

NON-DECOUPLING NEW PHYSICS

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THE SM EXISTS IN FIELD SPACE



We observe that Higgs physics is SM-like at our vacuum. We *assume* it is SM-like at the EW symmetric vacuum.

DECOUPLING NP GIVES SMALL EFFECTS EVERYWHERE

SMEFT is a Taylor expansion in Φ about $\Phi = 0$.

$$\mathcal{L} \approx |D\Phi|^2 + \frac{1}{\Lambda^2} |\Phi|^2 |D\Phi|^2 + \frac{1}{\Lambda^4} |\Phi|^4 |D\Phi|^2 + \dots$$



GEOMETRIC APPROACHES HELP HERE

(Alonso, Jenkins, and Manohar 2016)



We observe that Higgs physics is SM-like at our vacuum. We *assume* it is SM-like at the EW symmetric vacuum.

DECOUPLING NP (SMEFT) CORRELATES HIGGS OBS.

$$\mathcal{L} \approx |D\Phi|^2 + \frac{1}{\Lambda^2} |\Phi|^2 |D\Phi|^2 + \frac{1}{\Lambda^4} |\Phi|^4 |D\Phi|^2 + \dots$$

As can be seen in the broken phase

$$\begin{split} \mathcal{L} &\to \frac{1}{4} g_2^2 W^+ W^- \left[(v+h)^2 + \frac{1}{2\Lambda^2} (v+h)^4 + \frac{1}{4\Lambda^4} (v+h)^6 + \dots \right] \\ &\to \frac{1}{4} g_2^2 W^+ W^- \left[v^2 \left(1 + \frac{v^2}{2\Lambda^2} + \frac{v^4}{4\Lambda^4} + \dots \right) \right. \\ &\quad + 2vh \left(1 + \frac{v^2}{\Lambda^2} + \frac{3}{4} \frac{v^4}{\Lambda^4} + \dots \right) \\ &\quad + h^2 \left(1 + 3 \frac{v^2}{\Lambda^2} + \frac{15}{4} \frac{v^4}{\Lambda^4} + \dots \right) + \dots \right] \end{split}$$

Note $m_W \to 0$ when $v \to 0$ and correlation ($\kappa_V \approx \kappa_{2V} \approx \frac{v^2}{\Lambda^2}$).

WHEN DOES PHYSICS NOT DECOUPLE?

(Cohen, Craig, Lu, and Sutherland 2021)



Like a Laurent expansion

$$\mathcal{L} = \sum_{k=k_{\min}<0}^{\infty} c_k \frac{|\Phi|^{2k}}{\Lambda^{2k}} |D\Phi|^2$$

$$\mathcal{L} = \sum_{k=0}^{\infty} c_k \frac{|\Phi|^{2k}}{v^{2k}} |D\Phi|^2$$

Like a non-convergent expansion

1) Extra sources of electroweak 2 symmetry breaking t

2) New particles getting most of their mass from the Higgs.

Both cases have particles mass $m \lesssim 4\pi v$.

EXTRA SOURCES OF EWSB

(Englert, Naskar, and Sutherland 2023)



To get an interesting signal in κ_{2V} before κ_{V} , require significant mixing with $\lesssim 1 \,\mathrm{TeV}$ states.

$$\mathcal{L}_{UV} = |\partial \Phi|^2 + \frac{1}{2} (\partial S)^2 - \left(-\mu_{\Phi}^2 |\Phi|^2 + \lambda_{\Phi} |\Phi|^4 + \frac{1}{2} \underbrace{(m^2 + \kappa |\Phi|^2)}_{\equiv m^2(|\Phi|^2)} S^2 + \frac{1}{4} \lambda_S S^4 \right)$$

Choose $m^2, \kappa > 0$ so S does not get a vev

$$\mathcal{L}_{EFT} = |\partial \Phi|^{2} + \frac{1}{384\pi^{2}} \frac{\kappa^{2}}{m^{2} (|\Phi|^{2})} (\partial |\Phi|^{2})^{2} + \mu_{\Phi}^{2} |\Phi|^{2} - \lambda_{\Phi} |\Phi|^{4} + \frac{1}{64\pi^{2}} m^{4} (|\Phi|^{2}) \left(\ln \frac{\mu^{2}}{m^{2} (|\Phi|^{2})} + \frac{3}{2} \right)$$

