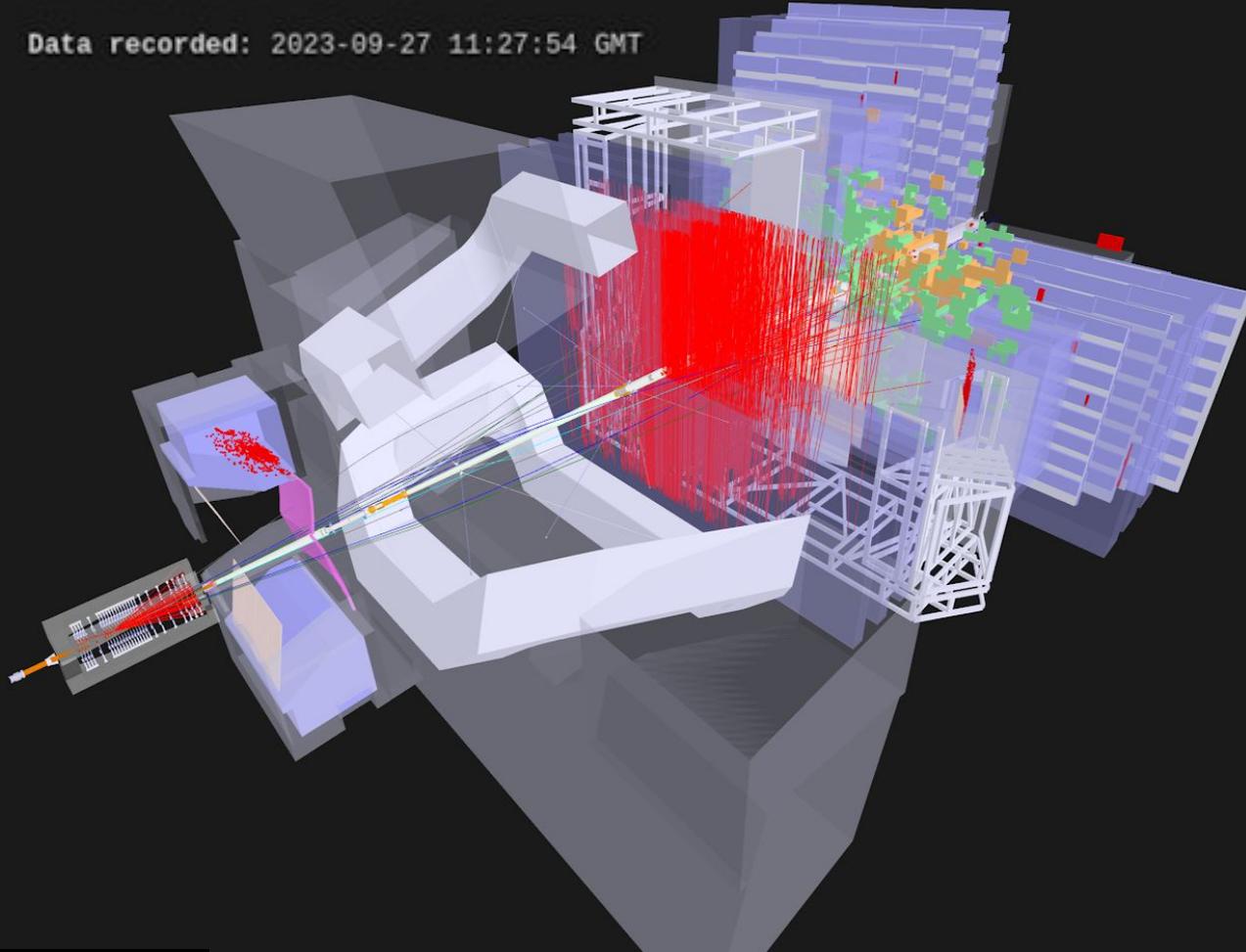


LHCb Experiment at CERN

Run / Event: 277491 / 5197775

Data recorded: 2023-09-27 11:27:54 GMT

Image from [CERN](#)



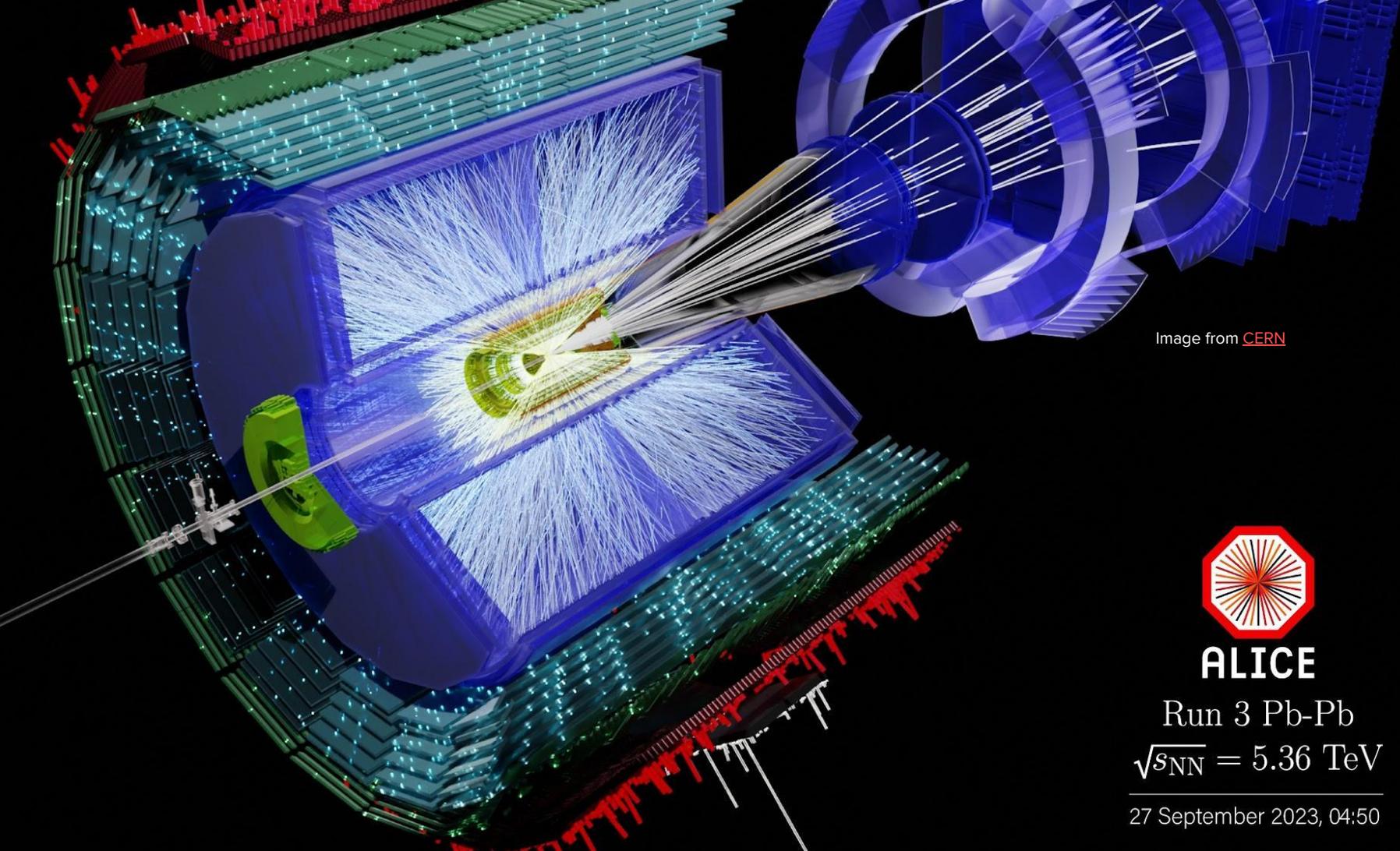


Image from [CERN](#)



**ALICE**

Run 3 Pb-Pb

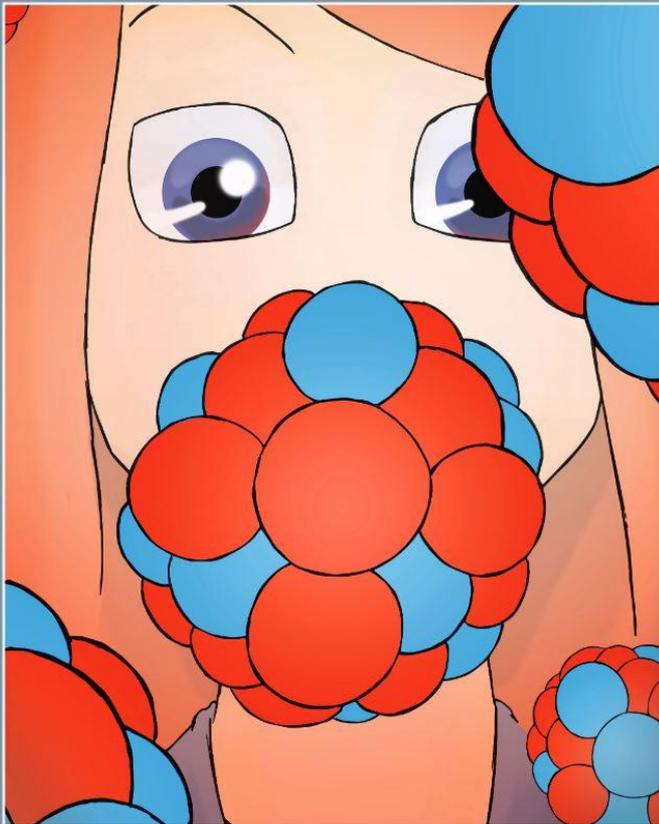
$\sqrt{s_{NN}} = 5.36 \text{ TeV}$

27 September 2023, 04:50

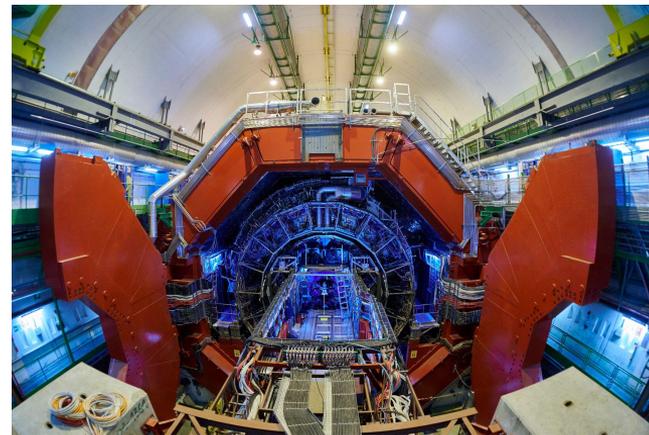
Who is ALICE?

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# ALICE



## The ALICE detector: heavy ions



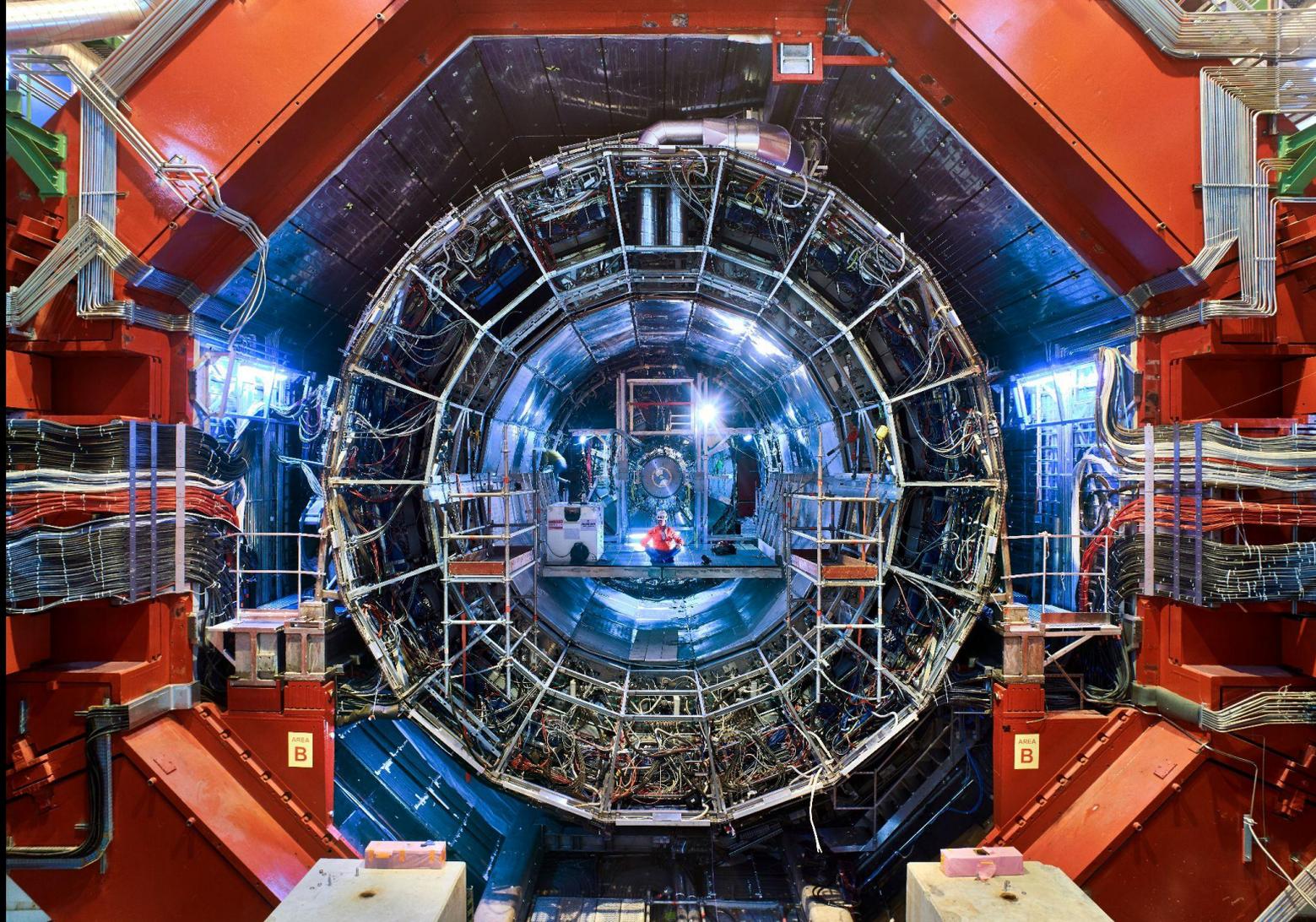
From

<https://alice.cern/brochures>

<https://cds.cern.ch/record/2809617>

From

<https://cds.cern.ch/record/2665476>



# A familiar diagram

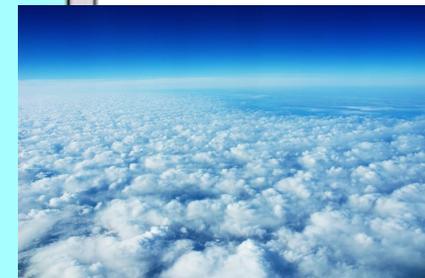
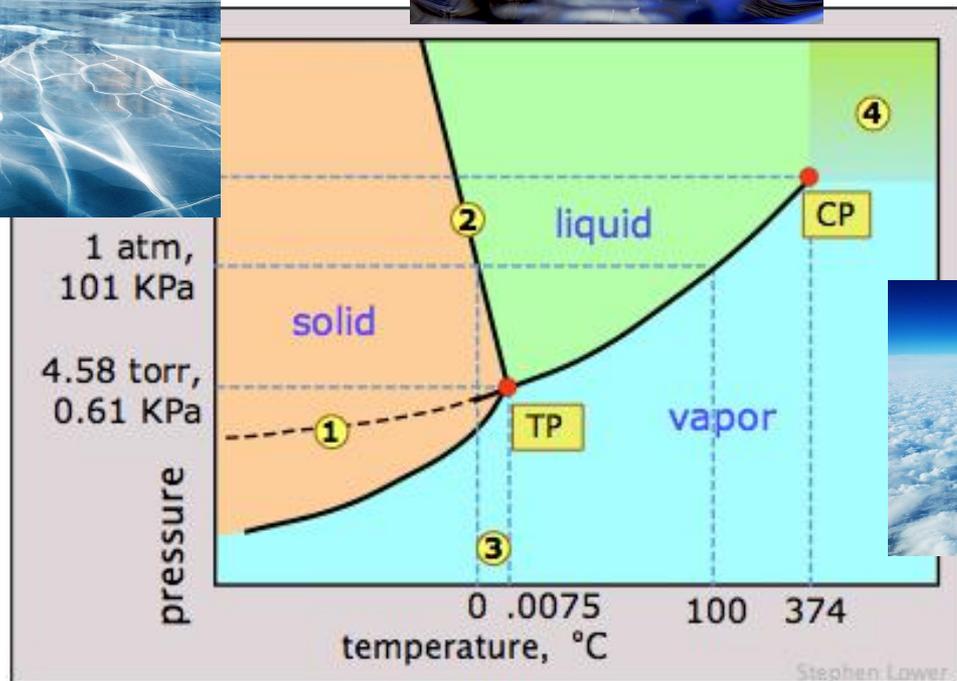


From <https://www.mindful.com/learn/ice-101>

From <https://en.wikipedia.org/wiki/Water>

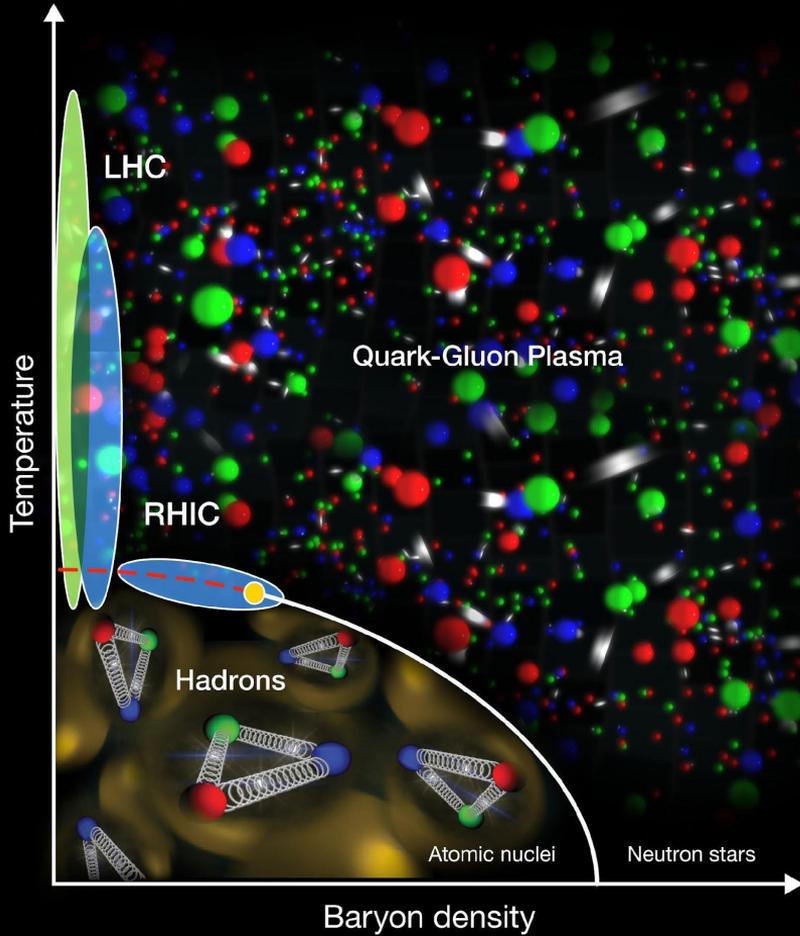


Matter comes in many different phases

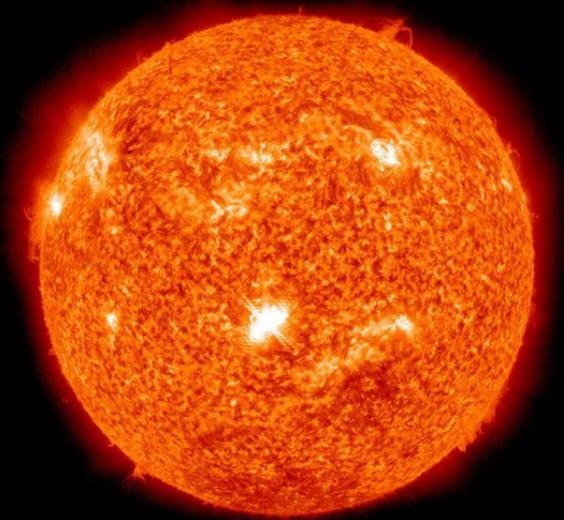


From <https://research.vias.org/jrc/2014/05/14/2014051400010109/Water-100-cm-in-the-upper-atmosphere-a-meteorological-mystery-09-09-2014>

# Quantum chromodynamics Phase diagram



Quark gluon plasma: more than 500,000 times hotter than the core of the sun!



