

Di-Pions in Ultra-Peripheral Collisions at LHCb



Amanda Donohoe
LHCb Collaboration
Hard Probes
27/09/2024

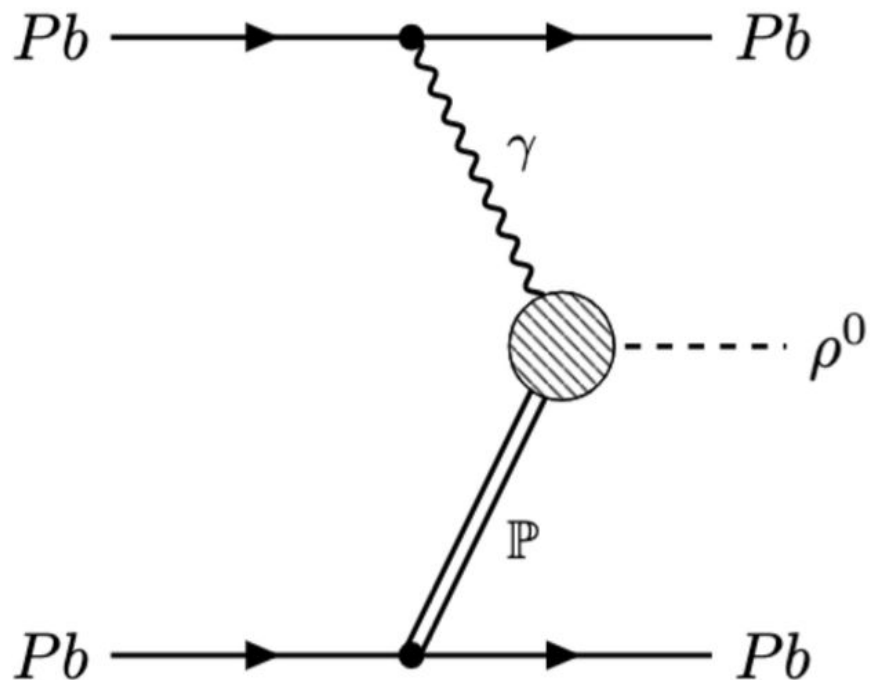


Outline

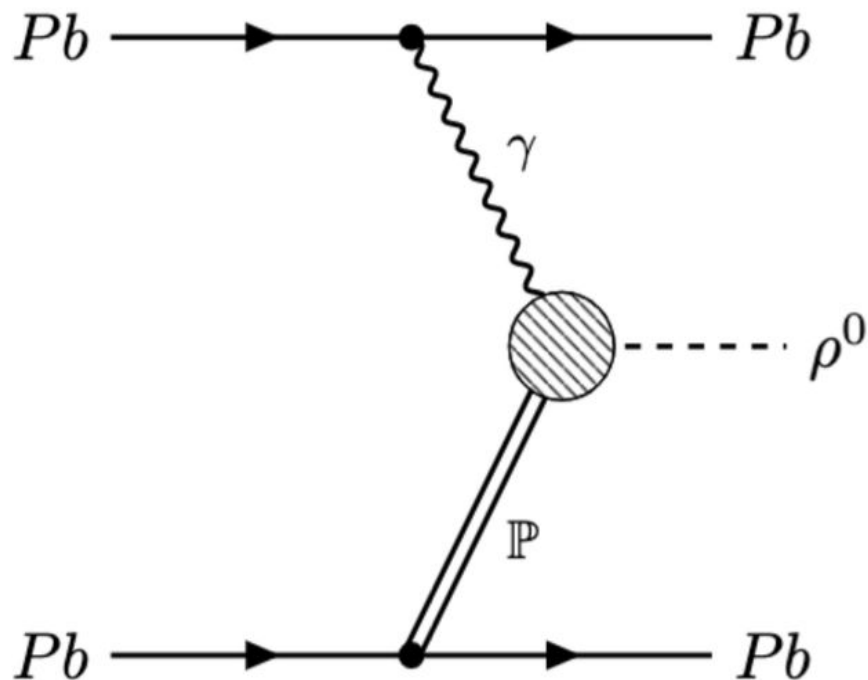
- Introduction
- Motivation
- Structure in the di-pion invariant mass spectrum
- Conclusion

Introduction

Process of Interest

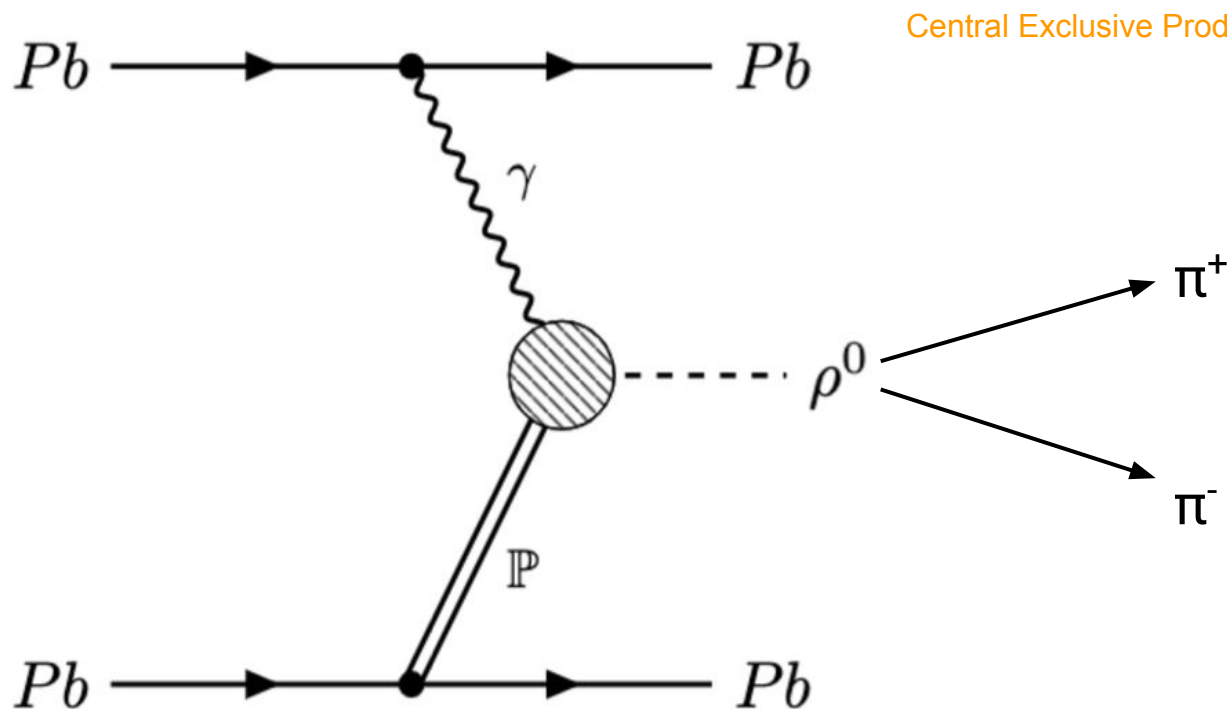


Process of Interest

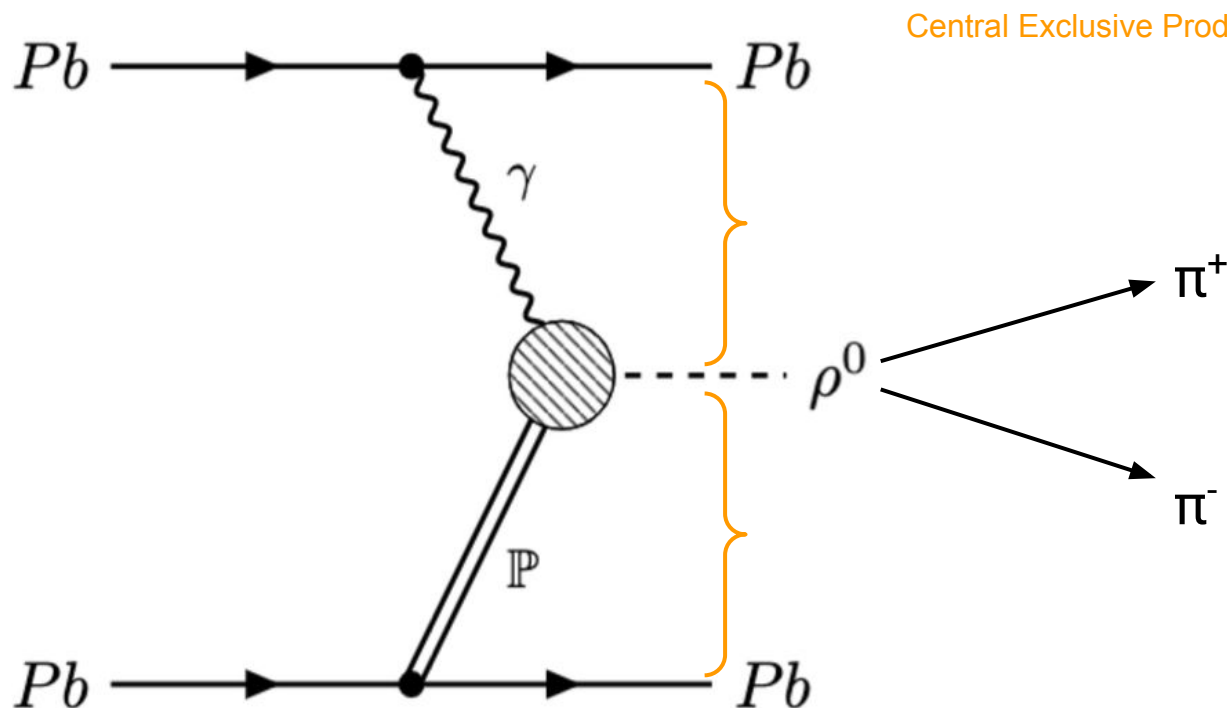


Central Exclusive Production (CEP)

