Cosmology from ACT DR6 lensing cross-correlated with DES Y3 galaxies

(in prep work)

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ACT & DES collaborations

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Introduction



This talk: Challenge of astrophysical contamination

Is growth of structure consistent over time?



Growth of structure from cosmological analyses



Why CMB lensing cross-correlations?

Model validity L = 4002.00 L = 80**CMB** lensing 1.75 1.50 1.25 Is there new physics or High z-bir **Redshifts** probed systematic effects? 1.00 0.75 0.50 Low z-bin 0.25 10^{-2} 10^{-1} 10^{-3} 100 10^{1} Modes probed

Challenge of astrophysical contamination

High-resolution CMB surveys are also LSS surveys, especially at the temperature channel





Contamination lowers amplitude of of cross-correlations!

CMB lensing reconstruction





Large lens modulates small scale CMB power spectrum -> look at shifts in the power spectrum to reconstruct the lens!

CMB lensing reconstruction





CMB lensing estimator thinks there is an underdensity! Galaxy on overdensity x underdensity -> **Negative Signal**

Extragalactic foregrounds: sims based



Improving mass maps for future cross-correlations



With future cross-correlations that include Simons Observatory, will need to use smaller and smaller scales.

Need a combination of methods to have competitive measurements.

Plus, better simulations.

Ongoing cross-correlations

DESI LRGs: Kim & Sailer et al. Sailer & Kim et al. Hang & Qu et al.

SDSS BOSS: Wenzl et al.

DES-Y3: Shaikh & Harrison et al. Pitocco et al.

DES-Y3 voids: Marques et al.

Planck CIB: Mheta et al.

Planck/ACT tSZ: Bolliet et al.



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