



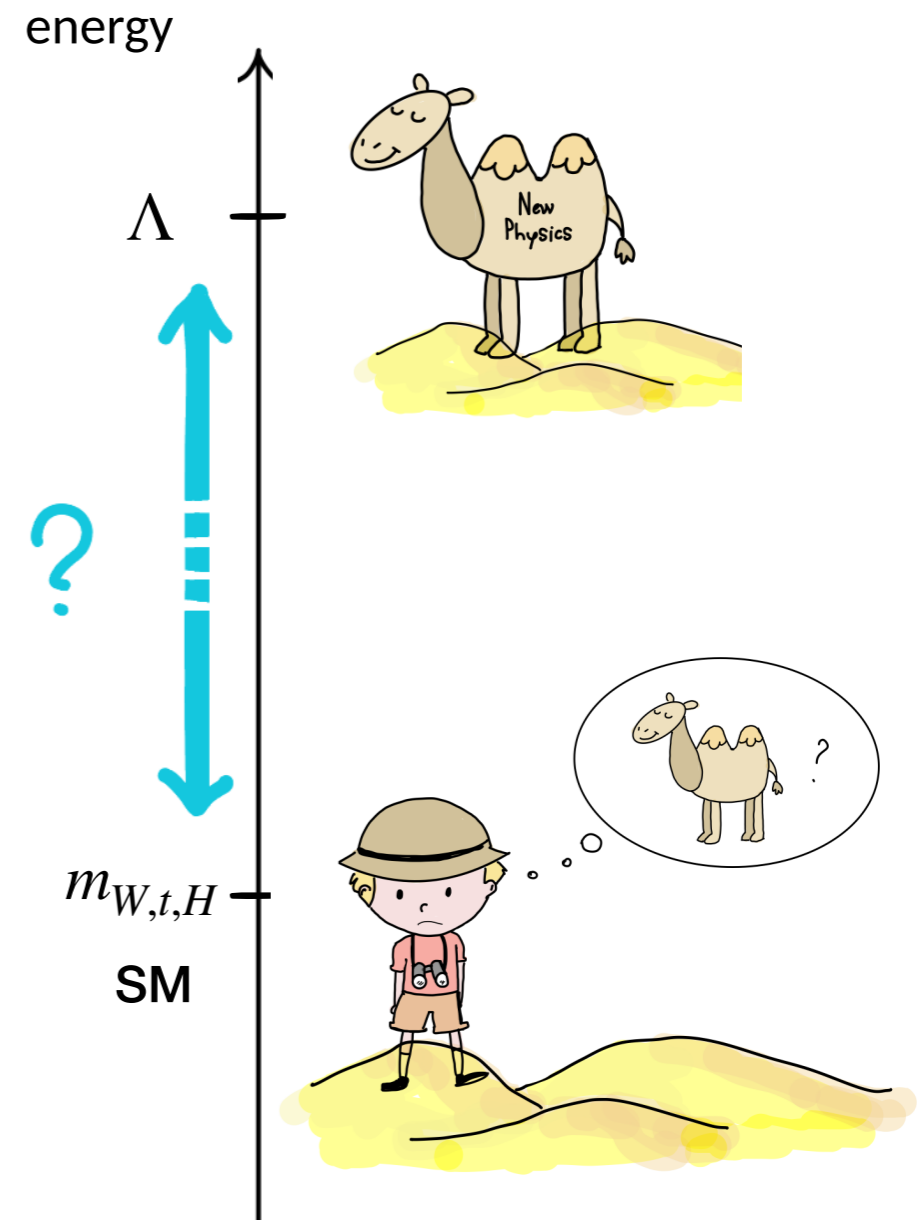
New Physics in the third generation: current status and future prospects

Claudia Cornella (JGU Mainz)

based on 2311.00020 and work in progress with L. Allwicher, G. Isidori, and B. Stefanek

The scale of New Physics

We have many reasons to think that the SM must be extended at higher energies. But **how high?**

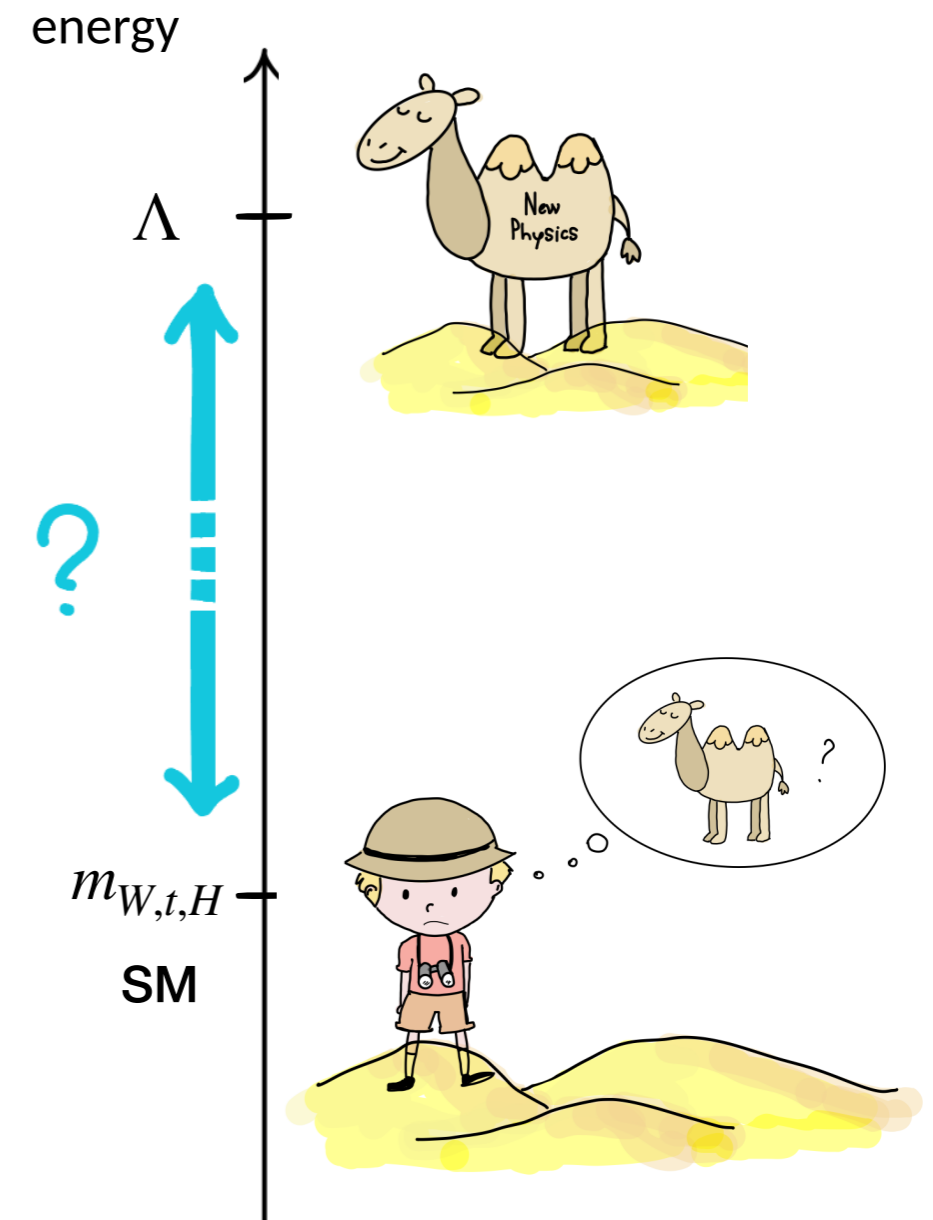


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With data we place constraints on the coefficients of SMEFT operators, and interpret them as constraints on the NP scale.



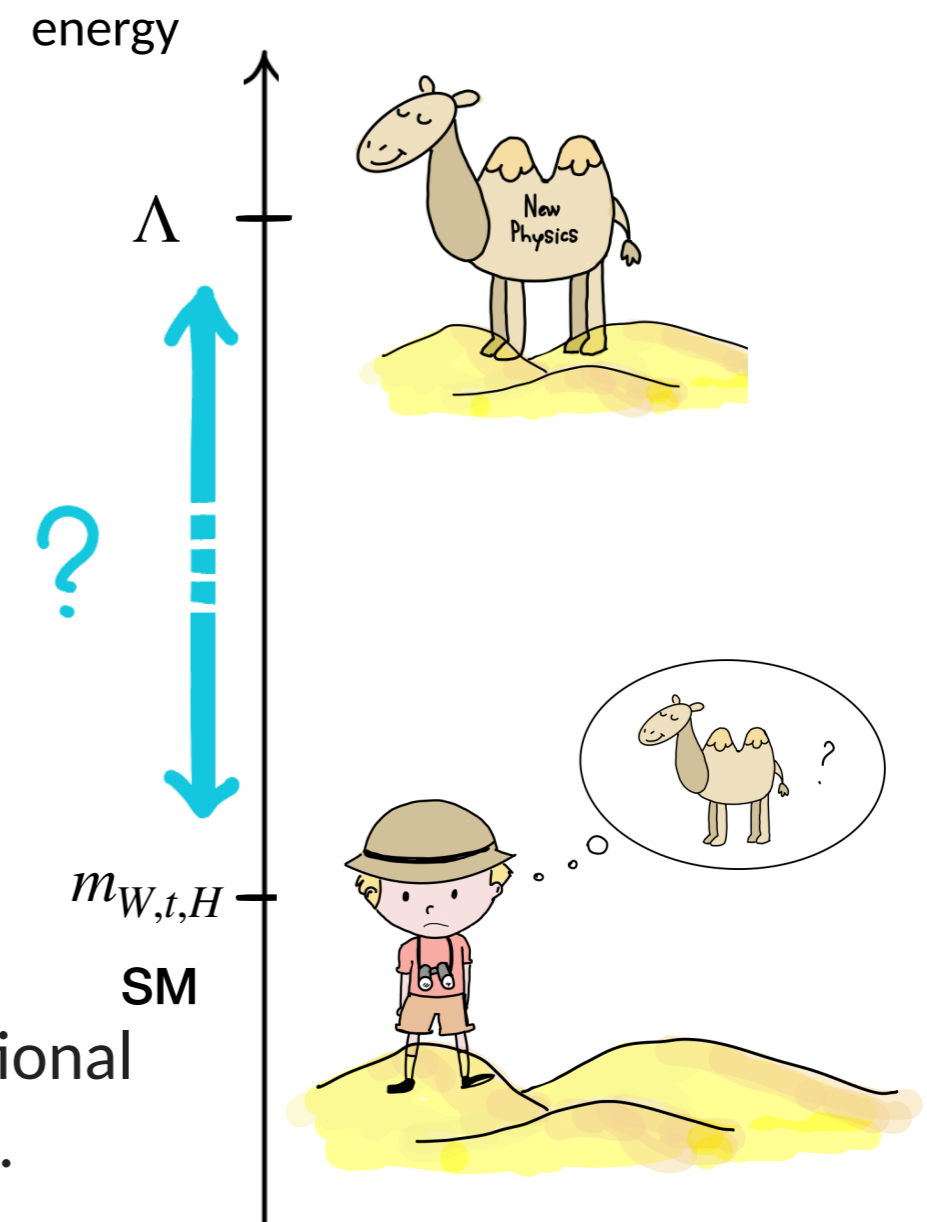
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However, interpreting these constraints without additional assumptions can lead to overly pessimistic estimates...



The scale of New Physics: an example from the past

Many thanks to ©Gino for suggesting this example!

Back in the 70s, the SM only had two generations of quarks.
CP was an accidental symmetry of the SM(2) Lagrangian.

