

A self-force primer

- The *Self-Force* on a “charge” arises from the interaction of the “charge” with its own “retarded field.” This includes the dissipative effects of radiation reaction and may also have some conservative effects as well.
- The acceleration from the self-force is proportional to q^2/m . For gravity this acceleration is proportional to m .
- Example: The orbital frequency of a circular Newtonian binary is

$$\Omega^2 = \frac{M}{R^3(1 + m/M)^2} \approx \frac{M}{R^3}(1 - 2m/M)$$

where R is the radius of the orbit of m , *not* the separation between the masses. The $-2m/M$ term is a consequence of the Newtonian gravitational self-force.

- For gravity, analysis of the self-force uses perturbation theory, $g_{ab} = g_{ab}^0 + h_{ab}$.

Some Details

- A *gauge choice in perturbation theory* is not just a choice of coordinates, but rather is a choice of how to associate an event on the background manifold with an event on the perturbed manifold. The change in h_{ab} from an infinitesimal change in coordinates ξ^a , is $\Delta h_{ab} - \nabla_a \xi_b - \nabla_b \xi_a$.
- Freedom in the choice of gauge adds confusion to a self-force calculation.
- The field from a charge is singular at the location of the charge — which is precisely where we need to know the field to calculate the self-force.
- *Regularizing* the field is the act of removing the singular part from the *actual* field leaving behind the regular part of the field, which is then used to calculate the self-force.
- Common ways of distinguishing between the singular and regular parts are

$$\left. \begin{array}{l} h_{ab}^{\text{act}} \\ h_{ab}^{\text{ret}} \\ h_{ab}^{\text{full}} \end{array} \right\} = \left\{ \begin{array}{ll} h^{\text{S}} & + h^{\text{R}} \\ h^{\text{direct}} & + h^{\text{tail}} \\ h^{\tilde{\text{S}}} & + h^{\tilde{\text{R}}} \end{array} \right.$$

Capra Issues

- Evolution of the Carter constant.
- Metric perturbations of the Kerr geometry.
- Gauge invariant descriptions of self-force effects.
- Formal analysis of second order metric perturbations with a point (small, black hole) source.
- A “practical” scheme for implementing the second order perturbation theory for a point source.

Capra Goals

- Better waveforms for LISA and LIGO.
- Comparison between perturbation analysis and the post-Newtonian approximation.
- Better understanding of the relativistic two body problem.