Field Theory and the Electroweak Standard Model

lecture 3 —

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SM on a mug

Lagrangian before SSB



Neutrinos

- Neutrinos are known to oscillate, which implies that (at least two) have (tiny) mass
- Canonical SM formulation with massless neutrinos clearly fails here
- However, simple extensions the SM that could yield neutrino masses have been proposed
 - Neutrino as a Dirac particle (Dirac spinors)
 - SU(2) singlet RH neutrinos ν_R, mimicking the construneutrino (only gravity)
 - neutrinos get mass through EW SSB, would require neutrino Yukawa couplings of (at least) 4-5 orders of magnitude smaller than electron Yukawa coupling
 - Neutrino as a Majorana particle (Majorana spinor -> Dirac spinor for particle = antiparticle)
 - Observed ν_L : LH component of a light Majorana neutrino with small mass generated by a seesaw mechanism?

→ see lectures by S. Lavignac







Abelian Higgs model:

$$\phi(x) = \frac{1}{\sqrt{2}} \left(v + \rho(x) \right) e^{i\xi(x)/v} = \frac{1}{\sqrt{2}} \left(v + \rho(x) + i\xi(x) + \dots \right)$$

Higgs Goldstone

Lagrangian in this gauge:

$$\mathscr{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \frac{1}{2}(\partial_{\mu}\rho)^{2} + \frac{1}{2}g^{2}\nu^{2}A_{\mu}A^{\mu} - \mu^{2}\rho^{2} + \frac{1}{2}g^{2}\nu^{2}A_{\mu}A^{\mu} - \frac$$

Advantage: physical fields!

Propagator of the A^{μ} field

What about other gauges?

Higgs Unitary gauge $\phi(x) \to \phi'(x) = e^{(-i\xi(x)/\nu)}\phi(x) = \frac{1}{\sqrt{2}} \left(\nu + \rho(x)\right)$ $A_{\mu}(x) \rightarrow A'_{\mu}(x) = A_{\mu}(x) - \frac{1}{gv} \partial_{\mu} \xi(x)$ $+\frac{1}{2}g^{2}A\mu A^{\mu}\rho^{2} + g^{2}vA_{\mu}A^{\mu}\rho - \lambda v\mu\rho^{3} - \frac{\lambda}{4}\rho^{4} + \frac{1}{4}\mu^{2}v^{2}$ $\frac{i\left(-g_{\mu\nu}+k_{\mu}k_{\nu}/m_{A}^{2}\right)}{k^{2}-m_{A}^{2}+i\epsilon}$ behaves as $\frac{1}{m_A^2}$ as $k \to \infty$; compare with $\frac{1}{k^2}$ for massless photon



Abelian Higgs model:

$$\phi(x) = \frac{1}{\sqrt{2}} \left(v + \rho(x) \right) e^{i\pi(x)/v} = \frac{1}{\sqrt{2}} \left(v + \rho(x) + i\pi(x) + \dots \right)$$

$$(D_{\mu}\phi)^{*}(D^{\mu}\phi) = \frac{1}{2}(\partial_{\mu}\rho)^{2} + \frac{1}{2}(\partial_{\mu}\pi)^{2} + \frac{1}{2}g^{2}v^{2}A_{\mu}A^{\mu}$$

Eliminate the kinetic mixing term $gv A_{\mu}\partial^{\mu}\pi$ by gauge fixing ($M_A = gv$)

Propagator of the A^{μ} field

 \sim

Propagator of the π field

 $^{\mu} + gvA_{\mu}\partial^{\mu}\pi + \text{interaction terms}$

 $\mathscr{L}_{GF} = -\frac{1}{2\zeta} (\partial^{\mu} A^{a}_{\mu} - \xi M_{A} \pi)^{2}$ $R_{\xi} \text{ gauges}$ (manifestly renormalizable) $\frac{i \left(-g_{\mu\nu} + (1-\xi)k_{\mu}k_{\nu}/(k^{2}-\xi m_{A}^{2})\right)}{k^{2}-m_{A}^{2}+i\epsilon} \text{ behaves as } \frac{1}{k^{2}} \text{ as } k \to \infty$ $\frac{i}{k^{2}-\xi m_{A}^{2}+i\epsilon}$





Propagator of the A^{μ} field

Propagator of the π field

$$\frac{1}{\zeta} (\partial^{\mu} A^{a}_{\mu} - \xi M_{A} \pi)^{2}$$

$$R_{\xi} \text{ gauges}$$
(manifestly renormaliza)

$$\frac{i\left(-g_{\mu\nu} + (1-\xi)k_{\mu}k_{\nu}/(k^2 - \xi m_A^2)\right)}{k^2 - m_A^2 + i\epsilon}$$

behaves as
$$\frac{1}{k^2}$$
 as $k \rightarrow$

$$\frac{\iota}{k^2 - \xi m_A^2 + i\epsilon}$$





Propagator of the A^{μ} field

 \sim

Propagator of the π field

itary gauge
eynman gauge
ar any other
$$R_{\xi}$$
)
tree level
loops
$$\frac{\zeta}{\zeta} (\partial^{\mu}A^{a}_{\mu} - \xi M_{A}\pi)^{2} R_{\xi} \text{ gauges} \\
\text{manifestly renormaliz} \\
\frac{i\left(-g_{\mu\nu} + (1-\xi)k_{\mu}k_{\nu}/(k^{2}-\xi m_{A}^{2})\right)}{k^{2}-m_{A}^{2}+i\epsilon} \text{ behaves as } \frac{1}{k^{2}} \text{ as } k \rightarrow \frac{i}{k^{2}-\xi m_{A}^{2}+i\epsilon}$$





