

Exploring DSNB boosted sub-GeV dark matter: insights from XENONnT and LZ experiments



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Anirban Majumdar

(In collaboration with V. D. Romeri, D. K. Papoulias and R. Srivastava)

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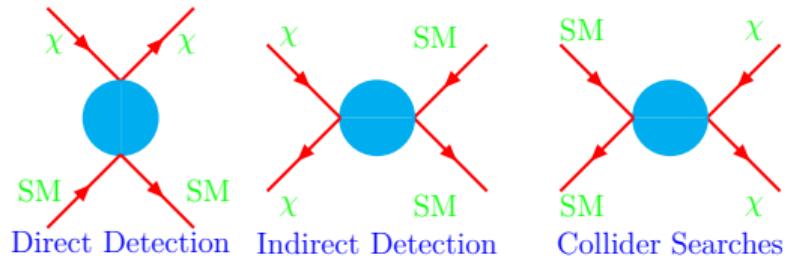
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Introduction



Motivation for searching the Dark Matter (DM)

- Cold DM: It is a non-luminous matter which occupies 27% of the mass and energy in the observable Universe. It does not interact with photons and interacts only “weakly” with ordinary matter.
- Astronomical and cosmological observations at various scales:
 - (i) Rotation curves of spiral galaxies and galaxy clusters
 - (ii) Gravitational lensing
 - (iii) Cosmic Microwave background (CMB) fluctuations



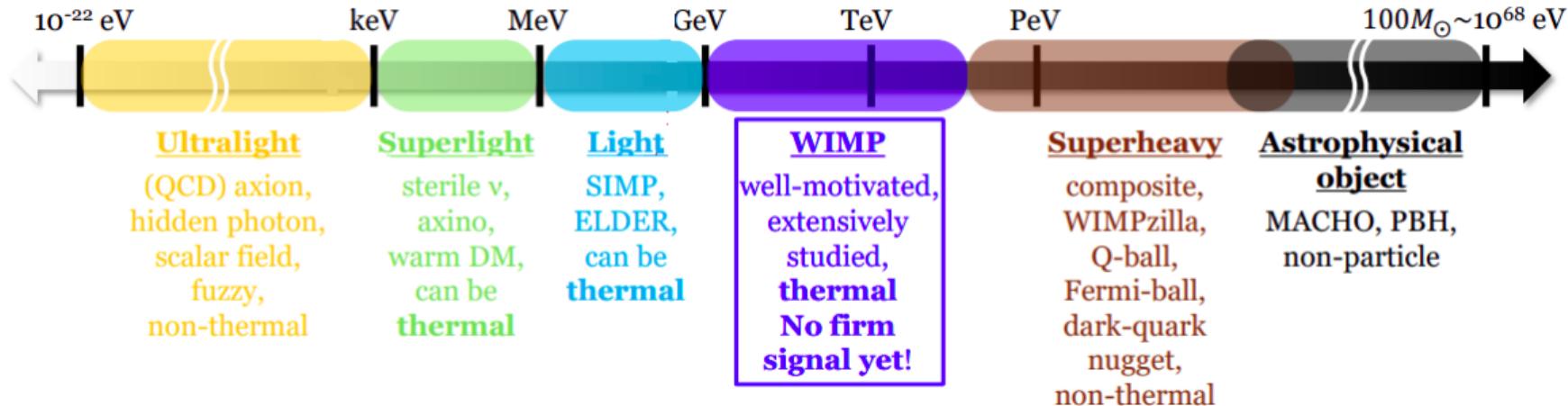
- **Direct Detection Experiments:** *XENONnT, LUX-ZEPLIN, Super-CDMS, Dark-Side, PandaX-4T, etc.*
- **Indirect Detection Experiment:** *IceCube, HESS, MAGIC, etc.*
- **Accelerator searches:** *ATLAS, CMS at CERN*

