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Dimensionally reduced EFTs for cosmological phase transitions

Higgs and Effective Field Theory (HEFT) 2024 · University of Bologna

Luis Gil [he/him]

Based on [2406.02667], by:

M. Chala, J. C. Criado, LG and J. López Miras

FTAE
High Energy Theory



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de Andalucía



Bubbles and gravitational waves

Thermally induced field phase transitions

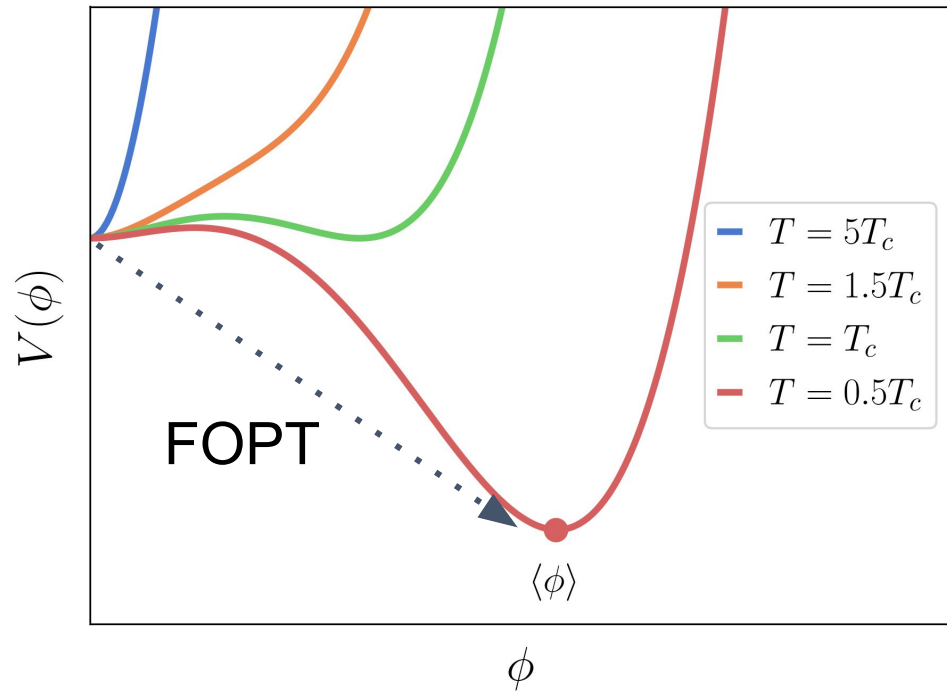


Fig. 1: Temperature evolution of scalar potential

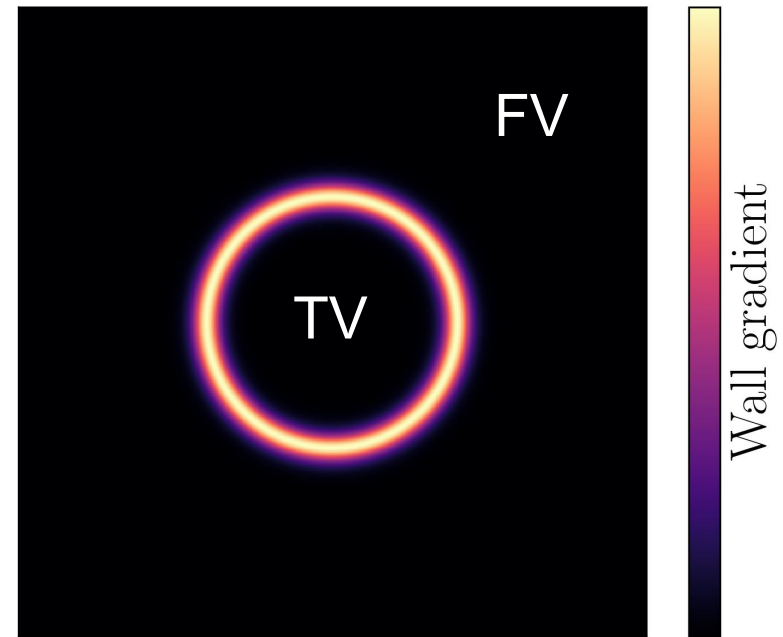
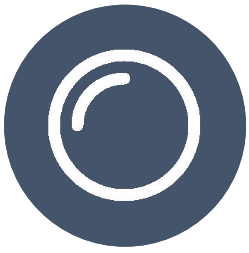