Analogous treatment for non-abelian gauge theories, b apart from unphysical Goldstone bosons also unphysic ghosts introduced

itary gauge
tree level
cynman gauge
r any other
$$R_{\xi}$$
)
 $i = \frac{i\left(-g_{\mu\nu} + (1-\xi)k_{\mu}k_{\nu}/(k^2 - \xi m_A^2)\right)}{k^2 - m_A^2 + i\epsilon}$
 $i = \frac{i\left(-g_{\mu\nu} + (1-\xi)k_{\mu}k_{\nu}/(k^2 - \xi m_A^2)\right)}{k^2 - m_A^2 + i\epsilon}$
behaves as $\frac{1}{k^2}$ as $k \to \frac{1}{k^2}$

 $k^2 - \xi m_A^2 + i\epsilon$



Counting parameters

- The Standard Model has 18 free parameters*
 - * 3 couplings g, g', g_s
 - * 2 parameters of the Higgs potential μ , λ
 - 9 (6 quark + 3 lepton) Yukawa couplings y_f
 - ✤ 4 parameters of the CKM matrix V_{CKM}

*though recollections may vary (one can additionally include non-canonical parameters, e.g. the Θ_{QCD} CP-violating angle or parameters of the neutrino sector bringing it up to 26 free parameters)



Counting parameters

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 Using relations between the parameters a more practical list can be obtained, e.g.:

* 2 coupling constants $\alpha = \frac{e^2}{4\pi}$ and $\alpha_s = \frac{g_s^2}{4\pi}$

Fermi constant
$$G_F = \frac{g^2}{4\sqrt{2}M_W^2}$$

- * 2 masses M_Z and M_H
- 9 fermion masses m_f
- ✤ 4 parameters of the CKM matrix V_{CKM}



SM input parameters



https://pdg.lbl.gov

- * All parameters of the SM have been experimentally measured (last unmeasured was M_H)
- Consult Particle Data Group for the most up-to-date values

*
$$\alpha^{-1} = 137.035999084(21)$$

 $\alpha_s(M_Z) = 0.1179(9)$
 $G_F = 1.1663788(6) \times 10^{-5} \text{ GeV}^{-1}$

* $M_Z = 91.1876 \pm 0.0021 \text{ GeV}$ $M_H = 125.25 \pm 0.17 \text{ GeV}$

*

 With input parameters known, one can make theoretical predictions for any SM observable!



The trouble with theoretical predictions

- The running of couplings \rightarrow depending on the strength of the coupling different methods applicable *
 - * expansion in the coupling constant(s) in the perturbative regime
 - Iattice gauge theory (e.g.) in the non-perturbative regime
- Feasibility of performing calculations in the sense of obtaining finite results (renormalisability) \rightarrow depends on the theory *
 - Problem (in perturbation theory): calculations of quantum loop corrections involve integration over unconstrained momenta of the virtual particle(s) in the loop(s)



The integration can yield \rightarrow UV singularities when $k \rightarrow \infty$ IR singularities when $k \rightarrow 0$



treated by the renormalisation program cancel against IR singularities from real emission diagrams



UV singularities and how to cure them

- 1. Acceptance. UV singularities can appear in the intermediate stages of calculations.

Most often used: dimensional regularisation (DR).

Integrals are calculated in $d \neq 4$ dimensions.

 $g^2 \rightarrow g^2 \mu^{4-d}$ to keep the action dimensionless, μ is an arbitrary scale

Singularities manifest as poles $\frac{1}{(4-d)^n}$

* DR preserves gauge symmetry and can be used for regularisation of both, UV and IR singularities.

to yield well-defined results for all observables at any order of perturbation theory.

2. Diagnosis. Infinities are given mathematical meaning. Achieved by introducing a regulator parameter in the expressions for the loops. The regulator makes the integrals well-defined, apart from a limit value of the regulator, for which the integral is singular. → **Regularisation**

$$\int d^4k \to \int d^dk$$

3. Treatment. The singularities are absorbed in the redefined parameters and fields of the theory. A finite number of redefinitions has \rightarrow Renormalisation



Example



Dimensional regularisation $(d = 4 - 2\epsilon)$

$$\int \frac{d^d k}{k^4} \to \int d\Omega_d \int_K^\infty dk \frac{k^{d-1}}{k^4} = \frac{(2\pi)^{2-\epsilon}}{\Gamma(2-\epsilon)} \left(-\frac{1}{2\epsilon}\right) \left[\frac{1}{k^{2\epsilon}}\Big|_{k=\infty} - \frac{1}{k^{2\epsilon}}\Big|_{k=K}\right] \propto \frac{1}{2\epsilon} - \log(K) + \mathcal{O}(\epsilon)$$

$$\frac{\text{Tr}[\dots]}{p((p-k)^2 - m^2)} \sim \int \frac{d^4k}{k^4} \sim \int \frac{dk}{k} \qquad \text{logarithmic singular}$$

$$\text{large } k$$



QED renormalisation

• Reinterpret the QED Lagrangian $\mathscr{L} = -\frac{1}{4}(\partial_{\mu}A_{\nu}$ as written in terms of bare (unrenormalised) parameters e_0 , m_0 and fields ψ^0 , A^0 and relate

$$\psi^0 = \sqrt{Z_2} \, \psi^R \qquad A^0_\mu = \sqrt{Z_3} \, A^R_\mu \qquad m_0 = Z_m m_R \qquad e_0 = Z_e e_R$$

