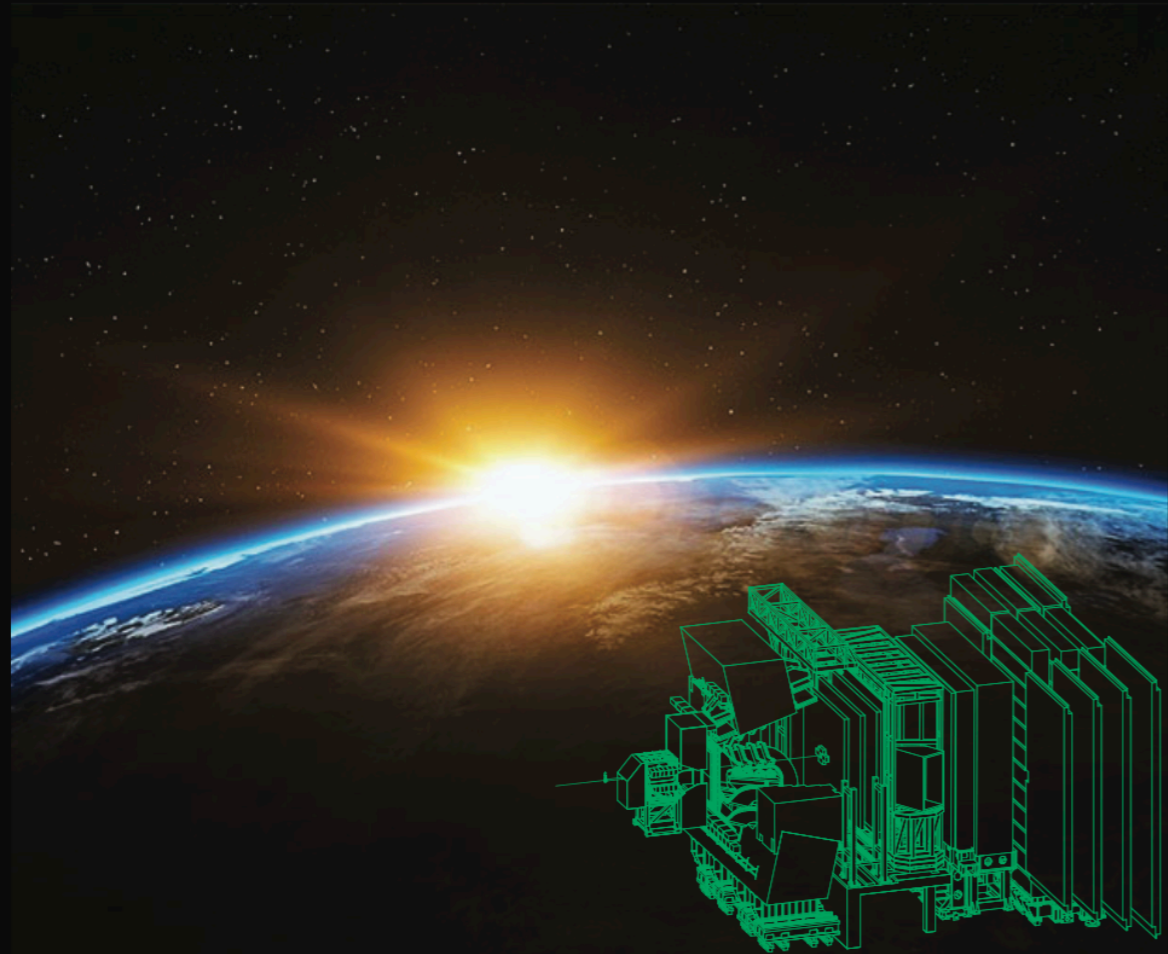




CERN/LHCC 2021-012  
LHCb TDR 23  
24 February 2022

# Framework LHCb UPGRADE II TDR



**Technical Design Report**

[LHCC-2021-012](#)



## Physics benchmarking for Scoping Document

*Matteo  
(INFN Frascati)*

**March 6<sup>th</sup> 2024**

[Upgrade 2 TWIKI](#)

## **1) *Luminosity scenarios***

- *quick recap of what we know from LHC*
- *present lumi figures for our 3 scenarios*



# Main LHC parameters

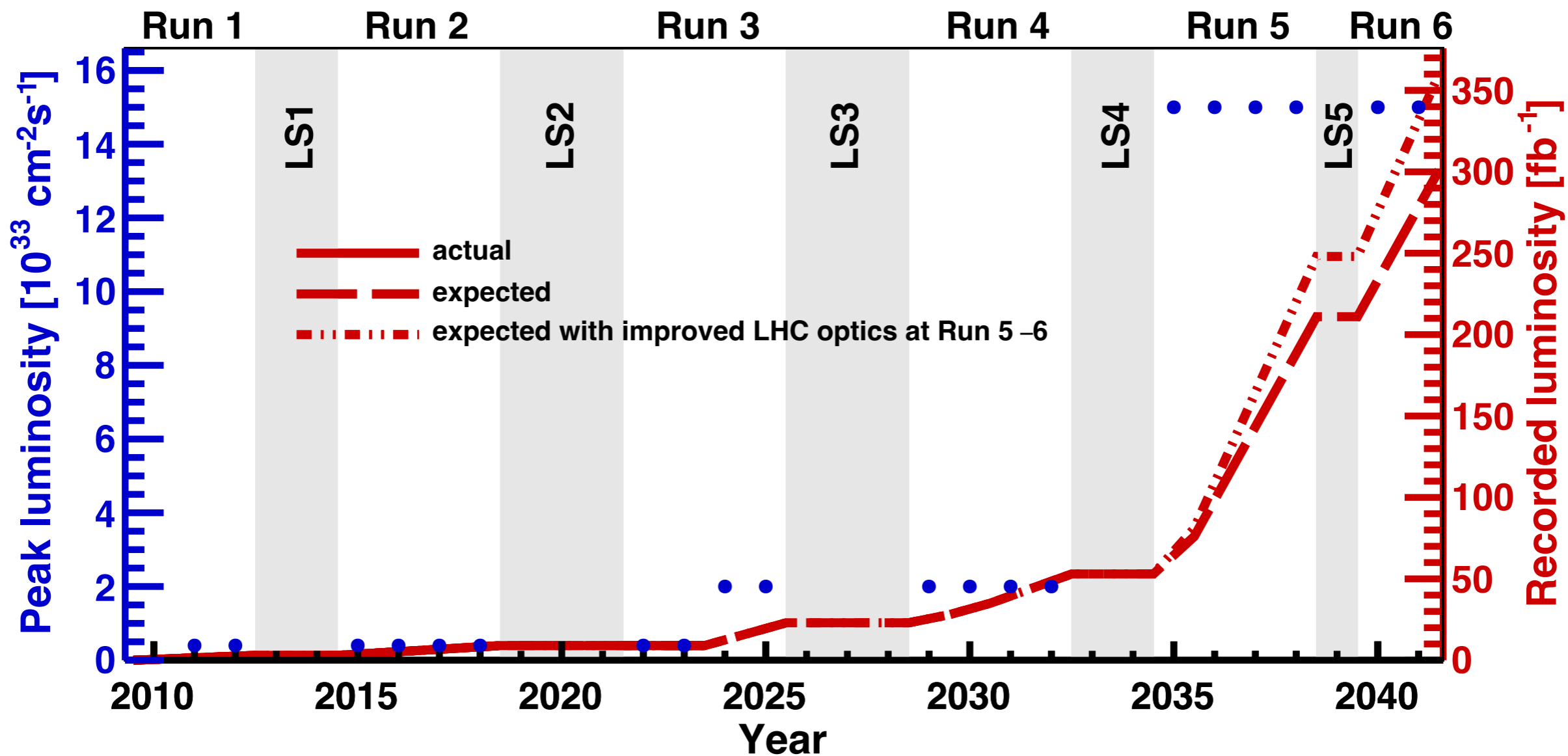
- **Flat optics** (see R. De Maria at LHCb week) would be our target for Scoping Document
  - > for the moment round optics still remains baseline for the machine, but studies are planned on flat optics
  - > we need to push for swapping the priority (but remain flexible in detector design)
- **Electron cloud:** baseline assumption from machine is to solve it
  - > 2574 bunches for LHCb (vs 2748 for ATLAS/CMS)
- **Luminosity region length** with flat optics: ~36mm (~45mm with round beams)
  - > less separation btw PVs
- **Machine operations:** 160 pp days/y (average Run5/6 with HI and YETS=15w), fill duration 8h + 3h turn-around, 50% operational efficiency
- **Effi(LHCb) for Upgrade II:** commissioning year at 50% + 5 nominal years at 90%
- **Recorded  $L_{int}$  (Run 1 - 4) ~53 fb<sup>-1</sup>**

**LHC note in preparation**

**we need it!!**

# The lumi plot

And this is what we get with levelled  $L_{peak} = 1.5 \times 10^{34} \text{ cm}^{-2}\text{s}^{-1}$



Upgrade 2 TWIKI



# Lumi scenarios: all numbers

All numbers to be checked with machine experts!!!

