



Jets:

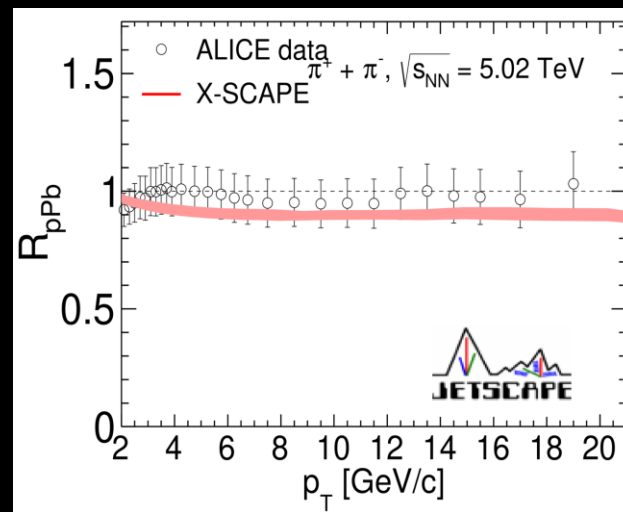
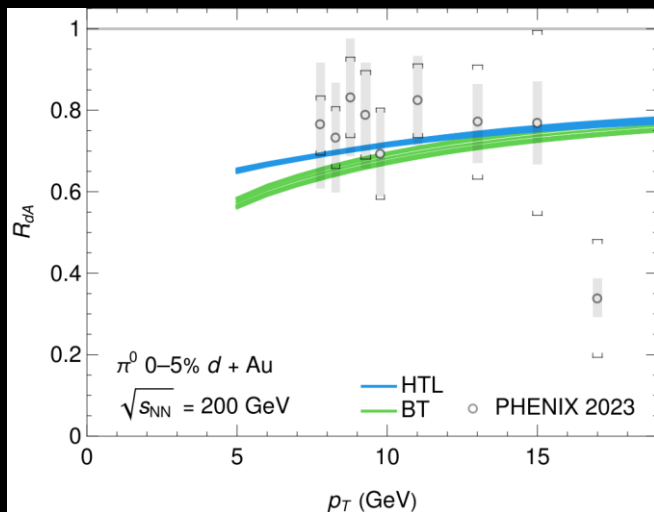
In-medium parton evolution in small systems

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Mandelstam Institute for Theoretical Physics (MITP)
SA-CERN Collaboration*

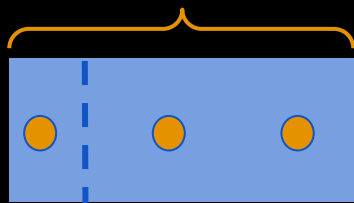
Thank you to Souvik Adhya, Carlota Andres, Jasmine Brewer, André Cordeiro, Dani Pablos, Wilke van der Schee, Ismael Soudi, Konrad Tywoniuk, Urs Wiedemann for advice and illuminating discussions.

Recent calculations



$$\frac{1}{\mu_D} \ll \Delta z \sim \lambda_{mfp} \ll L$$

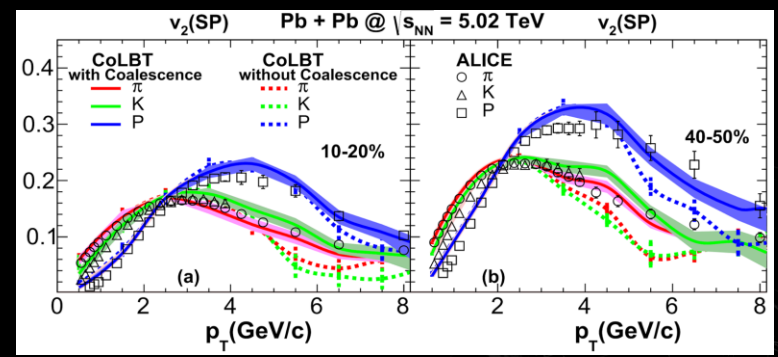
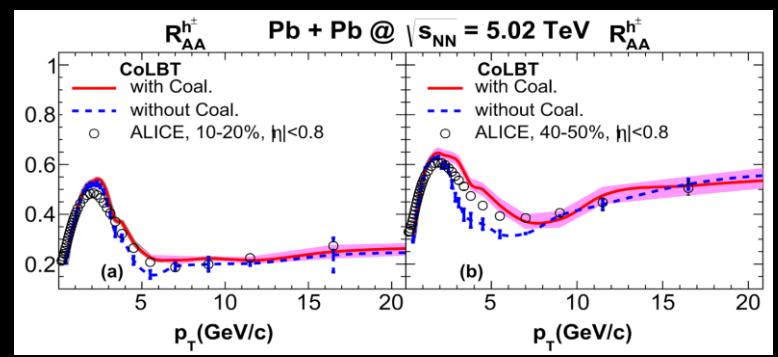
$$\frac{1}{\mu_D} \ll \lambda_{mfp}$$



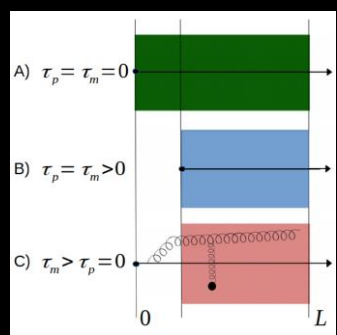
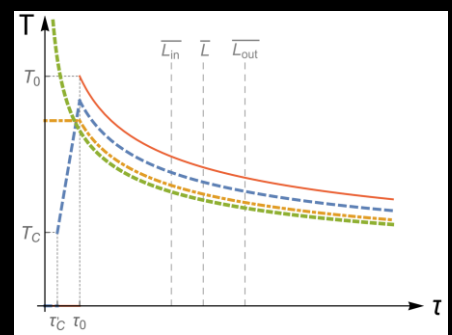
It's not the things you don't know that will hurt you..