• With $Z_1 \equiv Z_e Z_2 \sqrt{Z_3}$ and $Z_i = 1 + (Z_i - 1)$ obtain Lagrangian of the renormalised perturbation theory

$$\mathscr{L} = -\frac{1}{4} (\partial_{\mu}A_{\nu}^{R} - \partial_{\nu}A_{\mu}^{R})^{2} + \bar{\psi}^{R} \left(i\gamma^{\mu}\partial_{\mu} - e_{R}\gamma^{\mu}A_{\mu}^{R} - m_{R} \right) \psi^{R} - \frac{1}{4} (Z_{3} - 1)(\partial_{\mu}A_{\nu}^{R} - \partial_{\nu}A_{\mu}^{R})^{2} + i(Z_{2} - 1)\bar{\psi}^{R}\gamma^{\mu}\partial_{\mu}\psi^{R} - e_{R}(Z_{1} - 1)\bar{\psi}^{R}\gamma^{\mu}A_{\mu} - \int \left[(Z_{m} - 1) + (Z_{2} - 1) \right] m_{R}\bar{\psi}^{R} \psi^{R} - \frac{1}{4} (Z_{3} - 1)(\partial_{\mu}A_{\nu}^{R} - \partial_{\nu}A_{\mu}^{R})^{2} + i(Z_{2} - 1)\bar{\psi}^{R}\gamma^{\mu}\partial_{\mu}\psi^{R} - e_{R}(Z_{1} - 1)\bar{\psi}^{R}\gamma^{\mu}A_{\mu} - \frac{1}{4} (Z_{3} - 1)(\partial_{\mu}A_{\nu}^{R} - \partial_{\nu}A_{\mu}^{R})^{2} + i(Z_{2} - 1)\bar{\psi}^{R}\gamma^{\mu}\partial_{\mu}\psi^{R} - e_{R}(Z_{1} - 1)\bar{\psi}^{R}\gamma^{\mu}A_{\mu} - \frac{1}{4} (Z_{3} - 1)(\partial_{\mu}A_{\nu}^{R} - \partial_{\nu}A_{\mu}^{R})^{2} + i(Z_{2} - 1)\bar{\psi}^{R}\gamma^{\mu}\partial_{\mu}\psi^{R} - e_{R}(Z_{1} - 1)\bar{\psi}^{R}\gamma^{\mu}A_{\mu} - \frac{1}{4} (Z_{3} - 1)(\partial_{\mu}A_{\nu}^{R} - \partial_{\nu}A_{\mu}^{R})^{2} + i(Z_{2} - 1)\bar{\psi}^{R}\gamma^{\mu}A_{\mu} - \frac{1}{4} (Z_{3} - 1)(\partial_{\mu}A_{\nu}^{R} - \partial_{\nu}A_{\mu}^{R})^{2} + i(Z_{2} - 1)\bar{\psi}^{R}\gamma^{\mu}A_{\mu} - \frac{1}{4} (Z_{3} - 1)(\partial_{\mu}A_{\nu}^{R} - \partial_{\nu}A_{\mu}^{R})^{2} + i(Z_{2} - 1)\bar{\psi}^{R}\gamma^{\mu}A_{\mu} - \frac{1}{4} (Z_{3} - 1)(\partial_{\mu}A_{\nu}^{R} - \partial_{\nu}A_{\mu}^{R})^{2} + i(Z_{2} - 1)\bar{\psi}^{R}\gamma^{\mu}A_{\mu} - \frac{1}{4} (Z_{3} - 1)(\partial_{\mu}A_{\nu}^{R} - \partial_{\nu}A_{\mu}^{R})^{2} + i(Z_{3} - 1)(\partial_{\mu}A_{\nu}^{R} - \partial_{\mu}A_{\mu}^{R})^{2} + i(Z_{3} - 1)(\partial_{\mu}A_{\nu}^{R})^{2} + i(Z_{3} - 1)(\partial_{\mu}A_{\mu}^{R} - \partial_{\mu}A_{\mu}^{R})^{2} + i(Z_{3} - 1)(\partial_{\mu}A_{\mu}^{R})^{2} + i(Z_{3} - 1)(\partial_{\mu}A_{\mu}^{R} - \partial_{\mu}A_{\mu}^{R})^{2} + i(Z_{3} - 1)(\partial_{\mu}A_{\mu}^{R})^{2} + i(Z_{3} - 1)(\partial_{\mu}A$$

 \rightarrow additional Feynman rules for the counterterms, e.g.

* Fix renormalisation conditions defining Z_i 's

$$-\partial_{\nu}A_{\mu})^{2} + \bar{\psi}\left(i\gamma^{\mu}\partial_{\mu} - e\,\gamma^{\mu}A_{\mu} - m\right)\psi$$

renormalisation constants Z_i

counterterms

 \sim

$$-i(Z_3 - 1)(p^2 g^{\mu\nu} - p^{\mu\nu})$$



Example ctnd.

