



Gravitational Self-force on a Particle in the Schwarzschild background

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1. Introduction

Laser Interferometer Space Antenna (LISA)

One of the most promising wave source

Super massive black hole - Compact object binaries

Black hole perturbation:

Black hole background (mass: M)

+ Perturbation [point particle]

μ

(mass:)

$$g_{\mu\nu} = g_{\mu\nu}^{(b)} + h_{\mu\nu}^{\text{full}} .$$

We want to know the precise particle motion

which includes the self-force in a black hole space-time.

Point particle \rightarrow Self-force diverge \rightarrow Need regularization

MiSaTaQuWa force

Mino, Sasaki and Tanaka ('97), Quinn and Wald ('97)
Detweiler and Whiting ('03) **under Lorenz (L) gauge**

$$h_{\mu\nu}^{\text{full,L}} = h_{\mu\nu}^{\text{S(dir),L}} + h_{\mu\nu}^{\text{R(tail),L}},$$

Regularized gravitational self-force (reaction force)

$$\mu \frac{D^2 z^\mu(\tau)}{d\tau^2} = F^\mu(z)$$

$$F^\mu = -\frac{\mu}{2} (g_{(b)}^{\mu\nu} + u^\mu u^\nu) \left(2h_{\nu\beta;\alpha}^{\text{R,L}} - h_{\alpha\beta;\nu}^{\text{R,L}} \right) u^\alpha u^\beta,$$

$\{u^\alpha\}$: Four velocity of a particle

``R-part``: **Homogeneous solution** of linearized Einstein equation,
Depend on the history and the global structure of a space-time.
It is difficult to obtain this **directly**.

Regularization

Regularization: Subtract the singular part.

$$h_{\mu\nu}^{\text{R,L}} = h_{\mu\nu}^{\text{full,L}} - h_{\mu\nu}^{\text{S,L}} .$$

“S-part”: Possible to calculate around the particle location.
(under the Lorenz gauge condition)

“Full”: Regge-Wheeler-Zerilli formalism, Teukolsky formalism
(not the Lorenz gauge)

- 1) **Subtraction problem**: How do we subtract singular part?
We use the **spherical harmonics expansion**.
- 2) **Gauge problem**: Do we treat the gauge difference?
We consider an appropriate **gauge transformation**.

Strategy

* Schwarzschild background

Regge-Wheeler-Zerilli formalism for full metric perturbation

Sec. 2. Solution of the gauge problem: Finite gauge transformation

Sec. 3. Standard form for the regularization parameters (Singular part)

Sec. 4. New analytic regularization ($\tilde{S} + \tilde{R}$ decomposition)

[Hikida's talk]

→ Metric perturbation, Self-force

Sec. 5. Regularized self-force

Sec. 6. Discussion

2. Gauge Problem

We consider the Schwarzschild background.

$$ds^2 = - \left(1 - \frac{2M}{r}\right) dt^2 + \left(1 - \frac{2M}{r}\right)^{-1} dr^2 + r^2 (d\theta^2 + \sin^2 \theta d\varphi^2)$$

Full metric perturbation is calculated by

Regge-Wheeler-Zerilli formalism
under the Regge-Wheeler gauge condition.

We want to consider the regularization under the RW gauge analytically.

* R-part of the metric perturbation is the (homogeneous) solution of the linearized Einstein equation.

Finite gauge transformation: Gauge transformation for R-part

$$x_{\mu}^{\text{L}} \rightarrow x_{\mu}^{\text{RW}} = x_{\mu}^{\text{L}} + \xi_{\mu}^{\text{L} \rightarrow \text{RW}} [h_{\alpha\beta}^{\text{R,L}}]$$

Finite gauge transformation

Gauge transformation for **R-part**:

$$x_{\mu}^{\text{L}} \rightarrow x_{\mu}^{\text{RW}} = x_{\mu}^{\text{L}} + \xi_{\mu}^{\text{L} \rightarrow \text{RW}} [h_{\alpha\beta}^{\text{R,L}}]$$