It's the things you're sure you know,
that just ain't so. – Mark Twain

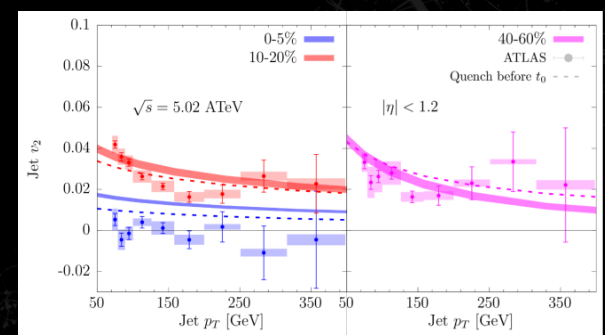
$R_{AA} \otimes v_2$ non-trivial even in AA



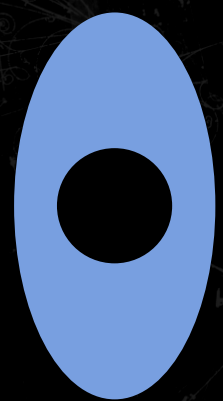
Sensitive to hadronization



Sensitive to late-time temperature profile?

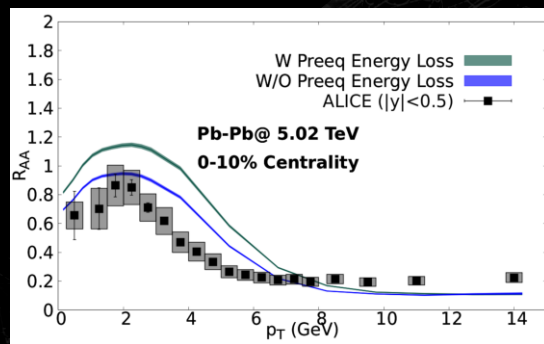
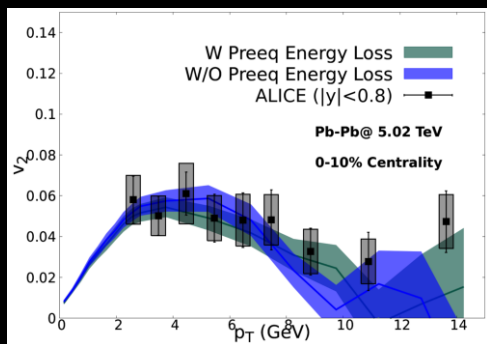
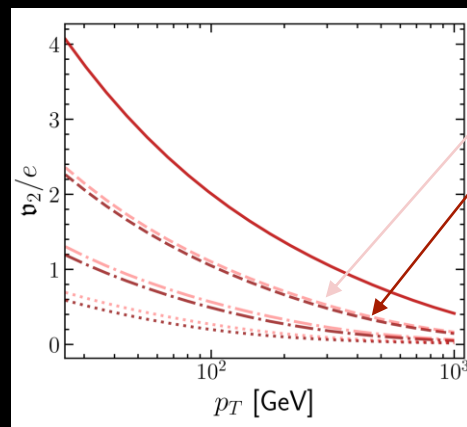
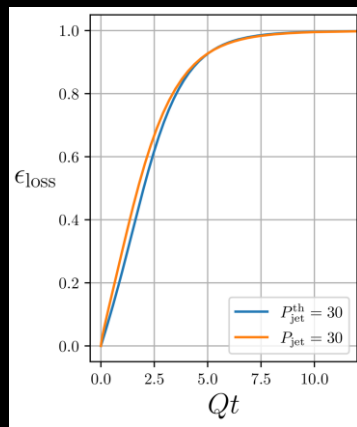
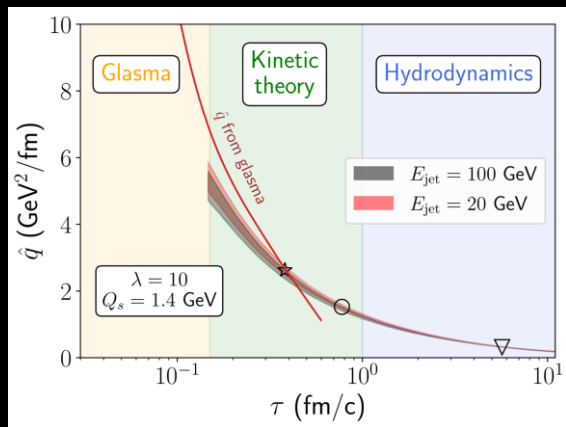


Sensitive to the nature of radiation in the early stages



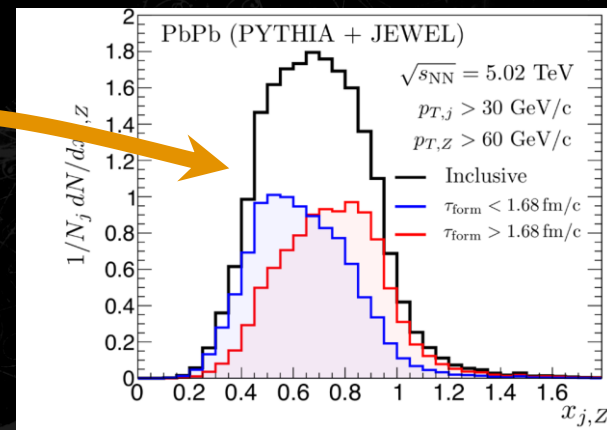
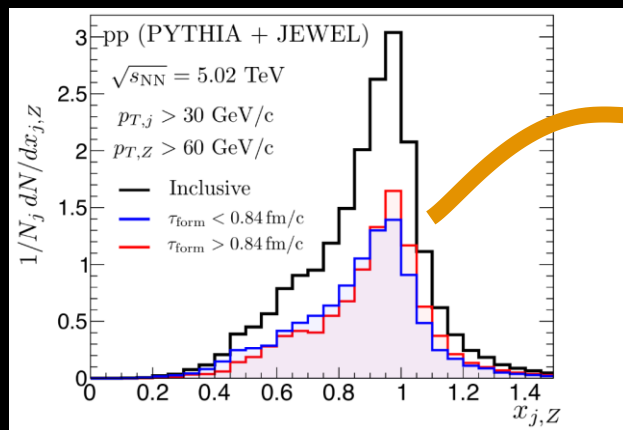
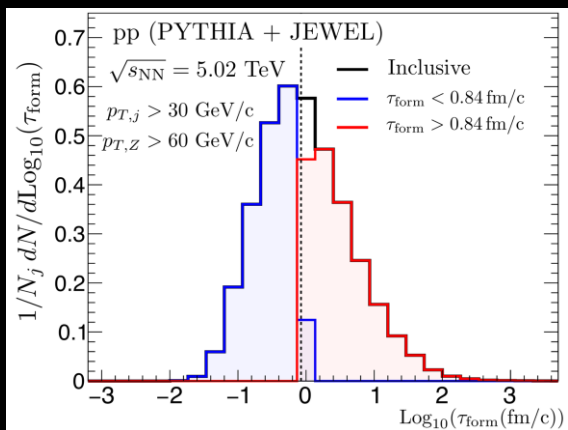
Dominated by the early stages

