

# Cosmological phase transitions in a dimensionally-reduced vector dark matter model

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## Abstract

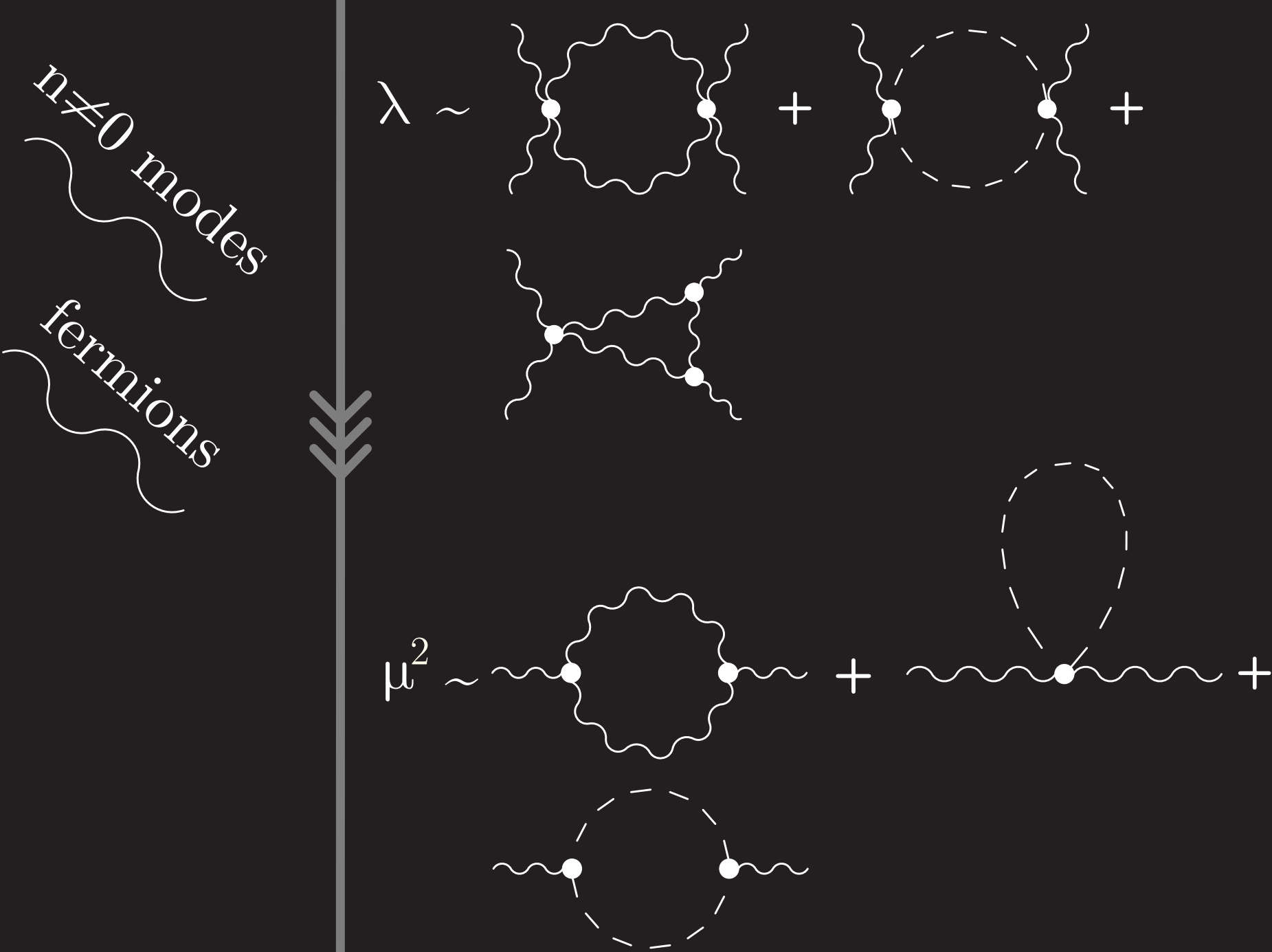
We explore potential gravitational signatures originating from a non-abelian vector dark matter framework, where interactions with the visible sector are mediated via a dark vector-like fermion. We examine the impact of the dark fermion on the phase transition and contrast it with a scenario involving a pure scalar-vector theory. To this effect, we have constructed a dimensionally reduced effective field theory, which has been shown to help mitigate troublesome uncertainties such as nonphysical renormalisation scale and gauge dependence.