$$\mu \sim (p^{2}g^{\mu\nu} - p^{\mu}p^{\nu}e^{2}) e^{2} \Pi_{2}(p^{2})$$

Counterterm contribution



Renormalisation condition

 $\Pi(0) = 0$

Up to order e^2

$$\Pi(p^2) = e^2 \Pi_2(p^2) + (Z_3 - 1) + \dots$$

$$\Pi(p^2) = e^2 \Pi_2(p^2) + (Z_3 - 1) + \dots \qquad \text{hence} \qquad Z_3 - 1 = -\frac{e^2}{6\pi^2} \left[\frac{1}{e} + \frac{1}{2} \log\left(\frac{\tilde{\mu}^2}{m^2}\right) \right]$$

$$\Rightarrow \ \Pi(p^2) = \frac{e^2}{2\pi^2} \int_0^1 dx \, x \, (1 - x) \log\left(\frac{m^2}{m^2 - p^2 x \, (1 - x)}\right) + \dots \qquad \text{finite result!}$$



$$\Pi_{2}(p^{2}) = \frac{1}{2\pi^{2}} \int_{0}^{1} dx \, x \, (1-x) \left[\frac{2}{\epsilon} + \log\left(\frac{\tilde{\mu}^{2}}{m^{2} - p^{2}x \, (1-x)}\right) \right]_{0}^{2} = 4\pi \epsilon$$
$$\tilde{\mu}^{2} = 4\pi \epsilon$$
$$-i(Z_{3} - 1) \, (p^{2}g^{\mu\nu} - p^{\mu}p^{\nu})$$

where $-i(p^2g^{\mu\nu} - p^{\mu}p^{\nu}) \Pi(p^2)$ is the sum of all 1PI contributions to the photon 2-point function



Renormalisability (1)

- Renormalisable theory: all UV divergencies can be cancelled with a finite number of counterterms to any order in * perturbation theory
 - QED renormalisation program at 1 loop:



- *
- that can be in principle calculated in perturbation theory
- * Renormalisation procedure introduces renormalisation scale *μ* as an artefact of the regularisation prescription





4 renormalisation conditions

Normalisation of the fields not observables, so can be rescaled. Normalisation of the parameters set by measured quantities. The principle of renormalisability is an essential condition for any viable physical theory \rightarrow observables are finite functions



Renormalisability (2)

theories with SSB was delivered a few years later, in 1971, by t'Hooft and Veltman

•

* Although Weinberg and Salam speculated that their theory is renormalizable, a proof of renormalizability of Yang-Mills



(Nobel Prize 1999)



Renormalisability (2)

- theories with SSB was delivered a few years later, in 1971, by t'Hooft and Veltman
- Renormalised QED has proven to be spectacularly successful
 - anomalous magnetic moment of the electron

calculated up to 5 loops in QED

 $a_{e}^{\exp} = 1159652.18059(13) \times 10^{-9}$ agrees with the SM to 1 part in 10^{12}

Although Weinberg and Salam speculated that their theory is renormalizable, a proof of renormalizability of Yang-Mills

(Nobel Prize 1999)

[Fan et al. Phys.Rev.Lett. 130 (2023) 7]

Renormalisability (2)

- theories with SSB was delivered a few years later, in 1971, by t'Hooft and Veltman
- Renormalised QED has proven to be spectacularly successful
 - anomalous magnetic moment of the muon

calculated up to 5 loops in QED

$$a_{\mu}^{\text{QED}} = 1165847.1893(10) \times 10^{-9}$$

 $a_{\mu}^{\text{SM}} = 1165918.100(430) \times 10^{-9}$

Although Weinberg and Salam speculated that their theory is renormalizable, a proof of renormalizability of Yang-Mills

(Nobel Prize 1999)

 $a_{\mu}^{\exp} = 1165920.591(221) \times 10^{-9}$ (world average 08'23)

Running coupling

consequence: running coupling. Not being an observable, the coupling can depend on this scale.

* The renormalisation process and the measurable quantities must be independent of the arbitrary renormalisation scale \rightarrow

Running coupling

- consequence: running coupling. Not being an observable, the coupling can depend on this scale.
- then from dimensional analysis after renormalisation $R = R(Q^2/\mu^2, \alpha(\mu^2))$
- Independence of *R* on μ implies *

$$\mu^2 \frac{d}{d\mu^2} R(Q^2/\mu^2, \alpha(\mu^2)) =$$

With
$$\beta(\alpha) \equiv \mu^2 \frac{\partial \alpha}{\partial \mu^2}$$
 $t \equiv \log\left(\frac{Q^2}{\mu^2}\right)$ $\alpha_\mu \equiv \alpha(\mu^2)$

Consider a dimensionless observable R, e.g. a ratio of two cross sections. If only one scale Q is relevant for this observable, $\alpha = g^2/(4\pi)$

$$= \left[\mu^2 \frac{\partial}{\partial \mu^2} + \mu^2 \frac{\partial \alpha}{\partial \mu^2} \frac{\partial}{\partial \alpha} \right] R(Q^2/\mu^2, \alpha(\mu^2)) = 0$$

renormalisation group equation

$$\left[-\frac{\partial}{\partial t} + \beta(\alpha_{\mu})\frac{\partial}{\partial\alpha_{\mu}}\right]R(e^{t},\alpha_{\mu}) = 0$$

Running coupling (2)

To solve

 $\left| -\frac{\partial}{\partial t} + \beta(\alpha_{\mu}) \frac{\partial}{\partial \alpha_{\mu}} \right|$

define (in an implicit way) a new function, the running coupling $\alpha(Q^2)$, through the equation $t = \int_{\alpha}^{\alpha(Q^2)} \frac{d\alpha}{\beta(\alpha)}$

By differentiating this equation wrt. t at fixed α_{μ} and wrt. α_{μ} at fixed t one finds $R(1, \alpha(Q^2))$ is a solution of (*)