## Round beams

	Low	Medium-a	Medium-b	High
levelled $L_{peak}$ ( $cm^{-2}s^{-1}$ )	<b><math>1.0 \times 10^{34}</math></b>	$1.2 \times 10^{34}$	<b><math>1.3 \times 10^{34}</math></b>	<b><math>1.5 \times 10^{34}</math></b>
levelled pile-up	<b>28</b>	34	<b>36</b>	<b>42</b>
levelling time (h)	<b>3.6</b>	2.9	<b>2.3</b>	<b>1.3</b>
<b>delivered</b> $L_{int}/y$ ( $fb^{-1}$ )	<b>42.5</b>	47.0	<b>48.0</b>	<b>49.9</b>
<b>delivered</b> Run 5-6 ( $fb^{-1}$ )	<b>234</b>	259	<b>264</b>	<b>274</b>
ratio $L_{int}(X) / L_{int}(high)$	<b>0,852</b>	0,936	<b>0,962</b>	<b>1.0</b>
<b>total recorded</b> $L_{int}$ Run 1-6	<b>263</b>	286	<b>291</b>	<b>300</b>

for detector

for physics

## Flat beams

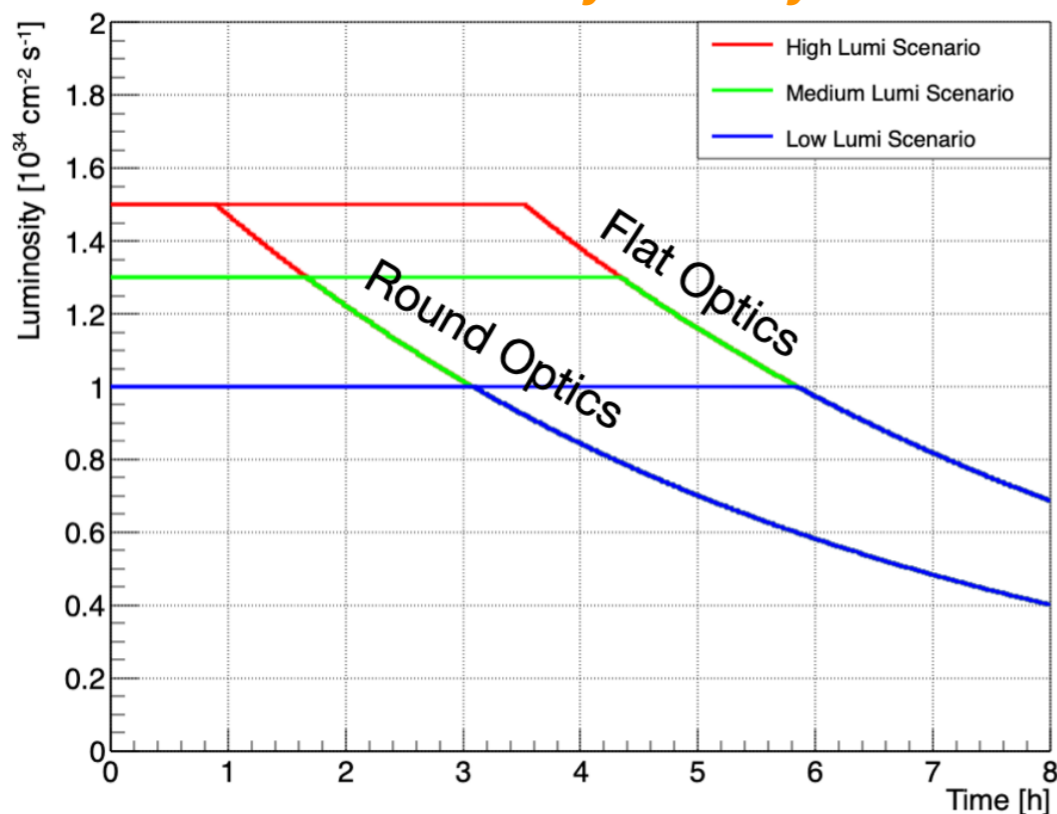
	Low	Medium-a	Medium-b	High
levelled $L_{peak}$ ( $cm^{-2}s^{-1}$ )	<b><math>1.0 \times 10^{34}</math></b>	$1.2 \times 10^{34}$	<b><math>1.3 \times 10^{34}</math></b>	<b><math>1.5 \times 10^{34}</math></b>
levelled pile-up	<b>28</b>	34	<b>36</b>	<b>42</b>
levelling time (h)	<b>5.7</b>	4.7	<b>4.2</b>	<b>3.4</b>
<b>delivered</b> $L_{int}/y$ ( $fb^{-1}$ )	<b>47.8</b>	54.3	<b>57.1</b>	<b>61.9</b>
<b>delivered</b> Run 5-6 ( $fb^{-1}$ )	<b>263</b>	299	<b>314</b>	<b>340</b>
ratio $L_{int}(X) / L_{int}(high)$	<b>0,772</b>	0,877	<b>0,922</b>	<b>1.0</b>
<b>total recorded</b> $L_{int}$ Run 1-6	<b>290</b>	322	<b>336</b>	<b>359</b>

for detector

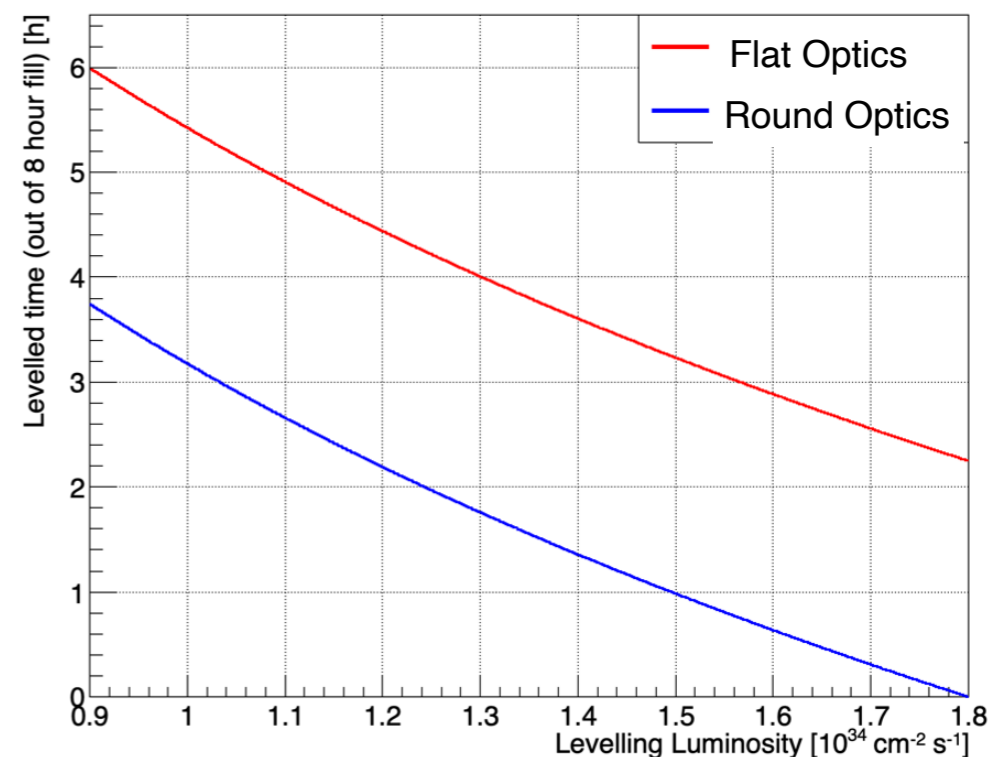
for physics

Scenarios at 1.2 and 1.3 have a small ~5% difference in lumi/occupancy conditions, but 1.3 looks better for FLAT optics, since it makes a better usage of machine improvement → **we've chosen 1.3**