Large \hat{q}



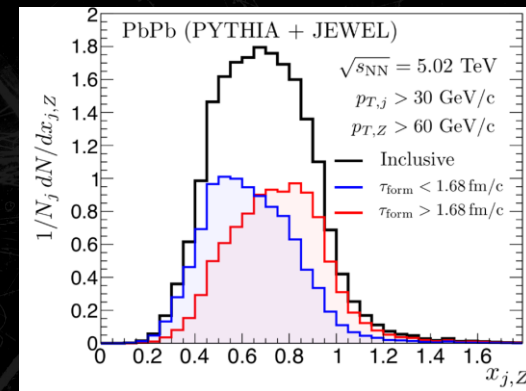
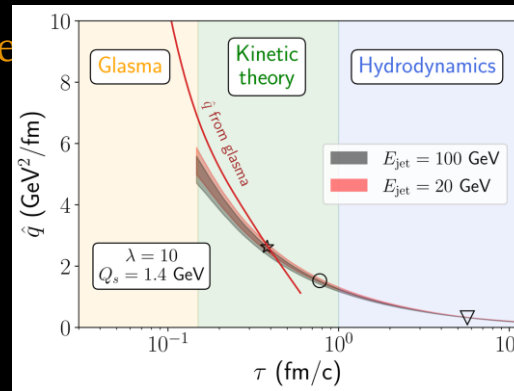
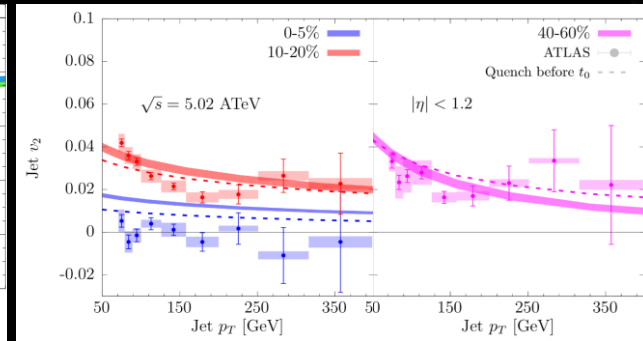
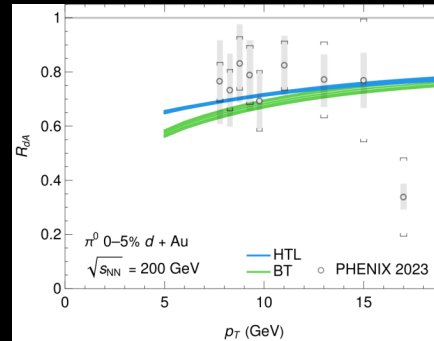
Need a space-time picture

Time reclustering: $d_{ij} = \min \left(p_{T,i}^{2p}, p_{T,j}^{2p} \right) \frac{\Delta R_{ij}^2}{R^2}$ $\xrightarrow{p=0.5}$ $p_{T,i} \theta^2 \sim \frac{1}{\tau_{form}}$



Conclusions

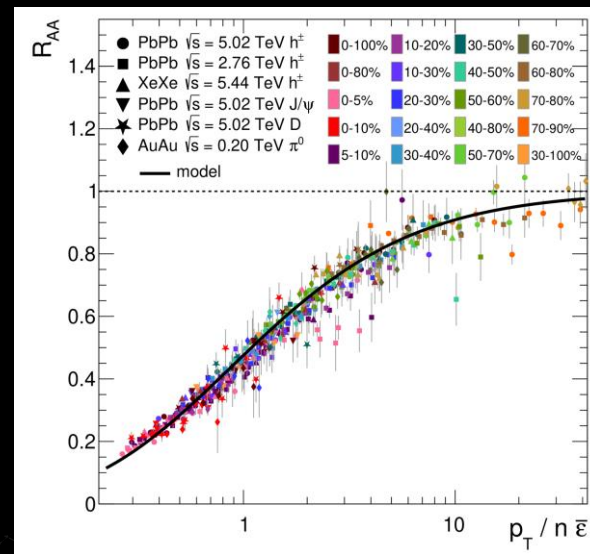
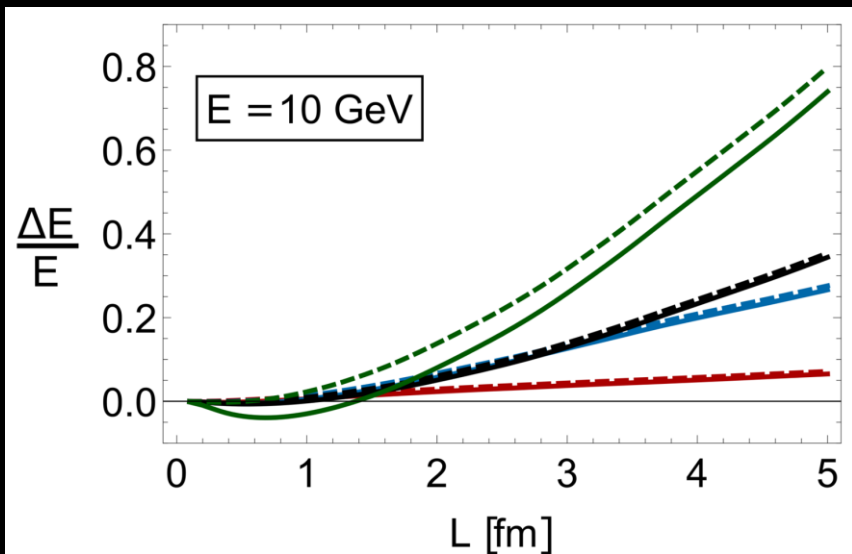
- Calculations exist
- Path-length dependence is not well known.
- Dominated by early stages
- Need a space-time picture



Thank You!