Process of Interest

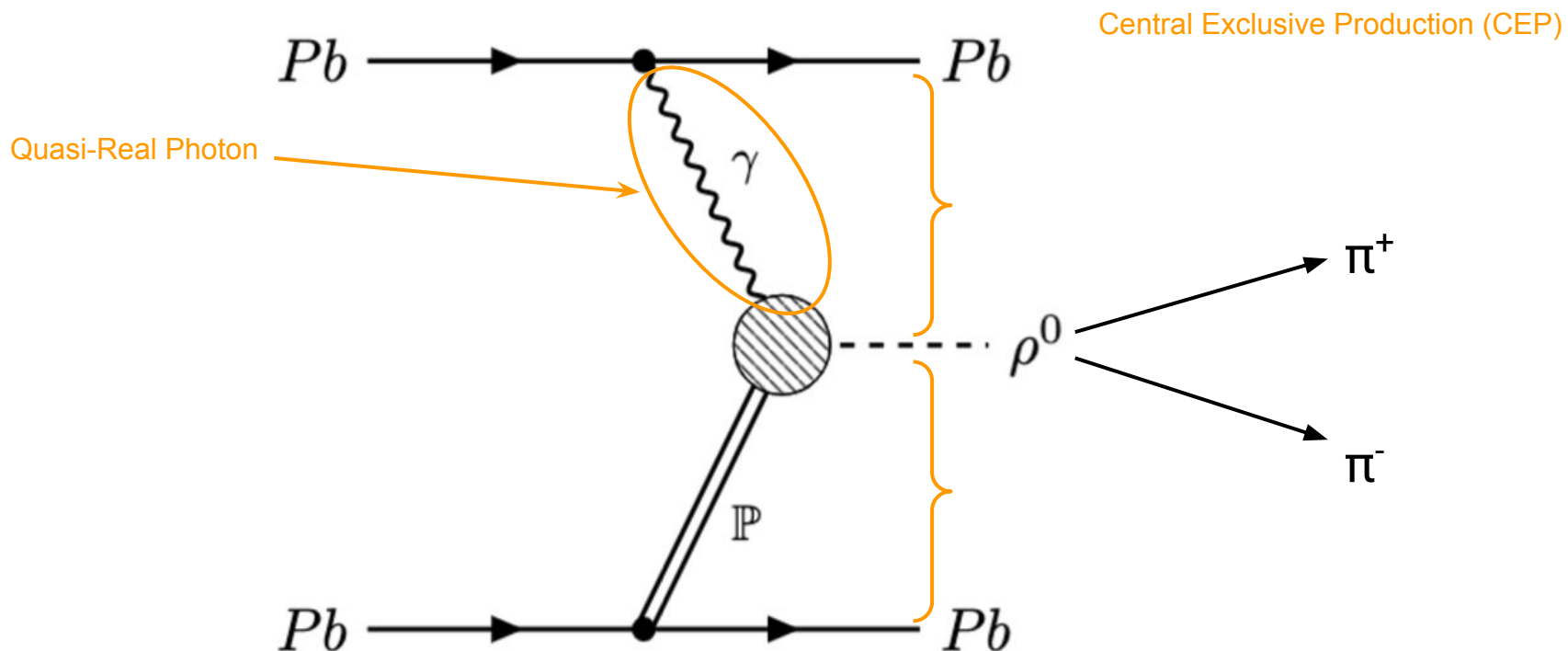


Process of Interest



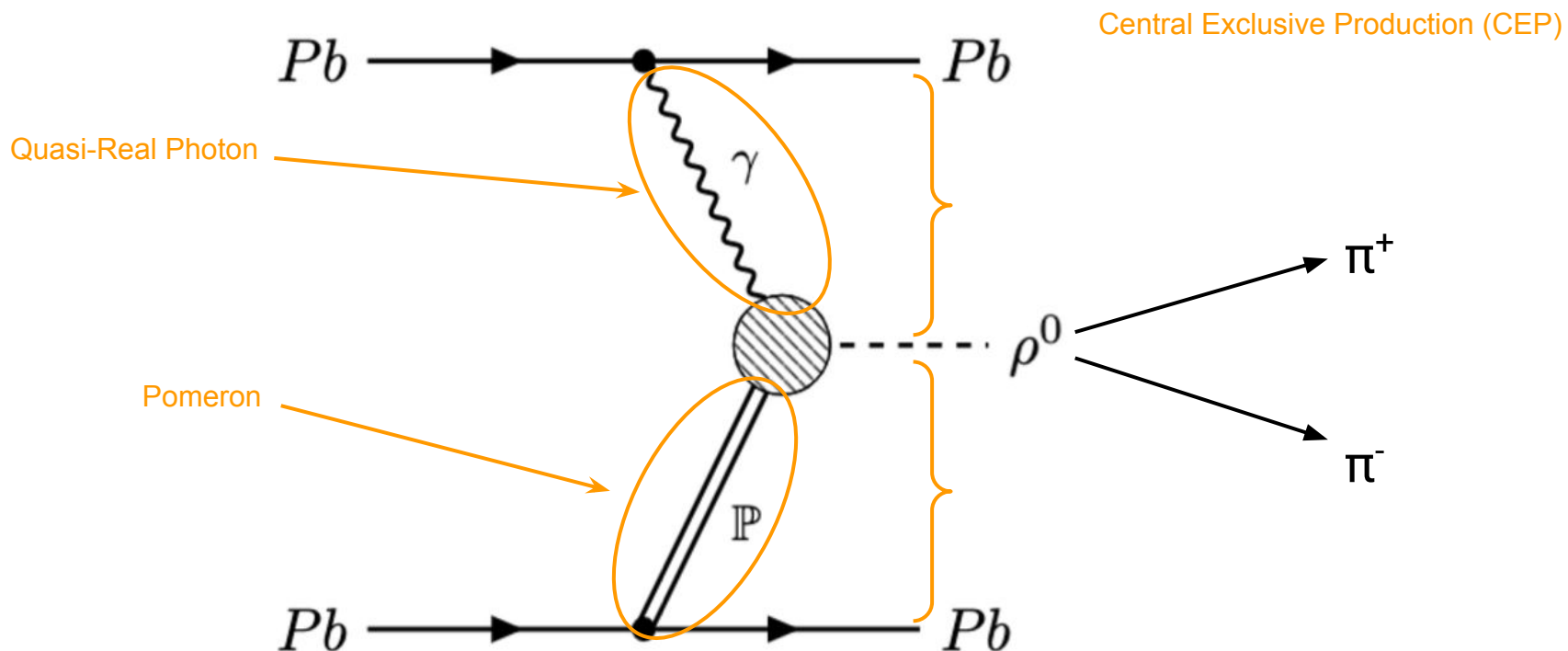
} = Rapidity Gap

Process of Interest



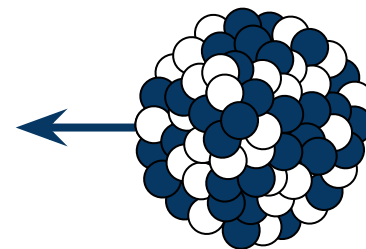
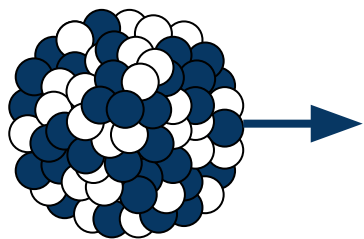
} = Rapidity Gap

Process of Interest

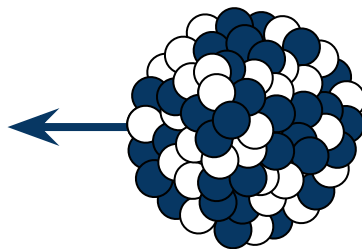
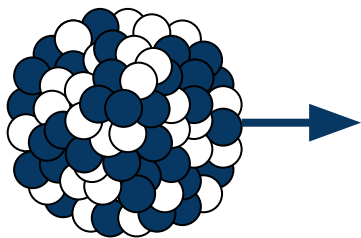