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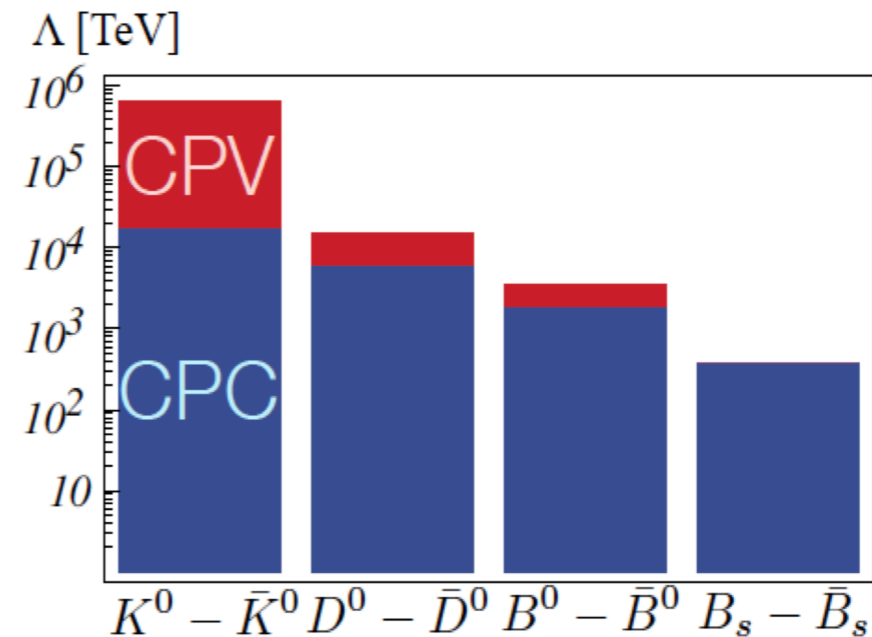
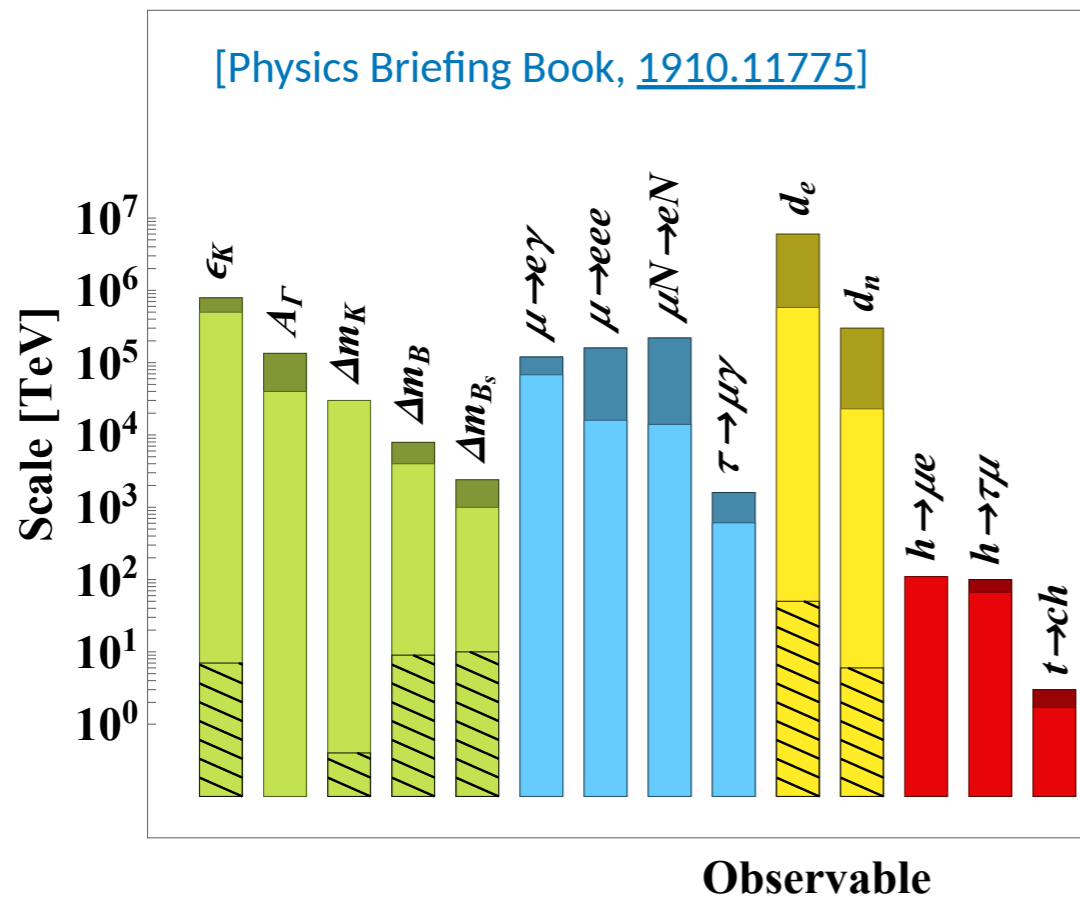
Now we know that the real characteristic scale of this interaction was much lower:

$$\frac{1}{\Lambda_{\text{CP}}^2} \sim \frac{(G_F m_t V_{ts} V_{td})^2}{4\pi^2}$$

The scale of New Physics

Similar caution is needed when interpreting SMEFT bounds.

With $O(1)$ NP couplings, bounds on flavor-violating operators point to huge scales:

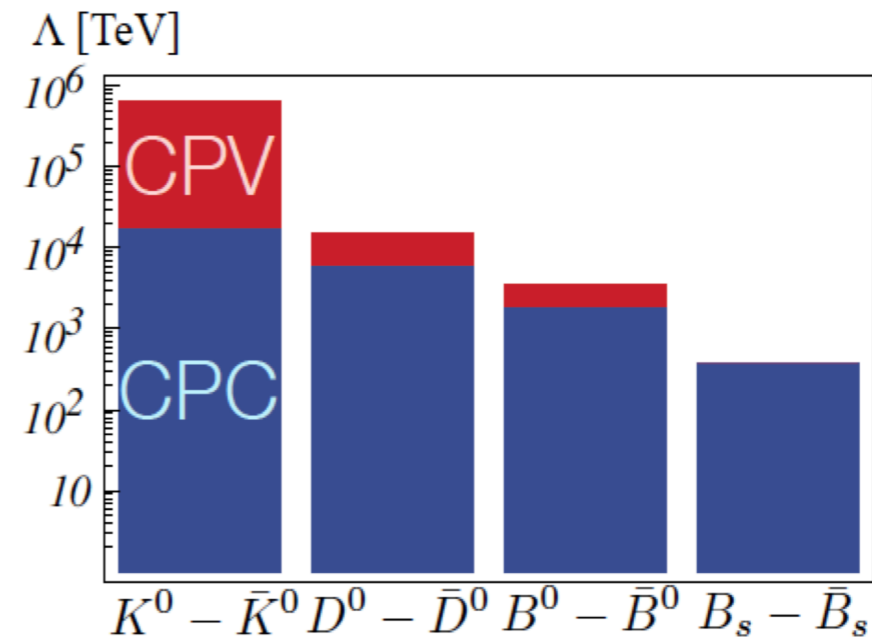
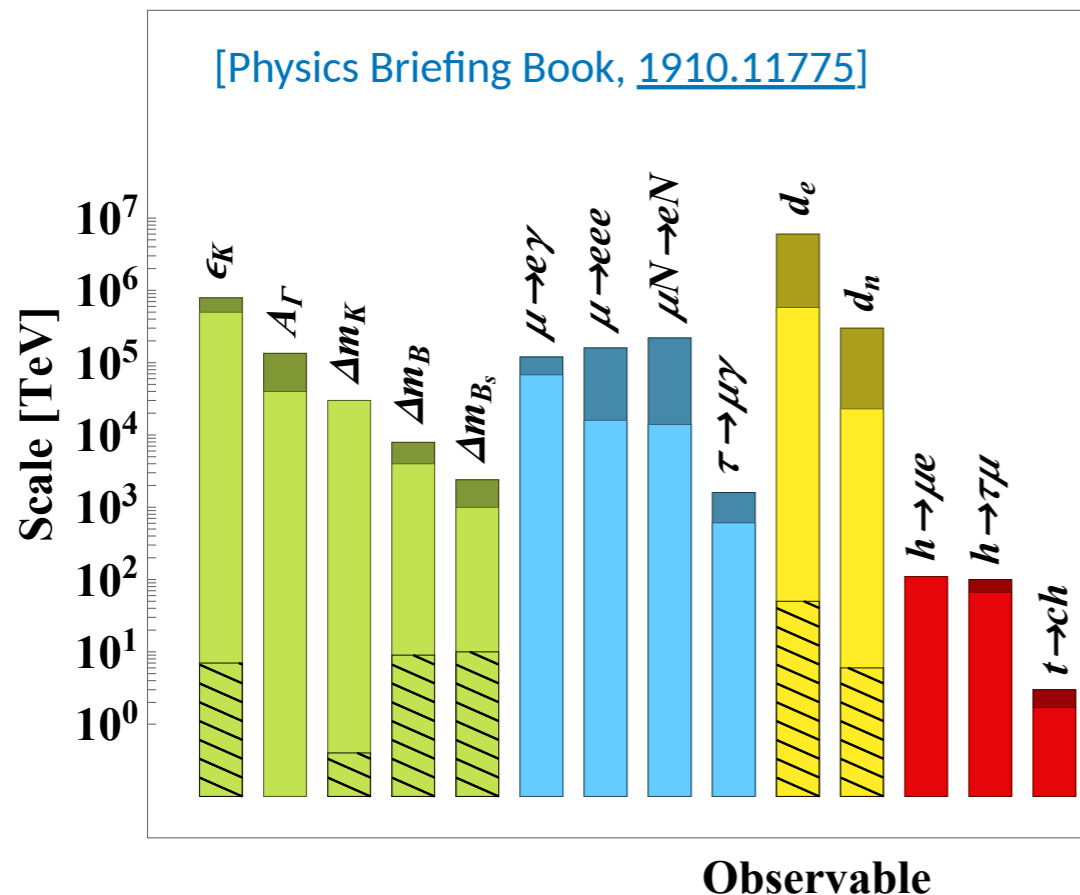


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Making **educated assumptions about the NP structure** and translating them into selection rules in the SMEFT can provide a more informative interpretation of bounds!

Goal and outline

Here: focus on **models where NP predominantly couples to the third generation.**

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- Outline
- Introduction of the flavor symmetry characterising these models, $U(2)^5$.
 - SMEFT + $U(2)^5$
 - Bounds on the $U(2)^5$ - symmetric SMEFT
 - Same, but for NP coupling mostly to the 3rd generation.
 - Future projections.

The SM flavor puzzle and $U(2)$ symmetry

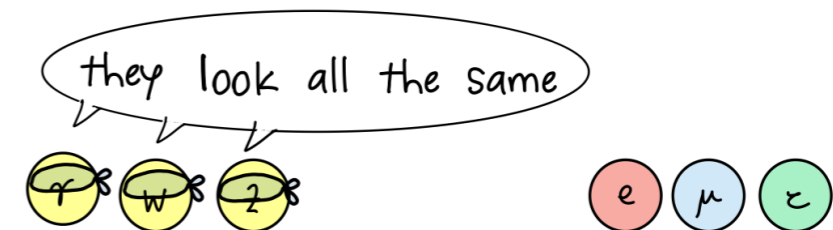
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The **gauge** sector of the SM is **flavor blind**, and has a large accidental symmetry:

$$\mathcal{G}_F = U(3)^5 \equiv U(3)_q \times U(3)_u \times U(3)_d \times U(3)_\ell \times U(3)_e$$



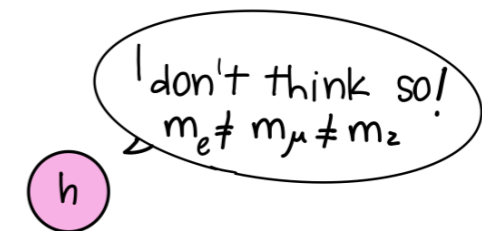
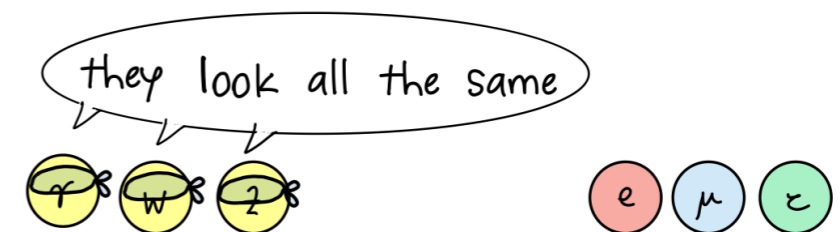
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Yukawa interactions **break** this symmetry in a specific way:



$$M_{e,d,u} = \begin{bmatrix} \text{light gray} & & \\ & \text{medium gray} & \\ & & \text{black} \end{bmatrix} \quad V_{\text{CKM}} = \begin{bmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{bmatrix} = \begin{bmatrix} \text{black} & \text{medium gray} & \text{light gray} \\ \text{medium gray} & \text{black} & \text{medium gray} \\ \text{light gray} & \text{medium gray} & \text{black} \end{bmatrix}$$

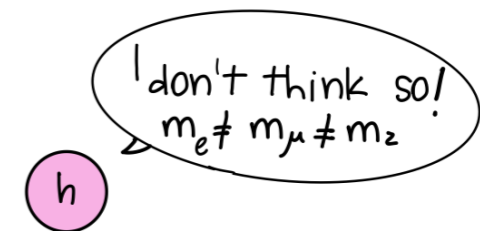
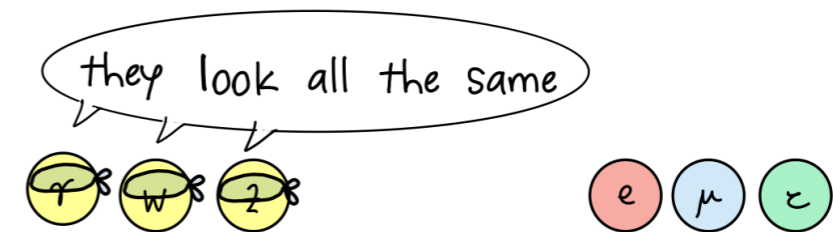
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$$U^5(3) \rightarrow U(2)^5 \equiv U(2)_q \times U(2)_u \times U(2)_d \times U(2)_\ell \times U(2)_e$$

$$\psi = (\psi_1 \psi_2) \psi_3$$

[Barbieri et al. 2022, Isidori, Straub 2012]

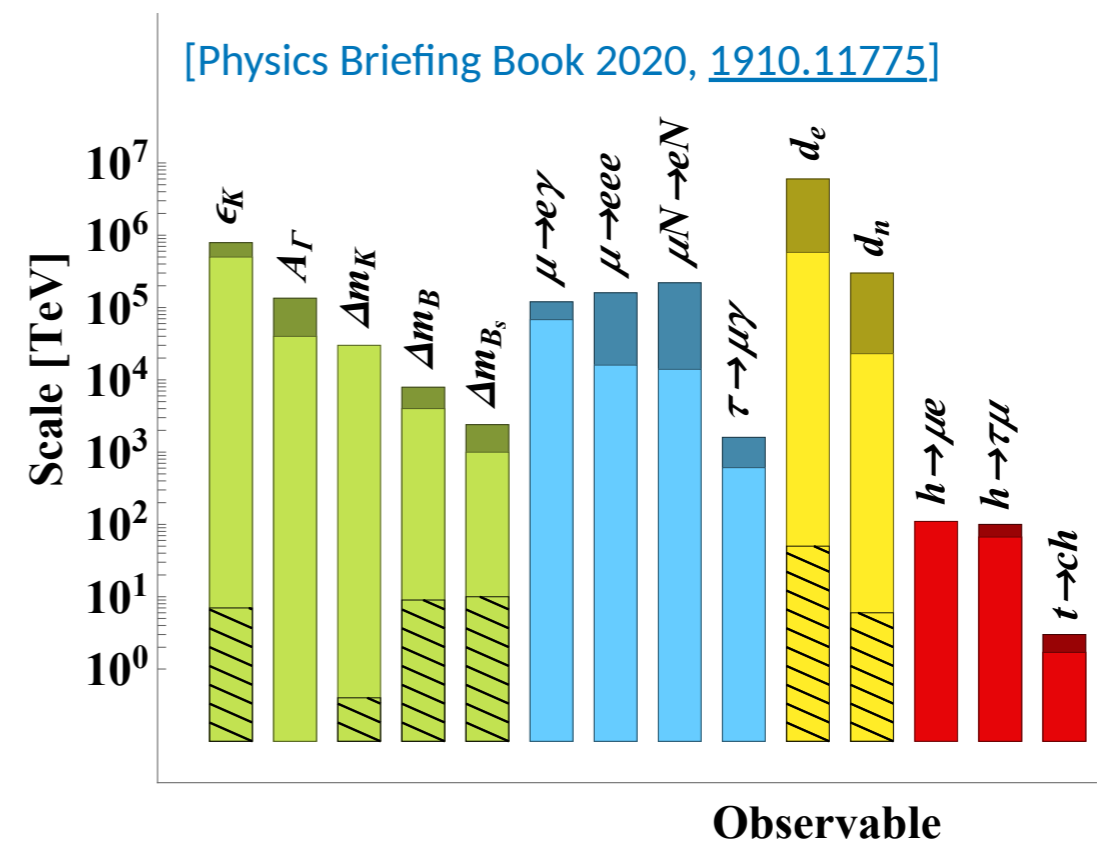
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The NP flavor puzzle:

Flavor is just an accidental symmetry: nothing forbids it to be badly violated in the UV. Then why don't we observe sizeable non-standard flavor-violating effects?