Fig. 2: $O(3)$ symmetric bounce solution



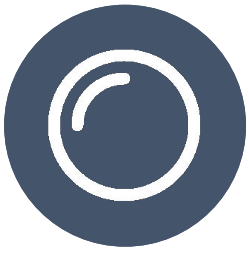
Bubbles and gravitational waves

From PTs to GWs

[Caprini *et al.* - 1512.06239]

As the transition rate grows, in a **first-order PT (FOPT)**:

- 1) Bubbles of true vacuum nucleate and expand in a hot plasma
- 2) Bubble fronts collide
- 3) Sound waves
- 4) Turbulence



Bubbles and gravitational waves

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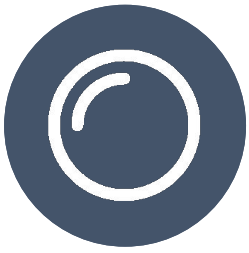
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$$h\Omega_{\text{GW}} \simeq h\Omega_{\phi} + h\Omega_{\text{sw}} + h\Omega_{\text{turb}}$$

Out of equilibrium



Bubbles and gravitational waves

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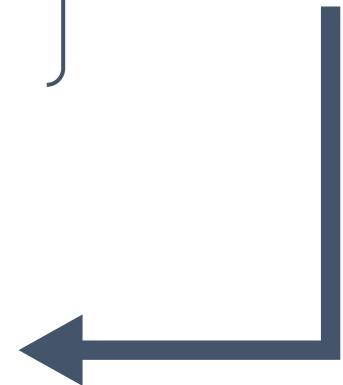
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Probe

Out of equilibrium

Answers for electroweak baryogenesis





Bubbles and gravitational waves

From PTs to GWs

How do we connect a QFT model to these GW spectra?



Bubbles and gravitational waves

From PTs to GWs

How do we connect a QFT model to these GW spectra?

Nucleation temperature

$$S_3[\varphi_c] \sim 100 - 4 \log \frac{T_*}{100 \text{ GeV}}$$

Strength parameter

$$\alpha \approx -0.03 \frac{V_3(\varphi_T)}{T_*^3}$$

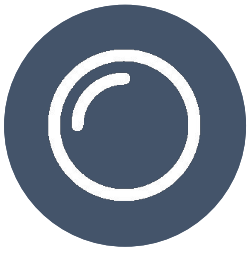
Inverse duration of PT

$$\frac{\beta}{H_*} = T_* \left. \frac{dS_3[\varphi_c]}{dT} \right|_{T_*}$$

Terminal bubble wall velocity

A whole different story...

[Laurent, Cline - 2204.13120]



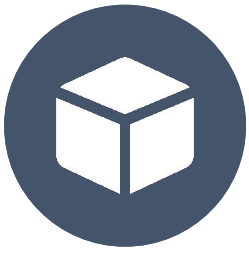
Bubbles and gravitational waves

What is the state of the art?

**This list is NOT
exhaustive!**

- Several studies of FOPTs in BSM extensions and in the SMEFT:
 - N⁴LO Higgs potential at finite T [Ekstedt *et al.* - 2405.18349]
 - EWPT in dim-6 SMEFT [Camargo-Molina *et al.* - 2103.14022]
 - EWPT in dim-6 XSM [Oikonomou *et al.* - 2403.01591]
 - EWPT in Σ SM [Niemi *et al.* - 1802.10500]
- Tools for PT-related computations:
 - *CosmoTransitions* [Wainwright - 1109.4189]
 - *FindBounce* [Guada *et al.* - 2002.00881]
 - *DRalgo* [Ekstedt *et al.* - 2205.08815]
 - *BubbleDet* [Ekstedt *et al.* - 2308.15652]
- Theoretical uncertainties persist from the particle physics side:
 - LISA [Caprini *et al.* - 1910.13125]