} = Rapidity Gap

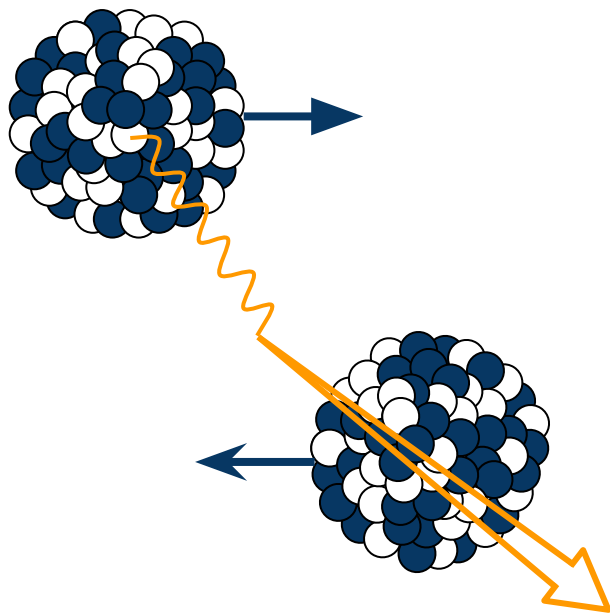
Ultra-Peripheral Collisions



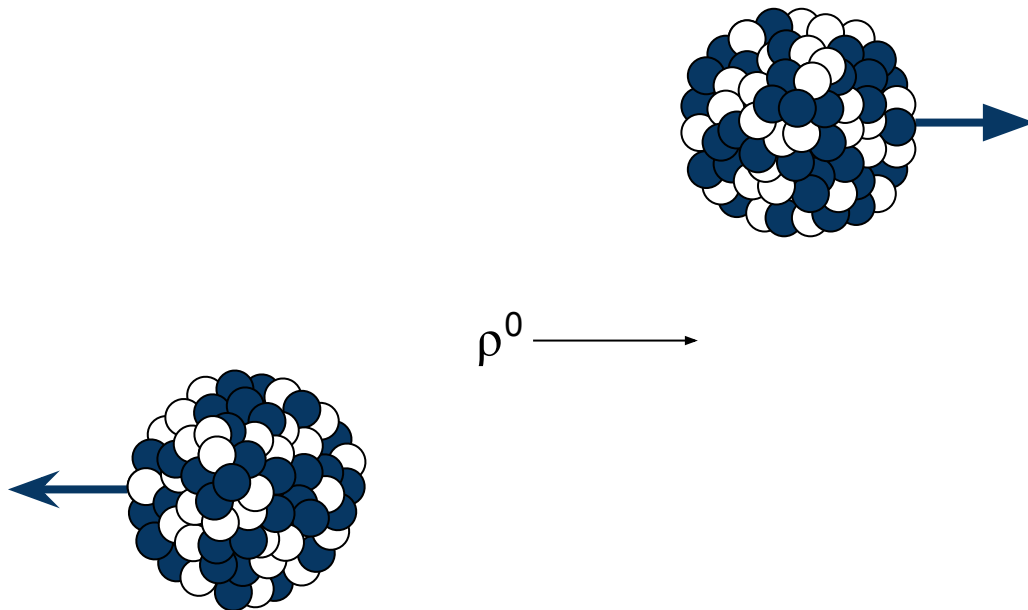
Ultra-Peripheral Collisions



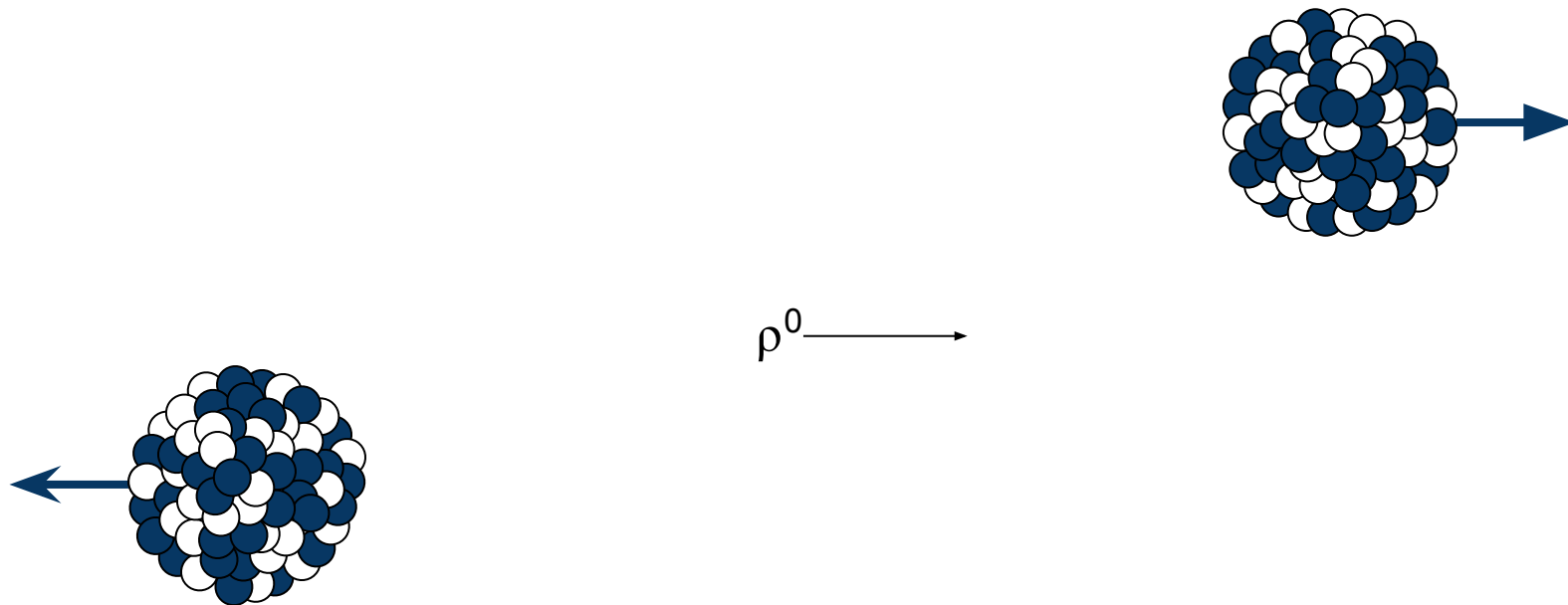
Ultra-Peripheral Collisions



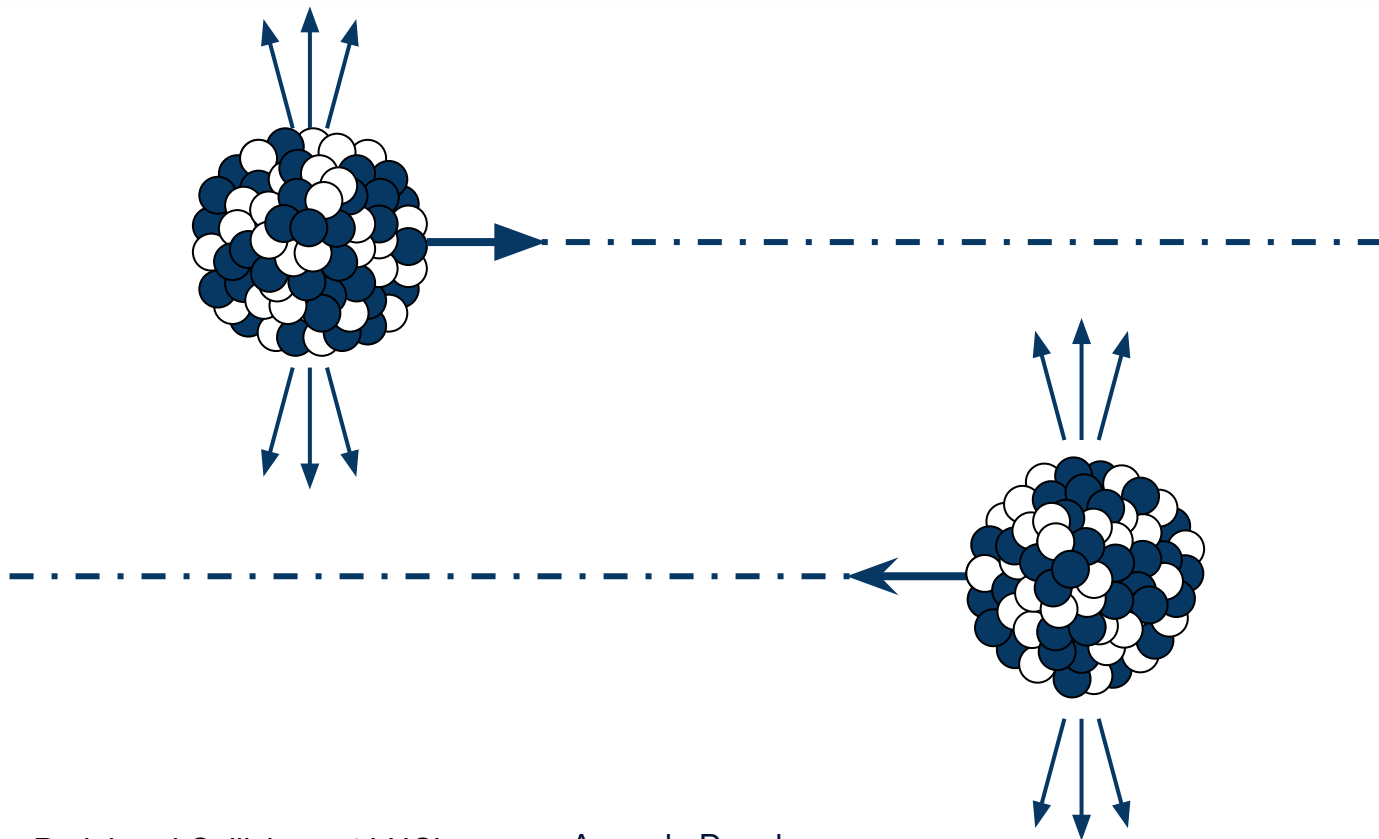
Ultra-Peripheral Collisions



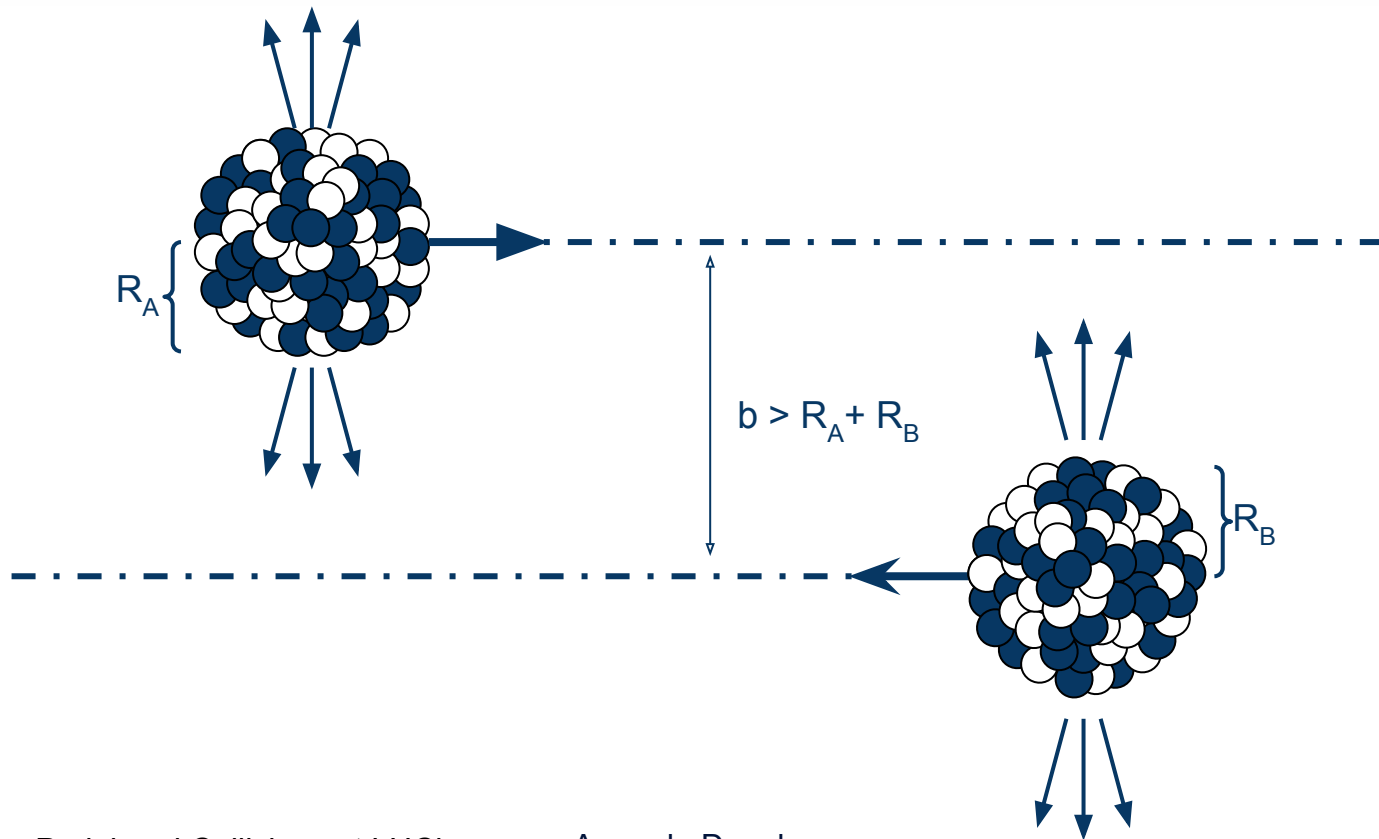
Ultra-Peripheral Collisions



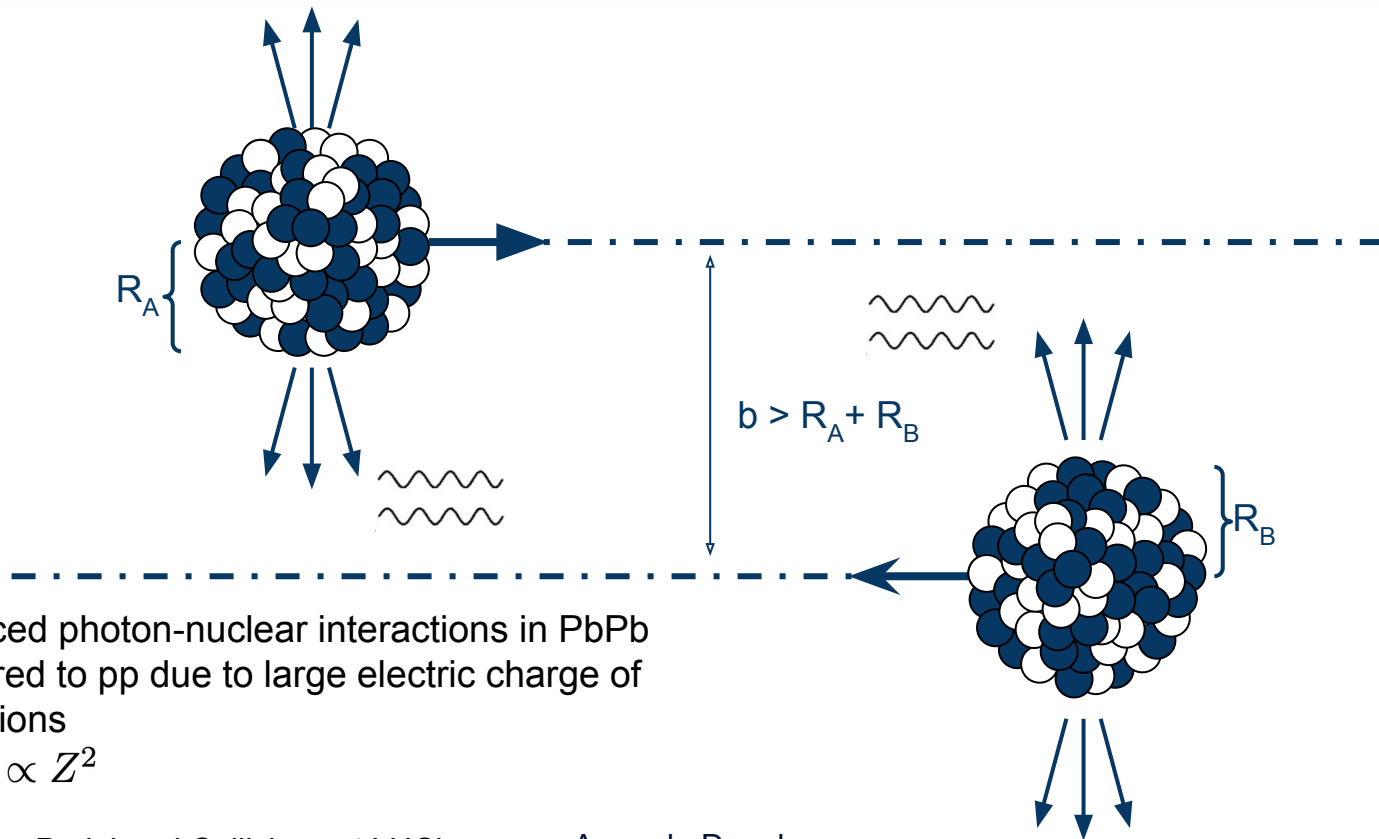
Ultra-Peripheral Collisions



Ultra-Peripheral Collisions



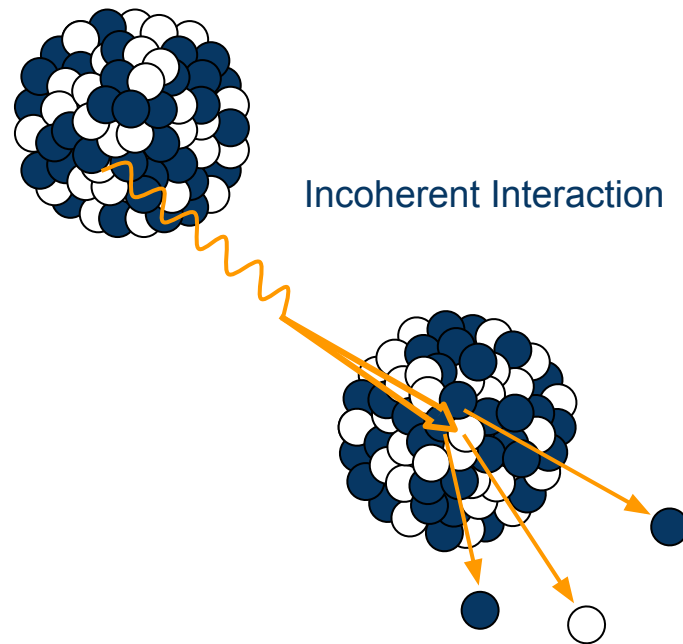
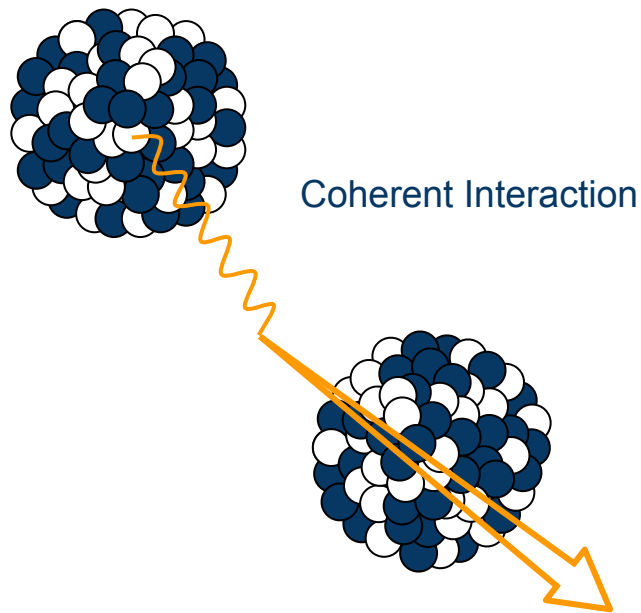
Ultra-Peripheral Collisions



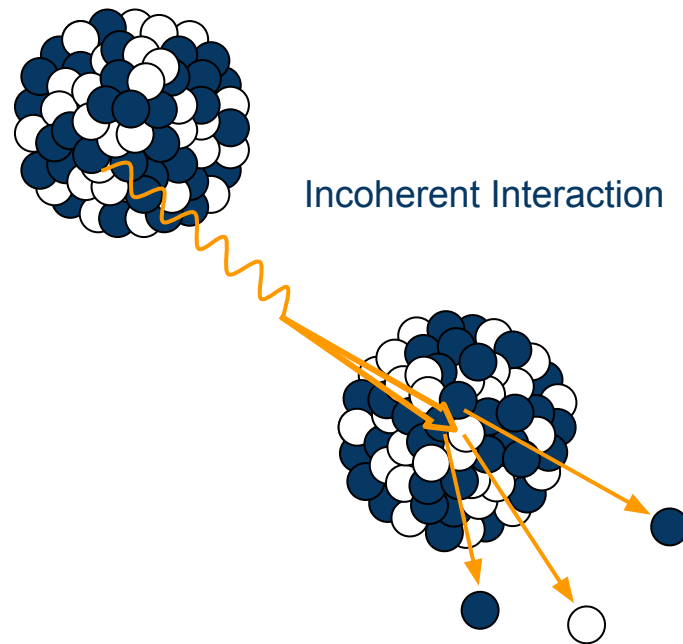
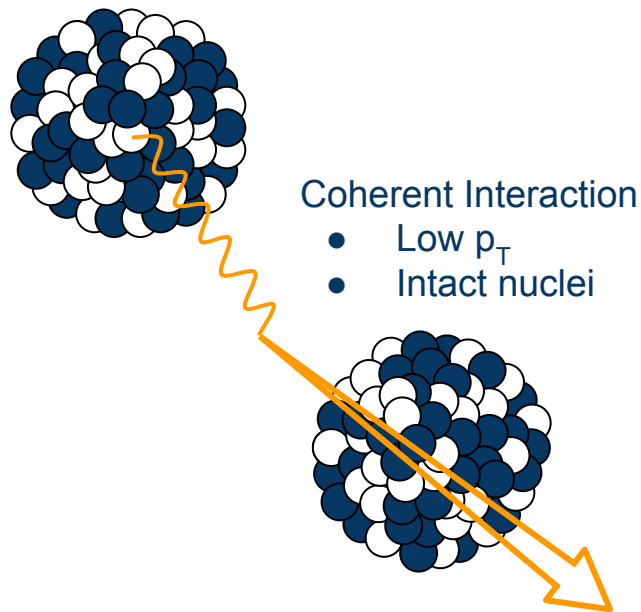
- Enhanced photon-nuclear interactions in PbPb compared to pp due to large electric charge of the Pb ions

$$\circ N_\gamma \propto Z^2$$

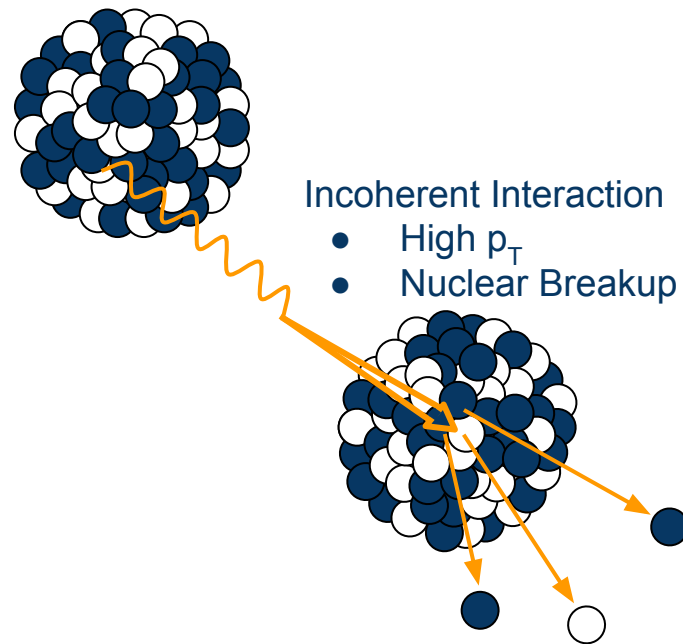
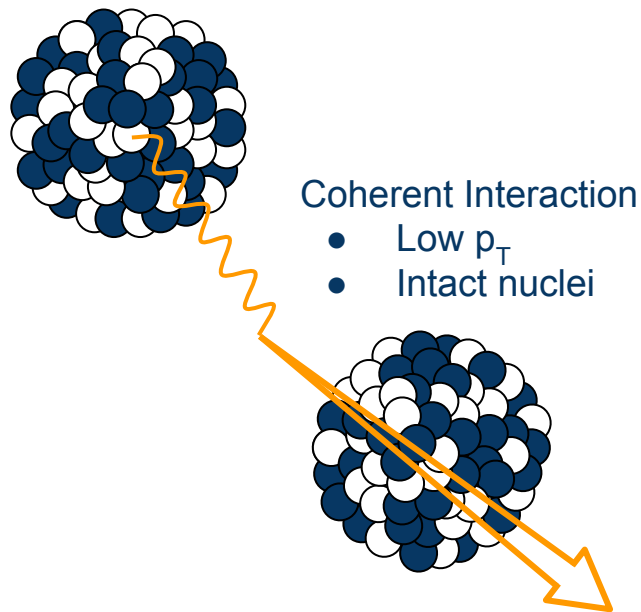
Coherent vs Incoherent Interactions