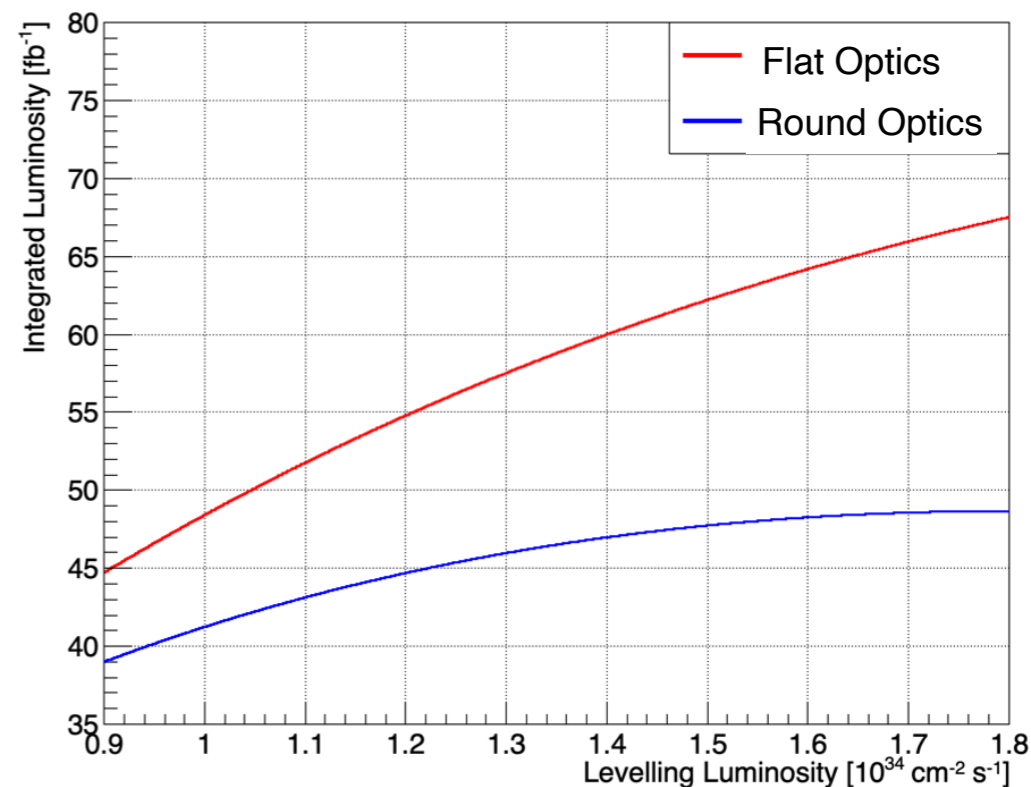
*luminosity decay*



*levelled time vs levelling lumi*



*Integrated luminosity per year*



*Due to its higher virtual peak luminosity ( $1.8 \rightarrow 2.8 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$ ), **flat optics** has much longer levelling times, higher integrated luminosity, and more stable data taking conditions*

# Relevant numbers for physics

	<i>Low</i>	<i>Medium</i>	<b>High</b>
<i>levelled <math>L_{peak}</math> (<math>cm^{-2}s^{-1}</math>)</i>	$1.0 \times 10^{34}$	$1.3 \times 10^{34}$	<b><math>1.5 \times 10^{34}</math></b>
<b><i>total recorded ROUND (<math>fb^{-1}</math>)</i></b>	263	291	<b>300</b>
<b><i>total recorded FLAT (<math>fb^{-1}</math>)</i></b>	290	336	<b>359</b>

- **Round optics** has a larger risk concerning the final integrate lumi target
- **Flat optics** provides a solid picture for reaching the Upgrade II minimal target of  $300 fb^{-1}$ , with operational risk considerably reduced (very valuable argument in the evaluation phase of our project)

*To make best use of it, however, we need a detector capable of running at highest possible peak luminosity, and to stand the slightly higher radiation dose*

*Within the timeline of Scoping Document we will not be certain of final beam configuration, but no show-stoppers are expected for FLAT optics, **so detector specifications need to be compliant with this***

*In summary, we need to define medium and low scenarios at 1.3 and 1.0, respectively, while matching the target cost savings of 15% and 30% of FTDR*

## **2) *Physics benchmarking***





# General considerations

- **High scenario:** Full breath of our physics programme, with highest possible margins on luminosity.
- **Medium scenario:** try to preserve as much as possible the physics programme, small loss in luminosity. But need to decrease some of the detector features in order to generate the target cost decrease.
- **Low scenario:** we should propose something that, even sacrificing a fraction of our physics programme, still allows key LHCb flagship measurements, in particular CKM phases, charm CPV and rare muons

*In addition: feedback also received by the CERN Scientific Policy Committee*

(december 2023)

Request to better understand which physics goals are truly unique in a global context, and what elements of the proposed upgrade are essential to address them

# Physics studies: what we need

This is how we presented so far our Upgrade II case

Observable	Current LHCb (up to 9 fb <sup>-1</sup> )	Upgrade I (23 fb <sup>-1</sup> )	Upgrade I (50 fb <sup>-1</sup> )	Upgrade II (300 fb <sup>-1</sup> )
<b>CKM tests</b>				
$\gamma (B \rightarrow DK, \text{ etc.})$	4° [9, 10]	1.5°	1°	0.35°
$\phi_s (B_s^0 \rightarrow J/\psi\phi)$	32 mrad [8]	14 mrad	10 mrad	4 mrad
$ V_{ub} / V_{cb}  (\Lambda_b^0 \rightarrow p\mu^-\bar{\nu}_\mu, \text{ etc.})$	6% [29, 30]	3%	2%	1%
$a_{\text{sl}}^d (B^0 \rightarrow D^-\mu^+\nu_\mu)$	$36 \times 10^{-4}$ [34]	$8 \times 10^{-4}$	$5 \times 10^{-4}$	$2 \times 10^{-4}$
$a_{\text{sl}}^s (B_s^0 \rightarrow D_s^-\mu^+\nu_\mu)$	$33 \times 10^{-4}$ [35]	$10 \times 10^{-4}$	$7 \times 10^{-4}$	$3 \times 10^{-4}$
<b>Charm</b>				
$\Delta A_{CP} (D^0 \rightarrow K^+K^-, \pi^+\pi^-)$	$29 \times 10^{-5}$ [5]	$13 \times 10^{-5}$	$8 \times 10^{-5}$	$3.3 \times 10^{-5}$
$A_\Gamma (D^0 \rightarrow K^+K^-, \pi^+\pi^-)$	$11 \times 10^{-5}$ [38]	$5 \times 10^{-5}$	$3.2 \times 10^{-5}$	$1.2 \times 10^{-5}$
$\Delta x (D^0 \rightarrow K_S^0\pi^+\pi^-)$	$18 \times 10^{-5}$ [37]	$6.3 \times 10^{-5}$	$4.1 \times 10^{-5}$	$1.6 \times 10^{-5}$
<b>Rare Decays</b>				
$\mathcal{B}(B^0 \rightarrow \mu^+\mu^-)/\mathcal{B}(B_s^0 \rightarrow \mu^+\mu^-)$	69% [40, 41]	41%	27%	11%
$S_{\mu\mu} (B_s^0 \rightarrow \mu^+\mu^-)$	—	—	—	0.2
$A_T^{(2)} (B^0 \rightarrow K^{*0}e^+e^-)$	0.10 [52]	0.060	0.043	0.016
$A_T^{\text{Im}} (B^0 \rightarrow K^{*0}e^+e^-)$	0.10 [52]	0.060	0.043	0.016
$\mathcal{A}_{\phi\gamma}^{\Delta\Gamma} (B_s^0 \rightarrow \phi\gamma)$	$\begin{matrix} +0.41 \\ -0.44 \end{matrix}$ [51]	0.124	0.083	0.033
$S_{\phi\gamma} (B_s^0 \rightarrow \phi\gamma)$	0.32 [51]	0.093	0.062	0.025
$\alpha_\gamma (\Lambda_b^0 \rightarrow \Lambda\gamma)$	$\begin{matrix} +0.17 \\ -0.29 \end{matrix}$ [53]	0.148	0.097	0.038
<b>Lepton Universality Tests</b>				
$R_K (B^+ \rightarrow K^+\ell^+\ell^-)$	0.044 [12]	0.025	0.017	0.007
$R_{K^*} (B^0 \rightarrow K^{*0}\ell^+\ell^-)$	0.12 [61]	0.034	0.022	0.009
$R(D^*) (B^0 \rightarrow D^{*-}\ell^+\nu_\ell)$	0.026 [62, 64]	0.007	0.005	0.002

1) Key observables in flavour physics

2) NOT ONLY flavour physics

LHCb, as a general purpose detector in forward region, will keep pursuing an ambitious programme in spectroscopy, EW precision and Higgs physics ( $\sim 2 - 3 \times y_{SM}^c$ ), dark sector and other exotic searches, heavy ions and fixed target physics ...