Backups

What is the pathlength dependence?

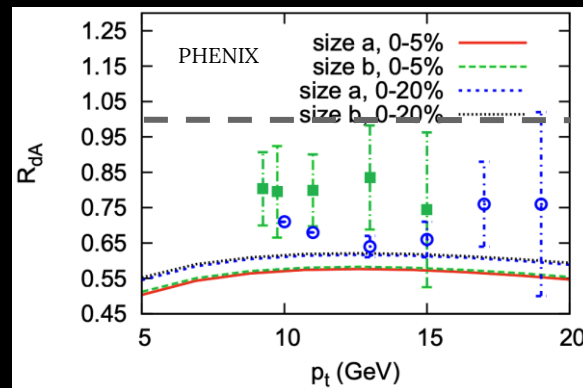
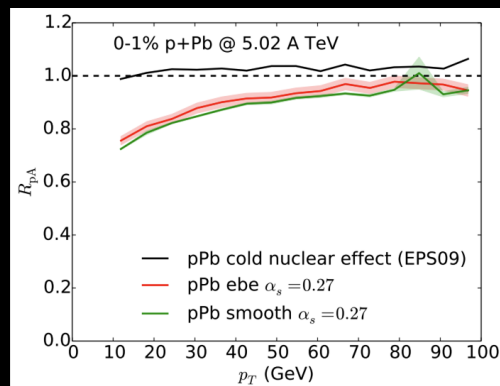


$$\langle \epsilon \rangle = L^\beta$$

$$\beta = 1.02^{+0.09}_{-0.06}$$

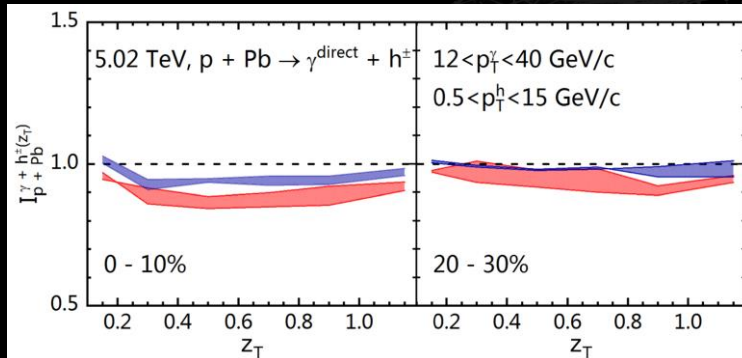
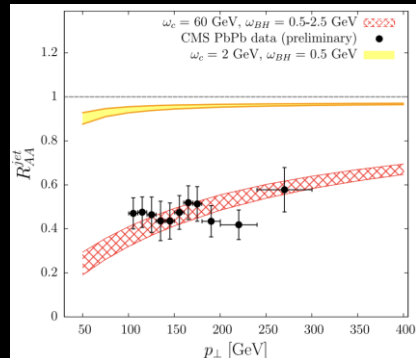
Start by varying the pathlength

Energetic partons +
MARTINI



Caveat:
Centrality
determination
is non-trivial
(see backups)

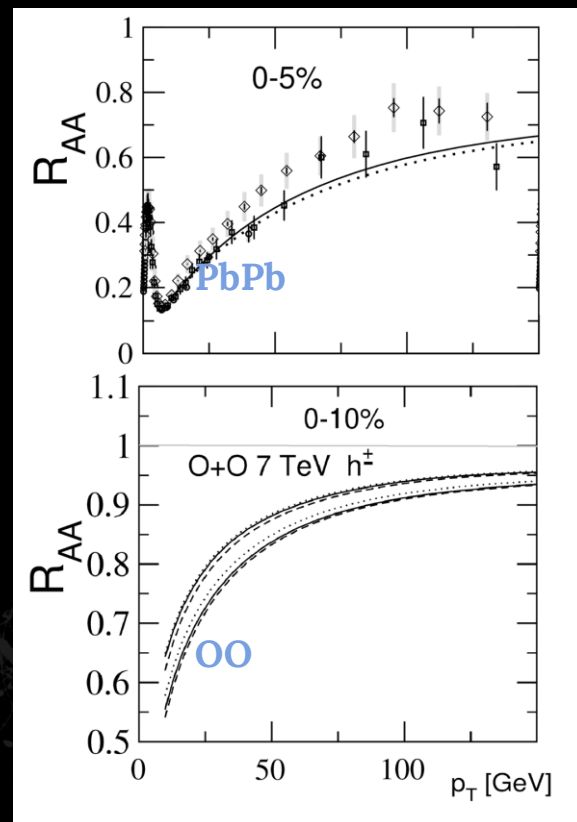
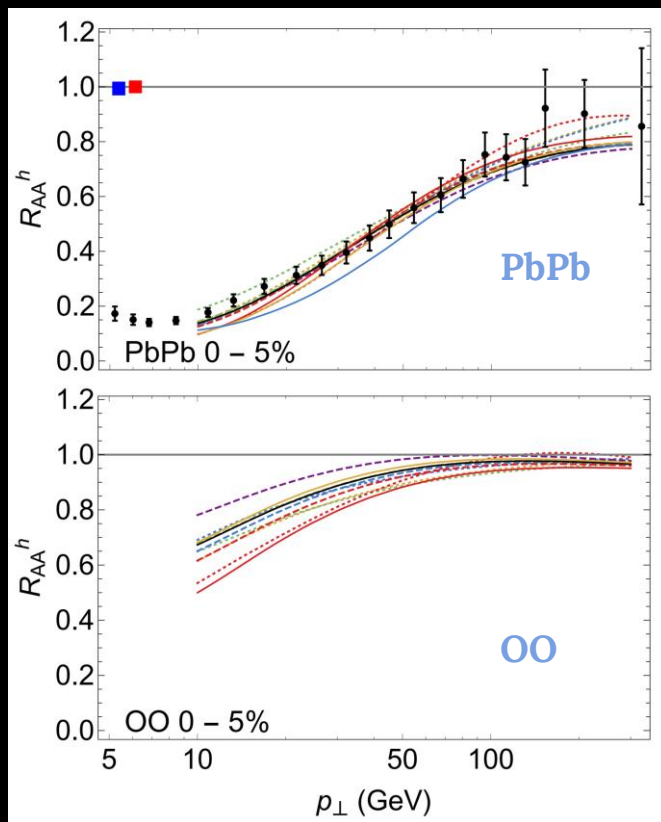
Estimate energy and
length scales,
multiple soft
scattering