Effective operators with
derivatives **not included**



3D EFT approach

Building our high-T EFT

See earlier talk by
M. Chala

UV theory in 4D Minkowski:

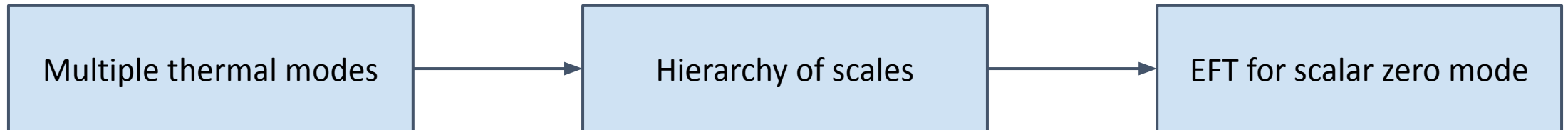
$$\mathcal{L}_4 = \frac{1}{2}(\partial_\mu\phi)^2 - \frac{1}{2}m^2\phi^2 - \kappa\phi^3 - \lambda\phi^4 + \bar{\Psi}i\not{\partial}\Psi - g\phi\bar{\Psi}\Psi \quad [\text{Gould, Xie - 2310.02308}]$$

At temperature T , in Euclidean space:

$$S_0 = (-i)\frac{1}{T} \int d^3\mathbf{x} \sum_{n=-\infty}^{\infty} \left[\frac{1}{2}(\partial_i\phi_n)^2 + \frac{1}{2} \left(m^2 + (2\pi nT)^2 \right) \phi_n^2 + \bar{\psi}_n \not{\partial} \psi_n + \left(2\pi \left(n + \frac{1}{2} \right) T \right)^2 \bar{\psi}_n \psi_n \right]$$

$$T^2 \gg m^2$$

Power counting





3D EFT approach

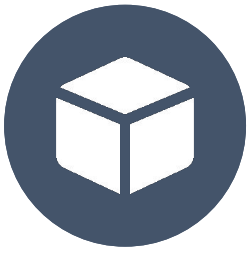
Building our high-T EFT

[Kajantie *et al.* - 9508379]
[Criado - 1901.03501]

We find an off-shell basis of operators up to (4D) dim-8 with the help of *BasisGen* and match at 1-loop (fermion only loops)¹:

$$\begin{aligned}\mathcal{L}_3 = & \frac{1}{2}K_3(\partial\varphi)^2 + \frac{1}{2}m_3^2\varphi^2 + \kappa_3\varphi^3 + \lambda_3\varphi^4 \\ & + \alpha_{61}\varphi^6 + \beta_{61}\partial^2\varphi\partial^2\varphi + \beta_{62}\varphi^3\partial^2\varphi \\ & + \alpha_{81}\varphi^8 + \alpha_{82}\varphi^2\partial_\mu\partial_\nu\varphi\partial^\mu\partial^\nu\varphi + \beta_{81}\varphi\partial^6\varphi + \beta_{82}\varphi^3\partial^4\varphi + \beta_{83}\varphi^2\partial^2\varphi\partial^2\varphi + \beta_{84}\varphi^5\partial^2\varphi\end{aligned}$$

¹ Scalar loop contributions to effective operators are subleading (see Appendix A in [2406.02667]).



3D EFT approach

Building our high-T EFT

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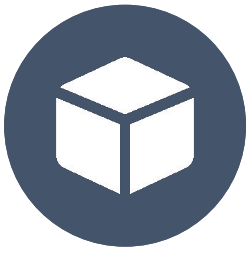
dim-6

Measure eff. ops. relevance

dim-8

Control EFT validity

¹ Scalar loop contributions to effective operators are subleading (see Appendix A in [2406.02667]).



