### DARWIN **DARWIN: On the Path to the Ultimate** Liquid-xenon Astroparticle Observatory

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### DARWIN WIMP search with Xenon

## Noble gases Time Projection Chamber

• Leading sensitivity @ mass range O(10-1000) GeV





### DARWIN WIMP search with Xenon

## Noble gases Time Projection Chamber

- Leading sensitivity @ mass range O(10-1000) GeV
- Liquid Xenon (LXe) detectors advantages:

  - SI and SD (<sup>129</sup>Xe, <sup>131</sup>Xe) interactions









# DARWIN XENON-LUX-ZEPLIN-DARWIN: XLZD

## Towards the ultimate LXe detectors

- New collaboration uniting the strengths of major actors (72 institutions and 163 senior scientists)
  - XENONnT, LZ demonstrated experience in largescale LXe TPCs
  - DARWIN Large-scale demonstrators, R&D: electrodes, HV, photosensors,...



xlzd.org



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## XLZD nominal design

- 60 t LXe TPC (~80 t total), early science with 40 t LXe
- 3" PMTs, 1182/array
- 2.98 m diameter, and 2.97 m electron drift (can vary [40, 80] t)
- Drift field: 240 290 V/vm; Extraction field: 6-8 kV/cm
- Double-walled low-background Ti cryostat + LXe "skin" surrounding the TPC







+ Passive and active muon and neutron shielding with gadolinium to enhance capture cross-section (ongoing) R&D with current generation of experiment)





Conceptual drawing of XLZD possible location at LNGS-Hall C, Adrian Schwenck, KIT





#### **Dark Matter**

WIMPs Sub-GeV Inelastic Axion-like particles Planck mass Dark photons

#### <u>Supernovae</u>

Early alert Supernova neutrinos Multi-messenger astrophysics







#### **Neutrino nature**

Neutrinoless double beta decay Neutrino magnetic moment Double electron capture

Physics Case JoPG, 50 013001 (2023)



pp neutrinos Solar metallicity <sup>7</sup>Be, <sup>8</sup>B, hep





#### DARWIN Science Goals

# Simultaneously explore WIMP space down to the "neutrino fog" and search for neutrinoless double- $\beta$ decay of <sup>136</sup>Xe



Design BookarXiv:2410.17137







- Drift/Extraction fields in a larger TPC
  - ➡ High-voltage delivery
  - Electrodes design/construction/test
  - Electric field homogeneity
- Liquid xenon purity
- Background mitigation (external/intrinsic)
- Light collection efficiency
- Photosensors performance





#### DARWIN Towards XLZD: DARWIN

## The DARWIN Collaboration

- ~200 members from 35 institutions
- Established structure and active working groups  $\bigcirc$
- Several large-scale demonstrators, as well as R&D setups





## Rich R&D program to tackle these challenges





DARWIN Xenoscope at UZH

## Vertical demonstrator

- Goals:
  - ➡ Electron drift over 2.6 m, ~400 kg of Xe
  - ➡ Electron cloud diffusion
  - ➡ Custom HV
  - Optical properties of Xe
- Phase 1: purity monitor
  - ➡ 53 cm single phase PM
  - Direct charge readout from electrodes

#### Phase 2: modular TPC

- ➡ 2.6 m dual-phase TPC
- Proportional scintillation light readout with a SiPM tiled array







PE/ns



Area: 57318.53 PE Length: 1.66 µs Position: 2.21 µs Amplitude: 61.30 PE/ns



### DARWIN Pancake at Uni Freiburg

## Full-scale Ø demonstrator

- Test components & concepts:
  - ➡ Test in: LXe, cold GXe, under HV
  - ➡ Probe: sagging, e<sup>-</sup> emission, large-scale cooling
- 5 t stainless steel & double-walled cryostat with 380 kg of xenon
- Flat floor design and possibility of using open top vessel

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## Full-scale Ø demonstrator

- Test components & concepts:  $\bigcirc$ 
  - → Test in: LXe, cold GXe, under HV
  - → **Probe**: sagging, e<sup>-</sup> emission, large-scale cooling
- 5 t stainless steel & double-walled cryostat with 380 kg of xenon
- Flat floor design and possibility of using open top vessel
- Successfully commissioned  $\bigcirc$
- Instrumented with PMTs & cameras  $\bigcirc$
- **Next step:** test of electrodes and HV  $oldsymbol{O}$ performances