## Dimensional reduction

4D Hard Theory ( $T = 0$ )

$$\mathcal{L}_{4D} = -\frac{1}{4}|F_{\mu\nu}^a|^2 + |D_\mu\phi|^2 + \mu^2|\phi|^2 + \lambda|\phi|^4$$

D=4  
 $\Lambda \sim \pi T$



3D Soft Theory ( $m/T < 1$ )

$$\mathcal{L}_S = -\frac{1}{4}|\mathcal{F}_{ij}^a|^2 + |D_i\mathcal{F}_0|^2 + |D_i\varphi|^2 + m_D^2|\mathcal{F}_0|^2 + \mu_3^2|\varphi|^2 + \lambda_3|\varphi|^4 + \lambda_V|\varphi|^2|\mathcal{F}_0|^2 + \lambda_T|\mathcal{F}_0|^4$$

D=3  
 $\Lambda \sim gT$

Temporal scalars

3D Ultra-Soft Theory ( $m/T < 1$ )

$$\mathcal{L}_{us} = -\frac{1}{4}|\bar{\mathcal{F}}_{ij}^a|^2 + |D_i\bar{\varphi}|^2 + \bar{\mu}_3^2|\bar{\varphi}|^2 + \bar{\lambda}_3|\bar{\varphi}|^4$$

D=3  
 $\Lambda \sim g^2T$

## Models

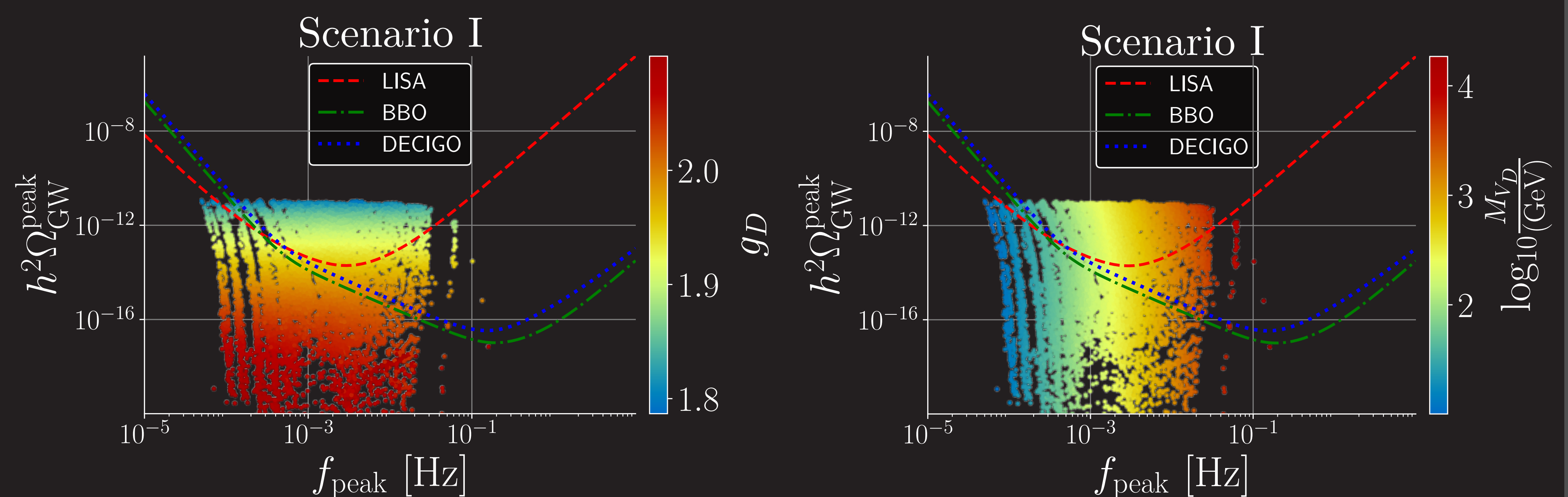
$$\mathcal{L}_{S-I} = -\frac{1}{4}(V_{\mu\nu}^a)^2 + |D_\mu\Phi_D|^2 - \mu_D^2\Phi_D^\dagger\Phi_D - \lambda_D(\Phi_D^\dagger\Phi_D)^2$$

