



Hyper-Kamiokande



CHARLES
UNIVERSITY

Status and Prospects of the Hyper-Kamiokande Experiment

Bedřich Roskovec

Charles University, Prague

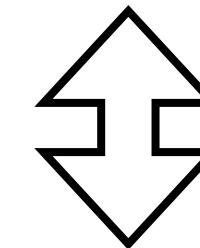
on behalf of the Hyper-Kamiokande collaboration



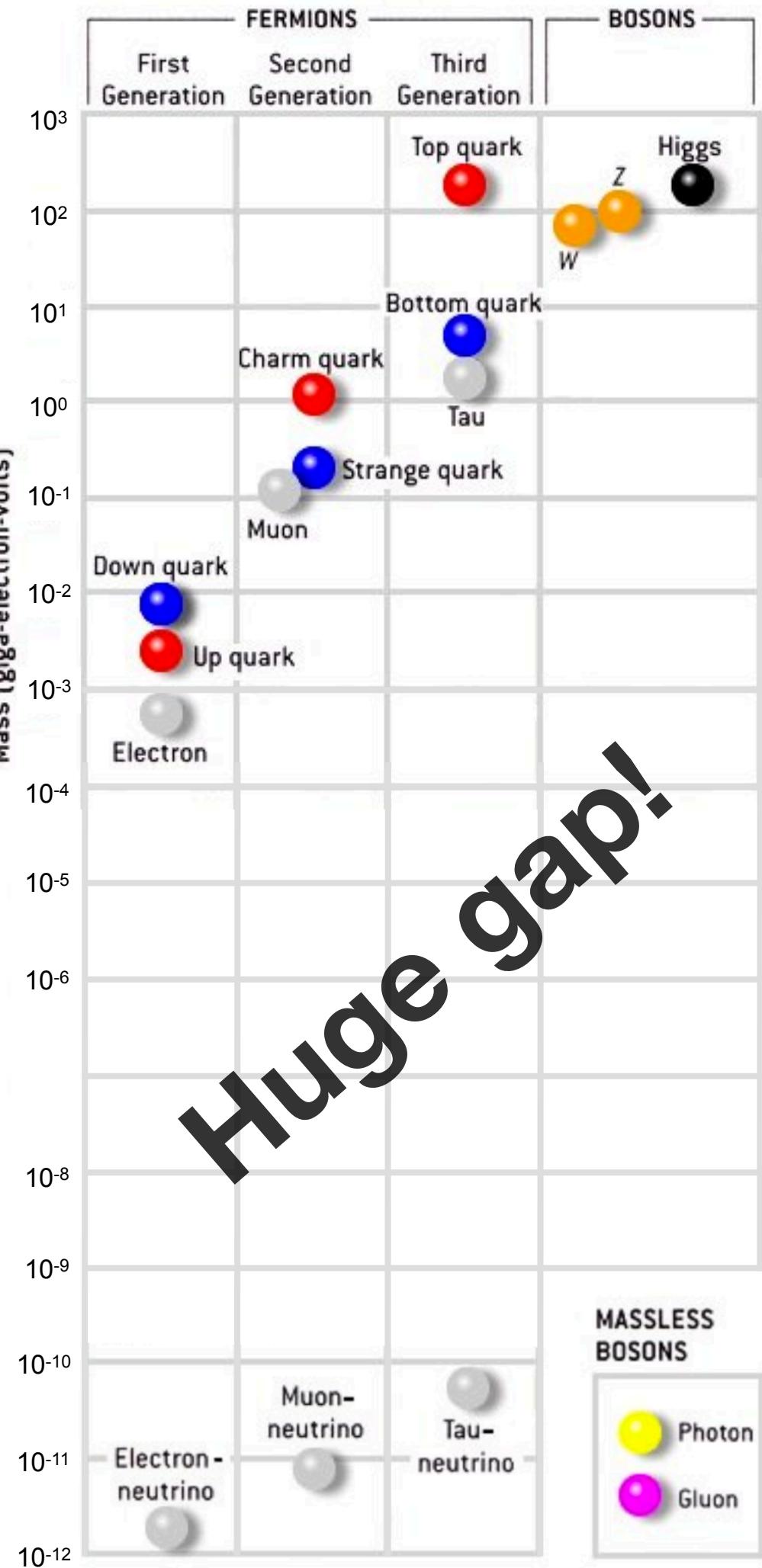
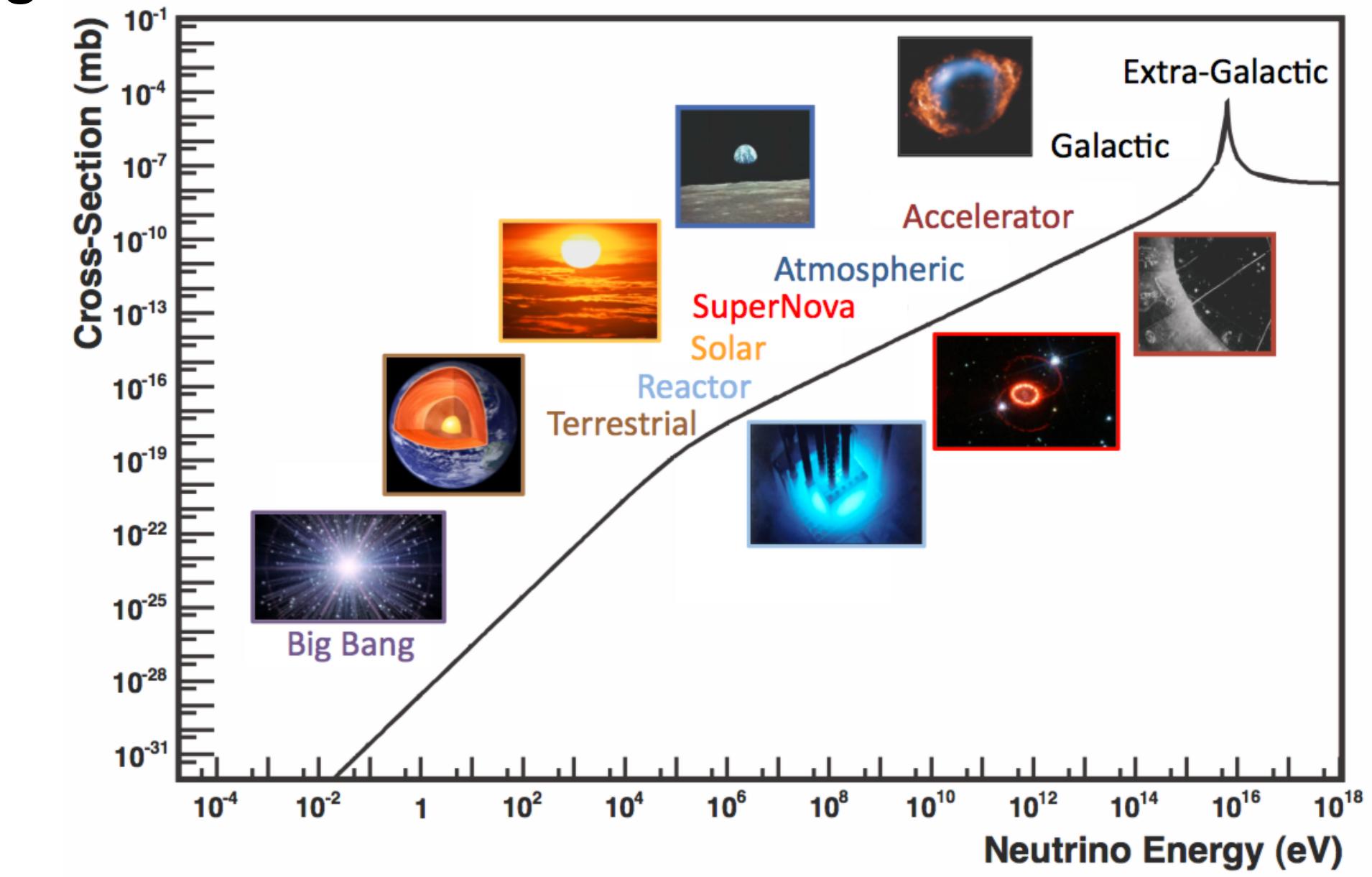
IAS Program on High Energy Physics 2024
HKUST, Hong Kong
24 January 2024

Neutrinos

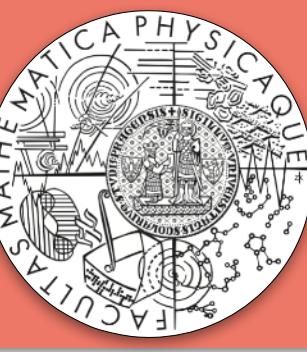
- Second most abundant particle in the universe
- Odd particle of the Standard Model
 - Very light compared to the other particles
 - Non-zero mass - does not originate from the interaction with the Higgs
→ evidence of the physics beyond the Standard Model
- Neutrinos come from various sources with broad range of energies
 - Bearers of key information
- Understanding the neutrinos



Understanding the universe



Basics of Neutrino Mixing and Oscillation (2v)

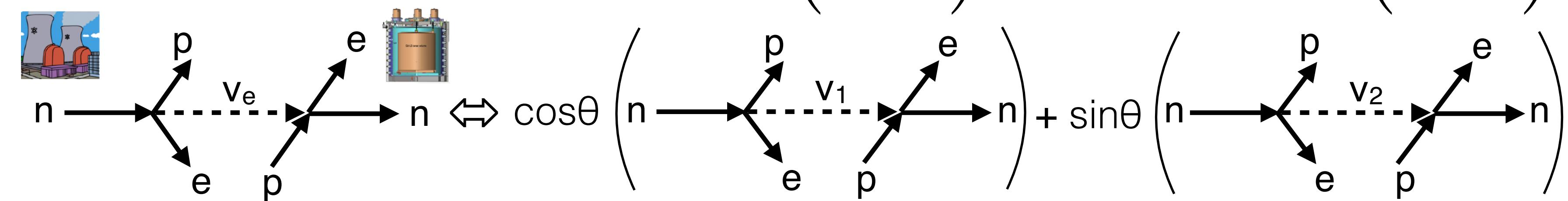


- Neutrino mixing: Flavor states \neq Mass states

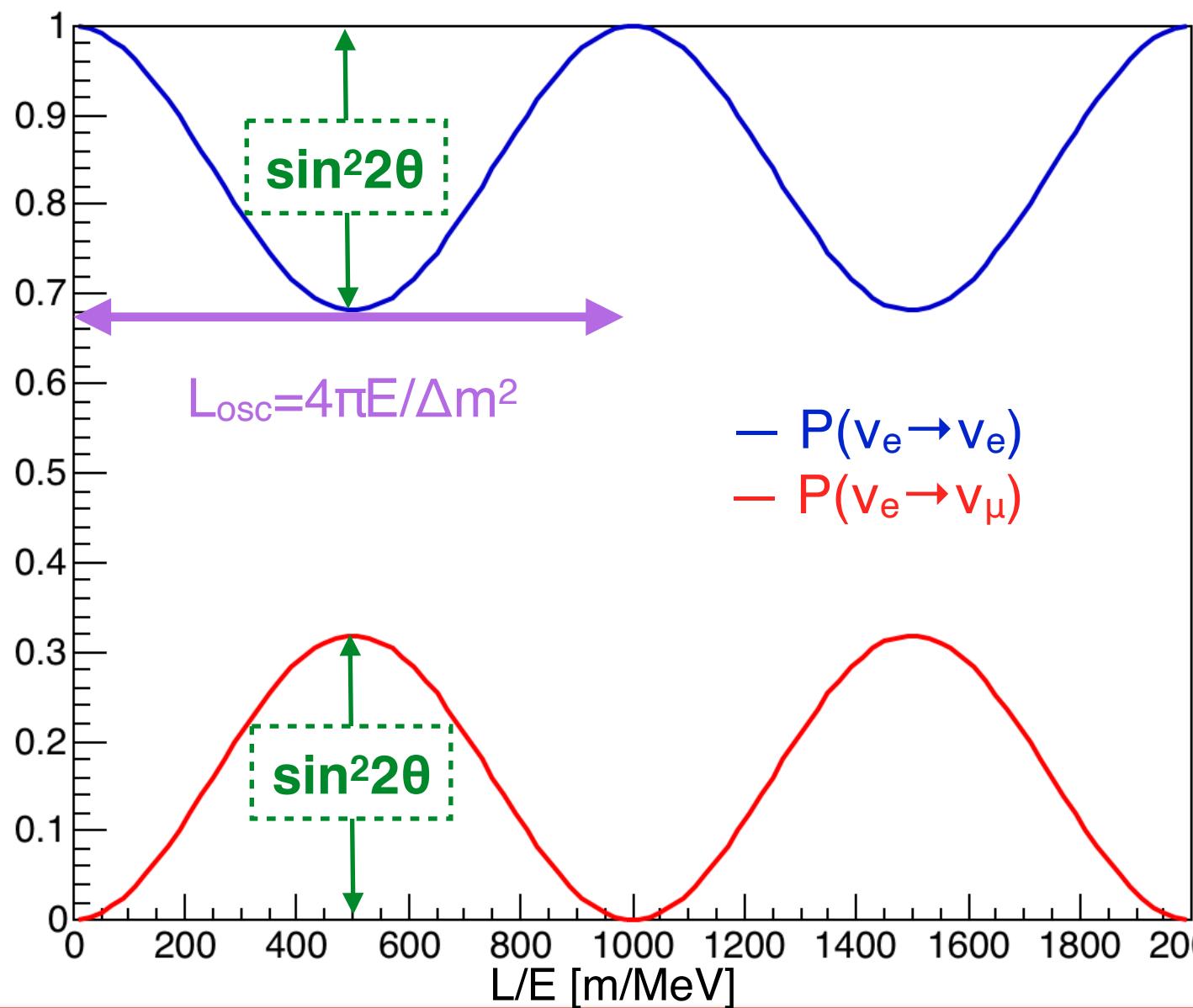
$$|\nu_e\rangle = \cos\theta |\nu_1\rangle + \sin\theta |\nu_2\rangle$$

$$|\nu_\mu\rangle = -\sin\theta |\nu_1\rangle + \cos\theta |\nu_2\rangle$$

$$|\nu_1(L)\rangle = \exp\left(-i\frac{m_1^2 L}{2E}\right) |\nu_1(0)\rangle \quad |\nu_2(L)\rangle = \exp\left(-i\frac{m_2^2 L}{2E}\right) |\nu_2(0)\rangle$$



- Neutrino Oscillation:



Disappearance Channel

$$P_{\nu_e \rightarrow \nu_e}(L, E) = 1 - \sin^2 2\theta \sin^2 \frac{\Delta m^2 L}{4E}$$

$$P_{\nu_e \rightarrow \nu_e} + P_{\nu_e \rightarrow \nu_\mu} = 1$$

Appearance Channel

$$P_{\nu_e \rightarrow \nu_\mu}(L, E) = \sin^2 2\theta \sin^2 \frac{\Delta m^2 L}{4E}$$

Oscillation essentials:

- Probability oscillation in L/E
- Mass squared difference $\Delta m^2 \equiv m_1^2 - m_2^2$
 - drives the frequency
- Mixing angle θ
 - drives the amplitude

Neutrino Mixing and Oscillation (3v)

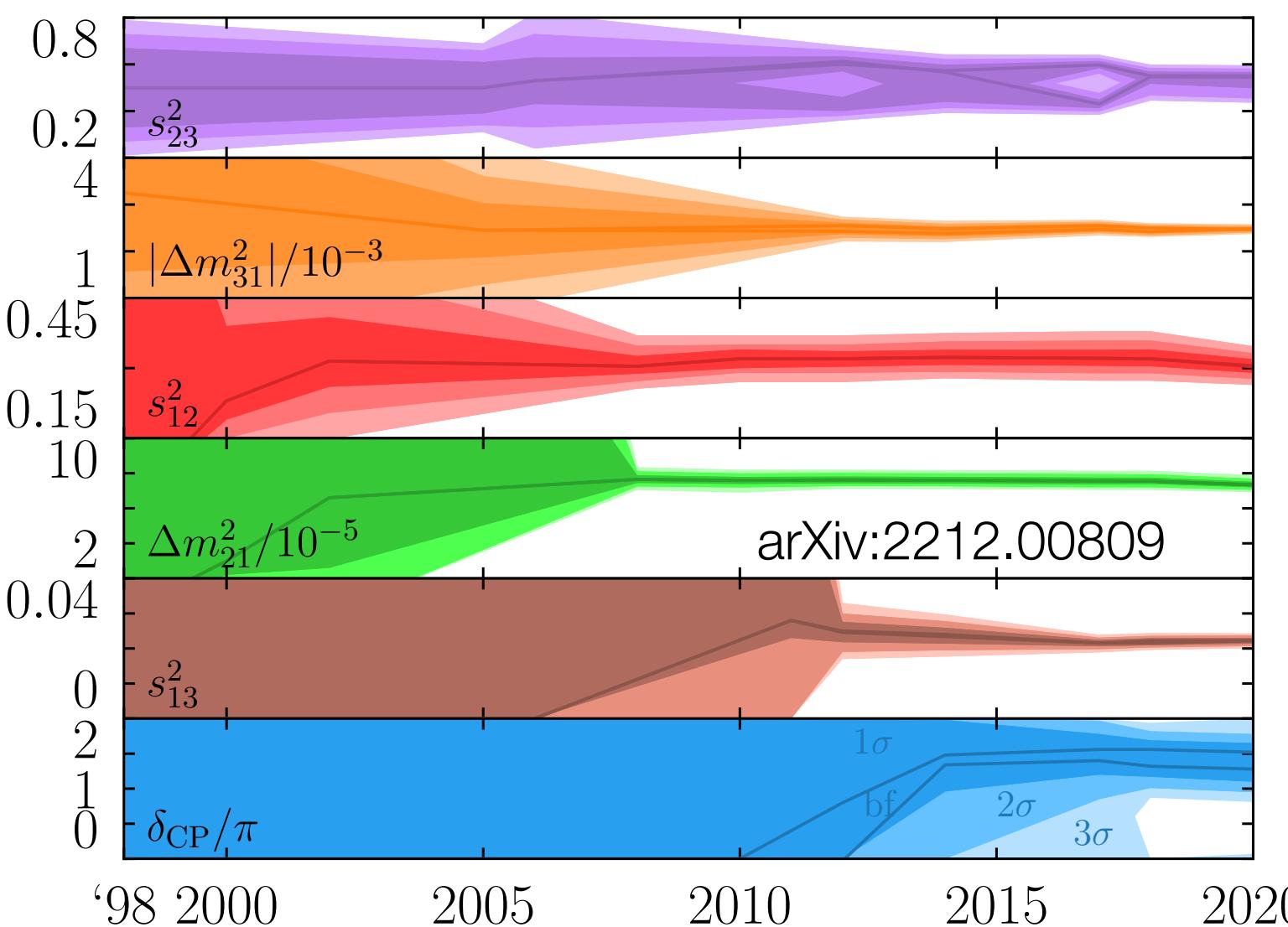
Three-neutrino mixing matrix:

$$\begin{array}{c}
 \text{Atmospheric, accelerator } \nu \\
 \text{Solar, reactor } L \sim 60 \text{ km } \nu \\
 \text{Reactor } L \sim 2 \text{ km, accelerator } \nu
 \end{array}$$

Flavor states $\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$ **Mass states**

$c_{ij} = \cos \theta_{ij}$
 $s_{ij} = \sin \theta_{ij}$
 $\Delta m_{ij}^2 \equiv m_i^2 - m_j^2$

Oscillation parameters:



Parameter	Value	Precision
θ_{23}	47°	3.8%
$ \Delta m_{32}^2 \approx \Delta m_{31}^2 $	$2.44 \times 10^{-3} \text{ eV}^2$	1.4%
θ_{12}	34°	4.2%
Δm_{21}^2	$7.53 \times 10^{-5} \text{ eV}^2$	2.4%
θ_{13}	9°	2.8 %
δ_{CP}	$\sim -\pi/2$	~30%

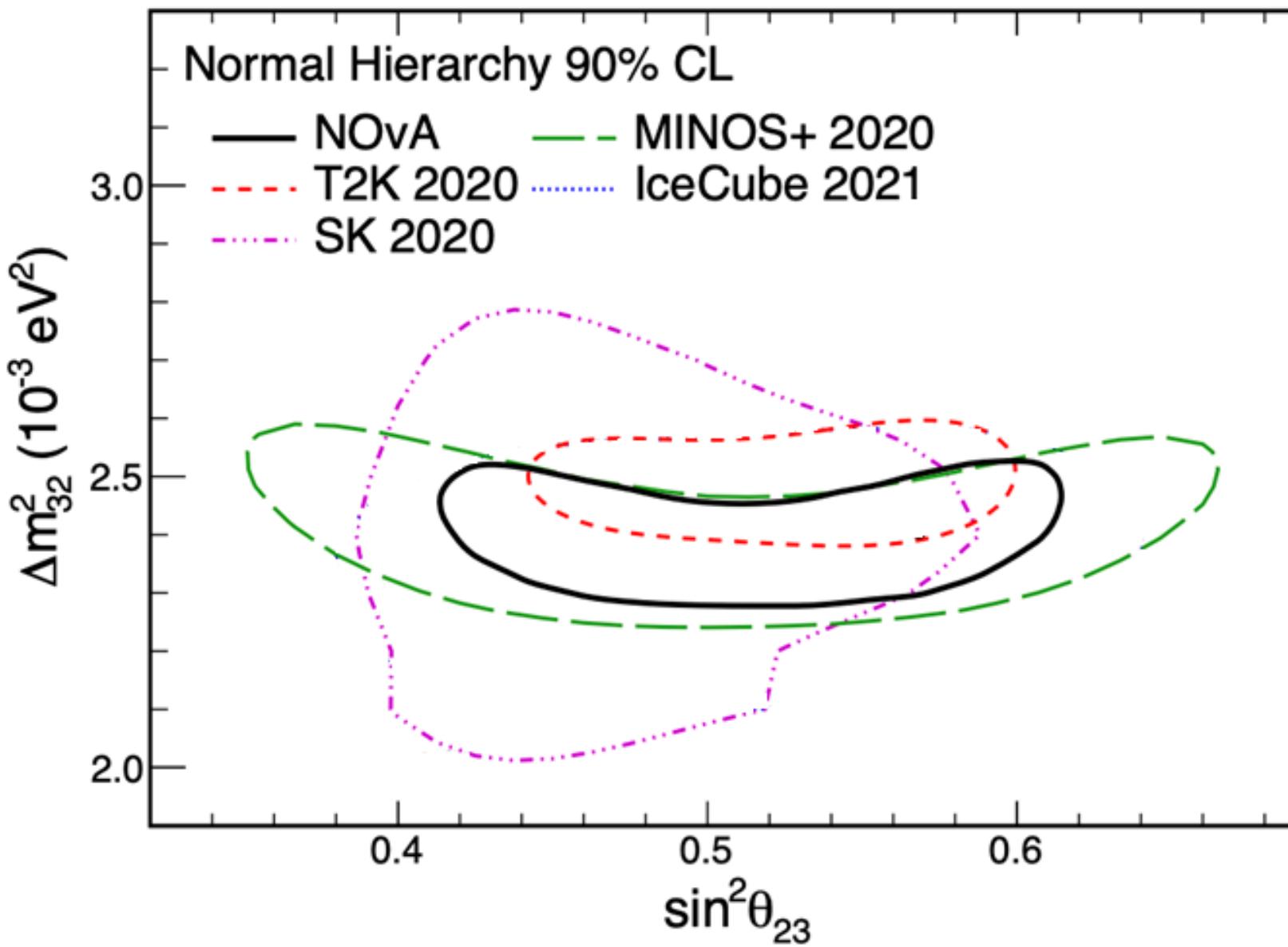
Open Questions

Three-neutrino mixing matrix:

$$\begin{array}{c}
 \text{Atmospheric, accelerator } \nu \\
 \text{Solar, reactor } L \sim 60 \text{ km } \nu \\
 \text{Reactor } L \sim 2 \text{ km, accelerator } \nu
 \end{array}$$

Flavor states $\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$ **Mass states**

$c_{ij} = \cos \theta_{ij}$
 $s_{ij} = \sin \theta_{ij}$
 $\Delta m^2_{ij} \equiv m_i^2 - m_j^2$



Parameter	Value	Precision
θ_{23}	47°	3.8%
$ \Delta m^2_{32} \approx \Delta m^2_{31} $	$2.44 \times 10^{-3} \text{ eV}^2$	1.4%
θ_{12}	34°	4.2%
$\Delta m^2_{21}{}^2$	$7.53 \times 10^{-5} \text{ eV}^2$	2.4%
θ_{13}	9°	2.8 %
δ_{CP}	$\sim -\pi/2$	$\sim 30\%$

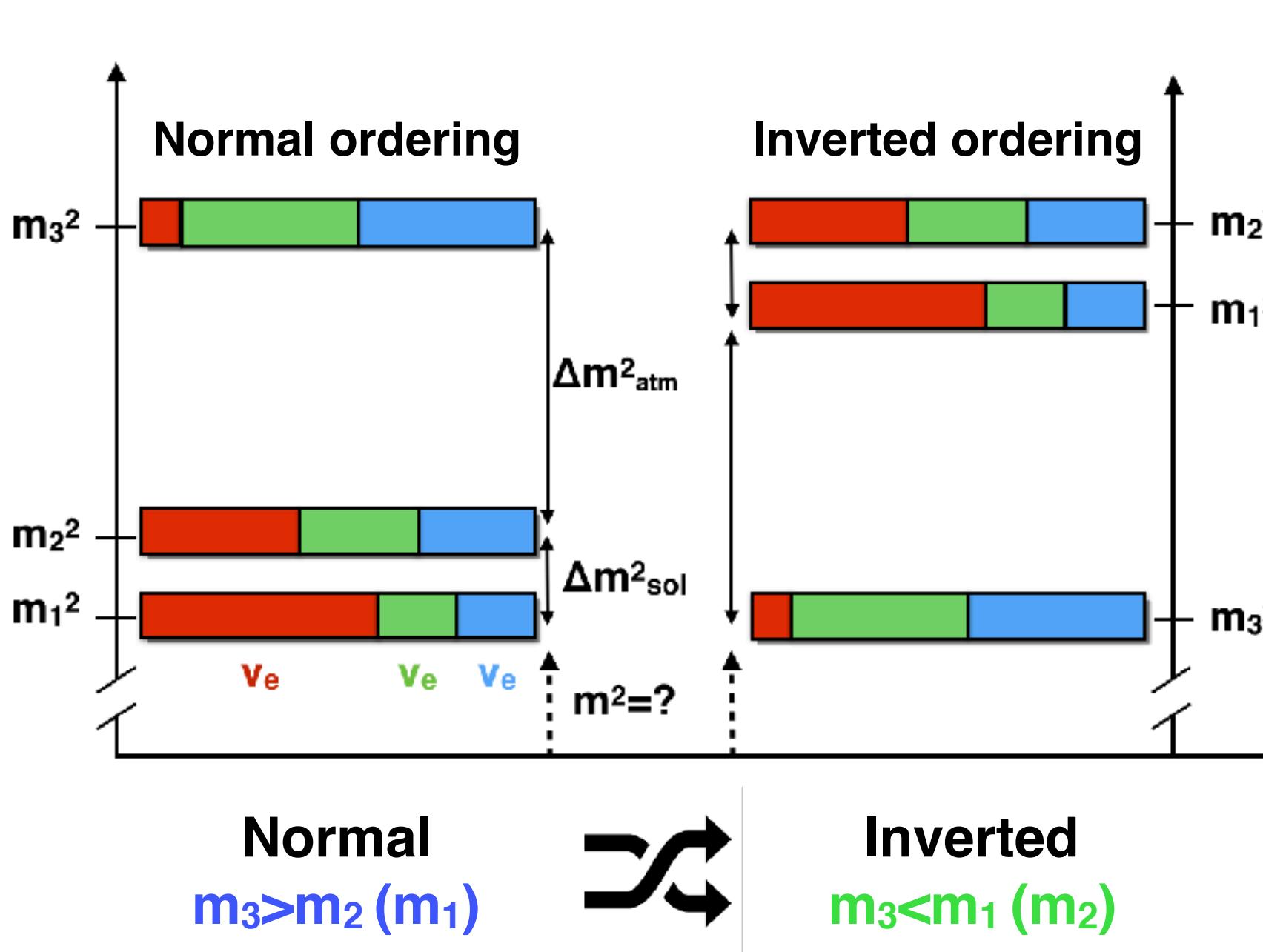
Open questions in ν -physics:

- θ_{23} octant $\Leftrightarrow \theta_{23} \gtrless 45^\circ$?