We need some quantitative statements (in the different detector scenarios) about some of the observables belonging to the table above, plus some discussion about the opportunities of extra-physics accessible (or lost) in a given scenario

# Scoping Document: the physics content

<b>1 Introduction</b>	<b>15p</b>	
1.1 Summary of physics programme	.....	←
1.2 Machine considerations	.....	
1.3 Detector overview	.....	
<b>2 Scoping scenarios</b>	<b>30p</b>	
2.1 General considerations	.....	
2.2 Detector options	.....	
2.3 Summary of scenarios and costs	.....	
<b>3 Impact on the LHCb Physics Programme</b>	<b>20p</b>	
3.1 Physics object performance	.....	←
3.2 Physics potential on selected channels	.....	←
<b>4 Project management and Schedule</b>	<b>15p</b>	
4.1 Project organisation	.....	
4.2 Schedule	.....	
<b>5 Summary</b>		
<b>A Options considered for the subdetectors</b>	<b>5p each</b>	
A.1 VELO	.....	
A.2 Upstream Tracker	.....	
A.3 Mighty Tracker	.....	
A.4 Magnet Stations	.....	
A.5 RICH Detectors	.....	
A.6 TORCH	.....	
A.7 PicoCal	.....	
A.8 Muon stations	.....	
A.9 RTA	.....	
A.10 Online	.....	
A.11 Infrastructure	.....	

**Sec. 1.1:** give the global context, remind again about unique LHCb capabilities and wide physics breath

**Sec. 3.1:** tracking and PID performance in different scenarios  
~mid April

**Sec. 3.2:** quantitative statements on how the above performance reflect into some physics channels, plus info on extra physics opportunities  
~September

*NB: both LHCC and FAs are expecting to see discussed quantitatively some physics examples, which they requested explicitly multiple times; they will not argue about the methodology we used to extract numbers, but they want to see some numbers*

## Proposed list of channels to study in full simulation

- $B^+ \rightarrow D[\rightarrow K_S^0 \pi^+ \pi^-] K^+$   
 $K_S^0$  reconstruction, PID ( $K/\pi$ ),  $\gamma$
- $B_s^0 \rightarrow J/\psi[\rightarrow \mu^+ \mu^-] \phi[\rightarrow K^+ K^-]$   
muons, PID ( $K/\pi$ ), flavour tagging, vertexing,  $\phi_s$
- $B^0 \rightarrow K^{*0}[\rightarrow K^+ \pi^-] e^+ e^-$  [low  $q^2$ ]  
electron reconstruction, acceptance, PID ( $e$ ),  $C_7^{(f)}$
- $\Lambda_b^0 \rightarrow \Lambda \gamma$   
non-trivial vertexing, photon reconstruction, acceptance,  $C_7^{(f)}$
- $D^{*+} \rightarrow D[\rightarrow K^+ K^-] \pi^+$   
slow  $\pi$  acceptance, PID ( $K/\pi$ ),  $A_\Gamma$
- $\Lambda_b^0 \rightarrow p \mu \nu$   
single muon, proton ID, missing energy/corrected mass,  $|V_{ub}|$

See Tim's presentations at [PPG](#) october 12 and in the [U2PG](#) october 18

[PPG](#) also produced a document with valuable inputs, available [here](#)

***This list does not define our physics programme. Rather it will be used to demonstrate the impact of different scoping scenarios. There's the attempt to cover a wide spectrum of different detector and reconstruction features***

*Other inputs are welcome: detectors will add their own specific channels, specific studies will be done for other physics topics, see e.g. [U2PG](#) on Heavy Ions nov 15*



# Planning the work

*Volunteers from the analysis groups kindly accepted to follow this task, but this requires strict coordination with people performing detector studies*

*~monthly U2PG meetings*

*Last meeting  
1 March*

<https://indico.cern.ch/event/1378464/>

## Round table on physics channels (short contributions discussing options/plans)


$B \rightarrow D(K\pi\pi)K$

**Speakers:** Resmi Puthumanai (University of Oxford (GB)), Sneha Sirirshkumar Malde (University of Oxford (GB))

 Update\_Resmi.pdf


$V_{ub}$

**Speakers:** Patrick Haworth Owen (University of Zurich (CH)), Ulrik Egede (Monash University (AU)), William Sutcliffe (University of Zurich (CH))

 Vub\_U2\_Update.pdf


$\phi_s$

**Speakers:** Peilian Li (CERN), Sara Celani (Heidelberg University (DE)), Veronika Georgieva Chobanova (University of A Coruna - UDC (ES))

 Phis\_u2pg.pdf

$B \rightarrow K^*e^+e^-$

**Speakers:** Marie-Helene Schune (Université Paris-Saclay (FR)), Martino Borsato (Universita & INFN, Milano-Bicocca (IT))


 U2PG\_B2KstEE\_202...

$\Lambda_b \rightarrow \Lambda\gamma$

**Speakers:** Arantza De Oyanguren Campos (Univ. of Valencia and CSIC (ES)), Miriam Calvo Gomez (La Salle, Ramon Llull University (ES))

Charm CPV

**Speakers:** Ao Xu (Universita & INFN Pisa (IT)), Michael J. Morello (SNS and INFN-Pisa (IT)), Tommaso Pajero (CERN)

 axu\_u2\_agamma\_2...

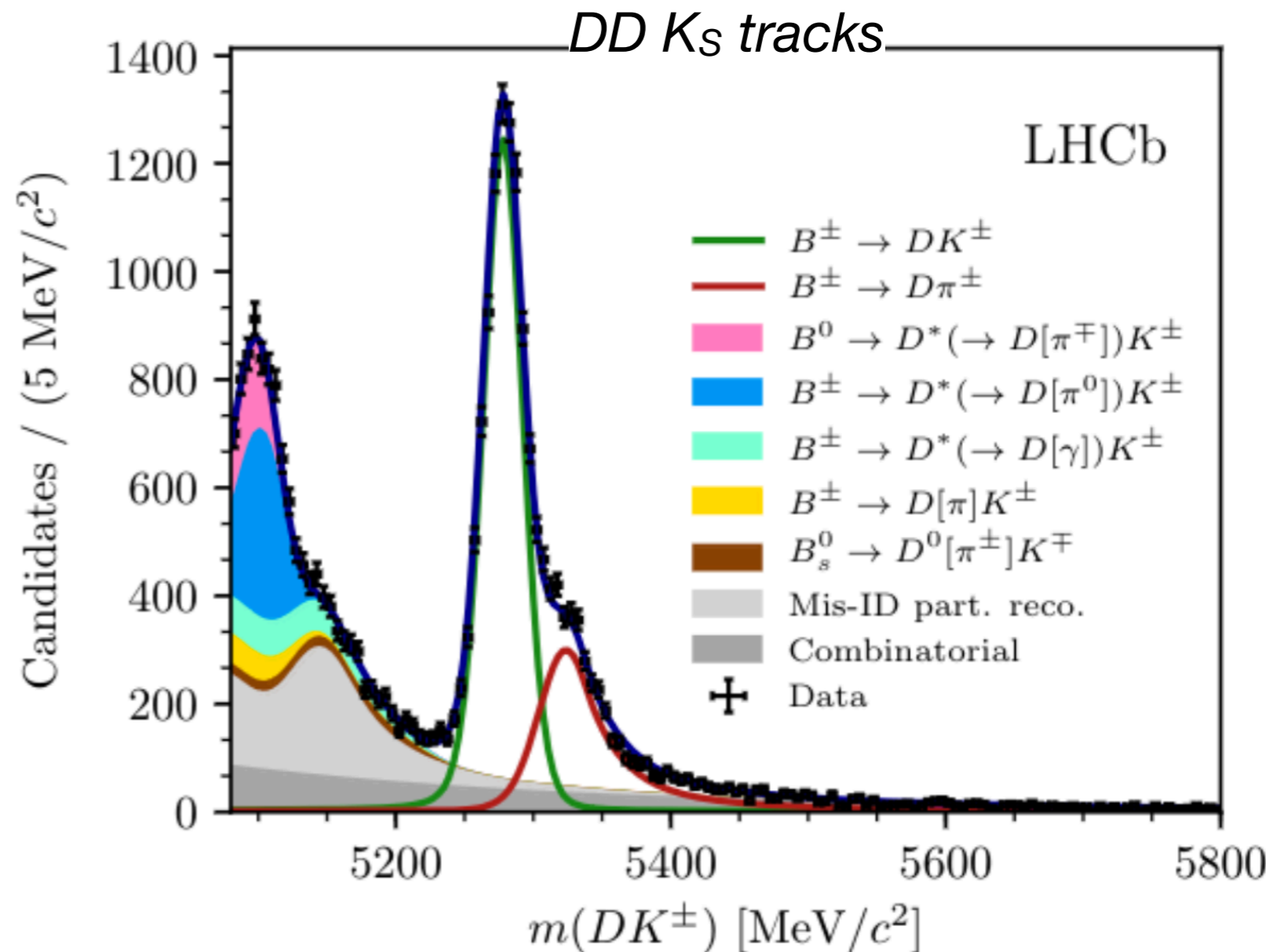
**Present status:** *prepare the work while waiting for scenarios to be defined and full simulation to be deployed; simplified approaches also useful in this phases, and to provide preliminary results*