DM landscape: a wide mass range

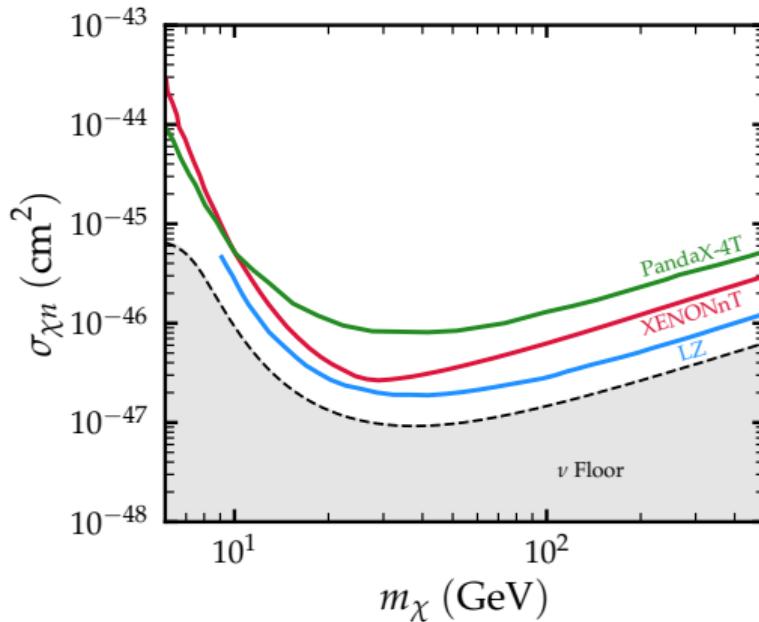


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DM landscape: a wide mass range

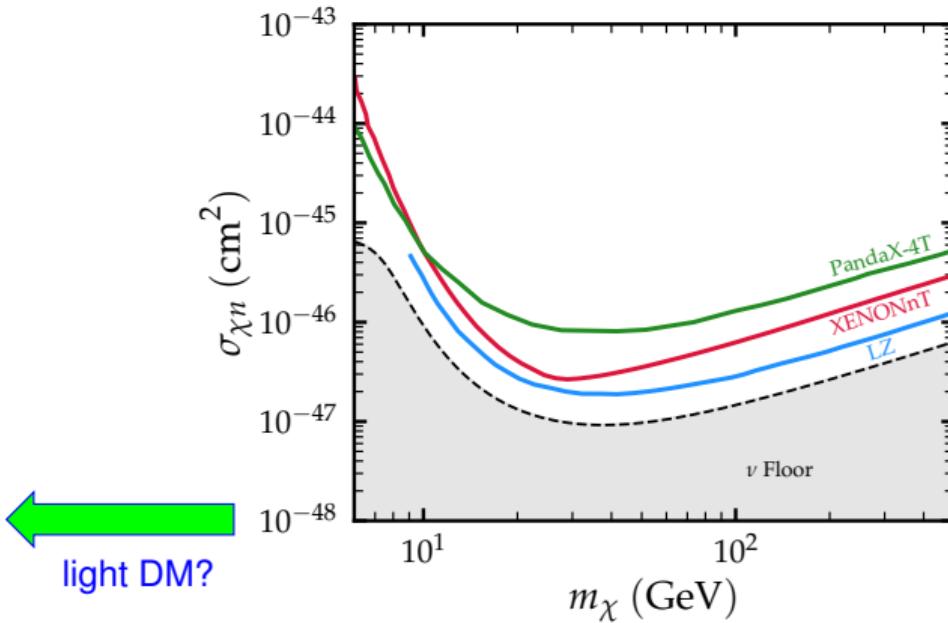


DM landscape: a wide mass range



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Thermal relic DM vs boosted DM



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The maximum recoil energy of the target:

$$T_r^{\max} \approx \frac{Q^2}{2m_T} \approx \frac{2m_X^2 m_T v_X^2}{(m_X + m_T)^2}$$

Thermal relic DM vs boosted DM

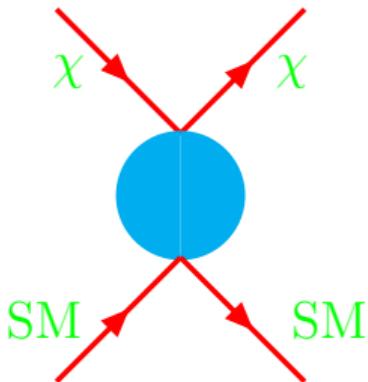


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$$T_r^{\max} \approx \frac{Q^2}{2m_T} \approx \frac{2m_\chi^2 m_T v_\chi^2}{(m_\chi + m_T)^2}$$