Open Questions

Three-neutrino mixing matrix:

	Atmospheric, accelerator ν	Solar, reactor $L \sim 60 \text{ km}$ ν
Flavor states	$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$	$\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$
	Reactor $L \sim 2 \text{ km}$, accelerator ν	Mass states $c_{ij} = \cos \theta_{ij}$ $s_{ij} = \sin \theta_{ij}$ $\Delta m^2_{ij} \equiv m_i^2 - m_j^2$



Parameter	Value	Precision
θ_{23}	47°	3.8%
$ \Delta m_{32}^2 \approx \Delta m_{31}^2 $	$2.44 \times 10^{-3} \text{ eV}^2$	1.4%
θ_{12}	34°	4.2%
Δm_{21}^2	$7.53 \times 10^{-5} \text{ eV}^2$	2.4%
θ_{13}	9°	2.8 %
δ_{CP}	$\sim -\pi/2$	$\sim 30\%$

Open questions in ν -physics:

- θ_{23} octant $\Leftrightarrow \theta_{23} \gtrless 45^\circ$?
- Mass ordering

Open Questions

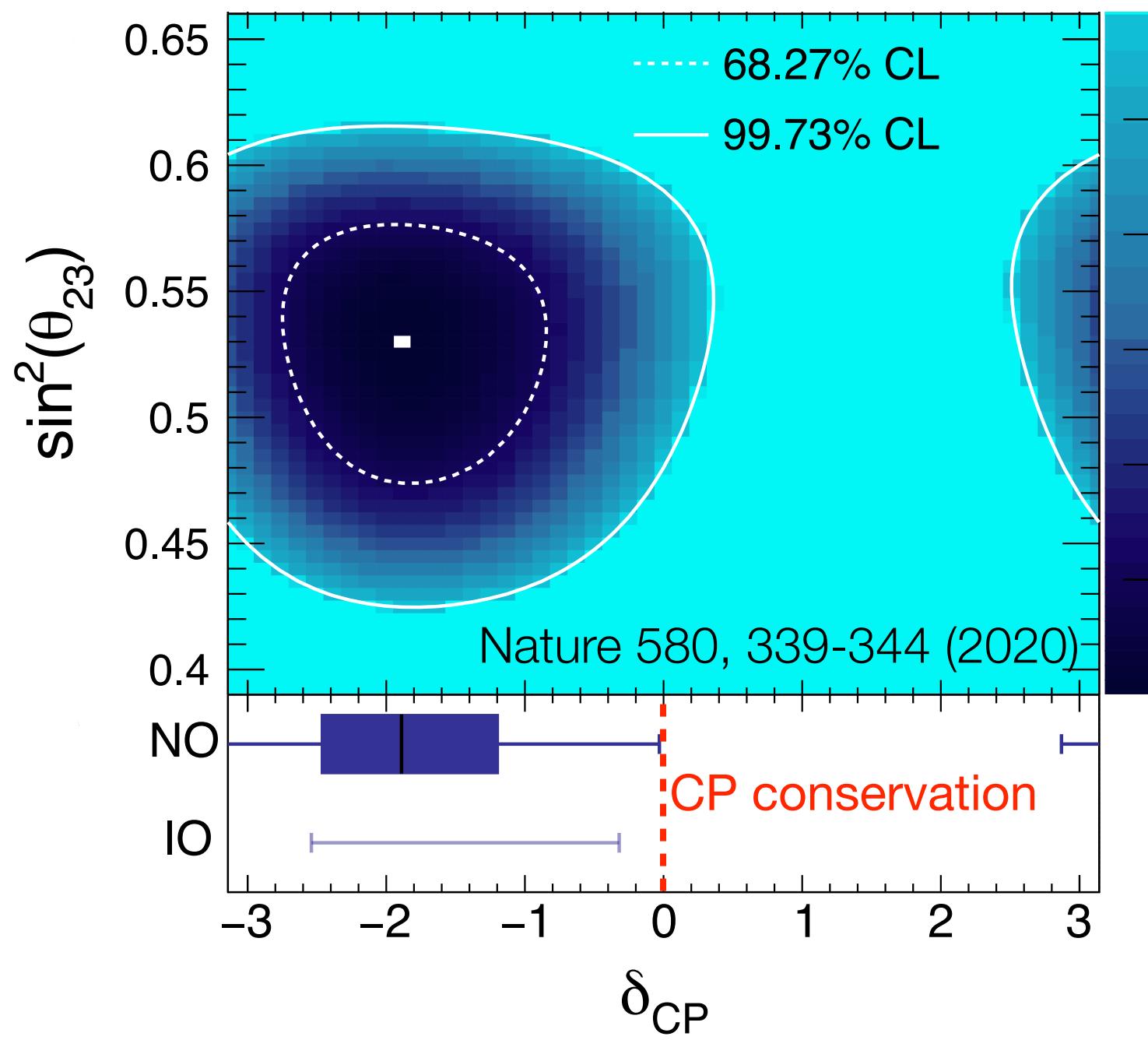


Three-neutrino mixing matrix:

$$\begin{array}{c}
 \text{Atmospheric, accelerator } \nu \\
 \text{Solar, reactor } L \sim 60 \text{ km } \nu \\
 \text{Reactor } L \sim 2 \text{ km, accelerator } \nu
 \end{array}$$

Flavor states $\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$ **Mass states**

$$\begin{aligned}
 c_{ij} &= \cos \theta_{ij} \\
 s_{ij} &= \sin \theta_{ij} \\
 \Delta m^2_{ij} &\equiv m_i^2 - m_j^2
 \end{aligned}$$



Parameter	Value	Precision
θ_{23}	47°	3.8%
$ \Delta m_{32}^2 \approx \Delta m_{31}^2 $	$2.44 \times 10^{-3} \text{ eV}^2$	1.4%
θ_{12}	34°	4.2%
Δm_{21}^2	$7.53 \times 10^{-5} \text{ eV}^2$	2.4%
θ_{13}	9°	2.8 %
δ_{CP}	$\sim -\pi/2$	$\sim 30\%$

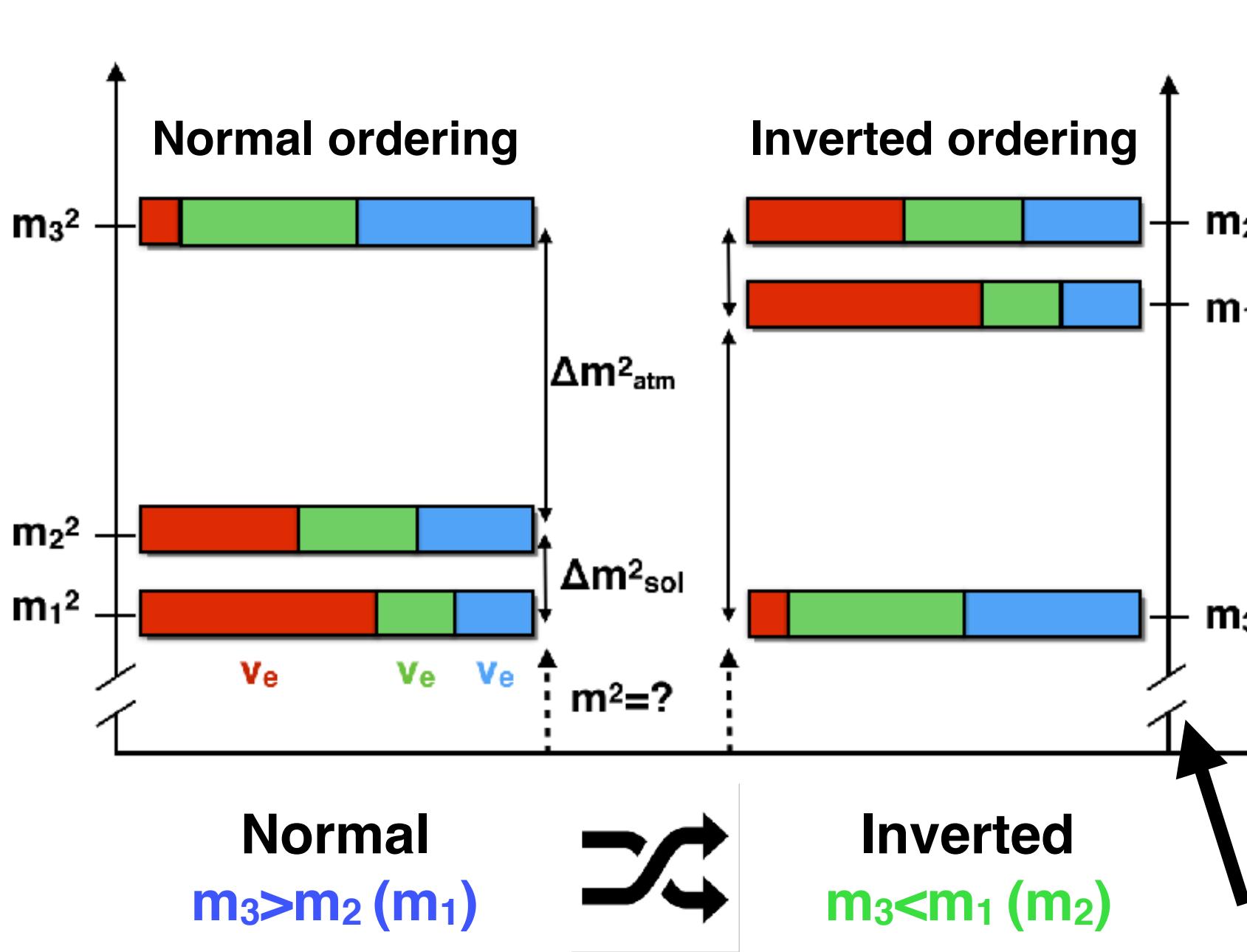
Open questions in ν -physics:

- θ_{23} octant $\Leftrightarrow \theta_{23} \gtrless 45^\circ$?
- Mass ordering
- Value of the δ_{CP}

Open Questions

Three-neutrino mixing matrix:

Flavor states	Atmospheric, accelerator ν	Solar, reactor $L \sim 60 \text{ km}$ ν	
$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$	Δm^2_{atm}	Δm^2_{sol}	$c_{ij} = \cos \theta_{ij}$ $s_{ij} = \sin \theta_{ij}$ $\Delta m^2_{ij} \equiv m_i^2 - m_j^2$
	Reactor $L \sim 2 \text{ km}$, accelerator ν		



Parameter	Value	Precision
θ_{23}	47°	3.8%
$ \Delta m_{32}^2 \approx \Delta m_{31}^2 $	$2.44 \times 10^{-3} \text{ eV}^2$	1.4%
θ_{12}	34°	4.2%
Δm_{21}^2	$7.53 \times 10^{-5} \text{ eV}^2$	2.4%
θ_{13}	9°	2.8 %
δ_{CP}	$\sim -\pi/2$	$\sim 30\%$

What is the absolute mass?

Open questions in ν -physics:

- θ_{23} octant $\Leftrightarrow \theta_{23} \gtrless 45^\circ$?
- Mass ordering
- Value of the δ_{CP}
- Absolute mass scale
- Are ν 's Dirac or Majorana nature?

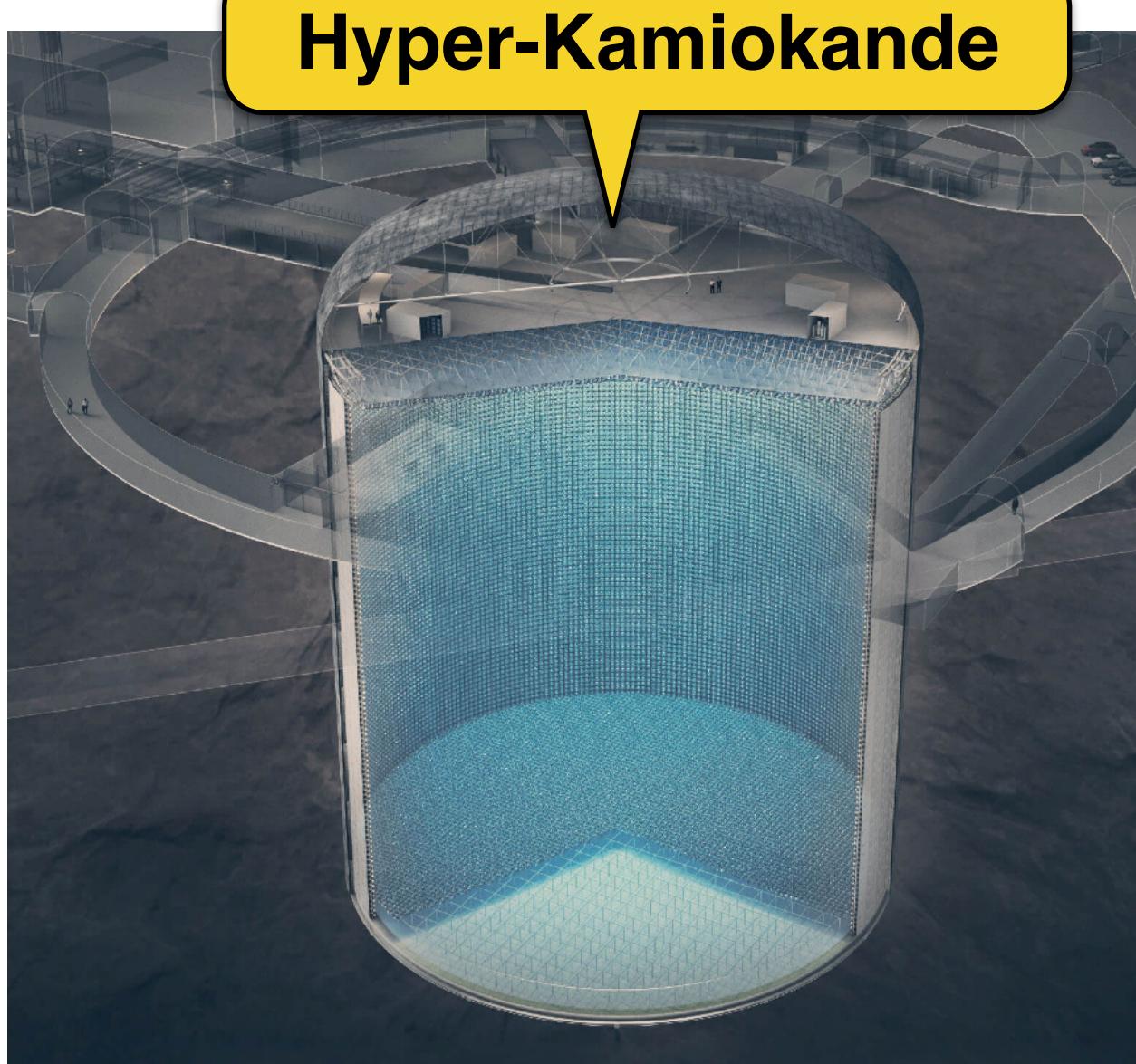
Open Questions

Three-neutrino mixing matrix:

$$\begin{array}{c}
 \text{Atmospheric, accelerator } \nu \\
 \text{Solar, reactor } L \sim 60 \text{ km } \nu \\
 \text{Reactor } L \sim 2 \text{ km, accelerator } \nu
 \end{array}$$

Flavor states $\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$ **Mass states**

$$\begin{aligned}
 c_{ij} &= \cos \theta_{ij} \\
 s_{ij} &= \sin \theta_{ij} \\
 \Delta m_{ij}^2 &\equiv m_i^2 - m_j^2
 \end{aligned}$$

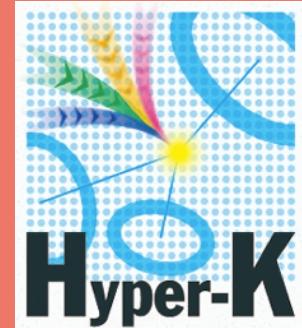


Parameter	Value	Precision
θ_{23}	47°	3.8%
$ \Delta m_{32}^2 \approx \Delta m_{31}^2 $	$2.44 \times 10^{-3} \text{ eV}^2$	1.4%
θ_{12}	34°	4.2%
Δm_{21}^2	$7.53 \times 10^{-5} \text{ eV}^2$	2.4%
θ_{13}	9°	2.8 %
δ_{CP}	$\sim -\pi/2$	~30%

Open questions in ν -physics:

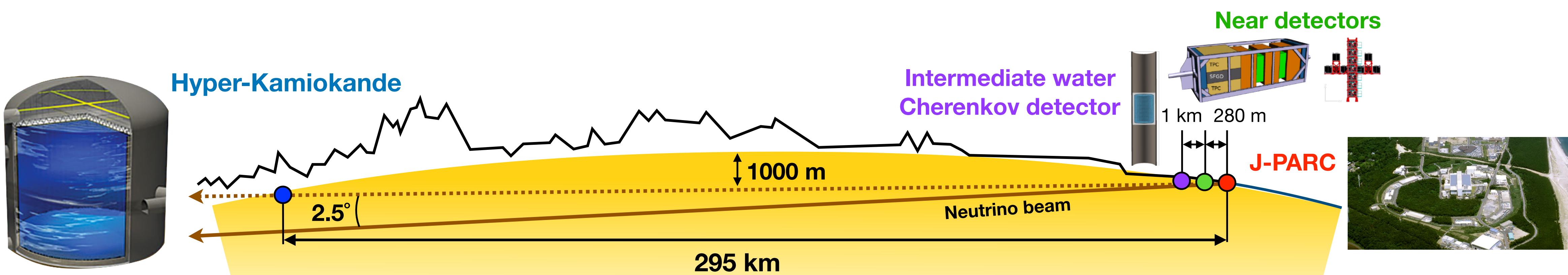
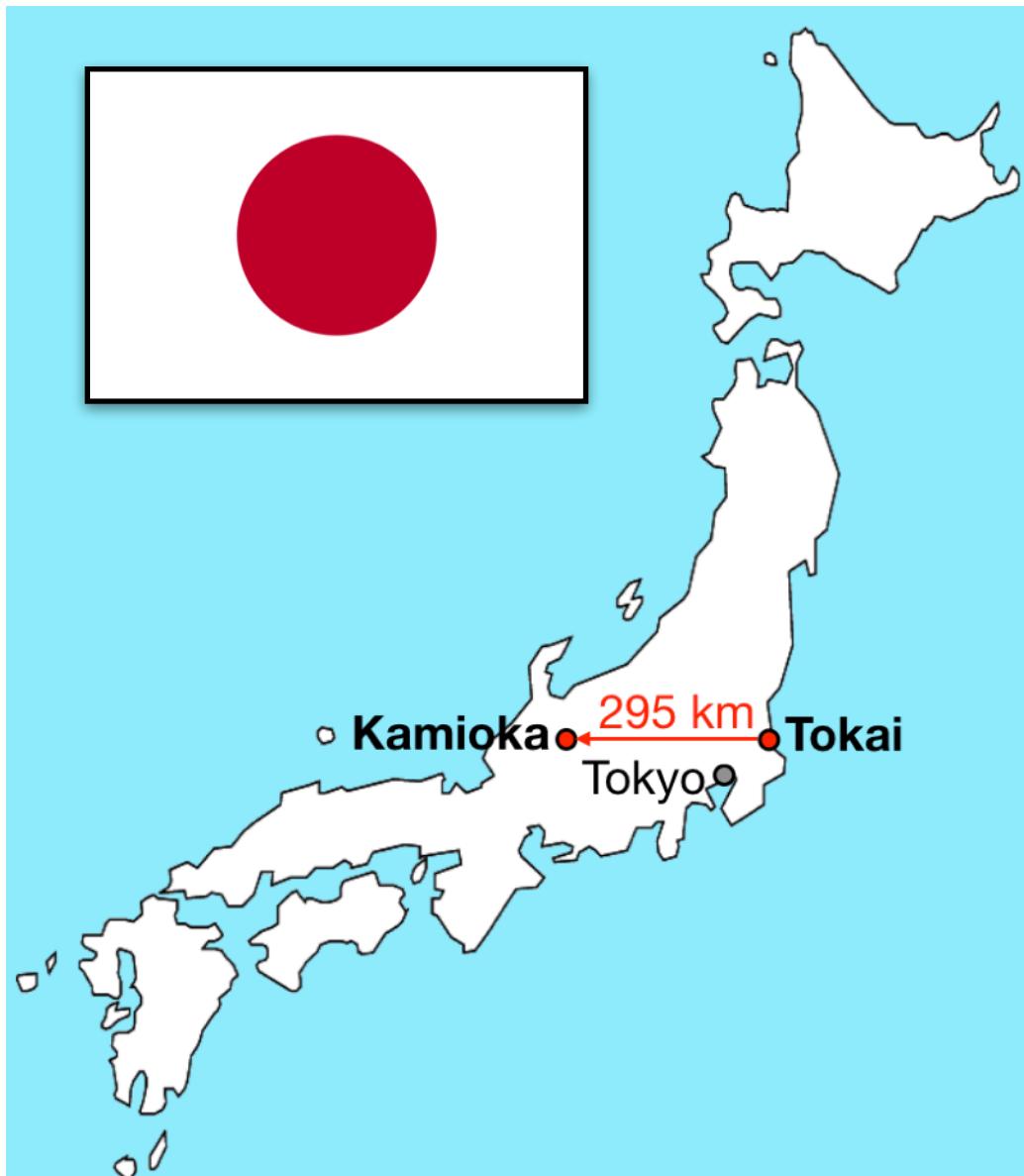
- θ_{23} octant $\Leftrightarrow \theta_{23} \geq 45^\circ$?
- Mass ordering
- Value of the δ_{CP}
- Absolute mass
- Are ν 's Dirac or Majorana nature?

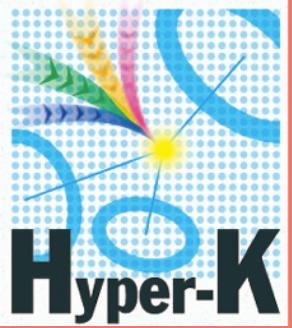
Will be addressed by
Hyper-Kamiokande



The Hyper-Kamiokande Experiment

- International multipurpose experiment under construction in Japan - long baseline neutrino oscillation as well as non-accelerator physics
- The world's most-intense neutrino beam from J-PARC (ultimately 1.3 MW)
- Near detectors to control the systematics (@J-PARC):
 - INGRID & ND280 (common with T2K) - further improved with upgrades
 - Intermediate Water Cherenkov detector - newly built movable detector
- Hyper-Kamiokande (far) detector will be the largest neutrino and nucleon decay detector in the world with 260 kt of water (@Kamioka)

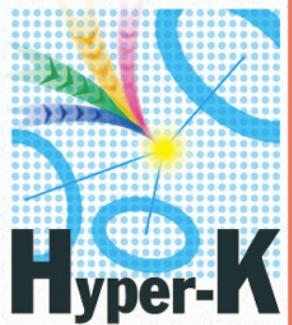




The Hyper-Kamiokande Collaboration

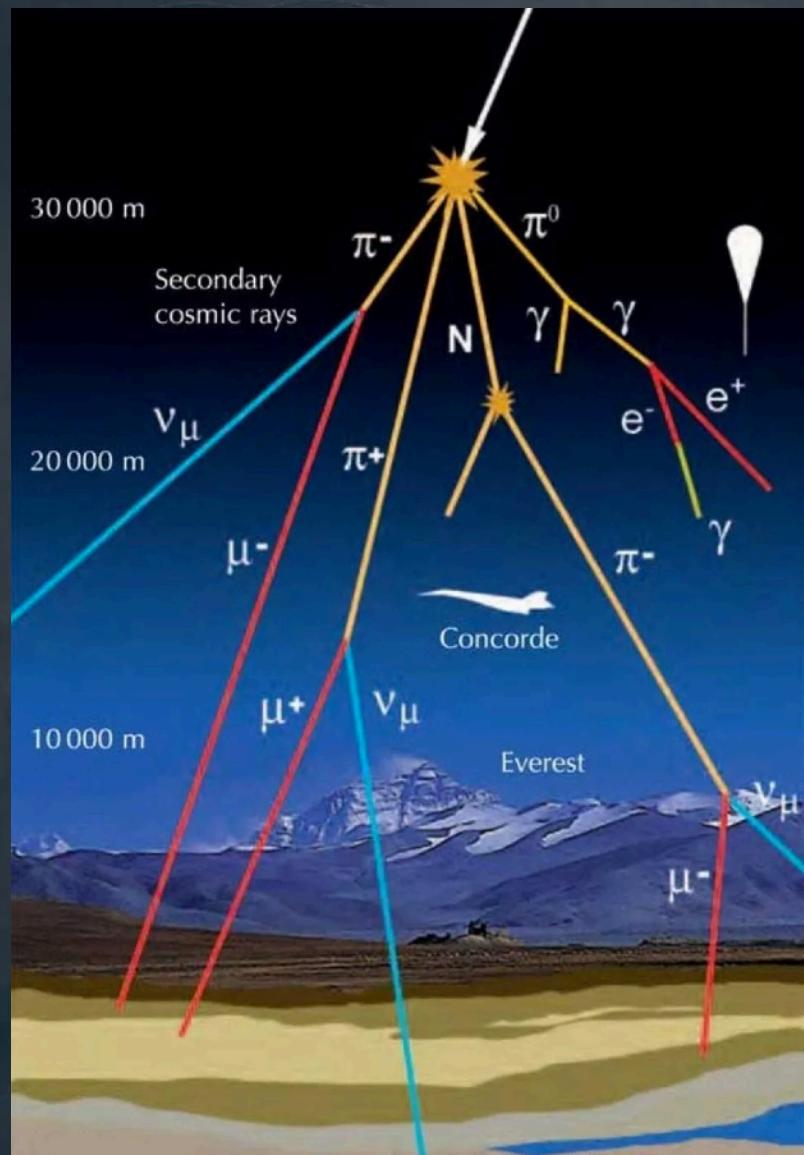
- About 560 scientists from 103 institutions from 22 countries





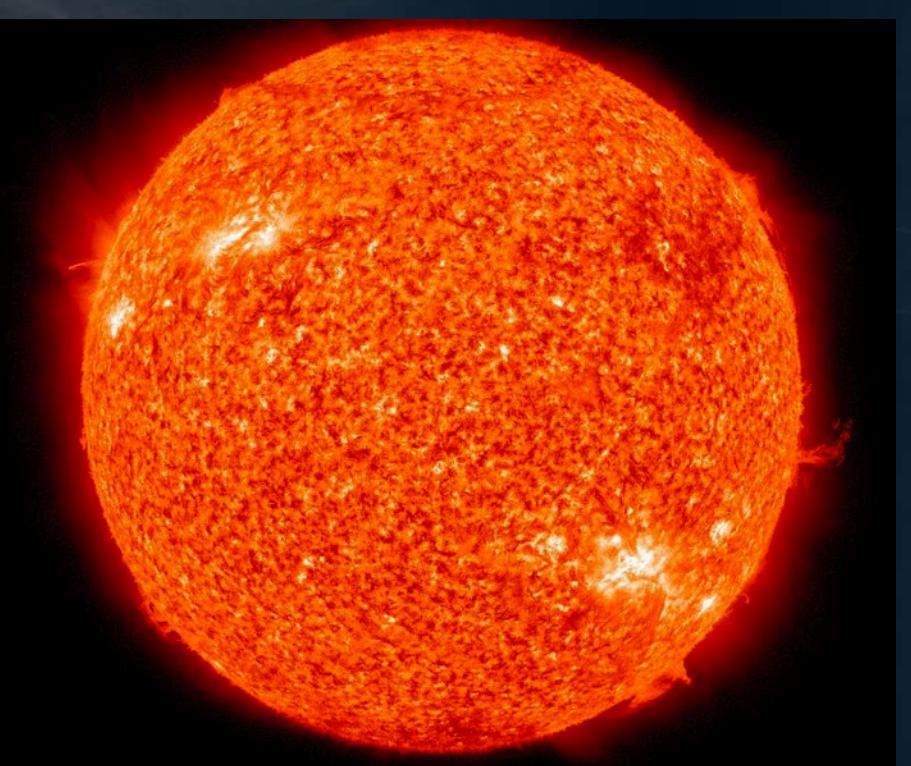
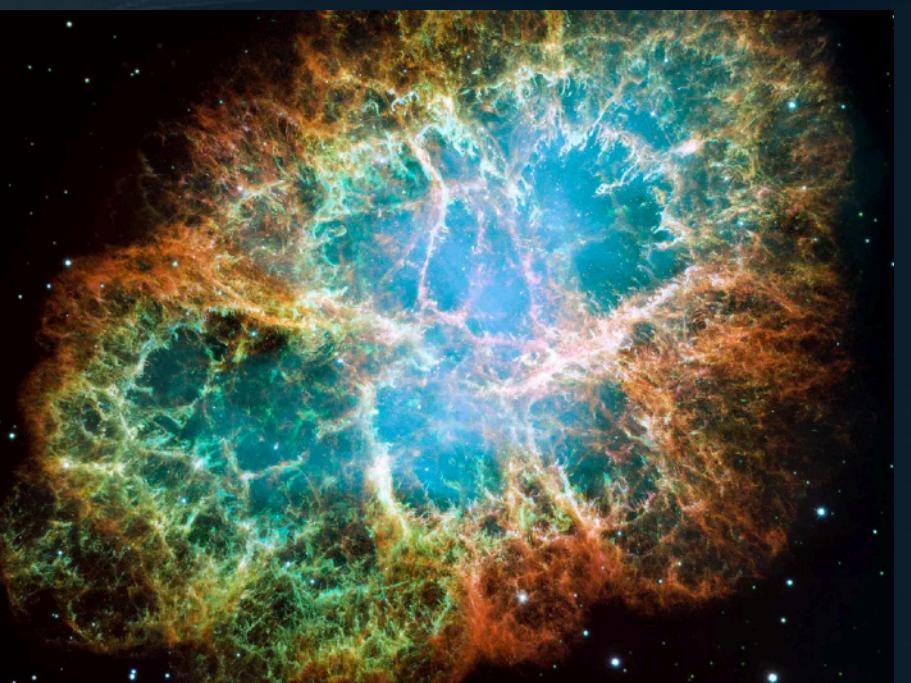
Physics Program in a Nutshell