*** We can define the (regularized) self-force under the RW gauge.**

$$\begin{aligned} F_{\alpha}^{\text{RW}}(\tau) &= \lim_{x \rightarrow z(\tau)} F_{\alpha} [h^{\text{R,RW}}] \\ &= \lim_{x \rightarrow z(\tau)} F_{\alpha} [h^{\text{R,L}} - 2 \nabla \xi^{\text{L} \rightarrow \text{RW}} [h^{\text{R,L}}]] (x) \\ &= \lim_{x \rightarrow z(\tau)} F_{\alpha} [h^{\text{full,L}} - h^{\text{S,L}} - 2 \nabla \xi^{\text{L} \rightarrow \text{RW}} [h^{\text{full,L}} - h^{\text{S,L}}]] (x) \\ &= \lim_{x \rightarrow z(\tau)} F_{\alpha} [h^{\text{full,L}} - 2 \nabla \xi^{\text{L} \rightarrow \text{RW}} [h^{\text{full,L}} \\ &\quad - h^{\text{S,L}} + 2 \nabla \xi^{\text{L} \rightarrow \text{RW}} [h^{\text{S,L}}]] (x) \\ &= \lim_{x \rightarrow z(\tau)} \left(F_{\alpha} [h^{\text{full,RW}}] (x) \right. \\ &\quad \left. - F_{\alpha} [h^{\text{S,L}} - 2 \nabla \xi^{\text{L} \rightarrow \text{RW}} [h^{\text{S,L}}]] (x) \right), \end{aligned}$$

3. Singular Part

3-1. S-part under the L gauge

S-part of the metric perturbation: calculated around the particle location.

$$\begin{aligned} \bar{h}_{\mu\nu}^{S,L}(x) = & 4\mu \left[\frac{\bar{g}_{\mu\alpha}(x, z_{\text{ret}})\bar{g}_{\nu\beta}(x, z_{\text{ret}})u^\alpha(\tau_{\text{ret}})u^\beta(\tau_{\text{ret}})}{\sigma_{;\gamma}(x, z_{\text{ret}})u^\gamma(\tau_{\text{ret}})} \right] \\ & + 2\mu(\tau_{\text{adv}} - \tau_{\text{ret}})\bar{g}_\mu{}^\alpha(x, z_{\text{ret}})\bar{g}_\nu{}^\beta(x, z_{\text{ret}})R_{\gamma\alpha\delta\beta}(z_{\text{ret}})u^\gamma(\tau_{\text{ret}})u^\delta(\tau_{\text{ret}}) \\ & + O(y^2), \end{aligned}$$

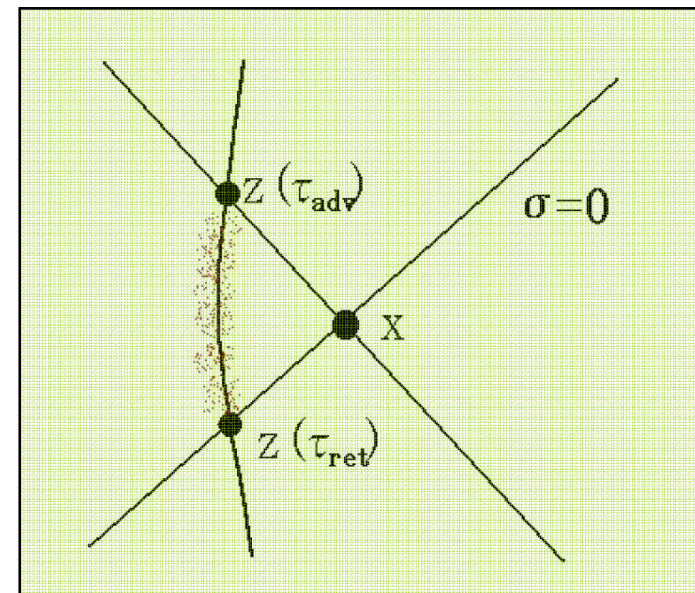
$\sigma(x, z)$: Bi-scalar of half
the squared geodesic distance

$\bar{g}_{\alpha\beta}(x, z)$: Parallel displacement bi-vector

$\tau_{\text{ret}}(x)$: Retarded time for x

y : Coordinate difference
between x and z_0 (small)

z_0 : Location of the particle



Local coordinate expansion

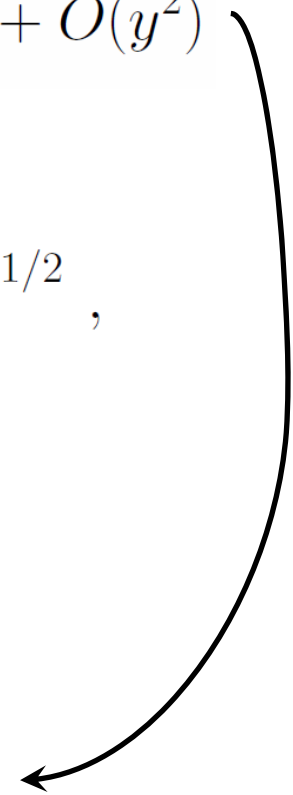
Metric component: around the particle location.

$$h_{\alpha\beta}^{\text{S,H}} = \mu \sum_{m,n,p,q,r} C_{\alpha\beta}^{m,n,p,q,r} \frac{T^m R^n \Theta^p \Phi^q}{\epsilon^r} + O(y^2)$$