Coherent vs Incoherent Interactions



Coherent vs Incoherent Interactions



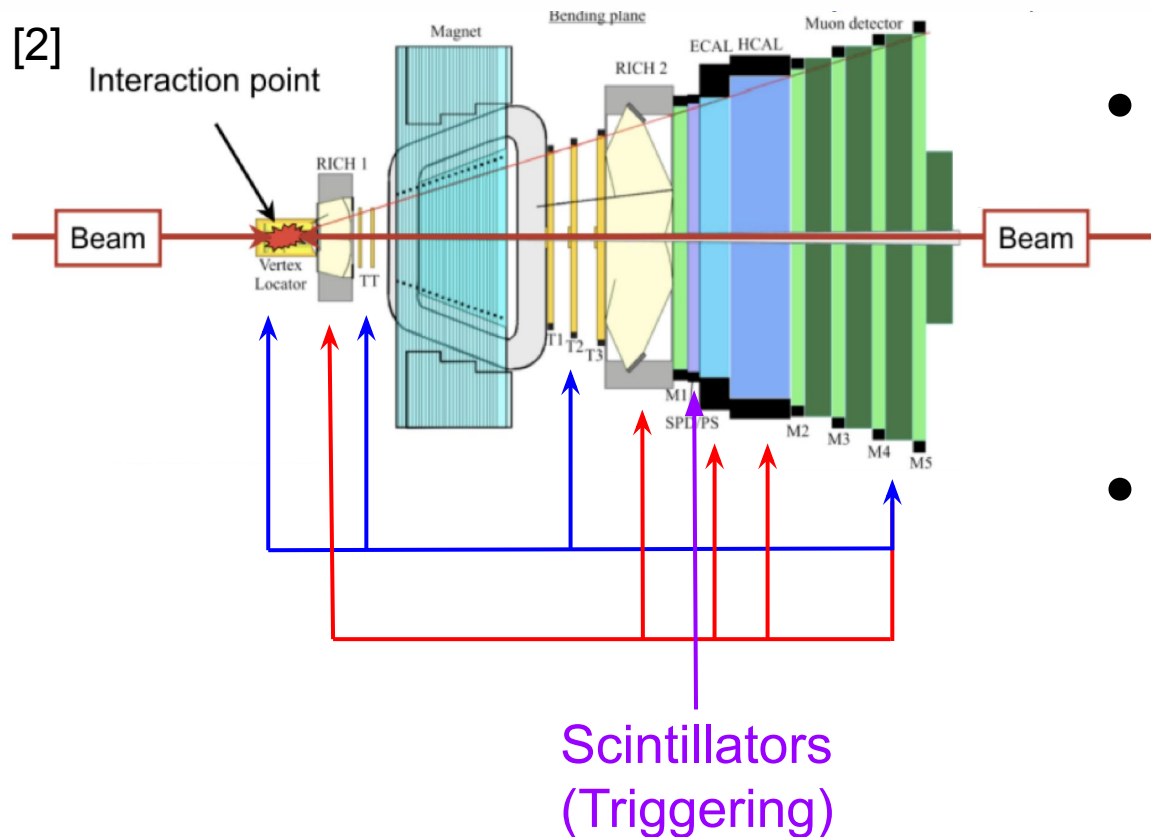
Motivation

Motivation

- Understanding the di-pion spectrum
- Measured ρ parameters appear process dependent (tau decay, e^+e^- , photoproduction)
- low-mass spectroscopy: continuum, ω , high excitations, interference.

Why LHCb?

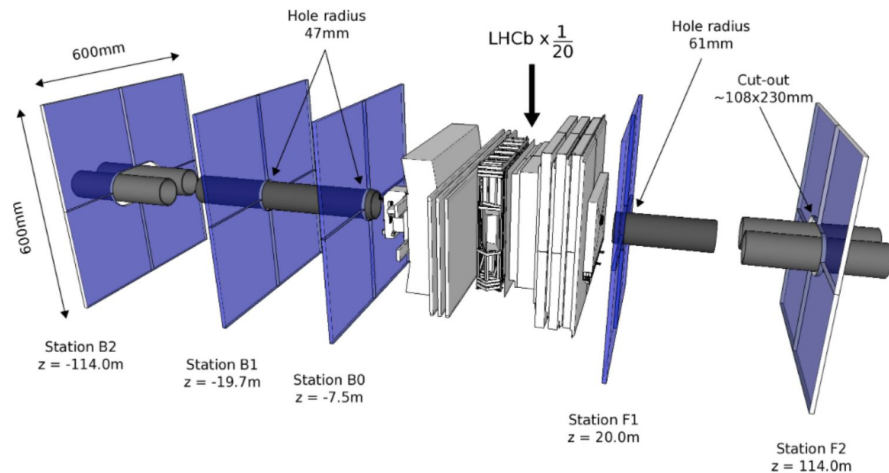
[2]



- A forward-arm spectrometer ($2 < \eta < 5$)
 - Unique acceptance and capabilities.
 - Constraints nPDFs down to $x \sim 10^{-6}$
- Precise Tracking and Full PID.
 - Mass resolution ($m_{\pi\pi\pi}$) ~ 5 MeV.
 - Low background levels

HeRSChEL Sub-Detector

[3]



- HeRSChEL
 - planes of scintillators up and downstream
 - extends coverage to $5 < |\eta| < 10$
- Enables distinction between coherent vs incoherent events due to ion dissociation

Di-Pion Lineshape

Data Selection

Data Used

- 2018 PbPb LHCb data
- $\sqrt{s_{NN}} = 5.02$ TeV
- Integrated luminosity: $228 \pm 10 \mu\text{b}^{-1}$

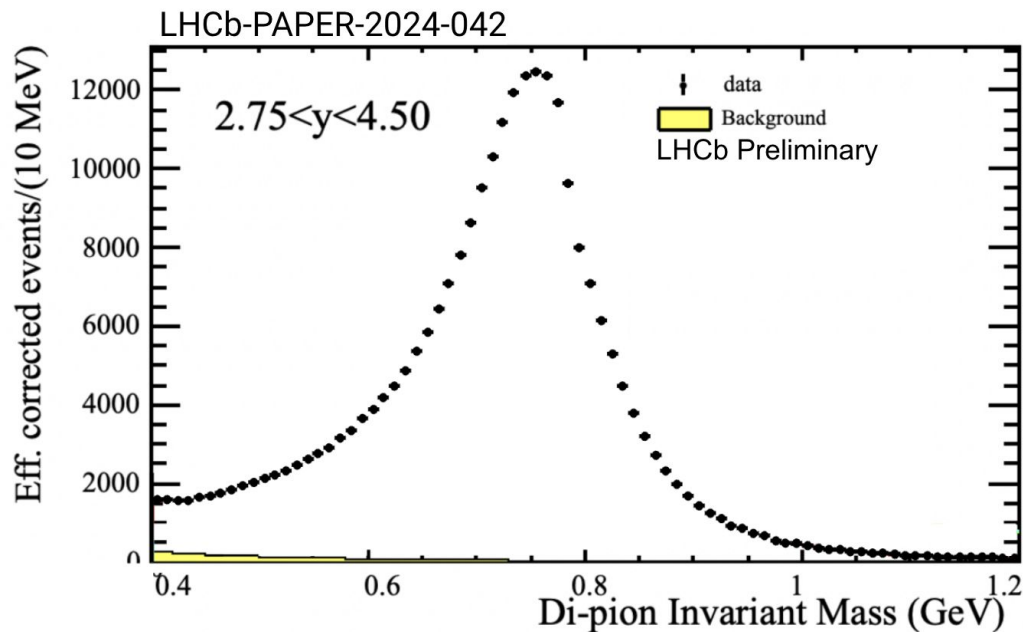
Selection Criteria

- Triggered Low Multiplicity Events ($\text{SPD} < 50$)
- Two oppositely charged tracks
- Fiducial region defined by $2.05 < y_{(\text{parent})} < 4.9$, $p_{T(\text{Track})} > 100$ MeV, and $2 < \eta_{\text{Track}} < 5$
- Invariant mass > 400 MeV
- Both tracks consistent with being a pion (using PID)
- Transverse momentum of the system < 100 MeV

Resultant Data

- Final sample contains ~12 million candidates

Backgrounds



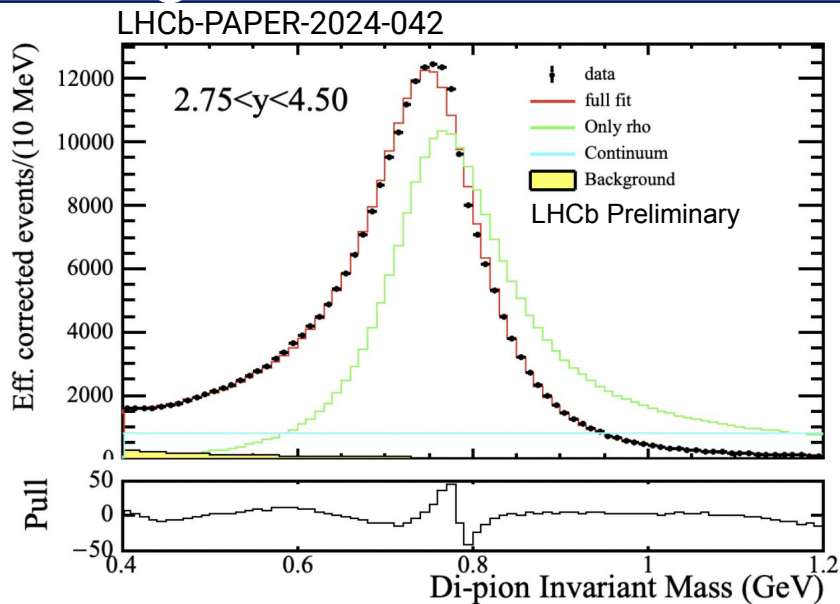
Backgrounds $\sim < 1\%$

- **Gamma-gamma production:**
 - Diphoton production is peaked at low masses and low p_T
 - Hadron cross-section is an order of magnitude lower than leptons.
 - The shape of the $\gamma\gamma \rightarrow ee/\mu\mu$ backgrounds were determined by fitting events identified as electrons or muons.
 - The amount of contamination was determined using simulation, the shape is scaled by the ratio of missing leptons compared to the rate for tagging both.
 - Contamination from $\gamma\gamma \rightarrow ee$ and $\gamma\gamma \rightarrow \mu\mu$ processes $\sim 0.5\%$.

Fitting the Invariant Mass Spectrum

- Söding basic
- Söding with ω (STAR)
- Söding with ω (H1)
- Söding with ω (H1) + form factor + continuum phase + ρ'

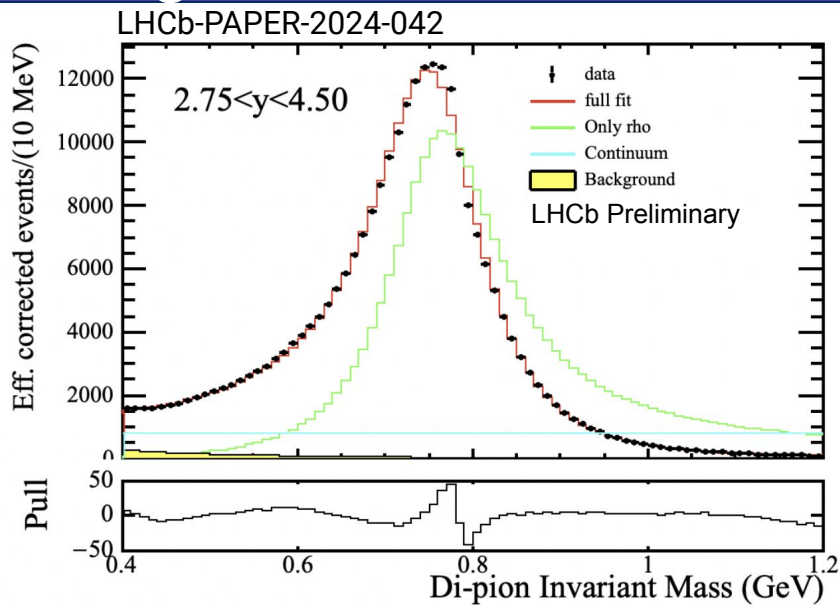
Söding basic



$$\mathcal{S}(M_{\pi\pi}) \propto \left| A \frac{\sqrt{M_{\pi\pi} M_\rho \Gamma}}{M_{\pi\pi}^2 - M_\rho^2 + i M_\rho \Gamma} + B \right|^2 \quad (1)$$

Experiment	Zeus [4]	H1 [5]	CMS [6]	ALICE [7]	LHCb Pb-Pb
Fit range [GeV]	[0.55,1.2]	[0.6,1.1]	[0.5,1.2]	[0.6,1.5]	[0.4,1.2]
M_ρ [MeV]	770 ± 2	769 ± 4	776 ± 1	$762^{+6.5}_{-3.8}$	771 ± 3
Γ_ρ [MeV]	146 ± 13	162 ± 8	154 ± 3	150^{+13}_{-7}	150 ± 4
$ B/A $	0.70 ± 0.04	0.57 ± 0.09 [45]	0.50 ± 0.05	$0.50^{+0.10}_{-0.06}$	0.72 ± 0.04

Söding basic

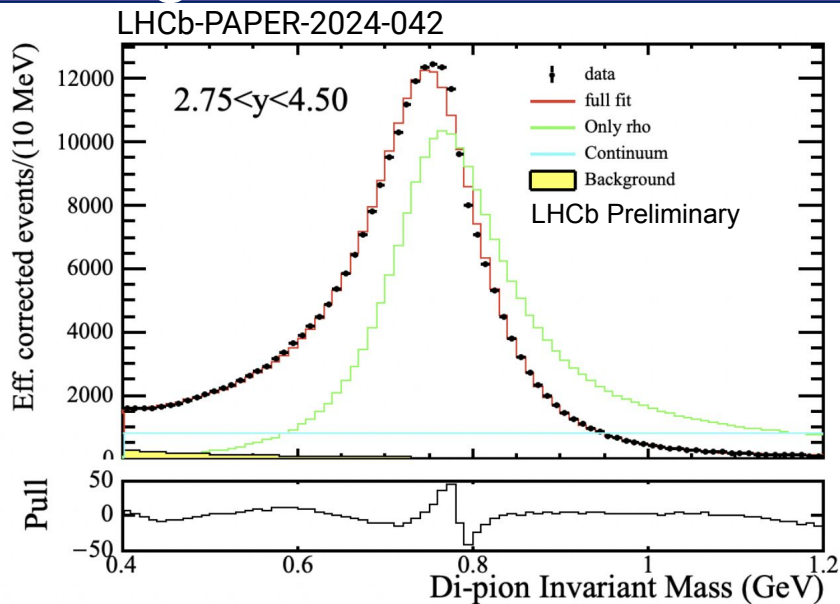


ρ resonance

$$\mathcal{S}(M_{\pi\pi}) \propto \left| A \frac{\sqrt{M_{\pi\pi} M_{\rho}} \Gamma}{M_{\pi\pi}^2 - M_{\rho}^2 + i M_{\rho} \Gamma} + B \right|^2 \quad (1)$$