Either because the scale of these interaction is astronomically high.....

Or because the couplings of these operators are small.



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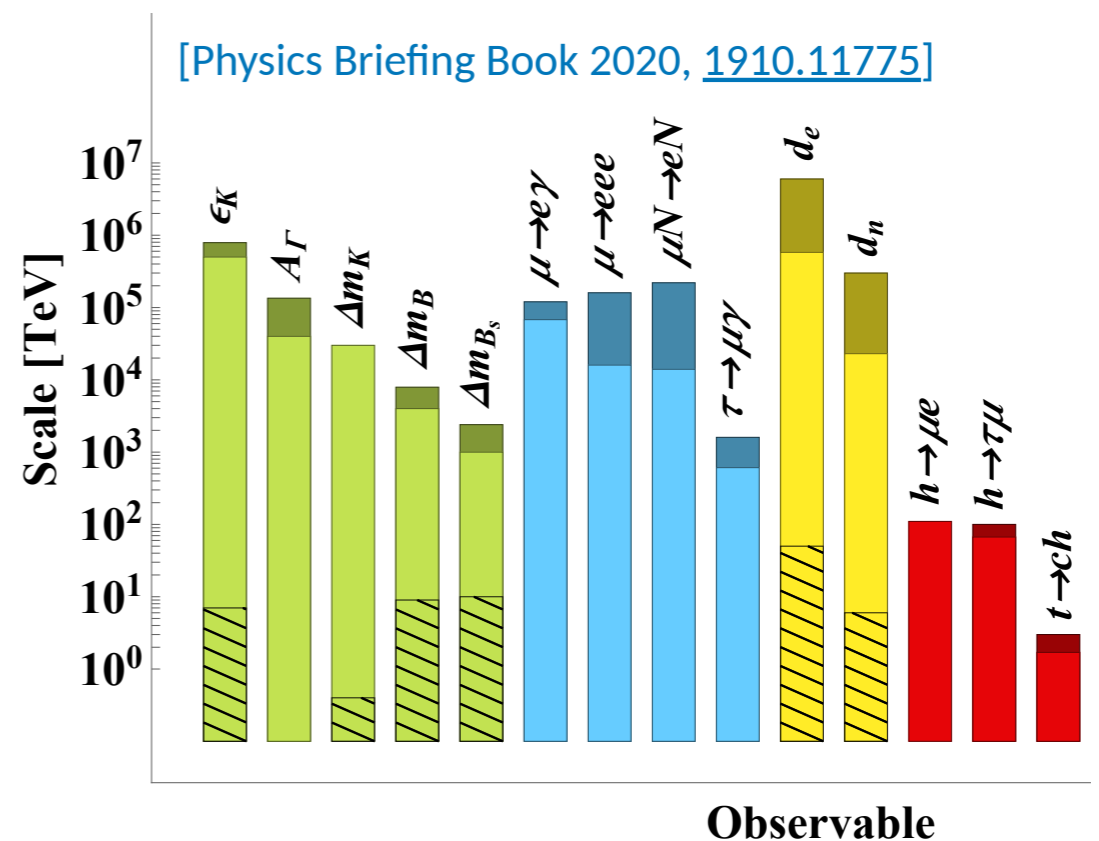
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Or because the couplings of these operators are small.

In either case, the only **unambiguous** message of these bounds is that **there is no large breaking of $U(2)^5$ at nearby scales.**

$U(2)^5$ is a good symmetry also of the SMEFT!



$U(2)^5$ vs MFV

Previously, the way to allow for TeV NP while protecting it from flavor bounds was to assume **Minimal Flavor Violation**.

- Yukawas are the only sources of $G_f=U(3)^5$ breaking also beyond the SM.
- by construction, MFV gives little to no effect in flavor-changing processes.
- MFV describes (perturbations around) **flavor-universal NP** $\textcircled{1} = \textcircled{2} = \textcircled{3}$

In particular, it does *not* suppress NP couplings to valence quarks....

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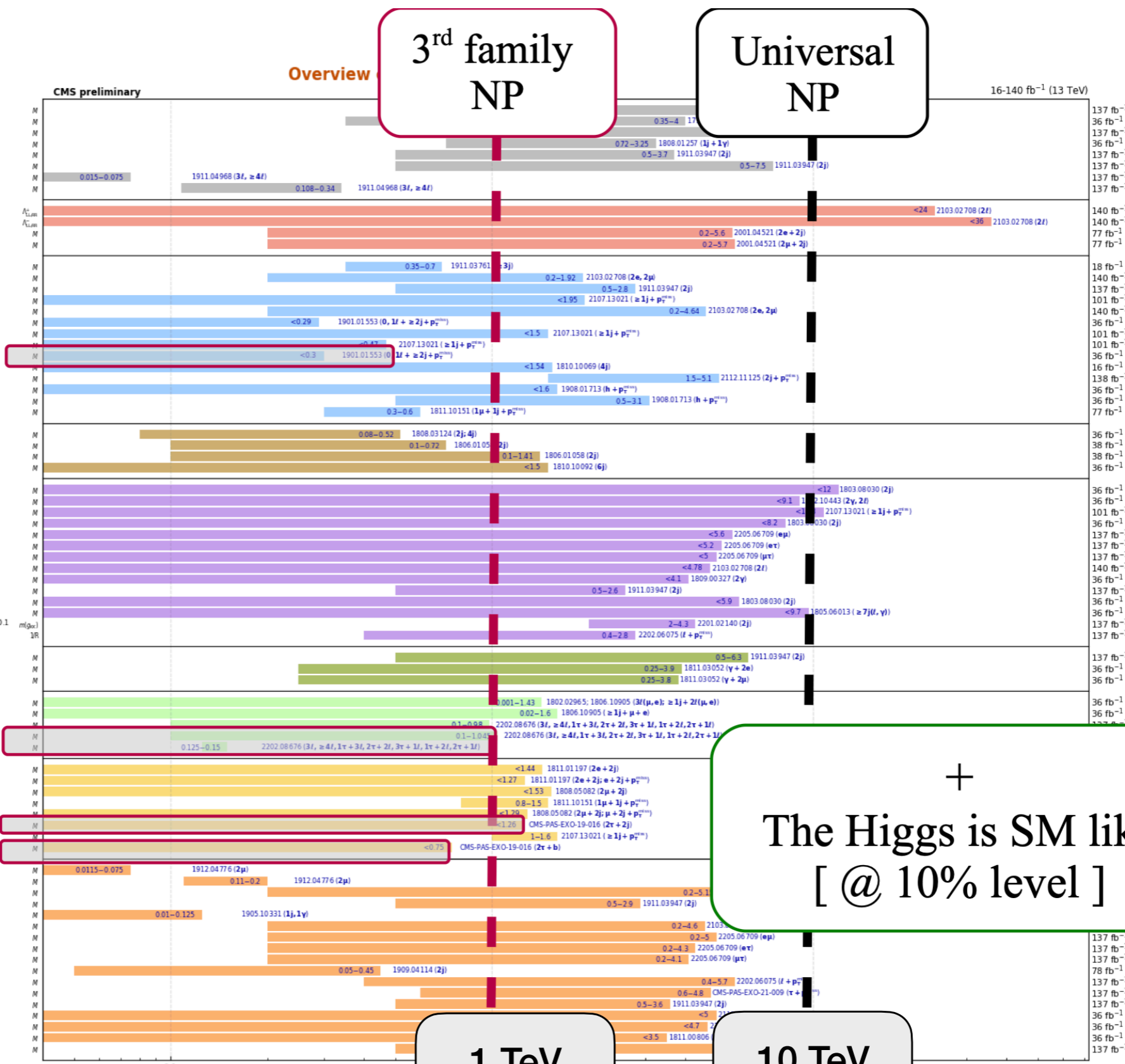
By contrast, $U(2)^5$ describes **flavor non-universal NP**, placing a clear distinction between light and heavy generations. $\textcircled{1} = \textcircled{2} \neq \textcircled{3}$

Different NP couplings for light families make it possible to suppress couplings to valence quarks and relax direct search bounds!

Status of high-energy searches

@ICHEP2022

Other	String resonance Zy resonance Wy resonance Higgs y resonance Color Octet Scalar, $k_2^2 = 1/2$ Scalar Diquark $t\bar{t} + \phi$, pseudoscalar (scalar), $g_{\text{top}}^2 \times \text{BR}(\phi \rightarrow 2t) > = 0.03(0.004)$ $t\bar{t} + \phi$, pseudoscalar (scalar), $g_{\text{top}}^2 \times \text{BR}(\phi \rightarrow 2t) > = 0.03(0.04)$
Contact Interactions	quark compositeness (ll), $\eta_{L,R} = 1$ quark compositeness (ll), $\eta_{L,R} = -1$ Excited Lepton Contact Interaction Excited Lepton Contact Interaction
Dark Matter	vector mediator (qq), $g_s = 0.25, g_{\text{DM}} = 1, m_\nu = 1$ GeV vector mediator (ll), $g_s = 0.1, g_{\text{DM}} = 1, g_t = 0.01, m_\nu > 1$ TeV (axial-)vector mediator (qq), $g_s = 0.25, g_{\text{DM}} = 1, m_\nu = 1$ GeV (axial-)vector mediator (ll), $g_s = 0.25, g_{\text{DM}} = 1, m_\nu = 1$ GeV (axial-)vector mediator (ll), $g_s = 0.1, g_{\text{DM}} = 1, g_t = 0.1, m_\nu > m_{\text{charm}}/2$ scalar mediator ($+t\bar{t}$), $g_s = 1, g_{\text{DM}} = 1, m_\nu = 1$ GeV scalar mediator (fermion portal), $\lambda_s = 1, m_\nu = 1$ GeV pseudoscalar mediator ($+t\bar{t}$), $g_s = 1, g_{\text{DM}} = 1, m_\nu = 1$ GeV pseudoscalar mediator ($+t\bar{t}$), $g_s = 1, g_{\text{DM}} = 1, m_\nu = 1$ GeV complex sc. med. (dark QCD), $m_{\text{dark}} = 5$ GeV, $c_{\text{dark}} = 25$ mm Z' mediator (dark QCD), $m_{\text{dark}} = 20$ GeV, $r_{\text{dark}} = 0.3, \alpha_{\text{dark}} = \alpha_{\text{SM}}$ Baryonic Z', $g_s = 0.25, g_{\text{DM}} = 1, m_\nu = 1$ GeV Z' - 2HDM, $g_s = 0.8, g_{\text{DM}} = 1, \tan\beta = 1, m_\nu = 100$ GeV Leptoquark mediator, $\beta = 1, B = 0.1, A_{\text{LQ}} = 0.1, 800 < M_{\text{LQ}} < 1500$ GeV
RPV	RPV stop to 4 quarks RPV squark to 4 quarks RPV gluino to 4 quarks RPV gluinos to 3 quarks
Extra Dimensions	ADD (jj) HLZ, $n_{\text{ED}} = 3$ ADD (yy, ll) HLZ, $n_{\text{ED}} = 3$ ADD G_{XX} emission, $n_{\text{ED}} = 2$ ADD QBH (jj), $n_{\text{ED}} = 6$ ADD QBH (jj), $n_{\text{ED}} = 4$ ADD QBH (jj), $n_{\text{ED}} = 4$ ADD QBH (jj), $n_{\text{ED}} = 4$ RS $G_{XX}(ll)$, $k/\bar{M}_P = 0.1$ RS $G_{XX}(yy)$, $k/\bar{M}_P = 0.1$ RS $G_{XX}(qq, gg)$, $k/\bar{M}_P = 0.1$ RS QBH (jj), $n_{\text{ED}} = 1$ non-rotating BH, $M_0 = 4$ TeV, $n_{\text{ED}} = 6$ 3-brane WED $g_{XX}(\phi + g + gg)$, $g_{\text{UV}} = 6, g_{\text{IR}} = 3, \epsilon = 0.5, m(\phi)/m(g_{XX}) = 0.1$ split-UED, $\mu \geq 2$ TeV
Excited Fermions	excited light quark (qq), $\Lambda = m_{\text{exc}}^*$ excited electron, $f_e = f = F = 1, \Lambda = m_{\text{exc}}^*$ excited muon, $f_\mu = f = F = 1, \Lambda = m_{\text{exc}}^*$
Heavy Fermions	$v\text{MSM}, V_{cb} ^2 = 1.0, V_{cb} ^2 = 1.0$ $v\text{MSM}, V_{cb} ^2/ V_{cb} ^2 + V_{cb} ^2 = 1.0$ Type-III seesaw heavy fermions, Flavor-democratic Vector like taus, Doublet Vector like taus, Singlet
Leptoquarks	scalar LQ (pair prod.), coupling to 1 st gen. fermions, $\beta = 1$ scalar LQ (pair prod.), coupling to 1 st gen. fermions, $\beta = 0.5$ scalar LQ (pair prod.), coupling to 2 nd gen. fermions, $\beta = 1$ scalar LQ (pair prod.), coupling to 2 nd gen. fermions, $\beta = 0.5$ scalar LQ (pair prod.), coupling to 3 rd gen. fermions, $\beta = 1$ scalar LQ (single prod.), coupling to 1 st gen. fermions, $\beta = 0, \lambda = 1$ scalar LQ (single prod.), coupling to 3 rd gen. fermions, $\beta = 1, \lambda = 1$
Heavy Gauge Bosons	Z ₁ , narrow resonance Z ₂ , narrow resonance SSM Z' (ll) SSM Z' (qq) Z' (qq) Superstring Z' ₁ LFV Z', BR($\mu\mu$) = 10% LFV Z', BR($\tau\tau$) = 10% LFV Z', BR($\mu\tau$) = 10% Leptophobic Z' SSM W' (lv) SSM W' ($\tau\nu$) SSM W' (qq) LRSM W ₁ (μN_c), $M_{W_1} = 0.5M_{\text{SM}}$ LRSM W ₂ ($e N_c$), $M_{W_2} = 0.5M_{\text{SM}}$ LRSM W ₃ (τN_c), $M_{W_3} = 0.5M_{\text{SM}}$ Axigluon, Coloron, $\cot\theta = 1$



3rd family NP

Universal NP

+
The Higgs is SM like
[@ 10% level]

1 TeV

10 TeV

Selection of observed exclusion limits at 95% C.L. (theory uncertainties are not included).

