Darkonia at Goffiders

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Charmonium





Appelquist and Politzer, PRL 34, 43 (1975)

Darkonia



- * 3 unknowns: dark matter spin, binding force, mediator
- We choose dark matter to be a Dirac fermion. The scalar dark matter case can be worked out in a similar way.
- * The three lightest darkonia are

$$\eta_d = \frac{\Upsilon_d}{1^{--}} = 0^{++}$$

 h_d



Dark force

 The mediator could be scalar, pseudo-scalar or vector boson

$$\begin{aligned} \mathscr{L}_{dark} \supset &- m_{\chi} \bar{\chi} \chi - \frac{1}{2} m_{d}^{2} S^{2} - g_{d} S \bar{\chi} \chi , \qquad (\mathsf{F}_{\mathsf{S}} \mathsf{ Model}) \\ \mathscr{L}_{dark} \supset &- m_{\chi} \bar{\chi} \chi - \frac{1}{2} m_{d}^{2} P^{2} - g_{d5} P \bar{\chi} i \gamma_{5} \chi , \qquad (\mathsf{F}_{\mathsf{P}} \mathsf{ model}) \\ \mathscr{L}_{dark} \supset &- m_{\chi} \bar{\chi} \chi - \frac{1}{2} m_{d}^{2} A_{d}^{\mu} A_{d\mu} - g_{d} A_{d}^{\mu} \bar{\chi} \gamma_{\mu} \chi , \qquad (\mathsf{F}_{\mathsf{V}} \mathsf{ model}) \end{aligned}$$

$$\alpha_d \equiv g_d^2 / (4\pi) \quad (\mathbf{F}_{\mathrm{S}}, \mathbf{F}_{\mathrm{V}}) \qquad \eta_d \quad \Upsilon_d \quad h_d$$

$$V(r) = -\frac{\alpha_d}{r} e^{-m_d r} \qquad \alpha_d \equiv \frac{g_{d5}^2}{4\pi} \times \frac{m_d^2}{4m_\chi^2} \quad (\mathbf{F}_{\mathrm{P}}) \qquad \eta_d$$

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Dark P & C



 The minimal models have dark parity and dark charge conjugation symmetries (similar to G-parity for pion)

$$P_d \chi_L P_d = -\chi_R \qquad P_d \chi_R P_d = -\chi_L$$
$$C_d \chi C_d = i\gamma^2 \chi^*$$

- * For the \mathbf{F}_{S} model, η_d is $P_d\text{-}\mathrm{odd}$ and Υ_d is $C_d\text{-}\mathrm{odd}$
- * For the F_P model, η_d and P are P_d -odd
- * For the ${\rm F}_{\rm V}$ model, P_d and C_d will not play a role for phenomenology



Binding energies



$$\frac{m_d}{\alpha_d m_{\chi}/2} < 1.19$$

 $\frac{m_d}{\alpha_d m_{\chi}/2}$ < 0.22



Wave functions



Portal interactions

 In general, the mediator particle to the SM may not be the force carrier. For simplicity, we assume that the force carrier can couple to the SM

$$\mathscr{L}_{\text{portal}} = -\mu_{S} S H^{\dagger} H - \lambda_{S} S^{2} H^{\dagger} H + \mu^{2} H^{\dagger} H - \lambda (H^{\dagger} H)^{2}, \quad (\mathsf{F}_{S})$$
$$\mathscr{L}_{\text{portal}} = -\lambda_{P} P^{2} H^{\dagger} H , \quad (\mathsf{F}_{P}) \qquad \text{[conserving dark parity]}$$
$$\mathscr{L}_{\text{portal}} = -\frac{1}{2} \frac{\epsilon}{c_{W}} B_{\mu\nu} F_{d}^{\mu\nu} \qquad (\mathsf{F}_{V})$$



Bound State EFT (BSEFT)

* Similar to NRQCD, one has an EFT with the cutoff scale

$$m_{\chi} v \sim \alpha_d m_{\chi} \ll \Lambda_d \ll m_{\chi}$$

* P_d - and C_d -conserving interactions up to three boundstates

$$\begin{aligned} \mathscr{L}_{\mathrm{F}_{\mathrm{S}},\mathrm{eff}} &= \frac{1}{2} \partial_{\mu} \eta_{d} \partial^{\mu} \eta_{d} - \frac{m_{\eta_{d}}^{2}}{2} \eta_{d}^{2} - \frac{1}{4} \Upsilon_{d}^{\mu\nu} \Upsilon_{d,\mu\nu} + \frac{m_{\Upsilon_{d}}^{2}}{2} \Upsilon_{d}^{\mu} \Upsilon_{d,\mu} \\ &+ \frac{1}{2} \partial_{\mu} S \partial^{\mu} S + \frac{1}{2} \partial_{\mu} h_{d} \partial^{\mu} h_{d} - \frac{m_{S}^{2}}{2} S S - \mu_{d}^{2} g_{d} h_{d} S - \frac{m_{h_{d}}^{2}}{2} h_{d}^{2} \\ &+ \lambda_{h} S S h_{d} + \lambda_{h}' S h_{d} h_{d} + \omega_{h} S S h_{d} h_{d} + \lambda_{\eta}' S \eta_{d} \eta_{d} + \omega_{\eta} S S \eta_{d} \eta_{d} + \omega_{h\eta} S h_{d} \eta_{d} + \xi_{\eta} h_{d} \eta_{d} \eta_{d} \\ &+ \lambda_{\Upsilon}' S \Upsilon_{d}^{\mu} \Upsilon_{d,\mu} + \omega_{\Upsilon} S S \Upsilon_{d}^{\mu} \Upsilon_{d,\mu} + \xi_{\Upsilon} h_{d} \Upsilon_{d,\mu}^{\mu} \Upsilon_{d,\mu} \end{aligned}$$

