



Reconstruction of Black Hole Metric Perturbations from Weyl Curvature II: The Regge-Wheeler gauge

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Motivations

- The Teukolsky Equation gives us Ψ_4 and Ψ_0 only.
- Metric Reconstruction.
 1. Self-Force computations
 2. Horizon structure
 3. Second order perturbations
- Time domain
 1. Computational efficiency
 2. Easier interface to NR
- With B. Whiting we did the metric reconstruction in the *radiation* gauges, in vacuum, in the Schwarzschild background.
- Here we do the metric reconstruction in the Regge-Wheeler gauge, still in the Schwarzschild background, but with arbitrary source terms.

Weyl Scalars

$$\psi_4 = \frac{1}{2} \left\{ \begin{aligned} &(\bar{\delta} + 3\alpha + \bar{\beta} - \bar{\tau})(\bar{\delta} + 2\alpha + 2\bar{\beta} - \bar{\tau})h_{\text{nn}} \\ &+ (\hat{\Delta} + \bar{\mu} + 3\gamma - \bar{\gamma})(\hat{\Delta} + \bar{\mu} + 2\gamma - 2\bar{\gamma})h_{\text{mmm}} \\ &- \left[2(\hat{\Delta} + \bar{\mu} + 3\gamma - \bar{\gamma})(\bar{\delta} - \bar{\tau} + 2\alpha) \right. \\ &\left. + 2(\bar{\delta} + 3\alpha + \bar{\beta} - \bar{\tau})(\bar{\mu}) \right] h_{(\text{nm})} \end{aligned} \right\}, \quad (46)$$

and

$$\psi_0 = \frac{1}{2} \left\{ \begin{aligned} &(\delta - \bar{\alpha} - 3\beta + \bar{\pi})(\delta - 2\bar{\alpha} - 2\beta + \bar{\pi})h_{\text{ll}} \\ &+ (\hat{D} - \bar{\rho} - 3\epsilon + \bar{\epsilon})(\hat{D} - \bar{\rho} - 2\epsilon + 2\bar{\epsilon})h_{\text{mmm}} \\ &- \left[2(\hat{D} - \bar{\rho} - 3\epsilon + \bar{\epsilon})(\delta + \bar{\pi} - 2\beta) \right. \\ &\left. - 2(\delta - \bar{\alpha} - 3\beta + \bar{\pi})(\bar{\rho}) \right] h_{(\text{lm})} \end{aligned} \right\}. \quad (47)$$

where $h_{\text{nm}} = n^\mu n^\nu h_{\mu\nu}$, $h_{\text{lm}} = l^\mu m^\nu h_{\mu\nu}$, etc.

Weyl Scalars (RW)

In the Regge-Wheeler gauge ($h_1^{\ell m} = h_0^{\ell m} = G^{\ell m} = 0 = {}^{(odd)}h_2^{\ell m}$), we get

$$\begin{aligned} \psi_4 \doteq \sum_{\ell m} \psi_4^{\ell m} {}_{-2}Y_{\ell m} &= - \sum_{\ell m} \sqrt{\frac{(\ell-2)!}{(\ell+2)!}} \\ &\times \left\{ -\frac{f}{16r^2} (H_0^{\ell m} - 2H_1^{\ell m} + H_2^{\ell m}) \right. \\ &\left. - \frac{i}{8r^2} [\partial_t - f\partial_r + f'] ({}^{(odd)}h_0^{\ell m} - f {}^{(odd)}h_1^{\ell m}) \right\} {}_{-2}Y_{\ell m}, \end{aligned} \quad (11)$$

and

$$\begin{aligned} \psi_0 \doteq \sum_{\ell m} \psi_0^{\ell m} {}_{+2}Y_{\ell m} &= - \sum_{\ell m} \sqrt{\frac{(\ell-2)!}{(\ell+2)!}} \\ &\times \left\{ -\frac{1}{4fr^2} (H_0^{\ell m} + 2H_1^{\ell m} + H_2^{\ell m}) \right. \\ &\left. - \frac{i}{2f^2r^2} [\partial_t + f\partial_r - f'] ({}^{(odd)}h_0^{\ell m} + f {}^{(odd)}h_1^{\ell m}) \right\} {}_{+2}Y_{\ell m}. \end{aligned} \quad (12)$$