Small quantities:

$$\begin{aligned} \epsilon &:= (r_0^2 + r^2 - 2 r_0 r \cos \Theta \cos \Phi)^{1/2}, \\ T &:= t - t_0, \quad R := r - r_0, \\ \Theta &:= \theta - \frac{\pi}{2}, \quad \Phi := \phi - \phi_0. \end{aligned}$$

Harmonics expansion: (Example)

$$\frac{1}{\epsilon} = \sum_{\ell m} \frac{1}{r_>} \left(\frac{r_<}{r_>} \right)^\ell Y_{\ell m}^*(\Omega_0) Y_{\ell m}(\Omega)$$


Mode sum regularization

S-part of the self-force under the Lorenz gauge:

$$F_{S,L}^{\mu(\pm)} \Big|_{\ell} = \pm A^{\mu} L + B^{\mu} + D_{\ell}^{\mu}, \quad L = \ell + \frac{1}{2}.$$

A^{μ} -term: Quadratic divergence

B^{μ} -term: Linear divergence

Standard form

* These terms are independent of ℓ .

D_{ℓ}^{μ} -term: Remaining finite contribution

$$D_{\ell}^{\mu} = \frac{d^{\mu}}{L^2 - 1} + \frac{e^{\mu}}{(L^2 - 1)(L^2 - 4)} + \frac{f^{\mu}}{(L^2 - 1)(L^2 - 4)(L^2 - 9)} + \dots$$

* D_{ℓ}^{μ} -term vanishes after summing over ℓ modes from $\ell = 0$ to ∞ .

We can consider regularization if we don't know the exact S-part.

Regularization under the L gauge

Regularization by using the S-force with the **standard form**:

$$\begin{aligned}
 F_{R,L}^\mu &= \sum_{\ell \geq 2} \left(F_{\text{full,L}}^\mu \Big|_\ell - F_{S,L}^\mu \Big|_\ell \right) \leftarrow \begin{array}{|l} \hline \text{The R-force is derived from} \\ \text{the homogeneous metric perturbation.} \\ \hline \end{array} \\
 &= \sum_{\ell \geq 2} \left(F_{\text{full,L}}^\mu \Big|_\ell - A^\mu L - B^\mu - D_\ell^\mu \right) \\
 &= \sum_{\ell \geq 2} \left(F_{\text{full,L}}^\mu \Big|_\ell - A^\mu L - B^\mu \right) + \sum_{\ell=0,1} D_\ell^\mu \\
 &= \sum_{\ell \geq 2} \left(F_{\text{full,L}}^\mu \Big|_\ell - A^\mu L - B^\mu \right) + \sum_{\ell=0,1} \left(F_{\text{full,L}}^\mu \Big|_\ell - A^\mu L - B^\mu \right) \\
 &= \sum_{\ell \geq 0} \left(F_{\text{full,L}}^\mu \Big|_\ell - A^\mu L - B^\mu \right) \cdot \quad (\text{ for } r > r_0)
 \end{aligned}$$

* We need to calculate the $\ell = 0$ and 1 modes.

3-2. About the RW gauge

Regge-Wheeler gauge condition:

* Some coefficients of the tensor harmonics expansion = 0

$$h_{2\ell m}^{S,RW} = h_{0\ell m}^{(e)S,RW} = h_{1\ell m}^{(e)S,RW} = G_{\ell m}^{S,RW} = 0.$$

Metric perturbation

$$\begin{aligned} h = \sum_{\ell m} & \left[f(r)H_{0\ell m}(t,r)\mathbf{a}_{\ell m}^{(0)} - i\sqrt{2}H_{1\ell m}(t,r)\mathbf{a}_{\ell m}^{(1)} + \frac{1}{f(r)}H_{2\ell m}(t,r)\mathbf{a}_{\ell m} \right. \\ & - \frac{i}{r}\sqrt{2\ell(\ell+1)}h_{0\ell m}^{(e)}(t,r)\mathbf{b}_{\ell m}^{(0)} + \frac{1}{r}\sqrt{2\ell(\ell+1)}h_{1\ell m}^{(e)}(t,r)\mathbf{b}_{\ell m} \\ & + \sqrt{\frac{1}{2}\ell(\ell+1)(\ell-1)(\ell+2)}G_{\ell m}(t,r)\mathbf{f}_{\ell m} + \left(\sqrt{2}K_{\ell m}(t,r) - \frac{\ell(\ell+1)}{\sqrt{2}}G_{\ell m}(t,r) \right) \mathbf{g}_{\ell m} \\ & - \frac{\sqrt{2\ell(\ell+1)}}{r}h_{0\ell m}(t,r)\mathbf{c}_{\ell m}^{(0)} + \frac{i\sqrt{2\ell(\ell+1)}}{r}h_{1\ell m}(t,r)\mathbf{c}_{\ell m} \\ & \left. + \frac{\sqrt{2\ell(\ell+1)(\ell-1)(\ell+2)}}{2r^2}h_{2\ell m}(t,r)\mathbf{d}_{\ell m} \right], \end{aligned}$$