Using $\beta(\alpha) = -\alpha^2 b_0 + \dots$ integration of (**) gives

$$t = \frac{1}{b_0} \left(\frac{1}{\alpha(Q^2)} - \frac{1}{\alpha_{\mu}} \right)$$

$$-\int_{u} R(e^{t}, \alpha_{\mu}) = 0$$

(**)

(*)

 \rightarrow entire scale dependence in *R* enters through $\alpha(Q^2)$

$$\Rightarrow \alpha(Q^2) = \frac{\alpha(\mu^2)}{1 + \alpha(\mu^2)b_0\log(Q^2/\mu^2)}$$

In QCD

Running of the SM parameters

SM gauge couplings

Hint of unification?

together with Higgs and Yukawa couplings

arXiv:1601.08143

Anomalies

- Anomalies appear when a symmetry of a classical system is not preserved at the quantum level
- Noether's theorem: continuous symmetries imply conserved currents *
- Currents associated with anomalous symmetries are not conserved \rightarrow in QFT gauged symmetries must be anomaly free \rightarrow * consistency check (global anomalies not problematic)
- In the SM, axial currents $J^A_\mu = \bar{\psi}\gamma_\mu\gamma_5 T^a\psi$ receive non-zero corrections to $\partial^\mu J^A_\mu$ from one-loop triangle diagrams •

In the SM only U(1)-SU(3)-SU(3), U(1)-SU(2)-SU(2), U(1)-U(1)-U(1) anomalies contribute

Anomalies (2)

$$\nu_L$$
 e_L
 e_R
 u_L
 d_L
 u_R
 d_R

 Y
 -1
 -1
 -2
 1/3
 1/3
 4/3
 -2/3

 Tr[$T^a\{T^b, T^c\}$]_L - Tr[$T^a\{T^b, T^c\}$]_R

✤ U(1)-SU(3)-SU(3), only quarks

U(1)-SU(2)-SU(2), only doublets

U(1)-U(1)-U(1) all fermions

 $3\left[2\left(\frac{1}{3}\right) + \frac{2}{3} - \frac{4}{3}\right] = 0$ $2\left[-1 + 3\left(\frac{1}{3}\right)\right] = 0$ $\left[2\left(-1\right)^{3} - \left(-2\right)^{3} + 6\left(\frac{1}{3}\right)\right]$

- number of particles and their hypercharges conspire to cancel
- works within 1 generation
- anomaly cancellation provides a strong restriction on potential new particles

$$2)^{3} + 6\left(\frac{1}{3}\right)^{3} - 3\left(-\frac{2}{3}\right)^{3} - 3\left(\frac{4}{3}\right)^{3}\right] = 0$$

SM@LHC(1)

SM@LHC(2)

				CMS prelim
_	inelastic	7 TeV	Phys. Lett. B 722 (2013) 5	
6CI	inelastic	13 TeV	JHEP 07 (2018) 161	
	jet	/ iev	FRD 90 (2014) 072000	
	γ	7 TeV	PRD 84 052011 (2011)	
	W	2.76 TeV	PLB 715 (2012) 66	
	w	5.02 lev 7 TeV	IHEP 10 (2011) 132	
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row	W	13 TeV	SMP-20-004	
ecti	Z	2.76 lev 5.02 TeV	JHEP 03 (2015) 022 SMP-20-004	
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	WZ WZ	8 TeV	EPJC 77 (2017) 236 IHEP 07 (2022) 032	
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	zz	7 TeV	JHEP 01 (2013) 063	
	ZZ	8 TeV	PLB 740 (2015) 250	
	22	12 164	LFJC 81 (2021) 200	
		13 TeV 13 TeV	PRL 125 151802 (2020)	
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1050	ZZZ	13 TeV	PRL 125 151802 (2020)	
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	VBF W VBF W	8 TeV 13 TeV	JHEP 11 (2016) 147 FPIC 80 (2020) 43	
	VBF Z	7 TeV	JHEP 10 (2013) 101	
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		13 TeV	EPJC 78 (2018) 589	
/BS	ex. $\gamma\gamma \rightarrow WW$	8 TeV	JHEP 08 (2016) 119	
∧ pc	EW qqWγ	8 TeV	JHEP 06 (2017) 106	
Far	EW qqWγ	13 TeV	PRD 108 032017	
٨B	EW os WW EW ss WW	13 lev 8 TeV	PLB 841 (2023) 137495 PRL 114 051801 (2015)	σ
	EW ss WW	13 TeV	PLB 809 (2020) 135710	
	EW qqΖγ	8 TeV	PLB 770 (2017) 380	<i>σ</i> (EW c
	EW qqZγ EW qqWZ	13 TeV	PRD 104 072001 (2021) PLB 809 (2020) 135710	σ(EW α
	EW qqWZ EW qqZZ	13 TeV	PLB 812 (2020) 135910	$\sigma(\text{EW qqZZ}) = 0.33$
	tt	5.02 TeV	JHEP 04 (2022) 144	
	tt ++	7 TeV	JHEP 08 (2016) 029	
	u tt	13 TeV	PRD 104 (2021) 092013	
	tt	13.6 TeV	Submitted to JHEP	
	t _{t - ch}	7 TeV	JHEP 12 (2012) 035	
	t - ch $t_{t - ch}$	o iev 13 TeV	190 (2014) אוונ (2014) PLB 72 (2017) 752	
	tW	7 TeV	PRL 110 (2013) 022003	
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•	tW t _{s-ch}	13 IeV 8 TeV	JHEP 10 (2018) 117 JHEP 09 (2016) 027	
Top	ttγ	8 TeV	JHEP 10 (2017) 006	
	ttγ	13 TeV	JHEP 05 (2022) 091	
	tZq tZa	8 TeV	JHEP 07 (2017) 003	
	ttZ	7 TeV	PRL 110 (2013) 172002	
	ttZ	8 TeV	JHEP 01 (2016) 096	
	ttZ	13 TeV	JHEP 03 (2020) 056	
	ιγ ttW	נע ieV 8 TeV	FRE 121 221802 (2018) JHEP 01 (2016) 096	
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		7 7-11		
	ggH ggH	7 ieV 8 TeV	ЕРЈС 75 (2015) 212 ЕРЈС 75 (2015) 212	
	ggH	13 TeV	Nature 607 60-68 (2022)	
	VBF qqH VBF qqH	7 TeV	EPJC 75 (2015) 212	
v	VBF qqH	13 TeV	Nature 607 60-68 (2022)	
ligg	VH	8 TeV	EPJC 75 (2015) 212	
-	WH 74	13 TeV	Nature 607 60-68 (2022)	
	∠⊓ ttH	8 TeV	EPJC 75 (2015) 212	
	ttH	13 TeV	Nature 607 60-68 (2022)	
	+H		Nature 607 60-68 (2022)	1