# A glimpse into the start of the universe

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# History of the Universe

Image: [Particle Data Group](#)

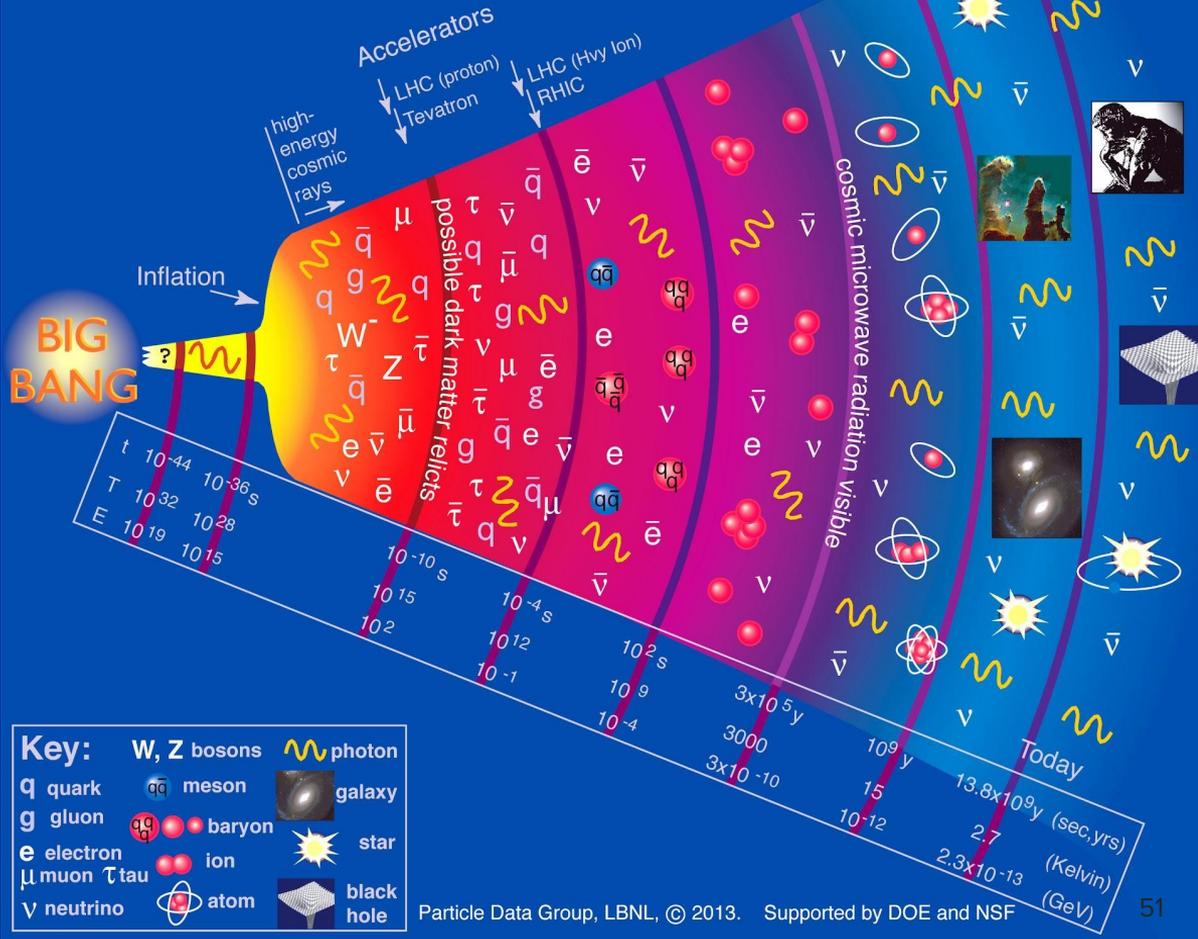
## Timeline

- $5.4 \cdot 10^{-44}$  s and  $T > 10^{19}$  GeV Planck time
- $10^{-43}$  s and  $T \sim 10^{19}$  GeV Grand Unification of the forces

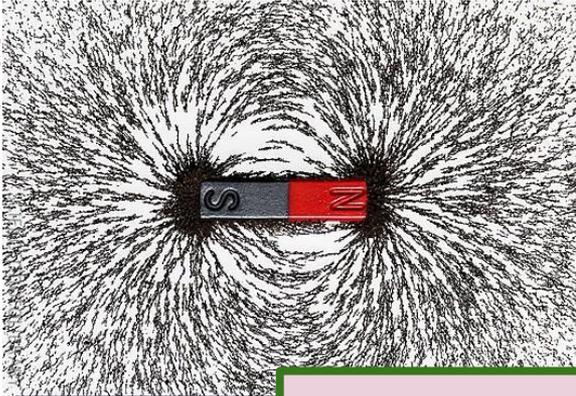
### Standard Model of Elementary Particles

	three generations of matter (fermions)			interactions / force carriers (bosons)	
	I	II	III		
mass span	=2.2 MeV/c <sup>2</sup> 1/6	=1.28 GeV/c <sup>2</sup> 1/6	=173.1 GeV/c <sup>2</sup> 1/6	0 1	=124.97 GeV/c <sup>2</sup> 0
QUARKS	<b>u</b> up	<b>c</b> charm	<b>t</b> top	<b>g</b> gluon	<b>H</b> higgs
	=4.7 MeV/c <sup>2</sup> 1/6	=96 MeV/c <sup>2</sup> 1/6	=4.18 GeV/c <sup>2</sup> 1/6	0 1	
	<b>d</b> down	<b>s</b> strange	<b>b</b> bottom	<b>γ</b> photon	
LEPTONS	=0.511 MeV/c <sup>2</sup> 1	=105.66 MeV/c <sup>2</sup> 1	=1.7768 GeV/c <sup>2</sup> 1	0 1	=91.18 GeV/c <sup>2</sup> 0
	<b>e</b> electron	<b>μ</b> muon	<b>τ</b> tau	<b>Z</b> Z boson	
	<1.0 eV/c <sup>2</sup> 0	=0.17 MeV/c <sup>2</sup> 1/6	=18.2 MeV/c <sup>2</sup> 1/6	0 1	=80.39 GeV/c <sup>2</sup> 1
	<b>ν<sub>e</sub></b> electron neutrino	<b>ν<sub>μ</sub></b> muon neutrino	<b>ν<sub>τ</sub></b> tau neutrino	<b>W</b> W boson	
					SCALAR BOSONS
					GAUGE BOSONS VECTOR BOSONS

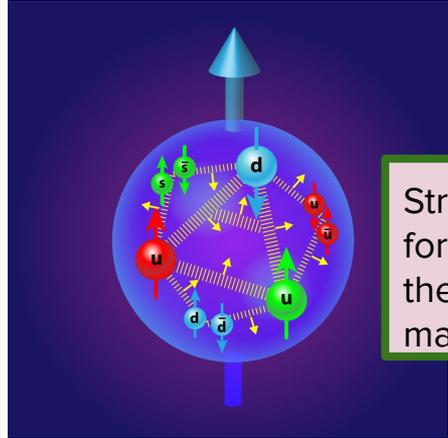
What are the four known forces in nature?



# Interactions: four known forces

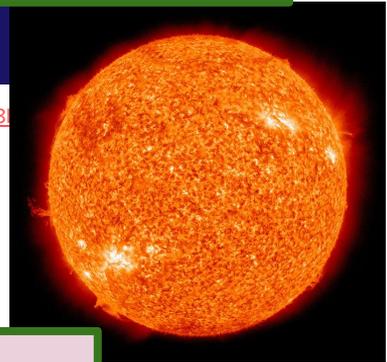
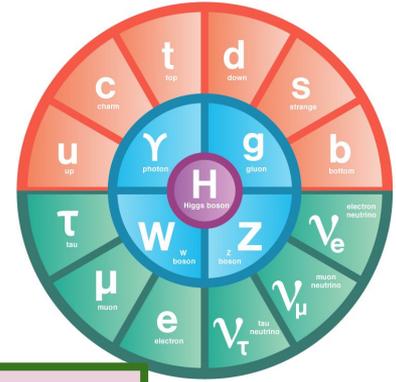


electromagnetism



[https://physics.aps.org/assets/89b4f0e0-b8b70d-d90f744d1790/e23\\_2.png](https://physics.aps.org/assets/89b4f0e0-b8b70d-d90f744d1790/e23_2.png)

Strong nuclear force: how does the proton get its mass?

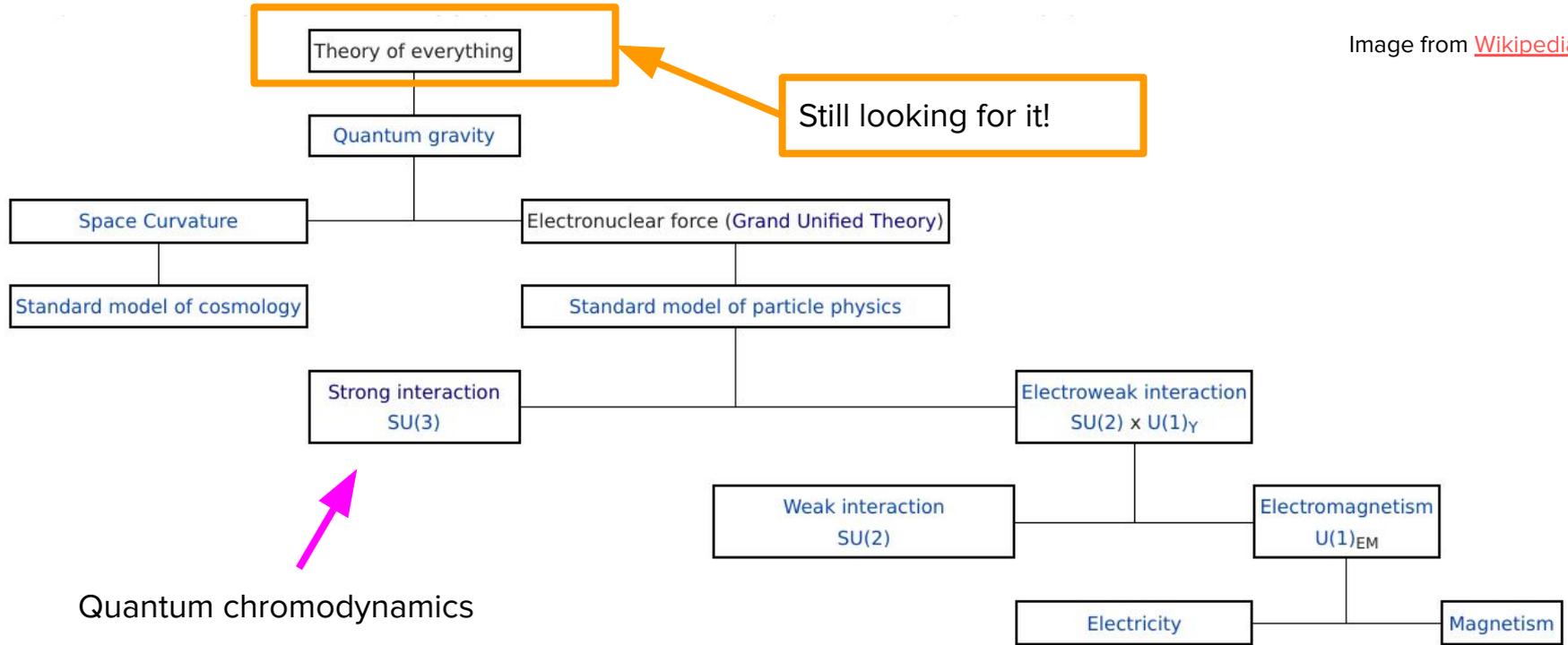


Weak nuclear force

What are we made of?  
How do particles get mass?  
**Gravity is not described by the Standard Model!**

[https://upload.wikimedia.org/wikipedia/commons/thumb/b/b4/The\\_Sun\\_by\\_the\\_Atmospheric\\_Imaging\\_Assembly\\_of\\_NASA%27s\\_Solar\\_Dynamics\\_Observatory\\_-\\_20100819.jpg/800px-The\\_Sun\\_by\\_the\\_Atmospheric\\_Imaging\\_Assembly\\_of\\_NASA%27s\\_Solar\\_Dynamics\\_Observatory\\_-\\_20100819.jpg](https://upload.wikimedia.org/wikipedia/commons/thumb/b/b4/The_Sun_by_the_Atmospheric_Imaging_Assembly_of_NASA%27s_Solar_Dynamics_Observatory_-_20100819.jpg/800px-The_Sun_by_the_Atmospheric_Imaging_Assembly_of_NASA%27s_Solar_Dynamics_Observatory_-_20100819.jpg)

# Hypothetical unification into a theory of everything



# History of the Universe

Image: [Particle Data Group](#)

## Timeline

- $5.4 \cdot 10^{-44}$  s and  $T > 10^{19}$  GeV Planck time
- $10^{-43}$  s and  $T \sim 10^{19}$  GeV Grand Unification
- $10^{-36}$  s and  $T \sim 10^{15}$  GeV Inflation
- $10^{-11}$  s and  $T \sim 200$  MeV Quark Gluon Plasma

QGP: a state where quarks and gluons are free: main ALICE goal.

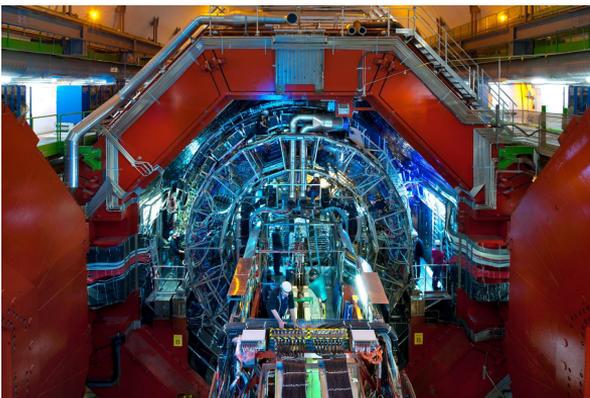


Image: [Wikipedia](#)

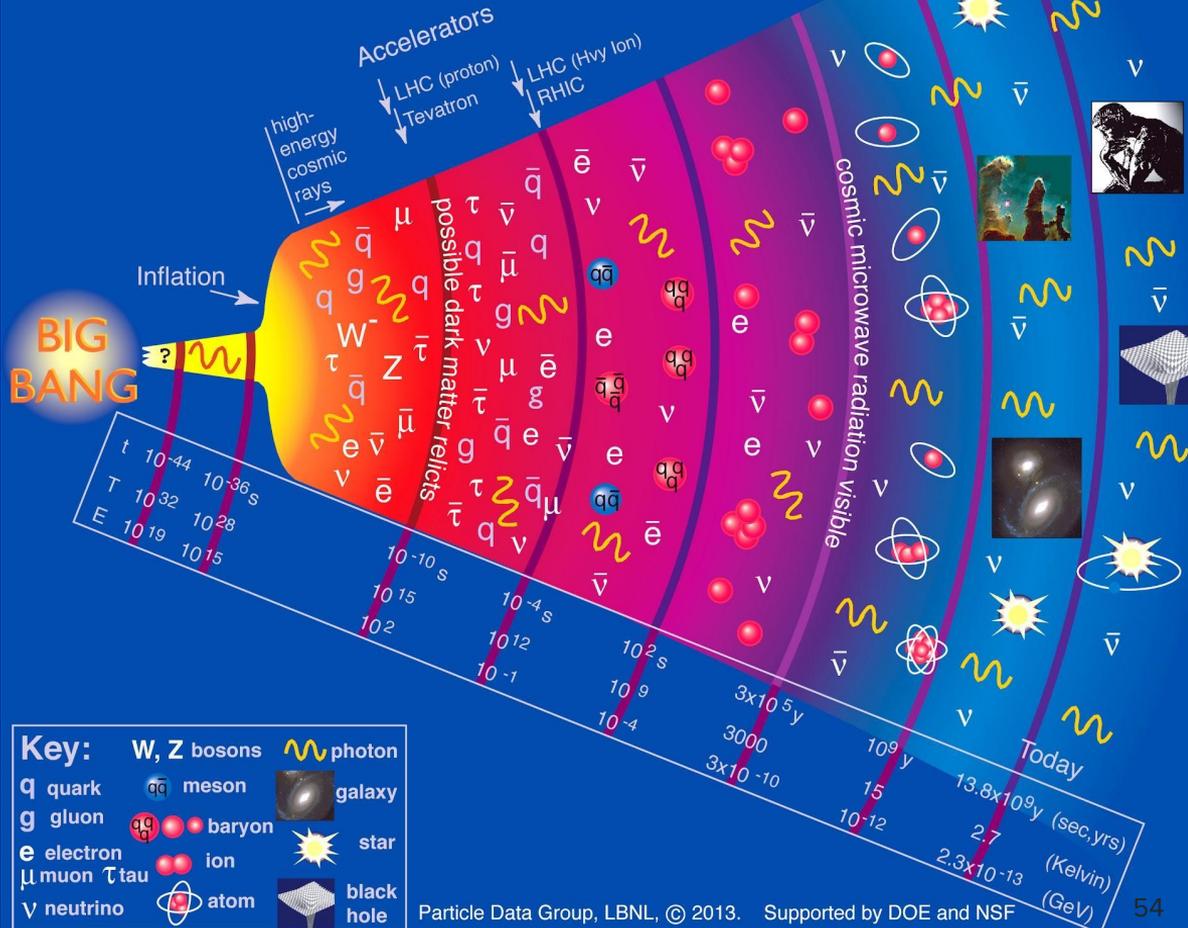
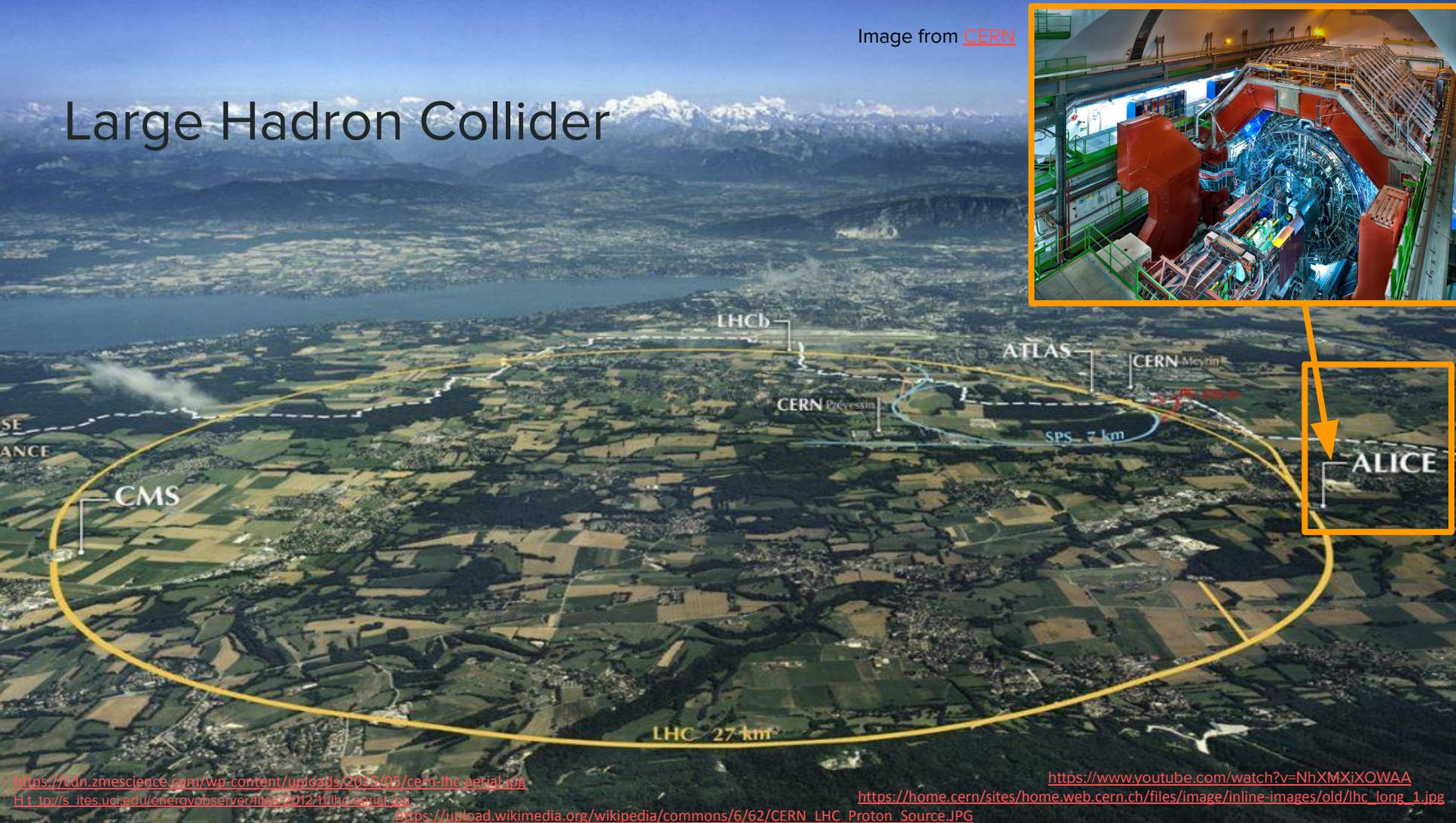


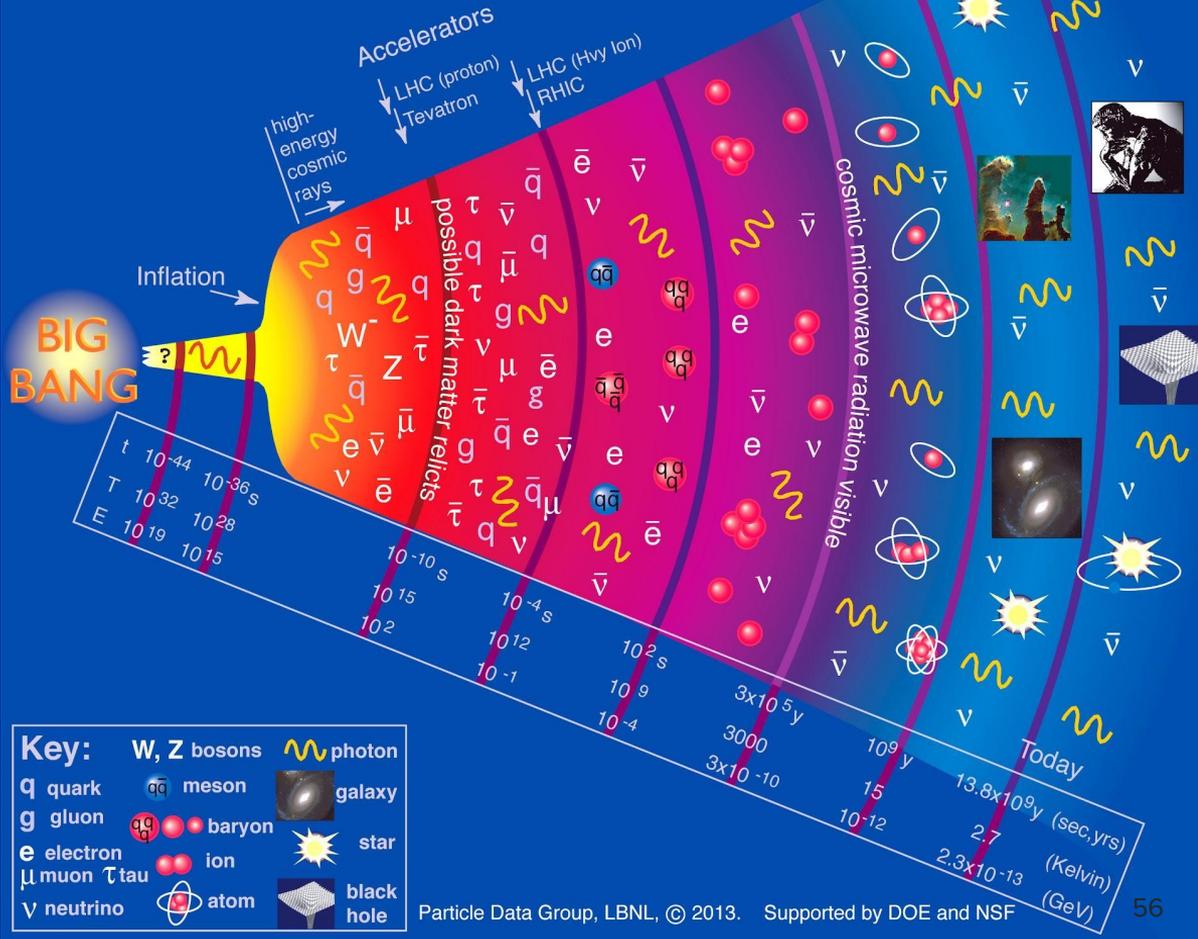
Image from [CERN](#)