Experiment	Zeus [4]	H1 [5]	CMS [6]	ALICE [7]	LHCb Pb-Pb
Fit range [GeV]	[0.55,1.2]	[0.6,1.1]	[0.5,1.2]	[0.6,1.5]	[0.4,1.2]
M_{ρ} [MeV]	770 ± 2	769 ± 4	776 ± 1	$762^{+6.5}_{-3.8}$	771 ± 3
Γ_{ρ} [MeV]	146 ± 13	162 ± 8	154 ± 3	150^{+13}_{-7}	150 ± 4
$ B/A $	0.70 ± 0.04	0.57 ± 0.09 [45]	0.50 ± 0.05	$0.50^{+0.10}_{-0.06}$	0.72 ± 0.04

Söding basic



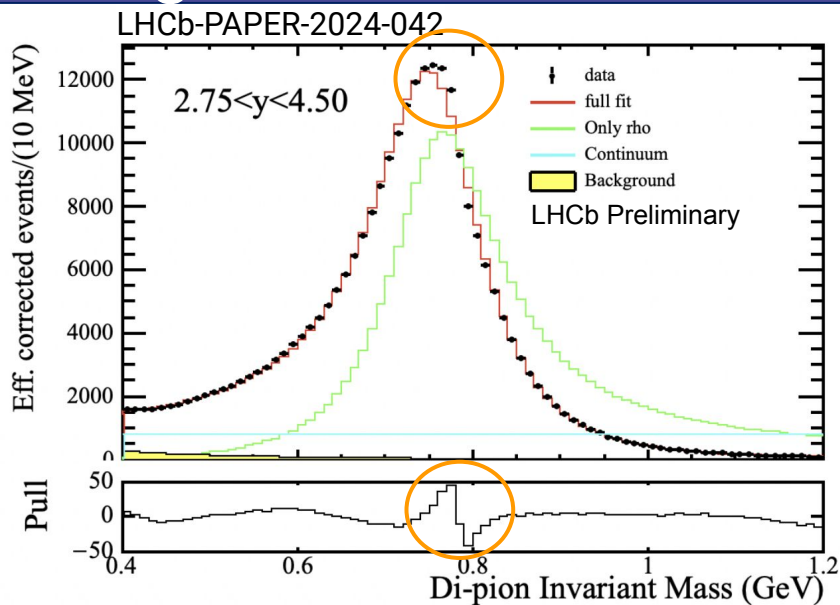
ρ resonance

$$\mathcal{S}(M_{\pi\pi}) \propto \left| A \frac{\sqrt{M_{\pi\pi} M_{\rho} \Gamma}}{M_{\pi\pi}^2 - M_{\rho}^2 + i M_{\rho} \Gamma} + B \right|^2 \quad (1)$$

flat continuum

Experiment	Zeus [4]	H1 [5]	CMS [6]	ALICE [7]	LHCb Pb-Pb
Fit range [GeV]	[0.55,1.2]	[0.6,1.1]	[0.5,1.2]	[0.6,1.5]	[0.4,1.2]
M_{ρ} [MeV]	770 ± 2	769 ± 4	776 ± 1	$762^{+6.5}_{-3.8}$	771 ± 3
Γ_{ρ} [MeV]	146 ± 13	162 ± 8	154 ± 3	150^{+13}_{-7}	150 ± 4
$ B/A $	0.70 ± 0.04	0.57 ± 0.09 [45]	0.50 ± 0.05	$0.50^{+0.10}_{-0.06}$	0.72 ± 0.04

Söding basic



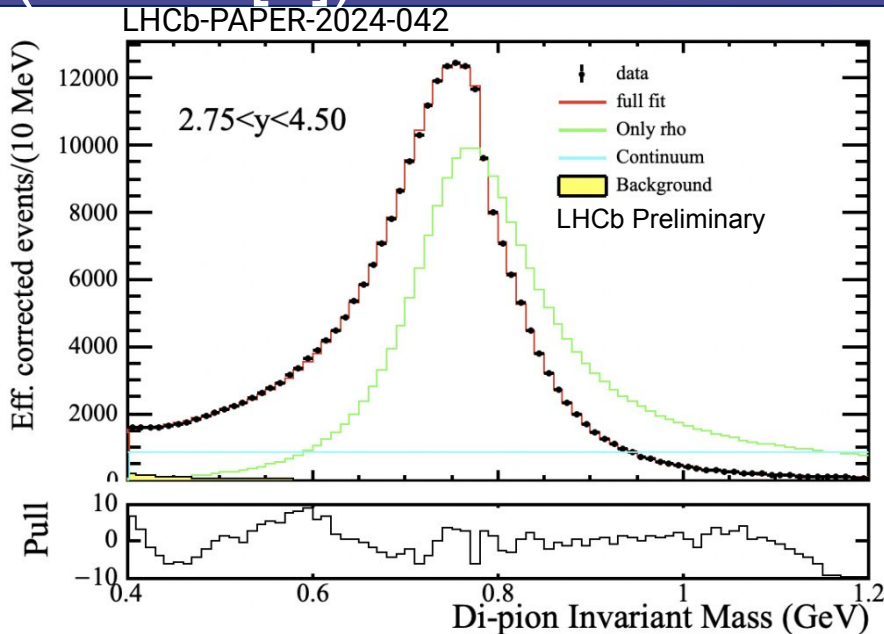
ρ resonance

$$\mathcal{S}(M_{\pi\pi}) \propto \left| A \frac{\sqrt{M_{\pi\pi} M_{\rho} \Gamma}}{M_{\pi\pi}^2 - M_{\rho}^2 + i M_{\rho} \Gamma} + B \right|^2 \quad (1)$$

flat continuum

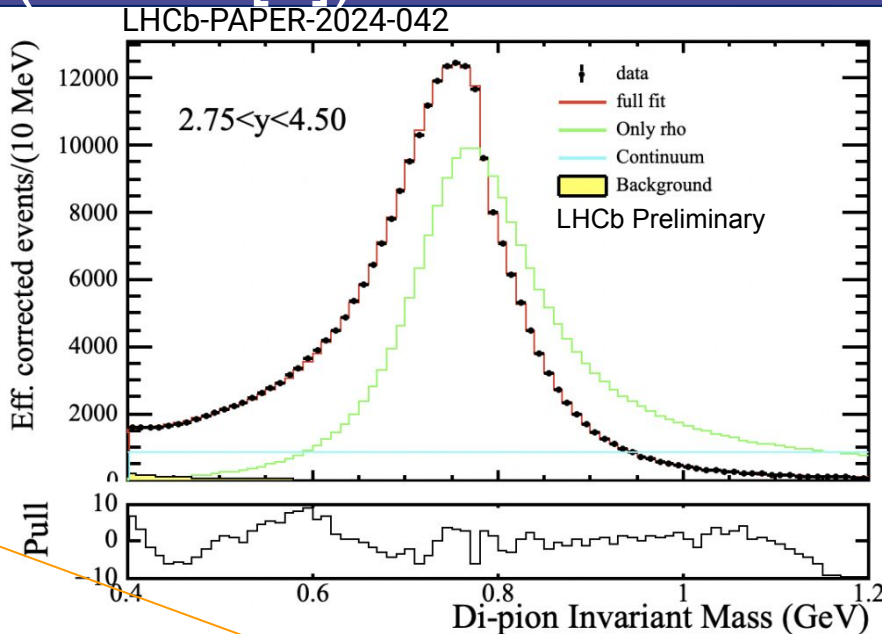
Experiment	Zeus [4]	H1 [5]	CMS [6]	ALICE [7]	LHCb Pb-Pb
Fit range [GeV]	[0.55,1.2]	[0.6,1.1]	[0.5,1.2]	[0.6,1.5]	[0.4,1.2]
M_{ρ} [MeV]	770 ± 2	769 ± 4	776 ± 1	$762^{+6.5}_{-3.8}$	771 ± 3
Γ_{ρ} [MeV]	146 ± 13	162 ± 8	154 ± 3	150^{+13}_{-7}	150 ± 4
$ B/A $	0.70 ± 0.04	0.57 ± 0.09 [45]	0.50 ± 0.05	$0.50^{+0.10}_{-0.06}$	0.72 ± 0.04

Söding + ω (STAR [8])



$$S = \left| A \frac{\sqrt{M_{\pi\pi} M_{\rho} \Gamma}}{M_{\pi\pi}^2 - M_{\rho}^2 + i M_{\rho} \Gamma} + B + C \exp(i\phi_{\omega}) \frac{\sqrt{M_{\pi\pi} M_{\omega} \Gamma_{\omega}}}{M_{\pi\pi}^2 - M_{\omega}^2 + i M_{\omega} \Gamma_{\omega}} \right|^2 \quad (2)$$

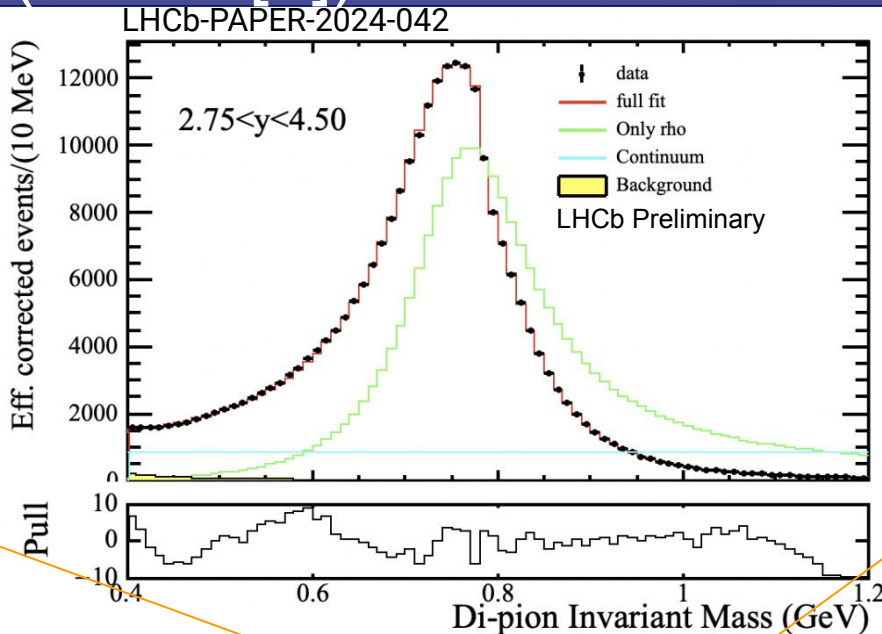
Söding + ω (STAR [8])



ω contribution

$$S = \left| A \frac{\sqrt{M_{\pi\pi} M_{\rho} \Gamma}}{M_{\pi\pi}^2 - M_{\rho}^2 + i M_{\rho} \Gamma} + B + C \exp(i\phi_{\omega}) \frac{\sqrt{M_{\pi\pi} M_{\omega} \Gamma_{\omega}}}{M_{\pi\pi}^2 - M_{\omega}^2 + i M_{\omega} \Gamma_{\omega}} \right|^2 \quad (2)$$

Söding + ω (STAR [8])

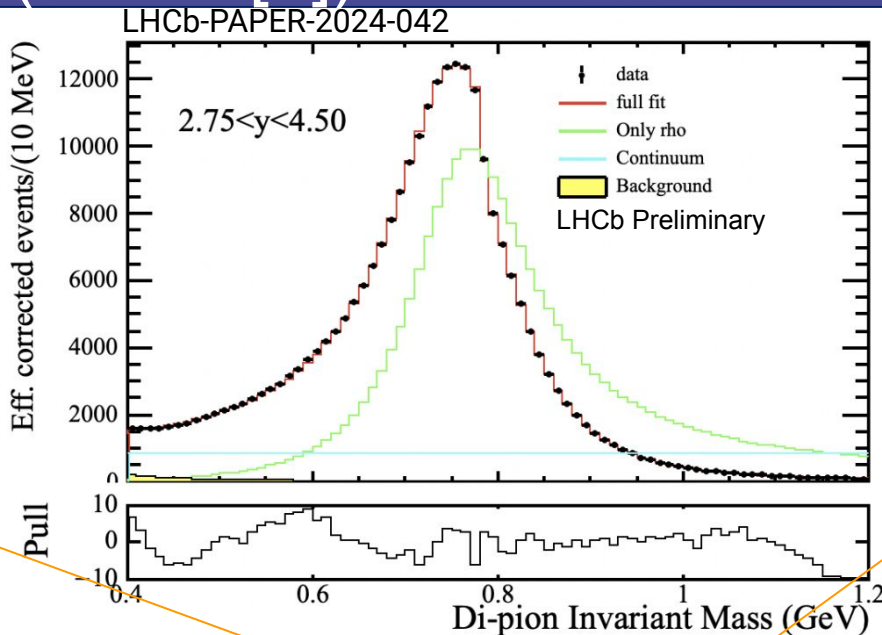


ω contribution

ω phase

$$S = \left| A \frac{\sqrt{M_{\pi\pi} M_{\rho} \Gamma}}{M_{\pi\pi}^2 - M_{\rho}^2 + i M_{\rho} \Gamma} + B + C \exp(i\phi_{\omega}) \frac{\sqrt{M_{\pi\pi} M_{\omega} \Gamma_{\omega}}}{M_{\pi\pi}^2 - M_{\omega}^2 + i M_{\omega} \Gamma_{\omega}} \right|^2 \quad (2)$$

Söding + ω (STAR [8])