3D EFT approach

Computing relevant PT magnitudes

Eff. ops. depend solely on the Yukawa (g)



Fix all other UV parameters in scan

[Chala *et al.* - 1905.00911]

&

Study all PT magnitudes as functions of the Yukawa

m^2, κ, λ

$T_*, \alpha, \beta/H_*$

See earlier talk by M. Chala

Bounce solution must be computed perturbatively!



Results

PT magnitudes and effective interactions

[Chala, Criado, LG, López Miras - 2406.02667]

■ $(m^2, \kappa, \lambda)_A = (20\,000 \text{ GeV}^2, -40 \text{ GeV}, 0.01)$

■ $(m^2, \kappa, \lambda)_B = (31\,643.5 \text{ GeV}^2, -71.1 \text{ GeV}, 0.045)$

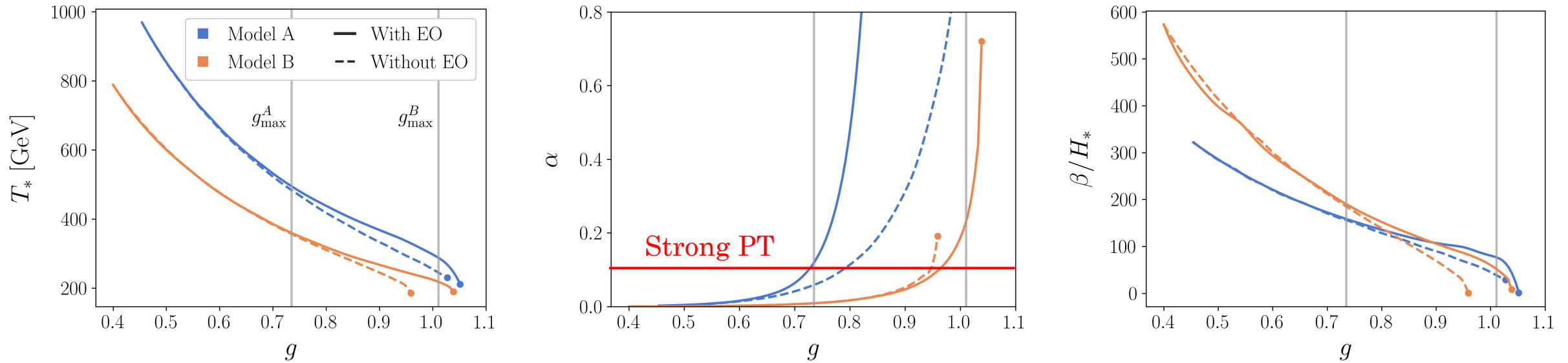


Fig. 5: PT magnitudes in two models, with and w/o effective operators



Results

[Chala, Criado, LG, López Miras - 2406.02667]

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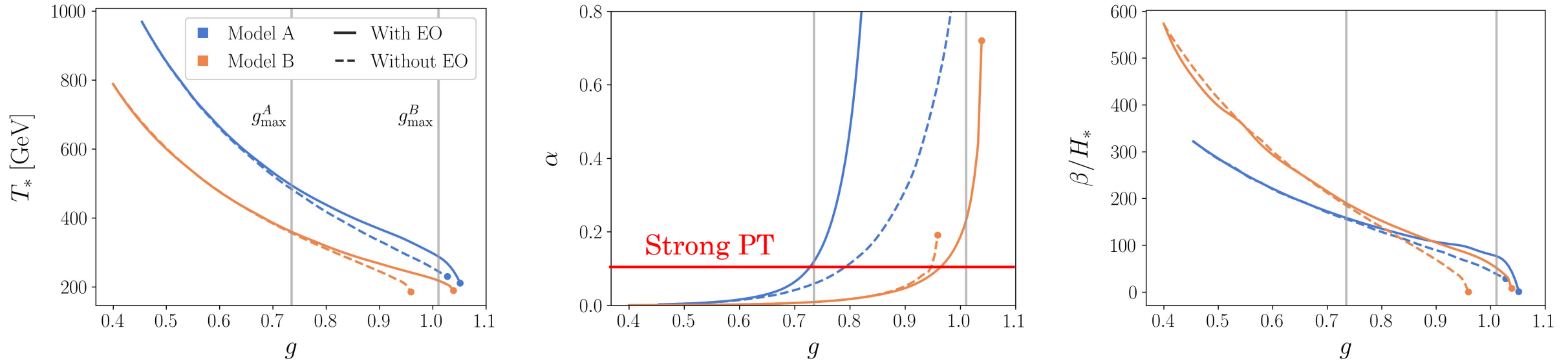