# Large Hadron Collider



# History of the Universe

Image: [Particle Data Group](#)



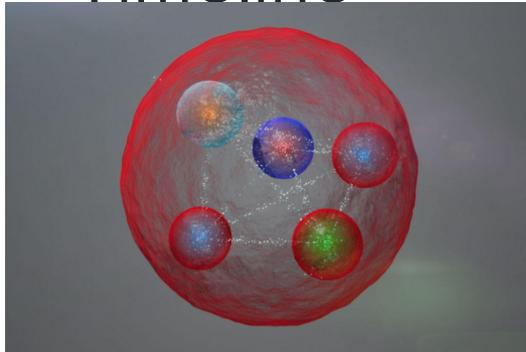
**Key:**

W, Z bosons	photon
q quark	meson
g gluon	baryon
e electron	ion
μ muon	τ tau
ν neutrino	atom
	galaxy
	star
	black hole

Particle Data Group, LBNL, © 2013. Supported by DOE and NSF

## Timeline

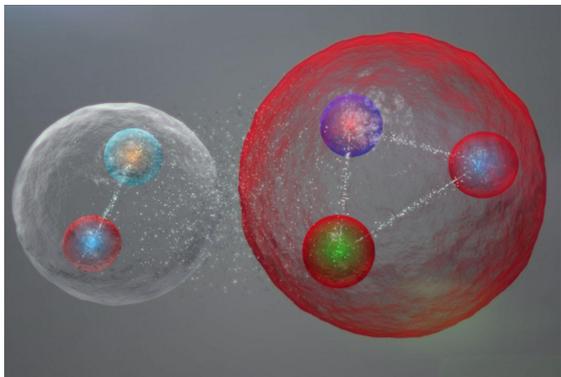
Images: [wired.com](#)



Penta quark

- 10<sup>-5</sup> s and T ~ **200 MeV** Hadronization and Baryogenesis

## Meson and Baryon

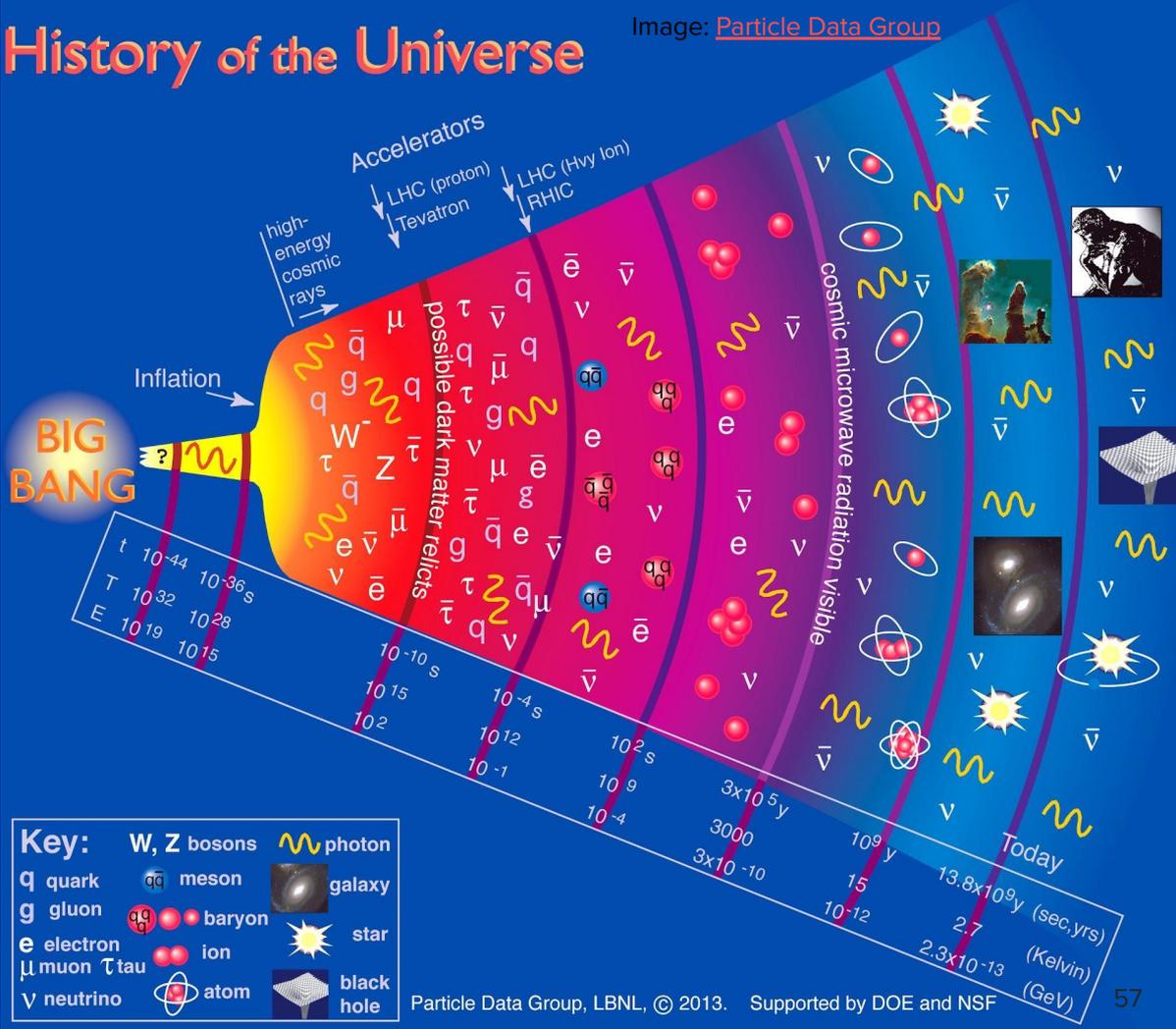


# History of the Universe

Image: [Particle Data Group](#)

## Timeline

- $5.4 \cdot 10^{-44}$  s and  $T > 10^{19}$  GeV Planck time
- $10^{-43}$  s and  $T \sim 10^{19}$  GeV Grand Unification
- $10^{-36}$  s and  $T \sim 10^{15}$  GeV Inflation
- $T \sim 10^1 - 10^3$  GeV dark matter freeze-out
- $10^{-11}$  s and  $T \sim 100$  GeV Quark Gluon Plasma
- $10^{-5}$  s and  $T \sim 200$  MeV Hadronization and Baryogenesis
- 1 s or  $T \sim 1$  MeV neutron freeze-out
- 1 s or  $T \sim 1$  MeV neutrino freeze-out
- 3 min and  $T \sim 100$  keV nucleosynthesis:  
D,  $^3\text{He}$ ,  $^4\text{He}$ , Li
- $T \sim 1$  eV structure formation
- $T \sim 0.4$  eV photon decoupling  $\rightarrow$  CMB
- $T = 2.7$  K  $\sim 10^{-4}$  eV today

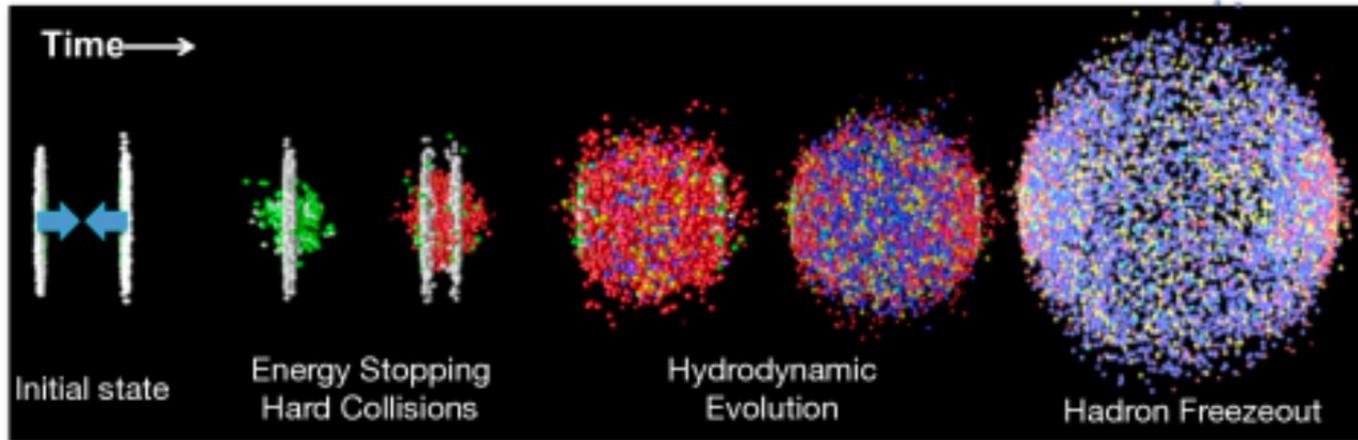


# Heavy ions for quark gluon plasma

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# Quark gluon plasma in the laboratory

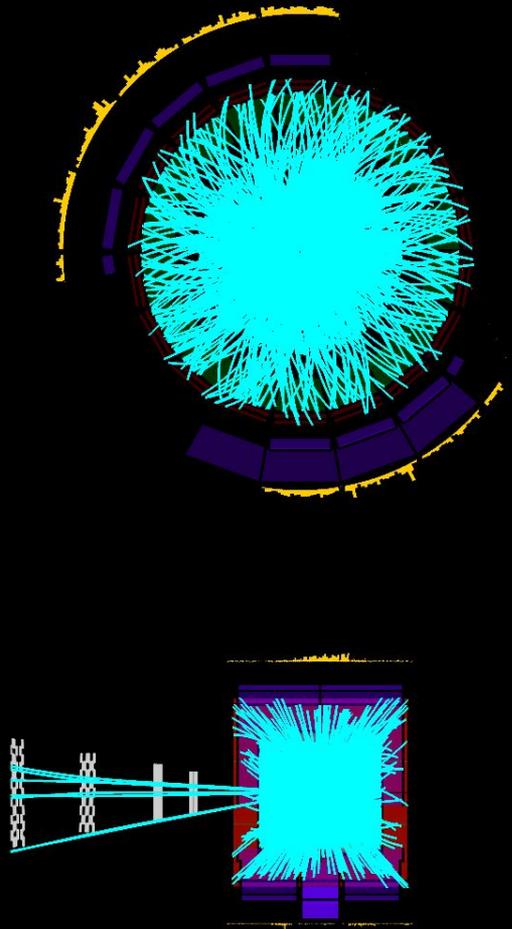
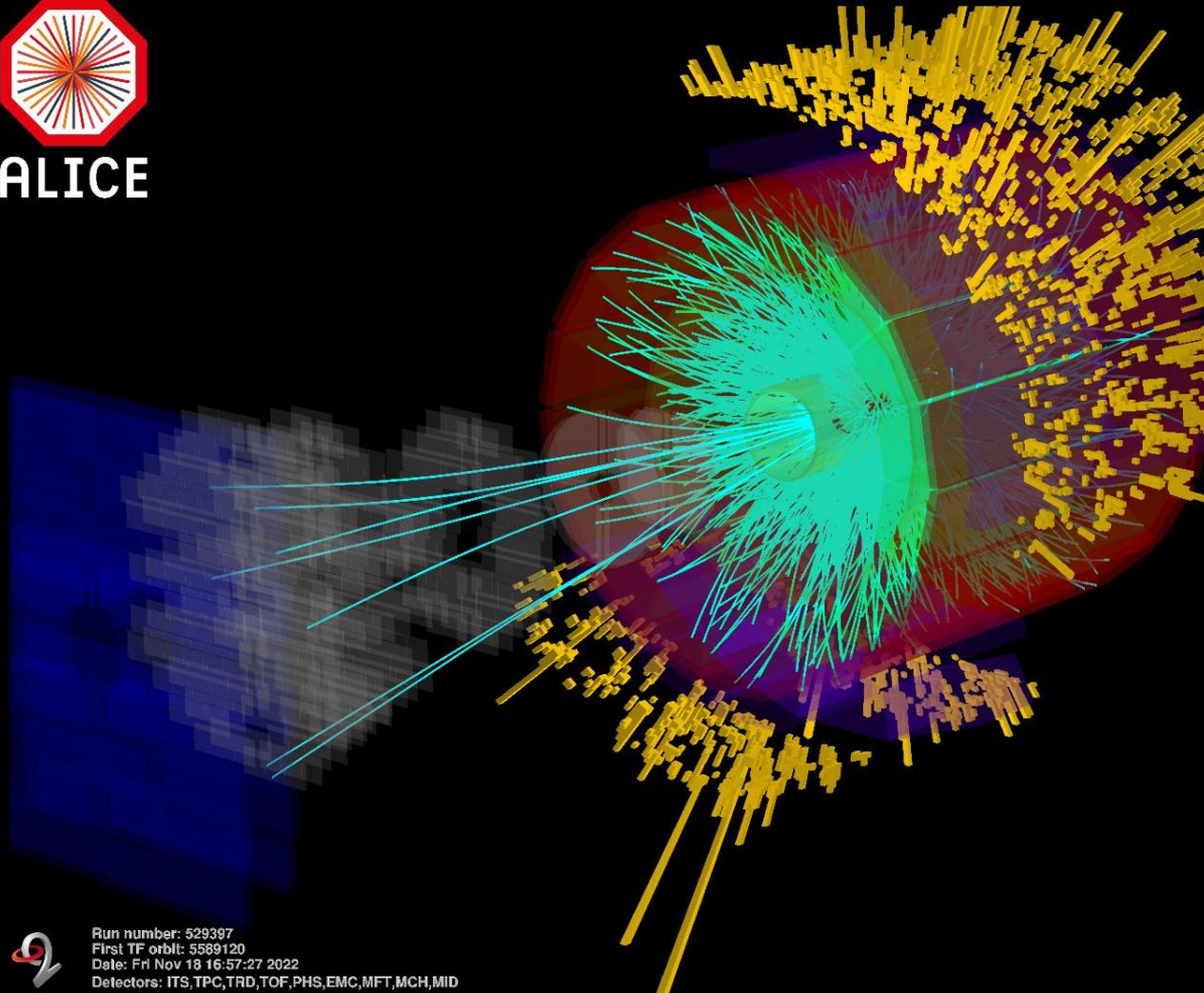
- 🔬 We collide heavy-ions, like lead on lead at speeds close to  $c$ , the speed of light ( $\sim 0.9998 c$ )



- 🔬 The energy density in the collision is large enough to produce QGP in the laboratory



ALICE



Run number: 529397  
First TF orbit: 5589120  
Date: Fri Nov 18 16:57:27 2022  
Detectors: ITS,TPC,TRD,TOF,PHS,EMC,MFT,MCH,MID



First ion collisions after long shutdown of several years!

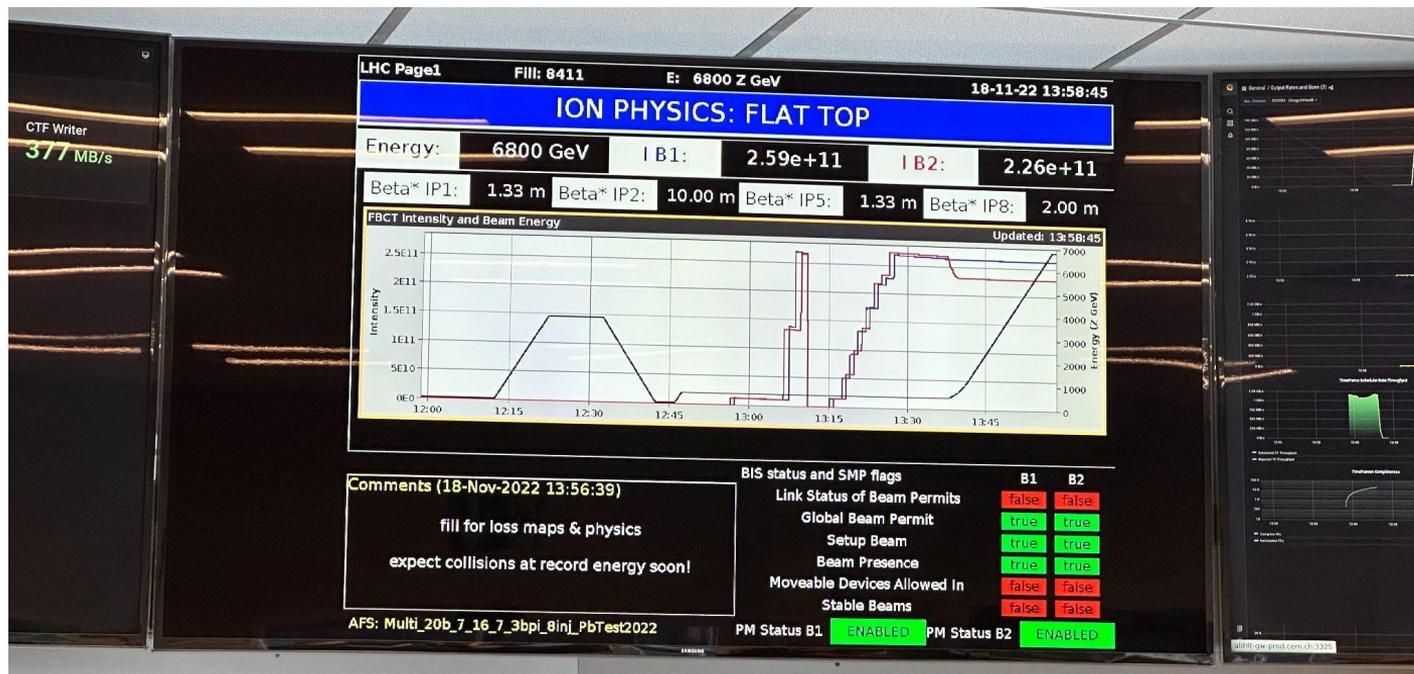
See also

<https://home.cern/news/news/experiments/first-lead-ion-collisions-lhc-record-energy>  
<https://alice.cern/node/169>



# What is the LHC doing right now?

Check out <https://op-webtools.web.cern.ch/vistar/vistars.php?usr=LHC1> or google “vistars LHC page 1”



Hoe detecteren we zo'n deeltje?

---

# ALICE inner tracking system:

10 m<sup>2</sup> of active silicon area  
nearly 13 billion pixels

The largest pixel detector ever built!

Nog  
dichter  
bij de  
oerknal

Deeltjeslab CERN, Genève. Als een geheimzinnige gouden halfpijp ligt een van de nieuwe onderdelen van deeltjesdetector ALICE in het assemblage-atelier. Dit is ITS, het Inner Tracking System dat het vederlichte hart van de detector gaat vormen. Een meterslange halve buis van ultralichte koolstofvezel vakwerkbalkjes met

Binnenin ITS ligt dan alleen nog de bundelpijp van de LHC-versneller, die middenin ALICE zware atoomkernen met de licht-

## A silicon PIXEL detector

daarop zijn vastgeplakt. De balkjes worden met de hand geplaatst en worden met een speciale machine aan elkaar vastgeplakt. Het is een zeer nauwkeurig proces. De balkjes worden dan in een speciale oven opgewarmd. Het proces duurt ongeveer een week. De balkjes worden dan in een speciale oven opgewarmd. Het proces duurt ongeveer een week.

de binnenin precies gaande is. De balkjes worden dan in een speciale oven opgewarmd. Het proces duurt ongeveer een week.

dezelfde plak silicium zitten. Dat scheidt kabels en elektronica in de detector.