→ **Model parameters:**  $M_{V_D}$ ,  $M_{H_D}$  and  $g_D$

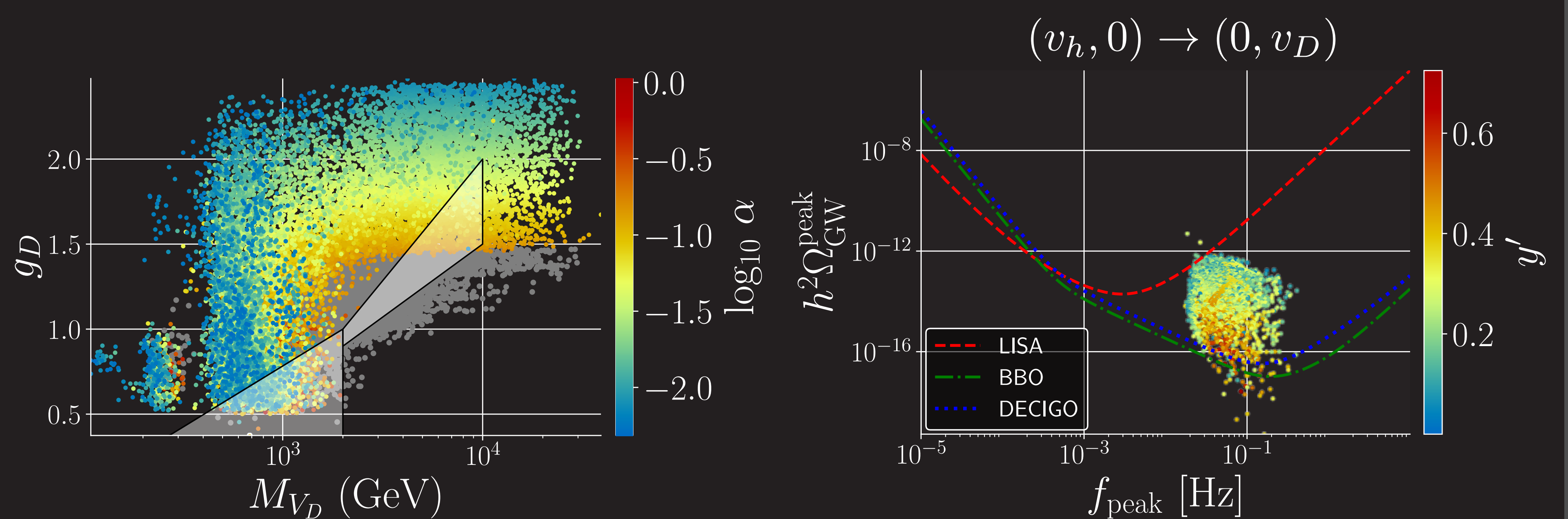
$$\mathcal{L}_{S-II} = \mathcal{L}_{S-I} - \frac{1}{4}(V_{\mu\nu}^i)^2|_{B,W^i,G^i} + \bar{f}^{\text{SM}}_i \not{D} f^{\text{SM}} + |D_\mu\Phi_H|^2 - \mu_H^2|\Phi_H|^2 - \lambda_H|\Phi_H|^4 - \lambda_{HD}|\Phi_H|^2|\Phi_D|^2 - (y\bar{f}_L^{\text{SM}}\Phi_H f_R^{\text{SM}} + y'\bar{\Psi}_L\Phi_D f_R^{\text{SM}} + \text{H.c.})$$

→ **Model parameters:**  $M_{V_D}$ ,  $M_{H_D}$  and  $g_D$ ,  $m_{f_D}$ ,  $m_F$  (physical  $\Psi$  mass),  $\sin\theta_S$  (scalar mixing angle)

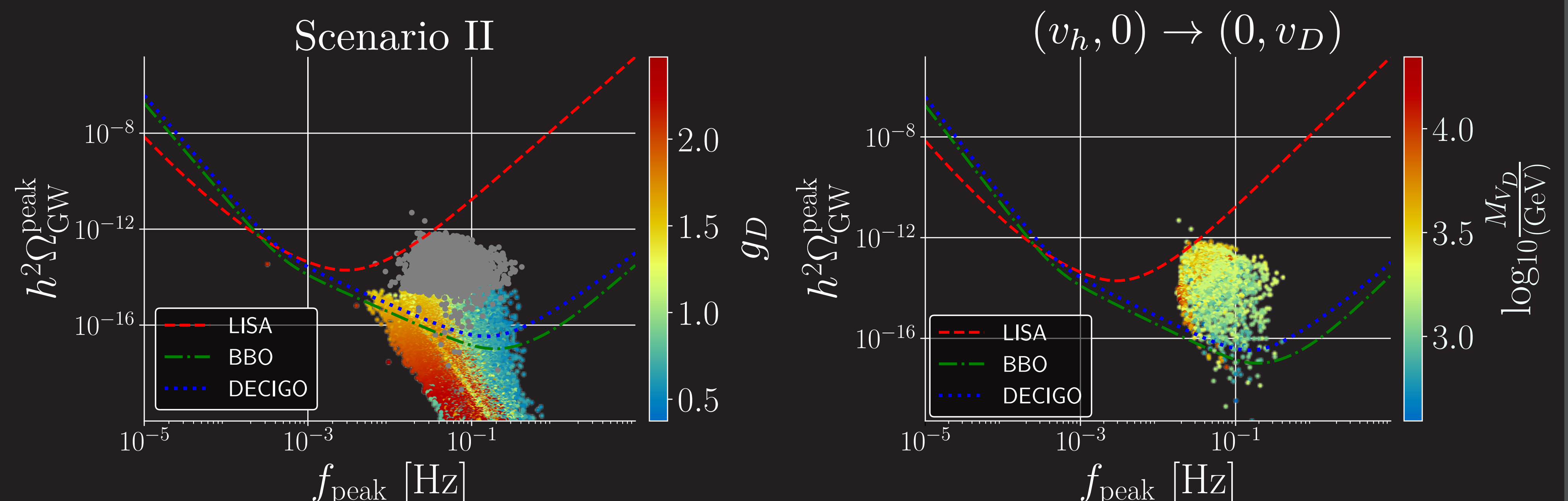
## Numerical results



Scenario I: →  $g_D$  controls the amplitude, while  $M_{V_D}$  controls the frequency.



→ Scenario II: similar behaviour with  $g_D$ .  $y'$  leads to weaker transitions. High- $T$  expansion breaks down ( $m(\phi)/T > 1$ ) for stronger transitions in scenario II.



→  $M_{V_D}$  controls the frequency in qualitative opposite behaviour between scenario I and II.

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