$$v_\chi \approx 10^{-3}c$$

Thermal relic DM vs boosted DM

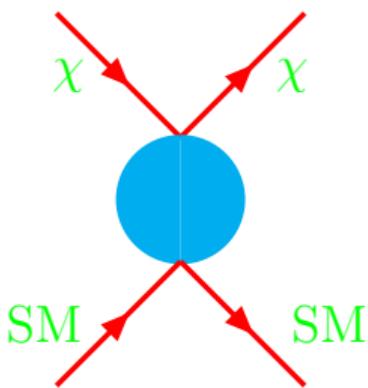


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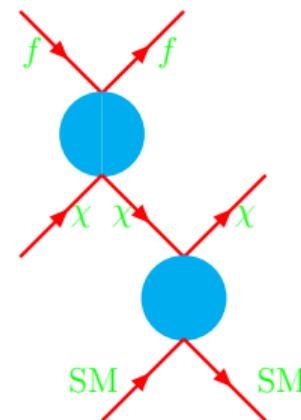
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$$v_\chi \approx 10^{-3}c$$



$$v_\chi \sim c$$

DSNB boosted dark matter

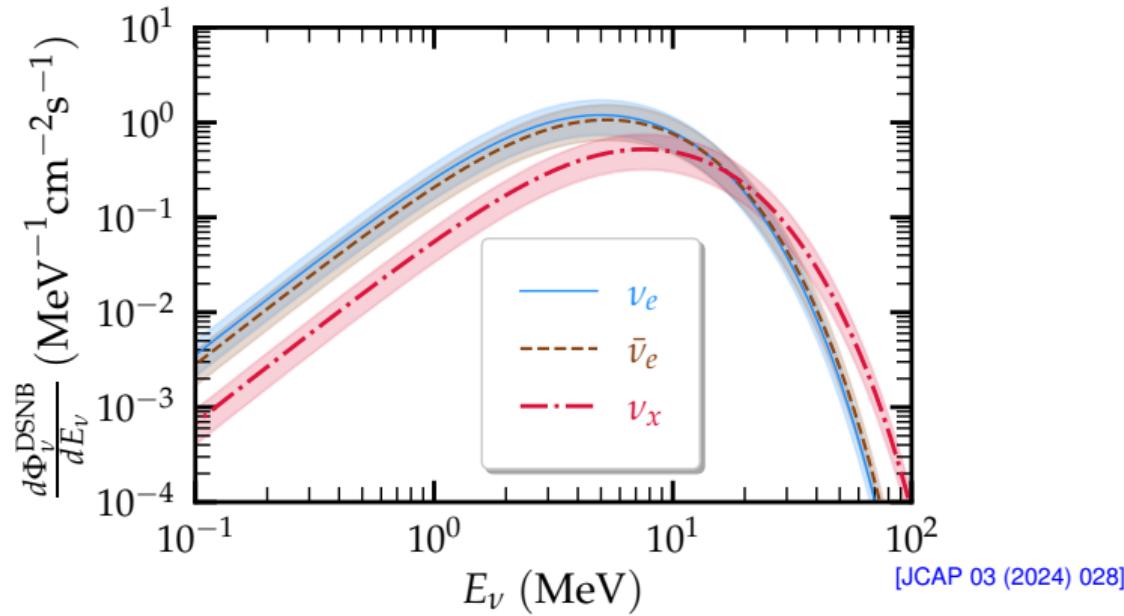
Diffuse Supernova Neutrino Background



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Right after the first star formation event, the Universe has been surrounded by an isotropic flux of MeV-energy neutrinos and antineutrinos of all flavors, produced from all supernovae events from the core-collapse explosions of huge stars throughout the Universe. This cumulative and isotropic flux of MeV neutrinos form DSNB.



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BDM Flux At The Underground Detectors

The DSNB-boosted DM differential flux,

$$\frac{d\Phi_\chi}{dT_\chi} = D_{\text{halo}} \sum_\alpha \int_{E_\nu^{\min}}^{E_\nu^{\max}} dE_\nu \frac{1}{m_\chi} \frac{d\sigma_{\nu\chi}}{dT_\chi} \frac{d\Phi_{\nu\alpha}^{\text{DSNB}}}{dE_\nu}$$

- D_{halo} encodes the line of sight integral of DM density within our galactic halo,

$$D_{\text{halo}} = \int_{\Delta\Omega} \frac{d\Omega}{4\pi} \int_0^{\ell_{\max}} \rho_{\text{NFW}}[r(\ell, \psi)] d\ell$$