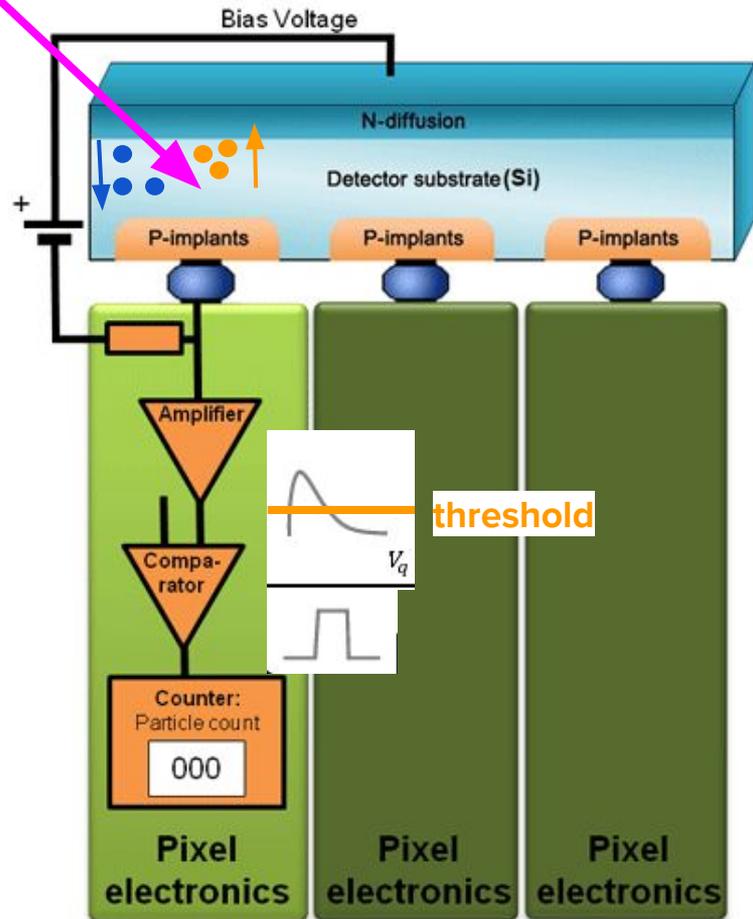
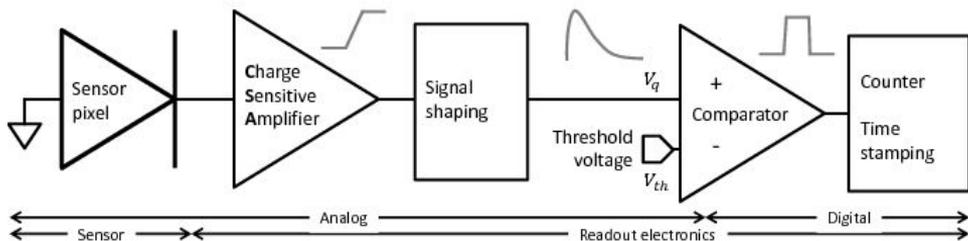
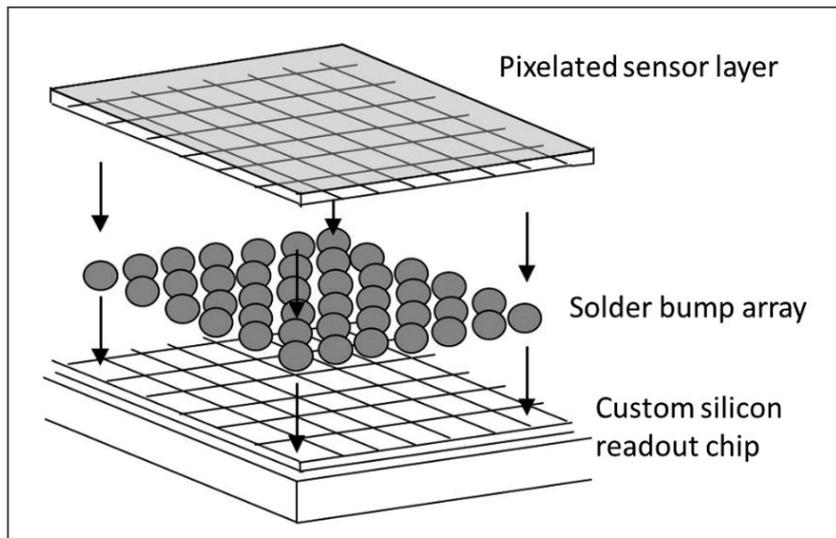
In de komende meetperiode kan ITS gemakkelijk honderd keer zoveel meetgegevens verzamelen als alles wat ALICE in

ne versies worden vervangen. Daar is ook de trigger-apparatuur bij die beslist welke botsingen bijzonder genoeg zijn om vast te leggen. Het computersysteem dat data verzamelt en toegankelijk maakt, wordt eveneens vernieuwd.

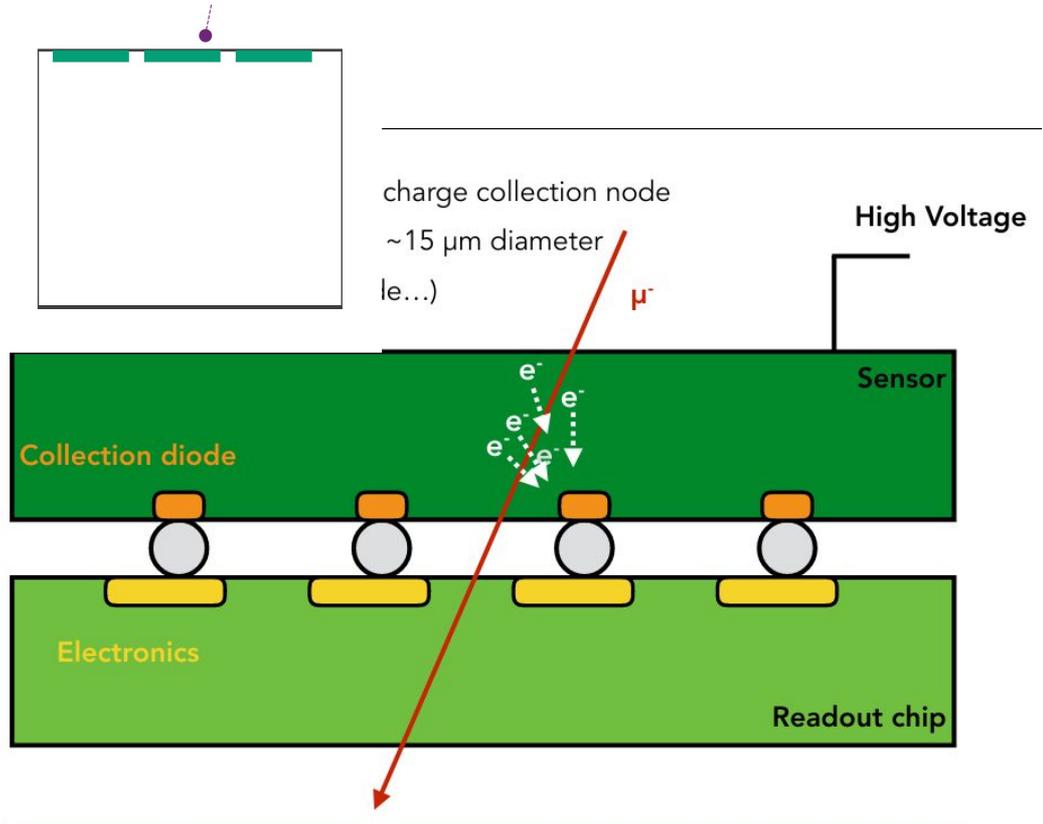
De upgrade-periode is een hectische tijd. Het binnenste van de grote ondergrondse detector is vorig jaar meteen

kleine honderd sensorduigen. Een kwart van alle duigen, die in de lagen nummer 6 en 7, zijn gemaakt op Nikhef in Amsterdam. Daar lijden leden van het ALICE-team met eindeloos geduld de koeling en de sensoren stuk voor stuk 64 handmatig op de ijle koolstofvezel dragers. Deze sensorduigen zijn vorig najaar al in trillingsvrije krachten van Amsterdam

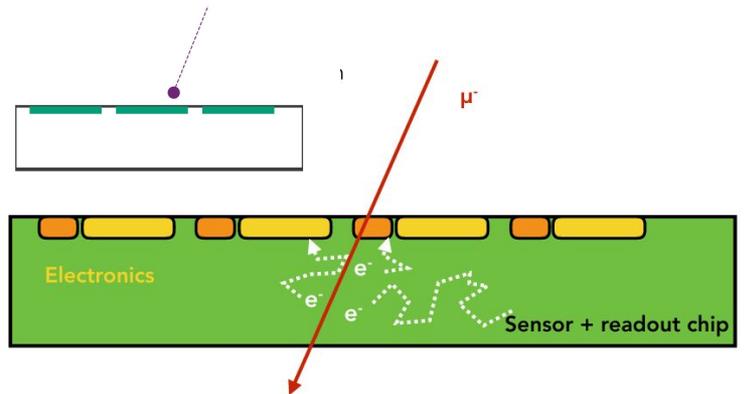
# A hybrid pixel detector



# Deeltjes maken lading los in een sensor: *ionisatie*



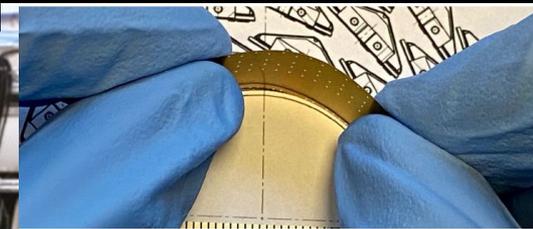
Ionisatie: atoom of molecuul raakt een elektron kwijt of krijgt er één bij.





Mijn passie: pixeldetectors van silicium -- ook in je telefoon!

Zo dun, we kunnen ze buigen!



Nog  
dichter  
bij de  
oerknal

Ik werk aan hele snelle pixel detectors: we gaan voor  
 $10 \text{ ps} = 0.0000000001 \text{ seconde!}$

Het goud is overigens geen goud, maar polyimide-folie met ragdunne koperen voedingskabels voor de sensoren. Dun genoeg om vrijkomende

zijn ontstaan. ITS moet de ontwerpen die uit die ziedende ontspannen en de fysici vertellen wat daarbinnen precies gaande is.

botsingen preciezer worden bekeken.

gemakkelijk honderd keer zoveel meetgegevens verzamelen als alles wat ALICE in

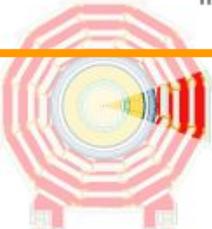
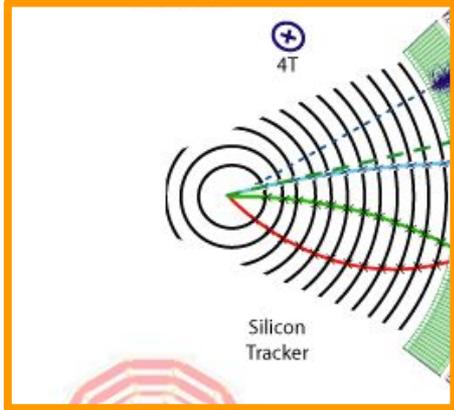
tijd. Het binnenste van de grote ondergrondse detector is vorig jaar meteen

gers. Deze sensorduigen zijn vorig najaar al in trillingsvrije kratten van Amsterdam

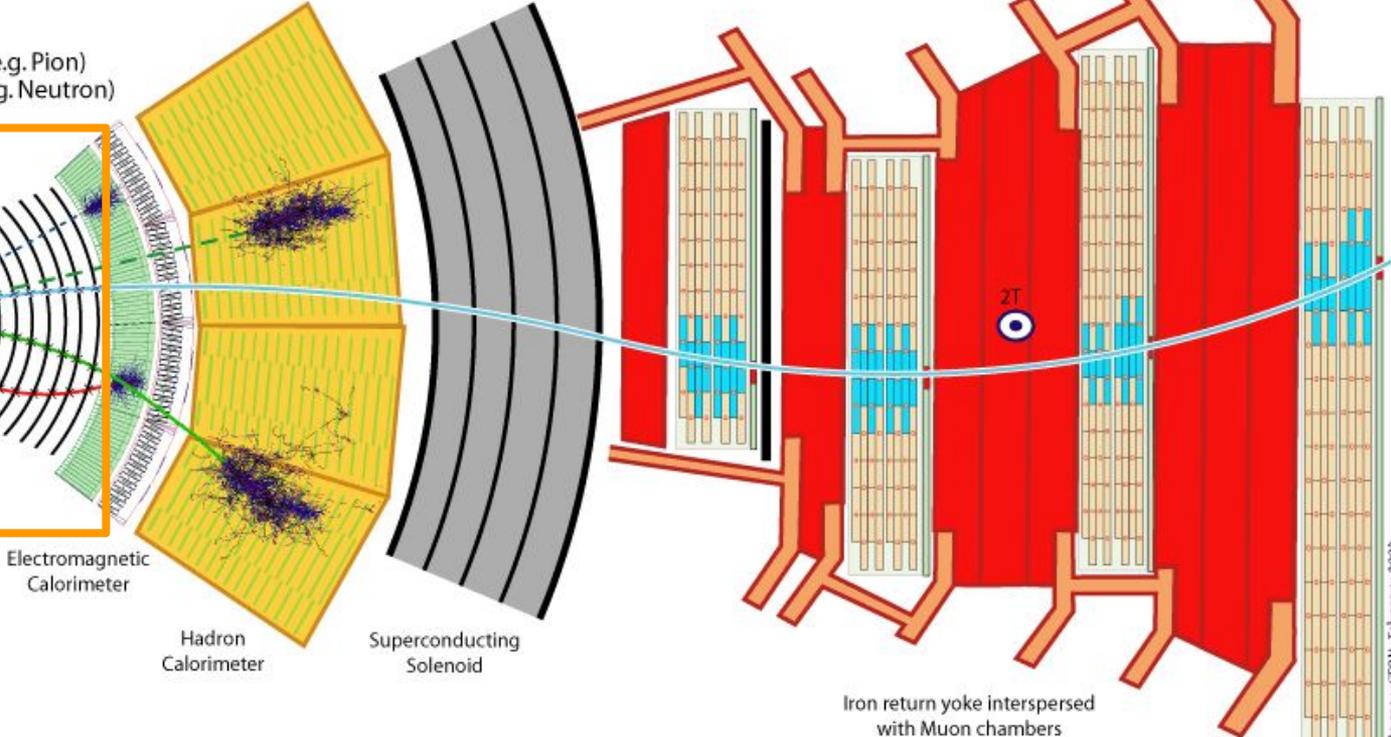
# Detectors at the large hadron collider: onion-like



- Key:
- Muon
  - Electron
  - Charged Hadron (e.g. Pion)
  - - - Neutral Hadron (e.g. Neutron)
  - - - Photon



Transverse slice through CMS



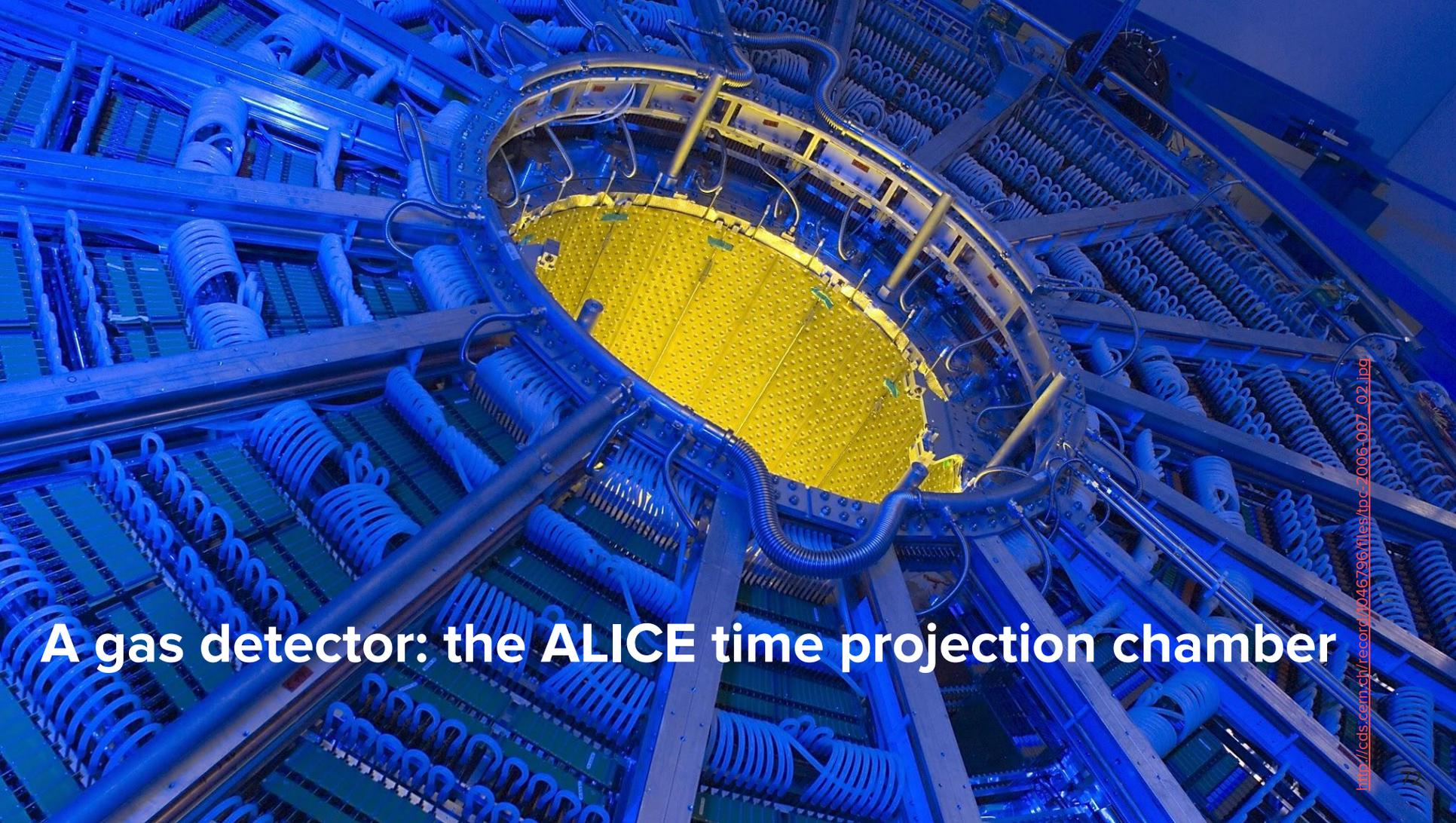
Example: CMS detector



# ALICE Time Projection Chamber

---

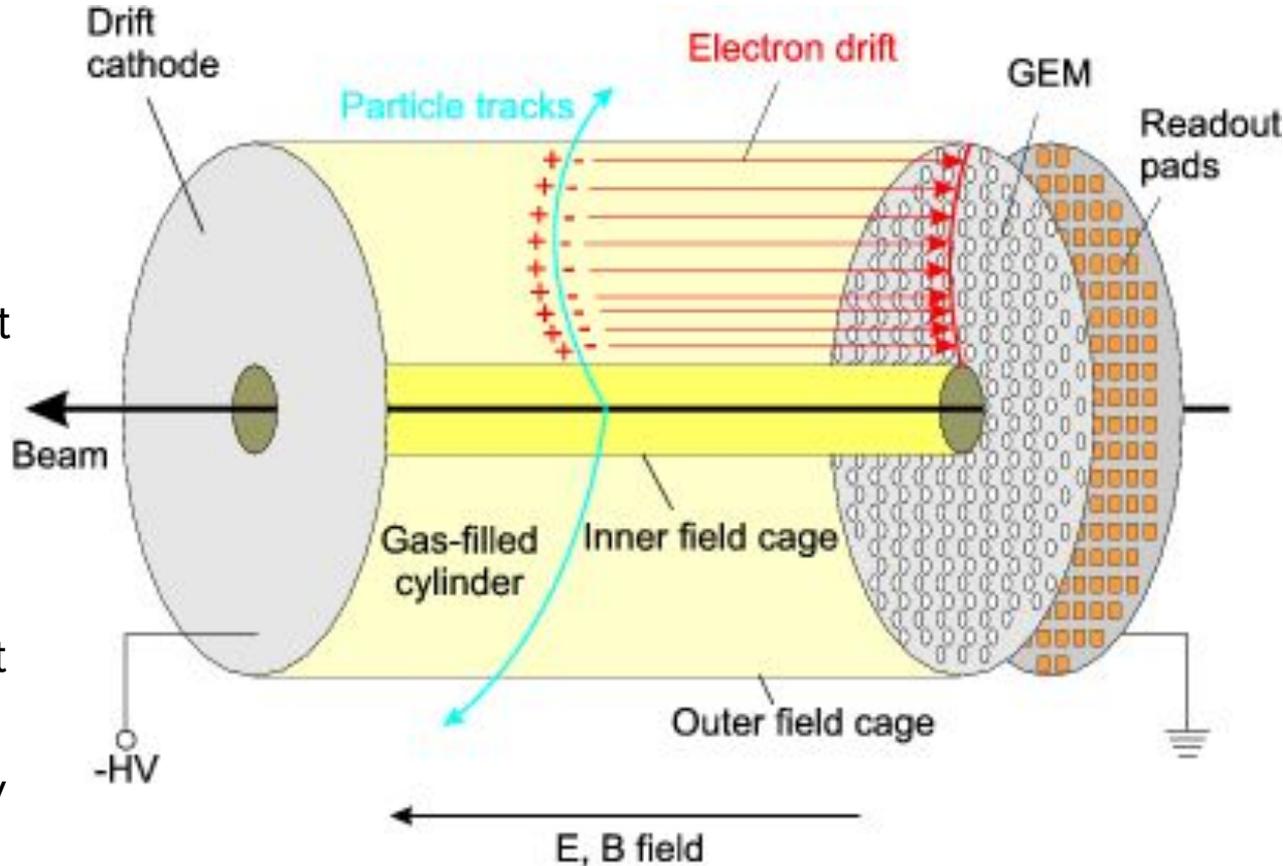
Een gas detector



A gas detector: the ALICE time projection chamber

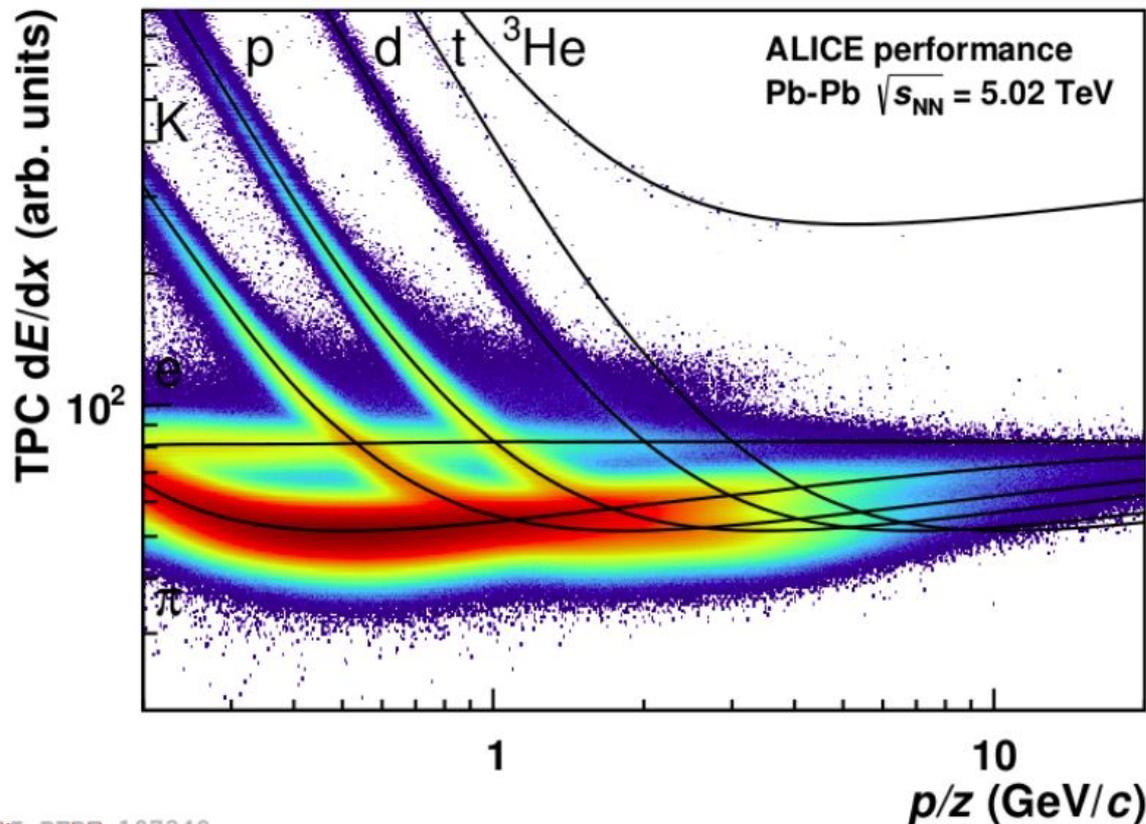
# TPC

1. Ionization of gas in **chamber** with electric field causes electron drift
2. Signal gets amplified, in this case by gas electron multipliers  $\rightarrow$  electron avalanche
3. Readout pads can detect signal that can be **projected** onto trajectory
4. z (along beam) information from **timing**