Even and Odd parity decomposition

We define

$$\psi^\pm = \frac{1}{2} \left[\psi^{\ell, m} \pm (-)^m \bar{\psi}^{\ell, -m} \right], \quad (13)$$

where for notational simplicity we dropped the ℓ, m indexes. Thus, given the symmetric nature of the even parity metric perturbations, Eqs. (11) and (12) take the form

Even parity:

$$\psi_4^+ = \frac{f}{16r^2} \sqrt{\frac{(\ell-2)!}{(\ell+2)!}} (H_0^{\ell m} - 2H_1^{\ell m} + H_2^{\ell m}), \quad (14)$$

and

$$\psi_0^+ = \frac{1}{4fr^2} \sqrt{\frac{(\ell-2)!}{(\ell+2)!}} (H_0^{\ell m} + 2H_1^{\ell m} + H_2^{\ell m}). \quad (15)$$

From Eqs. (14) and (15) we can obtain the ℓm components of the metric perturbations as follows

$$H_1^{\ell m}(r, t) = -\frac{4r^2}{f} \sqrt{\frac{(\ell+2)!}{(\ell-2)!}} \left[\psi_4^+ - \frac{f^2}{4} \psi_0^+ \right] \quad (16)$$

and

$$H_0^{\ell m}(r, t) + H_2^{\ell m}(r, t) = \frac{8r^2}{f} \sqrt{\frac{(\ell+2)!}{(\ell-2)!}} \left[\psi_4^+ + \frac{f^2}{4} \psi_0^+ \right] \quad (17)$$

We now bring into the play the Hilbert-Einstein's equations in the Regge-Wheeler gauge, Eq. (A.14) give,

8th

$$H_0^{\ell m}(r, t) - H_2^{\ell m}(r, t) = \frac{16\pi r^2}{\sqrt{2\lambda(\lambda+1)}} F^{\ell m}, \quad (18)$$



Using Eq. (A.12) we can solve for $\partial_r K^{\ell m}$ and then replace it in Eq. (A.8) to find $K^{\ell m}$ in terms of the other even parity metric coefficients and source terms

$$\begin{aligned}
K(r, t)^{\ell m} = & \\
& 2 \frac{(r - M) \frac{\partial}{\partial r} H_0^{\ell m}(r, t)}{\lambda} + \frac{(r - 2M) r \frac{\partial^2}{\partial r^2} H_0^{\ell m}(r, t)}{\lambda} \\
& - \frac{r^2 \frac{\partial^2}{\partial r \partial t} H_1^{\ell m}(r, t)}{\lambda} + \frac{M \frac{\partial}{\partial r} H_2^{\ell m}(r, t)}{\lambda} \\
& - \frac{(2r^2 - 8rM + 9M^2) H_0^{\ell m}(r, t)}{r\lambda (r - 2M)} \\
& + \frac{(-r^2\lambda + 2rM\lambda + 3M^2 - 2rM) H_2^{\ell m}(r, t)}{r\lambda (r - 2M)} \\
& + \frac{r(-3r + 7M) \frac{\partial}{\partial t} H_1(r, t)}{\lambda (r - 2M)} + 8 \frac{(r - 2M) \pi r^2 \frac{\partial}{\partial r} B^{\ell m}(r, t)}{\sqrt{\lambda + 1}\lambda} \\
& - 8 \frac{(7M - 4r) \pi r B^{\ell m}(r, t)}{\sqrt{\lambda + 1}\lambda} - \frac{A_0^{\ell m}(r, t) r^3}{\lambda (r - 2M)}. \tag{19}
\end{aligned}$$