- We consider Navarro-Frenk-White (NFW) profile for galactic DM density,

$$\rho_{\text{NFW}}(r) = \rho_\odot \left[\frac{r}{r_\odot} \right]^{-1} \left[\frac{1 + \frac{r_\odot}{r_s}}{1 + \frac{r}{r_s}} \right]^2$$

- For the differential cross section, we assume the constant cross section approximation, i.e.

$$\frac{d\sigma_{\nu\chi}}{dT_\chi} = \frac{\sigma_{\nu\chi}}{T_\chi^{\max}}$$

BDM Flux At The Underground Detectors



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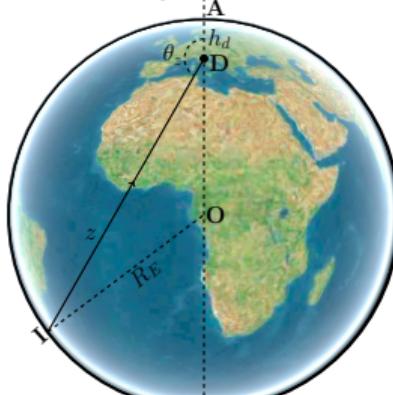
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The DSNB-boosted DM differential flux,

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DM flux gets attenuated by the elements of the atmosphere and Earth before reaching to the underground detector. However as $n_i^{\text{atm}} \ll n_i^{\text{Earth}}$, we have **neglected the attenuation due to elements of the atmosphere.**

$$\frac{dT_\chi^z}{dz} = -n_i \int_0^{T_i^{\max}(T_\chi^z)} \frac{d\sigma_{\chi i}}{dT_i} T_i dT_i$$



$$\frac{d\sigma_{\chi i}}{dT_i} = \frac{\sigma_{\chi i}}{T_i^{\max}}$$

$i \equiv \{e, N\}$

BDM Flux At The Underground Detectors



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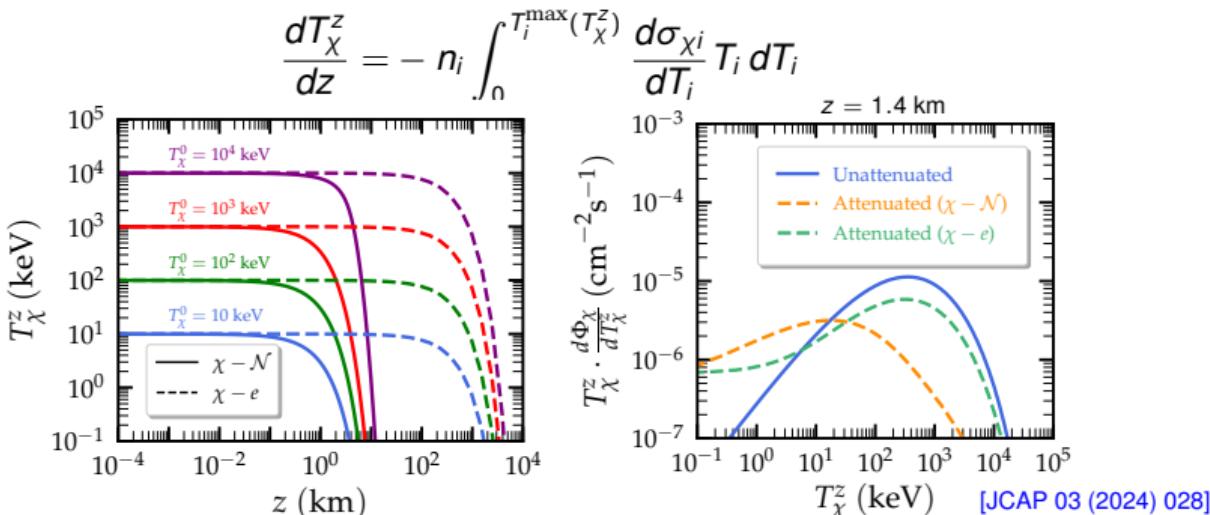
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$$\sigma_{\nu\chi} = \sigma_{\chi e} = \sigma_{\chi n} = 10^{-29} \text{ cm}^2$$