Flavor non-universal interactions

These considerations translate into model-building ideas!

For a long time, attempts to extend the SM implicitly assumed:

- TeV-scale **flavor-universal** NP (takes care of stabilising the Higgs)
- flavor dynamics originates at some $\Lambda \gg \text{TeV}$

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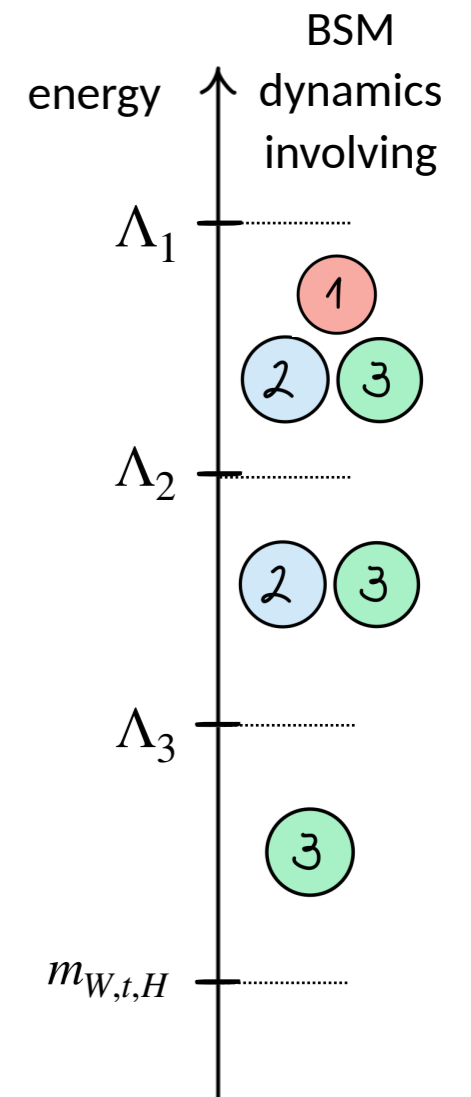
Now **flavor non-universal** interactions are gaining momentum.

[Dvali, Shifman, '00, Panico, Pomarol 1603.06609;...Bordone, CC, Fuentes, Isidori 1712.01368; Barbieri, 2103.15635; Davighi, Isidori, 2303.01520; Davighi, Stefanek, 2305.16280]

- The 3 families are *not* identical up to very high energies.

Multiscale picture: non-universal interactions acting on the i -th family switch on at $\Lambda_1 \gg \Lambda_2 \gg \Lambda_3 \gg m_W$

- interactions distinguishing light vs 3rd family emerge first @ Λ_3



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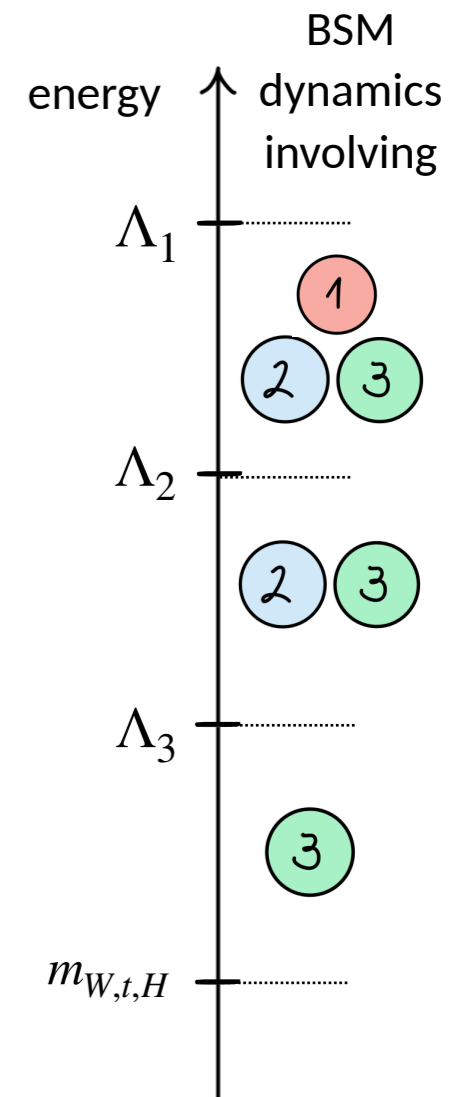
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acts on 3rd fam. & Higgs
acts on light fam.

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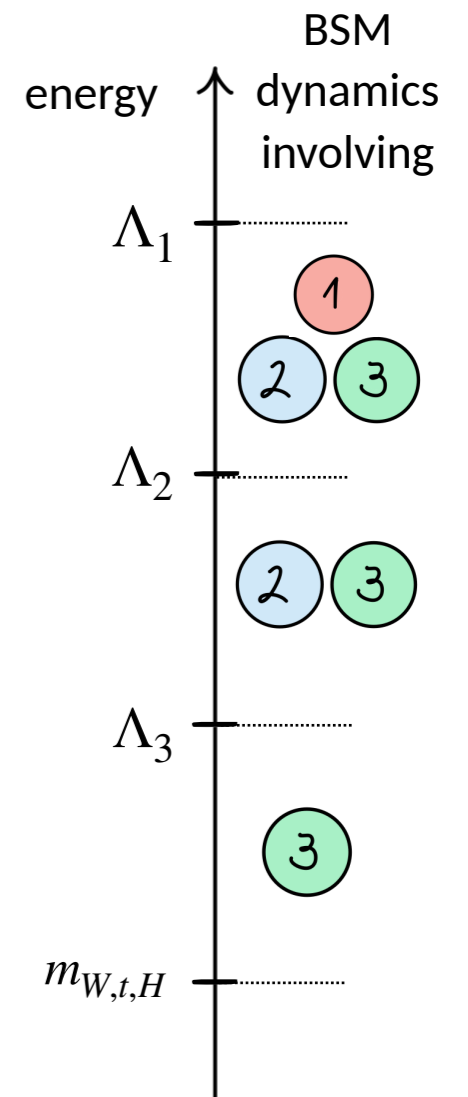
How low?

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The $U(2)$ symmetric SMEFT

$U(2)^5$ is an **efficient organising principle**:

- SMEFT with 3 generations has $1350 + 1149 = 2499$ independent WCs at dim-6.
- In the exact $U(2)^5$ limit, this is reduced to $124 + 23 = 147$ independent WCs.

Here we focus on the CP-conserving case.

Operators	$U(2)^5$ [terms summed up to different orders]													
	Exact		$\mathcal{O}(V^1)$		$\mathcal{O}(V^2)$		$\mathcal{O}(V^1, \Delta^1)$		$\mathcal{O}(V^2, \Delta^1)$		$\mathcal{O}(V^2, \Delta^1 V^1)$		$\mathcal{O}(V^3, \Delta^1 V^1)$	
Class 1–4	9	6	9	6	9	6	9	6	9	6	9	6	9	6
$\psi^2 H^3$	3	3	6	6	6	6	9	9	9	9	12	12	12	12
$\psi^2 XH$	8	8	16	16	16	16	24	24	24	24	32	32	32	32
$\psi^2 H^2 D$	15	1	19	5	23	5	19	5	23	5	28	10	28	10
$(\bar{L}L)(\bar{L}L)$	23	–	40	17	67	24	40	17	67	24	67	24	74	31
$(\bar{R}R)(\bar{R}R)$	29	–	29	–	29	–	29	–	29	–	53	24	53	24
$(\bar{L}L)(\bar{R}R)$	32	–	48	16	64	16	53	21	69	21	90	42	90	42
$(\bar{L}R)(\bar{R}L)$	1	1	3	3	4	4	5	5	6	6	10	10	10	10
$(\bar{L}R)(\bar{L}R)$	4	4	12	12	16	16	24	24	28	28	48	48	48	48
total:	124	23	182	81	234	93	212	111	264	123	349	208	356	215

Table 6: Number of independent operators in the SMEFT assuming a minimally broken $U(2)^5$ symmetry, including breaking terms up to $\mathcal{O}(V^3, \Delta^1 V^1)$. Notations as in Table 1.

[D. A. Faroughy, G. Isidori, F. Wilsch, K. Yamamoto, [arXiv:2005.05366](https://arxiv.org/abs/2005.05366)]

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An example:
$$Q_{He}^{[ij]} = (H^\dagger iD_\mu H)(\bar{e}_i \gamma^\mu e_j)$$

SMEFT

6 independent structures

U(2)⁵ - symmetric SMEFT

only 2 independent structures

$$Q_{He}^{[33]} = (H^\dagger iD_\mu H)(\bar{e}_3 \gamma^\mu e_3),$$
$$Q_{He}^{[ii]} = (H^\dagger iD_\mu H) \sum_{i=1,2} (\bar{e}_i \gamma^\mu e_i)$$

The flavor rotation

What is the third generation in the SMEFT?

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What is the **third generation** in the SMEFT?

Non-trivial to define for the LH quark doublet because of the CKM misalignment!

In the interaction basis where the dim-6 SMEFT operators are $U(2)^5$ symmetric, the **3rd generation quark doublet** is somewhere **in-between** the **down-aligned** and the **up-aligned** case.

$$\begin{pmatrix} t_L \\ V_{td}d_L + V_{ts}s_L + V_{tb}b_L \end{pmatrix} = \begin{matrix} q_t \\ \text{up-alignment } \varepsilon_F = 1 \\ \text{down-alignment } \varepsilon_F = 0 \end{matrix} \begin{matrix} q_3 \\ q_b \end{matrix} = \begin{pmatrix} V_{ub}^* u_L + V_{cb}^* c_L + V_{tb}^* t_L \\ b_L \end{pmatrix}$$

The flavor rotation

What is the **third generation** in the SMEFT?

Non-trivial to define for the LH quark doublet because of the CKM misalignment!

In the interaction basis where the dim-6 SMEFT operators are $U(2)^5$ symmetric, the **3rd generation quark doublet** is somewhere **in-between** the **down-aligned** and the **up-aligned** case.

$$\begin{pmatrix} t_L \\ V_{td}d_L + V_{ts}s_L + V_{tb}b_L \end{pmatrix} = \begin{matrix} q_t \\ q_b \end{matrix} = \begin{pmatrix} V_{ub}^* u_L + V_{cb}^* c_L + V_{tb}^* t_L \\ b_L \end{pmatrix}$$

In the spirit of minimally-broken $U(2)^5$, we describe this **misalignment** in terms of a single **angle** in the 2-3 sector, $\theta \sim V_{cb} \epsilon_F$.

Observables

EWPO

- W-pole observables [V. Bresó-Pla, A. Falkowski, M. González-Alonso, [2103.12074](#)]
- Z-pole observables [L. Allwicher, G. Isidori, J. M. Lizana, N. Selimovic, B. Stefanek, [2302.11584](#)]
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Flavor

- $\Delta F = 1$ ($B \rightarrow X_s \gamma$, $B \rightarrow K \nu \bar{\nu}$, $K \rightarrow \pi \nu \bar{\nu}$, $B \rightarrow K^{(*)} \mu^+ \mu^-$, $B_{s,d} \rightarrow \mu^+ \mu^-$)
- $\Delta F = 2$ ($B_{s,d}$ - mixing, K - mixing, D - mixing)
- Charged-current $b \rightarrow c, u$ transitions ($R_D, R_{D^*}, B_{u,c} \rightarrow \tau \nu$)

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Collider

- LHC Drell-Yan $pp \rightarrow \ell \ell$ and mono-lepton $pp \rightarrow \ell \nu$ [L. Allwicher, D. A. Faroughy, F. Jaffredo, O. Sumensari, F. Wilsch, [2207.10756](#)]
- LHC 4-quark observables
- LEP 4-lepton $ee \rightarrow \ell \ell$ [Ethier, Magni, Maltoni, Mantani, Nocera, Rojo, Slade, Vryonidou, Zhang, [2105.00006](#)]



Analysis strategy

- Run all WCs to a reference scale $\Lambda = 3 \text{ TeV}$.
- For LEFT running, LEFT-SMEFT matching and SMEFT running we use DSixTools, which allows us to work analytically in the WCs also beyond leading log.
- Once all observables have been expressed in terms of SMEFT WCs at the high scale, we impose the $U(2)^5$ symmetry.
- We construct the combined likelihood from collider, EW, and flavour observables as a function of the 124 WCs of the $U(2)^5$ -symmetric (and CP conserving) SMEFT, and switch them on one at a time to get lower bound on the NP scale.