# $\gamma$ from $B^\pm \rightarrow D(K_S^0 h^+ h^-) K^\pm$

Statistically limited at  $9 \text{ fb}^{-1}$ :  $\gamma = (68.7_{-5.1}^{+5.2})^\circ$ ; syst  $\sim 1^\circ$

*Key point: syst will remain under control if bkg is kept LOW*

- 2/3 of statistics from DD  $K_S$  tracks
- misID bkg: PID performance checks with scoping scenarios
- Comb bkg: from toys 40% error increase on CP observables with  $10 \times$  comb. bkg
- Part reco: plan to look at the effect of resolution



# $\phi_s$ from $B_s^0 \rightarrow J/\psi\phi$

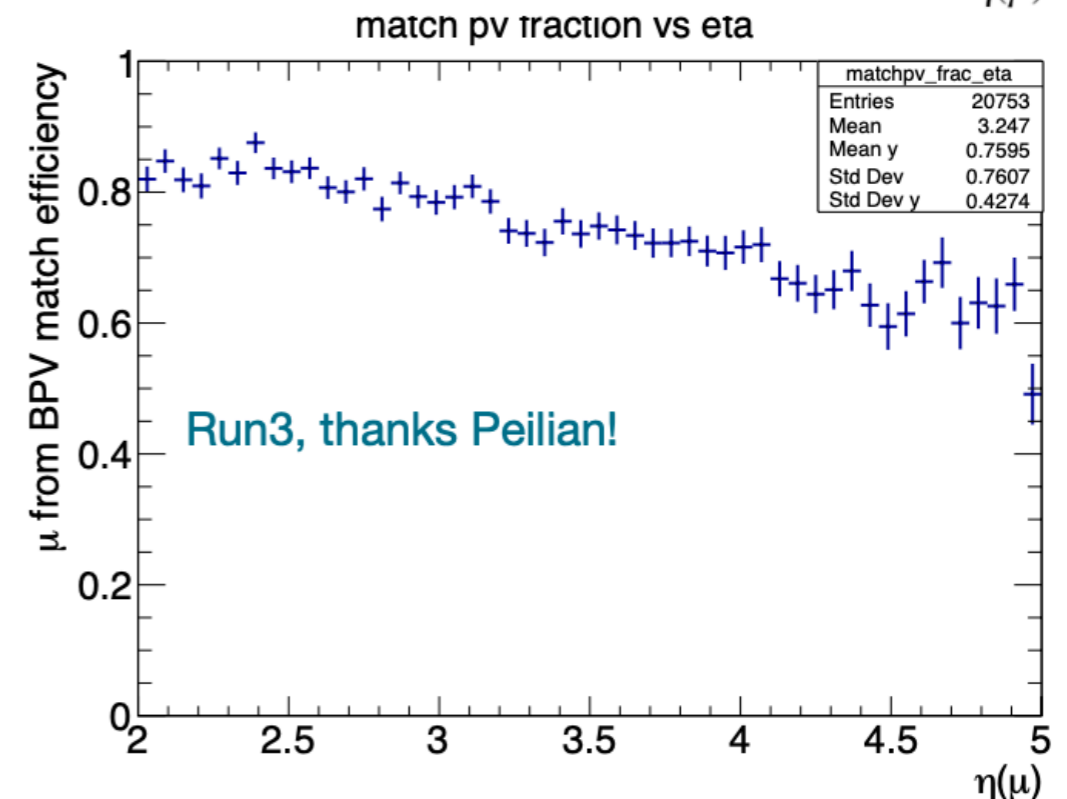
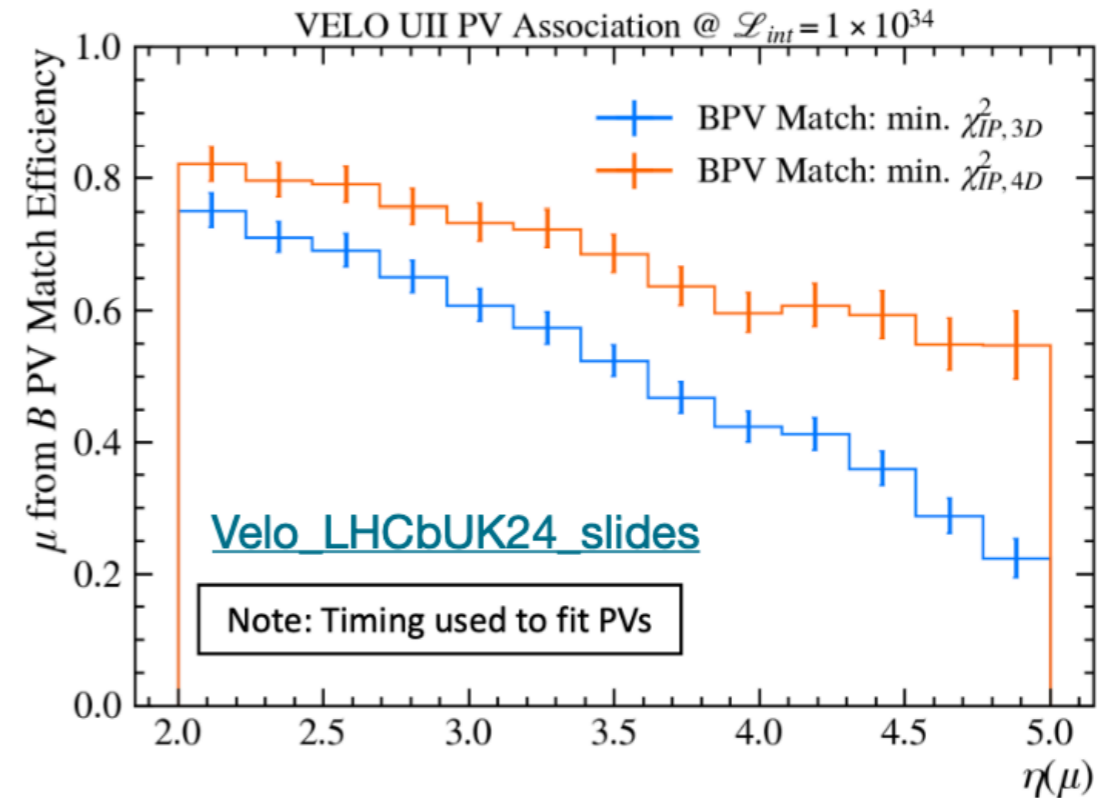
P.Li, S.Celani, Q.Fuehring,  
S.Menzemer, C.Langenbruch,  
V.Chobanova, D.M.Santos,  
C.Prouve, L.Uecker, M.Olocco

## Key point to study: flavour tagging

- 'Short' term plan: provide  $\varepsilon_{\text{tag}}, \omega, \bar{\varepsilon}$  as a function of possible U2 conditions (ghost probability, IP resolution, tracking efficiencies....)
- Starting using Run3 MC
  - ▶ But this means making some assumptions...
    - ❖ **PV matching** efficiency comparable to Run3
    - ❖ **IP resolution** of U2 at least as good as in Run3
      - If not, needed to smear the IP accordingly
    - ❖ **PID** efficiencies at least as good as Run3
      - If not, mis-identification ROC curves / efficiency maps wrt to  $p, p_T, \eta$  are needed to simulate the response

### In addition

- study dependence on event must by assigning random tracks from other events
- if FT found to be sensitive to ghost rates, take these into account



Key point to study:  $M_{corr}$  and  $q^2$  resolution

$$M_{corr} = \sqrt{p_{\perp}^2 + M_{p\mu}^2} + p_{\perp}$$

Dominated by PV and SV resolution;  
also important PV mis-association

- both effects have been studied with toy sim, ready to be updated with specific U2 scenario inputs