$$m_\chi = 300 \text{ MeV}$$



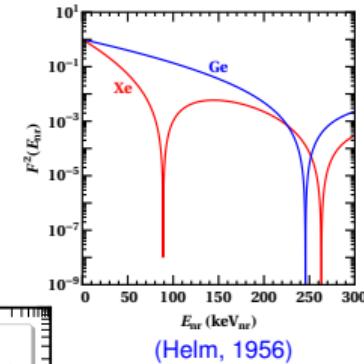
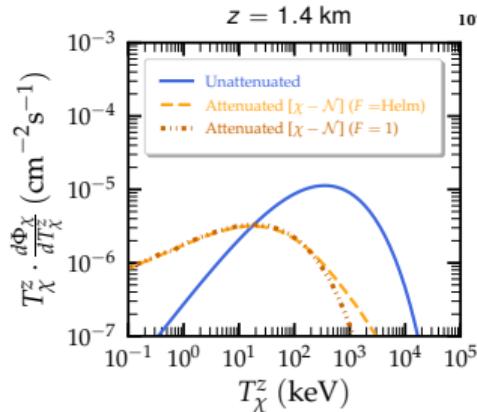
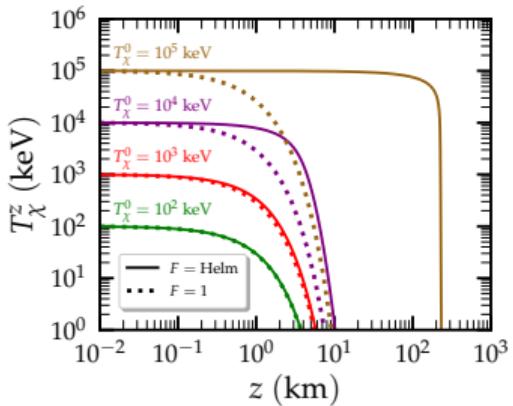
Implications of nuclear form factor



The spin independent DM-nuclei scattering cross section can be written as,

$$\sigma_{\chi N}^{\text{SI}}(q^2) = \frac{\mu_{\chi N}^2}{\mu_{\chi n}^2} A^2 \sigma_{\chi n} F^2(q^2)$$

$$\begin{aligned}\sigma_{\nu\chi} &= \sigma_{\chi e} = \sigma_{\chi n} = 10^{-29} \text{ cm}^2 \\ m_\chi &= 300 \text{ MeV}\end{aligned}$$



(Helm, 1956)

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DM signal at the underground detectors



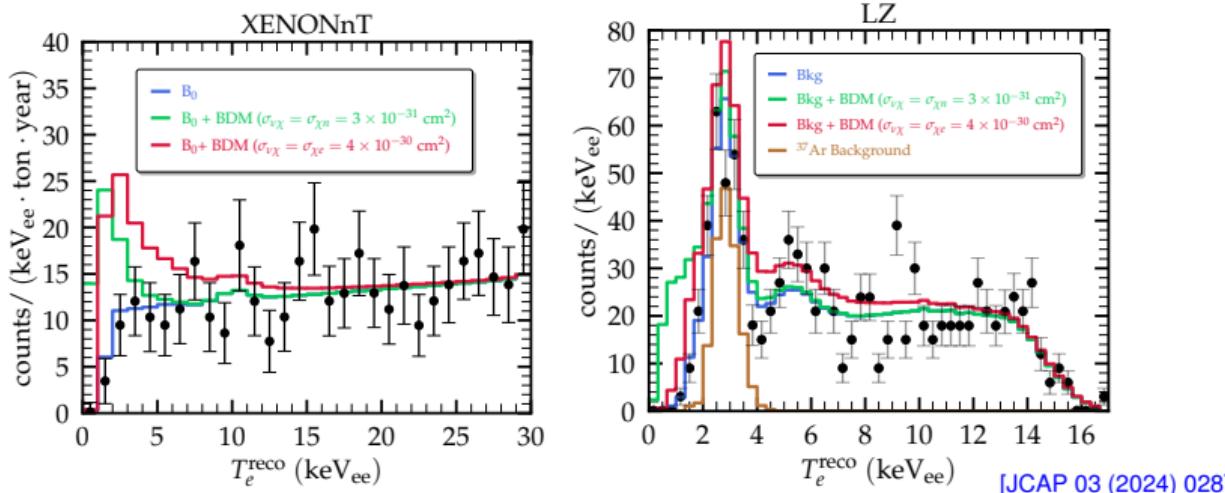
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After reaching the underground detector, the DSNB-boosted DM can scatter off both the electrons and nuclei of the target material, triggering both electronic and nuclear recoils. The differential event rate with respect to the recoil energy T_i can be written as,