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EFT CONVERGENCE

See also (Falkowski and Rattazzi 2019)

Expand Higgs potential, V, in SMEFT and HEFT



VIABLE NON-DECOUPLING MODELS

'Loryons' get most of their mass from the Higgs Interactions:



Signals:





pair production

WHITE SPACE MEANS EXPERIMENTALLY VIABLE

(Banta, Cohen, Craig, Lu, and Sutherland 2021)



Disallowed regions in colour: Orange, dotted: κ_{γ} or κ_g Blue, solid: perturb. unitarity $\lambda_{h\Phi}$ Green, dashed: Higgs decay Purple, dot-dash: Direct search

Plots: fraction of mass squared from Higgs (f_{max}) vs. total mass.

THESE MODELS PRODUCE A STRONGLY FIRST ORDER EWPT



Orange, dotted:

 κ_γ or κ_g expt. constraints

Blue, solid: perturb. unitarity

Green, dashed: strongly first-order phase transition

Red, solid

lower bound for stochastic gravitational wave background @ LISA

...a necessary condition for electroweak baryogenesis.

At HL-LHC, κ_g rules out coloured particles, κ_γ makes inroads, κ_λ approaches unitarity bound.



Orange, dotted: κ_{γ} or κ_{g} Blue, solid: perturb. unitarity $\lambda_{h\Phi}$ Green, dashed: Higgs cubic Purple, dot-dash: Direct search

Nightmare scenario of neutral scalar singlet remains open. $\kappa_{\lambda} \sim 5\%$ measurement of FCC-hh closes off everything.

THE FUTURE: LEPTONIC (1)



THE FUTURE: LEPTONIC (2)



Nature is SM-like at v = 246 GeV, but may be wildly different at v = 0, due to non-decoupling physics.

Gives generic signals in κ_h and κ_λ



Non-decoupling NP is a finite target space for future colliders

THANKS

BACKUP

HEFTY MODELS POORLY FIT BY DIMENSION 6 SMEFT

HL-LHC could probe the correlations of a single SMEFT operator across different Higgs multiplicities. (Henning, Lombardo, Riembau, and Riva 2019)

These correlations may be broken.





HEFTY PHYSICS BREAKS CORRELATIONS

(Abu-Ajamieh, Chang, Chen, and Luty 2020)