Results

Strong **complementarity** between 3 sectors.

Out of 124 bounds, 46 are dominated by **EWPO**, 42 by **collider**, 36 by **flavor**

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For **flavor-conserving** operators,

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	Λ_{EW}	$\Lambda_{\text{coll.}}$
$\mathcal{C}_{H\ell}^{(1)[33]}$	0.1	0.1	4.4	1.6
$\mathcal{C}_{H\ell}^{(1)[ii]}$	0.7	0.7	7.6	3.
$\mathcal{C}_{H\ell}^{(3)[33]}$	0.7	0.7	4.5	1.7
$\mathcal{C}_{H\ell}^{(3)[ii]}$	0.7	0.7	7.7	3.8
$\mathcal{C}_{He}^{[33]}$	-	-	3.8	1.5
$\mathcal{C}_{He}^{[ii]}$	0.9	0.9	6.6	2.7
$\mathcal{C}_{\ell q}^{(3)[3333]}$	0.7	1.5	1.4	1.
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$\mathcal{C}_{\ell q}^{(3)[iijj]}$	0.5	5.1	2.1	22.5

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Operators with 3rd-family fermions get milder bounds, ~ 1 TeV.

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- **Down alignment** can relax these bounds down to ~ few TeV.

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- **Importance of RG effects in the EW sector**

Without running, only 16 operators enter the EW fit.

With running, 123 out of 124 operators enter the EW fit.

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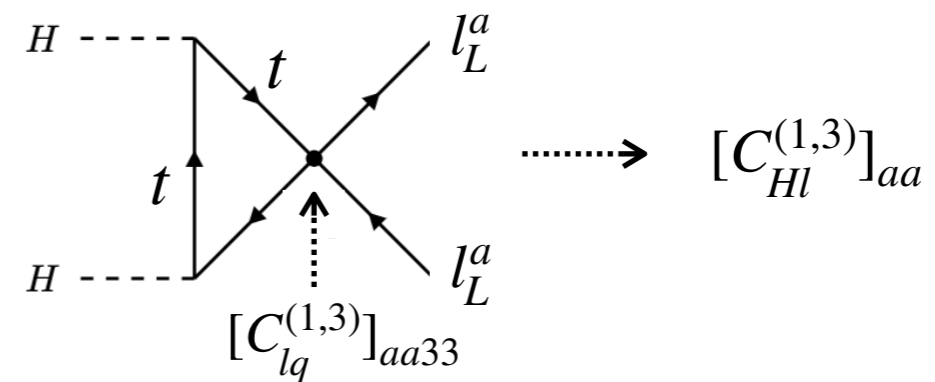
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44 get bounds stronger than 1 TeV!

these are operators w/ 3rd-family quarks running with y_t into operators directly constrained by Z-pole obs.



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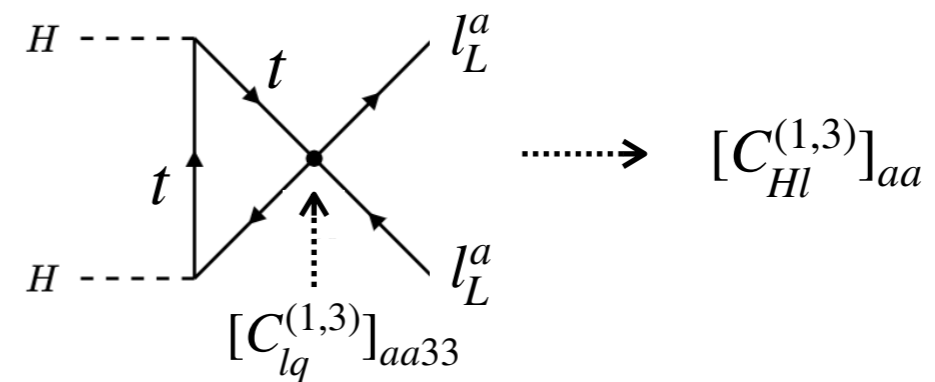
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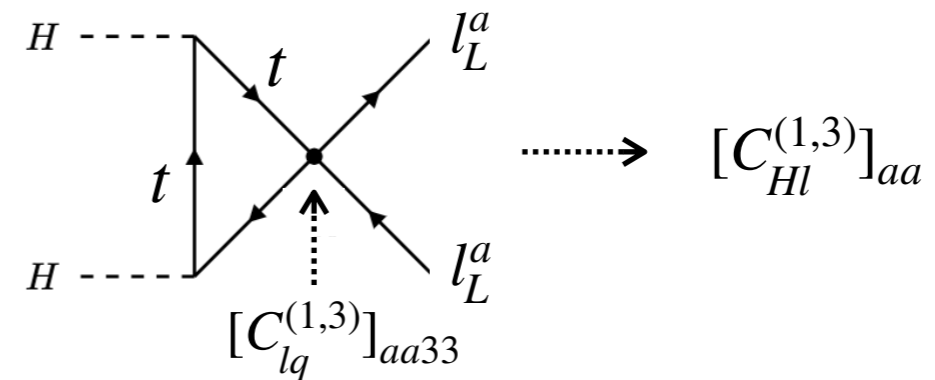
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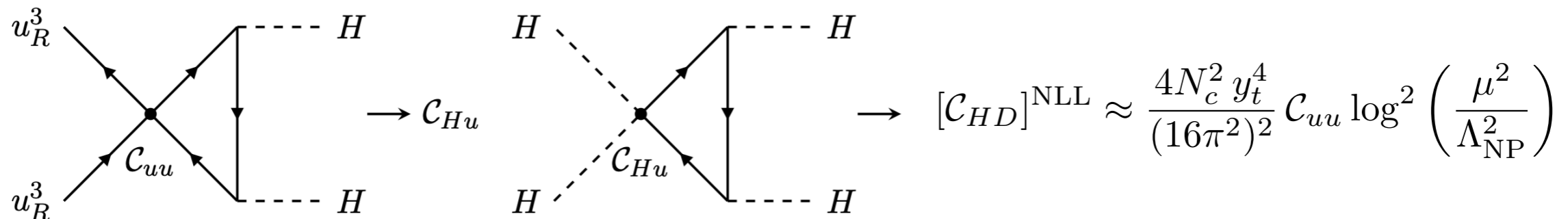
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Example: $[O_{uu}]_{3333}$ enters the EW fit only at NLL by mixing with O_{HD}



The hypothesis of NP in the 3rd generation

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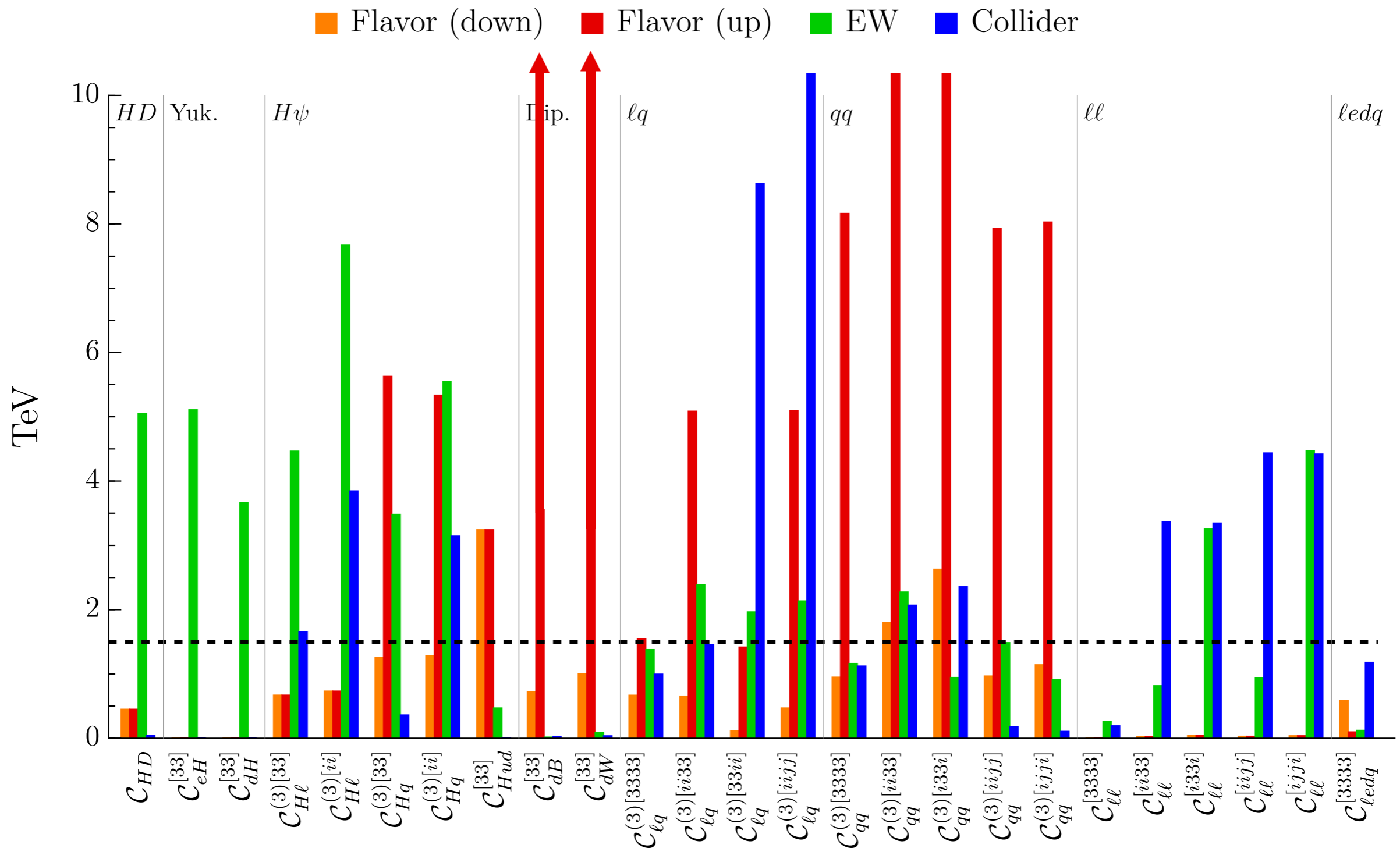
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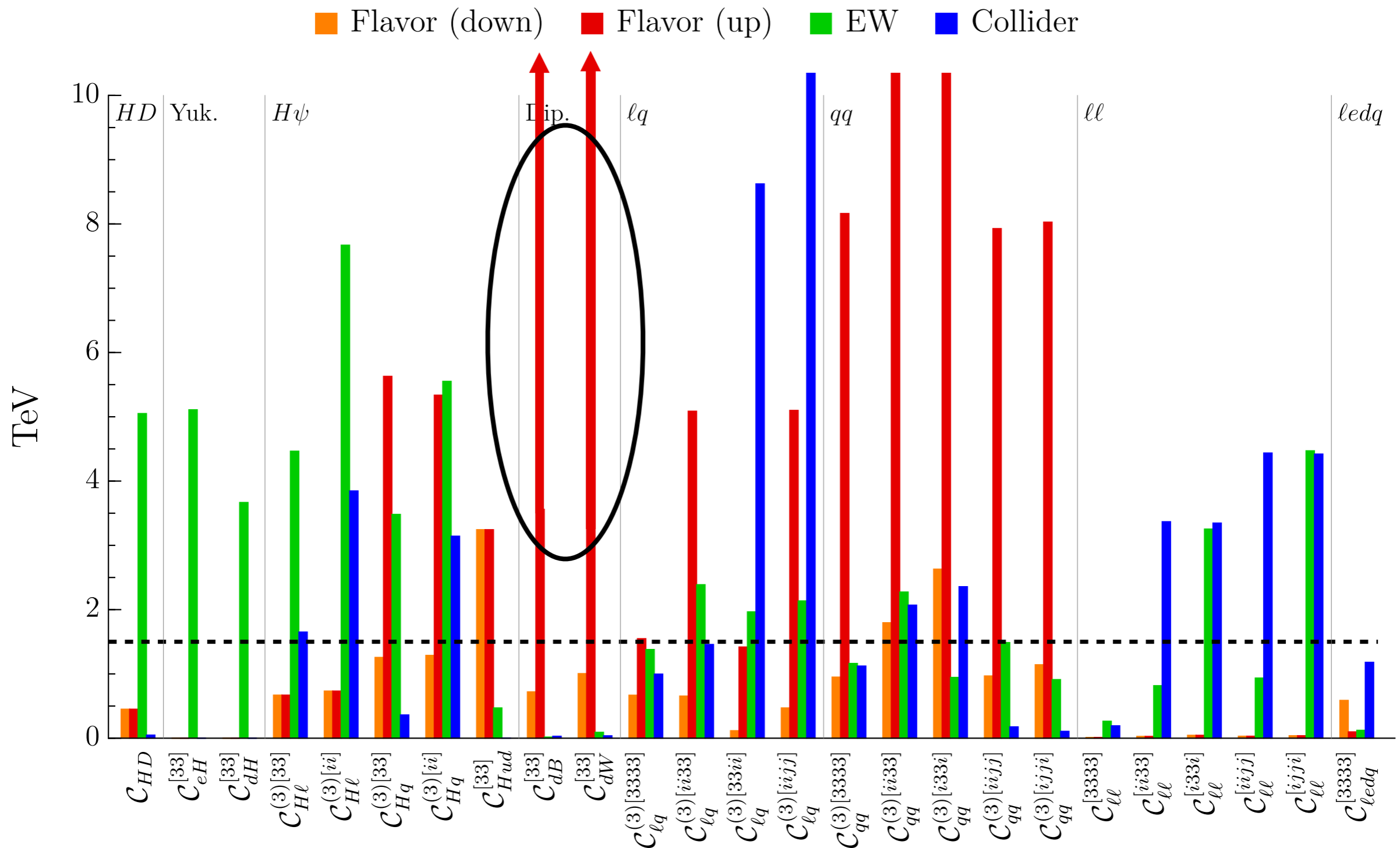
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Can we make the bounds on ALL other operators compatible with 1.5 TeV for reasonable values for the suppression factors ϵ_q, ϵ_l , and ϵ_H ?