Atmospheric neutrinos



Accelerator neutrino from J-PARC

Supernova neutrinos

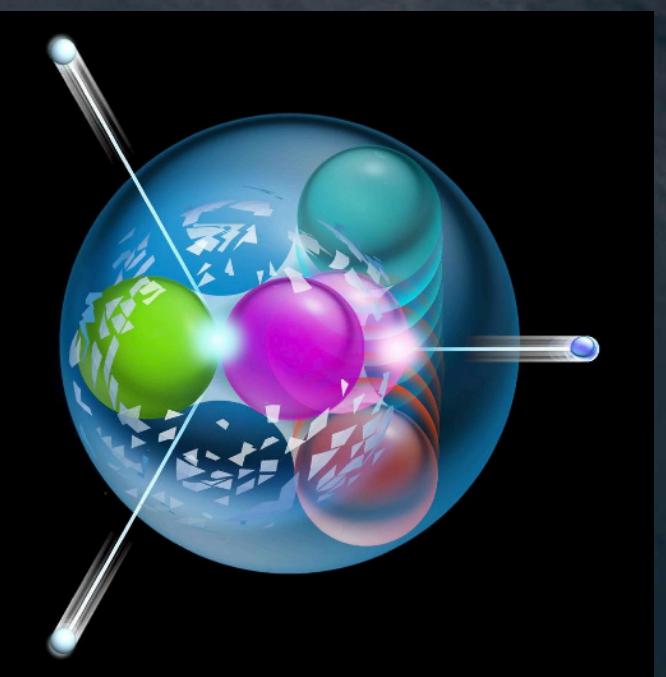


Solar neutrinos

all flavors

ν_e

ν_μ or $\bar{\nu}_\mu$

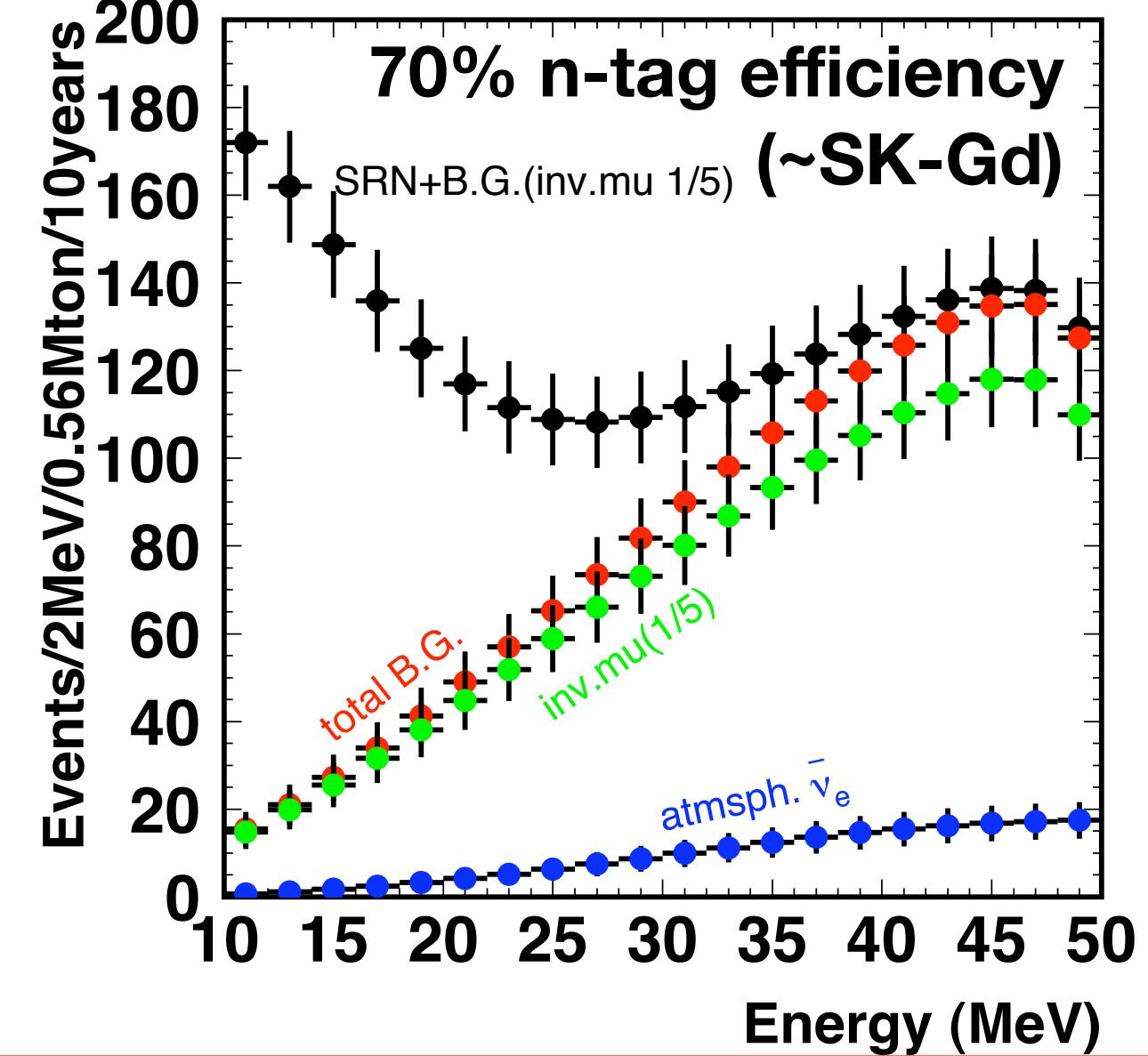
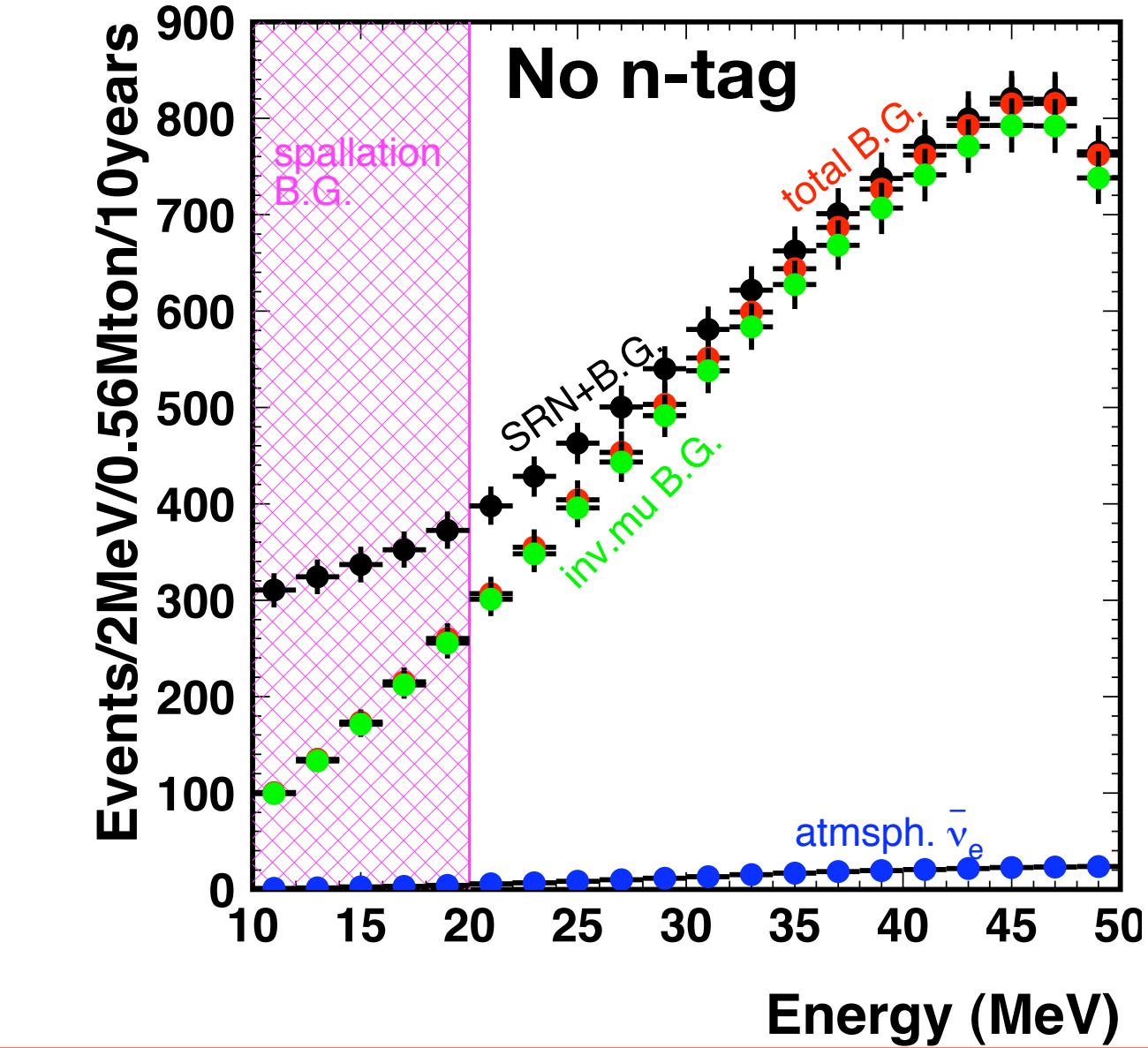
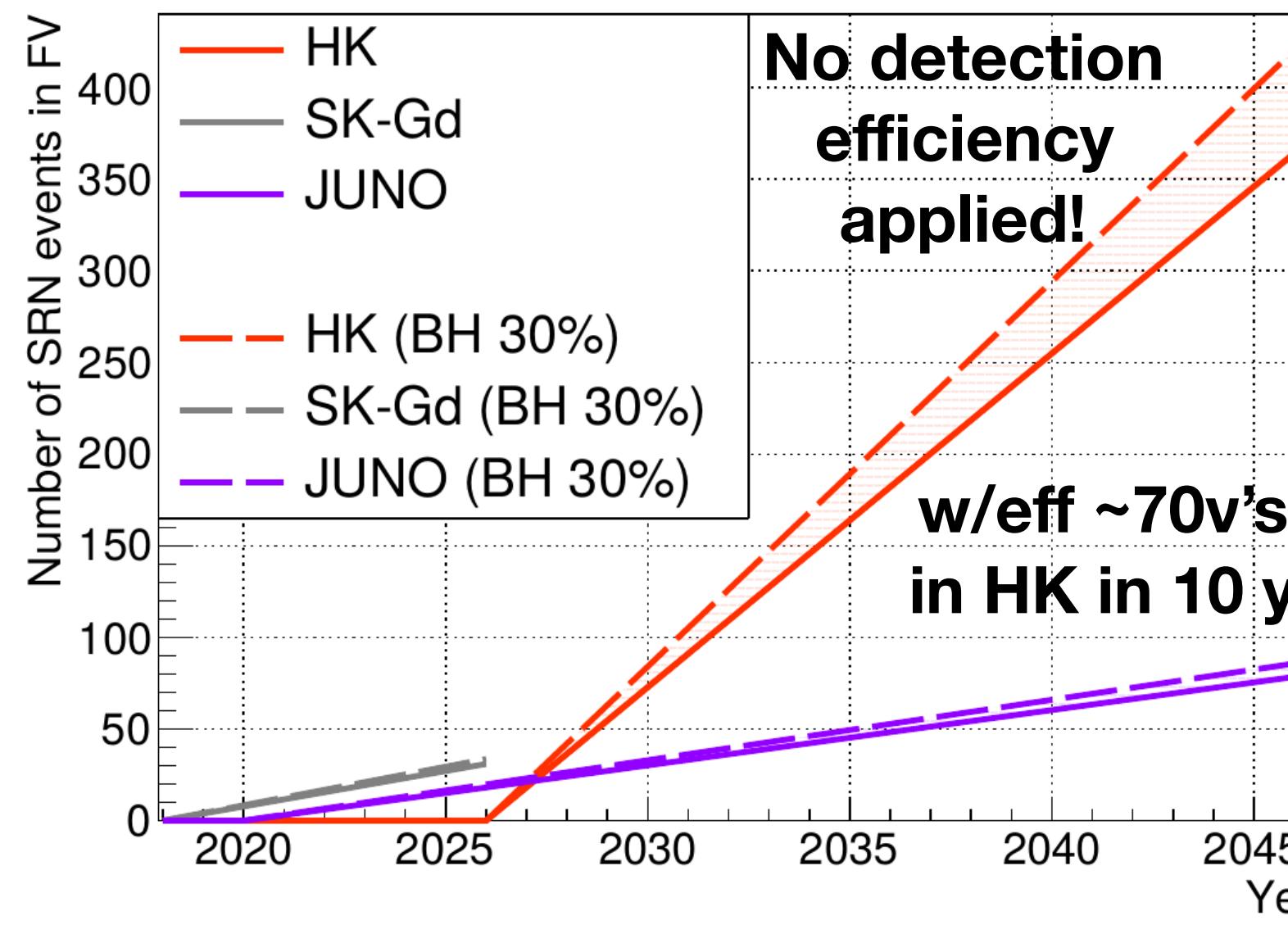


Nucleon decay

Neutrino Astrophysics - Supernova Neutrinos

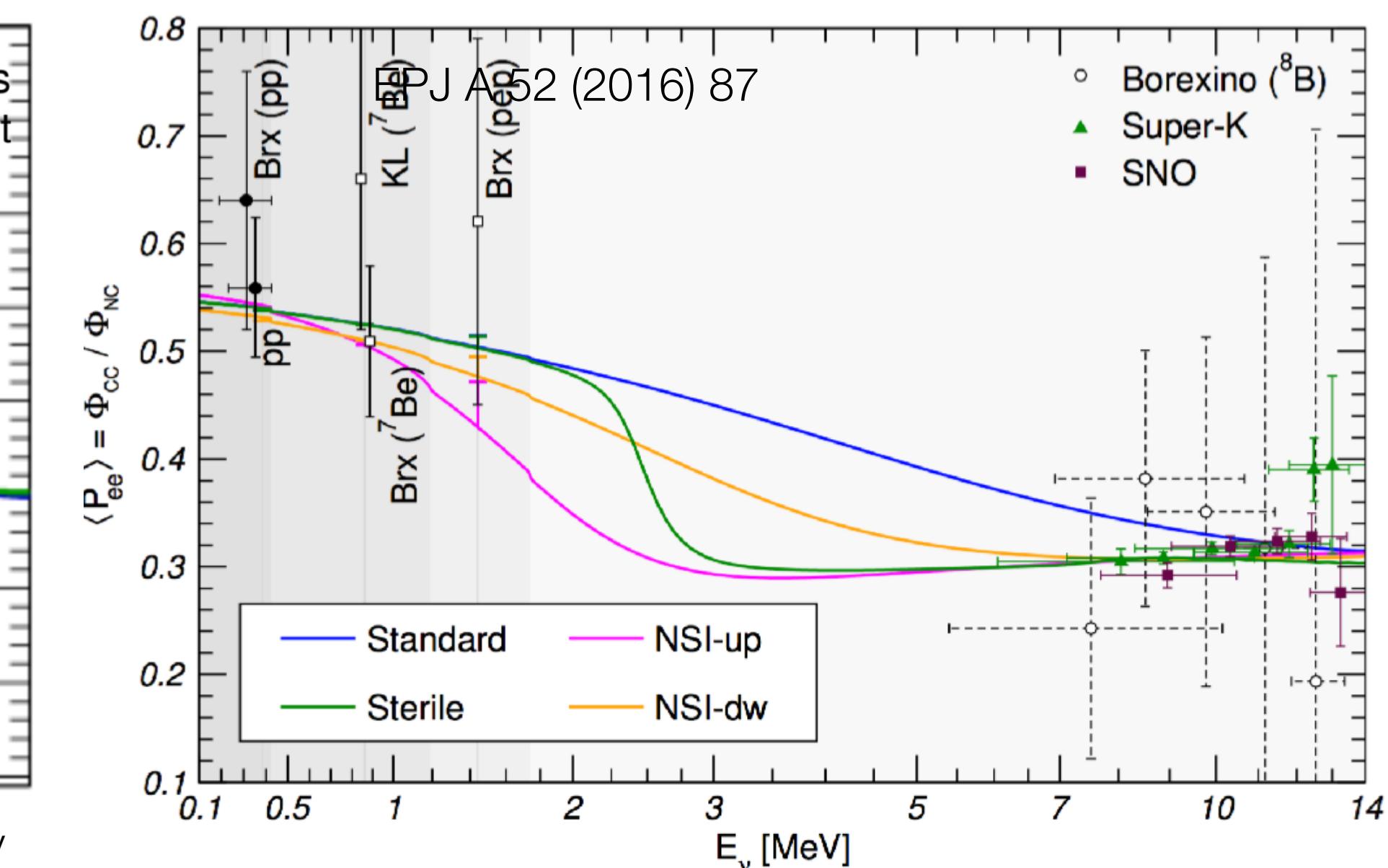
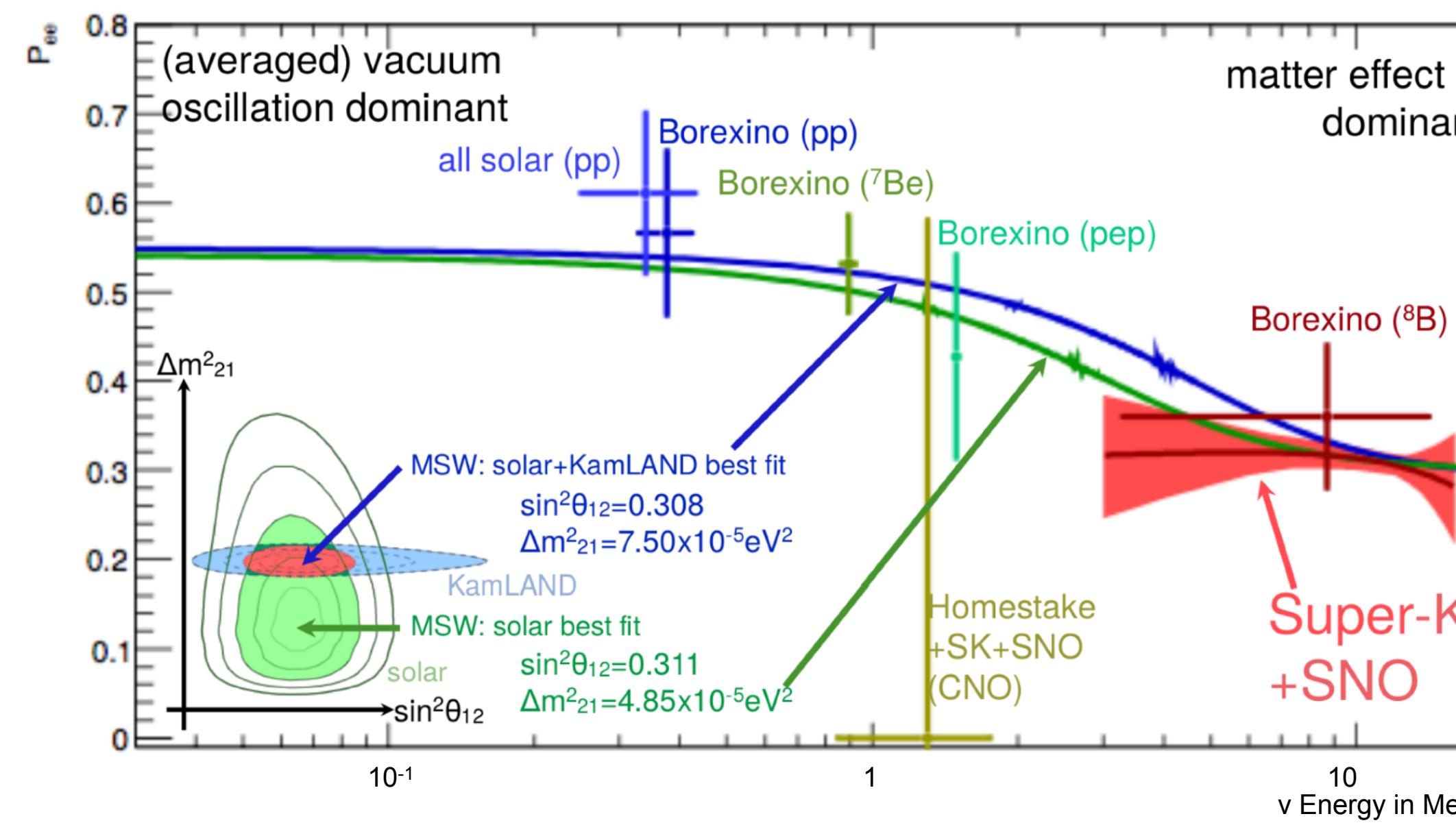


- 99% of the gravitational energy released in ν 's with energy up to ~ 50 MeV
- Supernova burst neutrinos - 54k-90k events in HK for supernova @10 kpc
 - Distinguishing explosion mechanism through time and energy spectra (ApJ 916 15 2021)
 - Event alert with $\sim 1^\circ$ direction precision - multi-messenger astronomy
- Supernova relic neutrinos - not yet observed - determine the past stellar collapse rate
 - Detection through inverse beta decay $\bar{\nu}_e + p \rightarrow e^+ + n$ - neutron tagging efficiency key aspect



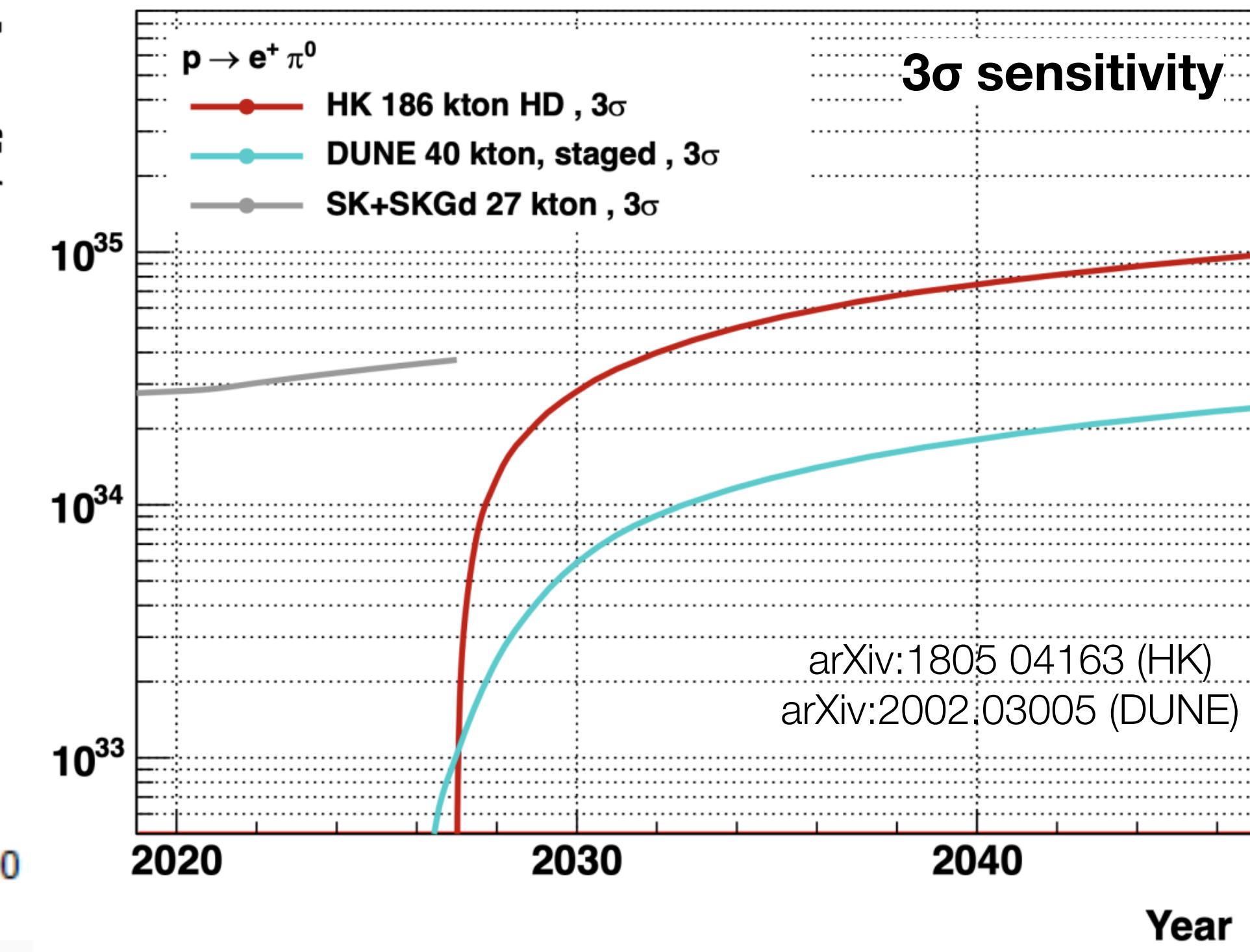
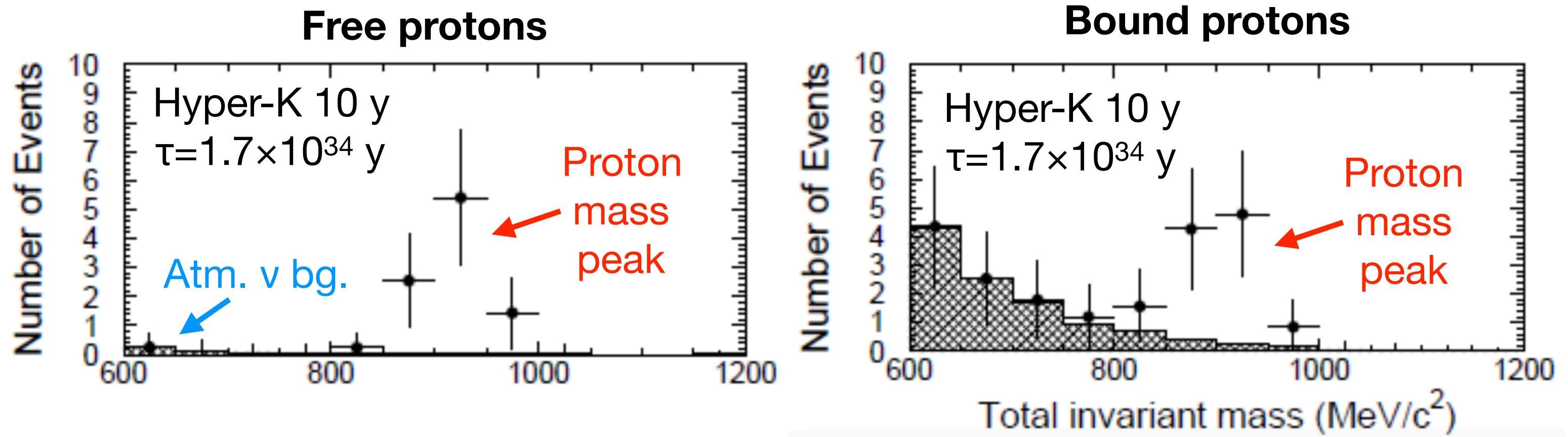
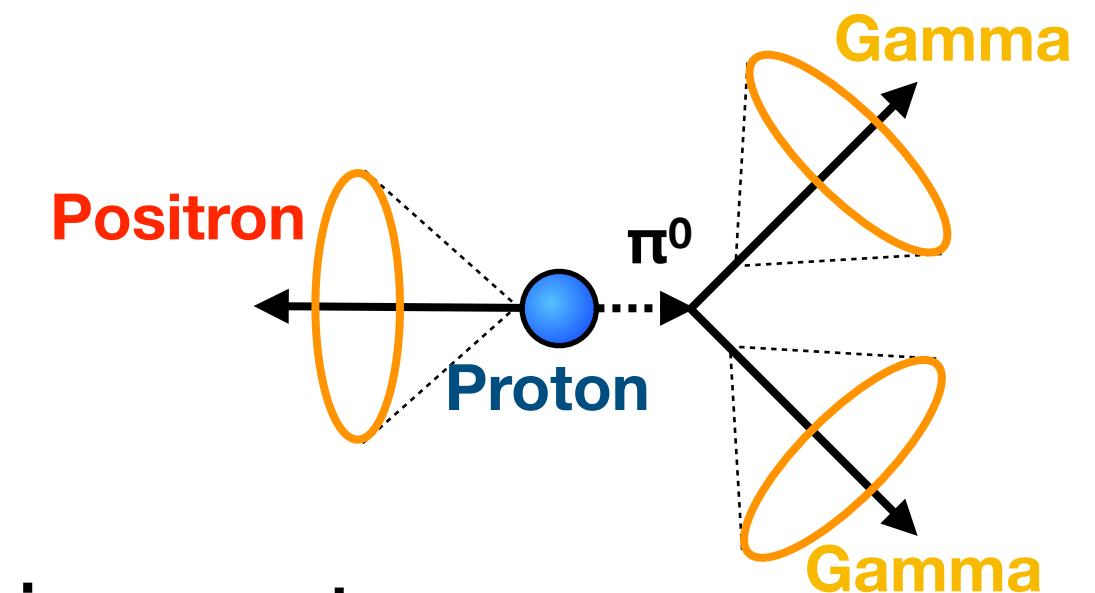
Neutrino Astrophysics - Solar Neutrinos

- So-called up-turn (transition between MSW-vacuum oscillations) - sensitive to various (non)standard parameters
 - Current tension between standard prediction and measurements
 - High-precision measurement by among others Hyper-Kamiokande separates between statistical fluke and other non-standard physics
- Hyper-K will further probe day/night asymmetry and measure high-energy solar neutrino flux (hep)