This form of the metric coefficient $K^{\ell m}$, involves second derivatives of the Weyl scalars. One can consider an alternative integral form (on the hypersurface $t = \text{constant}$) derived from Eq. (A.12)

$$\begin{aligned}
K^{\ell m} = & H_0^{\ell m} + \int_{2M}^r \frac{dr}{1 - \frac{2M}{r}} \left[-\frac{\partial H_1^{\ell m}}{\partial t} + \frac{2M}{r^2} H_0^{\ell m} \right. \\
& \left. - 16\pi (r - 2M) \frac{F^{\ell m}}{\sqrt{2\lambda(\lambda + 1)}} - \frac{8\pi(r - 2M)}{\sqrt{\lambda + 1}} B^{\ell m} \right] \tag{20}
\end{aligned}$$

Odd Parity

From Eq. (11) and (12), given the antisymmetric behaviour of the odd parity metric coefficients, we get

$$\psi_4^- = \frac{i}{8r^2} \sqrt{\frac{(\ell-2)!}{(\ell+2)!}} [\partial_t - f\partial_r + f'] ({}^{(odd)}h_0^{\ell m} - f {}^{(odd)}h_1^{\ell m}), \quad (21)$$

and

$$\psi_0^- = \frac{i}{2f^2r^2} \sqrt{\frac{(\ell-2)!}{(\ell+2)!}} [\partial_t + f\partial_r - f'] ({}^{(odd)}h_0^{\ell m} + f {}^{(odd)}h_1^{\ell m}). \quad (22)$$

A linear combination of these previous equations produces

$$\psi_4^- + \frac{f^2}{4}\psi_0^- = \frac{i}{4r^2} \sqrt{\frac{(\ell-2)!}{(\ell+2)!}} [\partial_t ({}^{(odd)}h_0^{\ell m} + (f\partial_r - f') (f {}^{(odd)}h_1^{\ell m}))], \quad (23)$$

and

$$\psi_4^- - \frac{f^2}{4}\psi_0^- = \frac{-i}{4r^2} \sqrt{\frac{(\ell-2)!}{(\ell+2)!}} [f\partial_t ({}^{(odd)}h_1^{\ell m} + (f\partial_r - f') ({}^{(odd)}h_0^{\ell m}))]. \quad (24)$$

Odd Parity (end)

which integrated produces

$$h_1^{lm} = \frac{1}{\sqrt{1 - \frac{2M}{r}}} \left\{ \int_{2M}^r S_1^{lm}(r', t) \sqrt{1 - \frac{2M}{r'}} dr' + C_1^{lm}(t) \right\}, \quad (26)$$

and

$$h_0^{lm} = \left(1 - \frac{2M}{r}\right) \left\{ \int_{2M}^r \frac{S_0^{lm}(r', t)}{1 - \frac{2M}{r'}} dr' + C_0^{lm}(t) \right\}. \quad (29)$$

where

$$S_1^{lm}(r, t) \doteq \frac{-2ir^2}{f^2} \sqrt{\frac{(\ell+2)!}{(\ell-2)!}} \left(\psi_4^- + \frac{f^2}{4} \psi_0^- \right) - \frac{2\pi ir^2 D_{\ell m}}{f \sqrt{\lambda(\lambda+1)}}, \quad (25)$$

$$S_0^{lm}(r, t) \doteq \frac{4ir^2}{f} \sqrt{\frac{(\ell+2)!}{(\ell-2)!}} \left(\psi_4^- - \frac{f^2}{4} \psi_0^- \right) - \partial_t h_1^{lm},$$

(28)

Waveforms' metric reconstruction: Even Parity

$$K^{\ell m} = \frac{6M^2 + 3M\lambda r + \lambda(\lambda + 1)r^2}{r^2(\lambda r + 3M)} \psi_{\text{even}}^{\ell m} + \left(1 - \frac{2M}{r}\right) \partial_r \psi_{\text{even}}^{\ell m} - \frac{8\pi r^3 A_{\ell m}^{(0)}}{(\lambda + 1)(\lambda r + 3M)}, \quad (\text{B.5})$$

