





# Photonuclear reactions with charged particles detection for nuclear astrophysics studies

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ISNS 2024, September 9-13, 2024

## Outline

- nuclear astrophysics with γ-ray beams & charged particle detection
- <sup>7</sup>Li(γ,t)<sup>4</sup>He reaction Li-problem or theoretical conundrum
- ${}^{16}O(\gamma,\alpha){}^{12}C$  reaction the Holy Grail
- VEGA System Variable Energy Gamma System



# nuclear astrophysics with **γ**-ray beams

# & charged particles detection





• the principle of detailed balance based on the notion of invariance under time reversal allows the calculation of the cross section  $\sigma_{capture}$  for the capture processes to the ground state by measuring the cross section  $\sigma_{\gamma}$  for the inverse, photo-disintegration reaction:

$$\mathsf{A}(\mathbf{x},\mathbf{\gamma})\mathsf{B} \to \mathsf{B}(\mathbf{\gamma},\mathbf{x})\mathsf{A} \qquad \sigma_{capture} = \frac{2(2j_B + 1)k_{\gamma}^2}{(2j_A + 1)(2j_x + 1)k_x^2}\sigma_{\gamma}$$

 photo-disintegration cross section is enhanced by a factor of 50 or more compared to the direct radiative capture process

 gamma-ray beams do not experience electronic energy loss as they pass through the target material

# why is ${}^{3}H(\alpha,\gamma)^{7}Li$ important?

- actions  ${}^{3}\text{He}(\alpha,\gamma){}^{7}\text{Be}$  and  ${}^{3}\text{H}(\alpha,\gamma){}^{7}\text{Li}$  is
- the study of the mirror alpha capture reactions  ${}^{3}\text{He}(\alpha,\gamma){}^{7}\text{Be}$  and  ${}^{3}\text{H}(\alpha,\gamma){}^{7}\text{Li}$  is needed to evaluate the primordial  ${}^{7}\text{Li}$  abundance in the universe •  ${}^{3}\text{H}(\alpha,\gamma){}^{7}\text{Li}$  could be responsible for  ${}^{7}\text{Li}$  production under different value of the baryonic density

 ${}^{3}H(\alpha,\gamma)^{7}Li$  experiments:

- conflicting results below 1 MeV
- lack of experimental data above 1.2 MeV
- most recent and complete data set by Brune et al in 1994
- measured angular distributions
- overall systematic uncertainty at 6%



from Nollet et al, PRD 61, 123505 (2000)

#### a bit of history on this project





D. Balabanski, 10/07/2014

An update on the big bang nucleosynthesis prediction for <sup>7</sup>Li: the problem worsens

#### Richard H Cyburt<sup>1</sup>, Brian D Fields<sup>2,3</sup> and Keith A Olive<sup>4</sup>

<sup>1</sup> Joint Institute for Nuclear Astrophysics (JINA), National Superconducting Cyclotron Laboratory (NSCL), Michigan State University, East Lansing, MI 48824. USA

<sup>2</sup> Department of Astronomy, University of Illinois, Urbana, IL 61801, USA <sup>3</sup> Department of Physics, University of Illinois, Urbana, IL 61801, USA

#### Primordial Nucleosynthesis

Centre de Spectrométrie Nucléaire et de Spectrométrie de Masse (CSNSM), CNRS/IN2P3, Université Paris Sud, UMR 8609, Bâtiment 104, F-91405 Orsay Campus, France E-mail: Alain.Coc@csnsm.in2p3.fr

Abstract. Primordial nucleosynthesis, or Big Bang Nucleosynthesis (BBN), is one of the three evidences for the lig-Bang model, together with the expansion of the Universe and the Cosmic Microwave Background. There is a good global agreement over a range of mise orders of magnitude between abundances of "He. D, "He and "Li deduced from observations and calculated in primordial nucleosynthesis. This comparison was used to determine the



- the mirror alpha capture reactions  ${}^{3}\text{He}(\alpha,\gamma){}^{7}\text{Be}$  and  ${}^{3}\text{H}(\alpha,\gamma){}^{7}\text{Li}$  are receiving a lot of theoretical attention recently
- theoretical models could provide the the capture cross section at lower energies where experiments are not possible
- Neff et al. calculate the radiative capture cross section in the fully microscopic fermionic molecular dynamics approach using a realistic effective interaction
- good agreement with measurements of  ${}^{3}\text{He}(\alpha,\gamma){}^{7}\text{Be}$
- no agreement with measurements of Brune et al for  ${}^{3}H(\alpha,\gamma)^{7}Li$