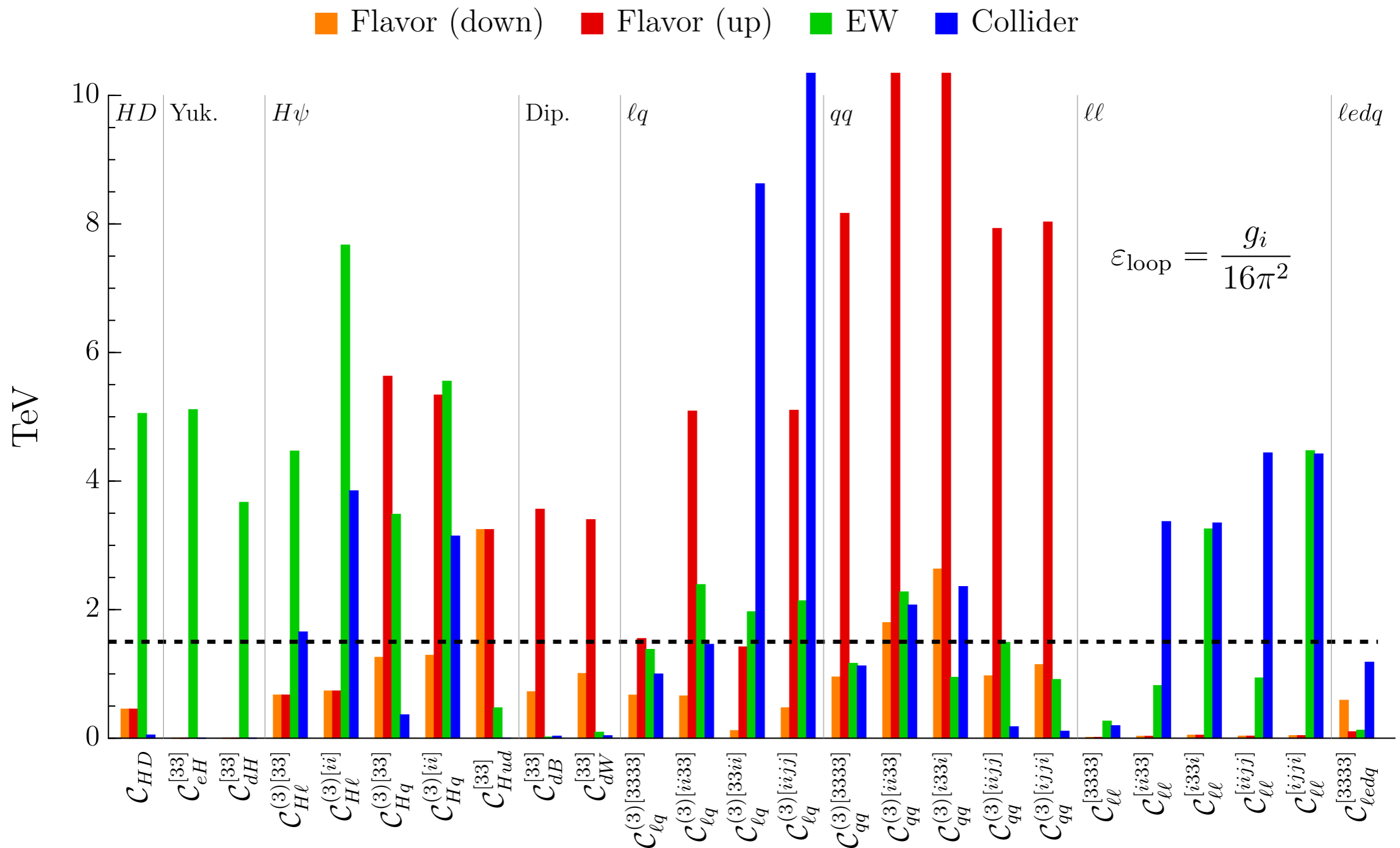
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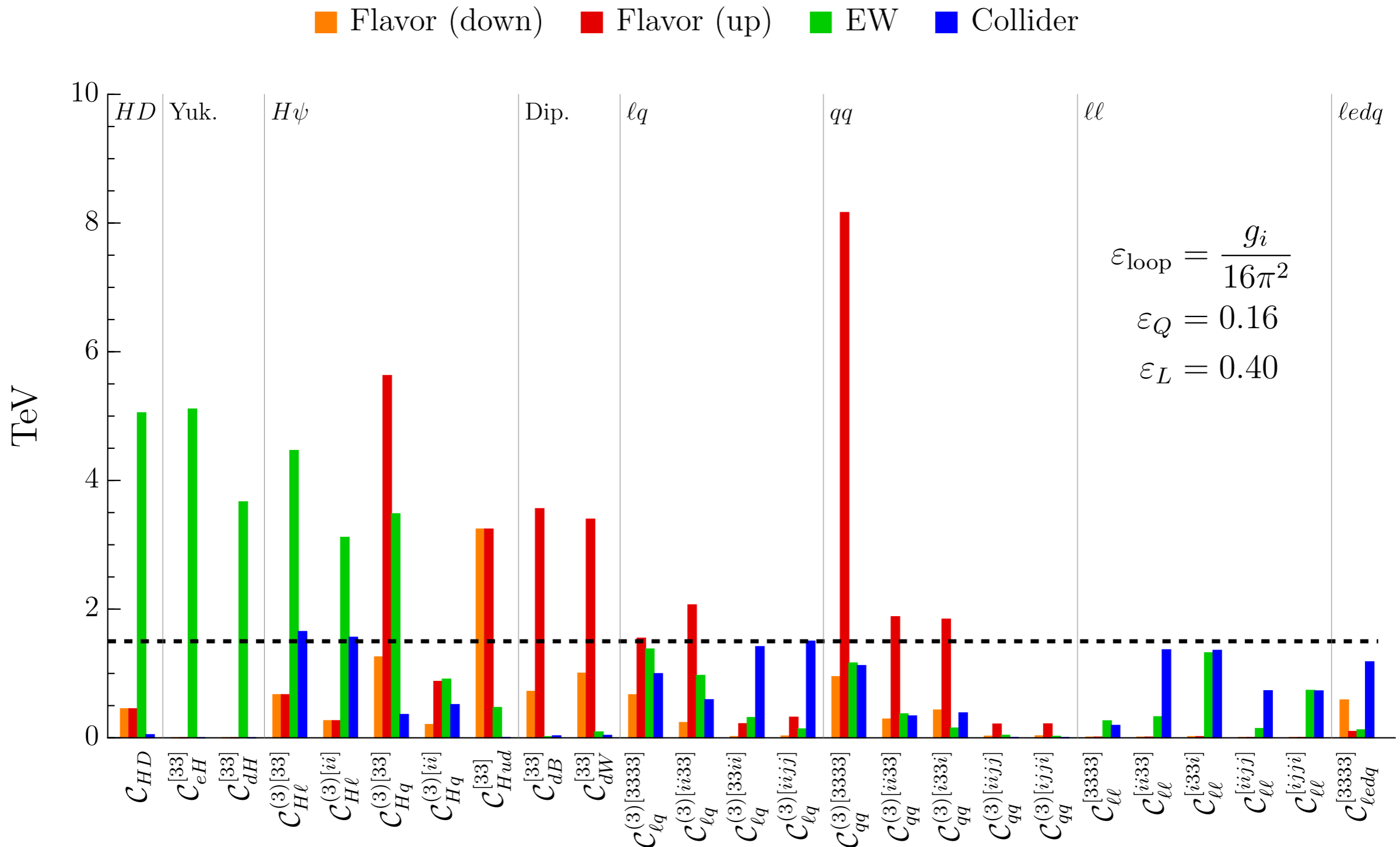
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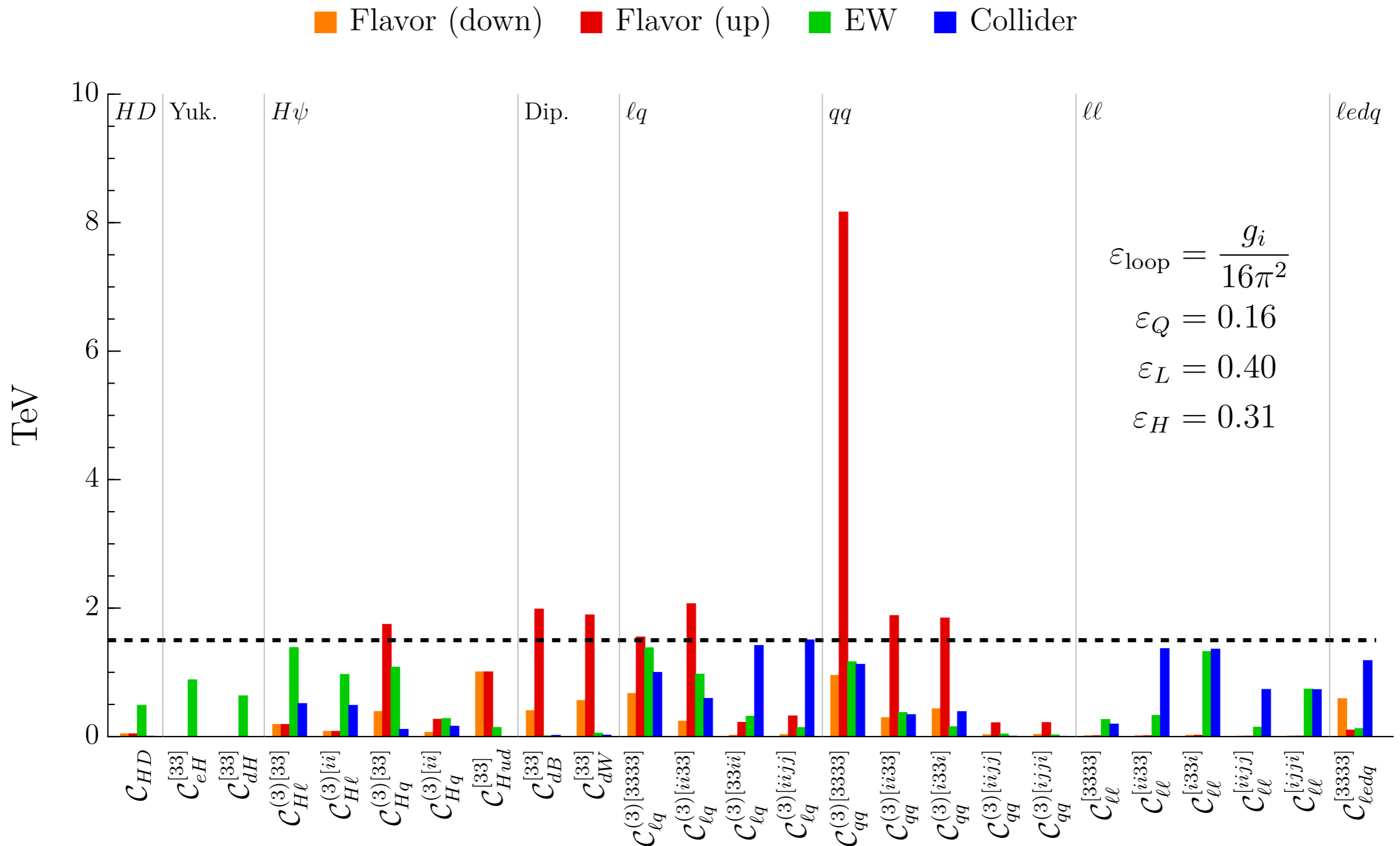
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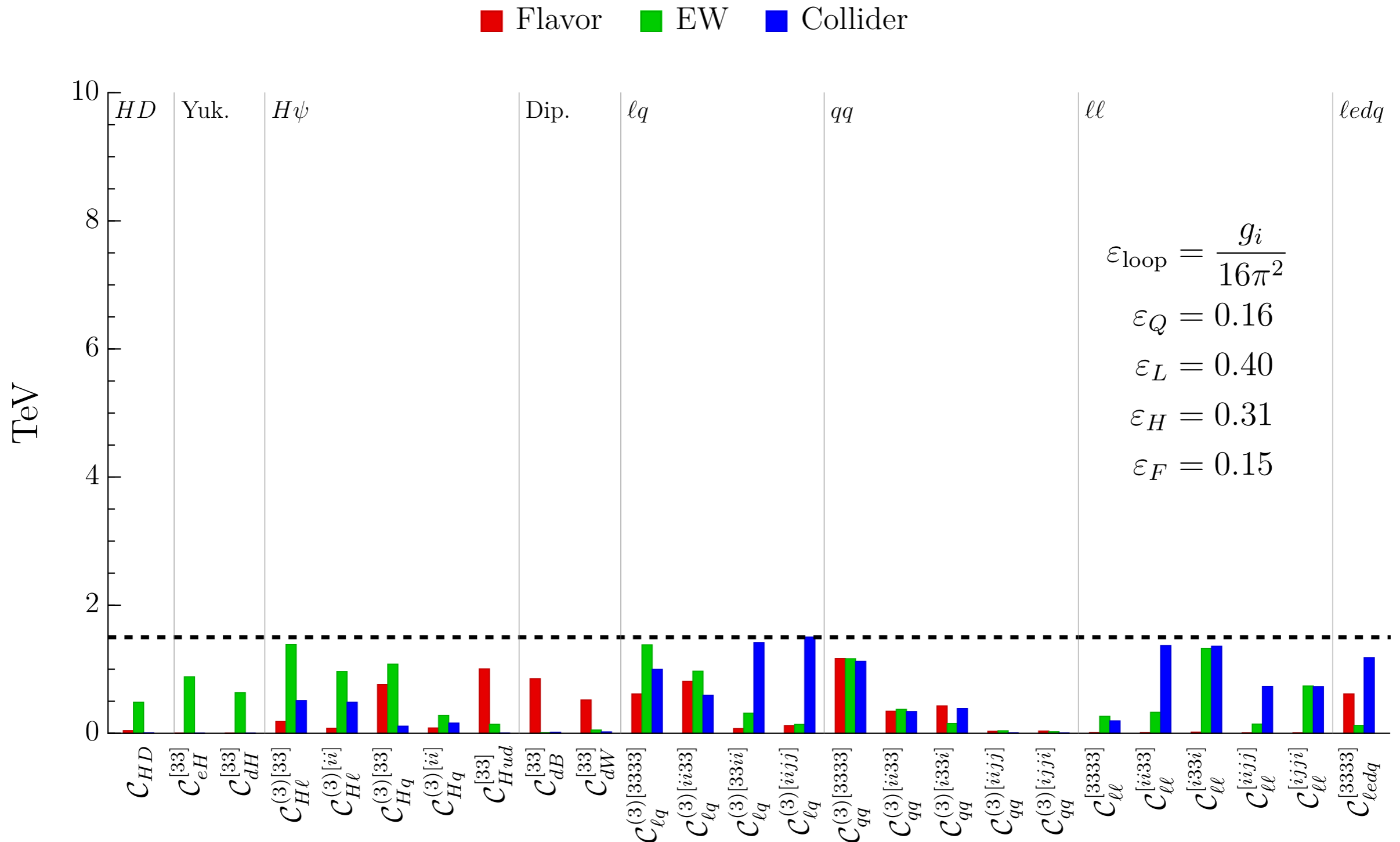
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New Physics mainly coupled to the 3rd generation compatible with all current data can exist at scales as low as 1.5 TeV under these conditions:

$$\varepsilon_q \leq 0.16, \quad \varepsilon_l \leq 0.40, \quad \varepsilon_H \leq 0.31, \quad \varepsilon_F \leq 0.15$$

The precise numbers are not “special”, but give a semi-quantitative **indication** of the general UV conditions NP models must meet to exist at nearby scales.