Measured cross sections and exclusion limits at 95% C.L. See here for all cross section summary plots

13 TeV Nature 607 60-68 (2022)

Overview of CMS cross section results

Inner colored bars statistical uncertainty, outer narrow bars statistical+systematic uncertaint Light to Dark colored bars: 2.76, 5.02, 7, 8, 13, 13.6 TeV, Black bars: theory prediction

EW tests over the years

* LEP: e^+e^- collider, in operation 1989-2000

- ✤ 4 experiments (ALEPH, DELPHI, L3, OPAL)
- * \sqrt{s} from 90 GeV to 209 GeV
- Two phases
 - ✤ LEP1: Z physics
 - LEP2: W physics, reaching the WW threshold and above

* largest and most powerful e^+e^- collider to date

* SLC, e^+e^- collider, in operation 1989-1998

* \sqrt{s} ~ 90 GeV , polarised beams

SppS, in operation 1981-1990

* $\sqrt{s} = 540, 630 \, \text{GeV}$

- discovery of W and Z bosons (UA1 &UA2)
- Tevatron, *pp* collisions (1987-2011)
 - * $\sqrt{s} = 1.8, 1.96 \text{ TeV}$
 - top quark discovery, mass measurement
 - * M_W measurements
- LHC, pp collisions (2008)

* ...

Higgs boson discovery

Testing the SM

What can be tested? *

* ...

- Theoretical predictions for observables: cross sections, differential distributions, decay widths,...
- "Properties" of the theory
 - built-in assumptions, e.g. number of generations
 - existence of the particles appearing in the model and their properties
 - existence of the interactions predicted by the model and their properties → e.g. Yukawa couplings, gauge boson interactions, gauge-fermion interactions
 - running of the coupling

* Overall consistency \rightarrow global precision fits

 \rightarrow most famous: Higgs boson but also top quark, τ neutrino, and even earlier W, Z bosons or b and c quarks,...

$e^+e^- \rightarrow f\bar{f}$

* $e^+e^- \to f\bar{f}$ cross section

at the Z resonance

$$\sigma = \frac{4\pi\alpha}{3s} \frac{1}{16sin^4\theta_W \cos^4\theta_W} (c_V^{e^2} + c_A^{e^2}) (c_V^{f^2} + c_A^{f^2}) \frac{1}{(s-x)^2} (s-x)^2 (s-x)^2$$

Partial width

$$\Gamma_f = \frac{\alpha M_Z}{12\sin^2\theta_W \cos^2\theta_W} (c_V^{f^2} + c_A^{f^2})$$

$$\Rightarrow \qquad \sigma = 12\pi \frac{\Gamma_e \Gamma_f}{M_Z^2} \frac{s}{(s - M_Z^2)^2 + (M_Z \Gamma_Z)^2} \qquad \text{and at } \sqrt{s} = N$$

 $(M_Z^2)^2 + (M_Z \Gamma_Z)^2$

$$\left(\Gamma_Z = \sum_f \Gamma_f\right)$$

$$A_Z \qquad \qquad \sigma = \frac{12\pi}{m_Z^2} \frac{\Gamma_e \Gamma_f}{\Gamma_Z^2}$$

Z line shape

Z resonance curve

$$\sigma = 12\pi \frac{\Gamma_e \Gamma_f}{M_Z^2} \frac{s}{(s - M_Z^2)^2 + (s \Gamma_Z / M_Z)^2} \text{ with peak at } \sigma_{\text{peak}} = \frac{1}{n}$$

The formula receives QED (and QCD if hadrons in the final state are considered) corrections, most importantly from the QED initial state radiation

	$m_{ m Z}[{ m GeV}]$	91.1891 ± 0.0031
	$\Gamma_{ m Z} [{ m GeV}]$	2.4959 ± 0.0043
	$\sigma_{ m h}^0[{ m nb}]$	41.558 ± 0.057
Г	$R_{ m e}$	20.690 ± 0.075
$R_l = \frac{\Gamma_{had}}{\Gamma}$	R_{μ}	20.801 ± 0.056
1	$R_{ au}$	20.708 ± 0.062