Enhanced energy
loss near T_c

Higher-twist
energy loss

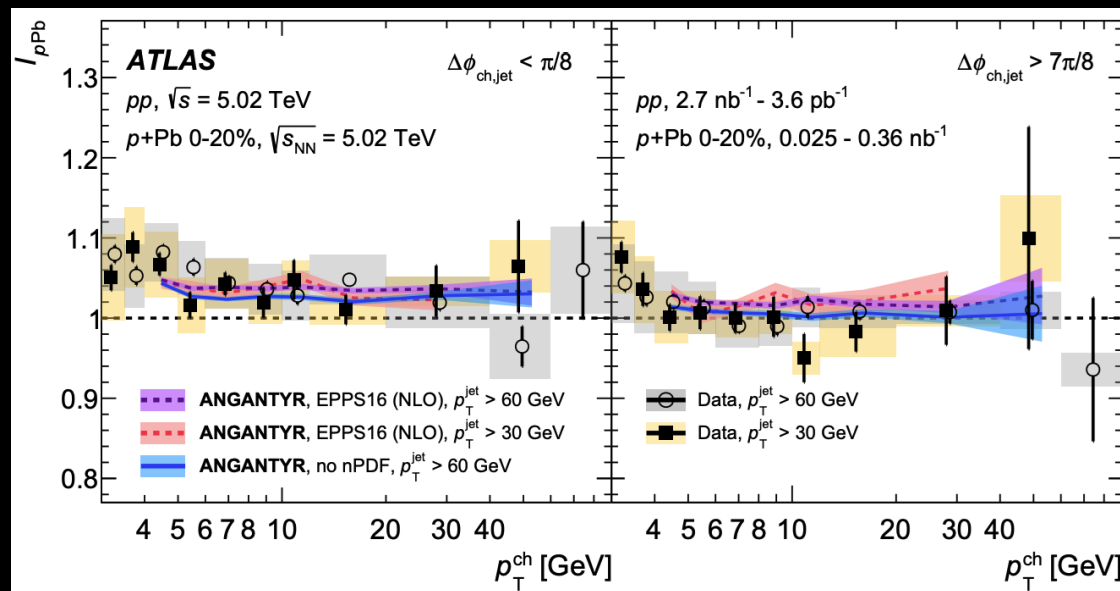
Lighter ions



No quenching?

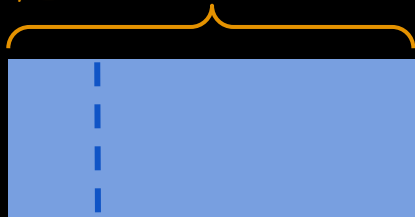
ANGANTYR with
string-shoving **OFF**

Caveat: 0-20% bin in
pPb is quantitatively
different to 0-5%



Relaxing large- L approximation

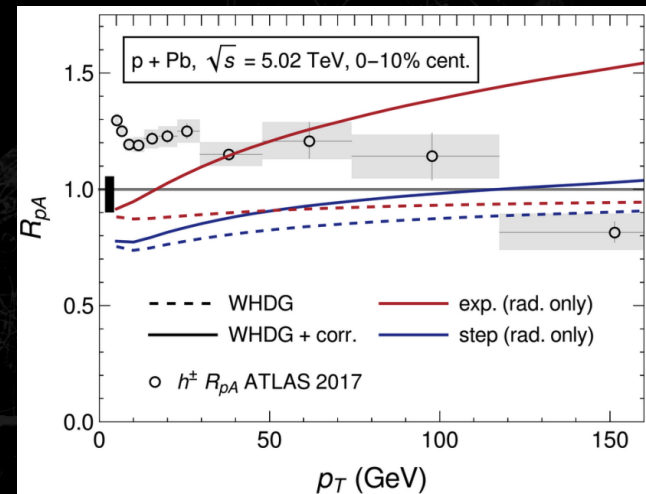
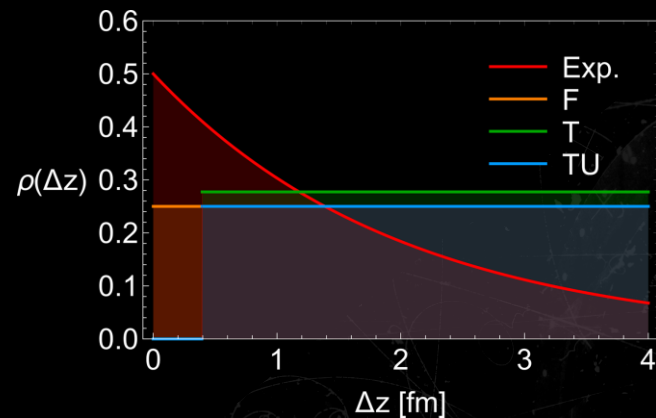
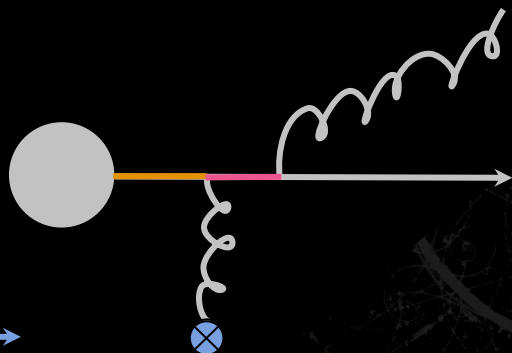
$$\frac{1}{\mu_D} \ll \Delta z \sim \lambda_{mfp} \ll L$$