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...How would these bounds look like with a future tera Z machine, like FCC-ee?

Projections for FCC-ee

The expected improvements for Z- and W-pole observables, Higgs and tau decays are available from the literature.

[J. De Blas, G. Durieux, C. Grojean, J. Gu and A. Paul, [1907.04311](#), A. Blondel and P. Janot, [2106.13885](#), Snowmass [2203.06520](#)]

Tera Z- pole run: **10^5 more Z bosons than LEP**, so statistics can improve by up to a factor 300.

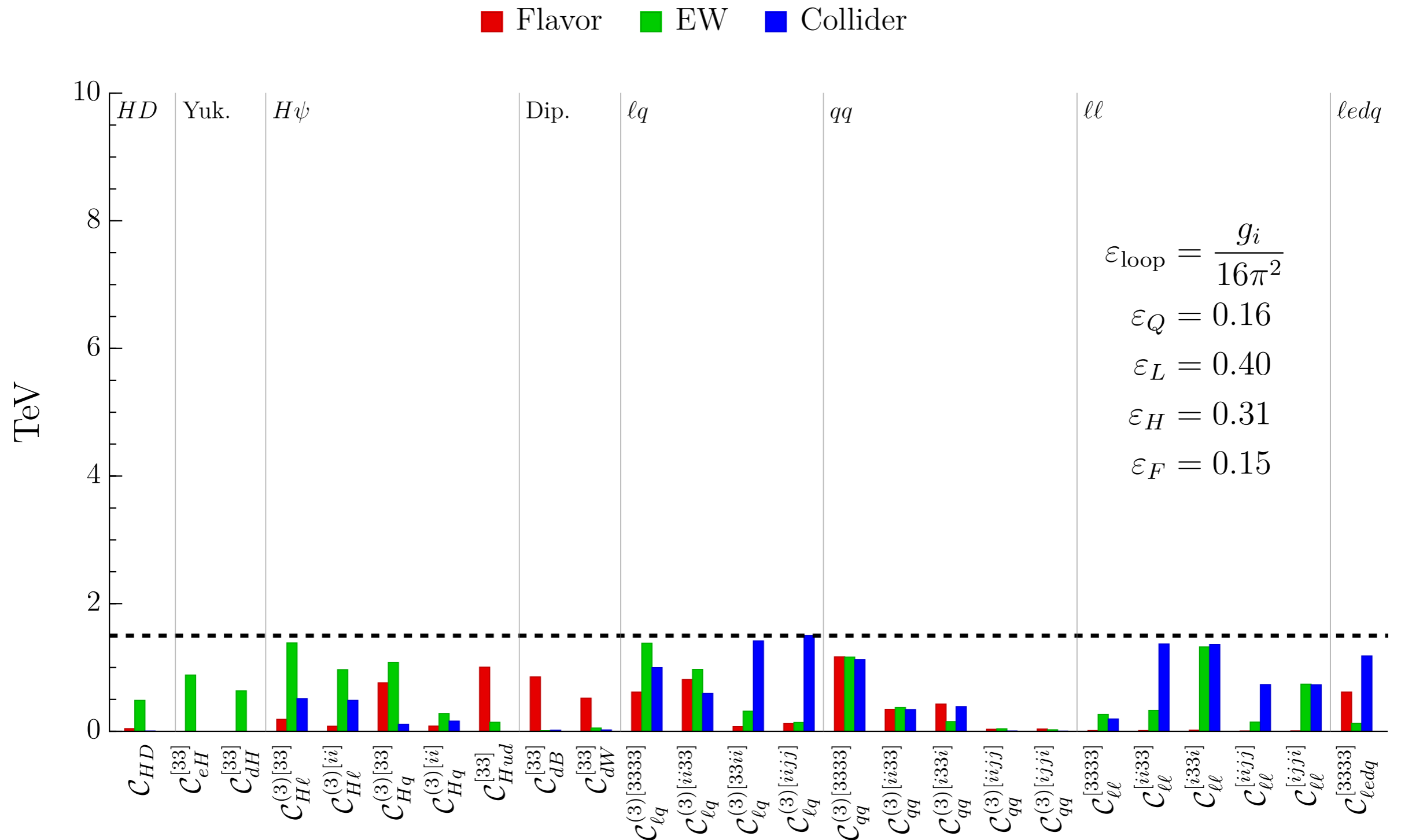
In practice, **leptonic** (**hadronic**) obs. improve by a factor **10-100** (**10**).

To build a projected EW likelihood for FCC-ee:

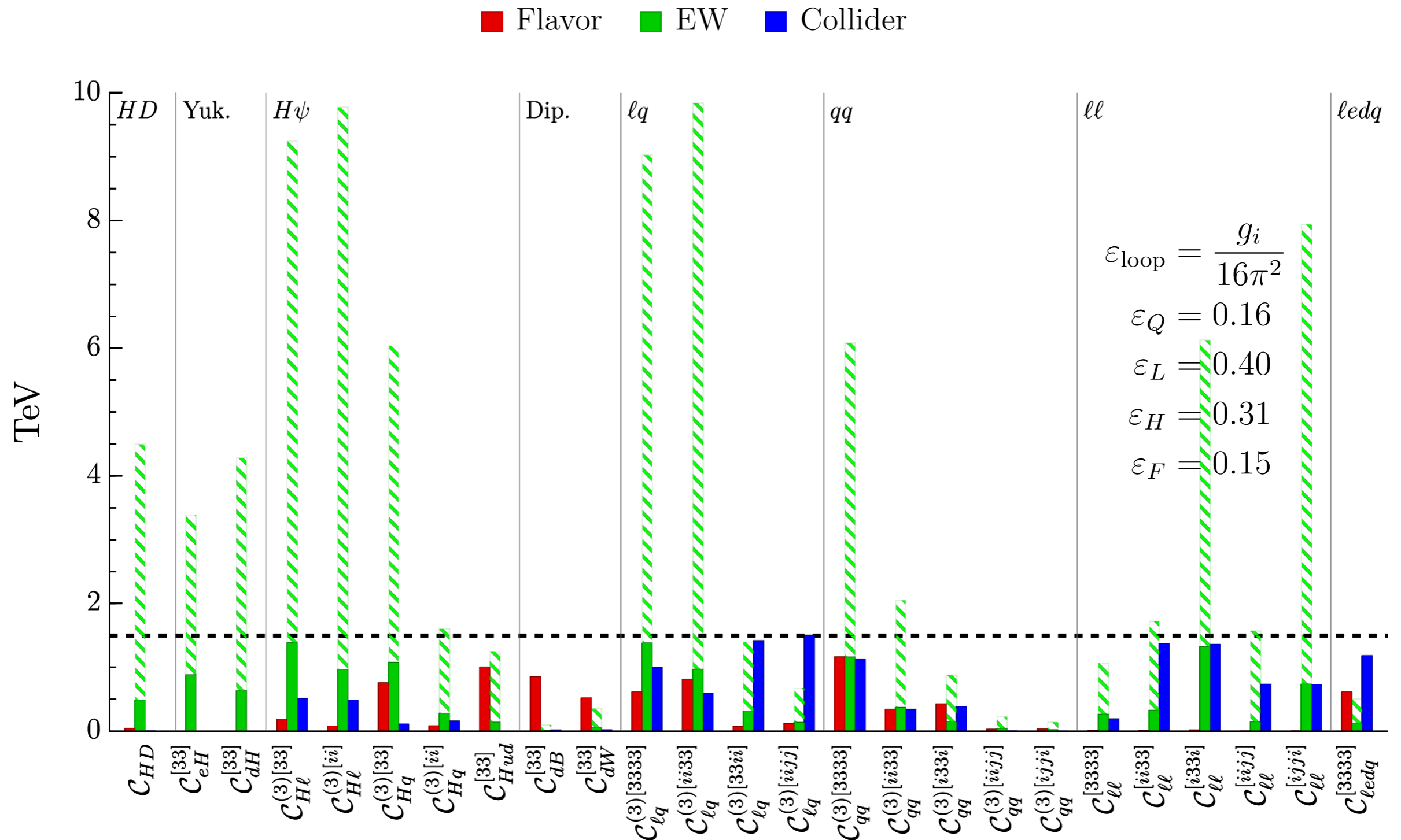
- Exp. values set to the SM
- error reduction as tabulated in the literature

Observable	Proj. Error Reduction
Γ_Z	23
σ_{had}^0	7.4
R_b	10.2
R_c	11.6
$A_{\text{FB}}^{0,b}$	15.5
$A_{\text{FB}}^{0,c}$	15.4
A_b	7.13
A_c	5.05
R_e	8.03
R_μ	31.8
R_τ	21.7
$A_{\text{FB}}^{0,e}$	30.8
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$A_{\text{FB}}^{0,\tau}$	21
A_e^{**}	130
A_μ^{**}	680
A_τ^{**}	340