#### Laser-Compton Scattering (LCS)



Accelerator Facility Facility/Project: HIGS 160 MeV Linac pre-injector Institution: TUNL 160 MeV-1.2 GeV Booster injector 240 MeV-1.2 GeV Storage ring Country: US FELs: OK-4 (lin), OK-5 (circ) Energy (MeV): 1-100 HIGS: two-bunch, 40-120 mA (typ) Accelerator: Storage Ring, 0.24–1.2 GeV Laser: FEL, 1060 - 190 nm (1.17-6.53 eV) Total flux: 10<sup>7</sup>-3x10<sup>10</sup>g/s(max ~10 MeV) Status: User Program Research: Nuclear physics, Astrophysics, National Security Storage Ring - ALLOND from HIgS website



#### <sup>7</sup>Li experiment @ HIGS

#### **SIDAR** array of YY1 detectors from ORNL





# <sup>7</sup>Li experiments @ HIgS

- <sup>7</sup>Li(g,t)<sup>4</sup>He during March-April 2017
- $E_v = 5 10 \text{ MeV}$
- 70 hours of beam
- LiF target on mylar (300 µg/cm<sup>2</sup>)
- SIDAR array (lamp-shade)
- 12 YY1 detectors (200 ch)



- Iarge collaboration led by ELI-NP
- ORNL, Rutgers, UNC, York U, INFN-LNS, Aarhus U, SKKU, U Michigan



#### during the experiment in March 2017





### <sup>7</sup>Li(γ,t)α w/ SIDAR at HIgS

nuclear physics

- using SIDAR array from ORNL
- $\bullet$  two lamp-shades of YY1: 300, 500, 1000  $\mu m$
- background proportional to YY1 thickness
- clean alpha-triton coincidence above 6 MeV





### <sup>7</sup>Li( $\gamma$ ,t) $\alpha$ w/ SIDAR below 6 MeV



1000&500 µm



Upstream detector (300 um)

Downstream detector (1000 um) 0<sup>L</sup> Upstream detector (500 um)





Unstream detector (300 µm)





Measurement of the 7Li( $\gamma$ ,t)4He ground-state cross section between E $\gamma$  = 4.4

and 10 MeV, M. Munch, C. Matei, S.D. Pain, M.T. Febbraro, K.A. Chipps, H.J. Karwowski, C.Aa. Diget, A. Pappalardo, S. Chesnevskaya, G.L. Guardo, D. Walter, D.L. Balabanski, F.D. Becchetti, C.R. Brune, K.Y. Chae, J. Frost-Schenk, M.J. Kim, M.S. Kwag, M. La Cognata, D. Lattuada, R.G. Pizzone, G.G. Rapisarda, G.V. Turturica, C.A. Ur, and Y. Xu, Phys. Rev. C 101, 055801 (2020)



## <sup>7</sup>Li experiments @ HIgS

- <sup>7</sup>Li(g,t)<sup>4</sup>He during April 2023
- $E_{\gamma} = 3.7 6 \text{ MeV}$
- 70 hours of beam
- . LiF target on mylar (80 µg/cm<sup>2</sup>)
- . SIDAR array (lamp-shade)
- 12 YY1 detectors (200 ch)
- . 100 & 65 µm detectors
- additional steps to lower the beaminduced background:
  - thinner entry flange
  - entry flange farther away
  - longer vacuum pipe before interaction
- large collaboration led by ELI-NP
- ORNL, Rutgers, INFN-LNS, Ohio U, UNC, SHU, SKKU



### $^{7}Li(\gamma,t)\alpha$ w/ SIDAR in 2023





work of PhD student: Ioana Kuncser

#### p-process at HIgS in 2023



- <sup>112</sup>Sn and <sup>102</sup>Pd are accessible test cases to constrain OMP
- OMP variations large effect on p-process abundances
- energy range 11-20 MeV
- P-17-19: D Lattuada (ELI-NP / INFN)
- P-15-19: K.A. Chipps (ORNL)
- carried out at HIgS for 200 hours during March 2023





- . SIDAR array (lamp-shade)
- · 24 YY1 detectors (384 ch)
- 12 dE detectors: 65 & 100 µm
- 12 E detectors: 1000 µm