$$\frac{1}{\mu_D} \ll \lambda_{mfp}$$

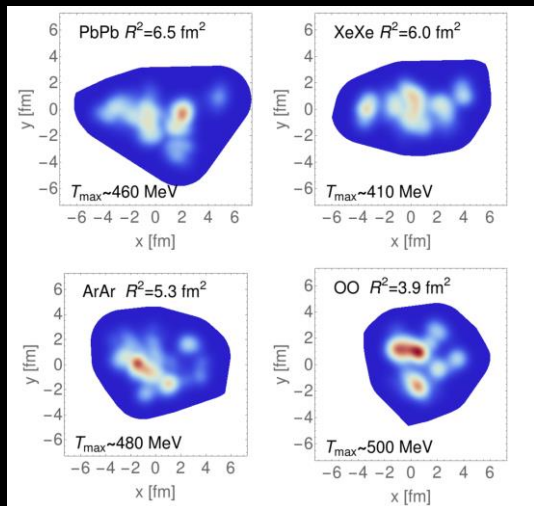
DGLV poles

corr. pole

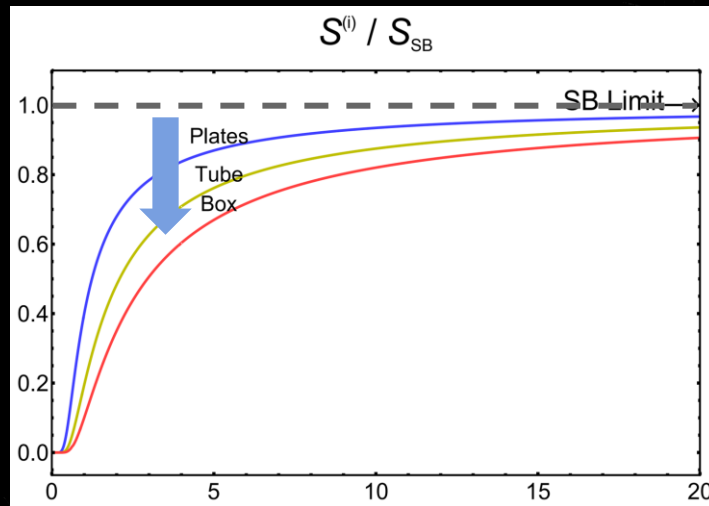


Small is not the only problem

$$\lambda_{mfp} \sim \frac{1}{\rho\sigma} \sim \frac{1}{g^2 T} \quad \mu_D \sim gT$$

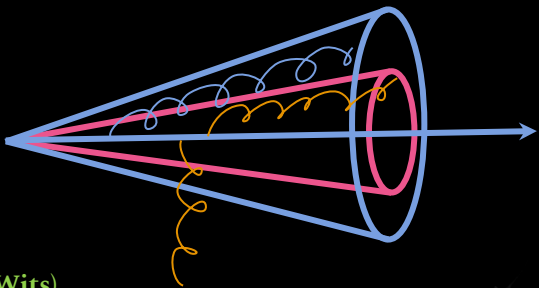
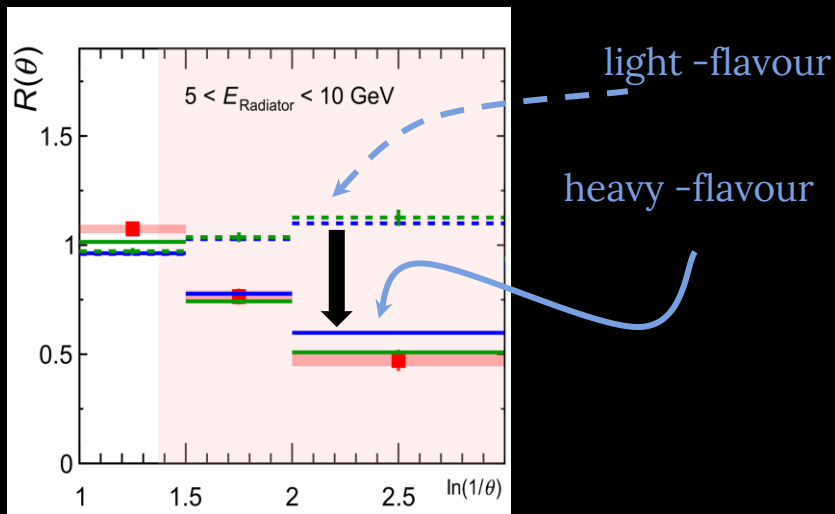