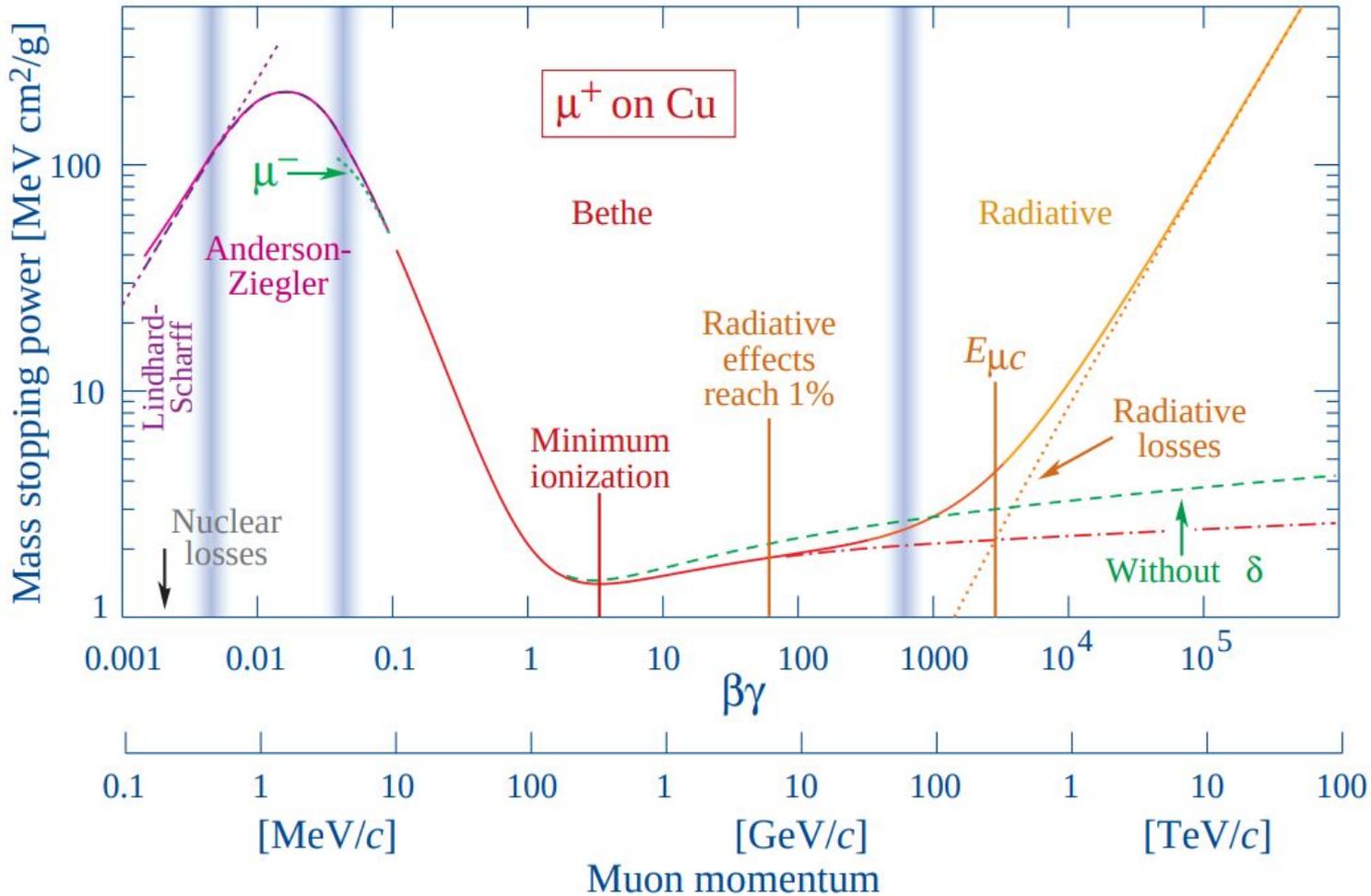


# ALICE time projection chamber: particle identification

- Every point is one measurement!
- Can identify particles for low momenta
- For higher momenta, all particles behave like a minimum ionizing particle (MIP)



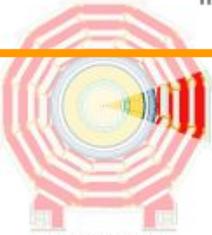
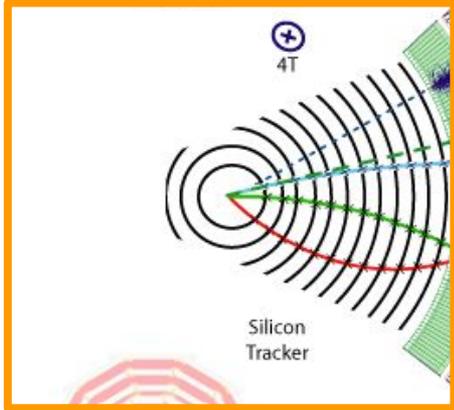
# Ionization loss



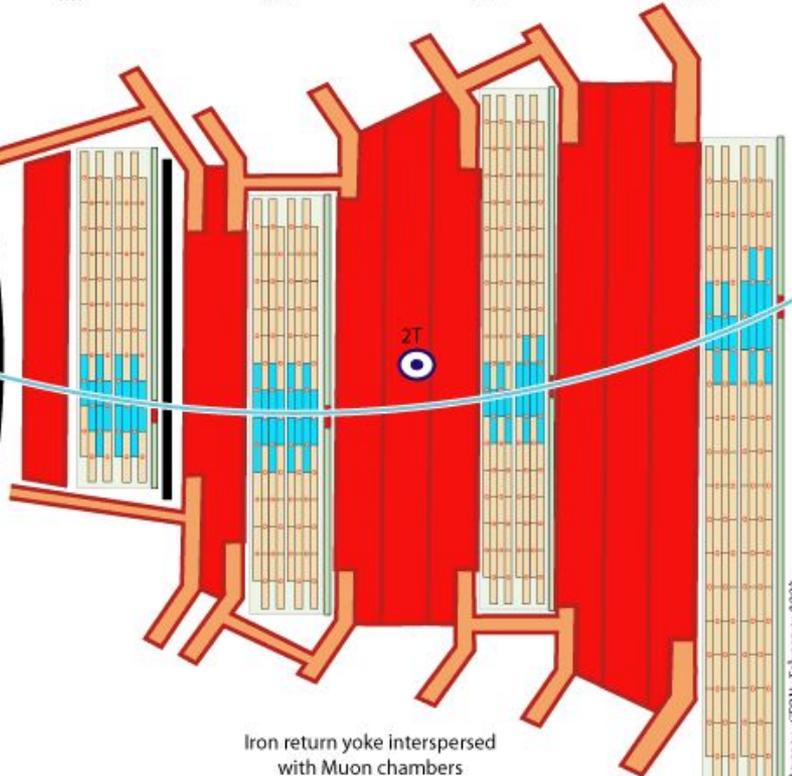
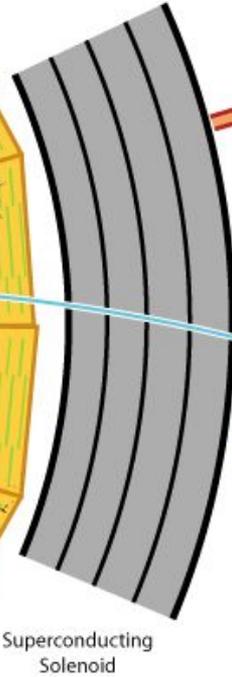
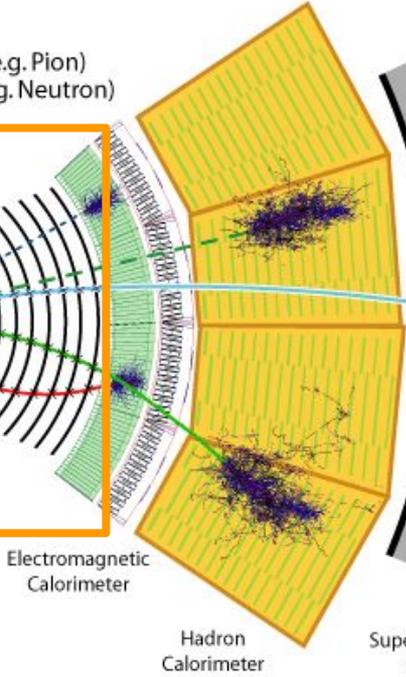
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  - - - Neutral Hadron (e.g. Neutron)
  - - - Photon



Transverse slice through CMS



[http://inspirehep.net/record/82.6852/files/EPS\\_CMS\\_Slice.png](http://inspirehep.net/record/82.6852/files/EPS_CMS_Slice.png)

# How to measure neutrinos?

---

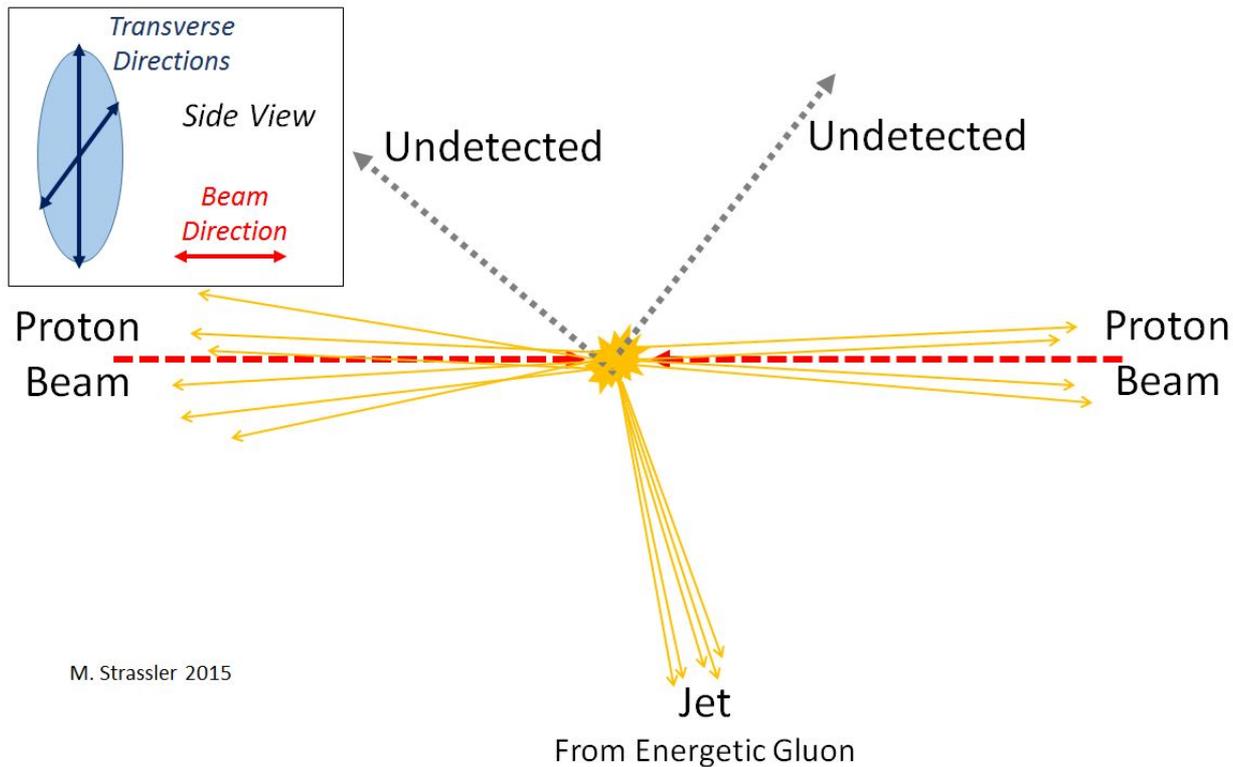
# Missing transverse momentum

Neutrinos?

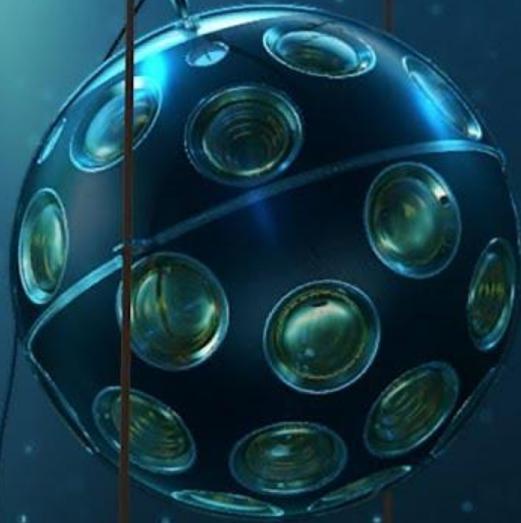
Mismeasurement?

Detector effect?

Dark matter?



- $10^9$  neutrinos /  $\text{cm}^2/\text{s}$
- De meeste van de zon en uit de atmosfeer
- Zeldzame gebeurtenissen van zwarte gaten, supernovae...

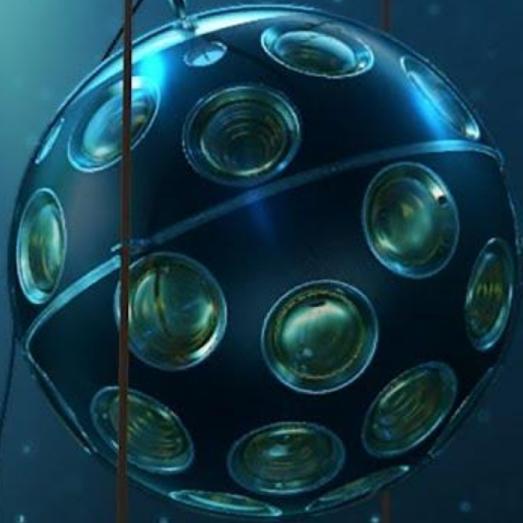


## **KM3NeT: cubic kilometer neutrino telescope**

- Tussen 2 en 4 km diep in Middellandse Zee (FR-IT-GR)
- 12000 digital optical modules (DOMs) aan 600 draden
- Cherenkov detectie met fotobuizen
- GeV, TeV, and PeV neutrinos

**Nederland speelt een grote rol in de constructie!**

- $10^9$  neutrinos /  $\text{cm}^2/\text{s}$
- Most from sun and atmosphere
- Rare events from black holes, supernovae...



## Neutrinojagers bouwen hun sensorbollen in Amsterdam



In de PIMU-hal van Nikhef op het Amsterdam Science Park is afgelopen zomer de massaproductie gestart van onderdelen voor de reusachtige internationale KM3NeT neutrino-telescoop die op de bodem van de Middellandse Zee wordt gebouwd.

Wekelijks kunnen daar acht tot twaalf

# Nieuwe detectoren op Nikhef en CERN

[Nikhef wetenschapsdag](#)

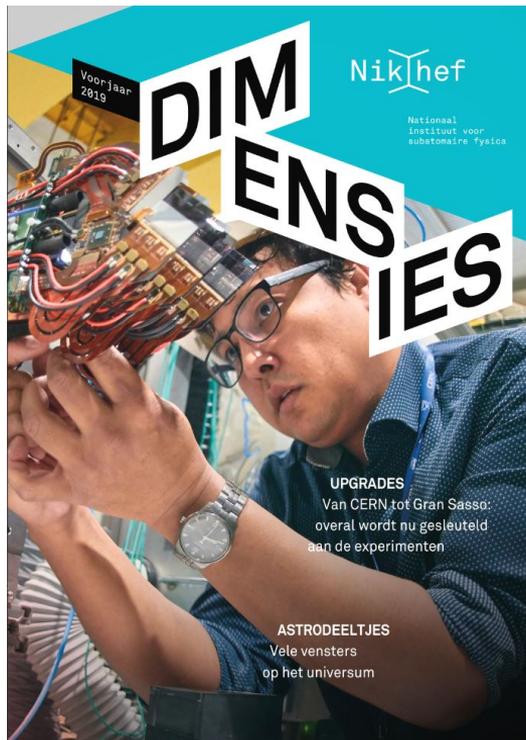
En [hier](#)

Zaterdag 5 oktober 12:00-17:00

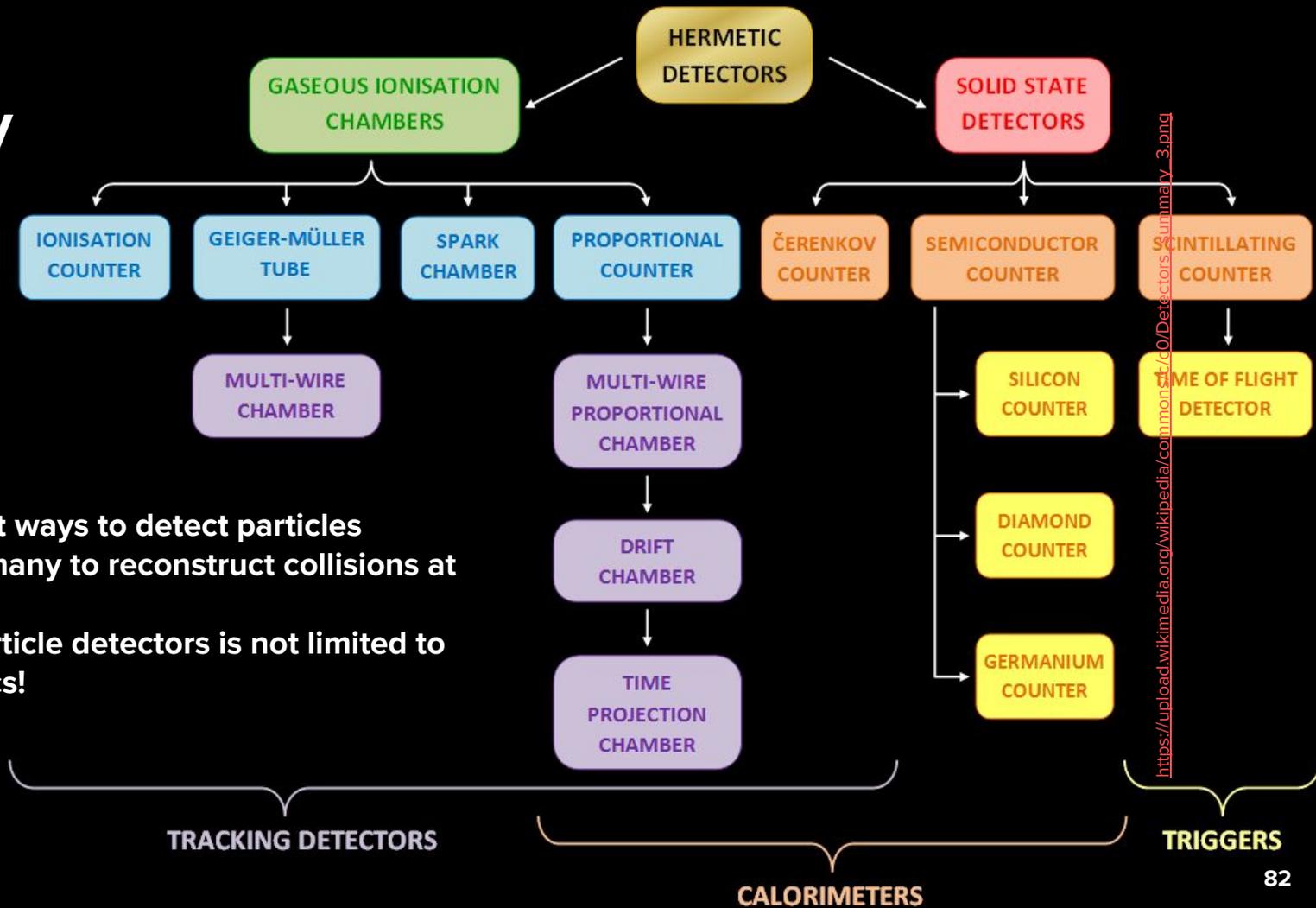
[Wetenschapsdag Amsterdam](#)

[Science Park](#)

Op Nikhef werken we mee aan detectoren die op CERN bij de LHC worden geïnstalleerd!



# Summary



- Many different ways to detect particles
- We combine many to reconstruct collisions at the LHC
- The use of particle detectors is not limited to collider physics!