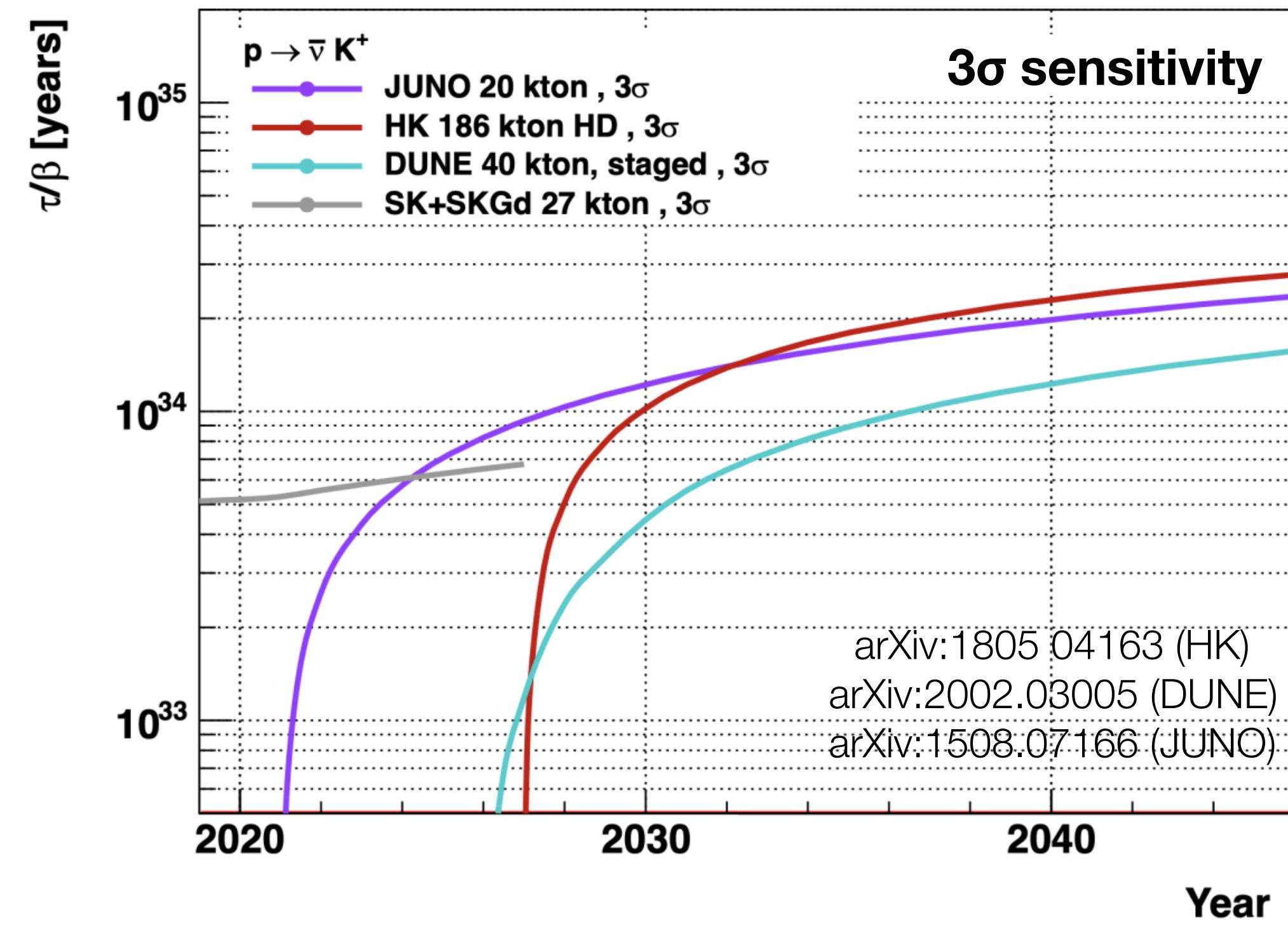
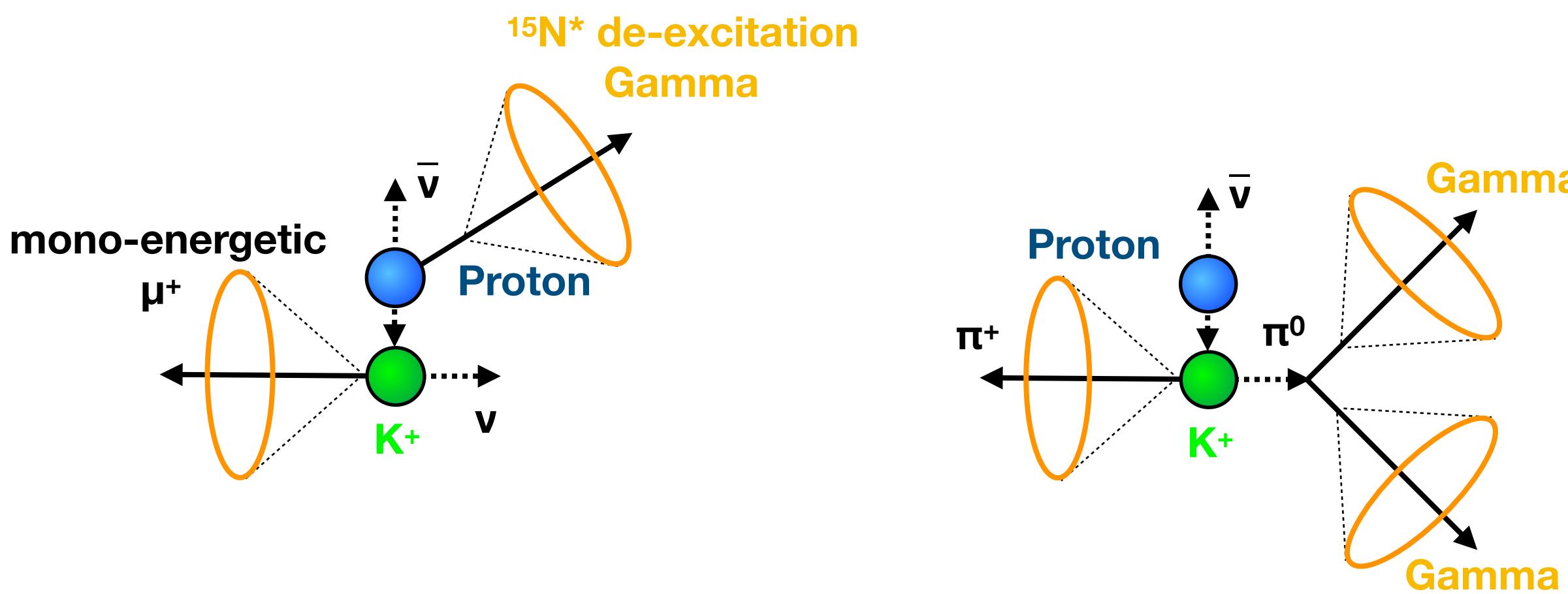
Nucleon Decay Search

- Nucleon decay present in majority of the Grand Unified Theories
- Very long lifetime - Hyper-Kamiokande balance this with very huge mass
- “Golden” channel $p \rightarrow e^+ + \pi^0$
 - Almost background free for free protons due to kinematics&topology requirements
 - Signal also from bound proton decays
- An order of magnitude improvement w.r.t. Super-K
- 3 σ sensitivity for $\tau \sim 10^{35}$ y



Nucleon Decay Search

- The $p \rightarrow K^+ + \bar{\nu}$ channel preferred by Supersymmetry theories
- Kaon decay signatures:
 - $K^+ \rightarrow \mu^+ + \nu_\mu$ (BR 64%) produces a mono-energetic 258 MeV muon and $^{15}\text{N}^*$ de-excitation gamma → prompt-delayed signal coincidence greatly suppresses the background
 - $K^+ \rightarrow \pi^+ + \pi^0$ (BR 21%) with a clear kinematic&topology pattern
- Hyper-Kamiokande the most sensitive experiment in this channel in about a year of data taking

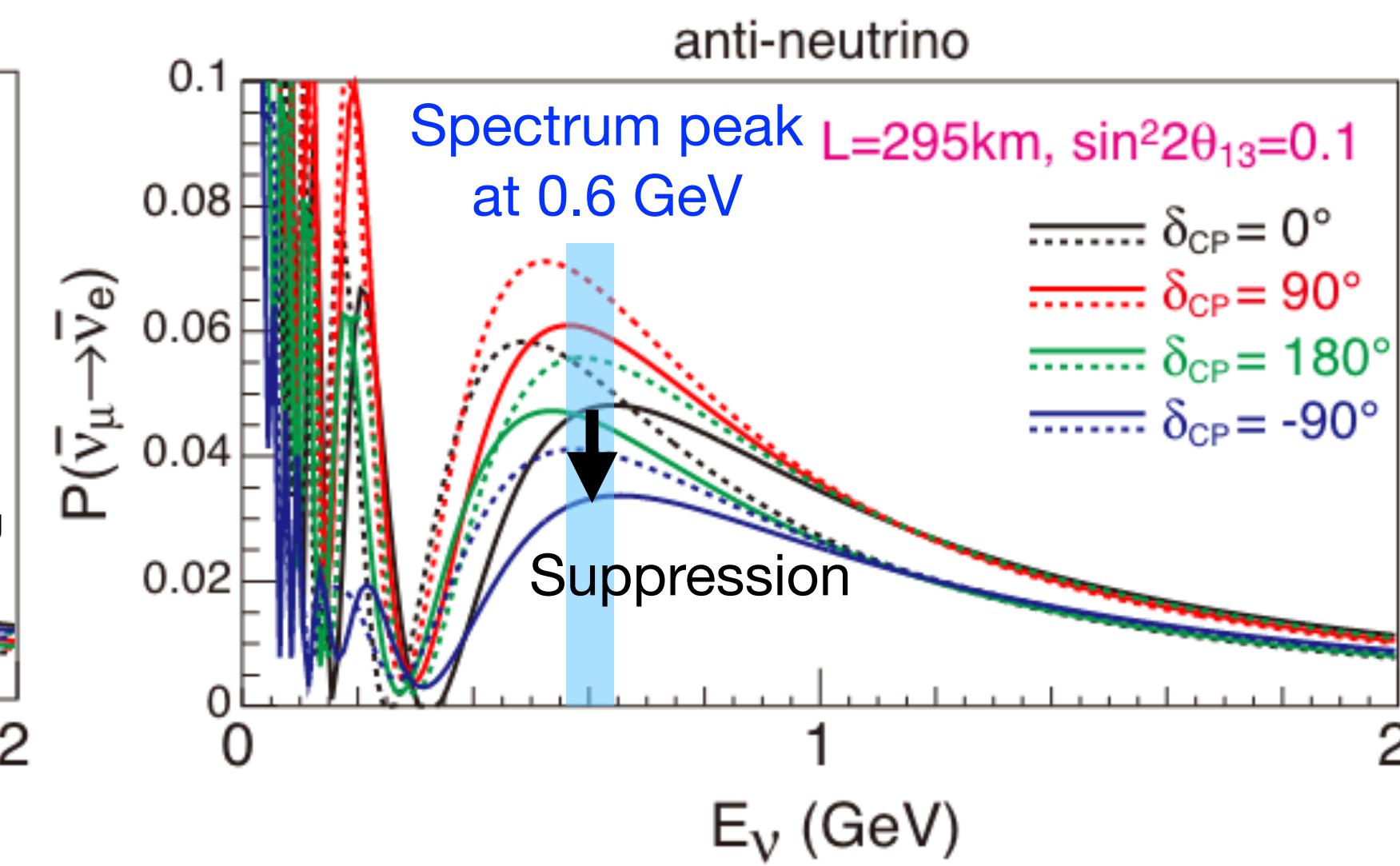
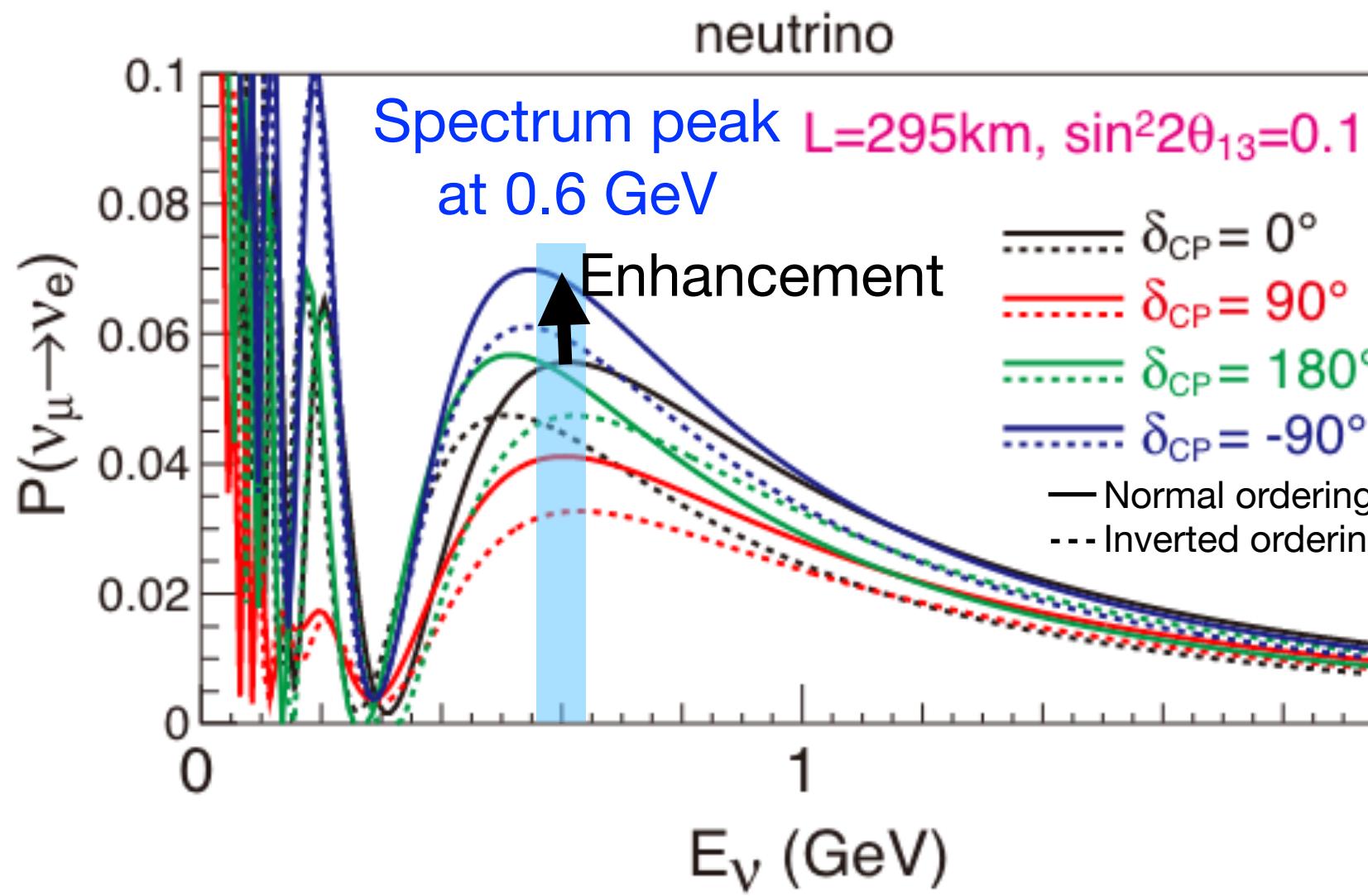


Accelerator Neutrino Long Baseline Measurement



- The narrow spectrum of 2.5° off-axis $\nu_\mu/\bar{\nu}_\mu$ beam peaks at 0.6 GeV (tuned to 295 km baseline)
- Strategy for the CP violation measurement - comparison of $P(\nu_\mu \rightarrow \nu_e)$ and $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$
- Example for $\delta_{CP}=-90^\circ$: probability enhancement for neutrinos vs. suppression for antineutrinos
- High statistics measurement (1.3 MW beam + huge HK det.) with $\sim 2k$ events both for ν_e and $\bar{\nu}_e$ (1:3)
- Improved systematics with new/upgraded near detectors

Error model	μ -like		e-like			$\nu/\bar{\nu}$ modes 0 d.e.
	ν -mode 0 d.e.	$\bar{\nu}$ -mode 0 d.e.	ν -mode 1 d.e.	$\bar{\nu}$ -mode 1 d.e.		
T2K 2020	3.0%	4.0%	4.7%	5.9%	14.1%	4.6%
Improved	1.2%	1.1%	2.1%	2.2%	5.2%	2.0%

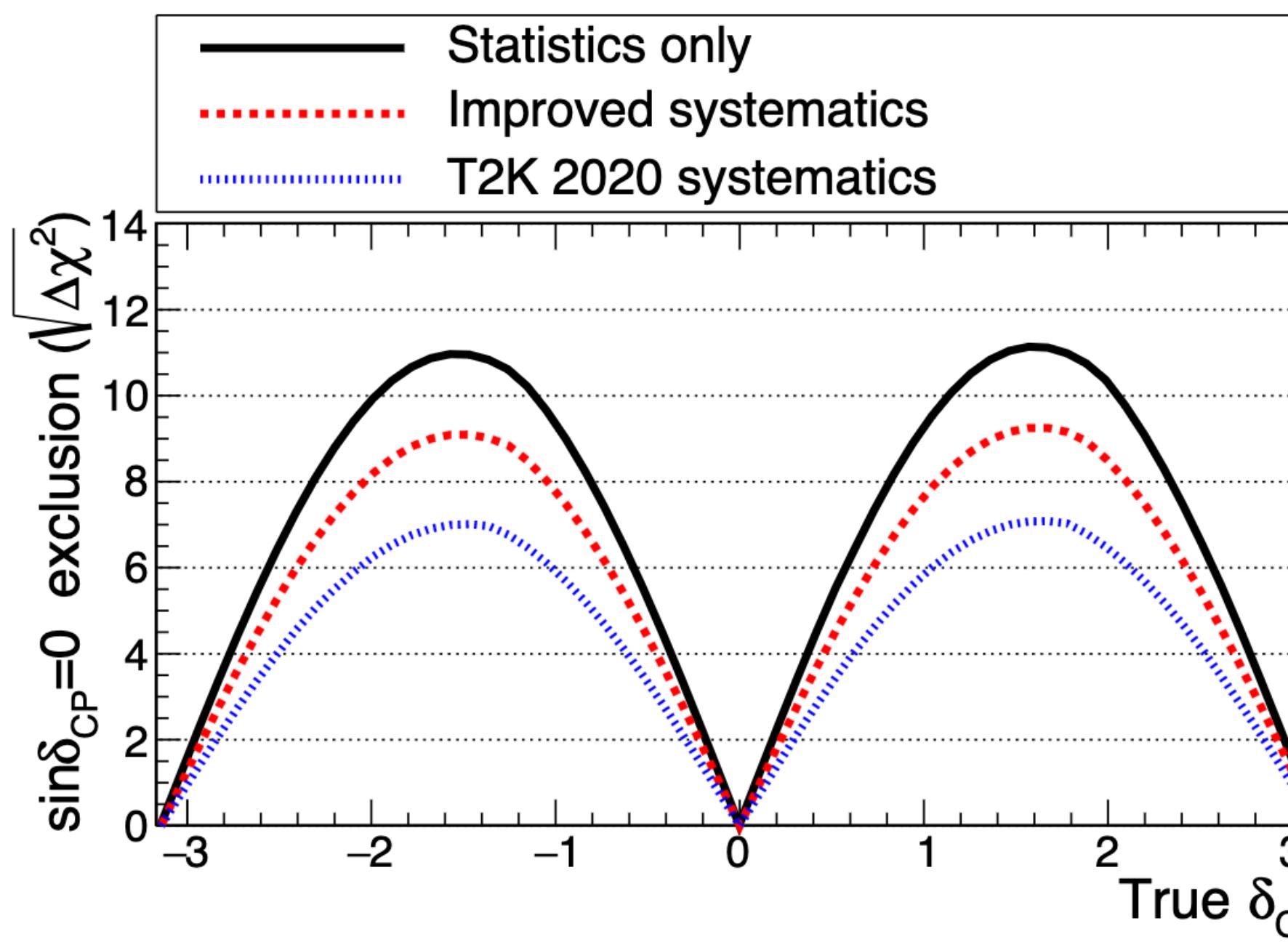


Expected event rate @ 10 years (2.7E22 POT), $\nu:\bar{\nu} = 1:3$, @ $\delta_{CP} = 0$

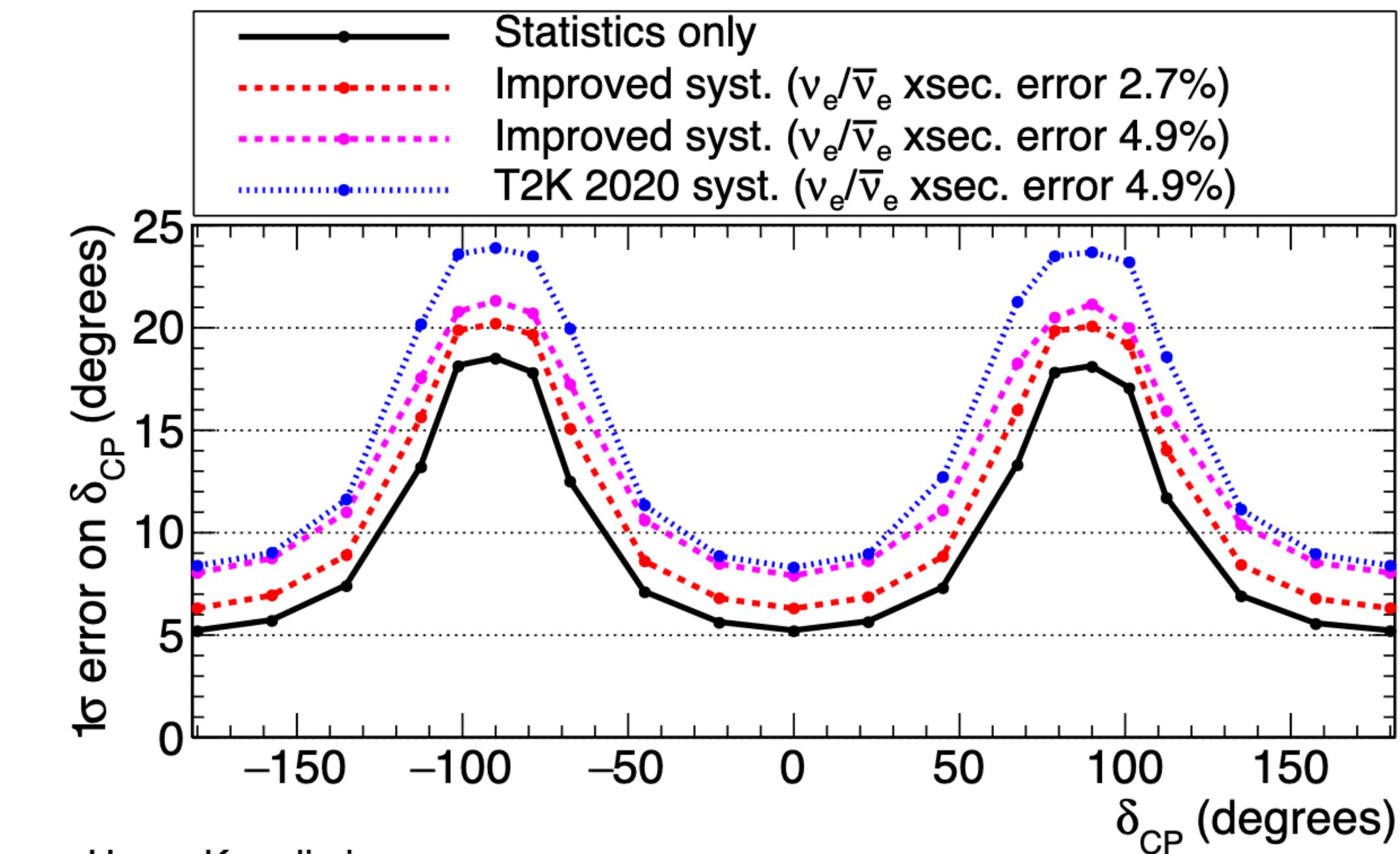
ν -mode, 1-ring μ -like	~ 8800
$\bar{\nu}$ -mode, 1-ring μ -like	~ 12000
ν -mode, 1-ring e-like + 0 decay e	~ 2100
$\bar{\nu}$ -mode, 1-ring e-like + 0 decay e	~ 1800
ν -mode, 1-ring e-like + 1 decay e	~ 300

CP Violation Sensitivity

- Excluding CP conservation for 62% of δ_{CP} parameter space @ 5σ in 10 y (assuming known mass ordering)
- Improved systematics has great impact
- 20.2° error at true $\delta_{\text{CP}}=-90^\circ$
- 6.3° error at true $\delta_{\text{CP}}=0^\circ$



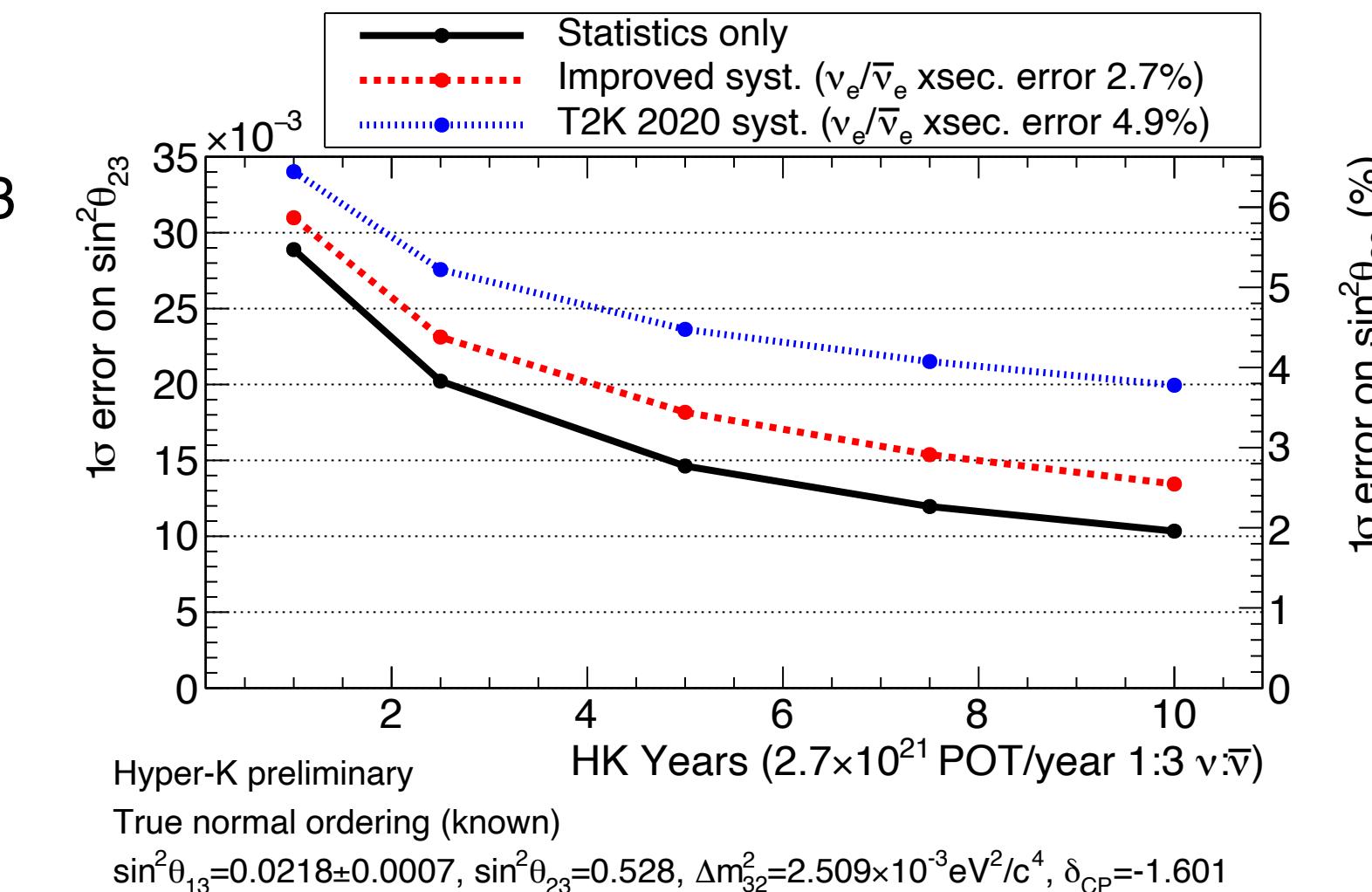
Hyper-K preliminary
True normal ordering (known), 10 years (2.7×10^{22} POT 1:3 v. \bar{v})
 $\sin^2\theta_{13}=0.0218\pm0.0007$, $\sin^2\theta_{23}=0.528$, $\Delta m_{32}^2=2.509\times10^{-3}\text{eV}^2/\text{c}^4$



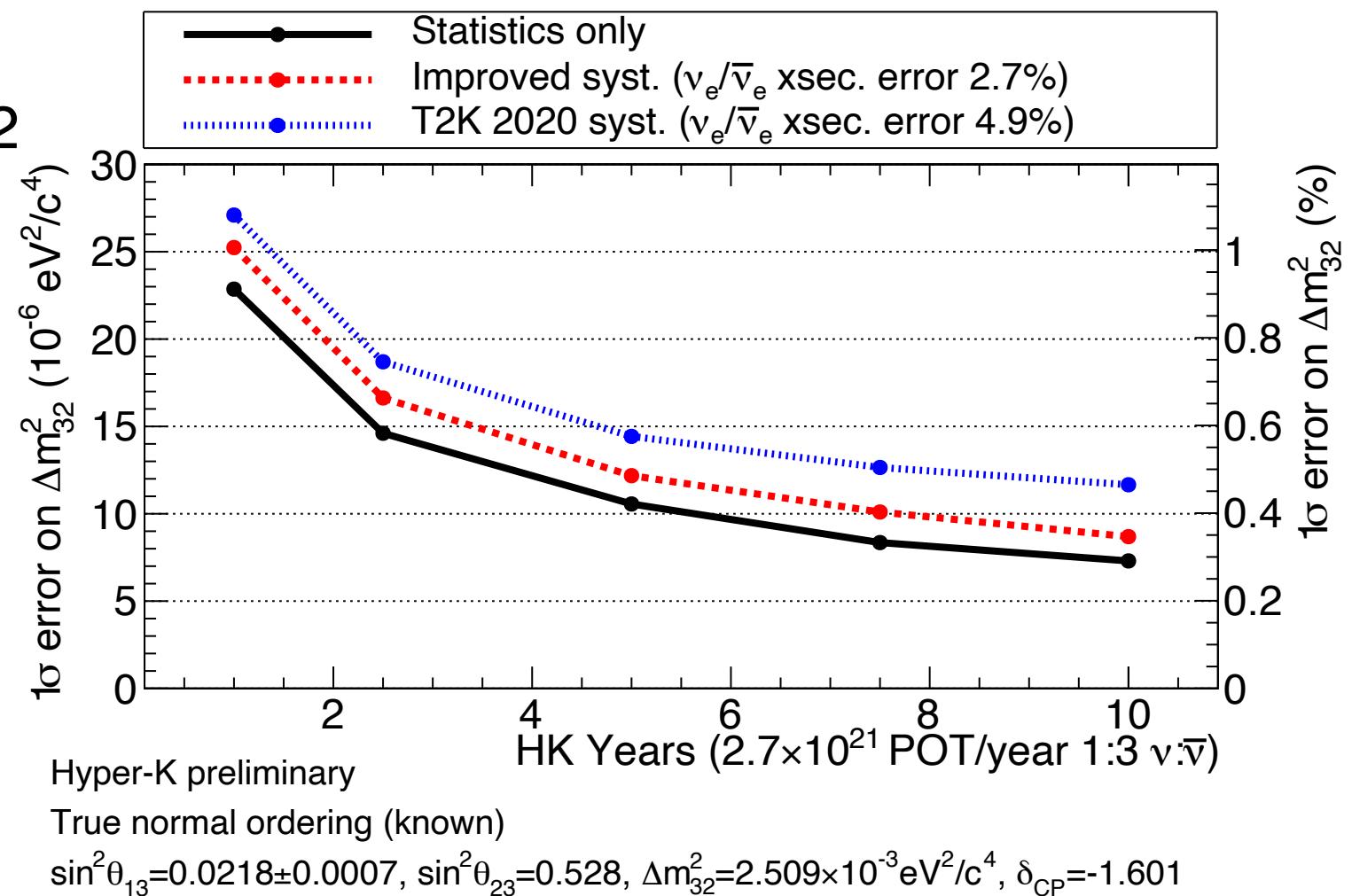
Hyper-K preliminary
True normal ordering (known), HK 10 Years (2.7×10^{22} POT 1:3 v. \bar{v})
 $\sin^2\theta_{13}=0.0218\pm0.0007$, $\sin^2\theta_{23}=0.528$, $\Delta m_{32}^2=2.509\times10^{-3}\text{eV}^2/\text{c}^4$