# Background mitigation

## To reach background level v-dominated

- Selection of radio-pure materials with low Rn-emanation
  - ➡ Material screening
- Reduce Xe target contamination from impurities
  - ➡ Fast LXe recirculation with radon-free filters and pumps
- Removal of intrinsic background sources
  - →  $^{85}$ Kr distillation → goal of 0.1 ppt natKr already achieved < 0.026 ppt
  - →  $^{222}$ Rn distillation → goal of 0.1 µBq/kg (achieved 0.8) µBq/kg) below ER from solar pp neutrinos. ERC LowRad
  - Coating techniques against radon emanation (electrochemical deposition of Cu)
- Study & Mitigation of accidental coincidences sources
  - Random pairing of isolated S1 and S2 signals



Eur. Phys. J. C (2017) 77:275

Top condense Package tube Input condenser



Eur. Phys. J. C (2022) 82:1104





VULCAN setup: Measure optical properties of materials (Fluorescence, Cherenkov emission,...)



### **DARWIN** Photosensors

## Nominal design with PMTs

- Established technology, low dark count rate (~0.02 Hz/mm<sup>2</sup>), high QE (30-40%)
- Radiopurity improvement on 3" PMTs, but still contribute via several decay chains.
- Testing of Square 2" PMT, lower buoyancy and sub-ns rise time
- Characterisation of SPE response, dark counts, light emission, after pulsing
- R&D & Study of other photosensors



Eur. Phys. J. C (2015) **75**: 546

#### Hamamatsu R12699-406-M4





















# Electrodes & HV Supply

## Efficiency and Robustness

- Electrical Field, optical & Mechanical simulations
  - Effects of electrode geometries on light collection efficiency
  - Mechanical design & stability; 2D/3D simulation studies
- Identification & Treatment of Features
  - Investigate stretching, sagging and flatness of meshes
  - Automatic feature detection with ML and repair with laser welding
  - Electrode surface treatment and coating
- 80 kg LXe TPC with multiple port access for diagnostic of HV components - up to -200 kV bias











#### DARWIN Summary

## DARWIN - R&D efforts towards the ultimate LXe astroparticle observatory (XLZD)

- - ➡ Electric fields
  - ➡ Xenon purity
  - Photosensors
  - Background mitigation

## XLZD (XENON-LZ-DARWIN): new international collaboration

- Aim to build & operate  $\geq$  60t LXe TPC
- Explore WIMP parameter space down to the "neutrino fog"
- $\bigcirc$ processes, and more...



Several large-scale demonstrators, as well as R&D setups to tackle the technical challenges:

Broad physics program with solar & Sn neutrinos,  $0\nu\beta\beta$ -decay and other Double-Weak decay



Back-up 



Precise measurements of electronic solar
neutrino survival probability and electroweak
mixing angle using pp neutrino





# $^{DARWIN}$ $0\nu\beta\beta$ of <sup>136</sup>Xe Projection Sensitivity

# Background consideration for different scenarios

- External bkg from screening  $\rightarrow$  Nominal
- x1/3 reduction factor  $\rightarrow$  Optimistic
- $^{137}$ Xe production at LNGS  $\rightarrow$  Nominal
- <sup>137</sup>Xe production at SURF  $\rightarrow$  Optimistic
- BiPo Tagging efficiency 99.95% → Nominal
- BiPo Tagging efficiency 99.99% → Optimistic
- Energy Resolution @  $Q_{\beta\beta}$  : 0.65%  $\rightarrow$  Nominal
- Energy Resolution @  $Q_{\beta\beta}$  : 0.60%  $\rightarrow$  Optimistic





### DARWIN Other Photosensors

- 12x12 mm<sup>2</sup> MPPC of VUV4 SiPMs
  - ➡ Low radioactivity
  - ➡ Cheaper
  - ➡ Higher buoyancy
  - ➡ Higher dark count rate
- Digital SiPMs
  - ➡ Can turn off single pixels
  - Output already digitised
- LDC VUV SiPMs
- Hybrid sensors (Abalone,...)



#### 48 12x12 mm<sup>2</sup> VUV4 MMPCs @ UZH







#### Map of DCR SPAD3





- Single phase TPC
  - Simplified TPC design, no liquid level control required
  - ➡ Reduce single electron emissions
- Hermetic TPC
  - Prevent radon impurity diffusion into inner volume
- $4\pi$  coverage with photosensors





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