Fig. 5: PT magnitudes in two models, with and w/o effective operators

Two observations

Eff. ops. **allow for PTs in a wider range** of values of the Yukawa

Including eff. ops. yields **very different estimations** at large Yukawas



Results

GW power spectra and effective interactions

[Chala, Criado, LG, López Miras - 2406.02667]

$$(m^2, \kappa, \lambda)_A = (20\,000 \text{ GeV}^2, -40 \text{ GeV}, 0.01)$$

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$$g \lesssim g_{\text{max}}$$

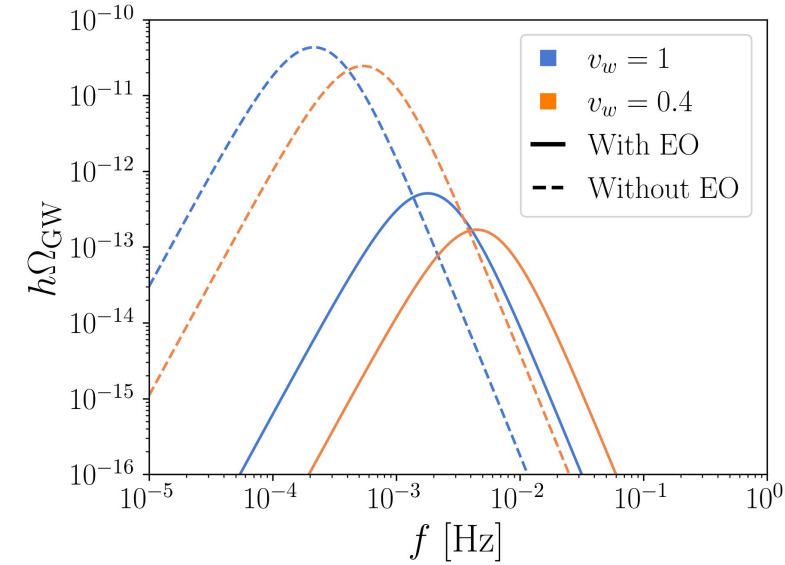
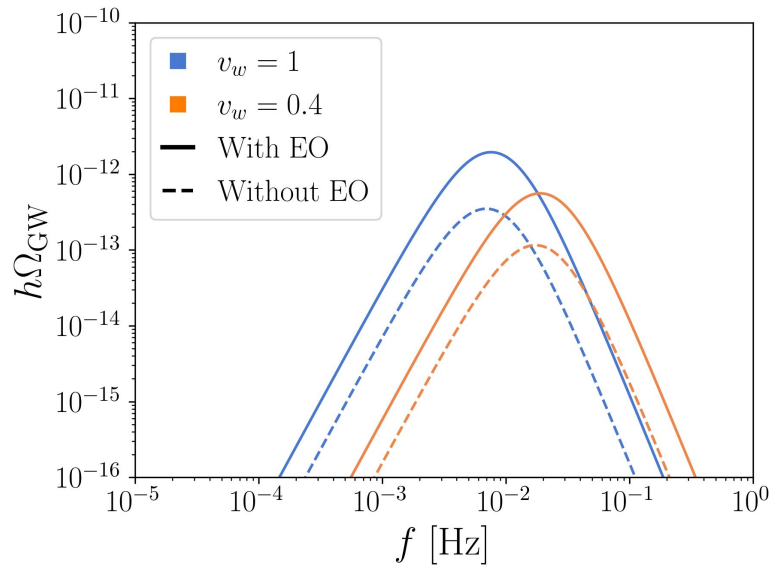