$$\frac{dR}{dT_i} = t_{\text{run}} N_{\text{target}}^i \mathcal{A} \int dT_X^z \frac{d\Phi_X}{dT_X^z} \frac{d\sigma_{\chi i}}{dT_i}$$

$$m_\chi = 300 \text{ MeV}$$

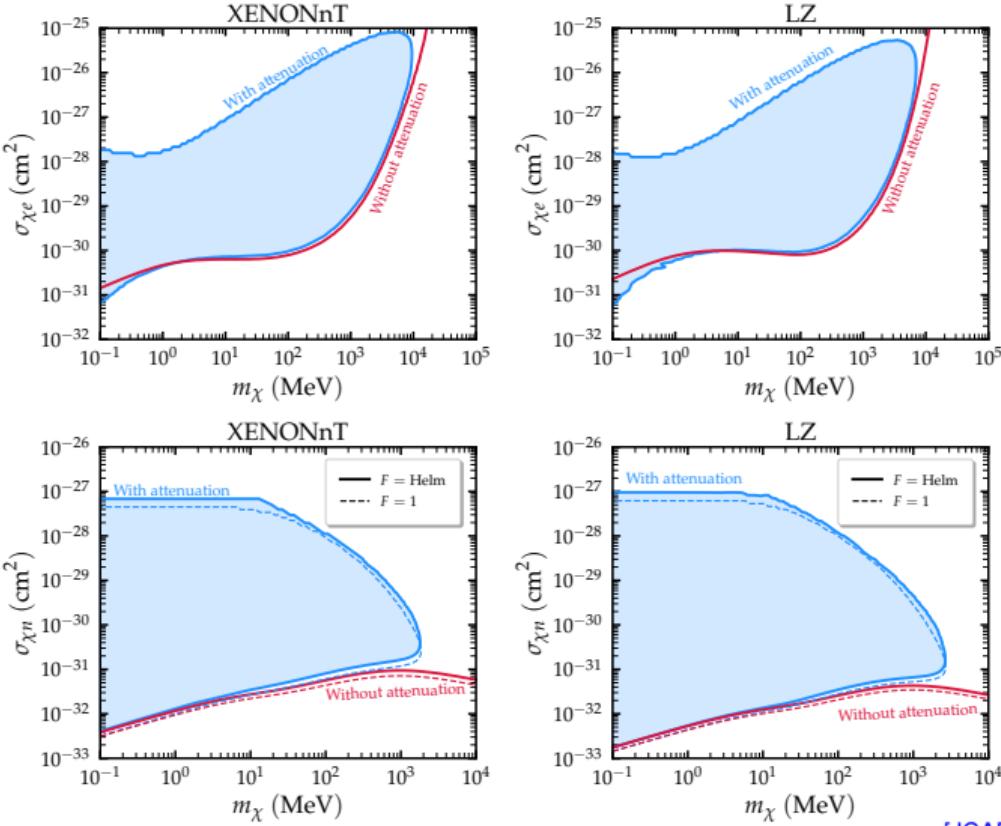


Effect of Earth attenuation in the resulting limits



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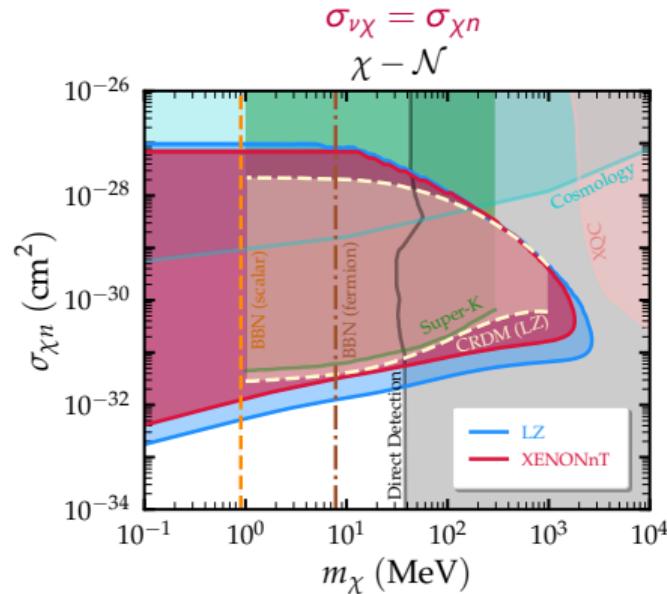
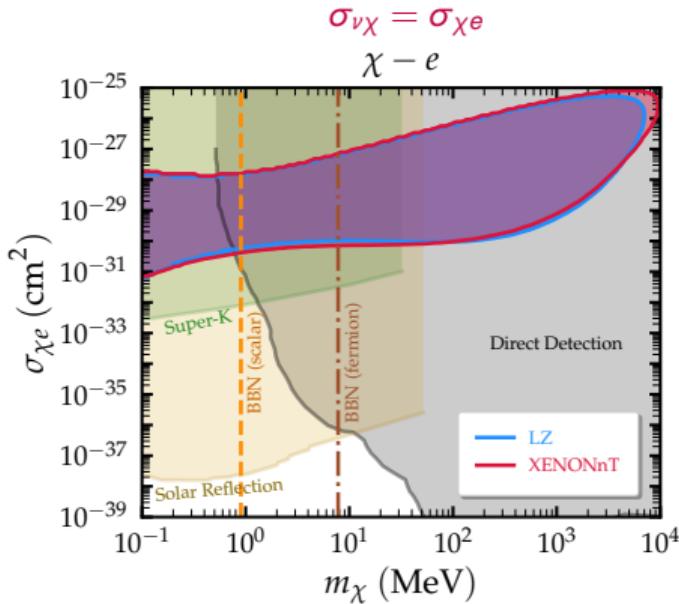
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Resulting Limits



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Conclusions

- ☞ DSNB Boosted DM produces a subdominant, semi-relativistic component of Galactic DM.
- ☞ Consideration of Earth attenuation is crucial for accurate interpretation of experimental results.
- ☞ *Although a significant part of our constraints lie in a region of parameter space already probed by other searches, these results highlight the complementarity and significance of the LZ and XENONnT data in probing the sub-GeV DM parameter space.*

THANK YOU

Extras

χ^2 function utilized



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- For the analysis of LZ data, we have performed a spectral analysis using the following Poissonian χ^2 function