# Extra materiaal

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# Credits

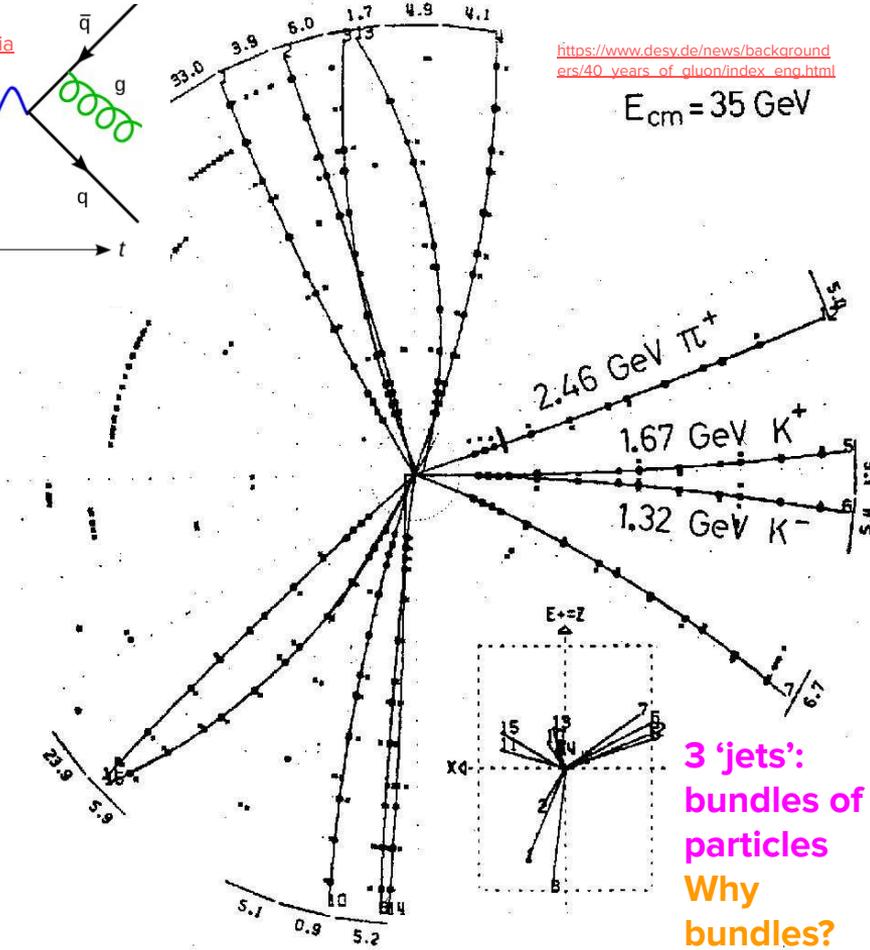
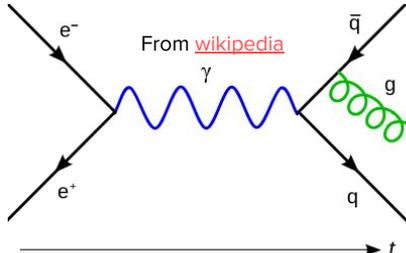
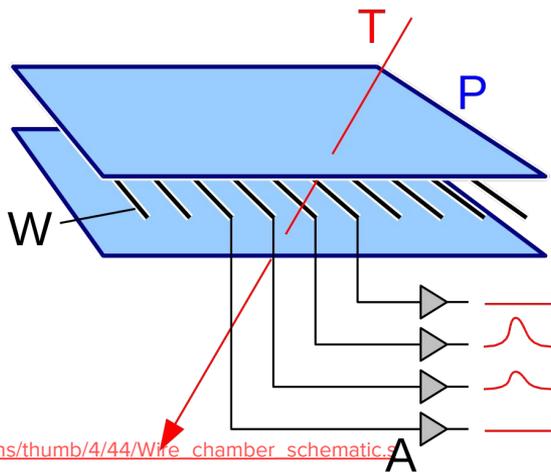
Many thanks to Erik Butz, Simon Spannagel, Freya Blekman, Peter Schleper, Erika Garutti  
**Wikipedia**

# Gluon discovery

Event in *drift chamber* of JADE experiment at PETRA collider at DESY.

Such events were used to prove the existence of gluons:  $e^+e^- \rightarrow qq\bar{g}$ .

**Wire chamber:** particle T passes through grounded plate P and ionizes gas in chamber. Charge drifts in electric field to high voltage wires W and is collected at an amplifier A.



[https://www.desy.de/news/backgrounders/40\\_years\\_of\\_gluon/index\\_eng.html](https://www.desy.de/news/backgrounders/40_years_of_gluon/index_eng.html)

$E_{cm} = 35 \text{ GeV}$

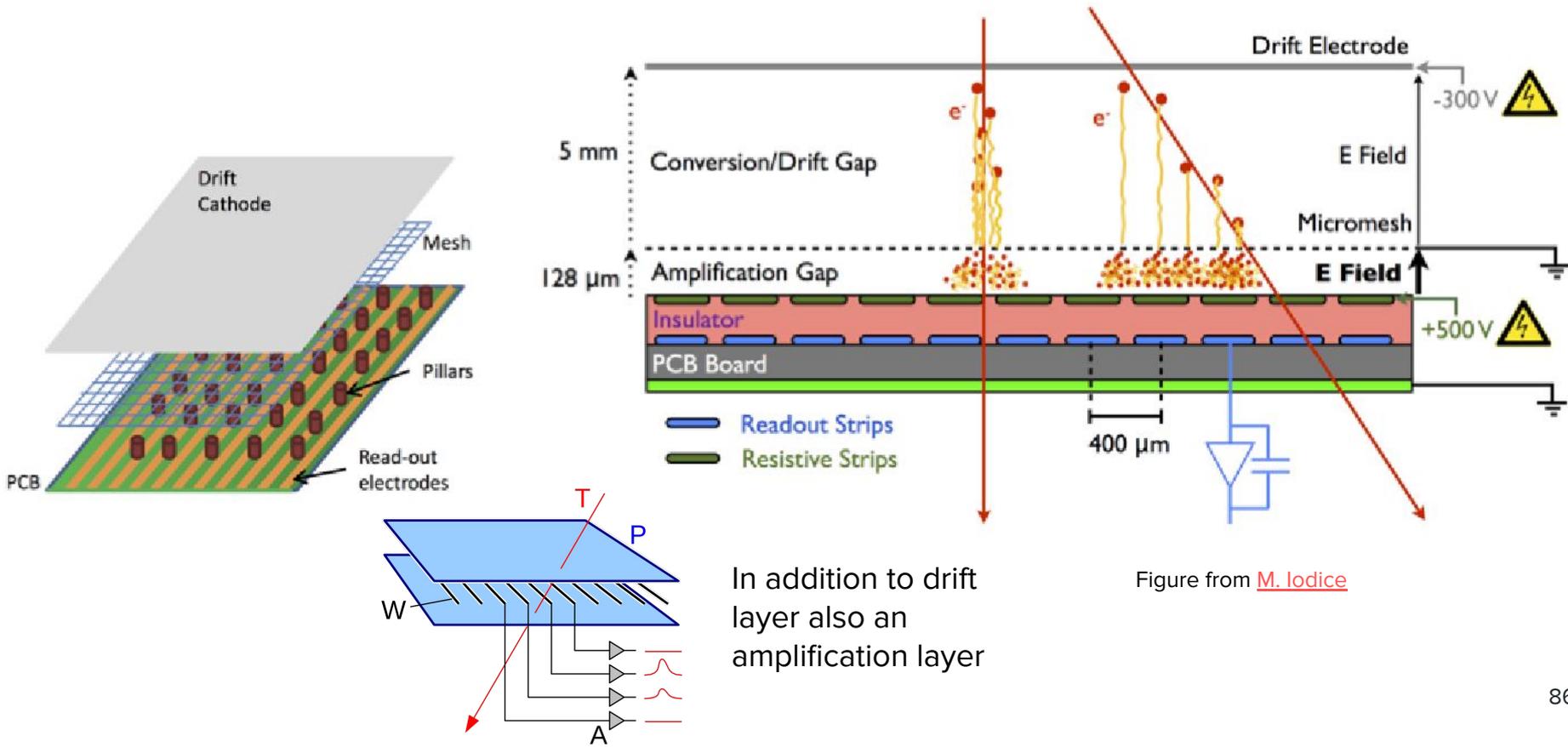
From [http://www.desy.de/sites2009/site\\_www-desy/content/e409/e287332/e287337/e287345/1980-09-22\\_TASSO-Event\\_Gluon\\_Entdeckung\\_sw\\_ger.jpg](http://www.desy.de/sites2009/site_www-desy/content/e409/e287332/e287337/e287345/1980-09-22_TASSO-Event_Gluon_Entdeckung_sw_ger.jpg)

[https://upload.wikimedia.org/wikipedia/commons/thumb/4/44/Wire\\_chamber\\_schematic.svg/1024px-Wire\\_chamber\\_schematic.svg.png](https://upload.wikimedia.org/wikipedia/commons/thumb/4/44/Wire_chamber_schematic.svg/1024px-Wire_chamber_schematic.svg.png)

22.9.80

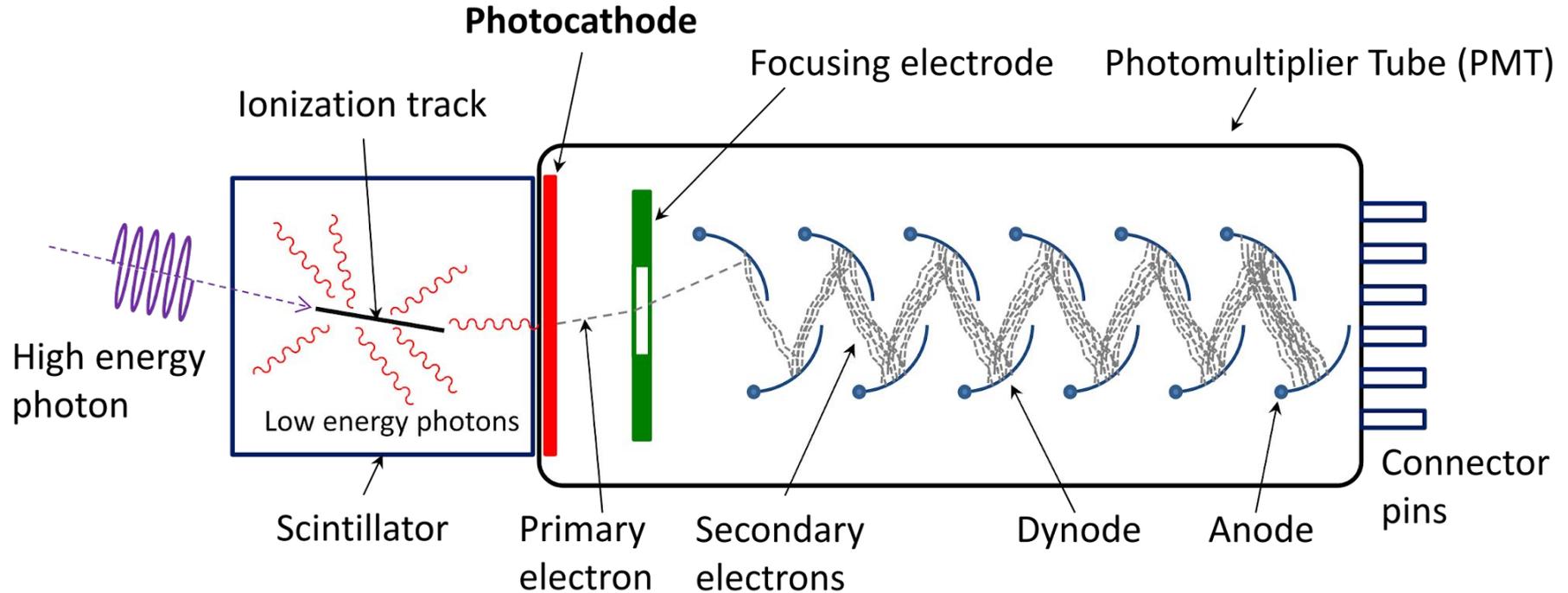
[https://www.desy.de/~schlepèr/lehre/physik5/WS\\_2018\\_19/Physik\\_5\\_72-95.pdf](https://www.desy.de/~schlepèr/lehre/physik5/WS_2018_19/Physik_5_72-95.pdf)

# Micromesh Gaseous Structure: Micromegas



# Photomultiplier tube

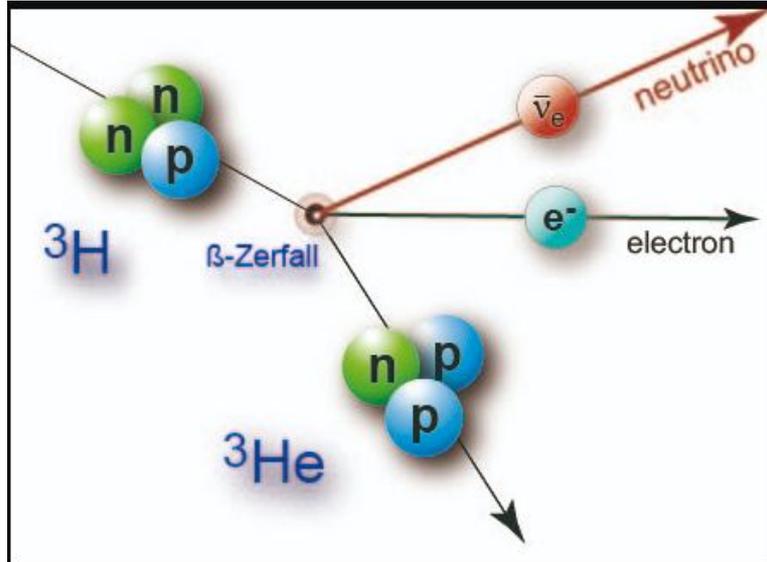
Image from [Wikipedia](#)



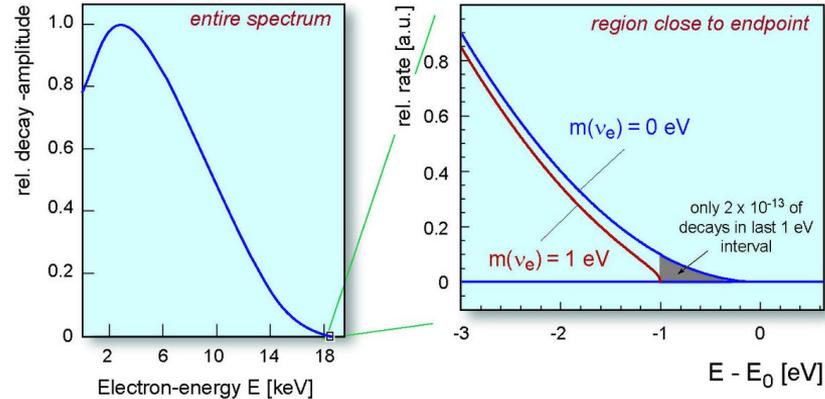


# KATRIN: neutrino mass measurement

Karlsruhe tritium neutrino experiment



$m_\nu < 1.1 \text{ eV}$  (90% confidence level)



Derive  
neutrino mass  
information  
from electron  
energy

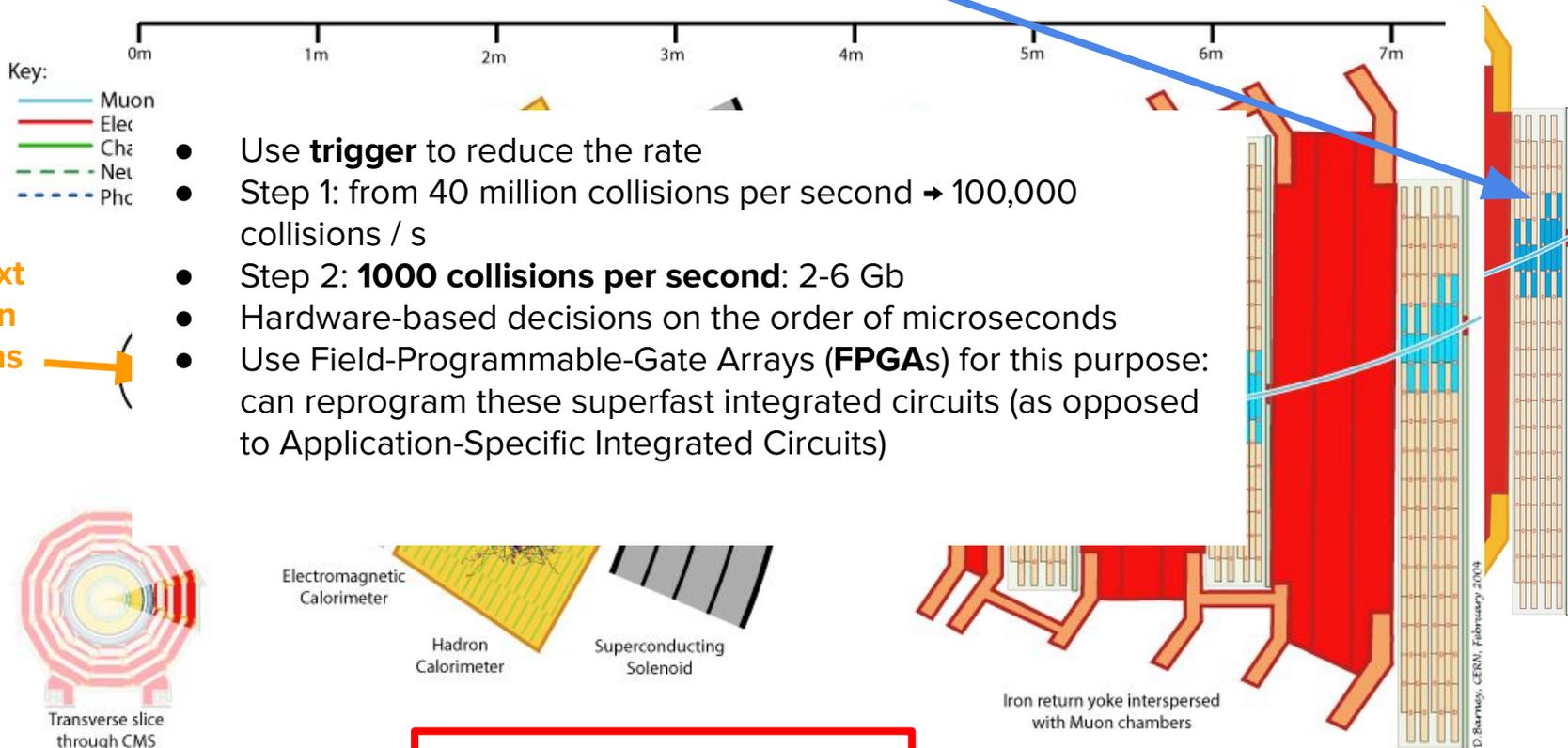
First experiment with sub-eV  
neutrino mass constraints:

<https://www.nature.com/articles/s41567-021-01463-1>



[https://www.katrin.kit.edu/img/spdctrum\\_rdx\\_1200x678.jpg](https://www.katrin.kit.edu/img/spdctrum_rdx_1200x678.jpg)

# Note when the muon arrives here



- Use **trigger** to reduce the rate
- Step 1: from 40 million collisions per second → 100,000 collisions / s
- Step 2: **1000 collisions per second**: 2-6 Gb
- Hardware-based decisions on the order of microseconds
- Use Field-Programmable-Gate Arrays (**FPGAs**) for this purpose: can reprogram these superfast integrated circuits (as opposed to Application-Specific Integrated Circuits)

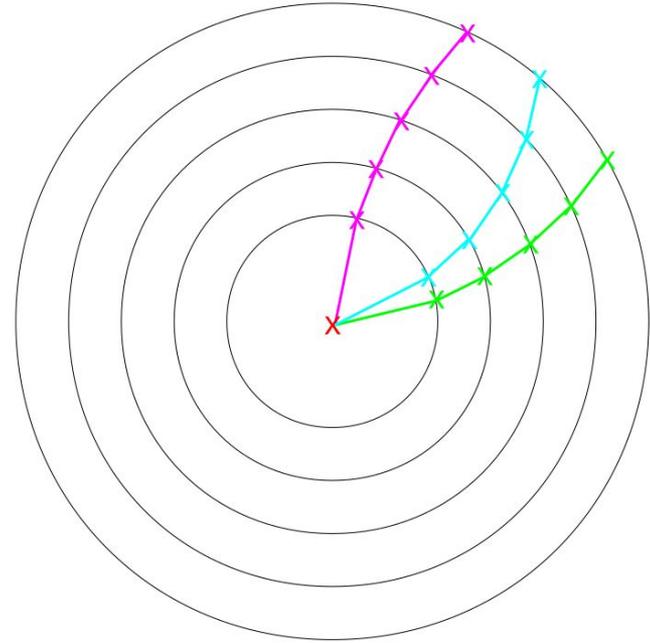
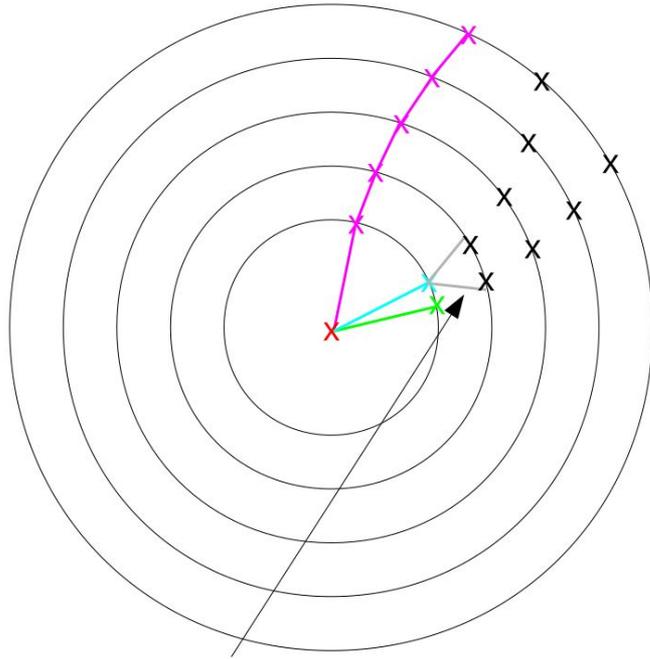
The next collision happens here:

$$25 \text{ ns} \cdot c \approx 7.5 \text{ m}$$

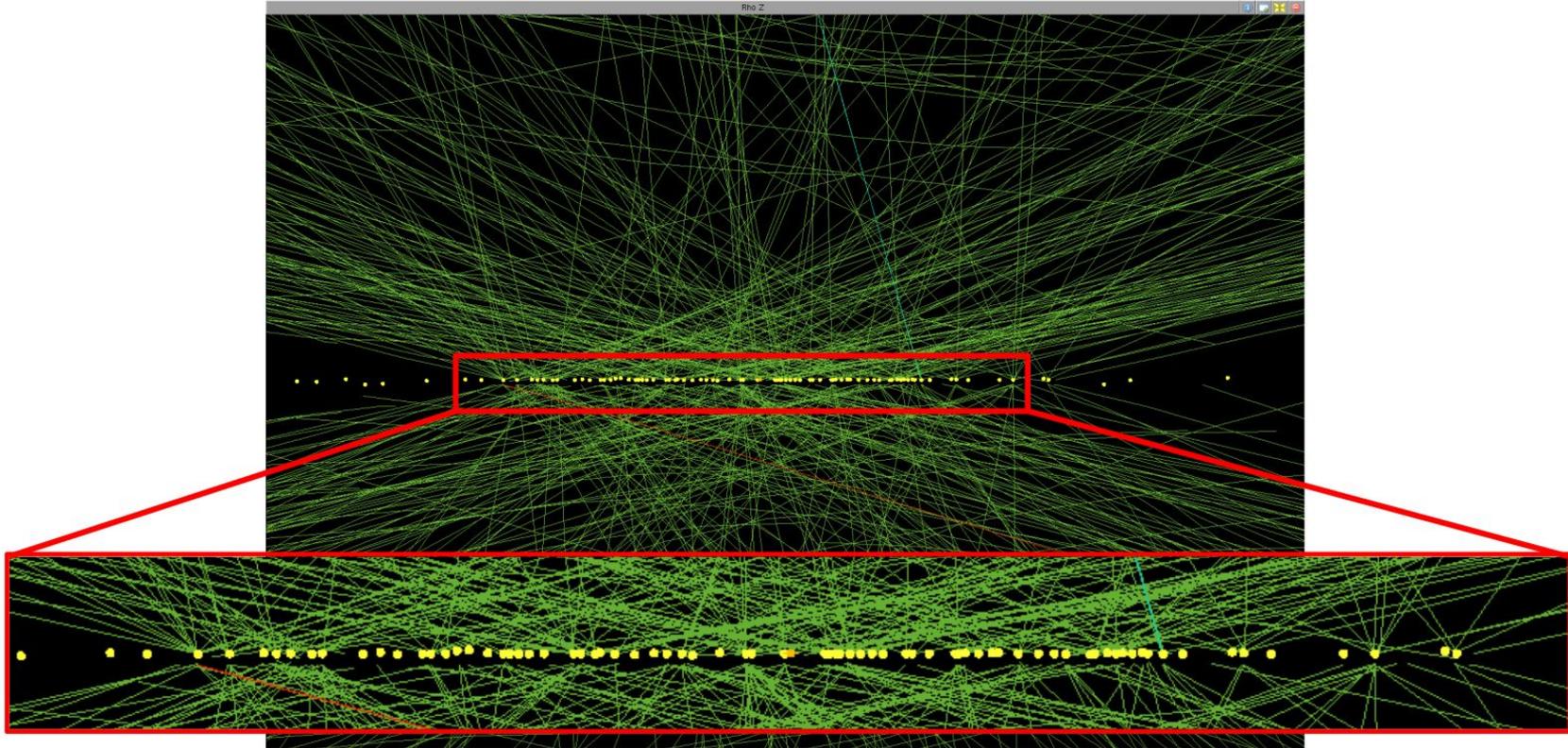
# Track reconstruction

---

# Track reconstruction: find hits that belong to track

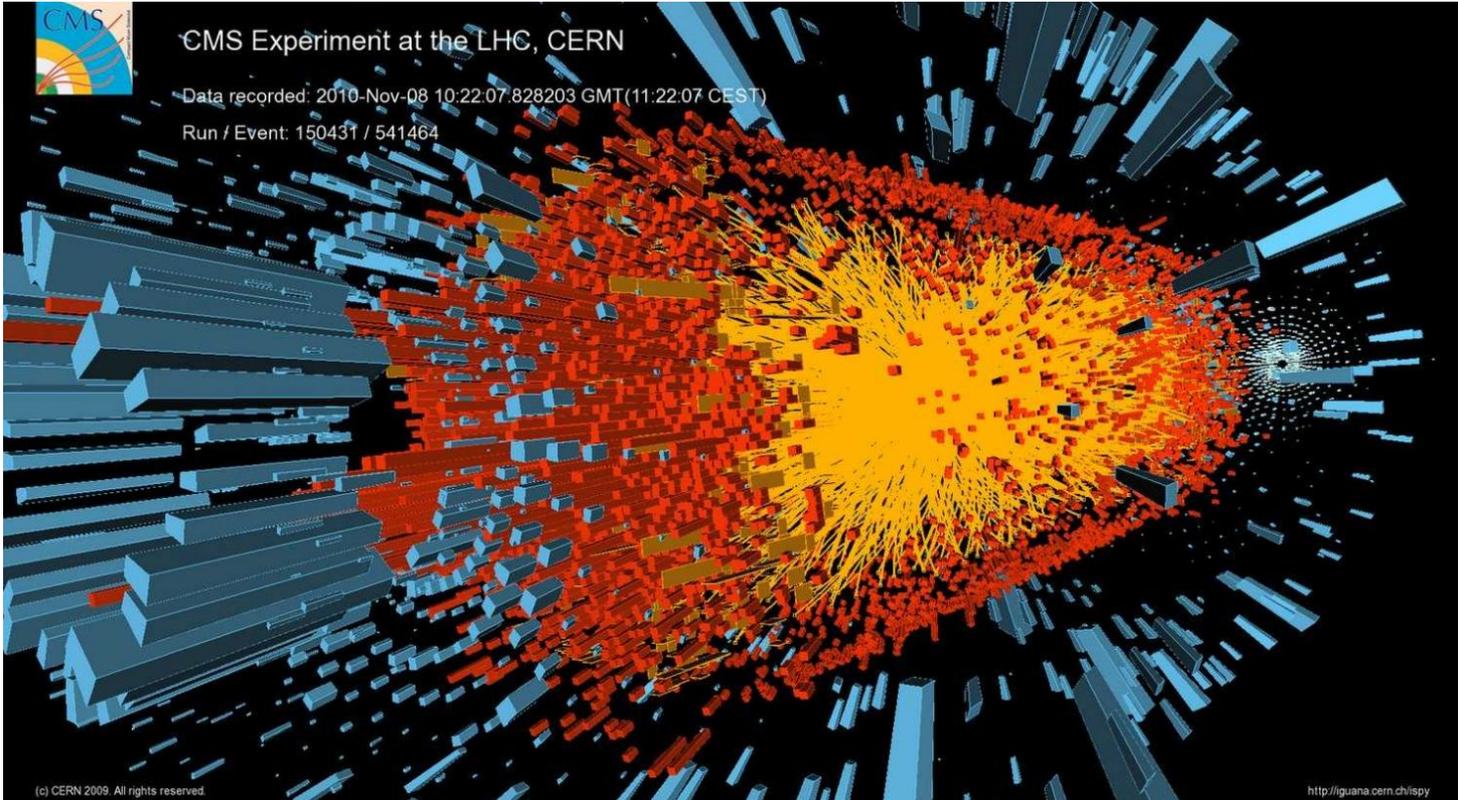


What if 78 interactions happen simultaneously?



# Or a collision of 2 lead nuclei?

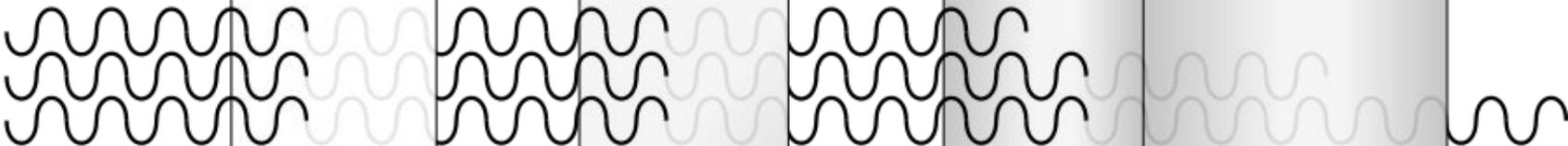
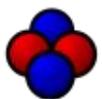
10000  
charged  
tracks!



# Calorimeters

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$\alpha$   
 $\beta$   
 $\gamma$



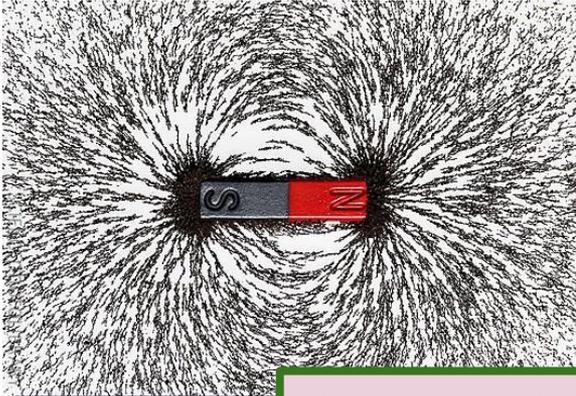
Paper

Aluminium

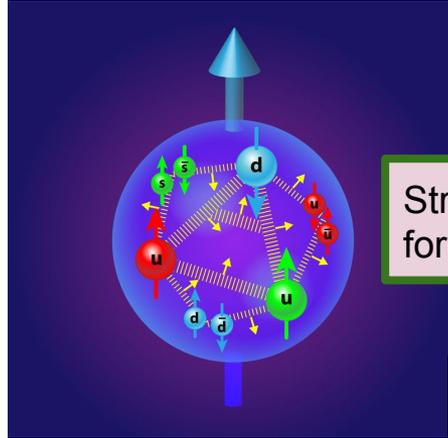
Lead

What are these particles?  
Why do some pass through material and others don't?

# Interactions: four known forces



electromagnetism

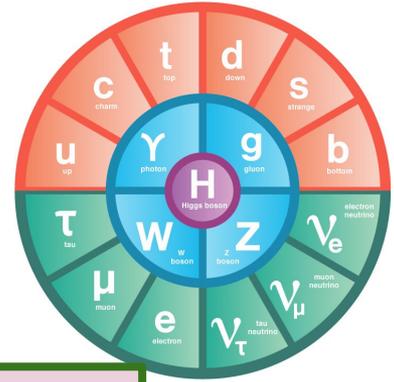
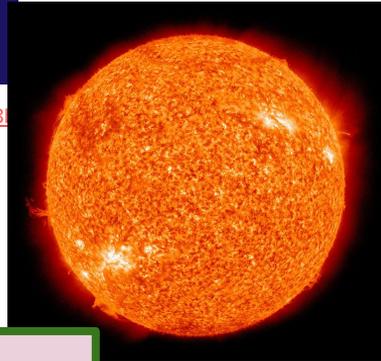


Strong nuclear force

[https://physics.aps.org/assets/89b4f0e0-b8b70d-d90f744d1790/e23\\_2.png](https://physics.aps.org/assets/89b4f0e0-b8b70d-d90f744d1790/e23_2.png)

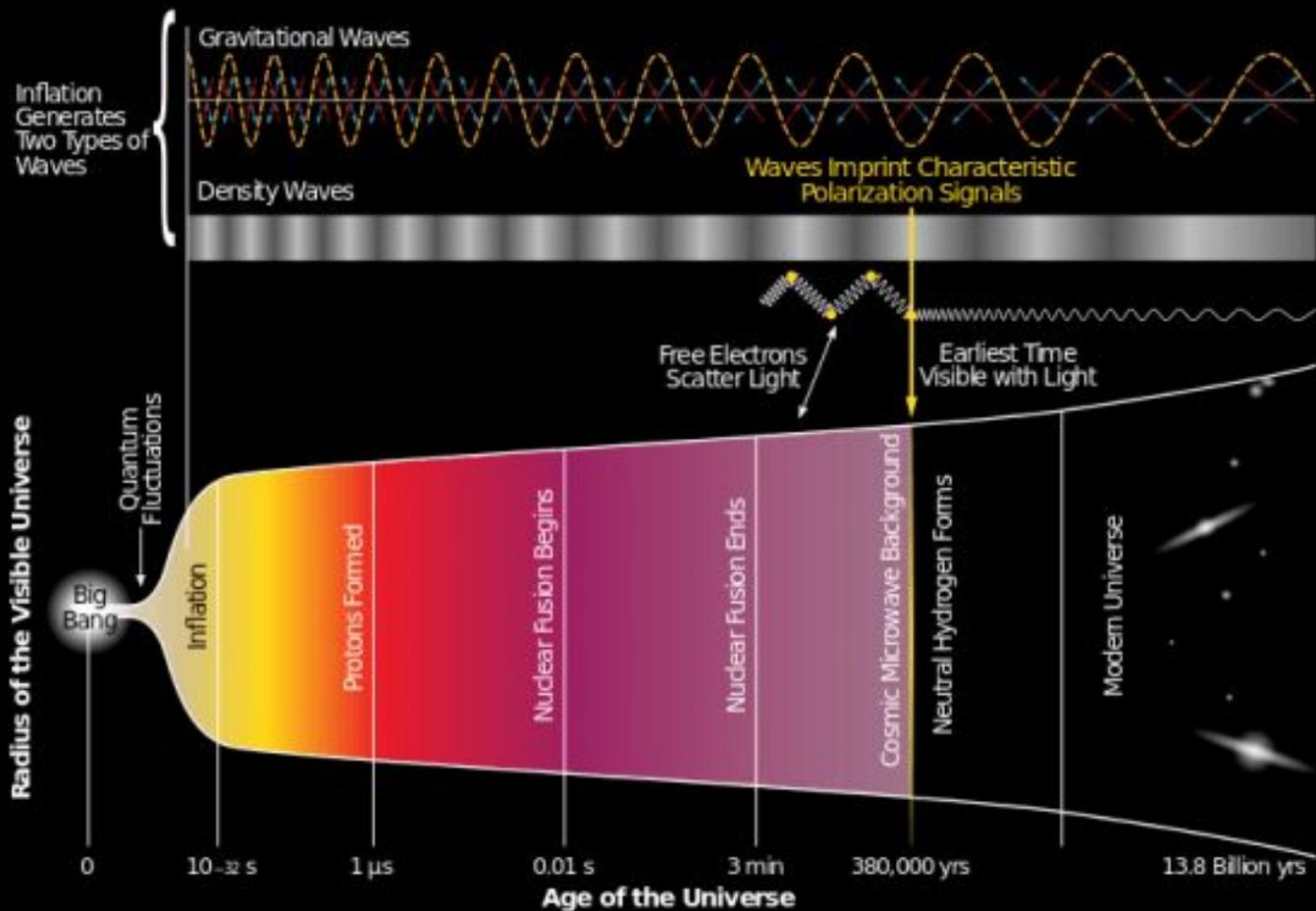
What are we made of?  
How do particles get mass?  
**Gravity is not described by the Standard Model!**

Weak nuclear force



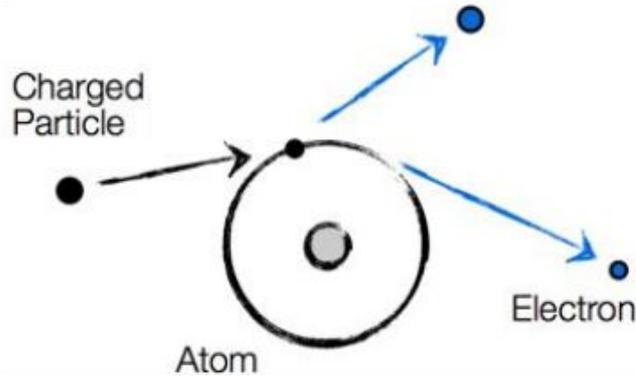
[https://upload.wikimedia.org/wikipedia/commons/thumb/b/b4/The\\_Sun\\_by\\_the\\_Atmospheric\\_Imaging\\_Assembly\\_of\\_NASA%27s\\_Solar\\_Dynamics\\_Observatory\\_-\\_20100819.jpg/800px-The\\_Sun\\_by\\_the\\_Atmospheric\\_Imaging\\_Assembly\\_of\\_NASA%27s\\_Solar\\_Dynamics\\_Observatory\\_-\\_20100819.jpg](https://upload.wikimedia.org/wikipedia/commons/thumb/b/b4/The_Sun_by_the_Atmospheric_Imaging_Assembly_of_NASA%27s_Solar_Dynamics_Observatory_-_20100819.jpg/800px-The_Sun_by_the_Atmospheric_Imaging_Assembly_of_NASA%27s_Solar_Dynamics_Observatory_-_20100819.jpg)

# History of the Universe

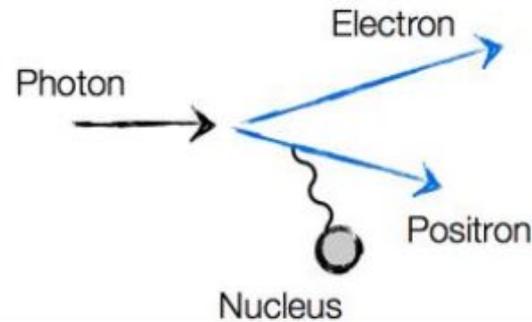


# Interaction with matter: destructive measurement

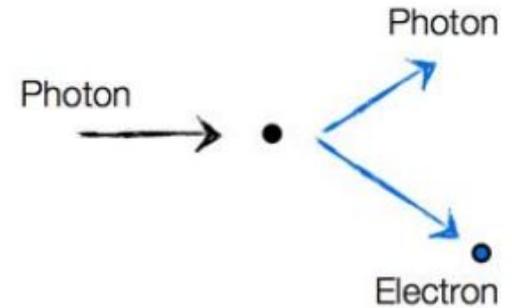
ionisation



Electron-positron  
pair production

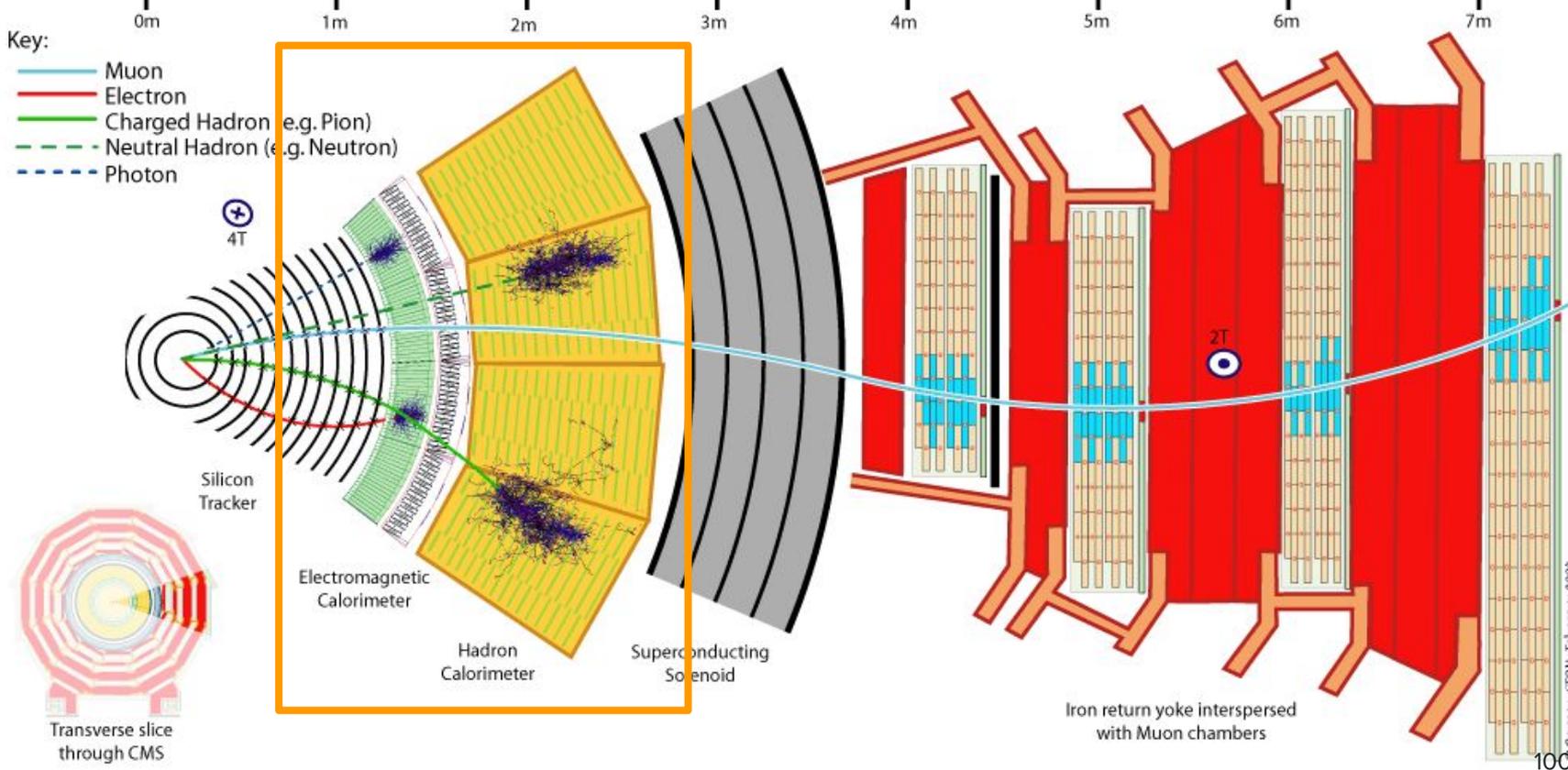


Compton  
scattering



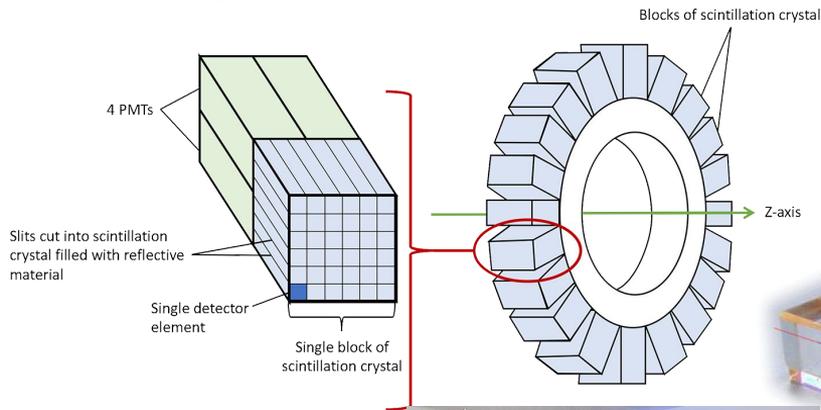
# Detectors at the large hadron collider: onion-like

[http://inspirehep.net/record/82.6852/files/EPS\\_CMS\\_Slice.png](http://inspirehep.net/record/82.6852/files/EPS_CMS_Slice.png)