Sensitivity for Other Oscillation Parameters

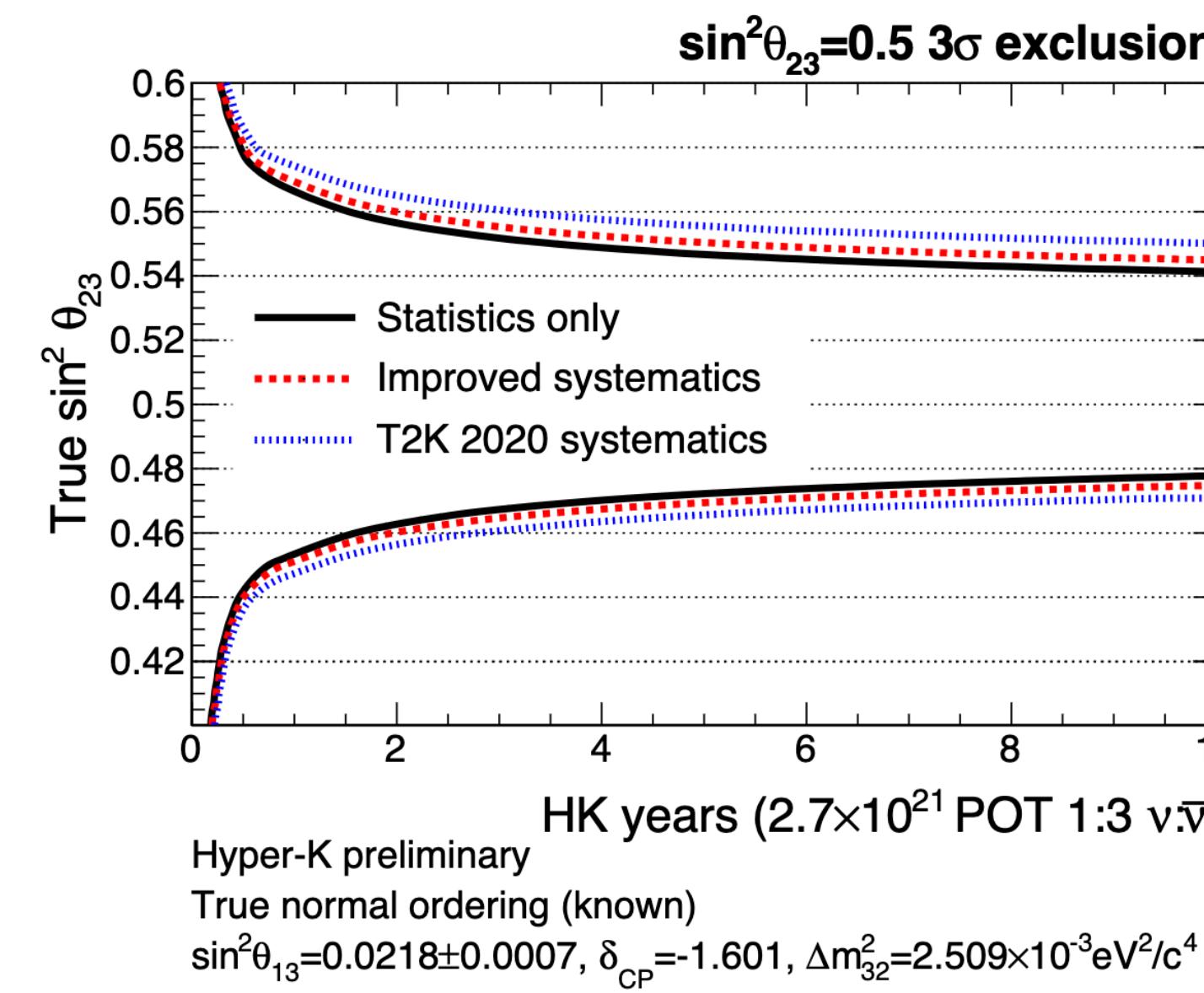
- Precision of θ_{23}
4% \rightarrow 2.5%



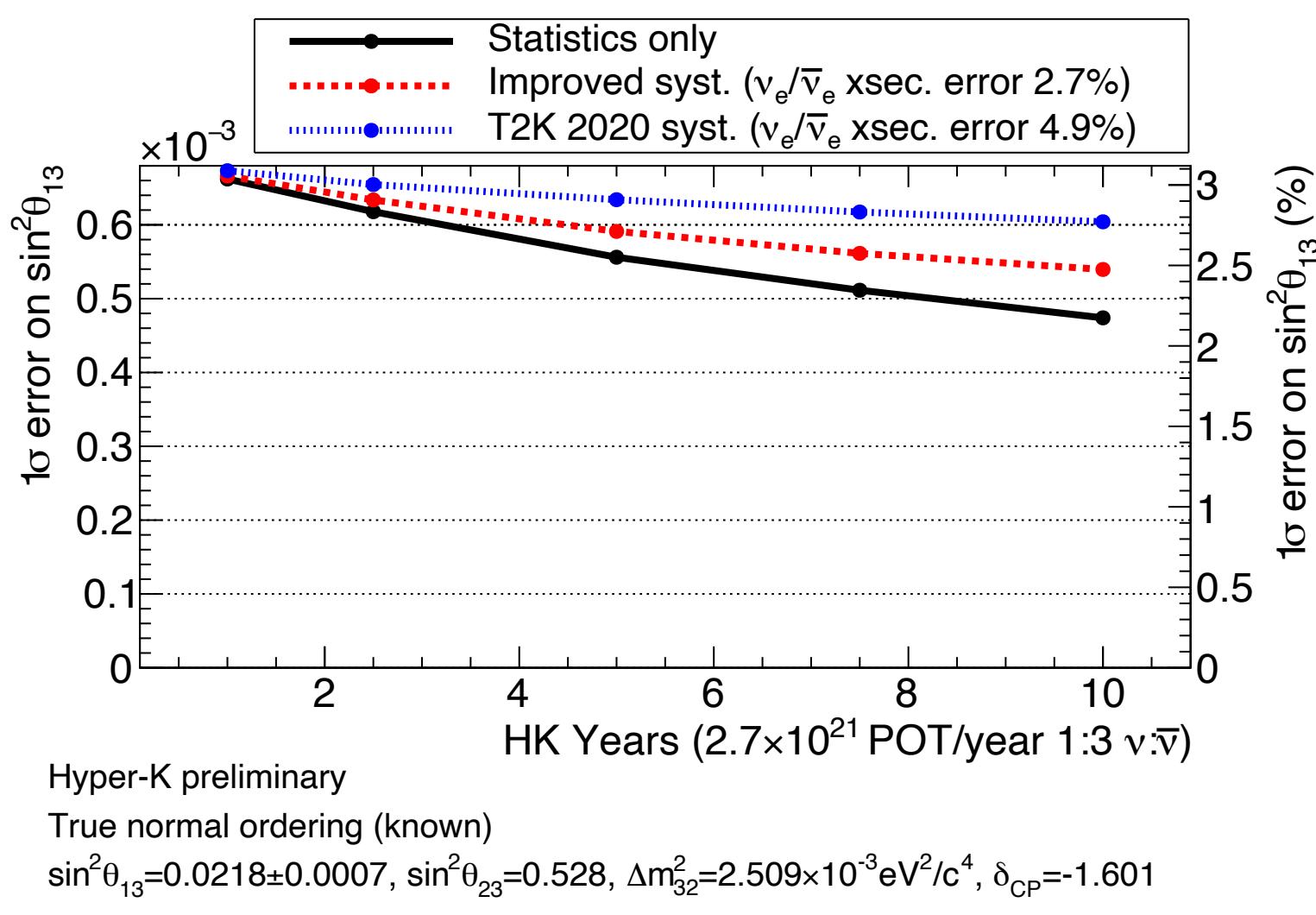
- Precision of Δm^2_{32}
 $\sim 2\%$ \rightarrow 0.36%
- JUNO more precise



- Octant of θ_{23} determined at 3σ for $\Delta\theta_{23} \gtrsim 2^\circ$

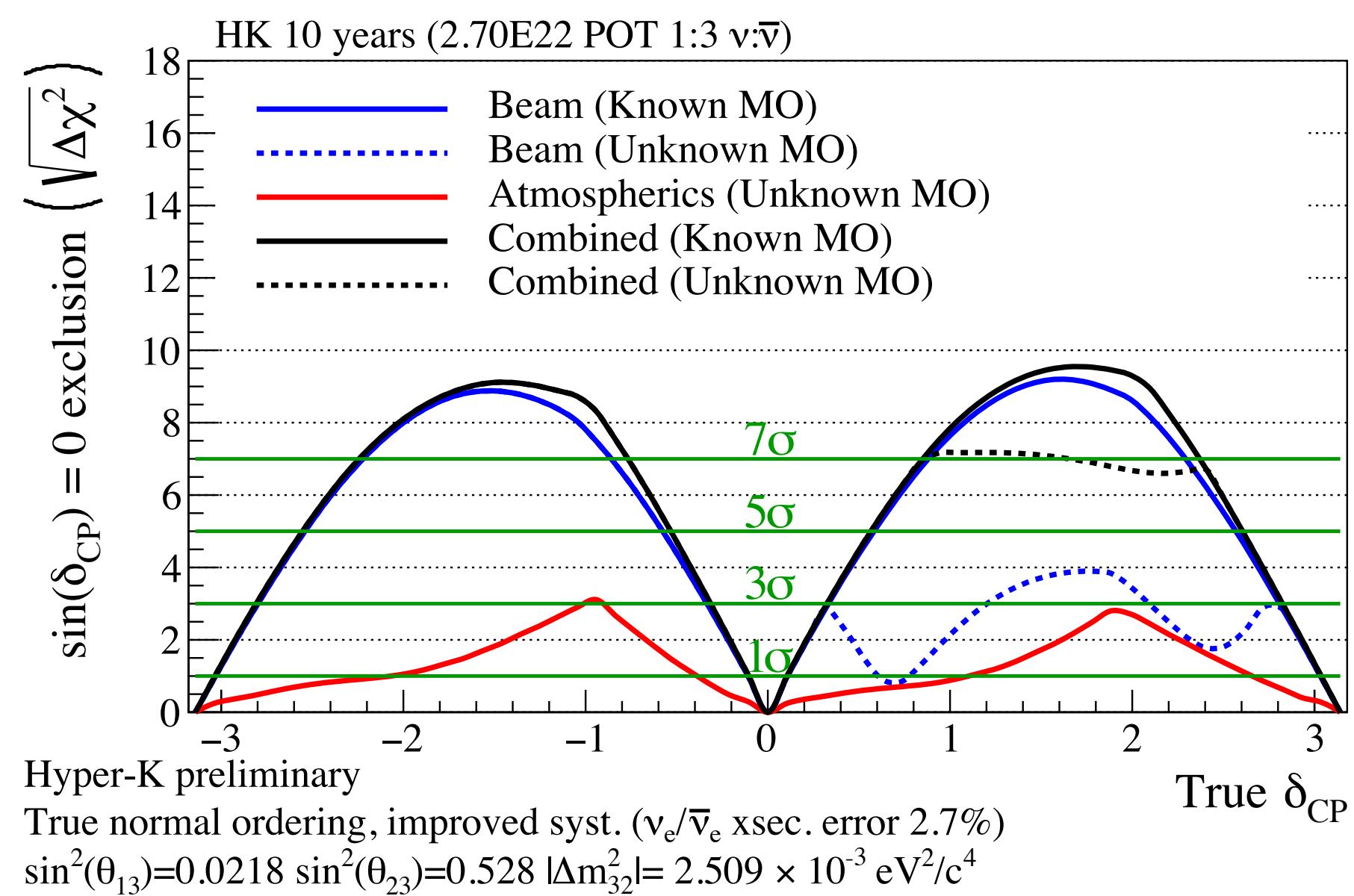
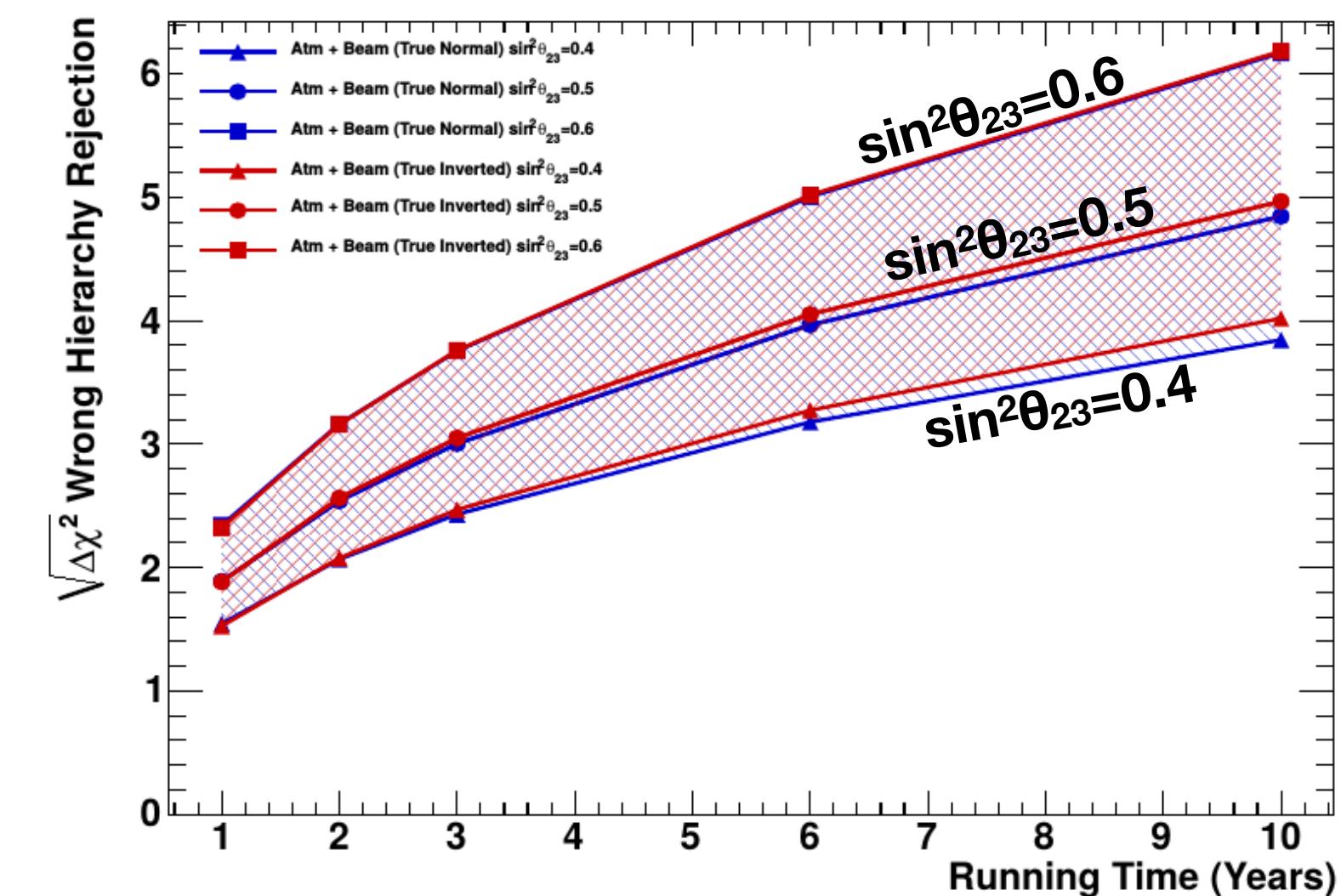
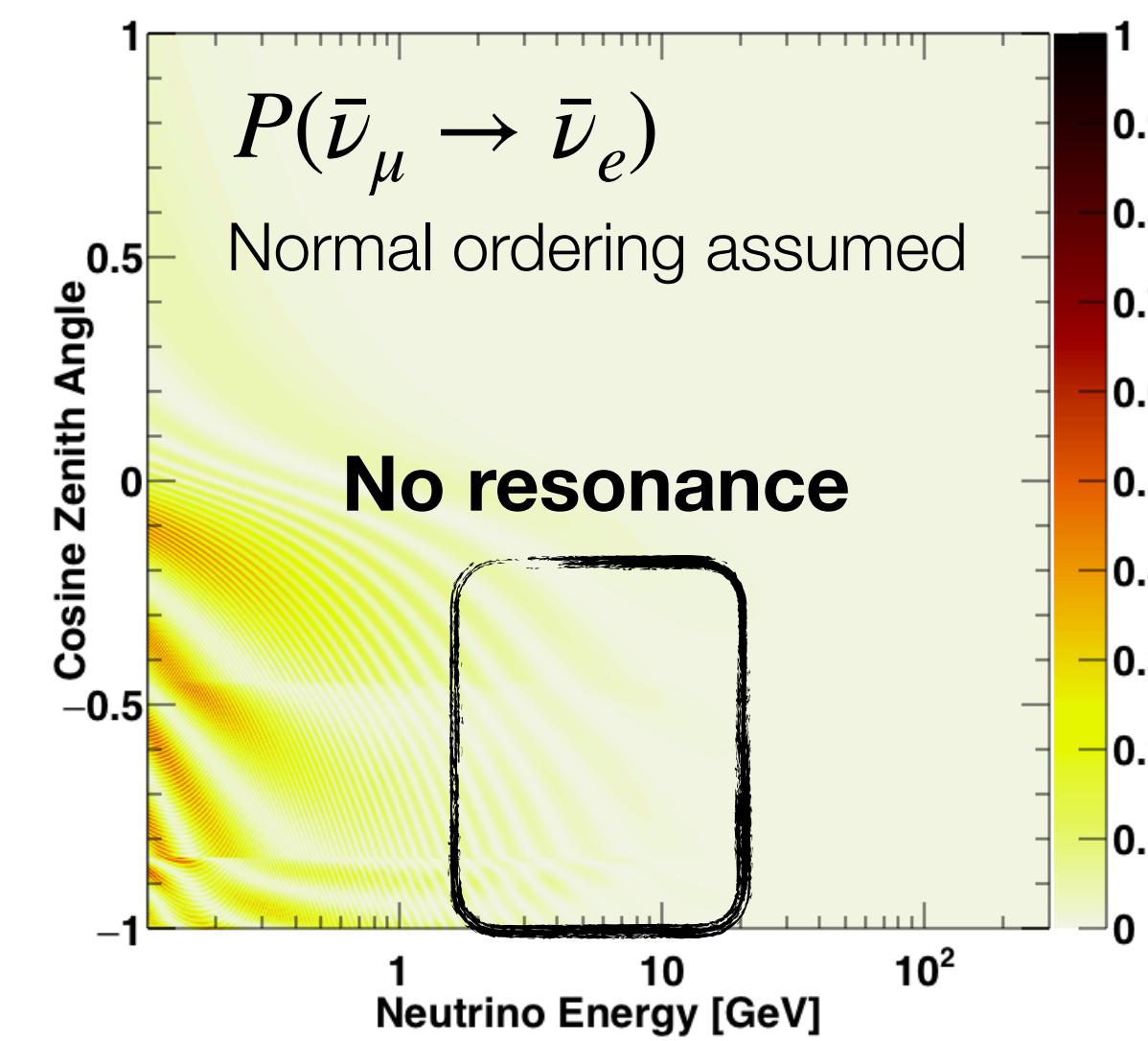
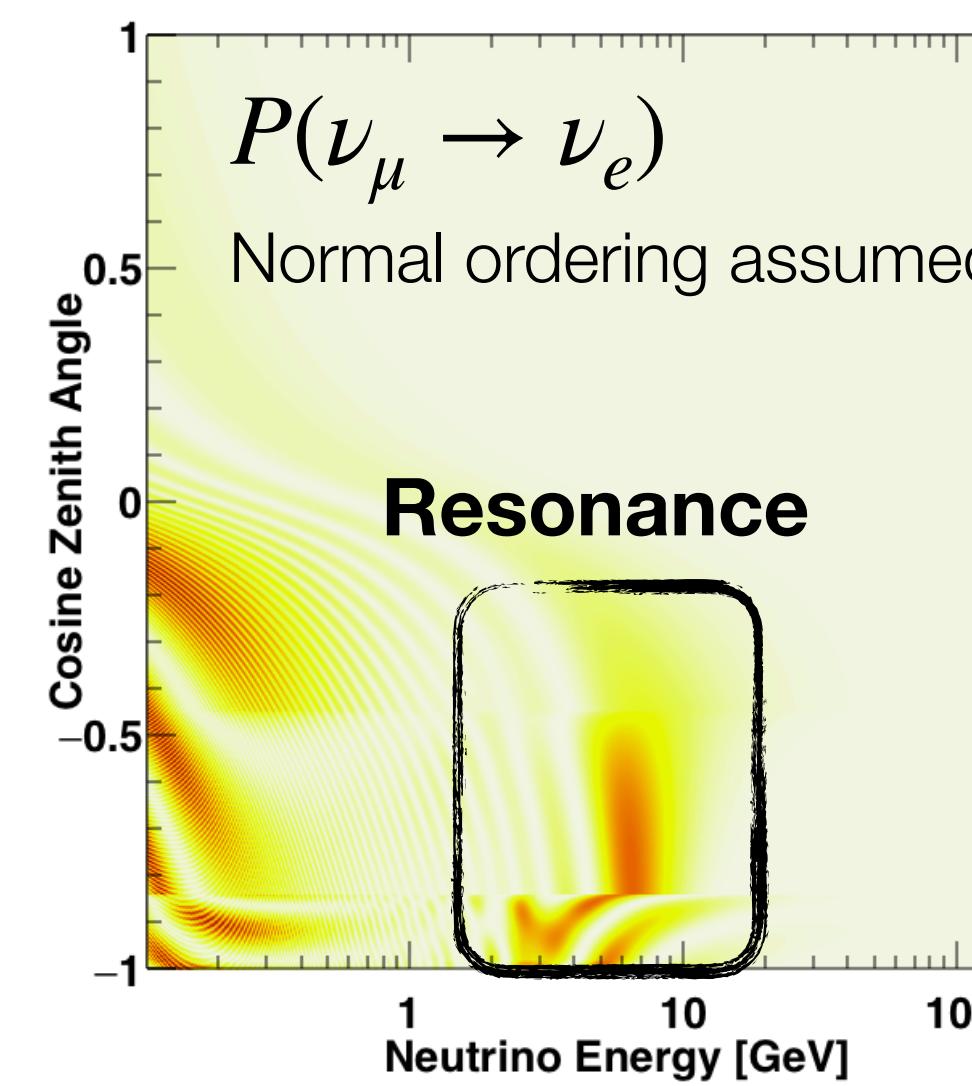


- Assuming react. constraints
- Precision of θ_{13}
 $\sim 2.8\%$ \rightarrow 2.5%
- No improvement w/o constraints due to θ_{13} - θ_{23} correlation

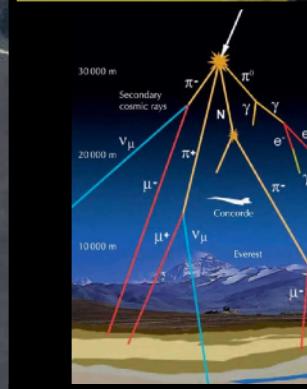
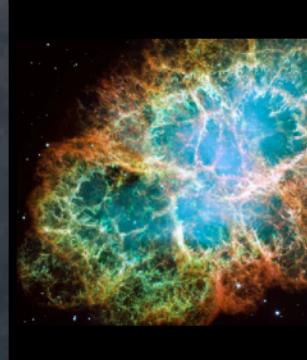
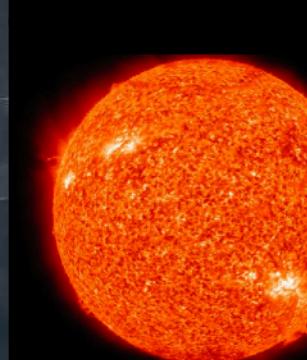
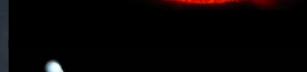
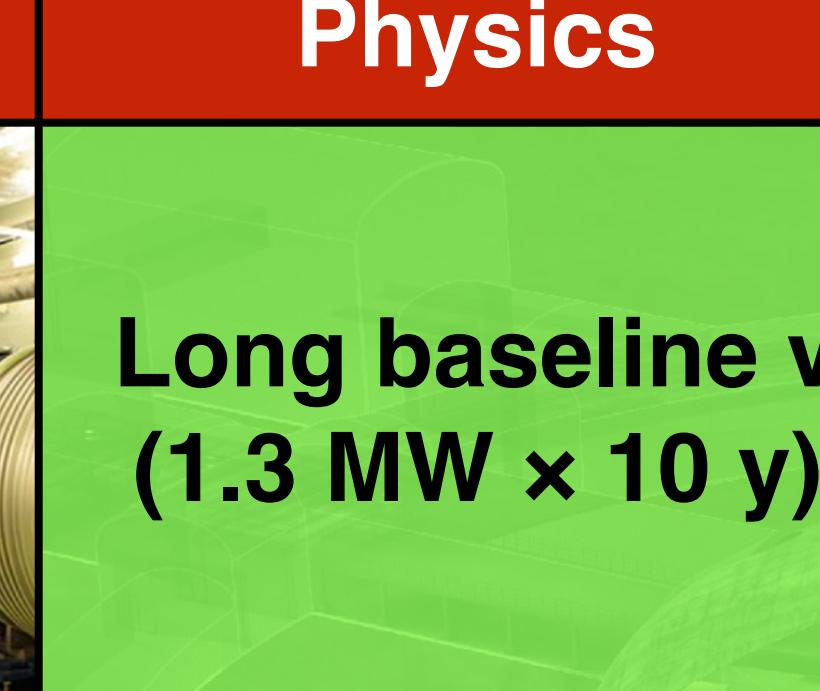


Atmospheric Neutrinos

- Atmospheric neutrinos sensitive to mass ordering through matter effect (higher energy + longer baseline in the matter)
- Breaks out degeneracy between mass ordering and δ_{CP}
 - Long baseline δ_{CP} measurement no longer relies on known ordering when combined with atmospheric neutrinos
 - Exclusion of wrong ordering on about $4\text{-}6\sigma$ in 10 y, depends on true value of θ_{23} ($>3\sigma$ in 6 y)



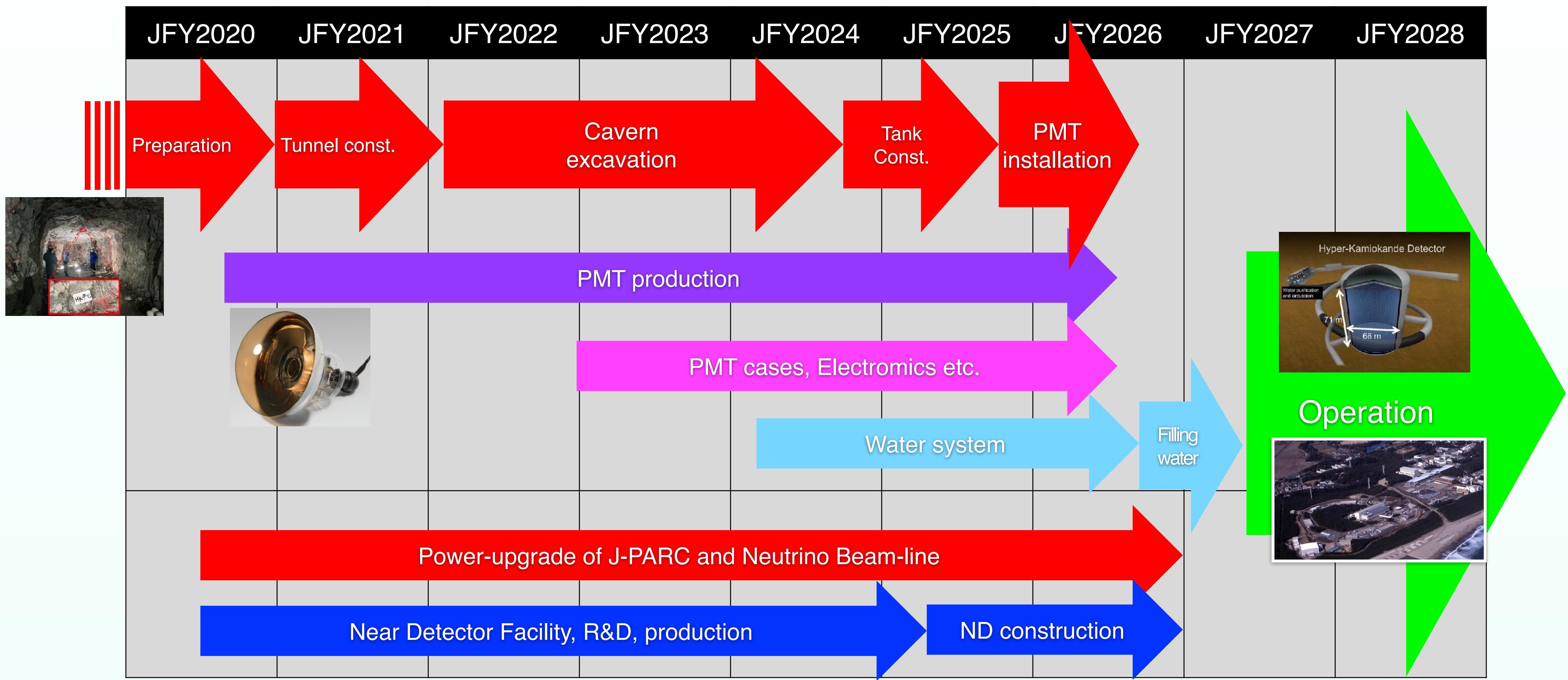
Physics Program Summary

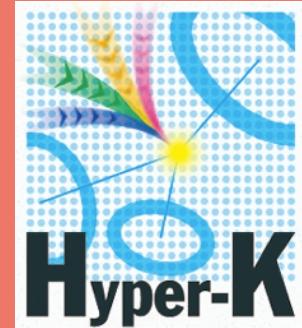
	Physics	Parameter	Sensitivity
 Long baseline ν (1.3 MW × 10 y)		δ_{CP} precision	7-20°
		δ_{CP} coverage (3/5σ)	76%/58%
		$\sin^2\theta_{23}$ error (for 0.5)	±0.017?
 Atm. ν + LBL ν (10 y)		Mass ordering	>3.8σ
		θ_{23} octant (3σ)	$\theta_{23}-45^\circ$ >2°
 Supernova ν		SN burst @ 10 kpc	54k-90k ν's
		SN Relic ν (10 y)	~70 ν's
 Solar ν (10 y)		Upturn	>3σ
		Day/Night asymmetry	8σ
 Proton decay (20 y)		τ for $p \rightarrow e^+ \pi^0$ (3σ)	1×10^{35} y
		τ for $p \rightarrow K^+ \nu$ (3σ)	3×10^{34} y