and

$$H_2^{\ell m} = -\frac{9M^3 + 9\lambda M^2 r + 3\lambda^2 M r^2 + \lambda^2(\lambda + 1)r^3}{r^2(\lambda r + 3M)^2} \psi_{\text{even}}^{\ell m} + \frac{3M^2 - \lambda M r + \lambda r^2}{r(\lambda r + 3M)} \partial_r \psi_{\text{even}}^{\ell m} + (r - 2M) \partial_r^2 \psi_{\text{even}}^{\ell m} - \frac{8\pi r^4}{(\lambda + 1)(\lambda r + 3M)} \partial_r A_{\ell m}^{(0)} + \frac{8\pi r^3 (\lambda^2 r^2 - 2\lambda r^2 + 10\lambda r M - 9r M + 27M^2)}{(\lambda + 1)(r - 2M)(\lambda r + 3M)^2} A_{\ell m}^{(0)}. \quad (\text{B.6})$$

Even Parity (Contd.)

From Eq. (A.9) and the expressions for $\partial_t K^{\ell m}$ and $\partial_t H_2^{\ell m}$ in terms of $\partial_t \psi_{\text{even}}^{\ell m}$, we find the $H_1^{\ell m}$ metric coefficient in the Regge-Wheeler gauge

$$H_1^{\ell m} = r \partial_r (\partial_t \psi_{\text{even}}^{\ell m}) + \frac{\lambda r^2 - 3M\lambda r - 3M^2}{(r - 2M)(\lambda r + 3M)} \partial_t \psi_{\text{even}}^{\ell m} - \frac{8\pi r^5}{(\lambda + 1)(r - 2M)(\lambda r + 3M)} \partial_t A_{\ell m}^{(0)} + \frac{4\sqrt{2}i\pi r^2}{(\lambda + 1)} A_{\ell m}^{(1)}. \quad (\text{B.7})$$

These equations together with

$$H_0^{\ell m} = H_2^{\ell m} + \frac{16\pi r^2 F_{\ell m}}{\sqrt{2\lambda(\lambda + 1)}}, \quad (\text{B.8})$$

Completes the four even parity metric coefficients.

Odd parity

One can use the field equations to write the metric perturbation in the Regge-Wheeler gauge

$$h_0^{\ell m}(r, t) = \frac{1}{2} \left(1 - \frac{2M}{r}\right) \partial_r (r \psi_{\text{odd}}^{\ell m}) + \frac{4\pi r^3 Q_{\ell m}^{(0)}}{\lambda \sqrt{(\lambda + 1)}} \quad (\text{B.21})$$

$$h_1^{\ell m}(r, t) = \frac{1}{2} \frac{r}{\left(1 - \frac{2M}{r}\right)} \partial_t \psi_{\text{odd}}^{\ell m} + \frac{4\pi i r^3 Q_{\ell m}}{\lambda \sqrt{(\lambda + 1)}}$$

And this completes the whole metric reconstruction in the RW gauge in terms of Moncrief waveforms.

Chandrasekhar transformations

$$\begin{aligned}
 \psi_4^+ = & \frac{1}{16r} \sqrt{\frac{(\ell-2)!}{(\ell+2)!}} \left\{ 2\psi_{,r^*r^*}^{\text{even}} - 2\psi_{,tr^*}^{\text{even}} + W^+(\psi_{,r^*}^{\text{even}} - \psi_{,t}^{\text{even}}) - V^+\psi^{\text{even}} \right. \\
 & + \frac{16\pi r^3}{(\lambda r + 3M)(\lambda + 1)} \left(\partial_t A_{lm}^{(0)} - \partial_{r^*} A_{lm}^{(0)} \right) - \frac{8i(r-2M)\sqrt{2\pi} A_{lm}^{(1)}(r,t)}{\lambda + 1} \\
 & + 16 \frac{\pi r (\lambda^2 r^2 - 2\lambda r^2 + 10\lambda rM - 9rM + 27M^2) A_{lm}^{(0)}(r,t)}{(\lambda + 1)(\lambda r + 3M)^2} \\
 & \left. - 8 \frac{F_{lm}(r,t)\sqrt{2\pi}(r-2M)}{\sqrt{\lambda}(\lambda + 1)} \right\}, \quad (C.1)
 \end{aligned}$$