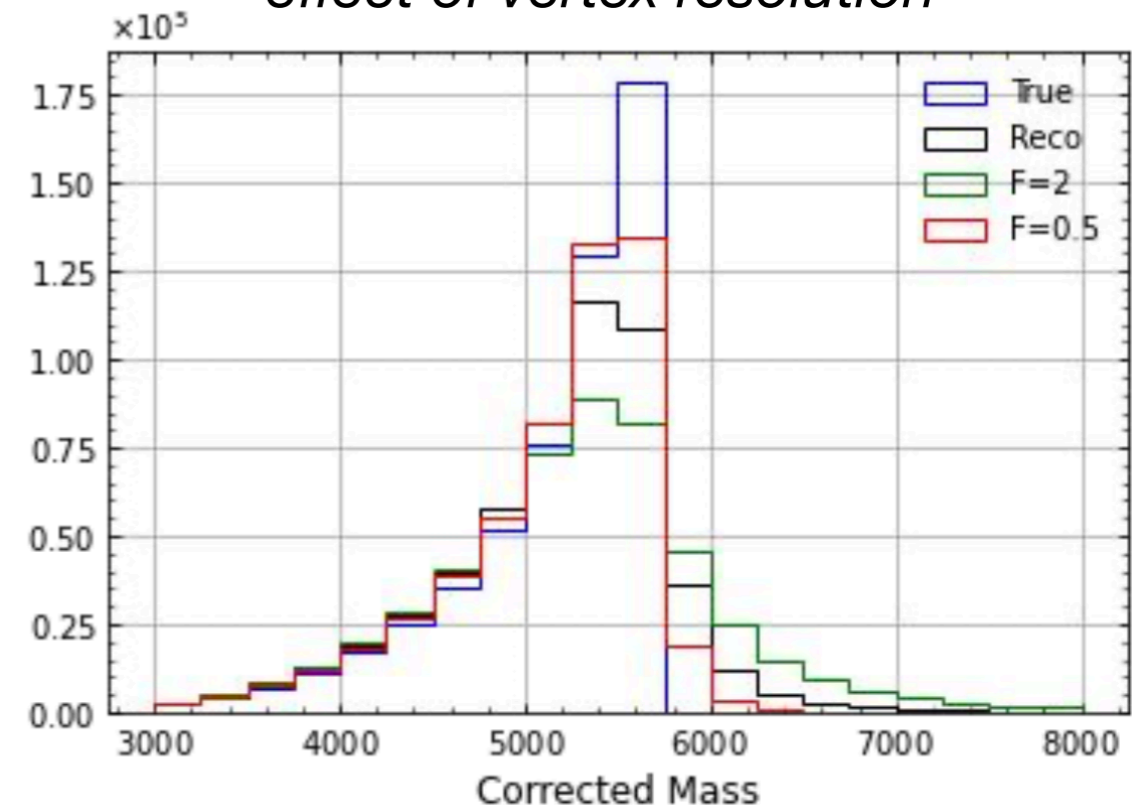
- can extend to include dominant bkg like  $\Lambda_b \rightarrow \Lambda_c\mu\nu$  and  $\Lambda_b \rightarrow N^*\mu\nu$

More studies:

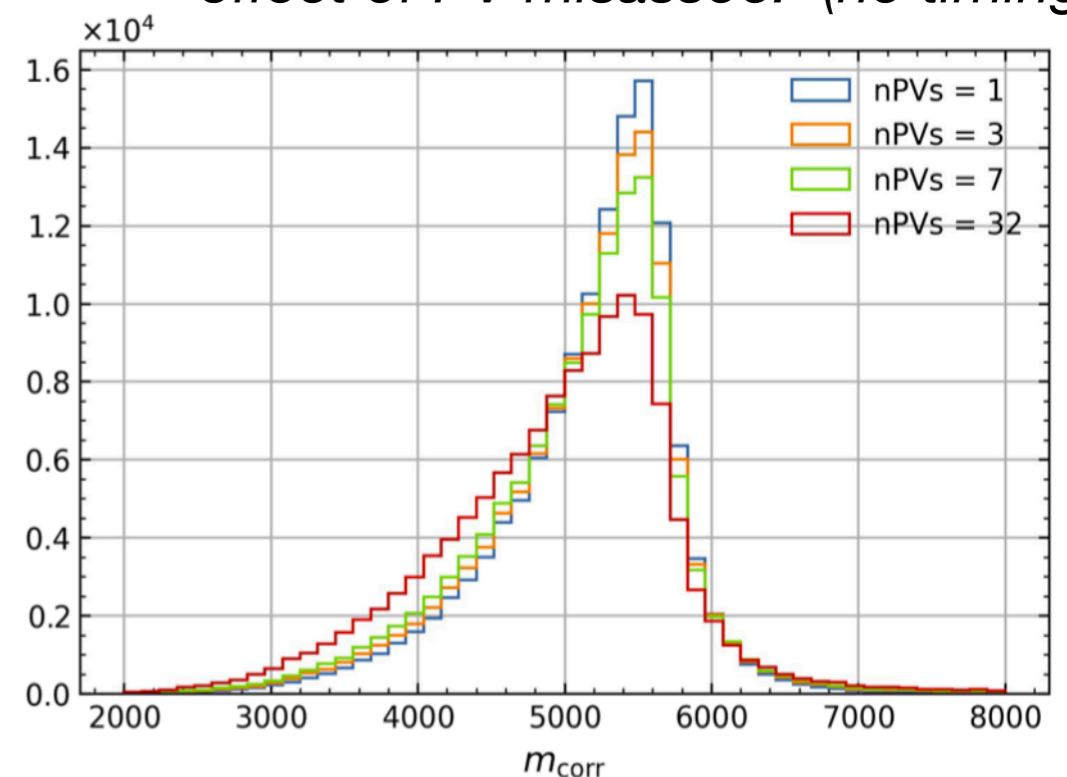
- impact of proton PID at low momenta, muon trigger efficiency

- more difficult: isolation, combinatorial  $p\mu$

effect of vertex resolution



effect of PV misassoc. (no timing)





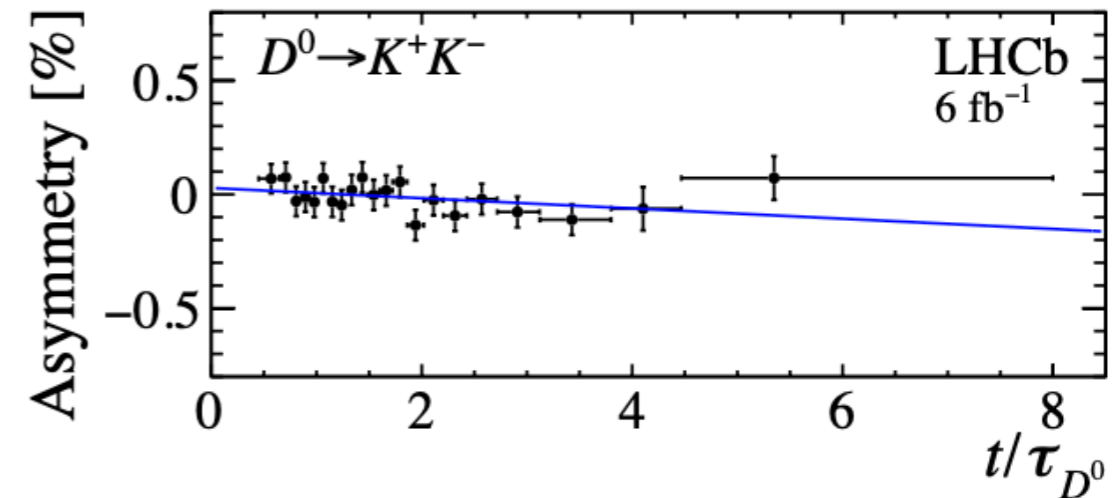
# $A_{\Gamma}$ from $D^0 \rightarrow K^+ K^-$

Slope of time-dependent asymmetry  $\rightarrow$  time reso of  $0.1 \tau(D^0)$  is enough

precision of  $10^{-4}$  already reached in Run 2

## Key points to address

- background level
- separation of primary and secondary (from-B) decays (mostly based on IP/IPCHI2)
- PV mis-ID (large time biases)
- size of detection asymmetries