Schedule



- Construction: 2020-2027 - on time
- Operation: 2027 -





J-PARC Beam Upgrade

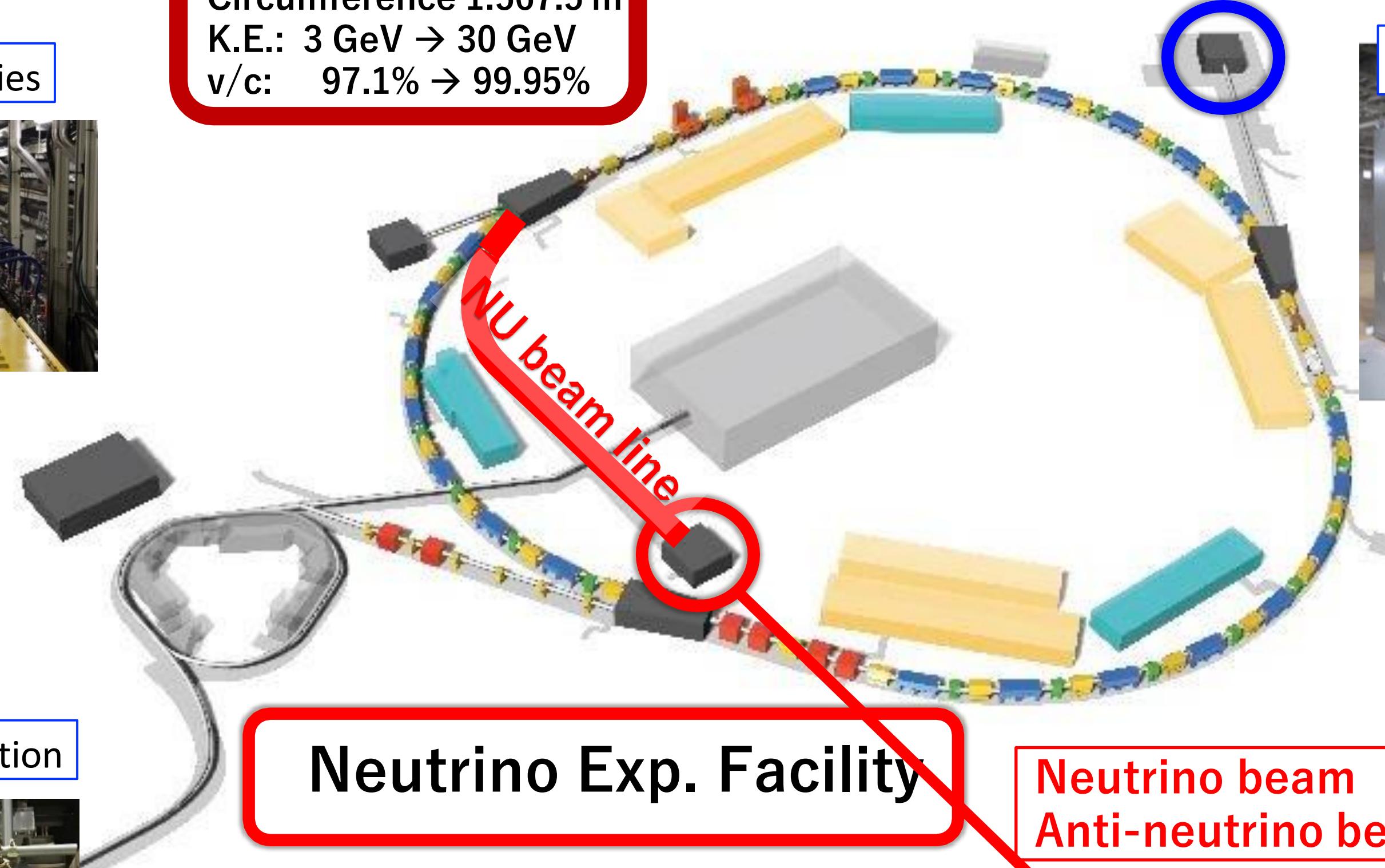


MR-RF cavities

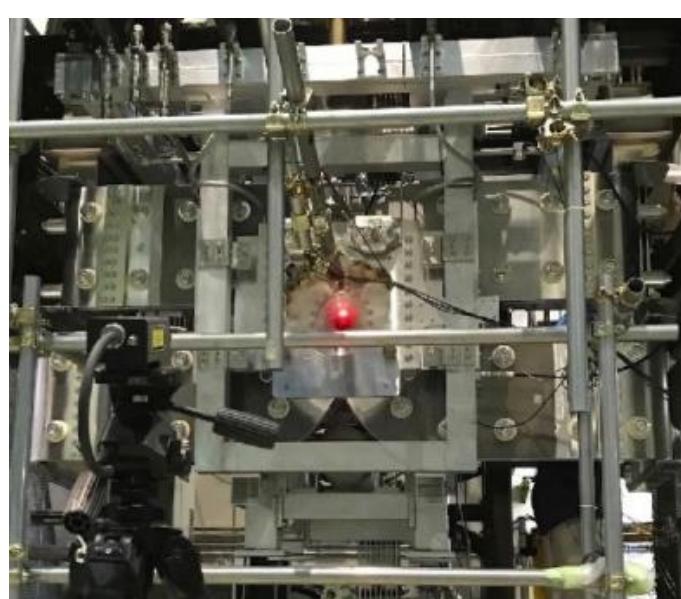


Main Ring

Circumference 1.567.5 m
K.E.: 3 GeV → 30 GeV
v/c: 97.1% → 99.95%



320kA horn operation

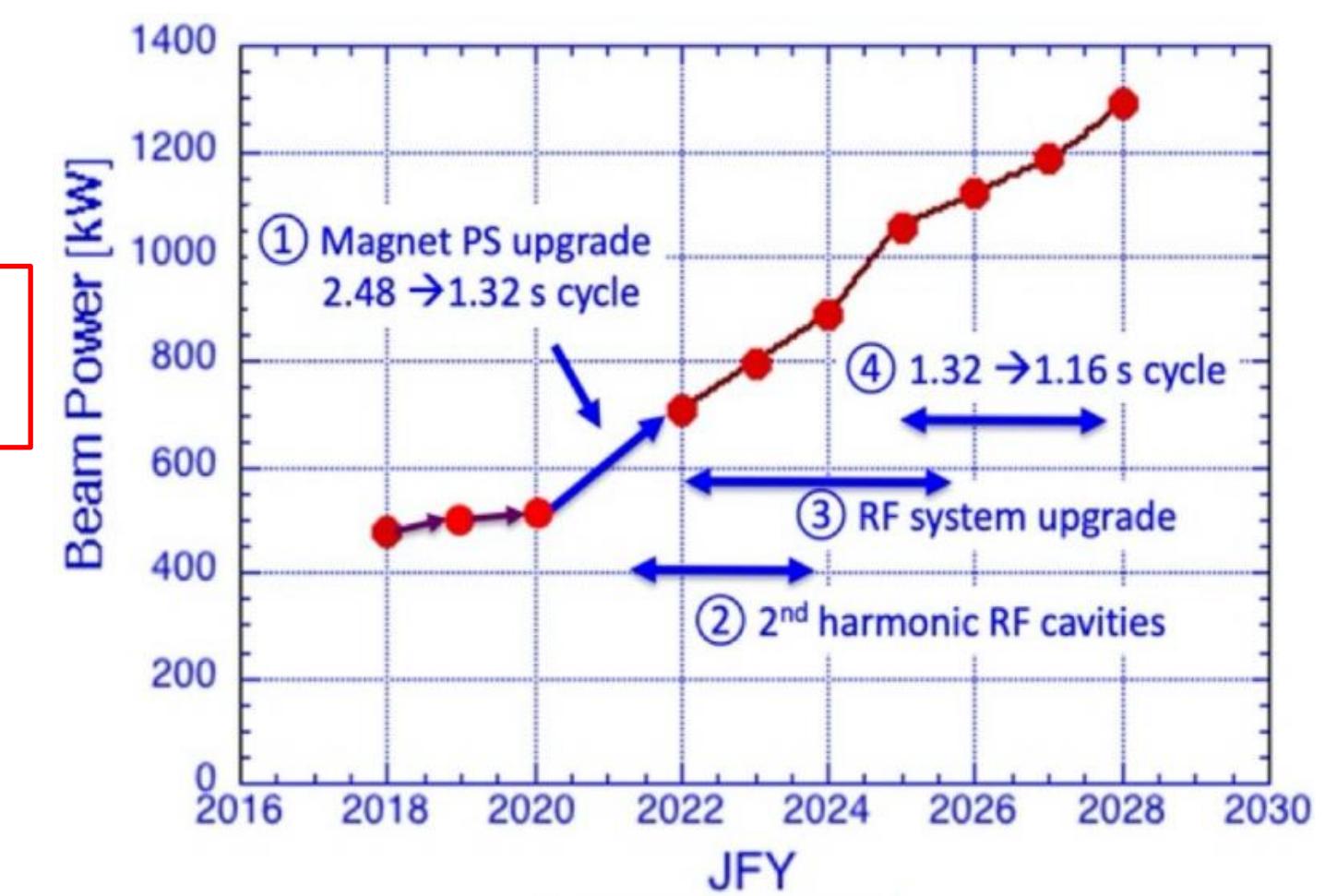


Neutrino Exp. Facility

Neutrino beam
Anti-neutrino beam

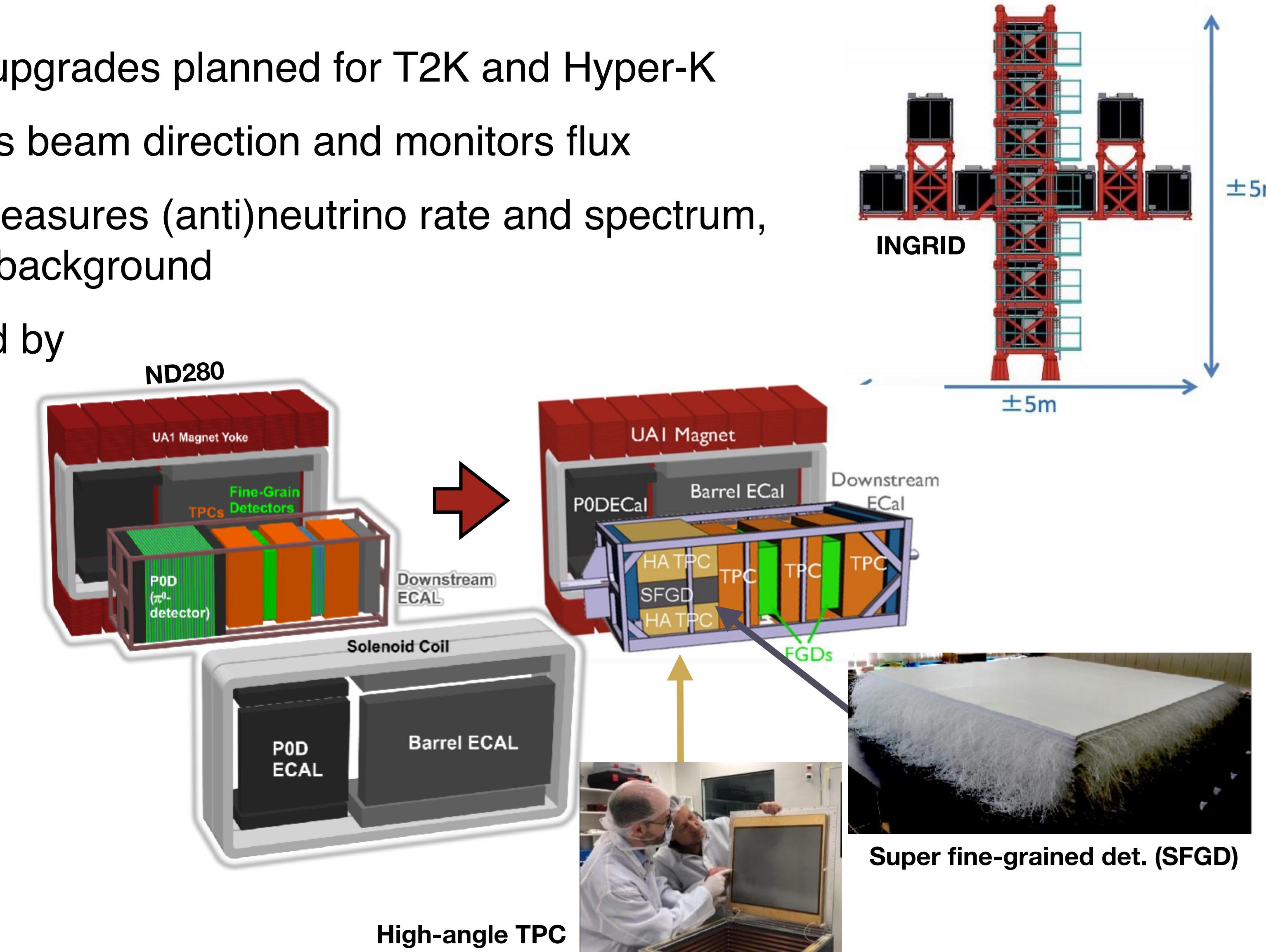
Achieved 515 kW in JFY2020
Aiming 1.3 MW by JFY2028

New main magnet PS for high rep. rate



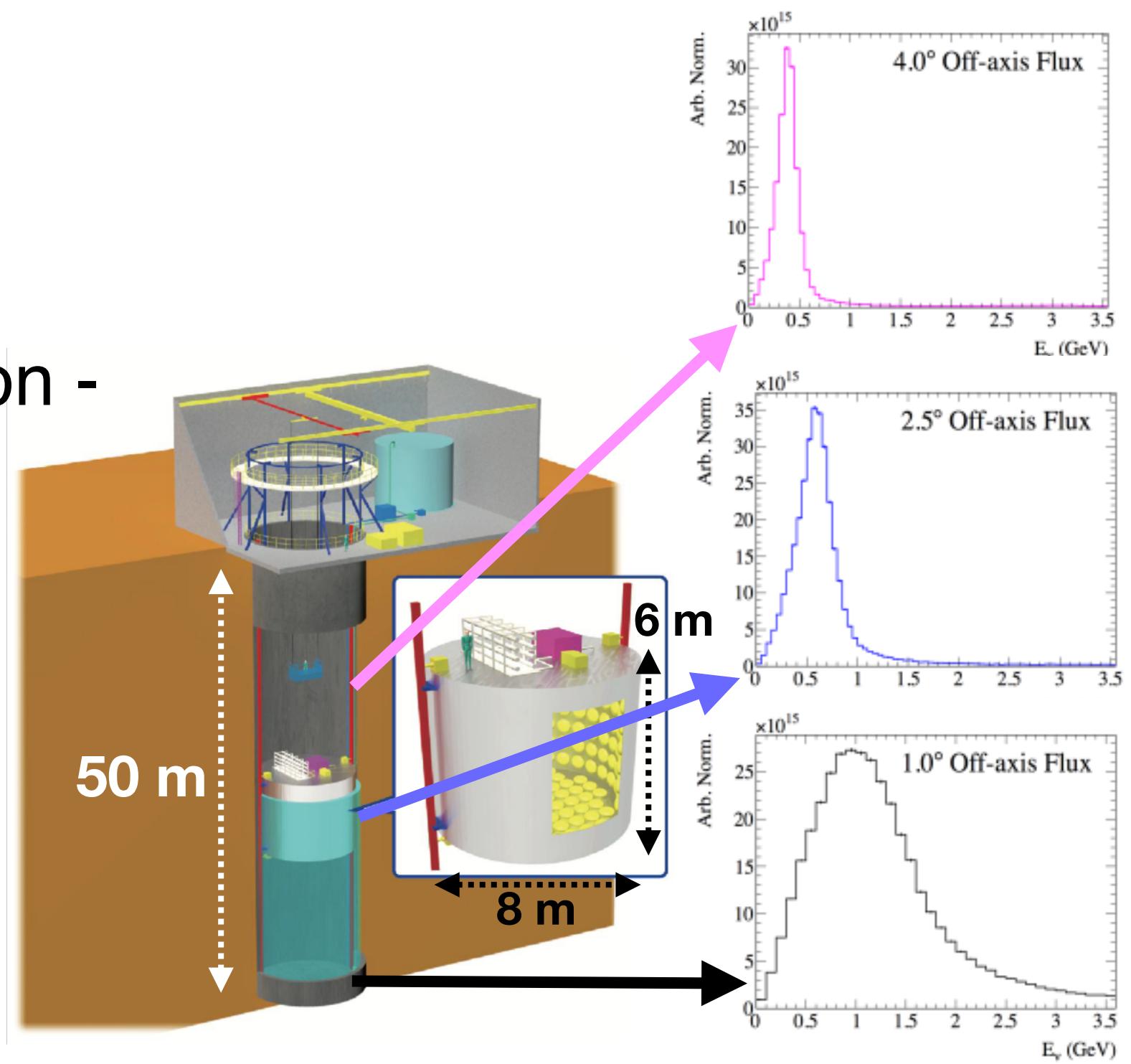
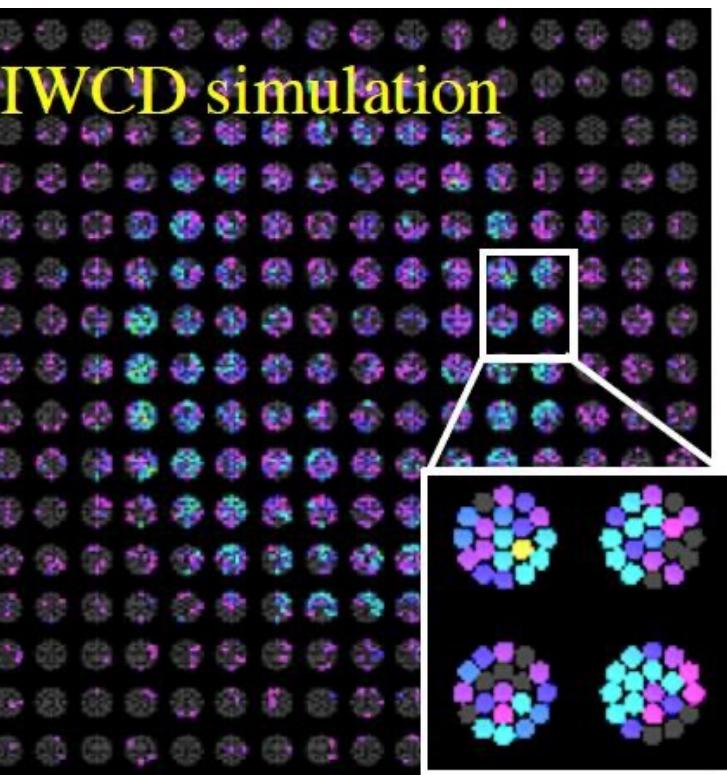
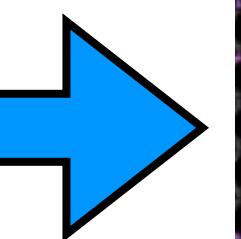
Near Detectors

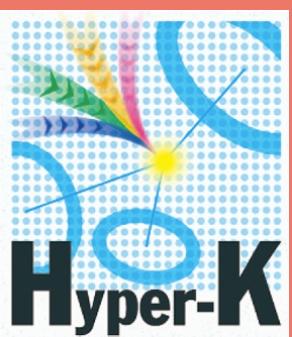
- Near detectors common with T2K - upgrades planned for T2K and Hyper-K
- On-axis detector INGRID - measures beam direction and monitors flux
- Off-axis detector ND280 - primary measures (anti)neutrino rate and spectrum, magnetized to separate wrong-sign background
 - Upgrade for/by T2K: P0D replaced by
 - Super fine-grained detector
 - High-angle TPCs
 - Time-of-flight planes
 - Further upgrades for Hyper-K era under investigation



Intermediate Water Cherenkov Detector

- Newly constructed movable (1° to 4° off-axis) 1 kt water Cherenkov detector at ~1 km from target
- Instrumented with 400 multi-PMT modules
 - 19 8 cm PMTs, dedicated fast electronics, etc.
 - High granularity and superb time resolution
- Precise cross-section measurement
 - Same target material and detection technique as Hyper-K
 - Reduction of dominant systematics of $\nu_e/\bar{\nu}_e$ interaction cross section - ND280+IWCD complementary reduce the systematics $\leq 3\%$ (alone $\leq 7.5\%$ and $\leq 3.7\%$ respectively)
- Precise measurement of the (anti)neutrino spectrum
 - Sampling at different off-axis angles to get a spectrum prediction for the Hyper-Kamiokande far detector



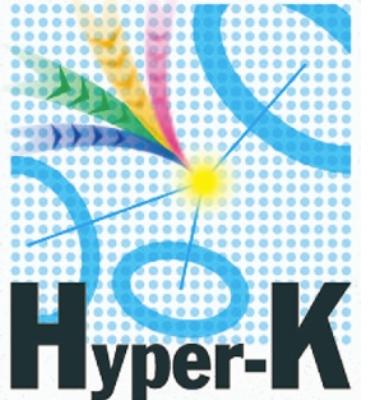


Far Detector - 3rd Generation of Successful Experiments



■ Hyper-Kamiokande: 2027-onwards

- Extend nucleon decay search
- Precision measurement of neutrino oscillation
- Determination of CP violation and mass ordering, ...
- Unknown unknowns?



$+8.4 =$



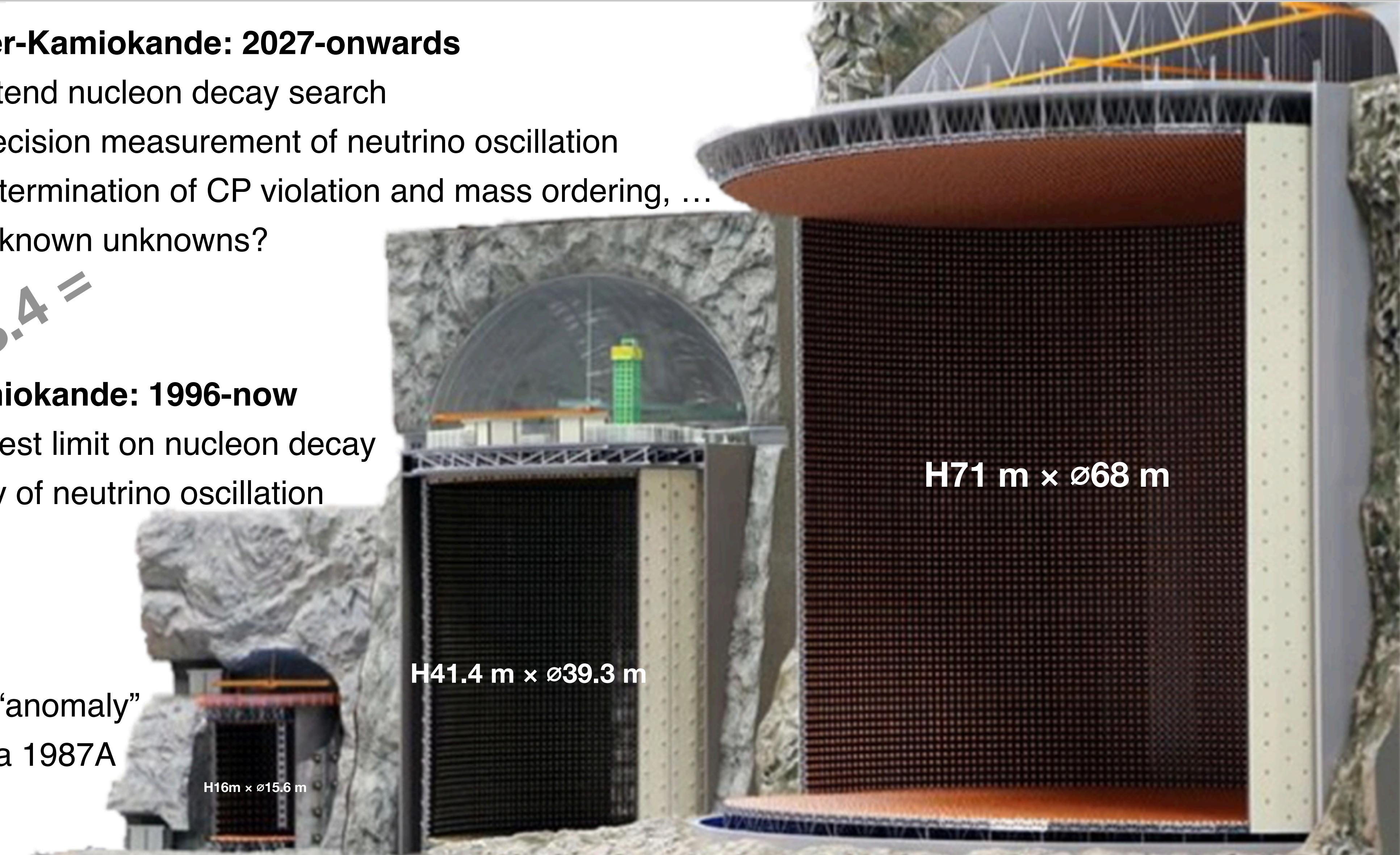
■ Super-Kamiokande: 1996-now

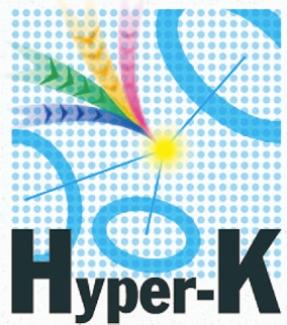
- World's best limit on nucleon decay
- Discovery of neutrino oscillation

$+20 =$

■ Kamiokande: 1983-1996

- Atm. and solar neutrino “anomaly”
- Neutrino from Supernova 1987A

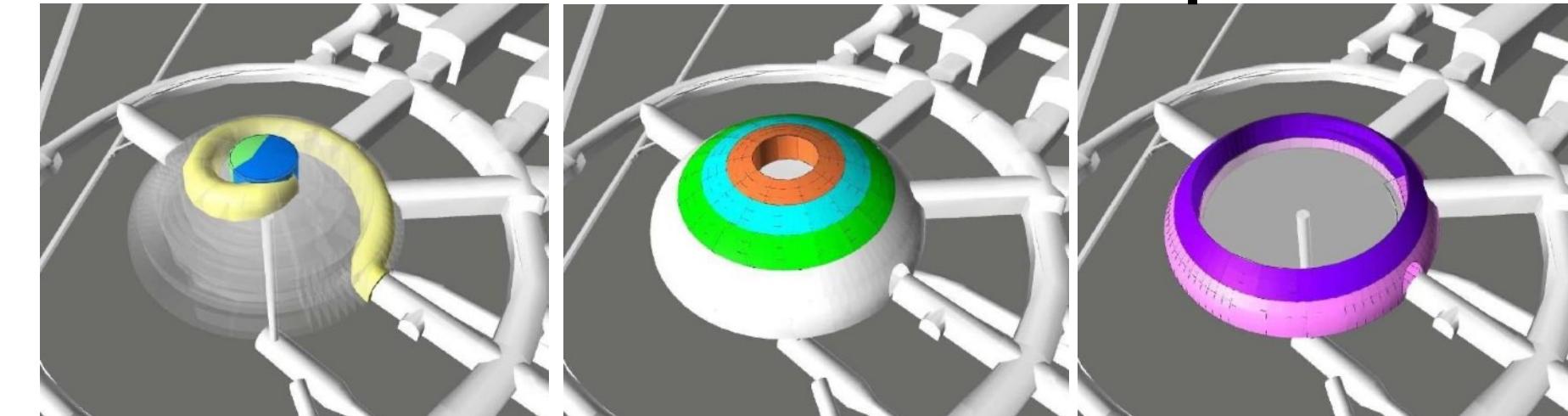




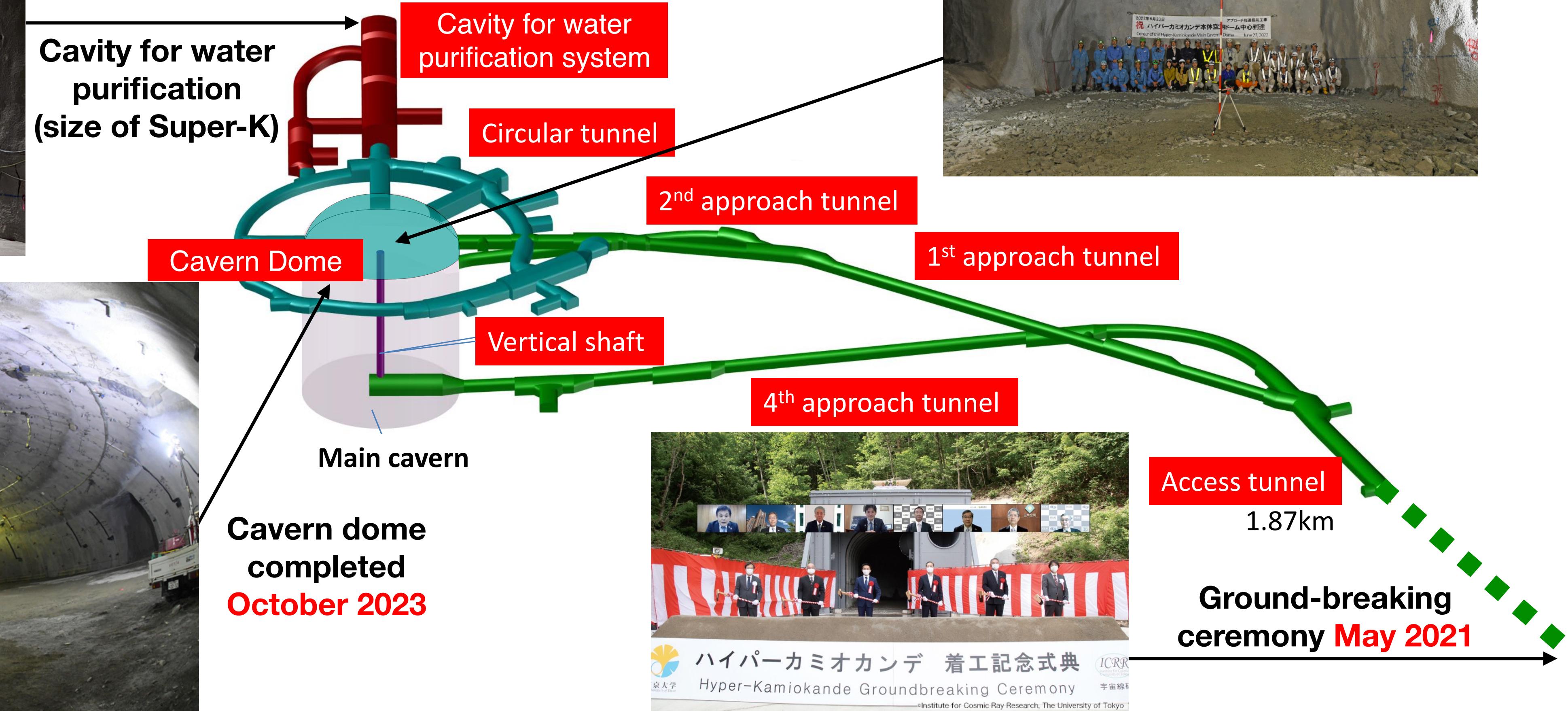
Far Detector Civil Construction on Schedule