Test of lepton couplings' universality

as measured by ALEPH

✤ Determination of c_V^f , c_A^f from forward-backward asymmetry
 (→ differential distributions)

Number of light neutrino generations

• The relation $\Gamma_Z = \Gamma_{had} + 3\Gamma_l + \Gamma_{inv}$

allows to determine the invisible partial width

$$\Gamma_{\rm inv} = N_{\nu} \Gamma_{\nu} \qquad \qquad \Gamma_{\nu} = \Gamma(Z \to \nu_i \bar{\nu}_i)$$

The number of neutrinos

$$N_{\nu} = \frac{\Gamma_{\text{inv}}}{\Gamma_{\nu}} = \left(\frac{\Gamma_{\text{inv}}}{\Gamma_{l}}\right) \left(\frac{\Gamma_{l}}{\Gamma_{\nu}}\right)$$

$$\uparrow \qquad \uparrow \qquad \uparrow$$

from exp. from SM

Combined result from the four LEP experiments

$$N_v = 2.984 \pm 0.008$$

Away from the Z pole

WW threshold

- * Measurements of $e^+e^- \rightarrow W^+W^- \rightarrow 4$ fermions allow to
 - probe gauge boson interactions
 - determine M_W from the dependence of the WW production threshold behaviour on M_W)

Global fits

image credit: J. de Blas

Global consistency of the SM

Predictions from global fits

mass prediction on basis of measurements at LEP and SLC

• With the top mass measured, one could make prediction for the Higgs mass \rightarrow the famous "blue band" plot

Just before turning on the LHC...

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Just before turning on the LHC...

5 December 2001

The legendary particle that physicists thought explained why matter has mass probably does not exist. So say researchers who have spent a year analysing data from the LEP accelerator at the CERN nuclear physics lab near Geneva.

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NEWSLETTERS

5 December 2001

The legendary particle that physicists thought explained why matter has mass probably does not exist. So say researchers who have spent a year analysing data from the LEP accelerator at the CERN nuclear physics lab near Geneva.

Not everyone is too bothered, yet. Frank Wilczek, a particle physics theorist at the Massachusetts Institute of Technology, points out that you could take the LEP results as evidence that the Higgs must be sitting at an improbably high energy. He says he'll start to get uncomfortable if the Higgs doesn't show up by about 130 GeV. "Then I would have a good long think," he says.

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e a few good reasons to keep on searching:	y matter has mass probably does no n the LEP accelerator at the CERN n
Unitarity	
Triviality	ysics theorist at the Massachusetts P results as evidence that the Higgs get uncomfortable if the Higgs does
Vacuum stability	ink," he says.

Unitarity bound

• Consider $2 \rightarrow 2$ elastic scattering $\frac{d\sigma}{d\Omega} = \frac{1}{64\pi^2 s} |A|^2$ * a_i are the spin *l* partial wave amplitudes $P_l(\cos\theta)$ are Legendre polynomials: $\int_{-1}^{1} dx P_l(x) P_{l'}(x) = \frac{2\delta_{l,l'}}{2l+1}$ $\Rightarrow \sigma = \frac{8\pi}{s} \sum_{l=0}^{\infty} (2l+1) \sum_{l'=0}^{\infty} (2l'+1) a_l a_{l'}^* \int_{-1}^{1} d\cos\theta$

Direct consequence of the unitarity of the S matrix: the optical theorem

 $\sigma = \frac{1}{s} \operatorname{Im} \left(A(\theta = 0) \right) \quad \text{so for each } l \qquad \operatorname{Im} \left(a_l \right) = |a_l|^2 \quad \text{and} \quad \left| \operatorname{Re} \left(a_l \right) \right| \le \frac{1}{2}$

Partial wave decomposition of amplitude

$$A = 16\pi \sum_{l=0}^{\infty} (2l+1) P_l(\cos\theta)a_l$$

$$\theta P_l(\cos\theta) P_{l'}(\cos\theta) = \frac{16\pi}{s} \sum_{l=0}^{\infty} (2l+1) |a_l|^2$$

Unitarity bound (2)

- Highly relativistic (boosted in the z-direction) gauge bosons dominated by their longitudinal polarisation
- Longitudinal degree of freedom ↔ Goldstone mode
- approximated by the scattering amplitude for the Goldstone bosons, up to $\mathcal{O}(M_W^2/E^2)$ corrections

Hence for $s \gg M_H^2$ from |Re a_0 | < $\frac{1}{2}$ it follows $M_H < 2\sqrt{\pi v} = 870$ GeV

• Equivalence theorem: at very high energies, $s \gg M_W^2$, the scattering amplitude for the longitudinally polarised W bosons can be

Triviality bound

Running of the Higgs self coupling

$$\frac{d\lambda}{d\log Q^2} = \frac{1}{16\pi^2} \left[12\lambda^2 + 6\lambda y_t^2 - 3y_t^4 - \frac{3}{2}\lambda\left(3g_2^2 + g_1^2\right) + \frac{3}{16}\left(2g_2^4 + \left(g_2^2 + g_1^2\right)^2\right) \right]$$

• In the limit of the strong coupling λ

$$\frac{d\lambda}{d\log Q^2} = \frac{3}{4\pi^2}\lambda^2$$

(Landau) pole at
$$1 = \frac{3}{4\pi^2} \lambda(v) \log\left(\frac{Q^2}{v^2}\right)$$