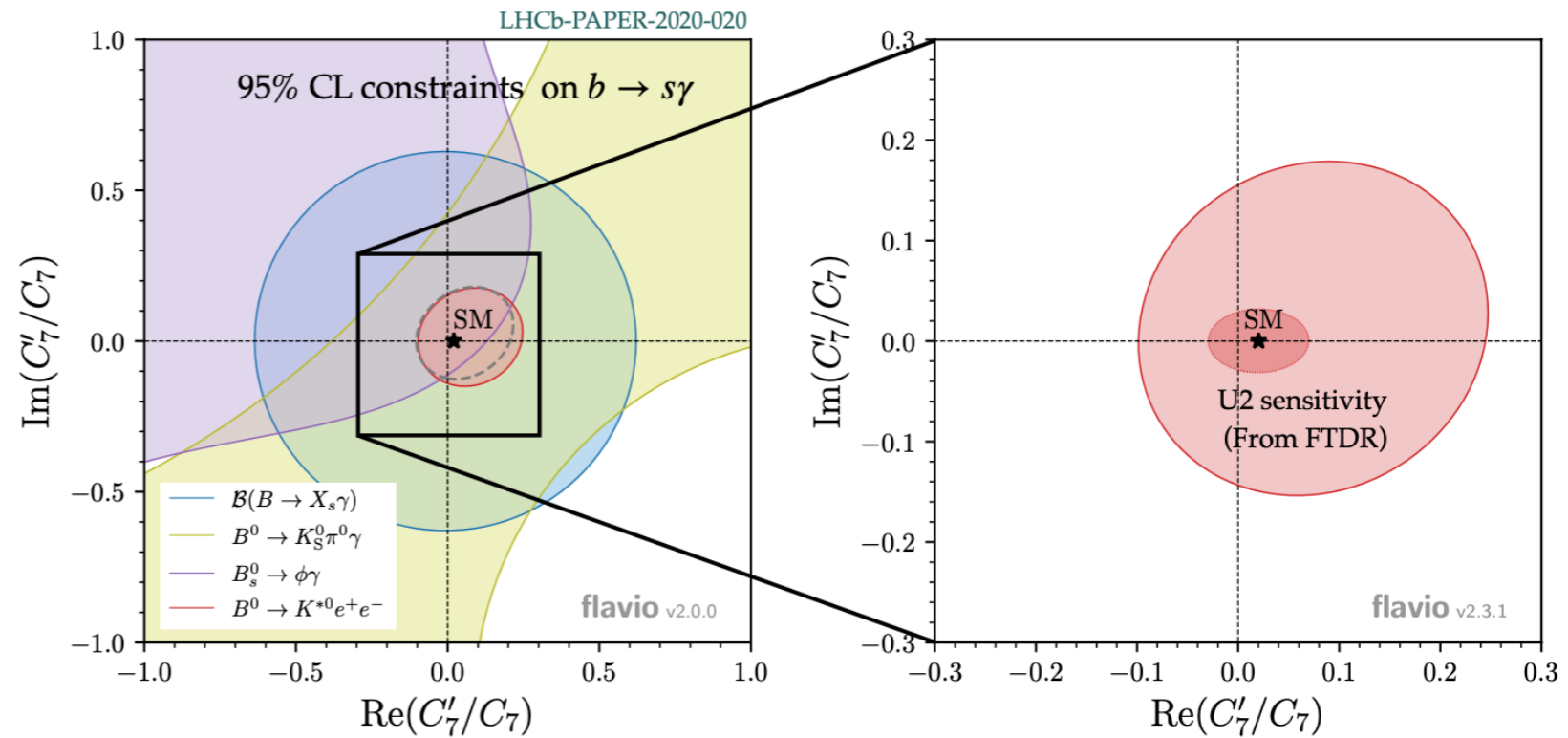


Final precision crucially depends on the trigger requirements that we can afford (S/B, bandwidth) with HLT1 playing the most important role until now

## How

- RapidSim + U2 momentum resolution smearing (Renato)
  - Mass resolution of  $D^0$  and  $D^*$
  - Inner/outer/MS geometrical acceptance
- Full simulation + Rec and Moore @ velo\_upgrade2 branch (Tim)
  - Efficiency vs. signal purity: IP/momentum thresholds, PID performance, MS
  - Prompt/secondary separation vs. IP resolution
  - Decay-time bias: PV association
  - Detection asymmetry

Strong potential for improved constraints at U2



## Key points

1. Electron **tracking** and PID (**ECAL**)
2. Low brem loss (upstream **material**)
3. Good brem recovery (**ECAL**)
4. Measurement of  $e^+e^-$  decay plane
  - Small  $e^+e^-$  angle (**VELO** resolution)
  - Multiple scattering in **VELO material**
5. Minimise bkg from  $\gamma$  conversions
  - **VELO material** budget
  - **VELO material** causes multiple scattering and worsens  $m(ee)$  resolution

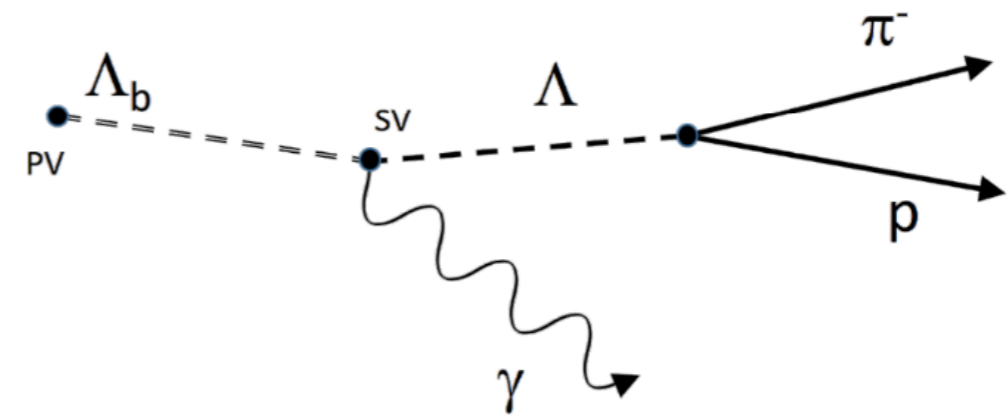
*Many of these will be discussed at this workshop*

## Key points

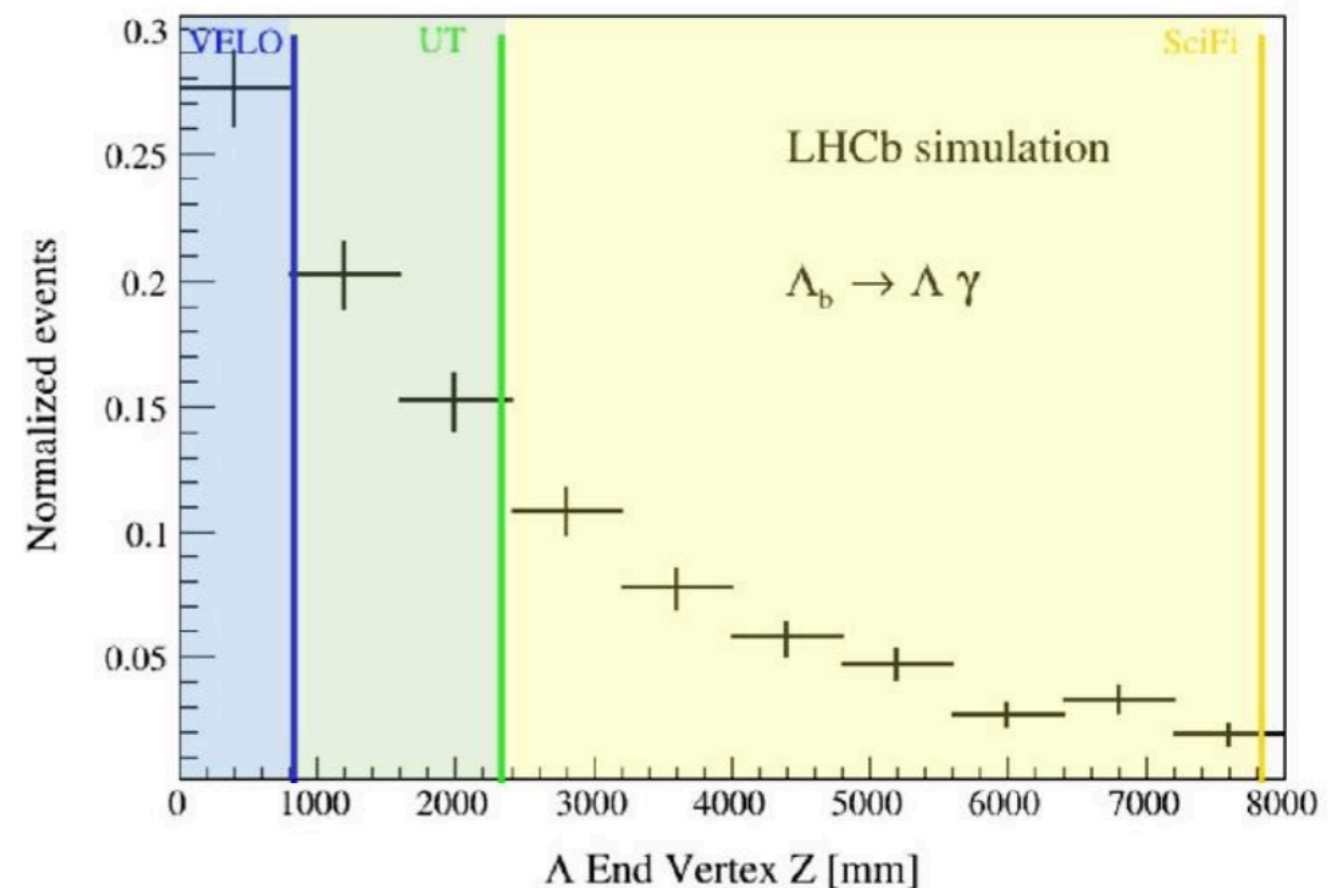
*Non trivial vertexing, photon reconstruction, acceptance*