Both η_d and Υ_d are stable



Bound State EFT (BSEFT)

$$\mathscr{L}_{\text{Fp,eff}} = \frac{1}{2} \partial_{\mu} \eta_{d} \partial^{\mu} \eta_{d} - \frac{m_{\eta}^{2}}{2} \eta_{d} \eta_{d} + \frac{1}{2} \partial_{\mu} P \partial^{\mu} P - \frac{1}{2} m_{P}^{2} P^{2} - \mu_{d}^{2} g_{d} \eta_{d} P$$
$$\mathscr{L}_{\text{portal}} \supset -\lambda_{P} (\cos \theta \hat{P} + \sin \theta \hat{\eta}_{d})^{2} \left(vh + \frac{1}{2}h^{2} \right)$$

The lighter one of \hat{P} and $\hat{\eta}_d$ is stable



Signatures at the LHC: $\boldsymbol{F}_{\boldsymbol{P}}$



Signatures at the LHC: $\boldsymbol{F}_{\boldsymbol{S}}$



$$pp \to h \xrightarrow{\hat{S}^*} \hat{S}\hat{h}_d \hat{h}_d \to (SM)_{\hat{S}} [\hat{S}\hat{S}]_{h_d} [\hat{S}\hat{S}]_{h_d} \to (SM)_{\hat{S}} [(SM)_{\hat{S}}(SM)_{\hat{S}}]_{\hat{h}_d} [(SM)_{\hat{S}}(SM)_{\hat{S}}]_{\hat{h}_d} \\ pp \to h \xrightarrow{\hat{S}^*} \hat{S}\hat{h}_d \hat{h}_d \to (SM)_{\hat{S}} [\eta_d \eta_d]_{h_d} [\eta_d \eta_d]_{h_d} \to (SM)_{\hat{S}} + \not{E} \\ pp \to h \xrightarrow{\text{mix}} \hat{S}\hat{h}_d \to (SM)_{\hat{S}} [\hat{S}\hat{S}]_{h_d} \to (SM)_{\hat{S}} [(SM)_{\hat{S}}(SM)_{\hat{S}}]_{\hat{h}_d} \\ pp \to h \xrightarrow{\text{mix}} \hat{S}\hat{h}_d \to (SM)_{\hat{S}} [\eta_d \eta_d]_{h_d} \to (SM)_{\hat{S}} + \not{E} \\ pp \to h \xrightarrow{s_\beta} \hat{h}_d \hat{h}_d \to [\hat{S}\hat{S}]_{h_d} [\hat{S}\hat{S}]_{h_d} \to [(SM)_{\hat{S}}(SM)_{\hat{S}}]_{\hat{h}_d} [(SM)_{\hat{S}}(SM)_{\hat{S}}]_{\hat{h}_d} \\ pp \to h \xrightarrow{s_\beta} \hat{h}_d \hat{h}_d \to [\hat{S}\hat{S}]_{h_d} [\hat{S}\hat{S}]_{h_d} \to [(SM)_{\hat{S}}(SM)_{\hat{S}}]_{\hat{h}_d} + \not{E}, \end{cases}$$

 $(\mathsf{SM})_{\hat{S}} = \{\ell^+ \ell^-, jj, \gamma\gamma\}$





$$pp \xrightarrow{1/\Lambda_d} \eta_d \gamma_d \to [(\mathrm{SM})_{\gamma_d} (\mathrm{SM})_{\gamma_d}]_{\eta} (\mathrm{SM})_{\gamma_d}$$
$$pp \xrightarrow{\epsilon} \Upsilon_d j \to [(\mathrm{SM})_{\gamma_d} (\mathrm{SM})_{\gamma_d} (\mathrm{SM})_{\gamma_d}]_{\Upsilon} j$$
$$pp \xrightarrow{\epsilon} \Upsilon_d j \xrightarrow{\Delta_{\Upsilon\eta}} [\eta_d \gamma_d^*]_{\Upsilon} j \to [[(\mathrm{SM})_{\gamma_d} (\mathrm{SM})_{\gamma_d}]_{\eta} \mathrm{SM}]_{\Upsilon} j,$$
$$(\mathrm{SM})_{\gamma_d} = \{\ell^+ \ell^-, jj\}$$

 Some light mass parameter space has been searched for at low-energy linear colliders BABAR, PRL 128 (2022) 2, 021802

Conclusions



 Dark matter could form bound states if it interacts through additional dark forces.

 Some models suggest that darkonia could remain stable not only at colliders but also on cosmological time scales.

 There are still many intriguing signatures at the LHC waiting to be explored.



Thanks!