Smaller systems are hotter at the same multiplicity

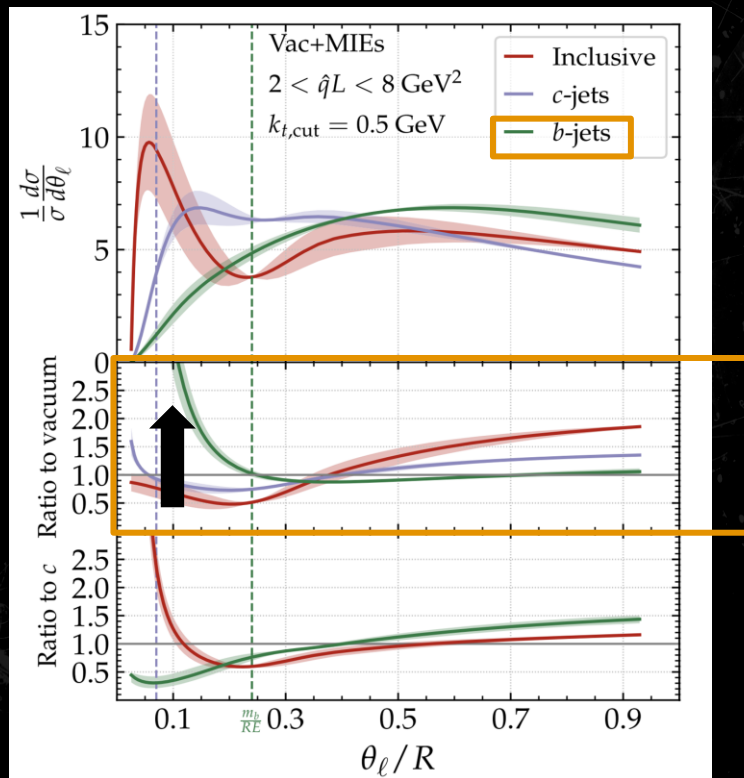


Single, massless, non-interacting, scalar field in a finite box

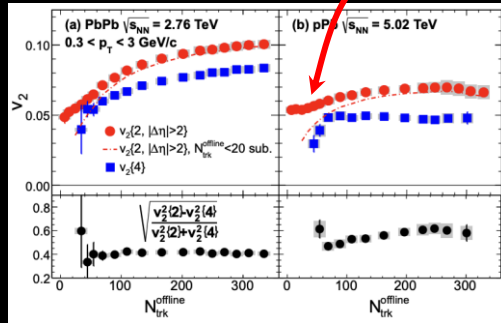
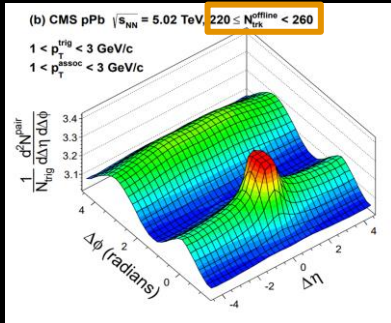
The dead cone



In-medium radiation fills the dead cone



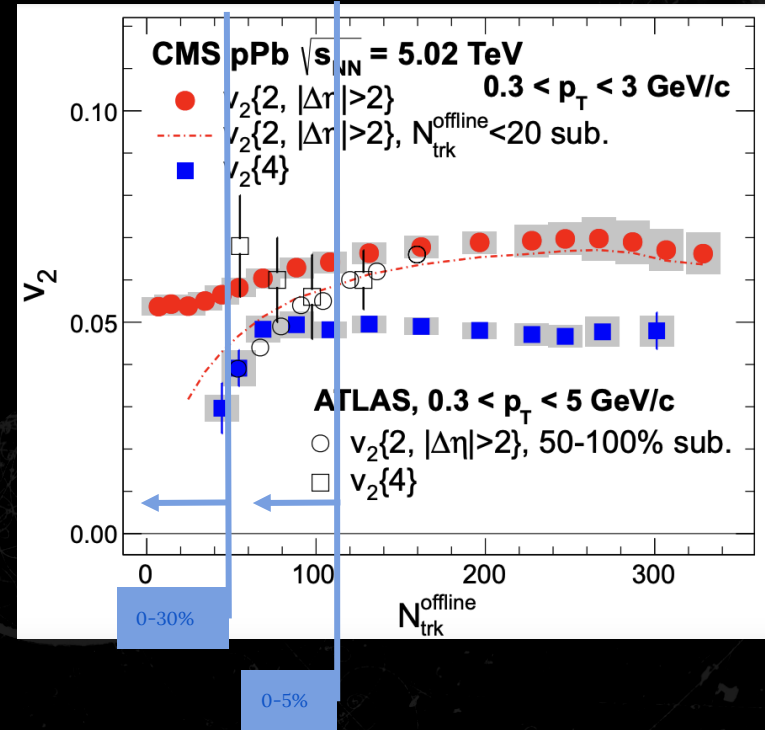
R_{AA} , v_2 , and Centrality



Subtract low mult-data (match ATLAS)

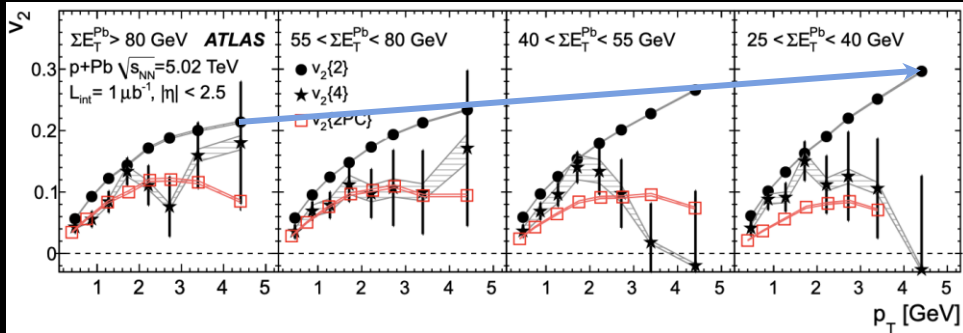
$N_{trk}^{offline}$ bin	PbPb data			pPb data		
	\langle Centrality $\rangle \pm$ RMS (%)	$\langle N_{trk}^{offline} \rangle$	$\langle N_{trk}^{corrected} \rangle$	Fraction	$\langle N_{trk}^{offline} \rangle$	$\langle N_{trk}^{corrected} \rangle$
[0, ∞)				1.00	40	50 \pm 2
[0, 20)	92 \pm 4	10	13 \pm 1	0.31	10	12 \pm 1
[20, 30)	86 \pm 4	24	30 \pm 1	0.14	25	30 \pm 1
[30, 40)	83 \pm 4	34		0.12	35	42 \pm 2
[40, 50)	80 \pm 4	44	0-50%	0.10	45	54 \pm 2
[50, 60)	78 \pm 3	54		0.09	54	66 \pm 3
[60, 80)	75 \pm 3	69	87 \pm 4	0.12	69	84 \pm 4
[80, 100)	72 \pm 3	89		0.07	89	108 \pm 5
[100, 120)	70 \pm 3	109	0-10%	0.03	109	132 \pm 6
[120, 150)	67 \pm 3	134		0.02	132	159 \pm 7
[150, 185)	64 \pm 3	167	210 \pm 9	4×10^{-3}	162	195 \pm 9
[185, 220)	62 \pm 2	202	253 \pm 11	5×10^{-4}	196	236 \pm 10
[220, 260)	59 \pm 2	239	299 \pm 13	6×10^{-5}	232	280 \pm 12
[260, 300)	57 \pm 2	279	350 \pm 15	3×10^{-6}	271	328 \pm 14
[300, 350)	55 \pm 2	324	405 \pm 18	1×10^{-7}	311	374 \pm 16