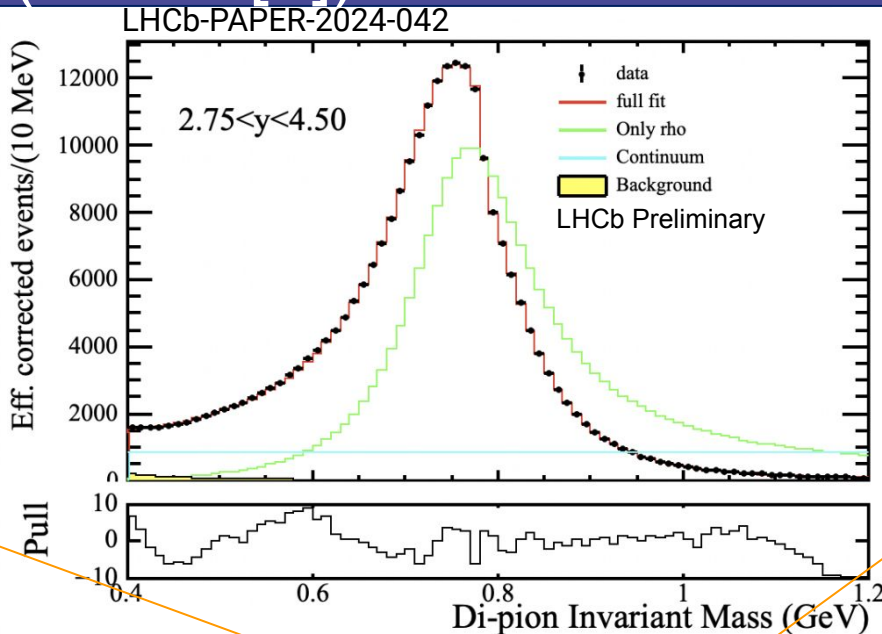
- ω mass and width are fixed to PDG values

ω contribution

ω phase

$$S = \left| A \frac{\sqrt{M_{\pi\pi} M_{\rho} \Gamma}}{M_{\pi\pi}^2 - M_{\rho}^2 + i M_{\rho} \Gamma} + B + C \exp(i\phi_{\omega}) \frac{\sqrt{M_{\pi\pi} M_{\omega} \Gamma_{\omega}}}{M_{\pi\pi}^2 - M_{\omega}^2 + i M_{\omega} \Gamma_{\omega}} \right|^2 \quad (2)$$

Söding + ω (STAR [8])



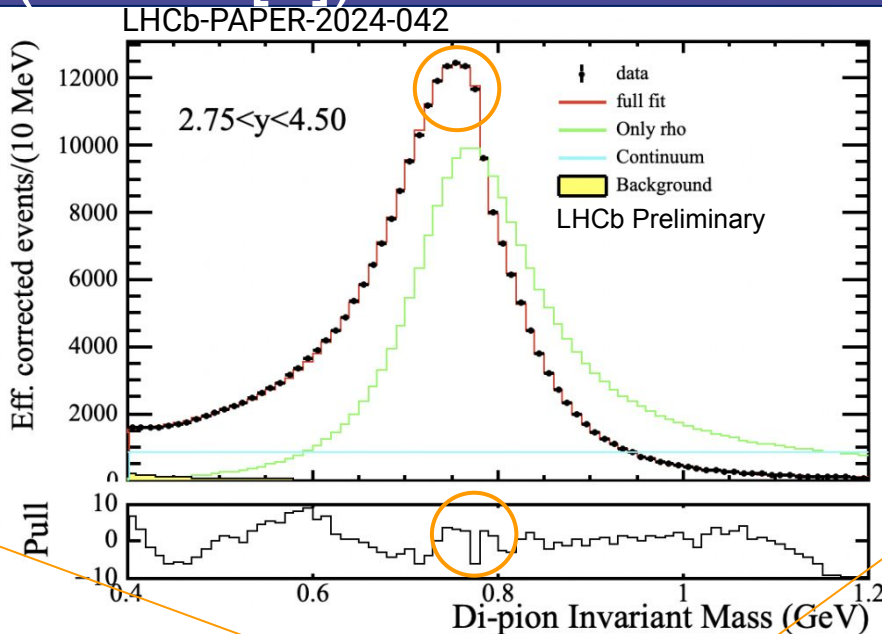
- ω mass and width are fixed to PDG values
- Small ω contribution, large effect from ω - ρ interference.

ω contribution

ω phase

$$S = \left| A \frac{\sqrt{M_{\pi\pi} M_{\rho} \Gamma}}{M_{\pi\pi}^2 - M_{\rho}^2 + i M_{\rho} \Gamma} + B + C \exp(i\phi_{\omega}) \frac{\sqrt{M_{\pi\pi} M_{\omega} \Gamma_{\omega}}}{M_{\pi\pi}^2 - M_{\omega}^2 + i M_{\omega} \Gamma_{\omega}} \right|^2 \quad (2)$$

Söding + ω (STAR [8])



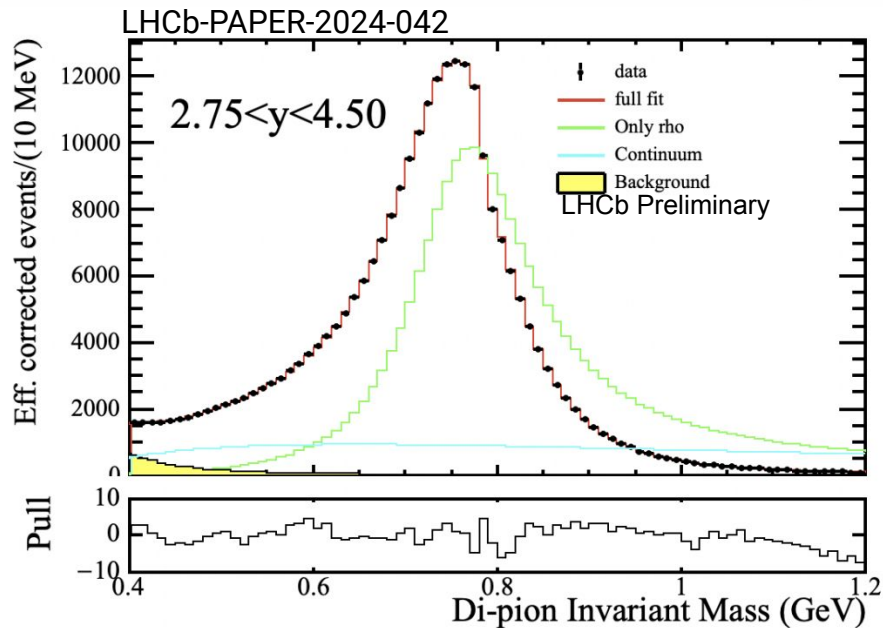
- ω mass and width are fixed to PDG values
- Small ω contribution, large effect from ω - ρ interference.

ω contribution

ω phase

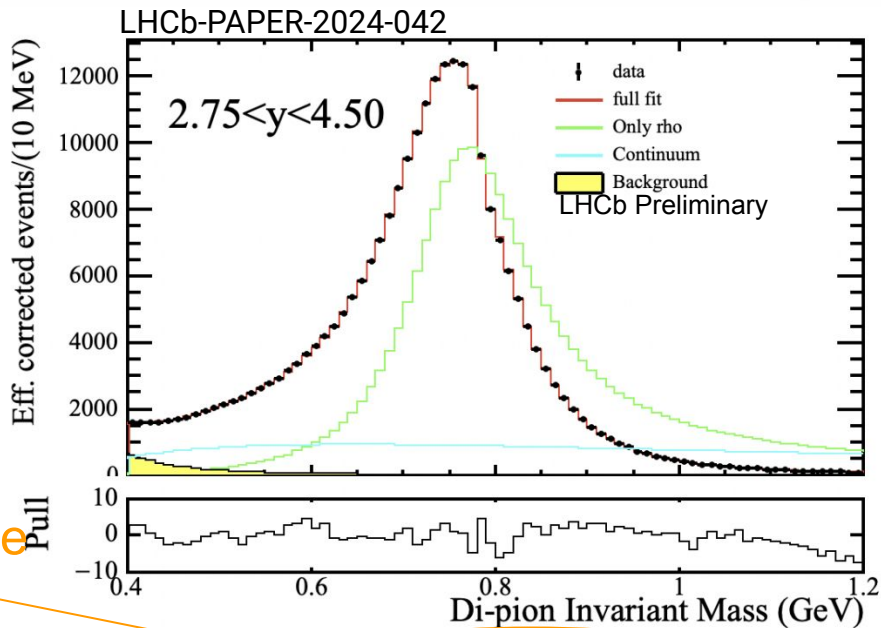
$$S = \left| A \frac{\sqrt{M_{\pi\pi} M_{\rho} \Gamma}}{M_{\pi\pi}^2 - M_{\rho}^2 + i M_{\rho} \Gamma} + B - C \exp(i\phi_{\omega}) \frac{\sqrt{M_{\pi\pi} M_{\omega} \Gamma_{\omega}}}{M_{\pi\pi}^2 - M_{\omega}^2 + i M_{\omega} \Gamma_{\omega}} \right|^2 \quad (2)$$

Söding + ω (H1 [9])



$$S = \frac{q^3(m_{\pi\pi})}{q^3(m_\rho)} \left| \mathcal{B}\mathcal{W}_\rho(m_{\pi\pi}) \left(1 + C \exp(i\phi_\omega) \frac{m_{\pi\pi}^2}{m_\omega^2} \mathcal{B}\mathcal{W}_\omega(m_{\pi\pi}) \right) + \frac{B}{(m_{\pi\pi}^2 - 4m_\pi^2 + \Lambda^2)^\delta} \right|^2 \quad (3)$$

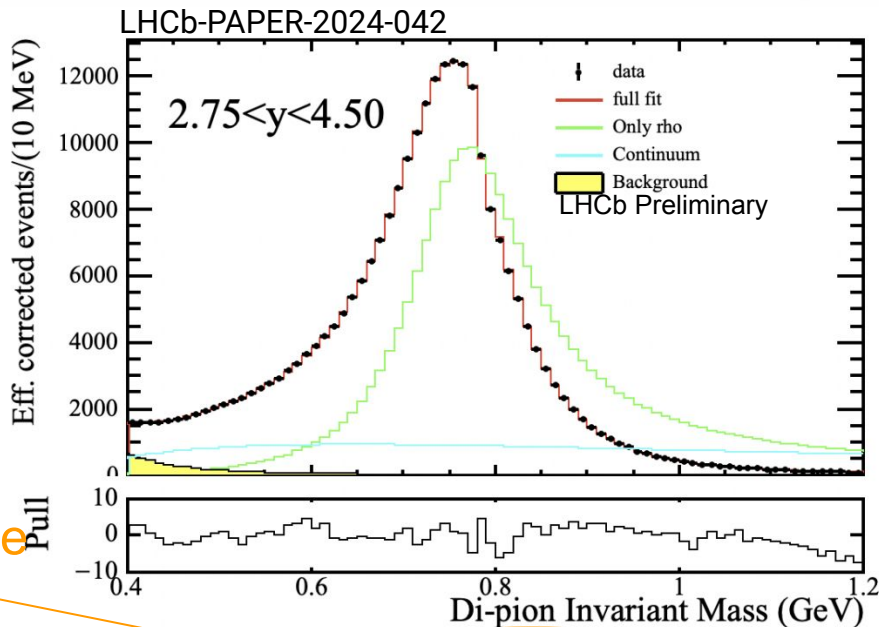
Söding + ω (H1 [9])



Modified ω
phase/interference

$$S = \frac{q^3(m_{\pi\pi})}{q^3(m_\rho)} \left| \mathcal{B} \mathcal{W}_\rho(m_{\pi\pi}) \left(1 + C \exp(i\phi_\omega) \frac{m_{\pi\pi}^2}{m_\omega^2} \mathcal{B} \mathcal{W}_\omega(m_{\pi\pi}) \right) + \frac{B}{(m_{\pi\pi}^2 - 4m_\pi^2 + \Lambda^2)^\delta} \right|^2 \quad (3)$$

Söding + ω (H1 [9])

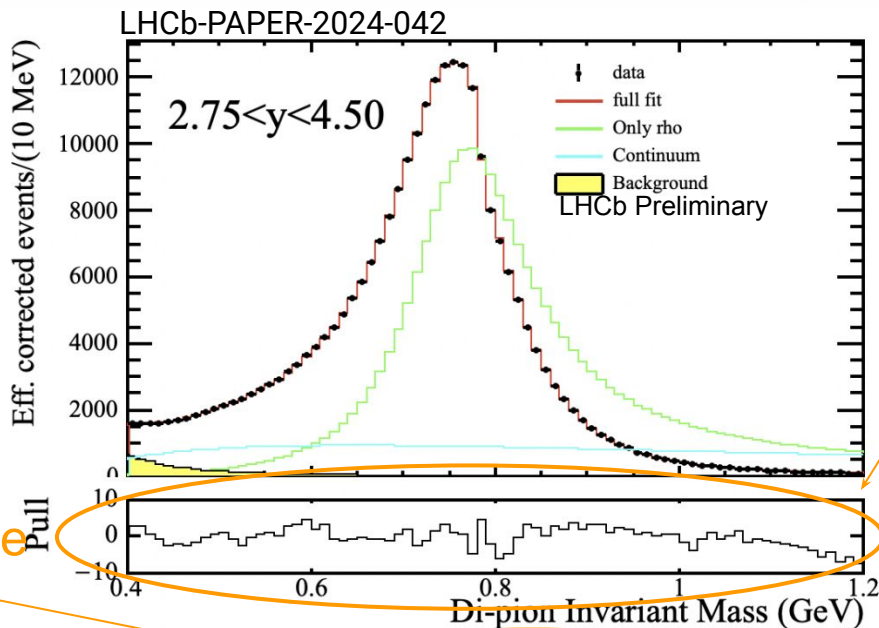


Modified ω
phase/interference

Modified
continuum

$$S = \frac{q^3(m_{\pi\pi})}{q^3(m_\rho)} \left| \mathcal{B} \mathcal{W}_\rho(m_{\pi\pi}) \left(1 + C \exp(i\phi_\omega) \frac{m_{\pi\pi}^2}{m_\omega^2} \mathcal{B} \mathcal{W}_\omega(m_{\pi\pi}) \right) \right|^2 + \left| \frac{B}{(m_{\pi\pi}^2 - 4m_\pi^2 + \Lambda^2)^\delta} \right|^2 \quad (3)$$

Söding + ω (H1 [9])