Fig. 6: GW power spectra in two models, with and w/o effective operators

(*) Generated with **PTPlot**
[Caprini *et al.* - 1910.13125]



Results

[Chala, Criado, LG, López Miras - 2406.02667]

GW power spectra and effective interactions

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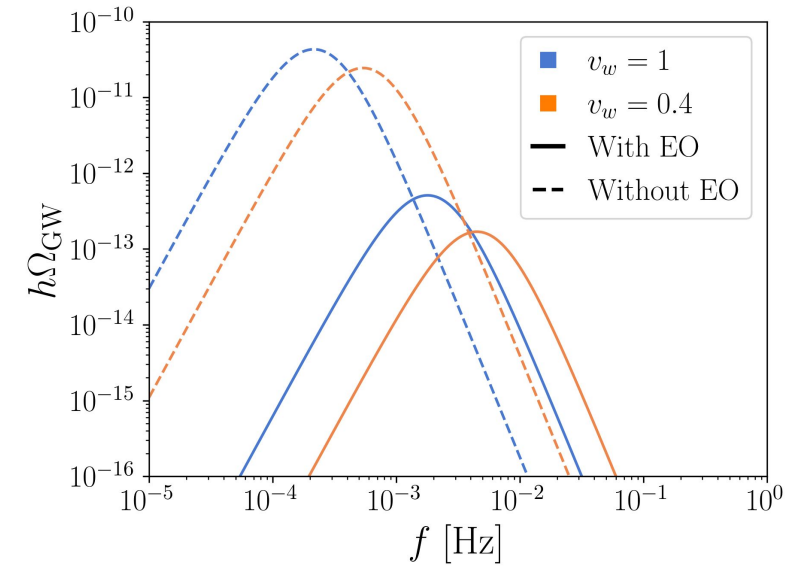
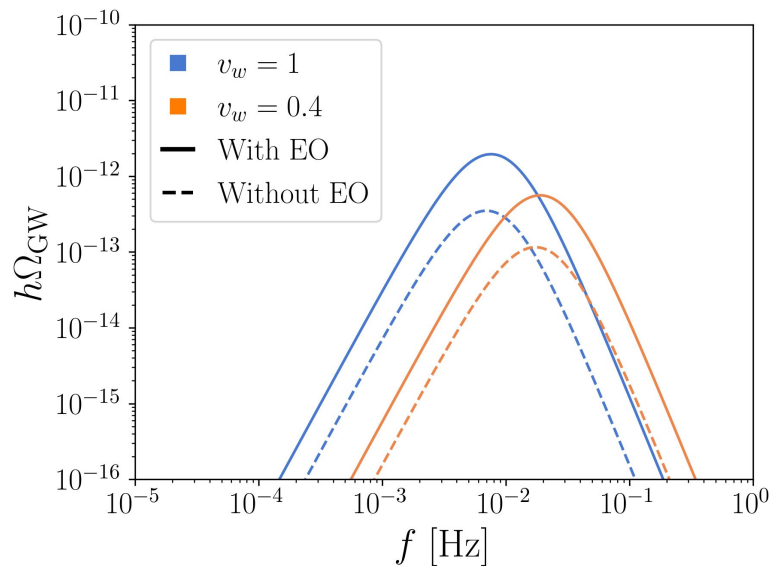


Fig. 6: GW power spectra in two models, with and w/o effective operators

Including eff. ops. can change the peak amplitude and frequency by **one order of magnitude**



Take home messages

We have learned

- Strong FOPTs occur in regions of parameter space **close the limit of validity** of the 3D EFT
- Higher dimensional effective operators **are thus relevant in the estimation of PT-related magnitudes** (order of magnitude differences in GW spectra)
- **Bounce solutions** can be obtained perturbatively in the presence of higher derivative terms

Future work

- Application to well-motivated BSM models
- Adding loop corrections to scalar zero mode

Thank you for your attention!

¡Gracias por vuestra atención!