0-0.00631% bin



R_{AA} , v_2 , and Centrality (Alternative - ATLAS)

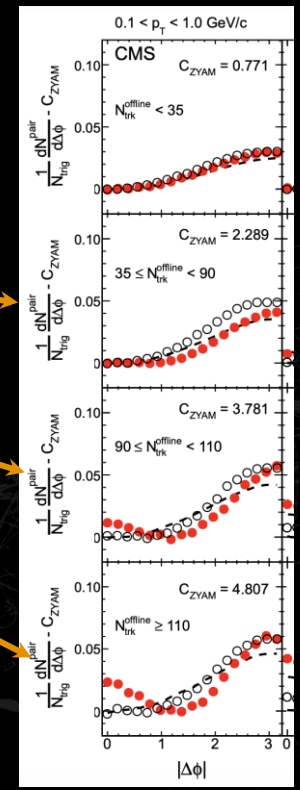
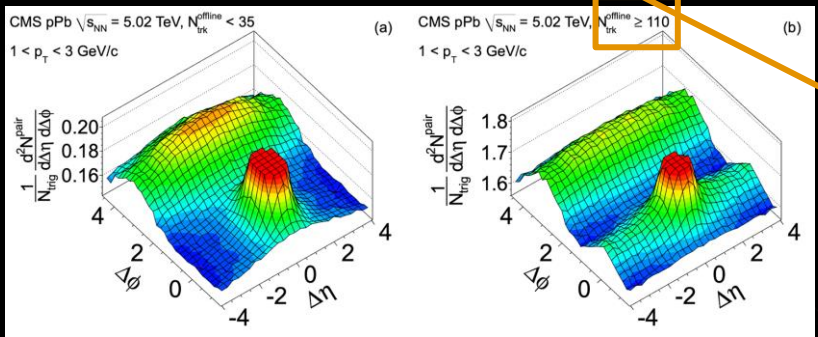
ΣE_T^{Pb} range [GeV]	$\langle \Sigma E_T^{\text{Pb}} \rangle$ [GeV]	range in fraction of events [%]	$\langle N_{\text{ch}}^{\text{rec}} \rangle$ (RMS)
> 80	93.7	0–1.9	134 (31)
55–80	64.8	1.9–9.1	102 (26)
40–55	46.7	9.1–20.0	80 (23)
25–40	31.9	20.0–39.3	60 (20)
10–25	16.9	39.3–70.4	37 (17)
< 10	4.9	70.4–100	16 (11)



R_{AA} , v_2 , and Centrality (Alternative - peripheral)

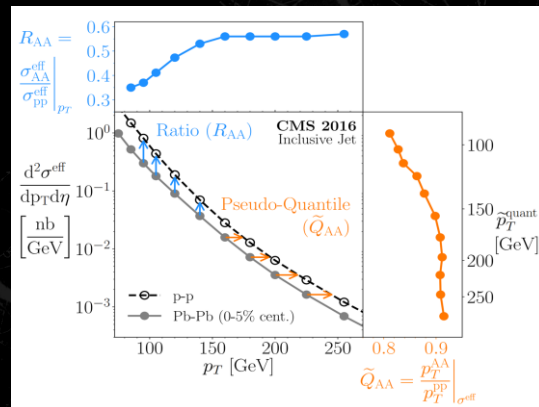
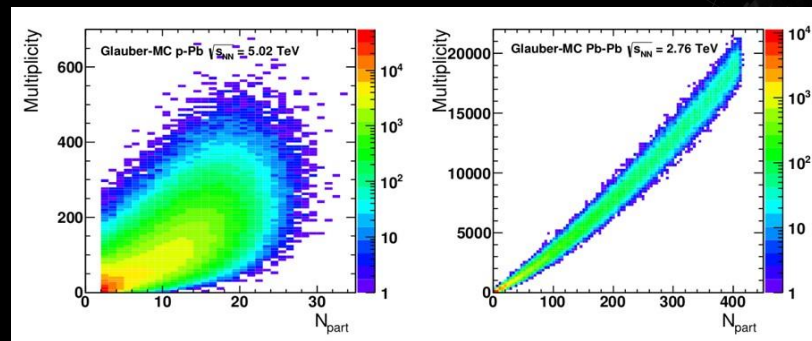
Multiplicity class ($N_{\text{trk}}^{\text{offline}}$)	Fraction (%)	$\langle N_{\text{trk}}^{\text{offline}} \rangle$	$\langle N_{\text{trk}}^{\text{corrected}} \rangle$
Minimum Bias	100.0	40.6	53.4 ± 2.9
$N_{\text{trk}}^{\text{offline}} < 35$	50.4	17.1	23.5 ± 1.3
$35 \leq N_{\text{trk}}^{\text{offline}} < 90$	41.9	56.3	75.6 ± 4.1
$90 < N_{\text{trk}}^{\text{offline}} < 110$	4.6	98.2	114.3 ± 6.2
$N_{\text{trk}}^{\text{offline}} \geq 110$	3.1	128.2	149.1 ± 8.1

Correlated yield



Why R_{AA} is not ideal for small systems

- Reliance on a reference system
- Steeply falling production spectrum
 - Sensitive only to large ΔE
 - Sensitive to PDFs and nPDFs
 - Species-dependent
- Sensitive to initial condition
 - Geometry
 - Momentum anisotropy
- Sensitive to jet fragmentation
- Supposed to quantify ΔE , but
 - $\Delta E \leftarrow L \leftarrow N_{coll}$: uncontrolled
 - $\Delta E = \Delta E(T)$: T is uncontrolled



Dead cone prediction in AA