$$f(r) = 1 - 2M/r$$

Tensor harmonics

$$\mathbf{a}_{\ell m}^{(0)} = \begin{pmatrix} Y_{\ell m} & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix},$$

$$\mathbf{a}_{\ell m}^{(1)} = (i/\sqrt{2}) \begin{pmatrix} 0 & Y_{\ell m} & 0 & 0 \\ Sym & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix},$$

$$\mathbf{a}_{\ell m} = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & Y_{\ell m} & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix},$$

$$\mathbf{b}_{\ell m}^{(0)} = ir[2\ell(\ell+1)]^{-1/2} \begin{pmatrix} 0 & 0 & (\partial/\partial\theta)Y_{\ell m} & (\partial/\partial\phi)Y_{\ell m} \\ 0 & 0 & 0 & 0 \\ Sym & 0 & 0 & 0 \\ Sym & 0 & 0 & 0 \end{pmatrix},$$

$$\mathbf{b}_{\ell m} = r[2\ell(\ell+1)]^{-1/2} \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & (\partial/\partial\theta)Y_{\ell m} & (\partial/\partial\phi)Y_{\ell m} \\ 0 & Sym & 0 & 0 \\ 0 & Sym & 0 & 0 \end{pmatrix},$$

$$\mathbf{c}_{\ell m}^{(0)} = r[2\ell(\ell+1)]^{-1/2} \begin{pmatrix} 0 & 0 & (1/\sin\theta)(\partial/\partial\phi)Y_{\ell m} & -\sin\theta(\partial/\partial\theta)Y_{\ell m} \\ 0 & 0 & 0 & 0 \\ Sym & 0 & 0 & 0 \\ Sym & 0 & 0 & 0 \end{pmatrix},$$

$$\mathbf{c}_{\ell m} = ir[2\ell(\ell+1)]^{-1/2} \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & (1/\sin\theta)(\partial/\partial\phi)Y_{\ell m} & -\sin\theta(\partial/\partial\theta)Y_{\ell m} \\ 0 & Sym & 0 & 0 \\ 0 & Sym & 0 & 0 \end{pmatrix},$$

$$\mathbf{d}_{\ell m} = ir^2[2\ell(\ell+1)(\ell-1)(\ell+2)]^{-1/2} \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & -(1/\sin\theta)X_{\ell m} & \sin\theta W_{\ell m} \\ 0 & 0 & Sym & \sin\theta X_{\ell m} \end{pmatrix},$$

$$\mathbf{g}_{\ell m} = (r^2/\sqrt{2}) \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & Y_{\ell m} & 0 \\ 0 & 0 & 0 & \sin^2\theta Y_{\ell m} \end{pmatrix},$$

$$\mathbf{f}_{\ell m} = r^2[2\ell(\ell+1)(\ell-1)(\ell+2)]^{-1/2} \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & W_{\ell m} & X_{\ell m} \\ 0 & 0 & Sym & -\sin^2\theta W_{\ell m} \end{pmatrix},$$

$$X_{\ell m} = 2\frac{\partial}{\partial\phi} \left(\frac{\partial}{\partial\theta} - \cot\theta \right) Y_{\ell m},$$

$$W_{\ell m} = \left(\frac{\partial^2}{\partial\theta^2} - \cot\theta \frac{\partial}{\partial\theta} - \frac{1}{\sin^2\theta} \frac{\partial^2}{\partial\phi^2} \right) Y_{\ell m}.$$

Gauge transformation

Generator of a gauge transformation from the L gauge to the RW gauge:

$$\xi_{\mu}^{(\text{odd})} = \sum_{\ell m} \Lambda_{\ell m}^{\text{S,L} \rightarrow \text{RW}}(t, r) \left\{ 0, 0, \frac{-1}{\sin \theta} \partial_{\phi} Y_{\ell m}(\theta, \phi), \sin \theta \partial_{\theta} Y_{\ell m}(\theta, \phi) \right\},$$