Requesting finite λ at a given $Q = \Lambda$, i.e. $\lambda(\Lambda)^{-1} > 0$ leads to

with a solution

$$\lambda(Q) = \frac{\lambda(v)}{1 - \frac{3}{4\pi^2}\lambda(v)\log\left(\frac{Q^2}{v^2}\right)}$$

o scale-dependent condition
$$M_H^2 < \frac{8\pi^2 v^2}{3\log\left(\frac{\Lambda^2}{v^2}\right)}$$

Vacuum stability bound

$$\frac{d\lambda}{d\log Q^2} = \frac{1}{16\pi^2} \left[12\lambda^2 + 6\lambda y_t^2 - 3y_t^4 - \frac{3}{2}\lambda\left(3g_2^2 + g_1^2\right) + \frac{3}{16}\left(2g_2^4 + \left(g_2^2 + g_1^2\right)^2\right) \right]$$

- negative value of the coupling \rightarrow unstable Higgs potential
- Considering only the dominant term with the top Yukawa coupling y_t *

$$\frac{d\lambda}{d\log Q^2} = -\frac{3m_t^4}{4\pi^2 v^4} \qquad \text{with the solution } \lambda(Q^2) = \lambda$$

* Imposing $\lambda(\Lambda) > 0$ leads to $M_H^2 > \frac{3m_t^4}{2\pi^2 v^2} \log\left(\frac{\Lambda^2}{v^2}\right)$

* In the limit of a weak Higgs self-coupling at the EW scale, the negative terms in the equation above can in principle lead to a

$$(v^2) - \frac{3m_t^4}{4\pi^2 v^4} \log\left(\frac{Q^2}{v^2}\right)$$

Constraints on M_H in the SM

Constraints on M_H in the SM

12 years with the Higgs @ LHC

Global signal strength $\mu = \sigma / \sigma_{SM}$

 $\mu = 1.05 \pm 0.06$ (ATLAS)

 $\mu = 1.00 \pm 0.06$ (CMS)

How to make the SM Higgs at the LHC

Gluon fusion

- Higgs boson does not couple directly to gluons
- * At LO the gluon fusion process is mediated via loop, with the biggest contribution from the top quark ($y_t \propto m_t$)
- Effective (loop-induced) ggH coupling in the heavy top mass $(m_t \to \infty)$ limit

$$\mathscr{L}_{ggH, eff} = -\frac{1}{4\nu} C_H H G^{\mu\nu} G_{\mu\nu}$$

In the heavy top limit approximation analogous to the Drell-Yan process!

Higgs boson interactions

* Higgs-gauge boson couplings \rightarrow test of the EWSB mechanism, Yukawa couplings \rightarrow mass of elementary particles

The scaling of the couplings with the mass of the particles is a central prediction of the theory

κ 's are rescaling factors of the SM couplings

 $-igM_Vg_{\mu\nu}\kappa_V$

 "kappa-framework" is very / too simple (problems with gauge invariance, kinematic information, consistency at higher orders,..), better approach is provided by effective field theory

Higgs self-coupling

$$V = \frac{M_H^2}{2}H^2 + \lambda v H^3 + \frac{\lambda}{4}H^4$$

- Value of λ decides the shape of the SM scalar potential
- * λ can be probed at hadron colliders through the Higgs pair-production processes
- Extremely low cross section at the LHC (1000 times * smaller than single Higgs production)

-2 10

 10^{3}

10 ²

10

arXiv:1910.00012

SM in EFT

physics at a scale $\Lambda \gg v \rightarrow SM$ as lowest order in EFT expansion of the full theory

$$\mathscr{L}_{SMEFT} = \mathscr{L}_{SM} + \frac{1}{\Lambda}\mathscr{L}_5 + \frac{1}{\Lambda}\mathscr{L}_6 + \dots$$

mechanism of EWSB via single SU(2) doublet field) apart from renormalizability: order-by-order

→ Dim-5: only 1 operator (Weinberg) $\mathcal{O}_5 = \frac{v^2}{2} \bar{\nu}_L^c \nu_L \rightarrow \text{Majorana mass term}$

Dim-6: 59 operators ("Warsaw basis"), 2499 if flavour structure considered

* All the principles same as behind SM (QFT, local gauge symmetries, matter content and the quantum numbers, Higgs

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* SM "works" fantastically well \rightarrow a very good approximation of an unknown BSM theory chosen by Nature \rightarrow all new

The future is bright

* Next stage of the LHC, after current Run 3: High-Luminosity LHC (HL-LHC): 20 times larger data sample and improved detectors

 However, precision measurements of the full Higgs sector: triple and quartic self-interactions, Yukawa couplings to light fermions etc. will require next generation colliders such as e.g. FCC (-ee).

Summary

- complete
- Results of collider physics experiments are in agreement with the EWSM it "works" amazingly well *
- * Yet, it can only be an effective theory the evidence for BSM physics is there
 - dark matter
 - neutrino masses
 - matter-antimatter symmetry
 - *
- The Higgs sector still needs to be fully tested
- The search for new physics relies on developments in both experiment and theory

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*

. . .

... and many theoretical questions remain:

- * where does the Higgs potential come from?
- why 3 generations? Why the observed mass * pattern?
- what protects the Higgs mass (naturalness * problem)?

Summary

- * With the discovery of a new boson 12 years ago, and then subsequent confirmation of it as the Higgs boson, the EW SM is now complete
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- * where does the Higgs potential come from?
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Good news: there is a lot to understand and discover!