Fewer Deviations
on Pull

Modified
continuum

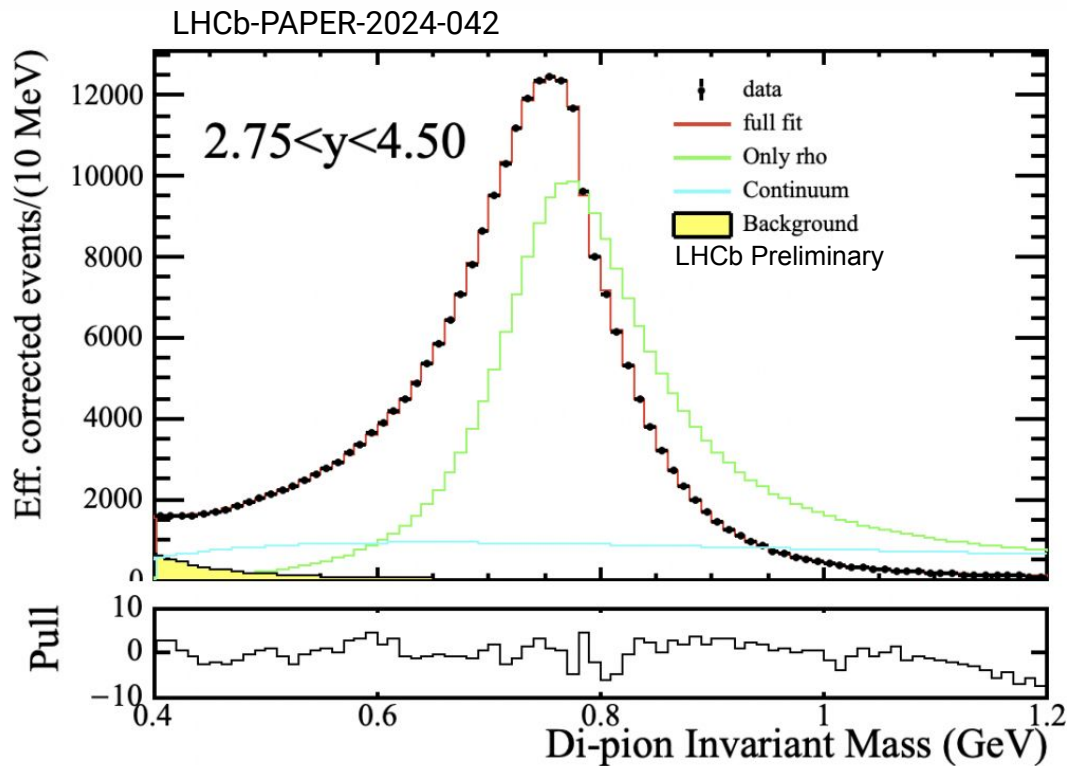
Modified ω
phase/interference

$$S = \frac{q^3(m_{\pi\pi})}{q^3(m_\rho)} \left| \mathcal{B} \mathcal{W}_\rho(m_{\pi\pi}) \left(1 + C \exp(i\phi_\omega) \frac{m_{\pi\pi}^2}{m_\omega^2} \mathcal{B} \mathcal{W}_\omega(m_{\pi\pi}) \right) + \frac{B}{(m_{\pi\pi}^2 - 4m_\pi^2 + \Lambda^2)^\delta} \right|^2 \quad (3)$$

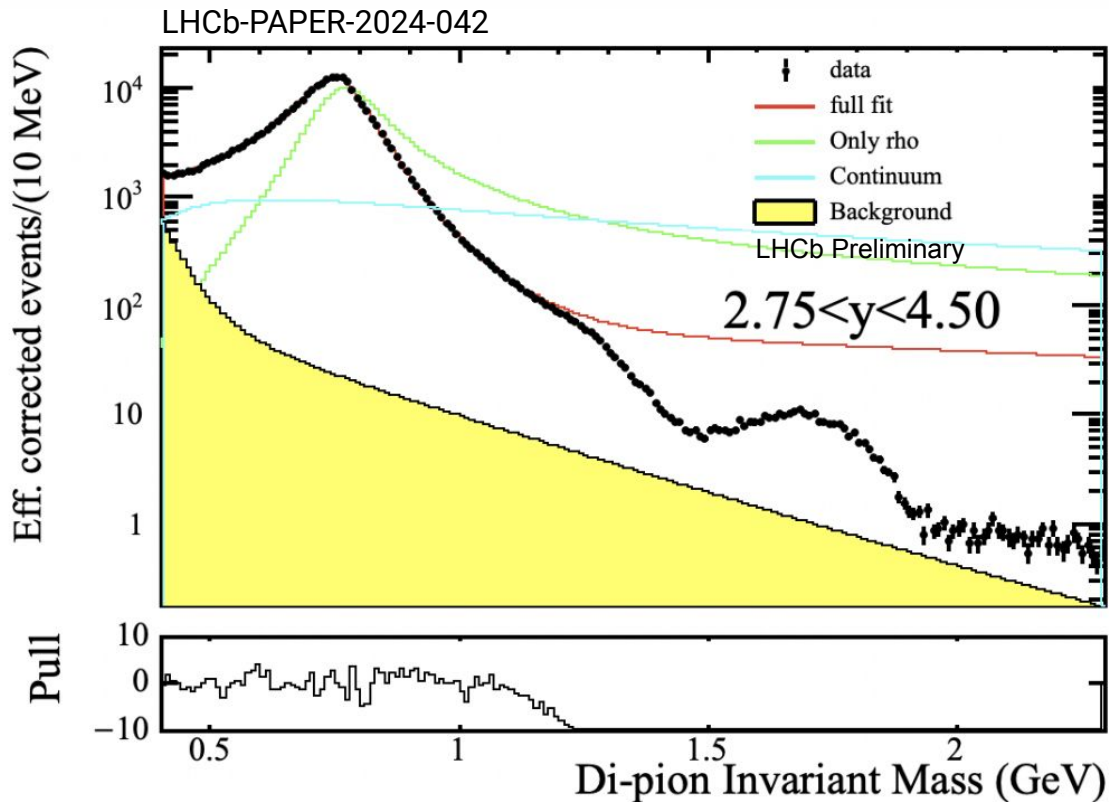
Fit Results

LHCb Preliminary LHCb-PAPER-2024-042	LHCb		STAR [8]	H1 [9]
	(Star eq. 2)	(H1 eq. 3)		
M_ρ [MeV]	774 ± 3	776 ± 3	776.2 ± 0.2	771 ± 3
Γ_ρ [MeV]	156 ± 3	153 ± 3	156 ± 1	151 ± 3
$ B/A $	0.73 ± 0.03	$0.19 \pm .02$	0.79 ± 0.08	0.19 ± 0.04
ϕ_ω [rad]	1.36 ± 0.03	$-0.23 \pm .04$	1.46 ± 0.11	-0.5 ± 0.3
$ C/A $	0.34 ± 0.03	$0.18 \pm .01$	0.36 ± 0.05	0.17 ± 0.02
Λ [MeV]	-	366 ± 110		180 ± 590
δ	-	$1.07 \pm .11$	-	0.76 ± 0.35

But What about Masses > 1.2 GeV?

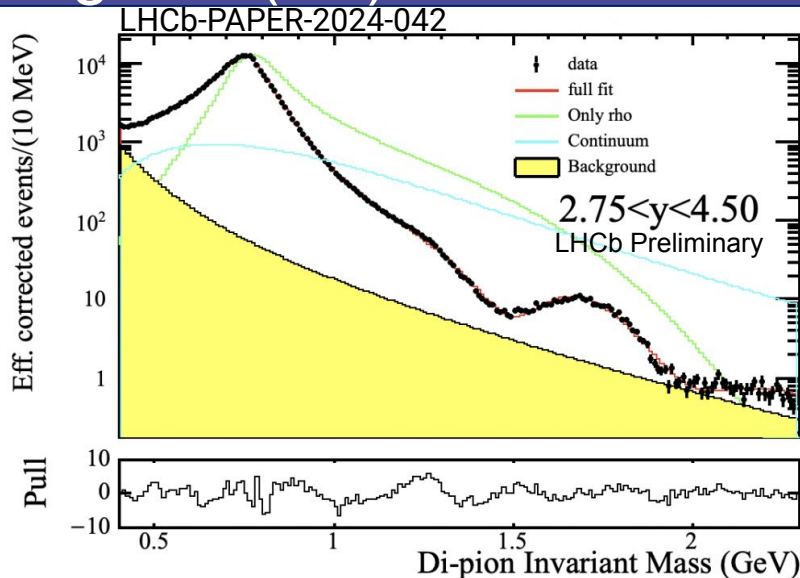


Extended Frame



- Data at high mass falls well below fit
- Does not account for clear structure around 1.7 GeV

Söding + ω (H1) + form factor + continuum phase + ρ'



Fit the data with:

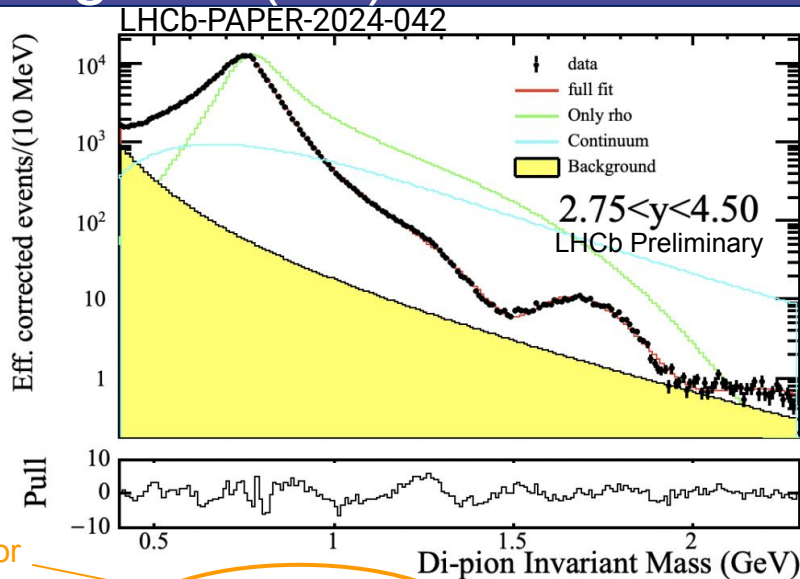
- Added form factor
- High-mass Breit-Wigner term
- Interference from continuum

Alternative parameterizations are possible

$$S = \frac{q^3(m_{\pi\pi})}{q^3(m_\rho)} \left| \exp\left(-\left(\frac{m_{\pi\pi}^2 - m_\rho^2}{\Delta^2}\right)^2\right) \mathcal{BW}_\rho(m_{\pi\pi}) \left(1 + C \exp(i\phi_\omega) \frac{m_{\pi\pi}^2}{m_\omega^2} \mathcal{BW}_\omega(m_{\pi\pi})\right) + \frac{B \exp(i\phi_c)}{(m_{\pi\pi}^2 - 4m_\pi^2 + \Lambda^2)^\delta} + D \exp(i\phi_{\rho'}) \mathcal{BW}_{\rho'} \right|^2 \quad (4)$$

(6)

Söding + ω (H1) + form factor + continuum phase + ρ'



Fit the data with:

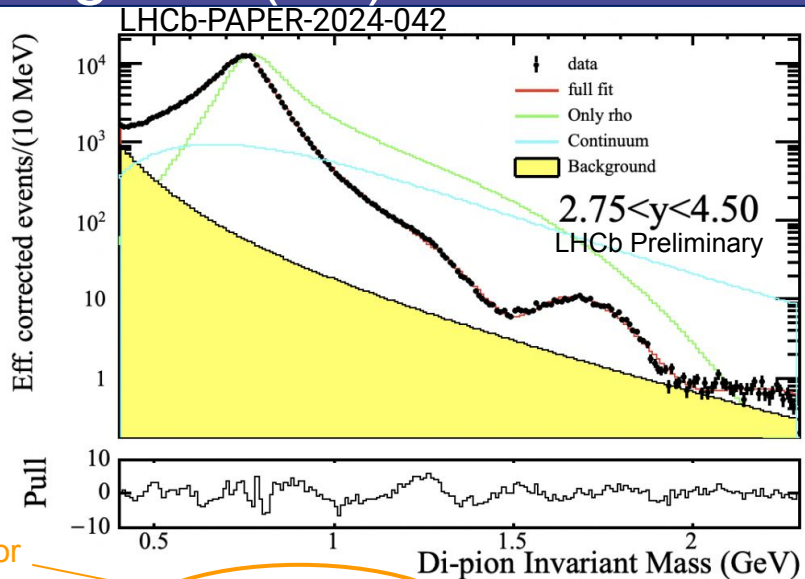
- Added form factor
- High-mass Breit-Wigner term
- Interference from continuum

Alternative parameterizations are possible

$$S = \frac{q^3(m_{\pi\pi})}{q^3(m_\rho)} \left| \exp\left(-\left(\frac{m_{\pi\pi}^2 - m_\rho^2}{\Delta^2}\right)^2\right) \mathcal{B}\mathcal{W}_\rho(m_{\pi\pi}) \left(1 + C \exp(i\phi_\omega) \frac{m_{\pi\pi}^2}{m_\omega^2} \mathcal{B}\mathcal{W}_\omega(m_{\pi\pi})\right) + \frac{B \exp(i\phi_c)}{(m_{\pi\pi}^2 - 4m_\pi^2 + \Lambda^2)^\delta} + D \exp(i\phi_{\rho'}) \mathcal{B}\mathcal{W}_{\rho'} \right|^2 \quad (4)$$

(6)

Söding + ω (H1) + form factor + continuum phase + ρ'



Fit the data with:

- Added form factor
- High-mass Breit-Wigner term
- Interference from continuum

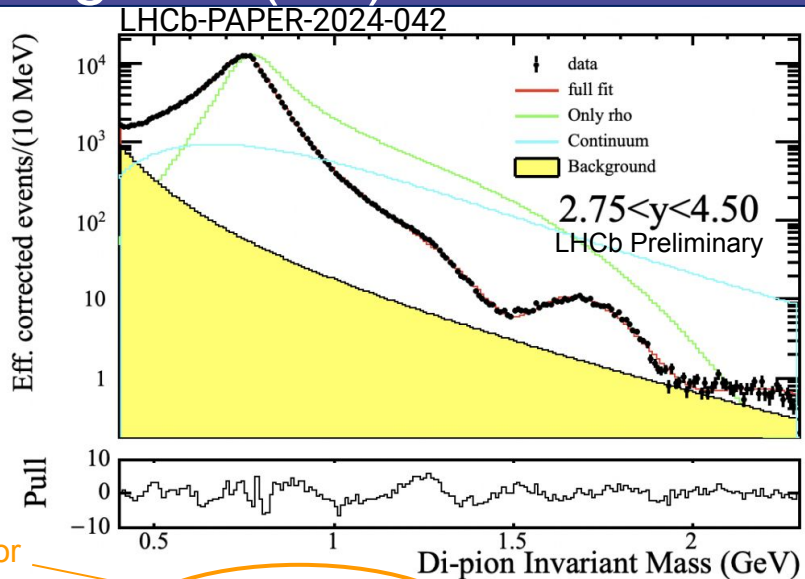
Alternative parameterizations are possible

$$S = \frac{q^3(m_{\pi\pi})}{q^3(m_\rho)} \left| \exp\left(-\left(\frac{m_{\pi\pi}^2 - m_\rho^2}{\Delta^2}\right)^2\right) \mathcal{B}\mathcal{W}_\rho(m_{\pi\pi}) \left(1 + C \exp(i\phi_\omega) \frac{m_{\pi\pi}^2}{m_\omega^2} \mathcal{B}\mathcal{W}_\omega(m_{\pi\pi})\right) + \frac{B \exp(i\phi_c)}{(m_{\pi\pi}^2 - 4m_\pi^2 + \Lambda^2)^\delta} + D \exp(i\phi_{\rho'}) \mathcal{B}\mathcal{W}_{\rho'} \right|^2 \quad (4)$$