$$\xi_{\mu}^{(\text{even})} = \sum_{\ell m} \left\{ M_{0\ell m}^{\text{S,L} \rightarrow \text{RW}}(t, r) Y_{\ell m}(\theta, \phi), M_{1\ell m}^{\text{S,L} \rightarrow \text{RW}}(t, r) Y_{\ell m}(\theta, \phi), \right. \\ \left. M_{2\ell m}^{\text{S,L} \rightarrow \text{RW}}(t, r) \partial_{\theta} Y_{\ell m}(\theta, \phi), M_{2\ell m}^{\text{S,L} \rightarrow \text{RW}}(t, r) \partial_{\phi} Y_{\ell m}(\theta, \phi) \right\}.$$

$$h_{2\ell m}^{\text{S,RW}} = h_{0\ell m}^{(e)\text{S,RW}} = h_{1\ell m}^{(e)\text{S,RW}} = G_{\ell m}^{\text{S,RW}} = 0.$$

Solution:

$$\Lambda_{\ell m}^{\text{S,L} \rightarrow \text{RW}}(t, r) = \frac{i}{2} h_{2\ell m}^{\text{S,L}}(t, r),$$

$$M_{2\ell m}^{\text{S,L} \rightarrow \text{RW}}(t, r) = -\frac{r^2}{2} G_{\ell m}^{\text{S,L}}(t, r),$$

$$M_{0\ell m}^{\text{S,L} \rightarrow \text{RW}}(t, r) = -h_{0\ell m}^{(e)\text{S,L}}(t, r) - \partial_t M_{2\ell m}^{\text{S,L} \rightarrow \text{RW}}(t, r),$$

$$M_{1\ell m}^{\text{S,L} \rightarrow \text{RW}}(t, r) = -h_{1\ell m}^{(e)\text{S,L}}(t, r) - r^2 \partial_r \left(\frac{M_{2\ell m}^{\text{S,L} \rightarrow \text{RW}}(t, r)}{r^2} \right).$$

* It is **not necessary** to calculate any integration with respect to time or radial coordinates.

Gauge transformation (cont.)

$$\begin{array}{l}
 \text{odd:} \\
 \text{even:}
 \end{array}
 \Lambda_{\ell m}^{S,L \rightarrow RW}(t, r) = \frac{i}{2} h_{2\ell m}^{S,L}(t, r),$$

$$M_{2\ell m}^{S,L \rightarrow RW}(t, r) = -\frac{r^2}{2} G_{\ell m}^{S,L}(t, r),$$

$$M_{0\ell m}^{S,L \rightarrow RW}(t, r) = -h_{0\ell m}^{(e)S,L}(t, r) - \partial_t M_{2\ell m}^{S,L \rightarrow RW}(t, r),$$

$$M_{1\ell m}^{S,L \rightarrow RW}(t, r) = -h_{1\ell m}^{(e)S,L}(t, r) - r^2 \partial_r \left(\frac{M_{2\ell m}^{S,L \rightarrow RW}(t, r)}{r^2} \right).$$

Metric perturbation under the RW gauge:

$$\text{odd:} \quad h_{0\ell m}^{S,RW}(t, r) = h_{0\ell m}^{S,L}(t, r) + \partial_t \Lambda_{\ell m}^{S,L \rightarrow RW}(t, r),$$

$$h_{1\ell m}^{S,RW}(t, r) = h_{1\ell m}^{S,L}(t, r) + r^2 \partial_r \left(\frac{\Lambda_{\ell m}^{S,L \rightarrow RW}(t, r)}{r^2} \right),$$

$$\text{even:} \quad H_{0\ell m}^{S,RW}(t, r) = H_{0\ell m}^{S,L}(t, r) + \frac{2r}{r-2M} \left[\partial_t M_{0\ell m}^{S,L \rightarrow RW}(t, r) - \frac{M(r-2M)}{r^3} M_{1\ell m}^{S,L \rightarrow RW}(t, r) \right],$$

$$H_{1\ell m}^{S,RW}(t, r) = H_{1\ell m}^{S,L}(t, r) + \left[\partial_t M_{1\ell m}^{S,L \rightarrow RW}(t, r) + \partial_r M_{0\ell m}^{S,L \rightarrow RW}(t, r) - \frac{2M}{r(r-2M)} M_{0\ell m}^{S,L \rightarrow RW}(t, r) \right],$$

$$H_{2\ell m}^{S,RW}(t, r) = H_{2\ell m}^{S,L}(t, r) + \frac{2(r-2M)}{r} \left[\partial_r M_{1\ell m}^{S,L \rightarrow RW}(t, r) + \frac{M}{r(r-2M)} M_{1\ell m}^{S,L \rightarrow RW}(t, r) \right],$$

$$K_{\ell m}^{S,RW}(t, r) = K_{\ell m}^{S,L}(t, r) + \frac{2(r-2M)}{r^2} M_{1\ell m}^{S,L \rightarrow RW}(t, r),$$

* The above calculation does not include factors

$$\frac{1}{\ell(\ell+1)} \text{ and } \frac{1}{(\ell-1)\ell(\ell+1)(\ell+2)}.$$

[Interruption factor for the standard form]

Does the standard form recovers under the RW gauge?