- *efficiency limited by track reconstruction*  
 ⇒ *evaluate per category (LL/DD/TT)*

- *mass resolution limited by photon reconstruction*  
 ⇒ *evaluate impact of PicoCal downscoping on resolution and photon ID*



12% (LL), 51% (DD), and 37% (TT)



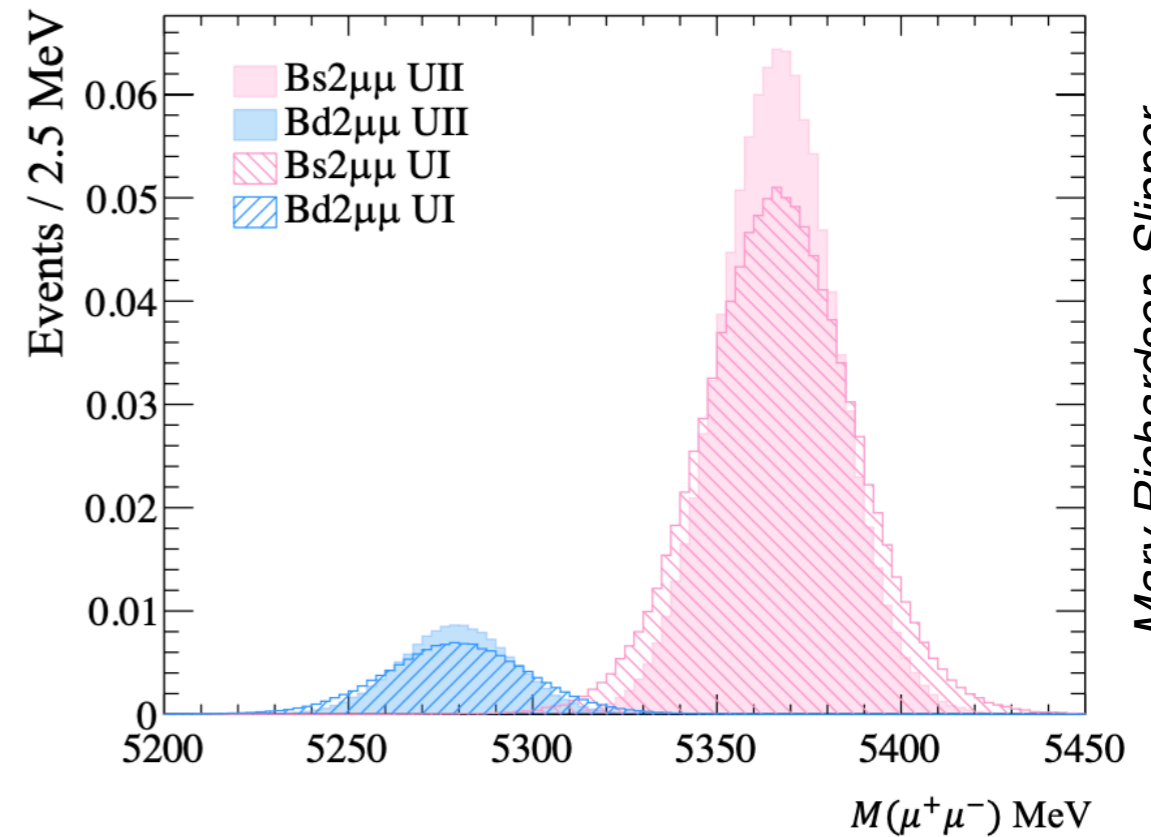
# ...and there's more

## Example 1

Flagship channels like  $B_{d,s} \rightarrow \mu^+ \mu^-$  can illustrate very well the power of an improved momentum resolution

Very good selling point for the project in ALL scenarios

$B_{d,s} \rightarrow \mu^+ \mu^-$  with  $U1/U2$  mass resolution

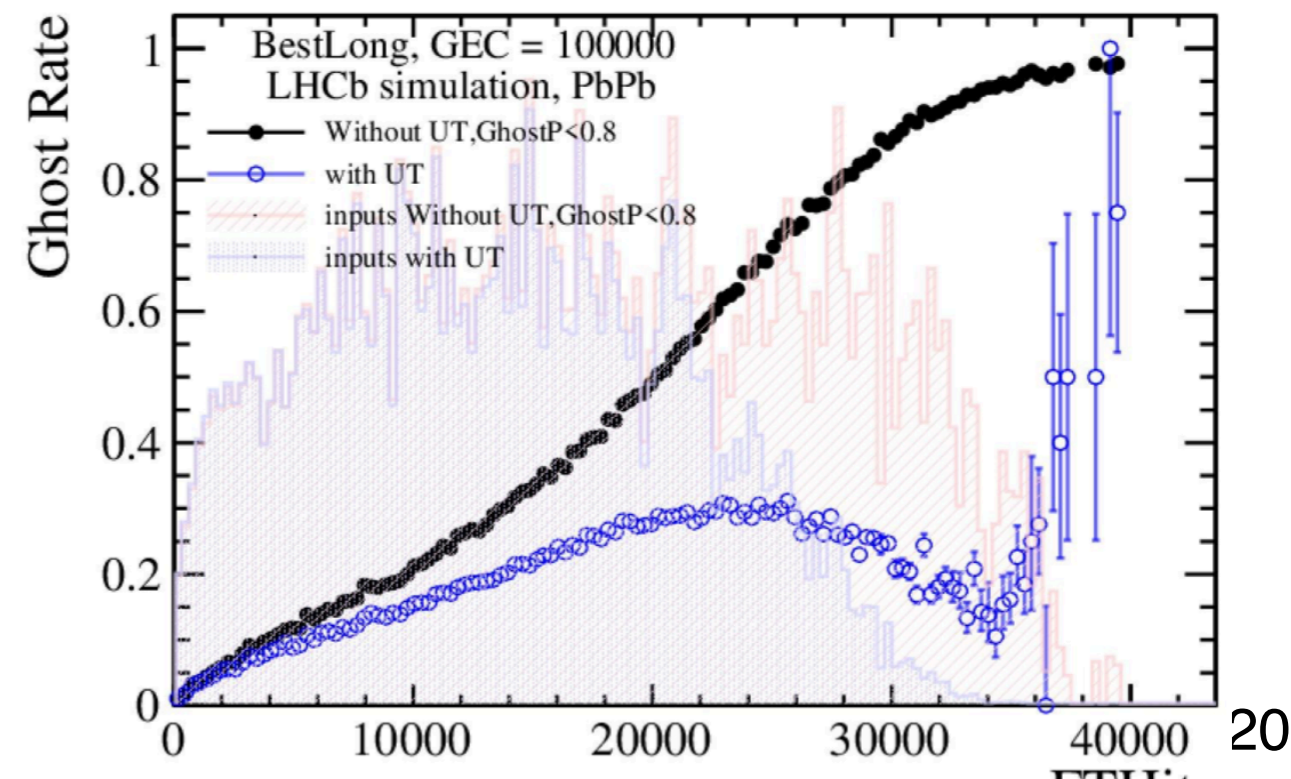


Mary Richardson-Slipper

## Example 2

Tracking with heavy ions down to very low centrality: critical metric ghost rate

Very good selling point for the HIGH scenario



# Conclusions

*We need quantitative statements on the impact of the detector scenarios on selected physics channels*

*Results are needed in time for the Scoping Document, which is supposed to be circulated within the collaboration end of June, and sent to LHCC beginning of September*

*Preliminary figures (e.g. effect of momentum resolution improvement) could be also shown at the April LHCC/RRB meetings, where we're supposed to present the U2 detector scenarios  $\Rightarrow$  this would have a positive impact on the discussion*

*Effort ongoing , needs good coordination with detector studies*

***A HUGE thanks to all people involved, and especially to the SIMULATION TEAM!!!***