- Ioana Kuncser, T. Petruse, A. Pappalardo, H. Pai, Yi Xu, D. Balabanski (ELI-NP)
- Steve Pain, Kelly Chipps, T. King, M Febbraro (ORNL)
- M. Grinder, S. Balakrishnan, <u>Heather Garland</u>, Jolie Cizewski (Rutger U)
- H. Karwowski, R.V. Janssens, T. Psaltis, C. Marshall (UNC Chapel Hill)
- <u>G. Restifo</u>, M. La Cognata, G.L. Guardo, S. Palmerini, L. Sergi, D. Lattuada, R.G. Pizzone, G. Rapisarda, A. Tumino (INFN-LNS)
- C.R. Brune, A. Voinov (Ohio U.)
- <u>O. Tindle</u>, <u>C. Haverson</u>, R. Smith (Shefield Hallam U, UK)
- K.Y. Chae, <u>Gyoungmo GU (SungKyunKwan U)</u>



#### more experiment approved at HIgS

- <sup>116</sup>Sn, <sup>120</sup>Sn as next targets for p-process program
- OMP variations large effect on p-process abundances
- energy range 11-20 MeV
- P-10-23: K.A. Chipps (ORNL)
- approved at HIgS by PAC 2023 for 140 hours

- <sup>24</sup>Mg as next target for understanding silicon burning
- ${}^{24}Mg(\gamma,\alpha){}^{20}Ne$  determine reaction rate
- energy range 10-13 MeV
- P-09-24: G.L. Guardo (INFN-LNS
- approved at HIgS by PAC 2024 for 72 hours



#### how we end experiments at HIgS





#### summary



- successful experiments w/ SIDAR @ HIgS
  - 7Li(γ,t)4He analysis underway, new results coming soon
  - analysis underway for ( $\gamma$ ,p) and ( $\gamma$ , $\alpha$ ) on <sup>112</sup>Sn and <sup>102</sup>Pd
  - beam-induced background manageable w/ SSDs
  - beam diagnostics is challenging & can be improved
  - approved by 2023 HIgS PAC for additional 160 hours of beam time for p-process
- ELI-NP is currently the most important research project implemented in Romania. The facility is a user facility open to international users:
  - HPLS operational since 2020
  - VEGA will become operational in the future



#### **Nuclear Photonics Training Center**

Spokesperson: N. Pietralla (TU Darmstadt), C.A. Ur (ELI-NP/UPB)

#### **IGK 2891 Nuclear Photonics**

Establishment Proposal



01.10.2023 to 30.09.2028 Technische Universität Darmstadt Prof. Dr. Dr. h.c. mult. Norbert Pietralla Dr. habil. Eng. Calin Alexandru Ur





- joint project between TU Darmstadt and ELI-NP / University Politechnica Bucharest
  - project approved in September 2023 by DFG & IFA
  - supports 3 cohorts of 20 PhD students between
    2023-2029
- open to international students for October 2024 admission
- topics in laser driven nuclear physics, accelerators, and photonuclear reactions

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# **Overview of VEGA System**

https://www.eli-np.ro/rd2.php





#### **ELI-NP overview & layout**





#### **VEGA System - overview**





#### **VEGA status – December 2022**

![](_page_25_Picture_1.jpeg)

![](_page_25_Picture_2.jpeg)

installed water cooling infrastructure in the basement

cable trays in E-P-10

installed waveguide supports in E-P-10

![](_page_25_Picture_6.jpeg)

![](_page_25_Picture_7.jpeg)

crated components

![](_page_25_Picture_9.jpeg)

#### VEGA status – July 2024

![](_page_26_Picture_1.jpeg)

all power & control for magnets / ion pumps / gate valves installed

ready for connecting accelerator structures to water cooling system

![](_page_26_Picture_4.jpeg)

Accelerator Bay 1

#### summary

![](_page_27_Picture_1.jpeg)

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![](_page_27_Picture_11.jpeg)

![](_page_28_Picture_0.jpeg)

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#### **JOIN US**

ELI-NP is building a team of dedicated, talented people willing to contribute to the operation and to perform laser-matter interaction experiments with the most powerful LASER in the world. If you are interested in working in an innovative, dynamic environment and share our passion for "pushing the limits", we would be thrilled to work with you.

ELI-NP offers researchers and engineers from various fields related to the activity of the new Research Center the opportunity to grow and reach their full potential in a multicultural and interdisciplinary environment. It aims to create an innovation a creativity lab for research and applications which benefit society at large. Our work ethic is based on respect, diversity, curiosity and integrity. We offer outstanding working conditions and competitive salaries and benefits.

We are very proud of our project and the team we have brought together. Our employees are our most valuable asset and the key means by which we can achieve our goals. For queries and further information: human.resources@eli-np.ro

#### ELI-NP Open positions

![](_page_29_Figure_6.jpeg)