Standard form:
$$D_\ell^\mu = \frac{d^\mu}{(2\ell - 1)(2\ell + 3)} + \frac{e^\mu}{(2\ell - 1)(2\ell + 3)(2\ell - 3)(2\ell + 5)} + \dots$$

At first glance, (last year)

S-part of the metric perturbation under the L gauge:

$$h_{0lm}^{(s)S,H}(t, r) \sim \frac{1}{l(l+1)},$$

$$h_{1lm}^{(e)S,H}(t, r) \sim \frac{1}{l(l+1)},$$

$$G_{\ell}^{S,H}(t, r) \sim \frac{1}{l(l+1)(l-1)(l+2)}.$$

These factor does not vanish in the self-force calculation under the RW gauge.

→ D-term does not vanish?

Standard form recovers under the RW gauge.

Tensor harmonics expansion:

Spherical **symmetry** of the Schwarzschild space-time

e.g.) Even parity ($t\theta$)-component of the metric perturbation $\rightarrow \partial_\theta Y_{\ell m}$

Tensor harmonics expansion of the S-part under the L gauge

$$\begin{aligned}\epsilon &= (r_0^2 + r^2 - 2 r_0 r \cos \Theta \cos \Phi)^{1/2} \\ &= \sum_{\ell m} [\text{S. F. of } D\text{-term}](r) Y_{\ell m}(\theta, \phi) Y_{\ell m}^*(\pi/2, \phi_0),\end{aligned}$$

Corresponding quantity

$$\partial_\theta \epsilon = \sum_{\ell m} [\text{S. F. of } D\text{-term}](r) \partial_\theta Y_{\ell m}(\theta, \phi) Y_{\ell m}^*(\theta_0, \phi_0).$$

$$h_{t\theta}^{(\text{even})} = \sum_{\ell m} h_{0\ell m}^{(e)}(t, r) \partial_\theta Y_{\ell m}(\theta, \phi).$$

This fact is trivial. (We can not obtain from the local quantity.)

* All components have no interruption factor ($1/\ell(\ell + 1)$ etc.).

Regularization under the RW gauge

Standard form recovers under the RW gauge.

→ $D_{\ell,RW}^\mu$ -term vanishes after summing over ℓ modes.

(The S-force is obtained in the form of the local coordinate expansion.)

Same regularization procedure as the L gauge:

$$\begin{aligned} F_{R,RW}^\mu &= \sum_{\ell \geq 2} \left(F_{\text{full},RW}^\mu \Big|_\ell - F_{S,RW}^\mu \Big|_\ell \right) \\ &= \sum_{\ell \geq 2} \left(F_{\text{full},RW}^\mu \Big|_\ell - A_{RW}^\mu L - B_{RW}^\mu - D_{\ell,RW}^\mu \right) \\ &= \sum_{\ell \geq 2} \left(F_{\text{full},RW}^\mu \Big|_\ell - A_{RW}^\mu L - B_{RW}^\mu \right) + \sum_{\ell=0,1} D_{\ell,RW}^\mu \\ &= \sum_{\ell \geq 2} \left(F_{\text{full},RW}^\mu \Big|_\ell - A_{RW}^\mu L - B_{RW}^\mu \right) + \sum_{\ell=0,1} \left(F_{\text{full},RW}^\mu \Big|_\ell - A_{RW}^\mu L - B_{RW}^\mu \right) \\ &= \sum_{\ell \geq 0} \left(F_{\text{full},RW}^\mu \Big|_\ell - A_{RW}^\mu L - B_{RW}^\mu \right) . \end{aligned}$$

* The regularization parameters for the $\ell \geq 2$ modes can be used for the $\ell = 0$ and 1 modes.

4. New Analytic Regularization

Hikida, Jhingan, Nakano, Sago, Sasaki and Tanaka ('04, '05)

[Hikida's talk]

Treatment of the Green function ($\tilde{S} + \tilde{R}$ decomposition)

$$g_{lm\omega}^{\text{full}}(r, r') = g_{lm\omega}^{\tilde{S}}(r, r') + g_{lm\omega}^{\tilde{R}}(r, r'),$$

\tilde{S} -part force: Possible to calculate for general orbits analytically.
(if we use **slow motion approximation**)

* We can extract the S-part.

\tilde{R} -part force: Generally Need to numerical calculation.