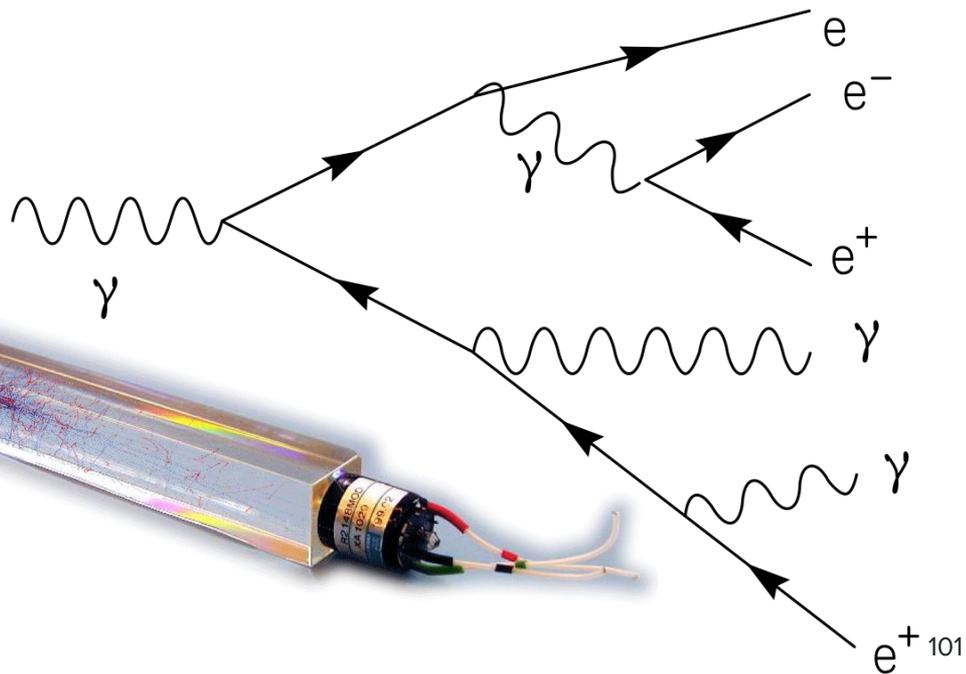


# Electromagnetic calorimeter

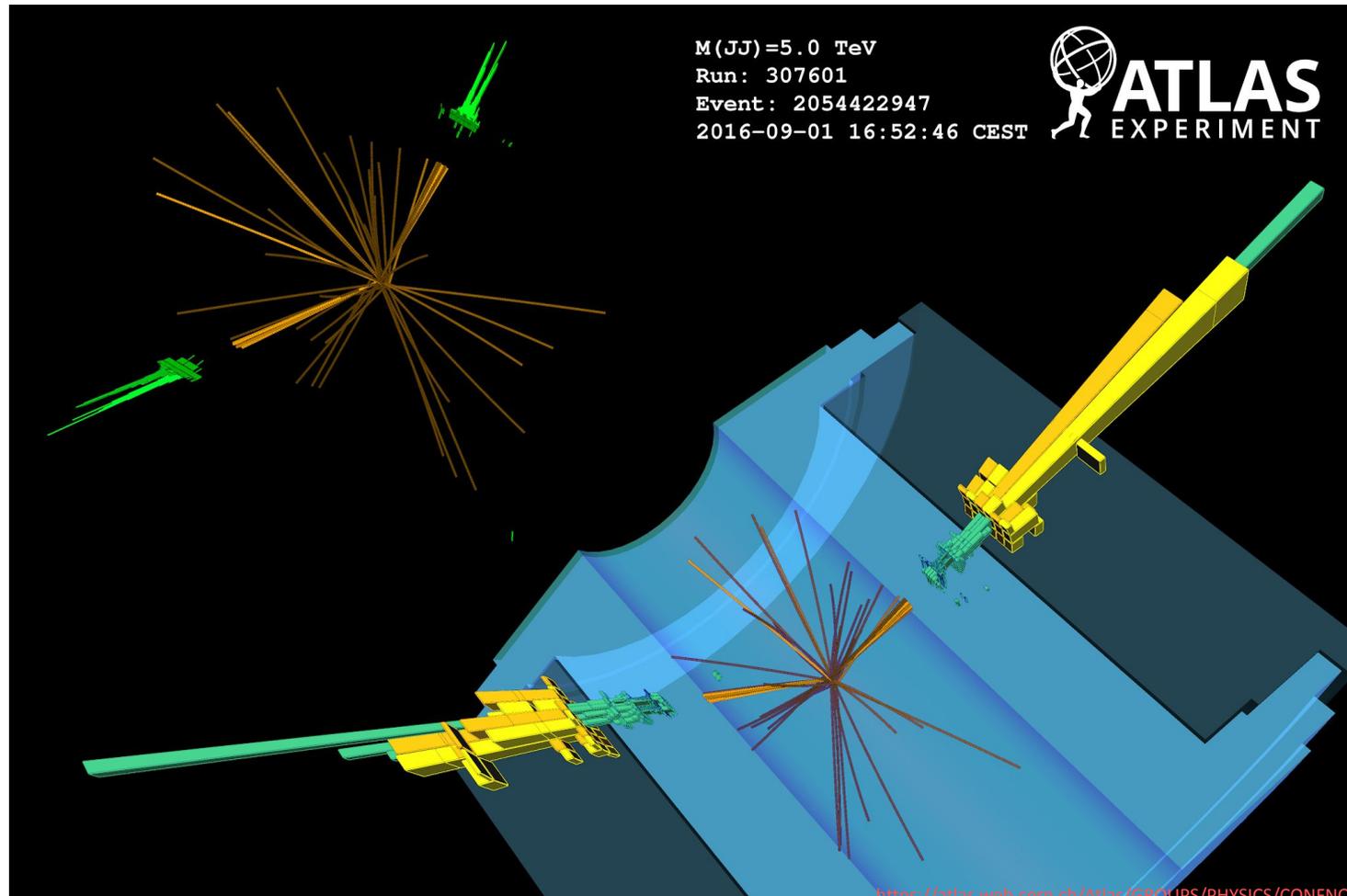
- Electromagnetic shower by interaction with material
- CMS uses scintillating lead tungstate crystals of 1.5 kg that can take 2 days to grow!
- Light detected with a photomultiplier tube



Same crystals used in PET scanners



# Diboson event: jets in the ATLAS detector

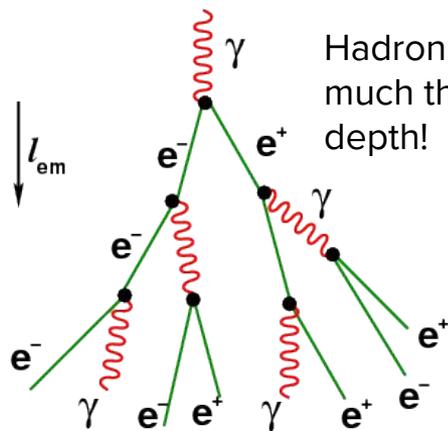


# Hadronic calorimeter

	$\lambda_{\text{int}}$ [cm]	$X_0$ [cm]
Szint.	79.4	42.2
LAr	83.7	14.0
Fe	16.8	1.76
Pb	17.1	0.56
U	10.5	0.32
C	38.1	18.8

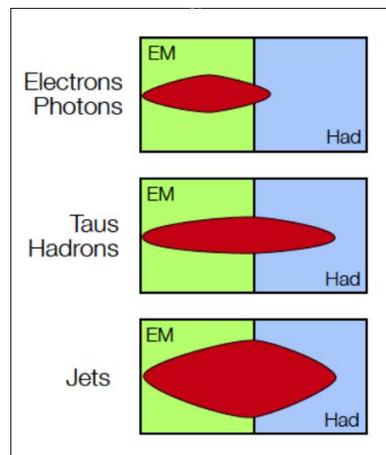
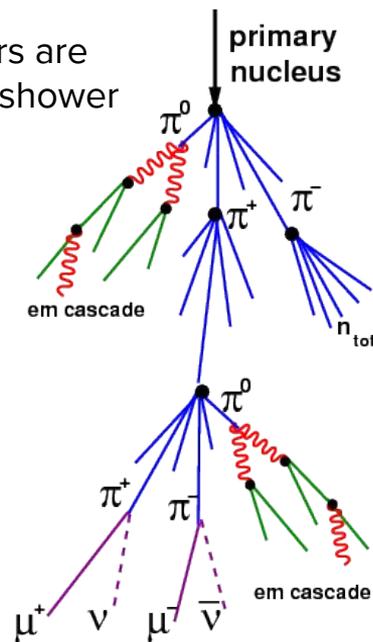
Need to consider fractions of energy in each calorimeter

## em cascade



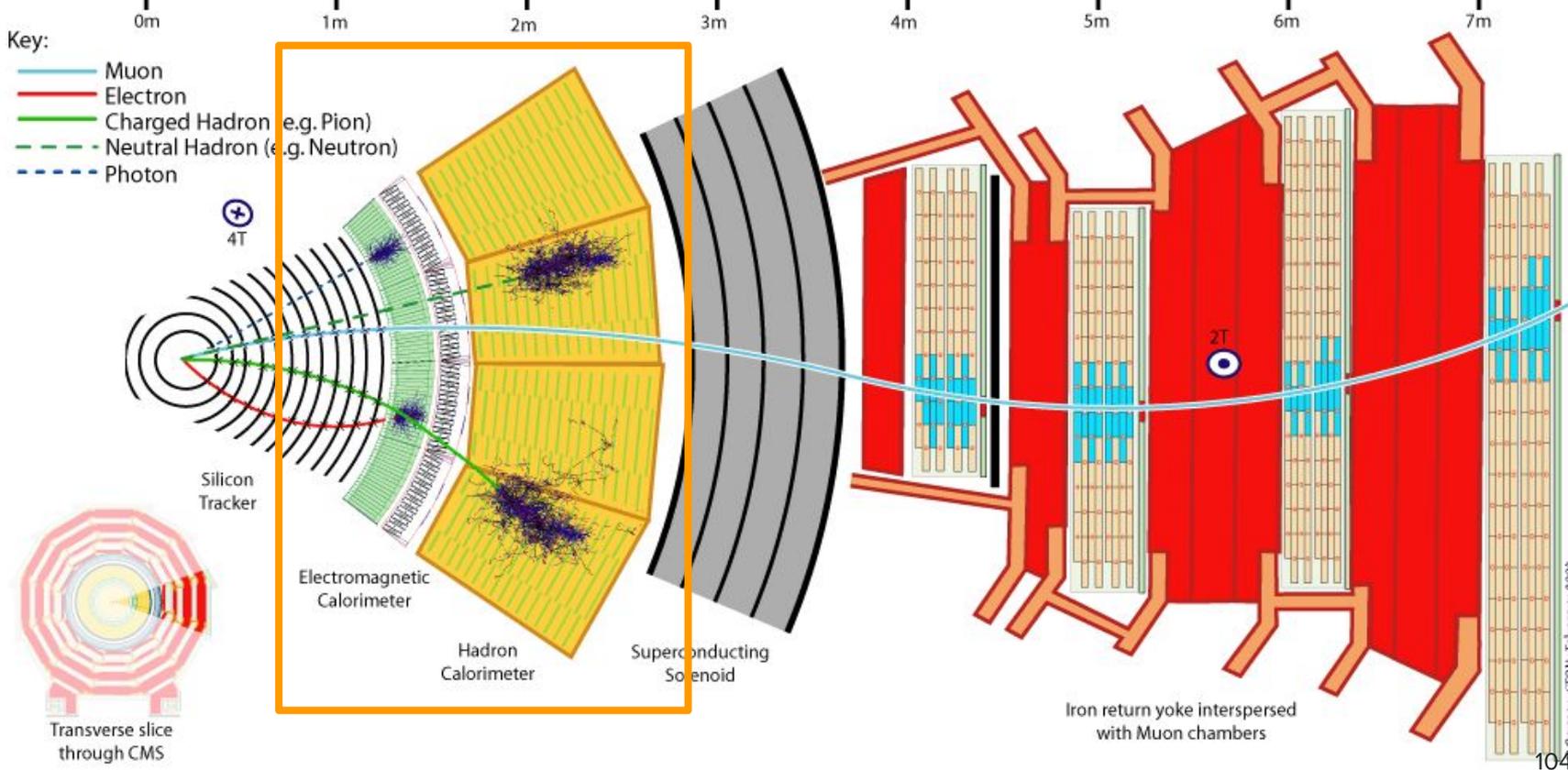
Hadronic calorimeters are much thicker: larger shower depth!

## hadronic cascade



# Detectors at the large hadron collider: onion-like

[http://inspirehep.net/record/82.6852/files/EPS\\_CMS\\_Slice.png](http://inspirehep.net/record/82.6852/files/EPS_CMS_Slice.png)



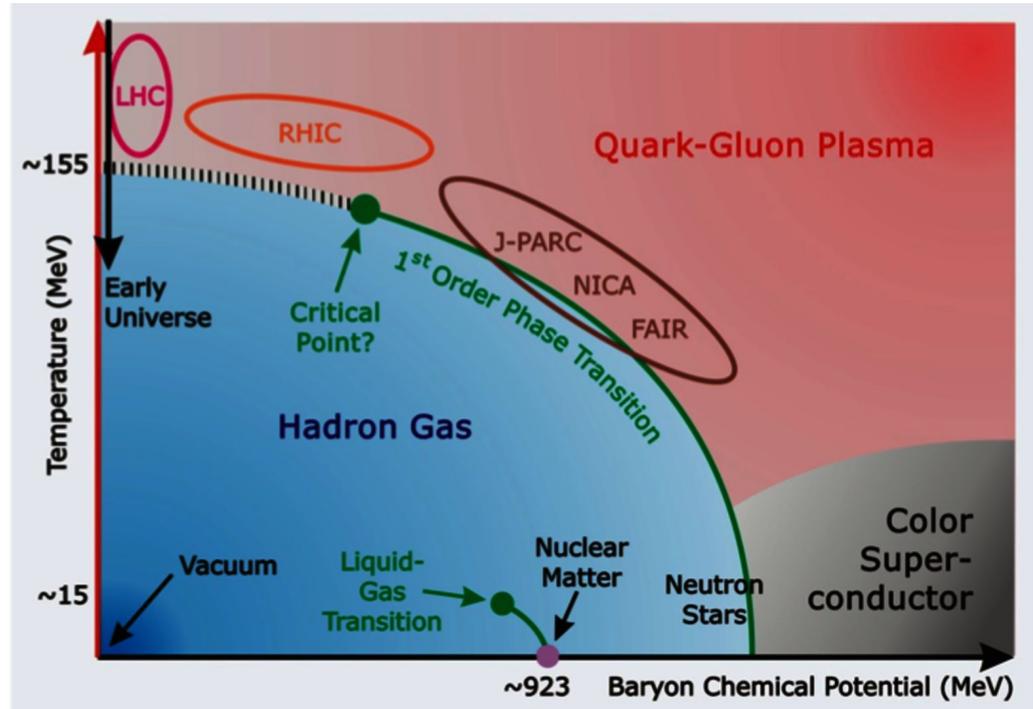
# CMS = compact muon solenoid

De CMS solenoid:

- Is de grootste supergeleidende magneet ooit gemaakt
- Weegt 12000 ton
- Is gekoeld tot 4.65 K, 2 graden warmer dan in de ruimte
- Is 100,000 keer sterker dan het magneetveld op aarde
- Heeft genoeg energie om 18 ton goud te smelten
- Heeft bijna twee keer zoveel ijzer als de Eiffeltoren

# Waarom zware ionen bestuderen?

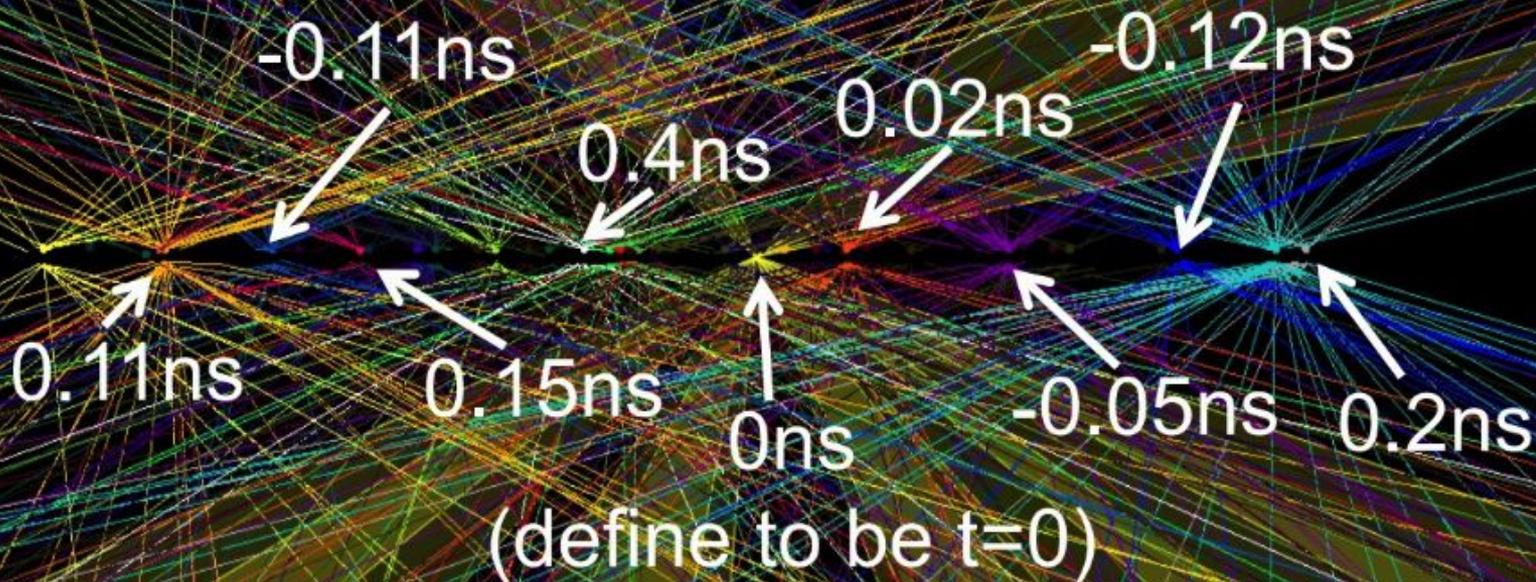
From [gsi.de](http://gsi.de)



CMS

E  
CMS Experiment at LHC, CERN  
Data recorded: Mon May 28 01:16:20 2012 CEST  
Run/Event: 195099 / 35438125  
Lumi section: 65  
Orbit/Crossing: 16992111 / 2295

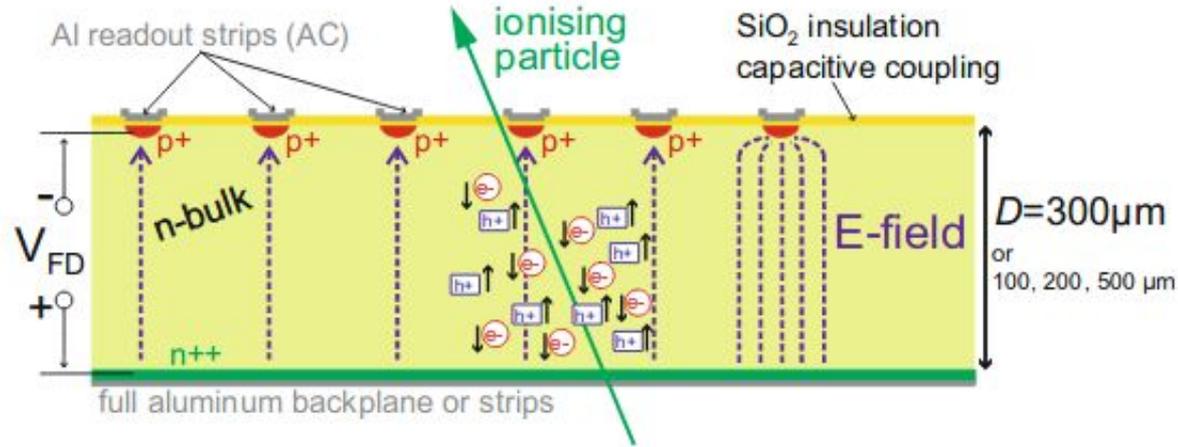
# LHC Bunch Crossing 1ns Clip



Raw  $\Sigma E_T \sim 2$  TeV  
14 jets with  $E_T > 40$   
Estimated PU  $\sim 50$

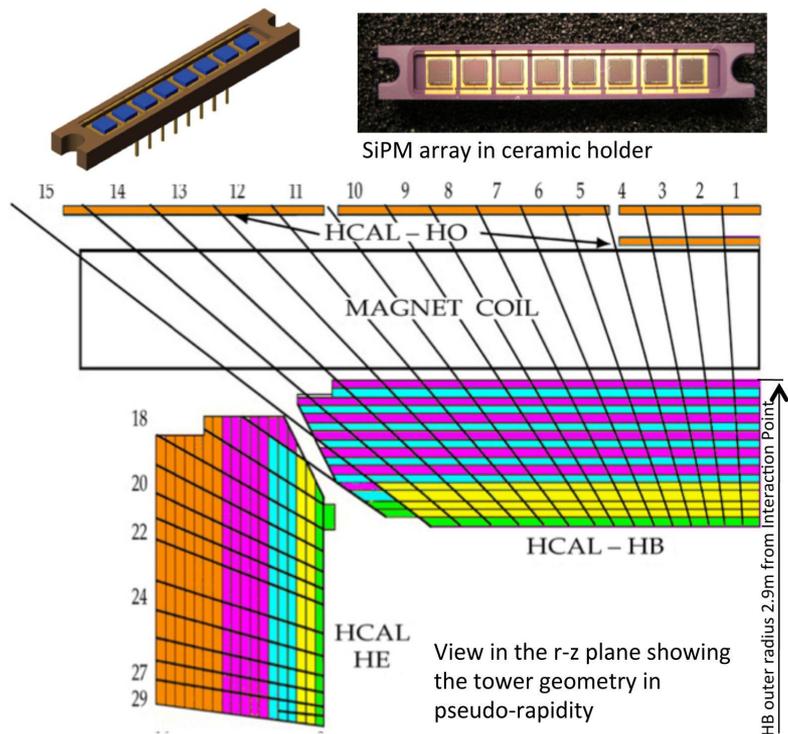
# Ideal signal detection with silicon sensors

- A minimum ionizing particle (MIP) traveling through a fully depleted region ( $V_{FD}$ ) creates electron hole pairs
- The charges drift to opposite directions under the electric field
- Within nanoseconds, charges are collected at the readout



p-in-n silicon sensor

# Photomultipliers → silicon photomultipliers



CMS upgraded their hadronic calorimeter:  
From hit information for 4 cells (1 tower) → hit information per cell