# FAST ION ACCELERATION IN 3D HYBRID SIMULATIONS OF QUASI-PERPENDICULAR SHOCKS

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### Particle acceleration at non-relativistic shocks

- Understanding particle acceleration at non-relativistic shocks is important for the origin of CRs.
- Energization via first order Fermi acceleration.
- Wide range of Mach numbers (Alfvénic  $M_A = v_{sh}/v_A \rightarrow M$ ) and relative inclination ( $\theta$ ) between the shock velocity and the unperturbed  $B_0$ .

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## Kinetic simulations

- PIC simulations: consist in iteratively moving particles on a grid according to the Lorentz force and selfconsistently adjusting the electromagnetic fields.
- Hybrid simulations treat  $e^-$  as a massless neutralizing fluid and ions as particles.
- No self-consistent kinetic simulation has reported large non-thermal tails of ions at quasi-perpendicular shocks ( $\theta \sim 80^{\circ}$ ).





### Measurements of efficient particle acceleration at quasi-perpendicular shocks

- •SN1006 shows a bilateral symmetry, correlated with the geometry of the background magnetic field.
- •Observations of SN1006 show a radio emission azimuthally symmetric (electrons at GeV energies).
- •Measurements of efficient ion acceleration in the quasi-perpendicular regions of Earth's Bow Shock  $(\theta > 45^{\circ}, M < 20).$





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## 2D and 3D simulations

- •2D-3D simulations (dHybrid, Gargaté et al. 2007)  $\theta = 80^{\circ}$ .
- •2D in-plane: downstream magnetic field amplification.
- •3D simulation: structures in the downstream similar to 2D in plane.

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2D and 3D simulations spectra



## Phenomenological implications

- We found large ion acceleration in kinetic simulation of quasiperpendicular shocks for the first time. Also *e*<sup>-</sup> should be injected.
- $e^-$  acceleration at  $\theta = 80^\circ$  could explain the radio emission ( $e^$ at GeV energies) detected from SN 1006 ( $M \sim 100$ ).
- Mechanism consistent with measurements of efficient ion acceleration at the Earth's Bow Shock (for  $\theta > 45^{\circ}$ ,  $M < 20, \epsilon \lesssim 10^{\circ}$ ).