$$\begin{aligned}
 \psi_4^- = & \frac{-i}{16r} \sqrt{\frac{(\ell-2)!}{(\ell+2)!}} \left\{ 2\psi_{,r^*r^*}^{\text{odd}} - 2\psi_{,tr^*}^{\text{odd}} + W^-(\psi_{,r^*}^{\text{odd}} - \psi_{,t}^{\text{odd}}) - V^-\psi^{\text{odd}} \right. \\
 & - \frac{16\pi r^2}{\lambda(\lambda + 1)} \left(\partial_t Q_{lm}^{(0)} - \partial_{r^*} Q_{lm}^{(0)} \right) + \frac{16i\pi r(r-2M)}{\lambda(\lambda + 1)} (\partial_t Q_{lm} - \partial_{r^*} Q_{lm}) \\
 & \left. - \frac{48i\pi(r-2M)^2 Q_{lm}}{\lambda\sqrt{\lambda+1}r} + \frac{16\pi(3r-8M)Q_{lm}^{(0)}}{\lambda\sqrt{\lambda+1}} - S^- \right\}, \quad (C.2)
 \end{aligned}$$

And similar expressions for Ψ_0

Inverse Chandrasekhar transformations

$$\begin{aligned}
 \psi^{\text{even}} = & \frac{r^2 (r - 2M) \frac{\partial^2}{\partial r^2} H_0(r, t)}{\lambda (\lambda + 1)} - \frac{r^3 \frac{\partial^2}{\partial r \partial t} H_1(r, t)}{\lambda (\lambda + 1)} \\
 & + \frac{r (rM\lambda - 3M^2 + r^2\lambda + 6rM) \frac{\partial}{\partial r} H_0(r, t)}{(\lambda + 1) \lambda (\lambda r + 3M)} \\
 & + \frac{(2r^2\lambda - 5rM\lambda - 21M^2 + 9rM) r^2 \frac{\partial}{\partial t} H_1(r, t)}{(\lambda + 1) (\lambda r + 3M) (-r + 2M) \lambda} \\
 & - \frac{(M^2\lambda r + 2r^2M\lambda^2 - 12r^2M - r^2M\lambda - 2r^3\lambda + 42rM^2 - r^3\lambda^2 - 63M^3) H_0(r, t)}{2(\lambda + 1) (\lambda r + 3M) (-r + 2M) \lambda} \\
 & - 4 \frac{r^2 \pi \sqrt{2\lambda + 2} (-5r^2\lambda - 12rM + 9rM\lambda + 21M^2) \sqrt{2} B(r, t)}{(\lambda + 1)^2 \lambda (\lambda r + 3M)} \\
 & + \frac{r^4 A^{(0)}(r, t)}{\lambda (\lambda + 1) (-r + 2M)} \\
 & + 4 \frac{\sqrt{2} \sqrt{\lambda (\lambda + 1)} r^2 \pi (2rM - 11M^2 - r^2\lambda + 2rM\lambda) F(r, t)}{(-r + 2M) (\lambda + 1)^2 \lambda^2} \\
 & - 8 \frac{r^3 (-r + 2M) \pi \frac{\partial}{\partial r} B(r, t)}{(\lambda + 1)^{3/2} \lambda} - 8 \frac{r^3 M \pi \sqrt{2} \frac{\partial}{\partial r} F(r, t)}{(\lambda + 1) \sqrt{\lambda (\lambda + 1)} \lambda}, \tag{C.10}
 \end{aligned}$$

Inverse Chandrasekhar transformations

So, finally

$$\psi^{\text{odd}} = \frac{r}{\lambda} \left\{ -\frac{2}{r} \left(1 - \frac{M}{r}\right) \int_{2M}^r \frac{S_0(r', t)}{1 - \frac{2M}{r'}} dr' + \frac{4ir^2}{f} \sqrt{\frac{(\ell+2)!}{(\ell-2)!}} \left(\psi_4^- - \frac{f^2}{4} \psi_0^- \right) - \frac{2}{\sqrt{1 - \frac{2M}{r}}} \left[\int_{2M}^r \partial_t S_1(r', t) \sqrt{1 - \frac{2M}{r'}} dr' \right] \right\}. \quad (\text{C.12})$$

Gives the odd parity waveforms.