$$\begin{aligned}\chi^2(\vec{\mathcal{S}}; \alpha, \beta, \delta) = & 2 \sum_{i=1}^{51} \left[R_{\text{pred}}^i(\vec{\mathcal{S}}; \alpha, \beta, \delta) - R_{\text{exp}}^i + R_{\text{exp}}^i \ln \left(\frac{R_{\text{exp}}^i}{R_{\text{pred}}^i(\vec{\mathcal{S}}; \alpha, \beta, \delta)} \right) \right] + \\ & \left(\frac{\alpha}{\sigma_\alpha} \right)^2 + \left(\frac{\beta}{\sigma_\beta} \right)^2 + \left(\frac{\delta}{\sigma_\delta} \right)^2,\end{aligned}$$

- The following Gaussian χ^2 function is used for the analysis of XENONnT data

$$\chi^2(\vec{\mathcal{S}}; \beta) = \sum_{i=1}^{30} \left(\frac{R_{\text{pred}}^i(\vec{\mathcal{S}}; \beta) - R_{\text{exp}}^i}{\sigma^i} \right)^2 + \left(\frac{\beta}{\sigma_\beta} \right)^2$$

Geophysical properties of Earth

We model the Earth's interior as a sphere of constant electron and nuclear densities ($n_e = 8 \times 10^{23} \text{ cm}^{-3}$ and $n_N = 3.44 \times 10^{22} \text{ cm}^{-3}$), based on the abundances of the main elements as shown in following table.

Element	Mass Number (A)	Relative Abundance (%)	$n_N (\text{cm}^{-3})$
Fe	55.845	32.1	6.11×10^{22}
O	15.999	30.1	3.45×10^{22}
Si	28.086	15.1	1.77×10^{22}
Mg	24.305	13.9	1.17×10^{22}
S	32.065	2.9	2.33×10^{21}
Ca	40.078	1.5	7.94×10^{20}
Al	26.982	1.4	1.09×10^{21}