* The ℓ mode convergence is good.

$$\begin{aligned} F_{\alpha}^{\tilde{R}} &= F_{\alpha}^{\text{full}} - F_{\alpha}^{\tilde{S}} \\ &= (F_{\alpha}^{\tilde{S}} - F_{\alpha}^{\tilde{S}}) + \sum_{\ell=0}^{\ell_{\text{max}}} F_{\alpha\ell}^{\tilde{R}} \end{aligned}$$

Formulation for scalar case \rightarrow Apply to Regge-Wheeler-Zerilli function.

\tilde{S} -part

Regge-Wheeler-Zerilli formalism

→ Fourier component: include $\frac{1}{\omega}$

* Time integration is needed.

Improvement of Jhingan and Tanaka ('03)

The \tilde{S} - part of the Green function has only positive power of ω .
We can calculate the \tilde{S} - part for general orbit.

* The \tilde{R} - part of the Green function has $\ln \omega$ terms.

Third post-Newtonian (3PN) order calculation

→ \tilde{R} - part is also written by the positive power of ω .

3PN order Calculation for general orbit

\tilde{S} -part force:

$$F_\ell^{t,\tilde{S}(-)} = \mu^2 u^r \left[\frac{\ell}{r_0^2} + \frac{2(8\ell^3 + 8\ell^2 - 6\ell - 15)M^2}{r_0^4(3+2\ell)(-1+2\ell)} - \frac{(\ell^2 + \ell + 3)\delta}{r_0^2(3+2\ell)(-1+2\ell)} \right. \\ - \frac{(8\ell^3 + 15\ell^2 + \ell + 6)L_z^2}{2r_0^4(3+2\ell)(-1+2\ell)} + \frac{(8\ell^3 + 8\ell^2 - 6\ell - 15)M}{r_0^3(3+2\ell)(-1+2\ell)} \\ - \frac{(32\ell^7 + 124\ell^6 + 36\ell^5 + 233\ell^4 + 784\ell^3 - 3447\ell^2 - 3942\ell + 720)ML_z^2}{r_0^5(\ell+1)\ell(5+2\ell)(-1+2\ell)(-9+4\ell^2)} \\ - \frac{8(\ell^4 + 2\ell^3 + 4\ell^2 + 3\ell - 45)\delta M}{r_0^3(5+2\ell)(-1+2\ell)(-9+4\ell^2)} \\ + \frac{(128\ell^5 + 455\ell^4 - 50\ell^3 + 2965\ell^2 + 3702\ell - 16020)L_z^4}{8r_0^6(-1+2\ell)(5+2\ell)(-9+4\ell^2)} \\ \left. - \frac{3(\ell^4 + 2\ell^3 + 54\ell^2 + 53\ell - 180)\delta L_z^2}{r_0^4(-1+2\ell)(5+2\ell)(-9+4\ell^2)} - \frac{3(\ell^4 + 2\ell^3 + 4\ell^2 + 3\ell - 45)\delta^2}{r_0^2(5+2\ell)(-1+2\ell)(-9+4\ell^2)} \right]$$

S -part force:

$$F_\ell^{t,S(-)} = \mu^2 u^r \left[\left(\frac{L_z^4}{r_0^6} - \frac{2ML_z^2}{r_0^5} - \frac{L_z^2}{r_0^4} + \frac{1}{r_0^2} + \frac{2M}{r_0^3} + \frac{4M^2}{r_0^4} \right) L \right. \\ - \frac{2M^2}{r_0^4} - \frac{3L_z^2}{8r_0^4} - \frac{3ML_z^2}{4r_0^5} + \frac{135L_z^4}{128r_0^6} - \frac{M}{r_0^3} \\ \left. - \frac{1}{2r_0^2} - \frac{M\delta}{2r_0^3} - \frac{3L_z^2\delta}{16r_0^4} - \frac{3\delta^2}{16r_0^2} - \frac{\delta}{4r_0^2} \right]$$

$$L_z = r_0^2 u^\phi \\ \delta = 1 - 1/E^2$$

5. Regularized gravitational self-force

3PN(v^6) order regularized gravitational self-force for general orbits:
 (We consider only for the $\ell \geq 2$ modes.)