Applications:

Numerical: wave extraction, tetrad, etc

Analytical: Initial Data, BC, QNM, etc

Kerr perturbations (RW1)

A possible generalization of the Regge-Wheeler gauge conditions for spherically symmetric backgrounds, but where perturbations are not decomposed into multipoles is [7]

$$(\sin \theta)^2 h_{\theta\theta} - h_{\phi\phi} = 0, \quad (31)$$

$$h_{\theta\phi} = 0, \quad (32)$$

$$\sin \theta \partial_\theta(\sin \theta h_{t\theta}) + \partial_\phi h_{t\phi} = 0, \quad (33)$$

$$\sin \theta \partial_\theta(\sin \theta h_{r\theta}) + \partial_\phi h_{r\phi} = 0. \quad (34)$$

The first equation above leads to the condition $G^{\ell m} = 0$. The second gives then $^{(odd)}h_2^{\ell m} = 0$. The other two differential conditions are chosen such that they lead to $^{(even)}h_0^{\ell m} = 0 = ^{(even)}h_1^{\ell m}$, but allow $^{(odd)}h_0^{\ell m} \neq 0$ and $^{(odd)}h_1^{\ell m} \neq 0$ be unconstrained.

Now we will consider the generalization of the Regge-Wheeler gauge in the Newman-Penrose formalism. In this formalism, the first two Regge-Wheeler conditions, Eqs. (31) and (32), have a simple generalization

$$h_{\mathbf{m}\mathbf{m}} = \mathbf{m}^\mu \mathbf{m}^\nu h_{\mu\nu} = 0. \quad (35)$$

Note that requiring that the real and imaginary parts vanish contains both conditions. Obviously,

$$h_{\bar{\mathbf{m}}\bar{\mathbf{m}}} = \bar{\mathbf{m}}^\mu \bar{\mathbf{m}}^\nu h_{\mu\nu} = 0, \quad (36)$$

also holds. Note that conditions (35) and (36) are invariant under type III (spin-boosts) transformations of the background tetrad

$$l \rightarrow A^2 l, \quad n \rightarrow A^{-2} n, \quad m \rightarrow e^{2i\Theta} m, \quad \bar{m} \rightarrow e^{-2i\Theta} \bar{m}. \quad (37)$$



Kerr perturbations (RW2)

To generalize the differential conditions (33) and (34) one can resort to the type III transformation properties of the δ and $\bar{\delta}$, as well as spin coefficient operators in the Kerr background acting on the metric coefficients $h_{(lm)}$ and $h_{(nm)}$. The objects

$$(\delta - 2\bar{\alpha})h_{(lm)} \rightarrow A^2(\delta - 2\bar{\alpha})h_{(lm)}, \quad (38)$$

and

$$(\bar{\delta} + 2\bar{\beta})h_{(nm)} \rightarrow A^{-2}(\bar{\delta} + 2\bar{\beta})h_{(nm)}, \quad (39)$$

transform as objects of spin- 0 and boost weight +1 and -1 respectively under type III transformations of the background tetrad (37).

In order to reproduce the differential conditions (33) and (34) in the Schwarzschild limit, one can then require

$$\Re [(\delta - 2\bar{\alpha})h_{(lm)}] = 0, \quad (40)$$

and

$$\Re [(\bar{\delta} + 2\bar{\beta})h_{(nm)}] = 0, \quad (41)$$

Where \Re is the real part.

Ambiguity by adding τ , π and its bars.

Discussion

1. We reconstructed the metric in the RW gauge in the Schwarzschild background from Ψ_4 and Ψ_0 including source terms.
2. To Continue Kerr program:
 - by decomposing into m-modes. Still Even/odd definition applies.
 - Use GR field equations projected along the tetrad.
 - Construct the metric.
3. Chandra and inverse-Chandra transformations in the time domain.