(6)

ρ' resonance

Söding + ω (H1) + form factor + continuum phase + ρ'



Fit the data with:

- Added form factor
- High-mass Breit-Wigner term
- Interference from continuum

Alternative parameterizations are possible

Form Factor

$$S = \frac{q^3(m_{\pi\pi})}{q^3(m_\rho)} \exp\left(-\left(\frac{m_{\pi\pi}^2 - m_\rho^2}{\Delta^2}\right)^2\right) \mathcal{B}\mathcal{W}_\rho(m_{\pi\pi}) \left(1 + C \exp(i\phi_\omega) \frac{m_{\pi\pi}^2}{m_\omega^2} \mathcal{B}\mathcal{W}_\omega(m_{\pi\pi})\right)$$

(6)

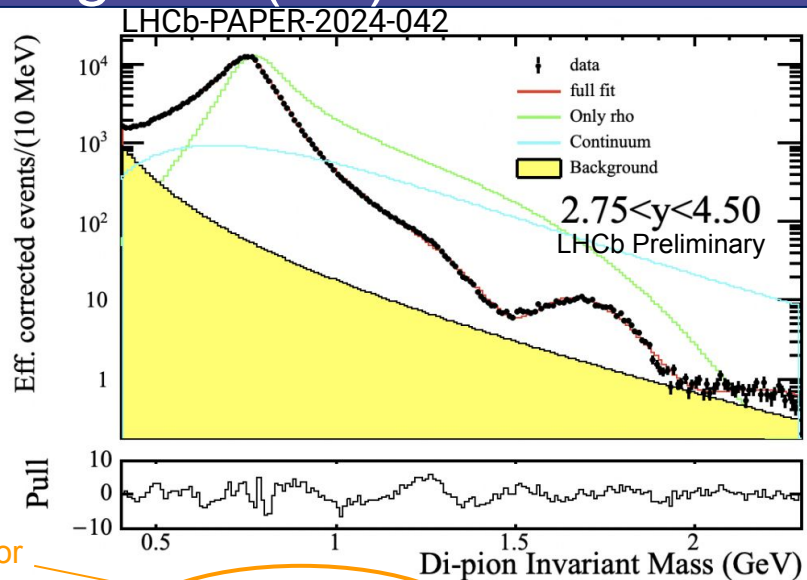
Continuum Phase Shift

$$+ \frac{B \exp(i\phi_c)}{(m_{\pi\pi}^2 - 4m_\pi^2 + \Lambda^2)^\delta} + D \exp(i\phi_{\rho'}) \mathcal{B}\mathcal{W}_{\rho'}^2$$

(4)

ρ' resonance

Söding + ω (H1) + form factor + continuum phase + ρ'



Form Factor

$$S = \frac{q^3(m_{\pi\pi})}{q^3(m_\rho)} \left| \exp\left(-\left(\frac{m_{\pi\pi}^2 - m_\rho^2}{\Delta^2}\right)^2\right) \mathcal{B}\mathcal{W}_\rho(m_{\pi\pi}) \left(1 + C \exp(i\phi_\omega) \frac{m_{\pi\pi}^2}{m_\omega^2} \mathcal{B}\mathcal{W}_\omega(m_{\pi\pi})\right) \right|$$

Continuum Phase Shift

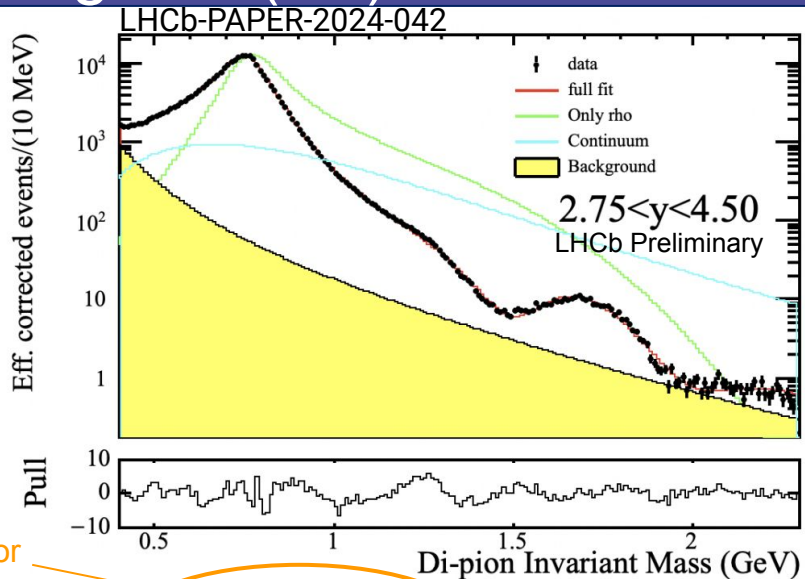
$$+ \frac{B \exp(i\phi_c)}{(m_{\pi\pi}^2 - 4m_\pi^2 + \Lambda^2)^\delta} + D \exp(i\phi_{\rho'}) \mathcal{B}\mathcal{W}_{\rho'} \quad (4)$$

 ρ' resonance

fit to eq. 4

M_ρ [MeV]	777.0 ± 0.1
Γ_ρ [MeV]	152.2 ± 0.2
ϕ_c	-0.52 ± 0.02
$ B/A $	1.1 ± 0.1
ϕ_ω [rad]	-0.14 ± 0.01
$ C/A $	0.160 ± 0.002
Λ [MeV]	1.04 ± 0.02
δ	2.98 ± 0.05
Δ [MeV]	1.49 ± 0.01
$ D/A $	$0.0124 \pm 0.0003,$
$M_{\rho'}$ [MeV]	1620 ± 8
$\Gamma_{\rho'}$ [MeV]	462 ± 10
$\phi_{\rho'}$	-0.26 ± 0.03

Söding + ω (H1) + form factor + continuum phase + ρ'



Form Factor

$$S = \frac{q^3(m_{\pi\pi})}{q^3(m_\rho)} \left| \exp\left(-\left(\frac{m_{\pi\pi}^2 - m_\rho^2}{\Delta^2}\right)^2\right) \mathcal{B}\mathcal{W}_\rho(m_{\pi\pi}) \left(1 + C \exp(i\phi_\omega) \frac{m_{\pi\pi}^2}{m_\omega^2} \mathcal{B}\mathcal{W}_\omega(m_{\pi\pi})\right) \right|$$

Continuum Phase Shift

$$+ \frac{B \exp(i\phi_c)}{(m_{\pi\pi}^2 - 4m_\pi^2 + \Lambda^2)^\delta} + D \exp(i\phi_{\rho'}) \mathcal{B}\mathcal{W}_{\rho'} \quad (4)$$

 ρ' resonance

fit to eq. 4

M_ρ [MeV]	777.0 ± 0.1
Γ_ρ [MeV]	152.2 ± 0.2
ϕ_c	-0.52 ± 0.02
$ B/A $	1.1 ± 0.1
ϕ_ω [rad]	-0.14 ± 0.01
$ C/A $	0.160 ± 0.002
Λ [MeV]	1.04 ± 0.02
δ	2.98 ± 0.05
Δ [MeV]	1.49 ± 0.01
$ D/A $	$0.0124 \pm 0.0003,$
$M_{\rho'}$ [MeV]	1620 ± 8
$\Gamma_{\rho'}$ [MeV]	462 ± 10
$\phi_{\rho'}$	-0.26 ± 0.03

} = Strongly Model Dependent

Conclusions and Outlook

Conclusions

- Very clean sample with approximately 1% background
- Extended measurement range to 0.4 - 2.3 GeV
 - Previous results from 0.6-1.0 GeV (H1 [9]) and 0.6-1.3 GeV (STAR [8])
- Distinct and well-resolved resonance observed at 1.7 GeV
- Full description of the di-pion spectrum requires phenomenological input.

Outlook

- Will feed into cross section measurements.

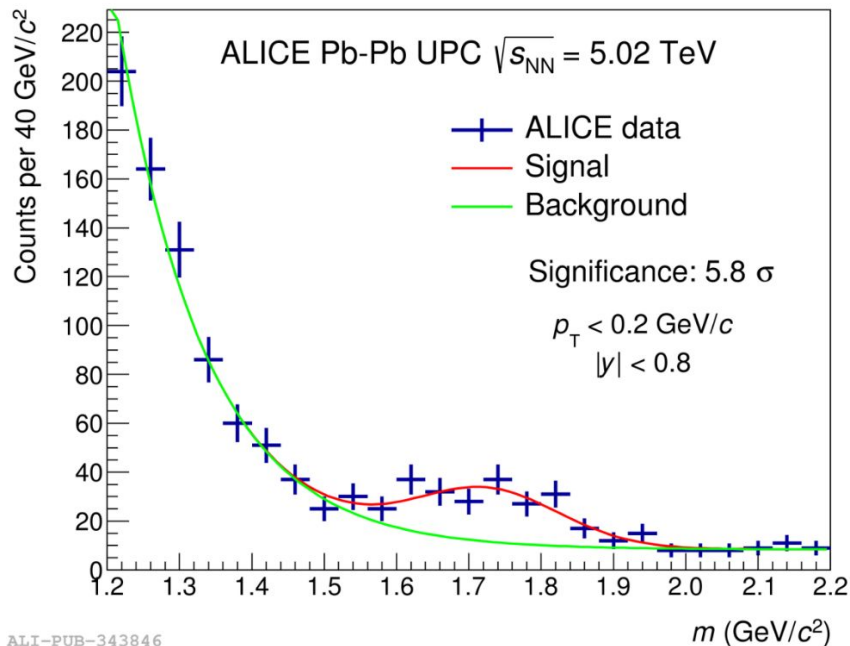
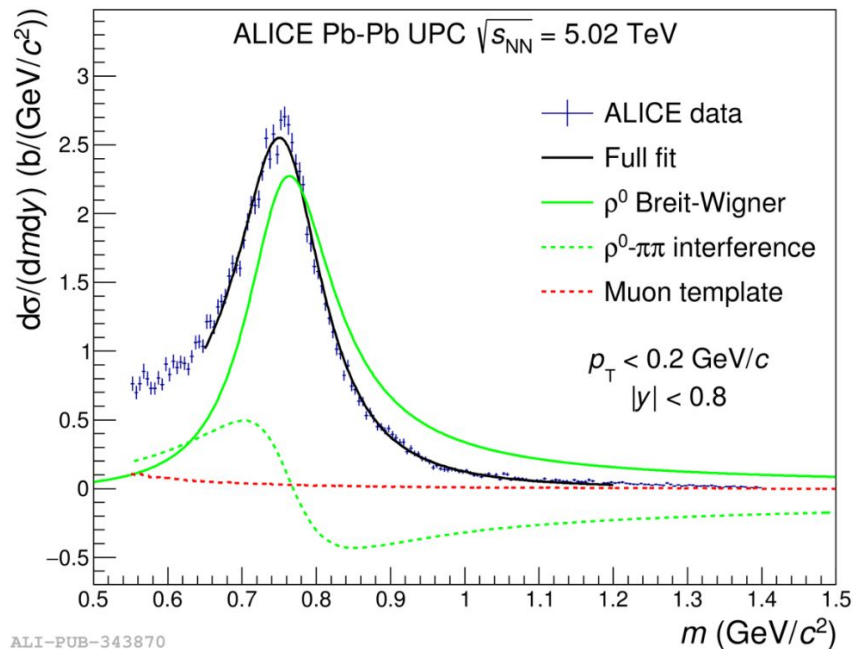
References

- [1] M. Davier, A. Hoecker, B. Malaescu, and Z. Zhang, Reevaluation of the hadronic $3\pi^0$ vacuum polarisation contributions to the Standard Model predictions of the muon $g - 2$ and $\alpha(m_Z)$ using newest hadronic cross-section data, Eur. Phys. J. C 77 (2017) 387–827, arXiv:1706.09436.
- [2] Int. J. Mod. Phys. A 30 (2015) 1530022
- [3] JINST 13 (2018) no.04, P04017
- [4] ZEUS, J. Breitweg et al., Elastic and proton dissociative ρ^0 photoproduction at HERA, Eur. Phys. J. C 2 (1998) 247, arXiv:hep-ex/9712020.
- [5] H1, F. D. Aaron et al., Diffractive Electroproduction of rho and phi Mesons at HERA, Eur. Phys. J. C 45 (2010) 032, arXiv:0910.5831.
- [6] CMS, A. M. Sirunyan et al., Measurement of exclusive $\rho(770)^0$ photoproduction in ultraperipheral pPb collisions at $\sqrt{s_{NN}} = 5.02$ TeV, Eur. Phys. J. C 79 (2019) 702, 457 arXiv:1902.01339.
- [7] ALICE, J. Adam et al., Coherent ρ^0 photoproduction in ultra-peripheral Pb-Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV, JHEP 09 (2015) 095, arXiv:1503.09177.
- [8] STAR, L. Adamczyk et al., Coherent diffractive photoproduction of ρ^0 mesons on gold nuclei at 200 GeV/nucleon-pair at the Relativistic Heavy Ion Collider, Phys. Rev. C 96 (2017) 054904, arXiv:1702.07705.
- [9] H1, V. Andreev et al., Measurement of Exclusive $\pi^+\pi^-$ and ρ^0 Meson Photoproduction at HERA, Eur. Phys. J. C 80 (2020) 1189, arXiv:2005.14471.
- [10] ALICE collaboration, Coherent photoproduction of 0^- vector mesons in ultra-peripheral Pb-Pb collisions at $\sqrt{s_{NN}} = 5.02$ TeV, JHEP 06 (2020) 035 [2002.10897].

Thank You

back-up

Alice ρ UPC in PbPb [10]



What is the Structure?

ALICE fitted an enhancement in this region with a Gaussian and obtained a mass of 1725 ± 17 MeV and a width of 143 ± 21 MeV compatible with the $\rho(1700)$ while STAR, with a similar fit, obtained a mass of 1653 ± 10 MeV and a width of 164 ± 15 MeV, compatible with the spin-3 $\rho_3(1690)$.