Cavern dome excavation sequence



Cavity for water
purification
(size of Super-K)



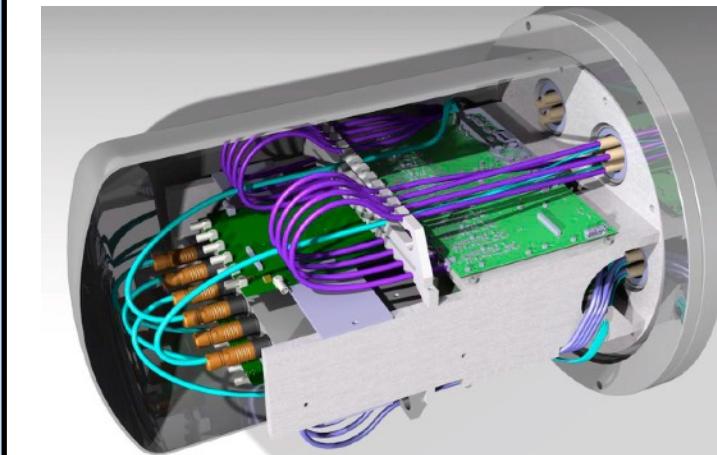
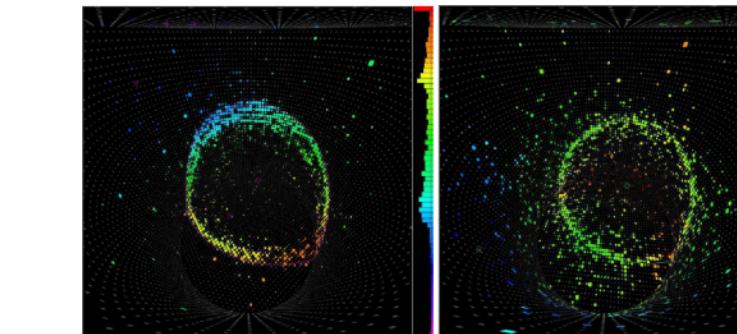
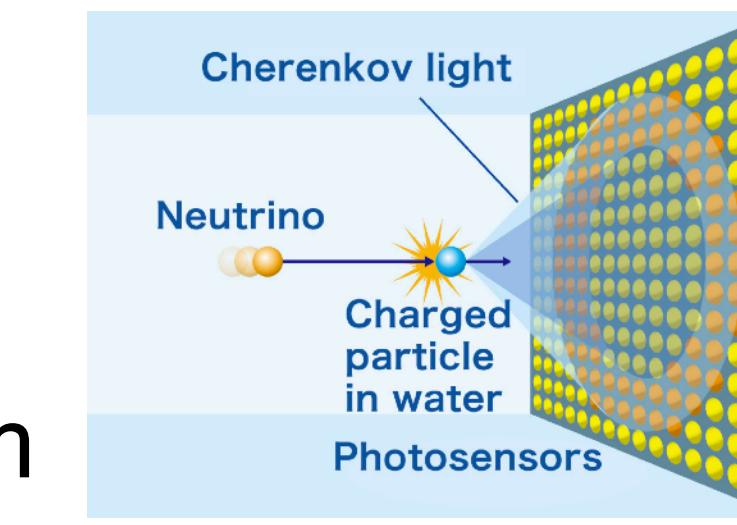
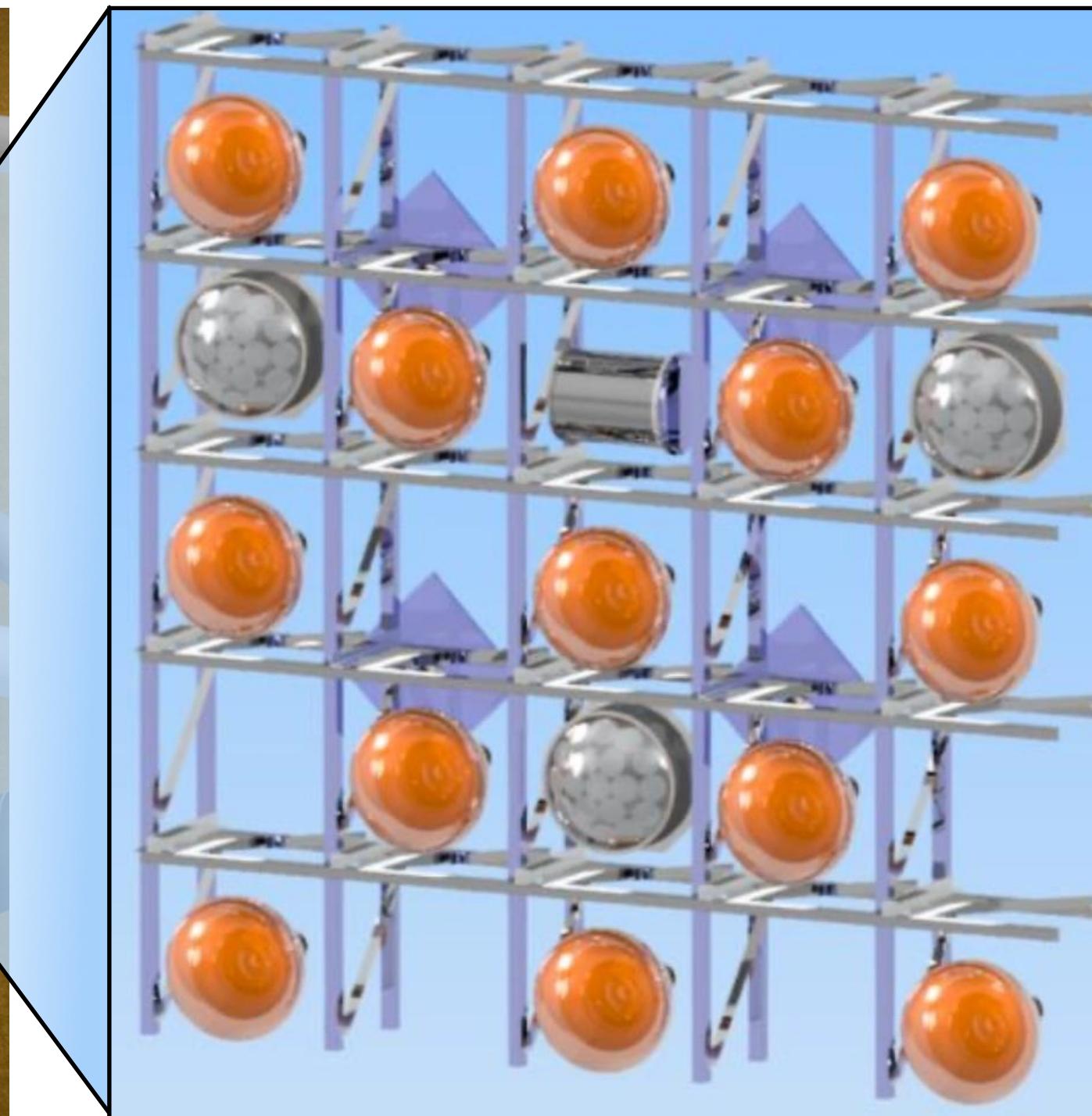
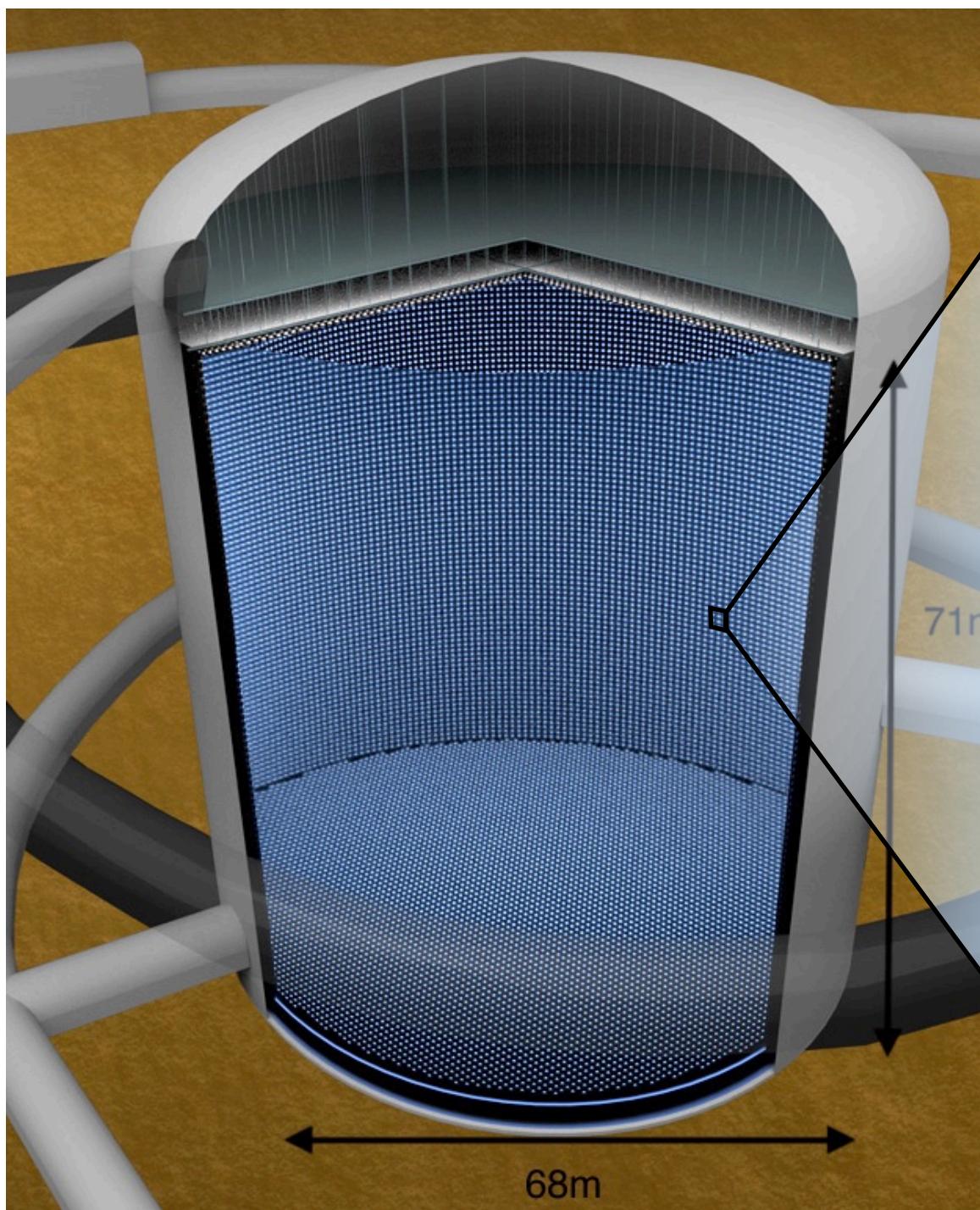
Center of the cavern dome reached **July 2022**



ハイパーカミオカンデ 着工記念式典
Hyper-Kamiokande Groundbreaking Ceremony
Institute for Cosmic Ray Research, The University of Tokyo

Hyper-Kamiokande Far Detector Overview

- 260kt water Cherenkov detector
 - 1700 m.w.e. underground
 - Inner detector (ID) H64.8 m × ø65.8 m
 - Fiducial mass of 190 kt (8.4×SK)
 - Outer detector (OD) 1m/2m thick



- 20k 50 cm PMTs
- Main detection system of ID
- 20% photo-coverage
- ~800 mPMTs modules
- Direction and time resolution
- Direct/scattered light
- 3600 OD modules
- 8cm PMT and WLS plate
- Provides veto against muons
- Underwater electronics vessels
- ID & Hybrid (ID+OD) vessel
- mPMT signal-collecting vessel

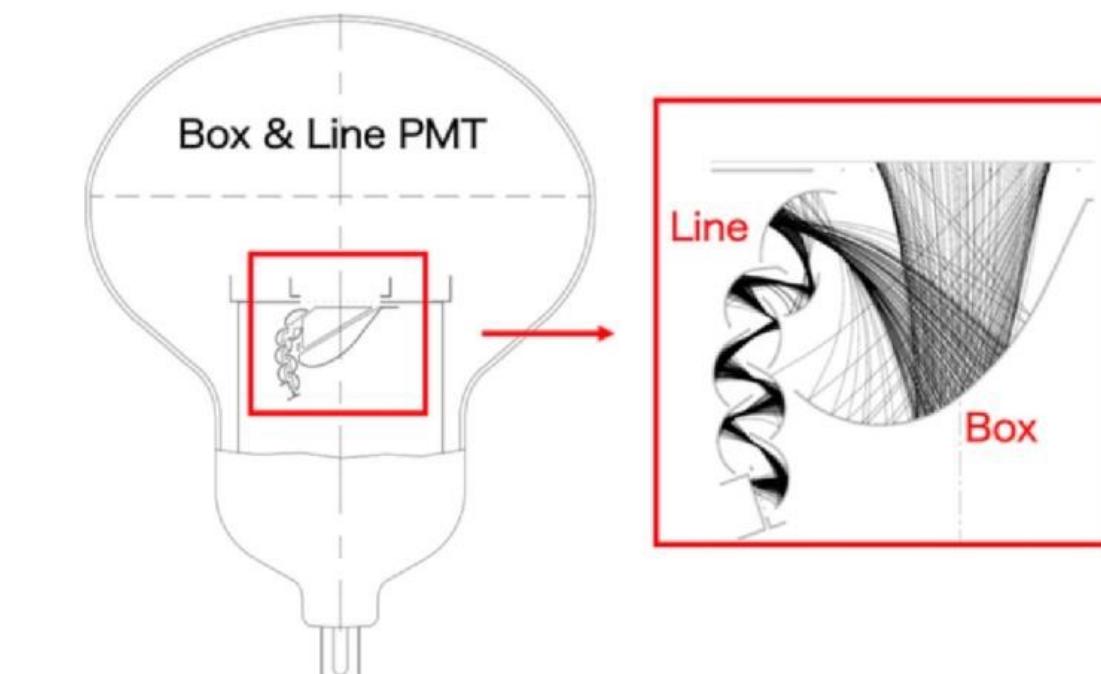
50 cm PMTs



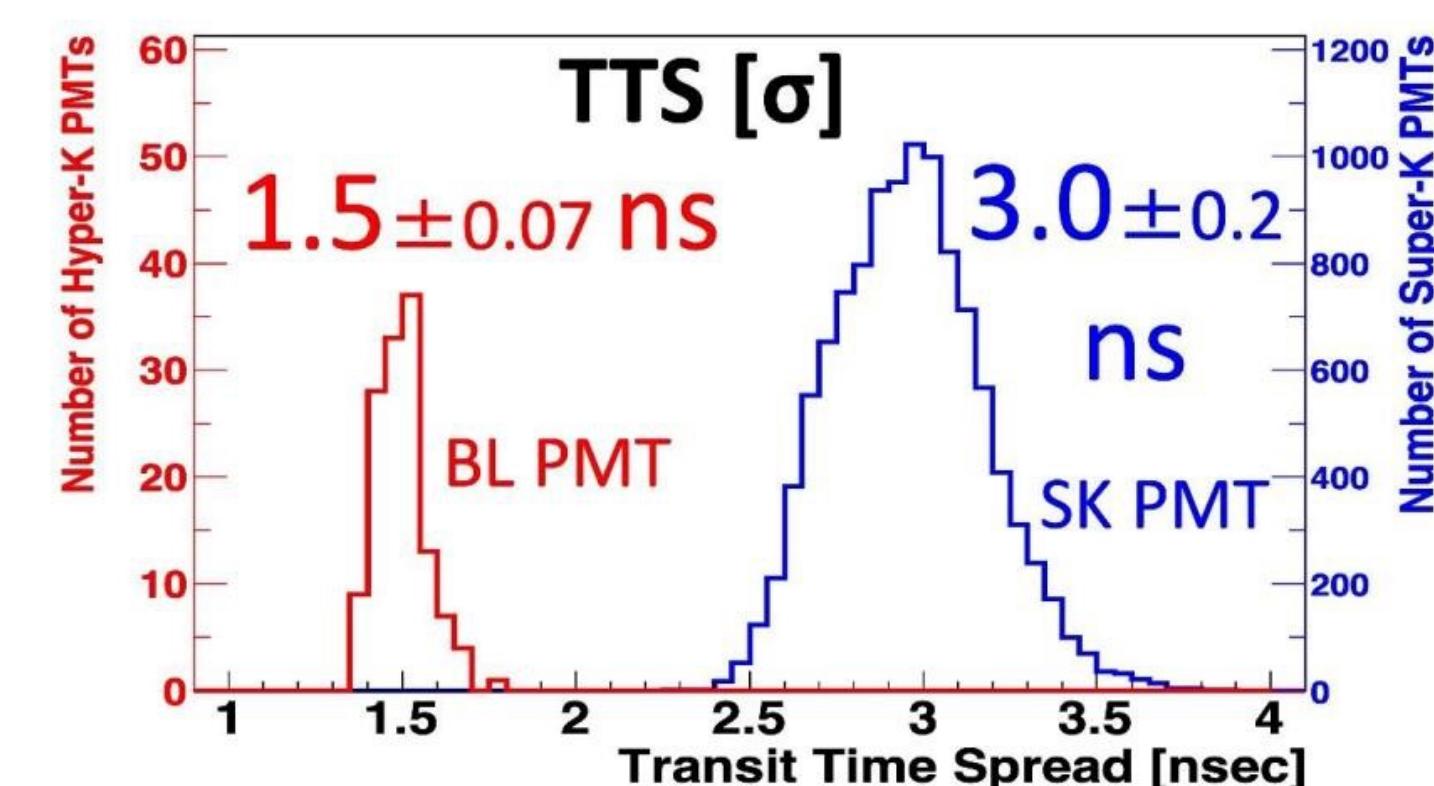
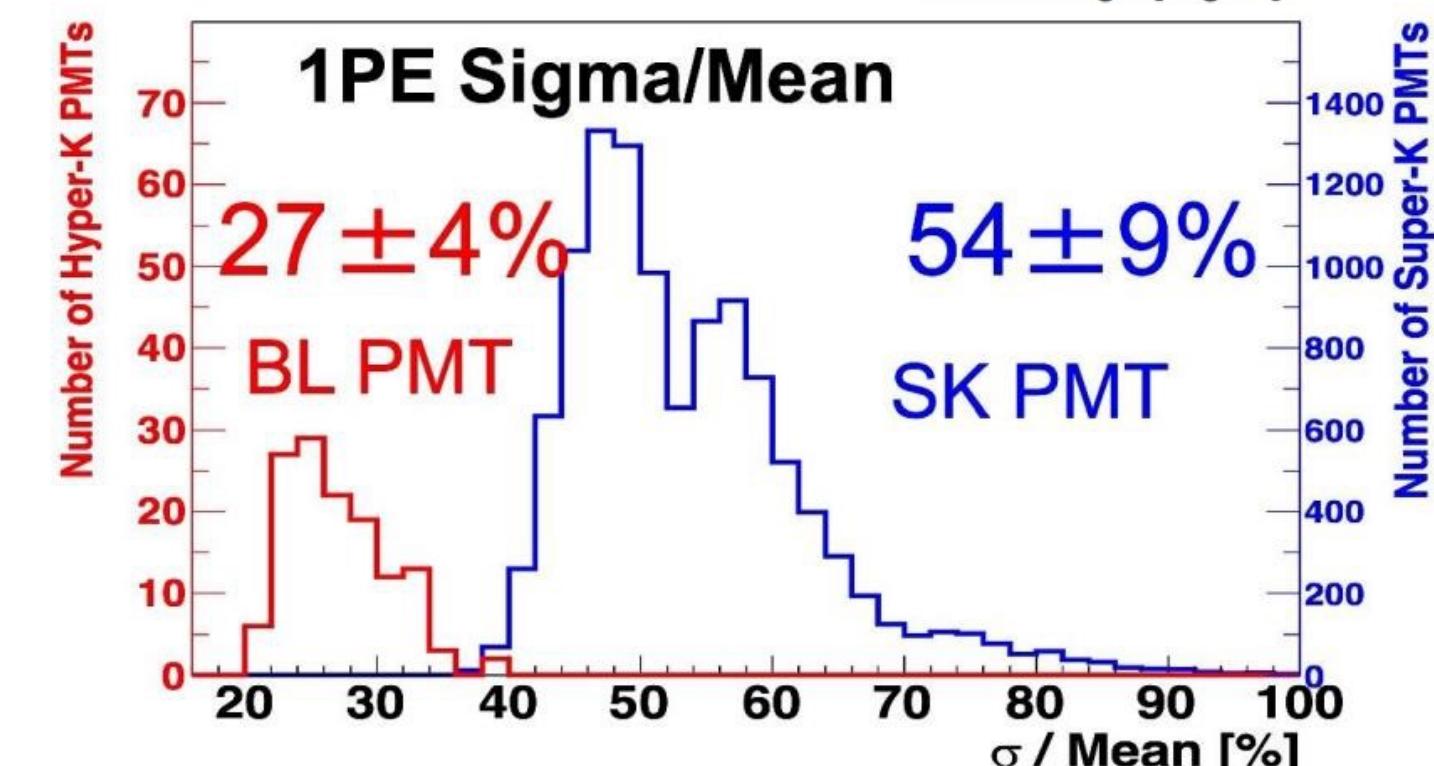
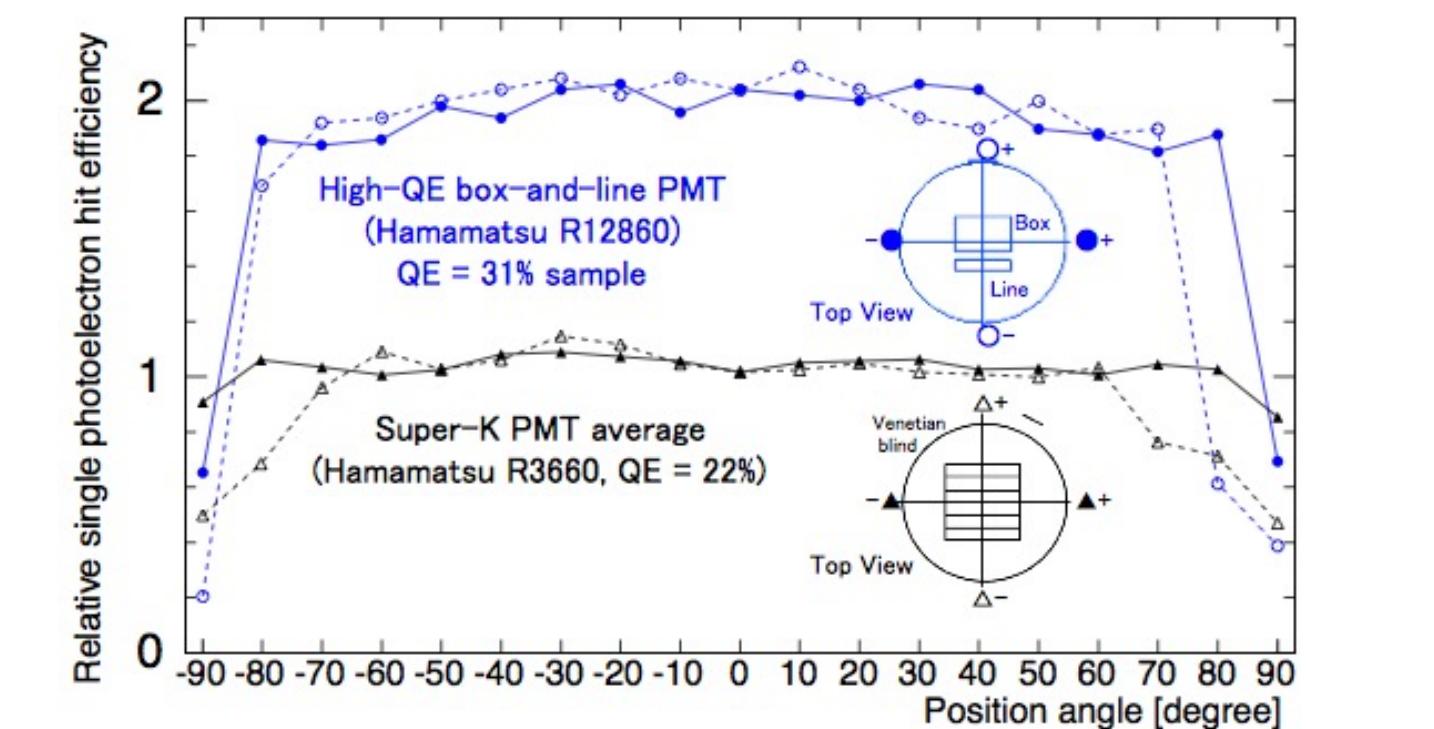
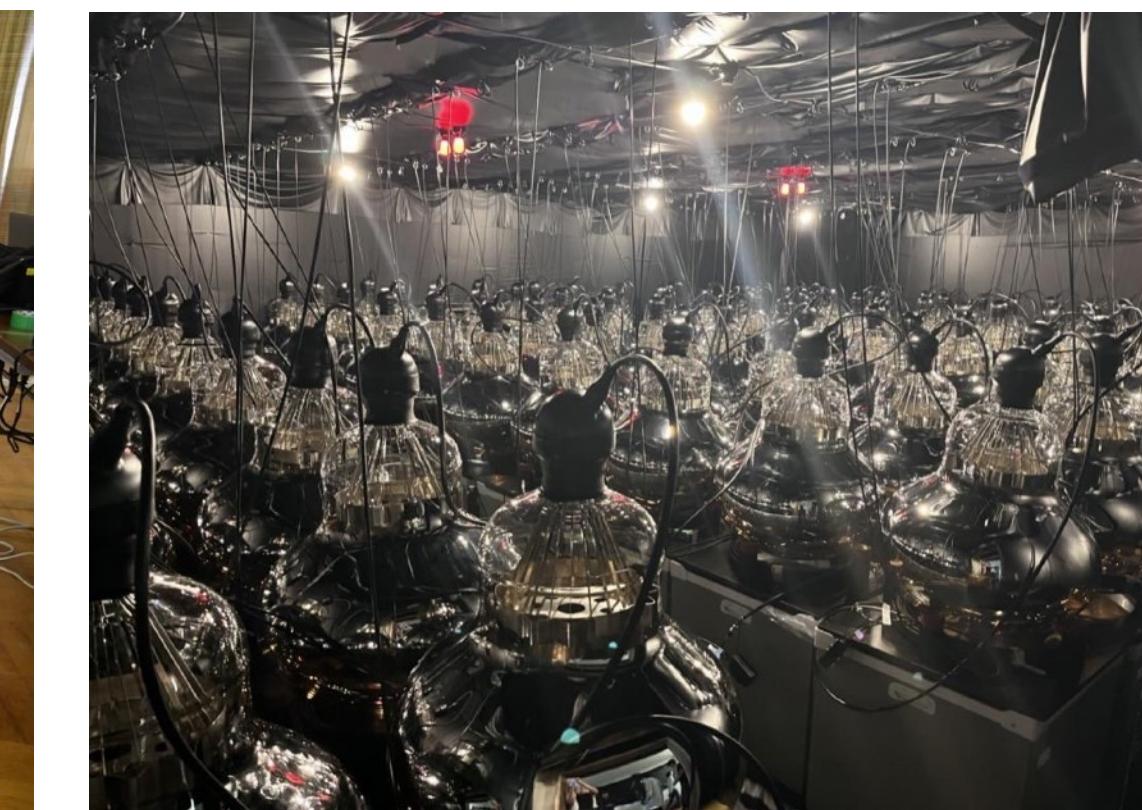
■ Hamamatsu Photonics R12860 Box&Line dynode PMTs

- $\times 2$ better photon detection efficiency
- $\times 2$ better single photoelectron resolution
- $\times 2$ better time resolution
- $\times 2$ higher pressure tolerance
- Low dark rate of 4 kHz

(all w.r.t. SK
Hamamatsu
R3660)