$$\begin{split} \mathcal{L} &= \mathcal{L}_{\rm SM} - \delta_3 \frac{m_h^2}{2v} h^3 - \delta_4 \frac{m_h^2}{8v^2} h^4 - \sum_{n=5}^{\infty} \frac{c_n}{n!} \frac{m_h^2}{v^{n-2}} h^n + \cdots \\ &+ \delta_{Z1} \frac{m_Z^2}{v} h Z^{\mu} Z_{\mu} + \delta_{W1} \frac{2m_W^2}{v} h W^{\mu +} W_{\mu}^- + \delta_{Z2} \frac{m_Z^2}{2v^2} h^2 Z^{\mu} Z_{\mu} + \delta_{W2} \frac{m_W^2}{v} h^2 W^{\mu +} W_{\mu}^- \\ &+ \sum_{n=3}^{\infty} \left[\frac{c_{Zn}}{n!} \frac{m_Z^2}{v^n} h^n Z^{\mu} Z_{\mu} + \frac{c_{Wn}}{n!} \frac{2m_W^2}{v^n} h^n W^{\mu +} W_{\mu}^- \right] + \cdots \\ &- \delta_{t1} \frac{m_t}{v} h \bar{t} t - \sum_{n=2}^{\infty} \frac{c_{tn}}{n!} \frac{m_t}{v^n} h^n \bar{t} t + \cdots \\ &- \delta_{t1} \frac{m_t}{v} h \bar{t} t - \sum_{n=2}^{\infty} \frac{c_{tn}}{n!} \frac{m_t}{v^n} h^n \bar{t} t + \cdots \\ \frac{P \operatorname{rocess}}{h^{2Z \to B^2}} \frac{\times \frac{g_{tr}}{152v^{4v}}}{\left[4\delta_{V1} - 2\delta_{V2} + \frac{1}{2}c_{V3} \right]} \\ h^{2Z \to A^2} U^{WW} - \frac{-\frac{1}{\sqrt{2}} \left[4\delta_{V1} - 2\delta_{V2} + \frac{1}{2}c_{V3} \right]}{h^{2W \to WWWW}} - \frac{1}{\sqrt{2}} \left[4\delta_{V1} - 2\delta_{V2} + \frac{1}{2}c_{V3} \right]} \\ h^{2W \to WWWW} + \left[36\delta_{V1} - 13\delta_{V2} + \frac{1}{2}c_{V3} \right] \\ h^{2W \to WWWW} + hWW + \left[36\delta_{V1} - 13\delta_{V2} + 2c_{V3} \right] \\ hW^{WW \to MWWW} - \frac{1}{\sqrt{2}} \left[36\delta_{V1} - 13\delta_{V2} + 2c_{V3} \right]}{hW^{WW} - \delta MW^{WW}} - \frac{1}{\sqrt{2}} \left[36\delta_{V1} - 13\delta_{V2} + 2c_{V3} \right] \\ hW^{WW \to MWWW} + \left[28\delta_{V1} - 13\delta_{V2} + 2c_{V3} \right] \\ hW^{WW \to MWWW} - \frac{1}{\sqrt{2}} \left[36\delta_{V1} - 13\delta_{V2} + 2c_{V3} \right]}{\sqrt{2}} \\ \frac{1}{\delta_{Z} \to hW^{WW}} - \frac{1}{\sqrt{2}} \left[32\delta_{V1} - 11\delta_{V2} + \frac{1}{2}c_{V3} \right] \\ \frac{1}{\delta_{Z} \to hW^{WW}} - \frac{1}{\sqrt{2}} \left[32\delta_{V1} - 11\delta_{V2} + \frac{1}{2}c_{V3} \right] \\ \frac{1}{\delta_{Z} \to hW^{WW}} - \frac{1}{\sqrt{2}} \left[32\delta_{V1} - 11\delta_{V2} + \frac{1}{2}c_{V3} \right]}{\frac{1}{\delta_{Z} \to hW^{WW}}} - \frac{1}{\sqrt{2}} \left[32\delta_{V1} - 11\delta_{V2} + \frac{1}{2}c_{V3} \right] \\ \frac{1}{\delta_{Z} \to hW^{WW}} - \frac{1}{\sqrt{2}} \left[32\delta_{V1} - 11\delta_{V2} + \frac{1}{2}c_{V3} \right] \\ \frac{1}{\delta_{Z} \to hW^{WW}} - \frac{1}{\sqrt{2}} \left[32\delta_{V1} - 11\delta_{V2} + \frac{1}{2}c_{V3} \right]}{\frac{1}{\delta_{Z} \to hW^{WW}}} - \frac{1}{\sqrt{2}} \left[32\delta_{V1} - 11\delta_{V2} + \frac{1}{2}c_{V3} \right] \\ \frac{1}{\delta_{Z} \to hW^{WW}} - \frac{1}{\sqrt{2}} \left[32\delta_{V1} - 11\delta_{V2} + \frac{1}{2}c_{V3} \right] \\ \frac{1}{\delta_{Z} \to hW^{WW}} - \frac{1}{\sqrt{2}} \left[32\delta_{V1} - 11\delta_{V2} + \frac{1}{2}c_{V3} \right] \\ \frac{1}{\delta_{Z} \to hW^{WW}} - \frac{1}{\sqrt{2}} \left[32\delta_{V1} - 10\delta_{V2} + \frac{1}{2}c_{V3} \right] \\ \frac{1}{\delta_{Z} \to hW^{WW}} - \frac{1}{\sqrt{2}} \left[32\delta_{V1}$$

HEFTY MODELS PRODUCE A STRONGLY FIRST ORDER EWPT

(Banta 2022)



Orange, dotted: κ_{γ} or κ_{g} expt. constraints Blue, solid: perturb. unitarity

Green, dashed:

strongly first-order phase transition

Red, solid

lower bound for stochastic gravitational wave background @ LISA

HEFTY (CUSTODIALLY SYMMETRIC) FERMIONS

(Banta, Cohen, Craig, Lu, and Sutherland 2021)



Plots: fraction of mass from Higgs (f_{max}) vs. total mass. Assuming no mass splitting among components of multiplet m_z^2 pieces in s



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