$$\begin{aligned}
 F^{t,R} &= F^{t,\tilde{S}-S} + \sum_{\ell} F_{\ell}^{t,\tilde{R}} \\
 &= \frac{\mu^2}{r_0^2} \left[\frac{64 M^3}{3 r_0^3} + \frac{160 M^2 L_z^2}{3 r_0^4} + \frac{256 M^2 \delta}{15 r_0^2} - \frac{40 M L_z^4}{r_0^5} + \frac{136 M L_z^2 \delta}{5 r_0^3} + \frac{16 M \delta^2}{5 r_0} \right] \\
 &\quad + \frac{\mu^2 u^r}{r_0^2} \left[-\frac{4 M^2}{r_0^2} - \frac{11 M L_z^2}{2 r_0^3} - \frac{2 M}{r_0} - \frac{M \delta}{r_0} + \frac{1095 L_z^4}{64 r_0^4} - \frac{3 L_z^2}{4 r_0^2} - \frac{51 L_z^2 \delta}{8 r_0^2} \right. \\
 &\quad \quad \left. - \frac{\delta}{2} - \frac{3 \delta^2}{8} \right], \\
 F^{r,R} &= F^{r,\tilde{S}-S} + \sum_{\ell} F_{\ell}^{r,\tilde{R}} \\
 &= \frac{\mu^2}{r_0^2} \left[\frac{490 M^3}{3 r_0^3} - \frac{41 M^3 \pi^2}{4 r_0^3} - \frac{68 M^2 L_z^2}{15 r_0^4} + \frac{369 L_z^2 M^2 \pi^2}{64 r_0^4} + \frac{196 M^2 \delta}{3 r_0^2} \right. \\
 &\quad - \frac{123 \delta M^2 \pi^2}{32 r_0^2} - \frac{3647 M L_z^4}{120 r_0^5} - \frac{15 M L_z^2}{r_0^3} + \frac{186 M L_z^2 \delta}{5 r_0^3} - \frac{2 M}{r_0} - \frac{3 M \delta^2}{4 r_0} \\
 &\quad - \frac{M \delta}{r_0} - \frac{27 L_z^2 \delta}{4 r_0^2} - \frac{39 L_z^2 \delta^2}{4 r_0^2} + \frac{651 L_z^4}{64 r_0^4} - \frac{\delta}{2} - \frac{5 \delta^2}{8} - \frac{7065 L_z^6}{256 r_0^6} \\
 &\quad \left. - \frac{11 \delta^3}{16} + \frac{3 L_z^2}{4 r_0^2} + \frac{1703 \delta L_z^4}{64 r_0^4} \right] \\
 &\quad + \frac{\mu^2 u^r}{r_0^2} \left[\frac{32 M^2}{3 r_0^2} + \frac{16 M L_z^2}{r_0^3} + \frac{16 M \delta}{5 r_0} \right].
 \end{aligned}$$

Circular limit

Energy and angular momentum of a point particle

$$E = \frac{1 - 2M/r_0}{\sqrt{1 - 3M/r_0}}$$
$$L = \sqrt{\frac{Mr_0}{1 - 3M/r_0}}$$

Contribution of only the $\ell \geq 2$ modes

$$F^{t,R} = -\frac{32 \mu^2 M^3}{5 r_0^5},$$
$$F^{r,R} = -\frac{3 \mu^2 M}{4 r_0^3} \left[1 - \frac{97 M}{16 r_0} + \frac{(164 \pi^2 - 739) M^2}{192 r_0^2} \right].$$

* We can **systematically** calculate higher PN order.
(same as the scalar case, 18PN or higher?)

$$\omega = m \sqrt{\frac{M}{r_0^3}}$$

6. Discussion

Gravitational self-force on a particle
around the Schwarzschild black hole (for general orbits)

Subtraction problem:

Spherical harmonics expansion

Mode sum regularization + New analytic regularization
($\tilde{S} + \tilde{R}$ decomposition)

Gauge problem:

Full metric perturbation can be obtained the RW gauge.

Gauge transformation for the R-part

→ Finite gauge transformation

Standard form recovers under the RW gauge.

To complete this self-force, we need the $\ell = 0$ and 1 modes.

* $\ell = 0$ and 1 odd modes satisfy the RW gauge automatically.

An appropriate choice of gauge

Full metric perturbation under the Zerilli gauge

→ Retarded causal boundary condition in the L gauge

Detweiler and Poisson ('04) Low multipole contribution in the L gauge

$\ell = 0$ mode

$\ell = 1$ odd mode → analytically calculated

$\ell = 1$ even mode → need numerical calculation

* S-part of the self-force (regularization parameter) for the $\ell \geq 2$ modes can be used.

Transform to the RW gauge and subtract the S-part.

Next step

- * Higher order PN calculation. (Next Capra meeting by Hikida?)

Self-force \longrightarrow dissipative / conservative part

 Adiabatic radiation reaction [Sago's talk]

- * Regularization for the master variable.
(Regge-Wheeler-Zerilli function, Teukolsky function)
- * Extend to Kerr background.
- * Second order metric perturbation and derive the wave form.
(gauge invariant)
Compare and combine with the standard PN method.