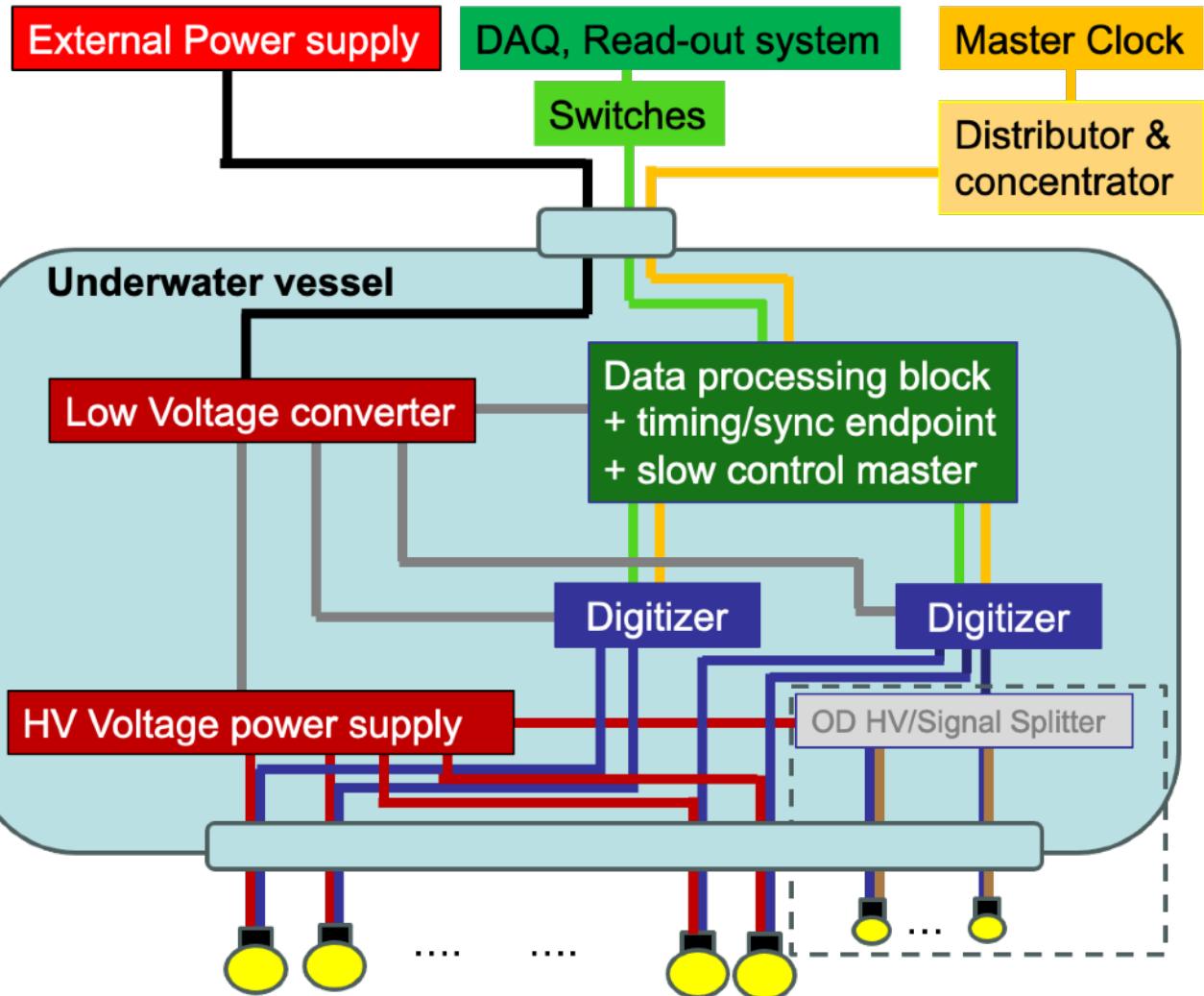
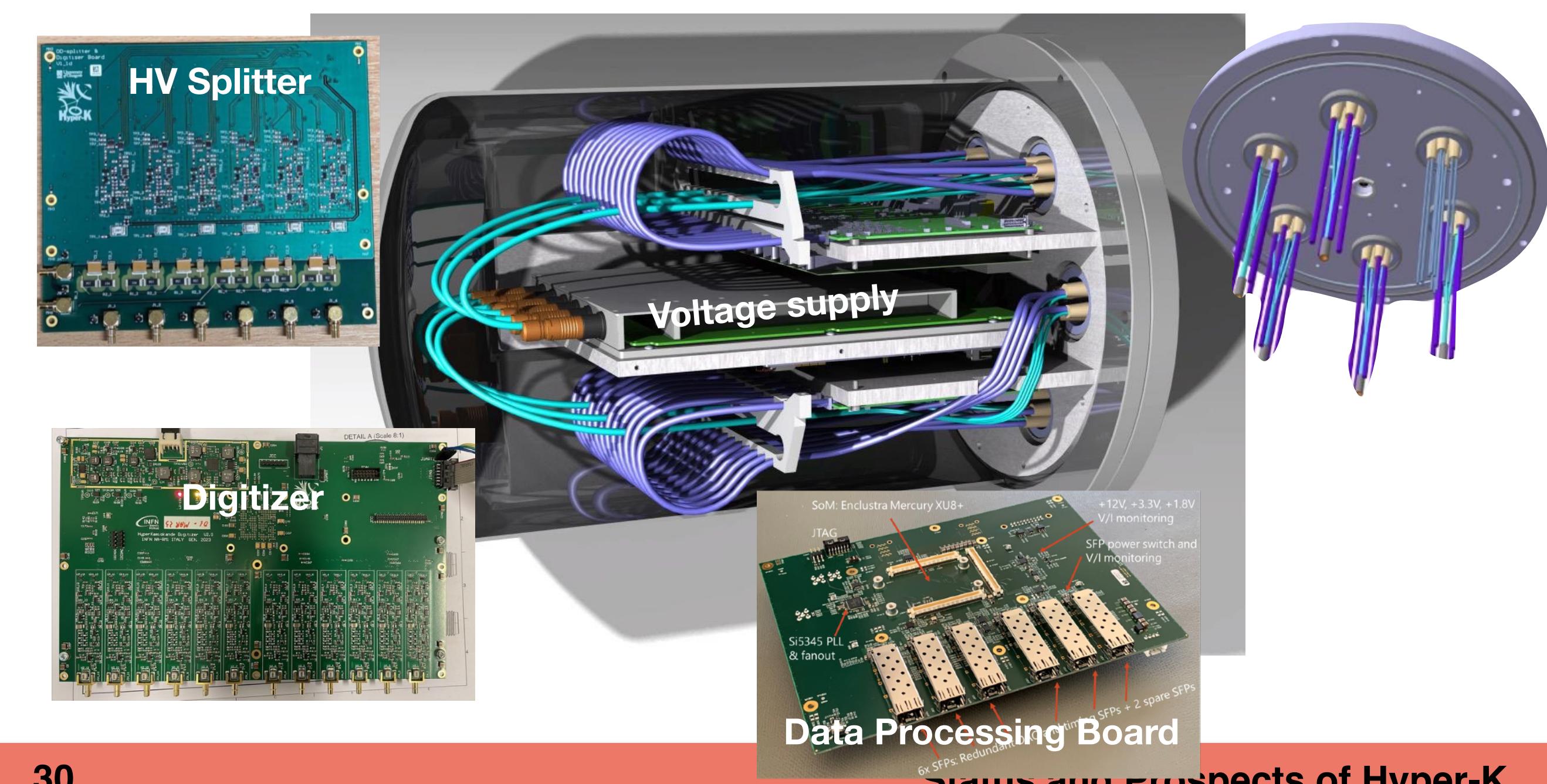
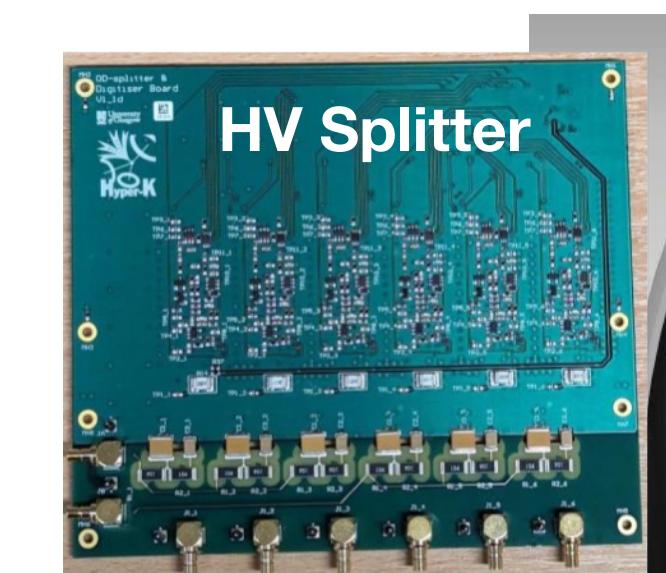
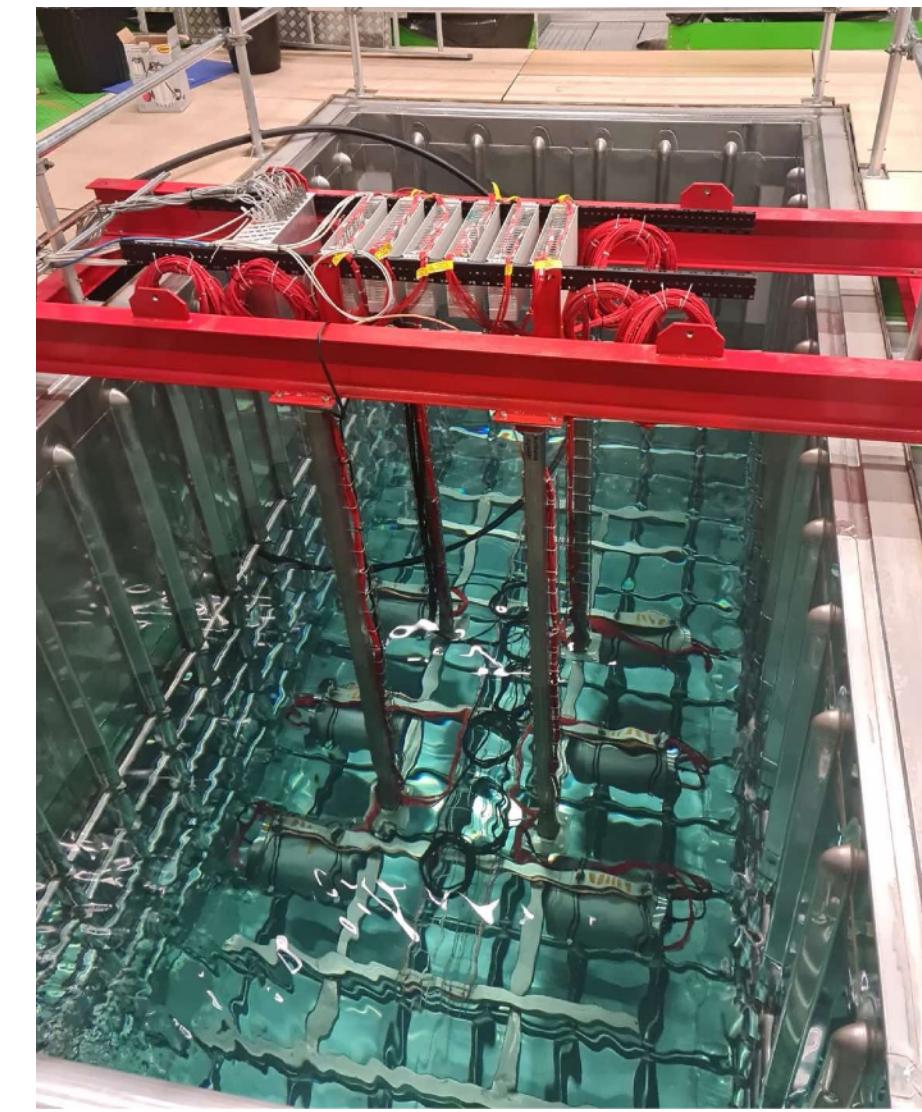


- ## ■ Production resumed - 6000+ delivered and tested @Kamioka
- Previous issues due to flashers successfully resolved



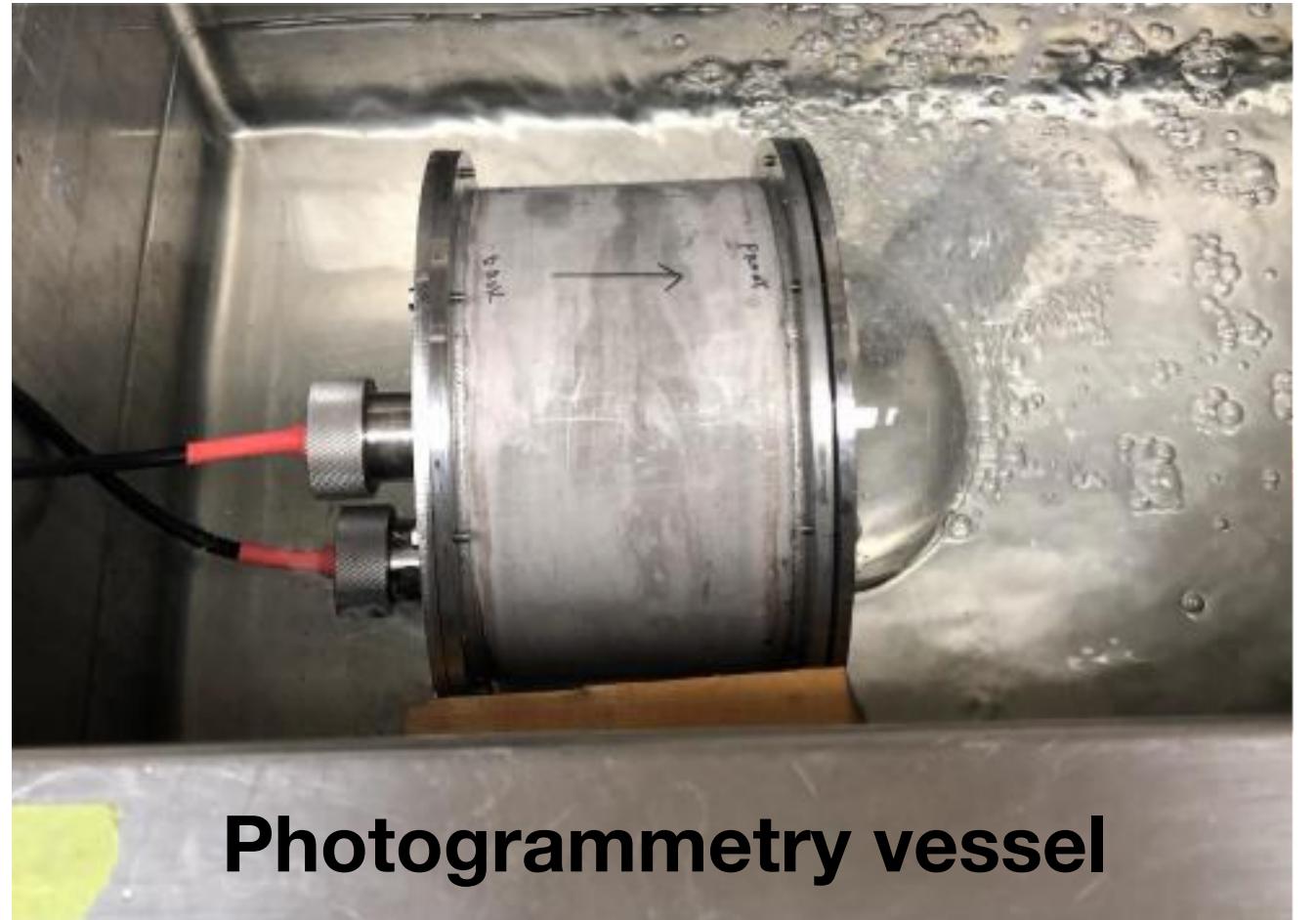
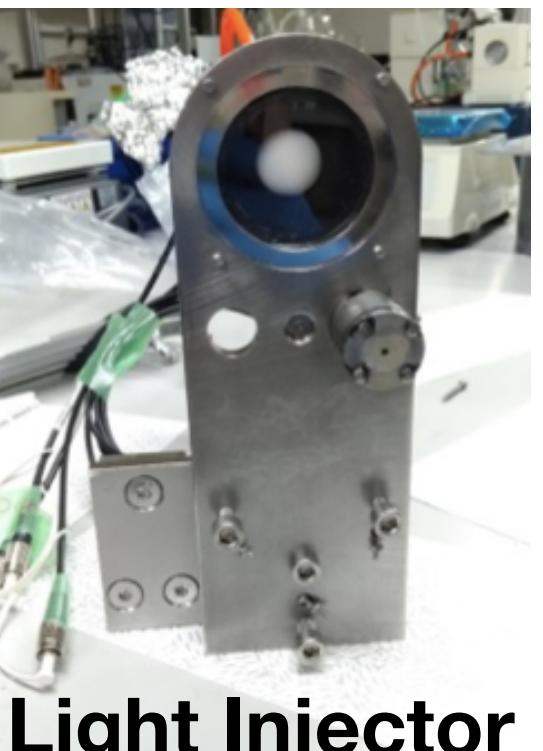
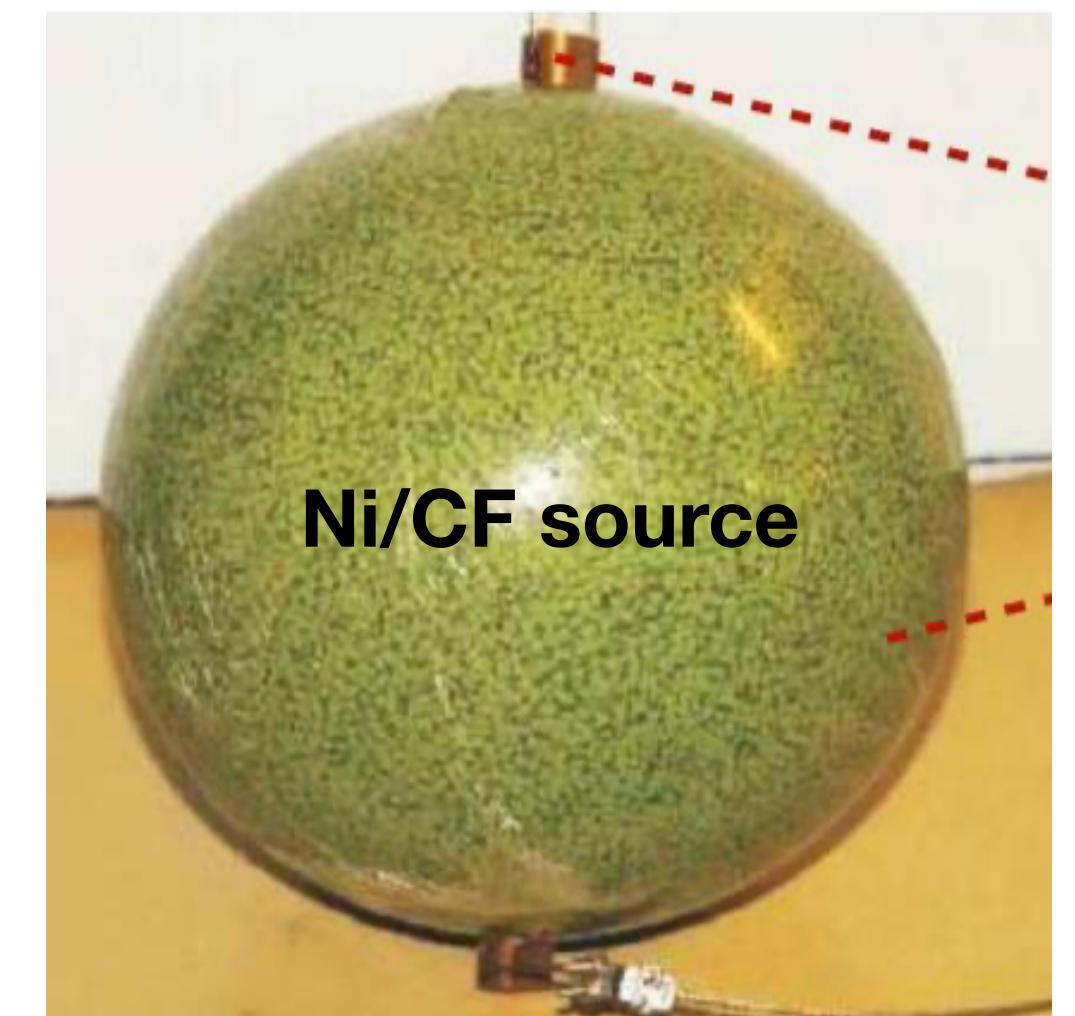
Underwater Electronics

- Underwater electronics to reduce the cable length (signal quality, cost)
- El. vessel includes low-voltage & high-voltage power supply, data processing board, digitizer board
- Two configurations:
 - ID vessel services 24 ID PMTs
 - Hybrid (ID+OD) vessel services 20 ID PMTs and 12 OD modules
- Long-term underwater test conducted at CERN



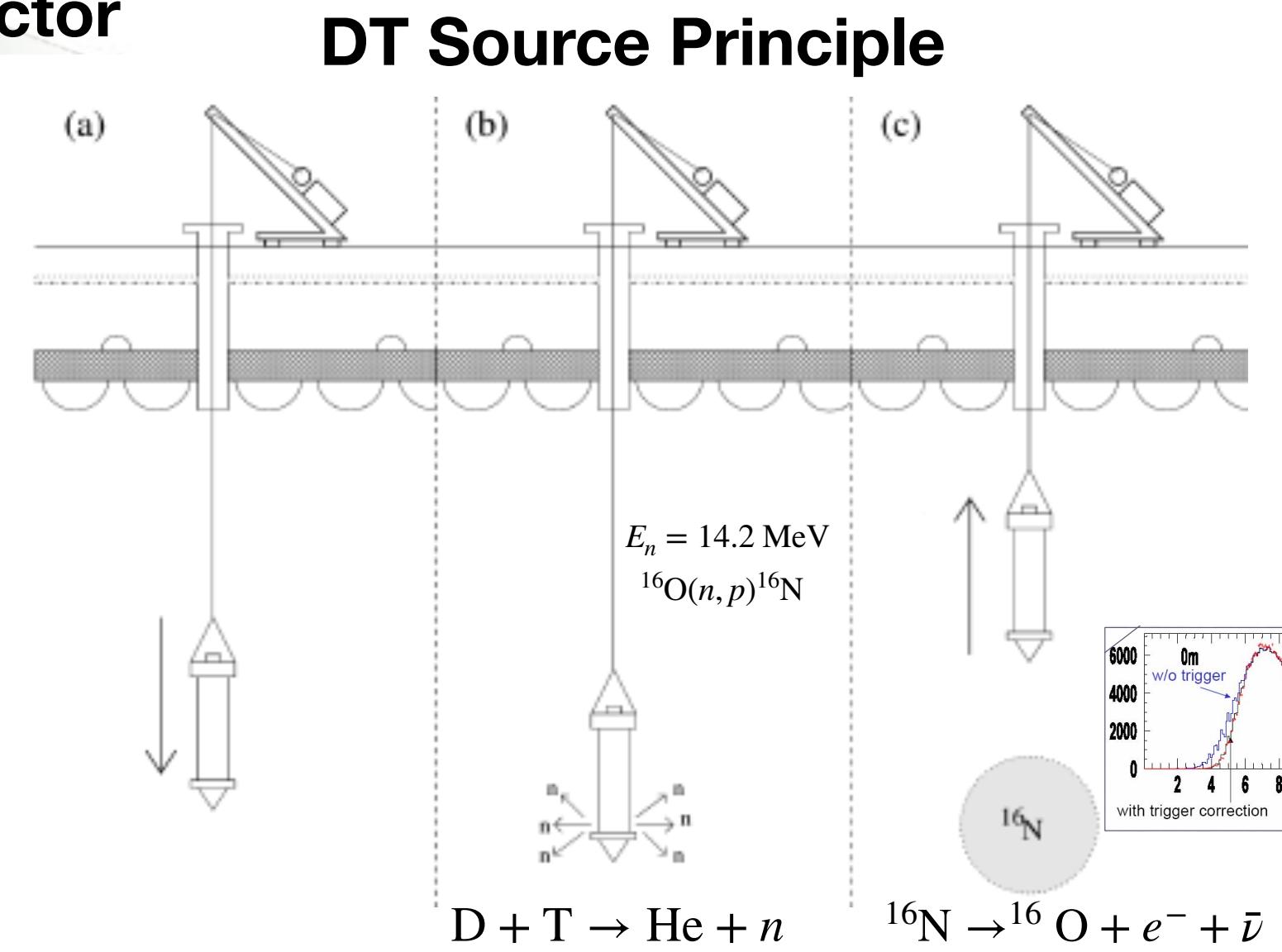
Far Detector Calibration

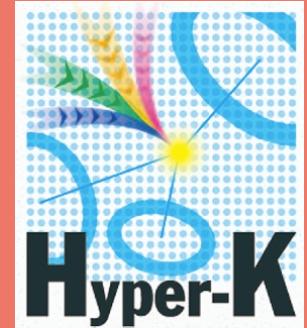
- Great experience from Super-K
- Calibration through optical&radioactive sources and control sample
- Pre-calibration: calibrated PMTs and photogrammetry
- Regular calibration
 - Light injection through diffusers and collimators
 - mPMT flashing LEDs
 - Outer detector light injectors
- Electron LINAC: 3-24 MeV electrons
- Radioactive sources
 - DT source
 - AmBe+BGO - tagged neutrons
 - Ni/Cf 9 MeV γ cascade



Light Injector

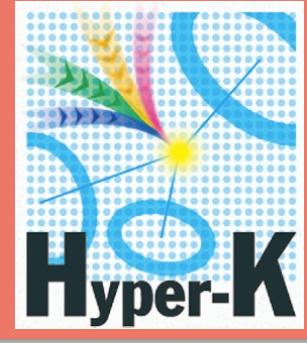
Photogrammetry vessel



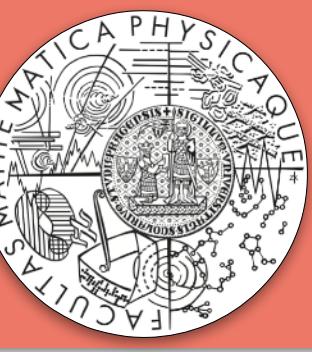


Conclusions

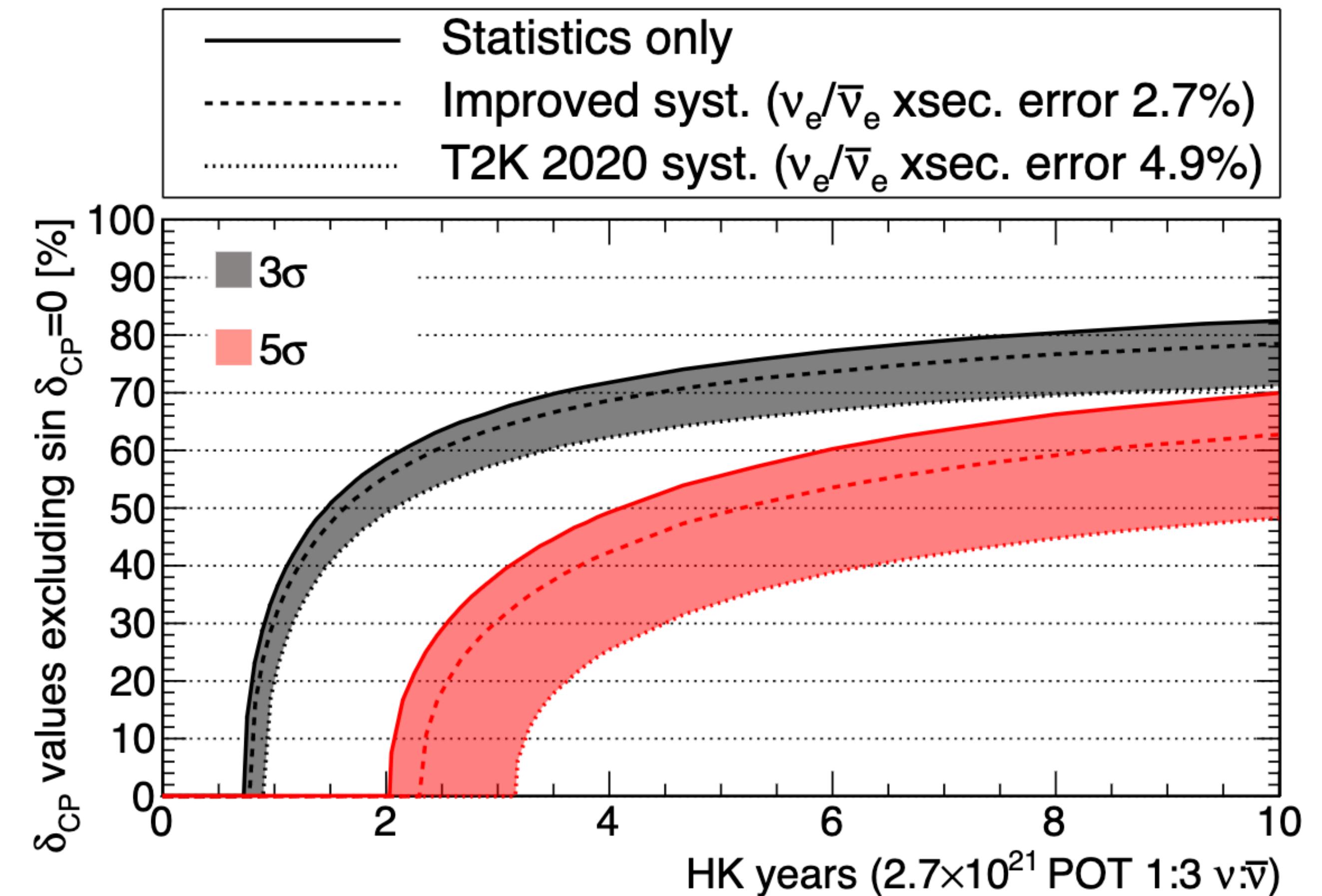
- Hyper-Kamiokande is a 3rd generation water Cherenkov experiment @ Kamioka
 - World's largest accelerator long baseline neutrino oscillation and nucleon decay detector
 - Built on well-proven technology
- Upgrade of the near detectors + introduction of the Intermediate Water Cherenkov Detector to reduce systematics
- Construction on time and on the way to start operation in 2027
- Many results in 20 years of data taking
 - CP violation in neutrino oscillation discovery for 62% of true δ_{CP} @ 5σ
 - Nucleon decay explored up to $\tau > 10^{35}$ y
 - A lot more...



Extras



CP Conservation Exclusion with Time



Hyper-K preliminary

True normal ordering (known)

$$\sin^2 \theta_{13} = 0.0218 \pm 0.0007, \sin^2 \theta_{23} = 0.528, \Delta m_{32}^2 = 2.509 \times 10^{-3} \text{ eV}^2/c^4$$

Atm. + LBL Both Orderings