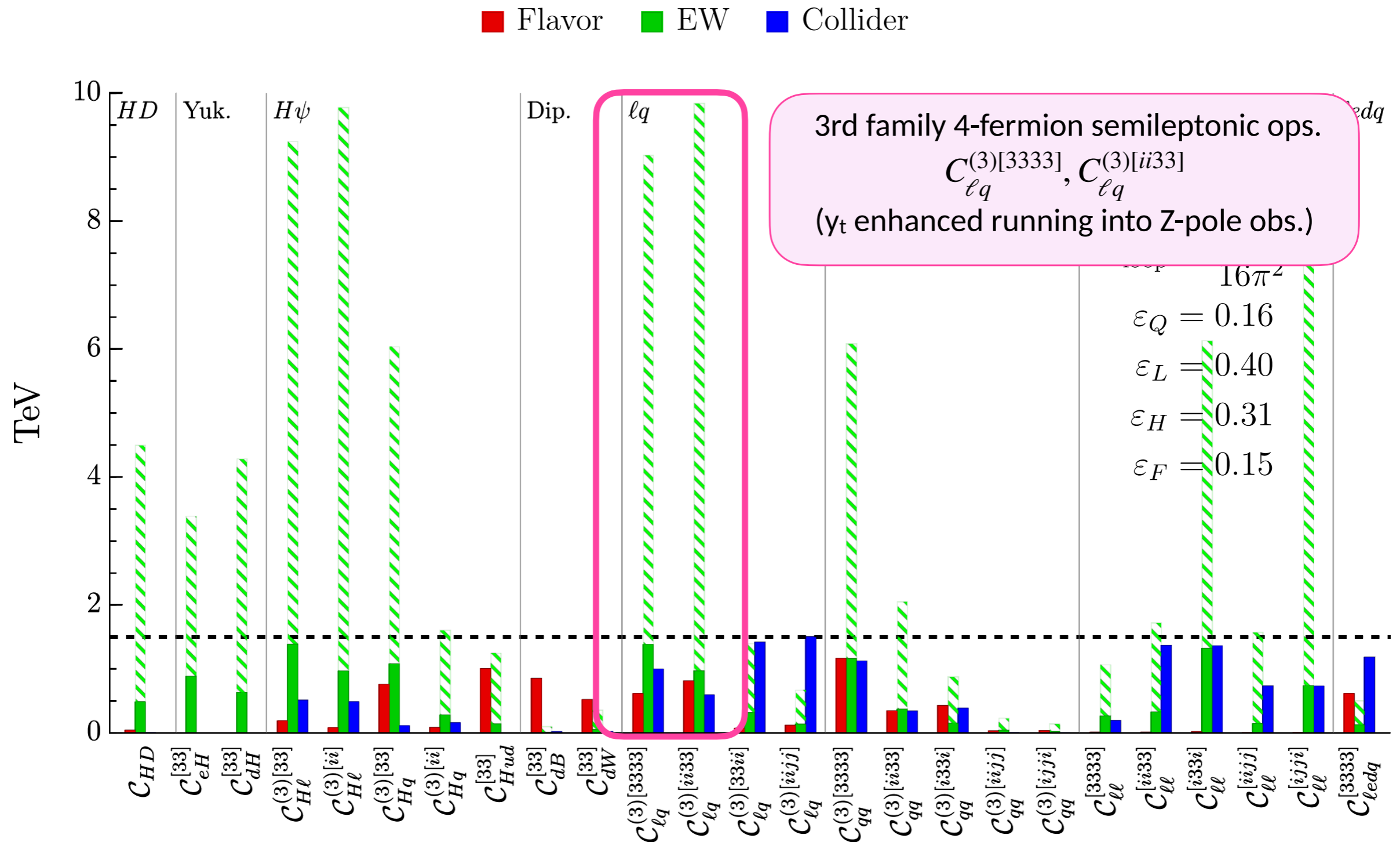
FCC and 3rd generation NP



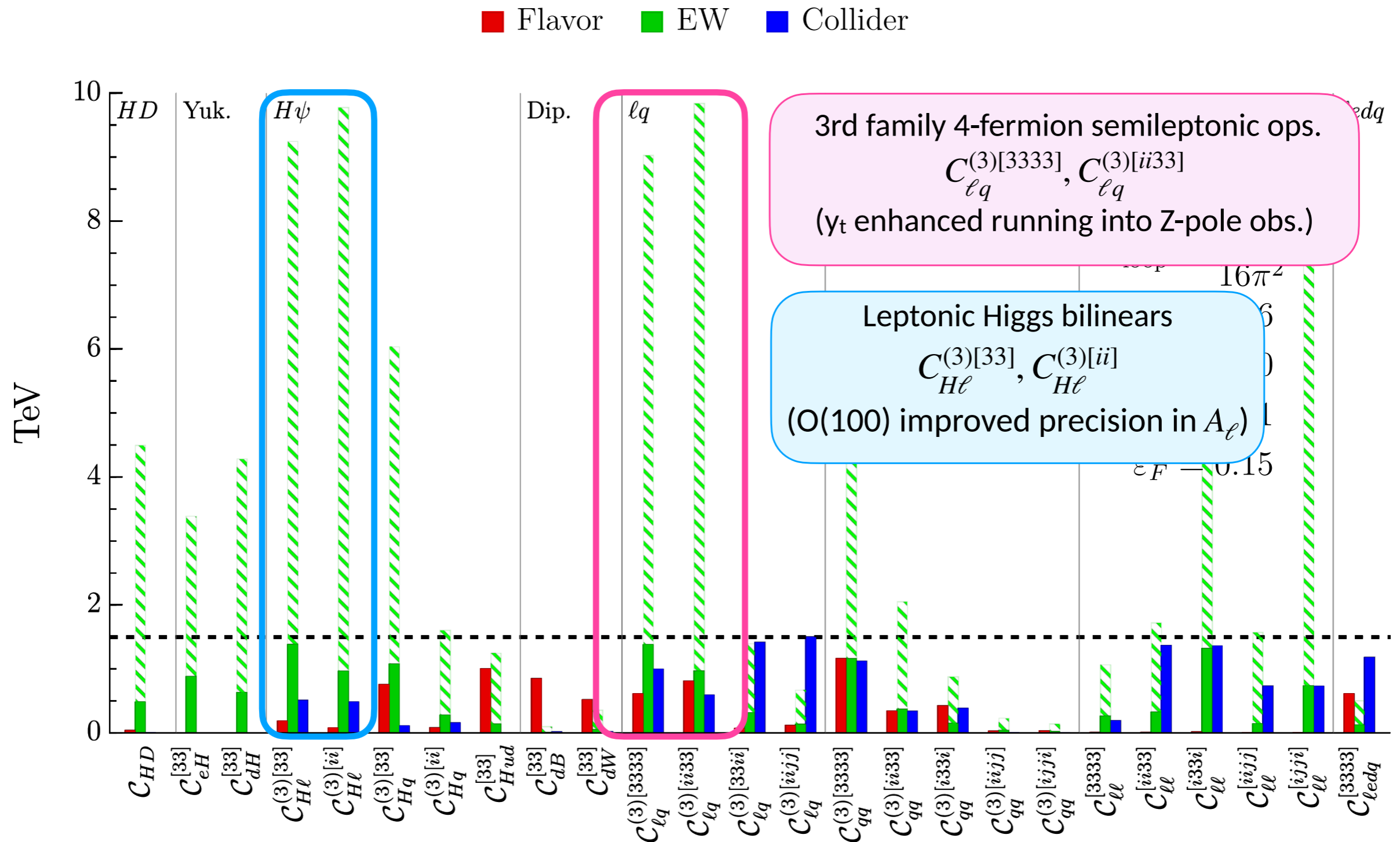
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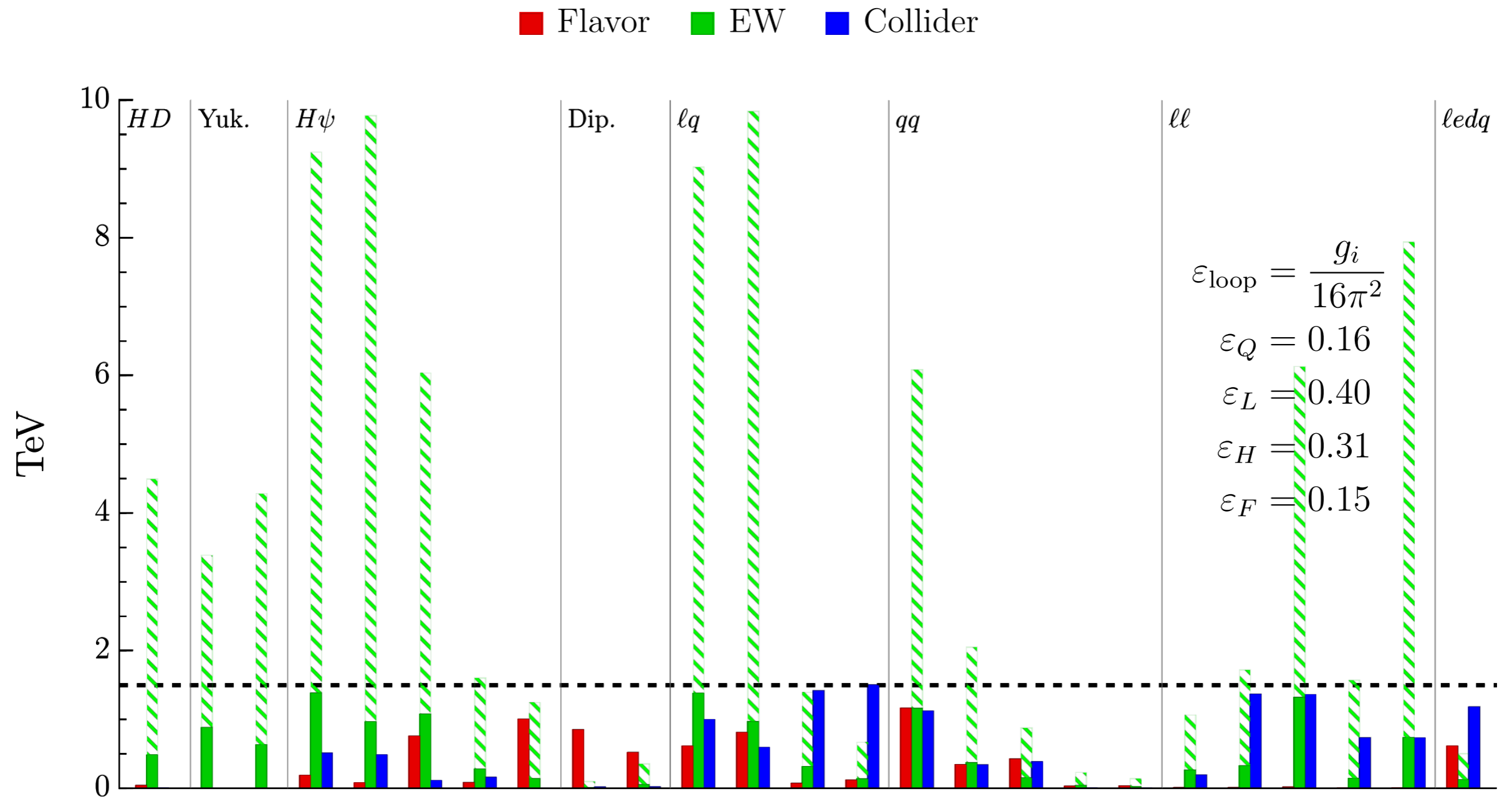
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FCC-ee could probe third-generation New Physics up to ~ 10 TeV!

Rare decays and 3rd generation NP

More short-term, improvements in flavor and collider observables can help us probe this scenario. Consider the **rare decays** $B \rightarrow K\nu\bar{\nu}$ and $K \rightarrow \pi\nu\bar{\nu}$.

$$\frac{\mathcal{B}(B^+ \rightarrow K^+ \nu\bar{\nu})_{\text{exp}}}{\mathcal{B}(B^+ \rightarrow K^+ \nu\bar{\nu})_{\text{SM}}} = 2.8 \pm 0.8,$$



[Exp: combination from Belle II @EPS 2023]

~3 σ tension with the SM

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[Exp: NA62 2021; SM: Buras et al. 2015]

Compatible with the SM at 1 σ



- theoretically clean
- significant improvements expected in the next years:
Belle II will measure $B \rightarrow K\nu\bar{\nu}$ @ 1%, and NA62(HIKE) $K \rightarrow \pi\nu\bar{\nu}$ @ 15%(5%)

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- sensitive to a limited number of EFT operators: $C_{\ell q}^{(3)[3333]}$, $C_{\ell q}^{(1)[3333]}$

Rare decays and 3rd generation NP

More short-term, improvements in flavor and collider observables can help us probe this scenario. Consider the **rare decays** $B \rightarrow K\nu\bar{\nu}$ and $K \rightarrow \pi\nu\bar{\nu}$.

$$\frac{\mathcal{B}(B^+ \rightarrow K^+ \nu\bar{\nu})_{\text{exp}}}{\mathcal{B}(B^+ \rightarrow K^+ \nu\bar{\nu})_{\text{SM}}} = 2.8 \pm 0.8,$$



[Exp: combination from Belle II @EPS 2023]

~3 σ tension with the SM

$$\frac{\mathcal{B}(K^+ \rightarrow \pi^+ \nu\bar{\nu})_{\text{exp}}}{\mathcal{B}(K^+ \rightarrow \pi^+ \nu\bar{\nu})_{\text{SM}}} = 1.23 \pm 0.39$$

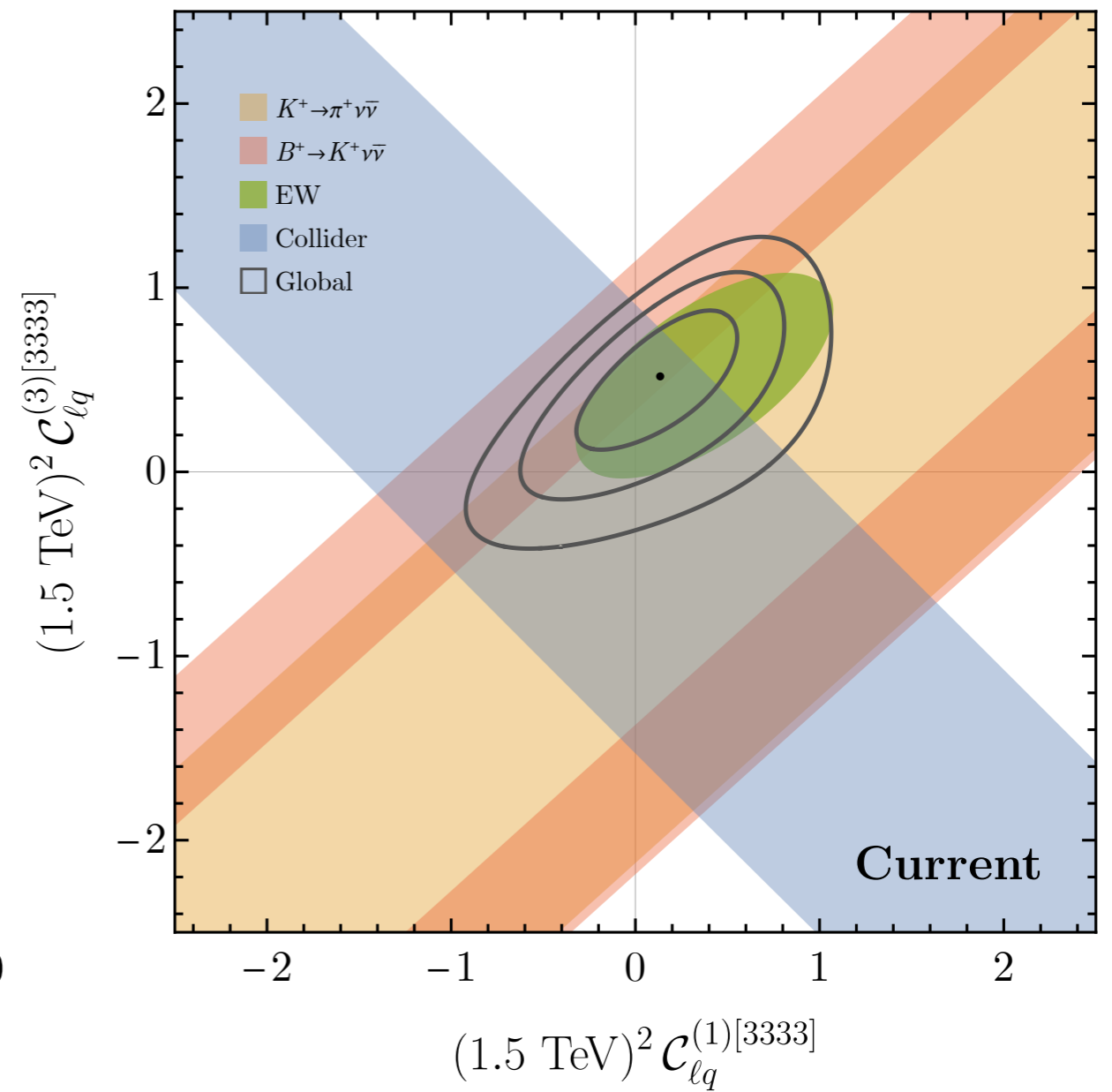
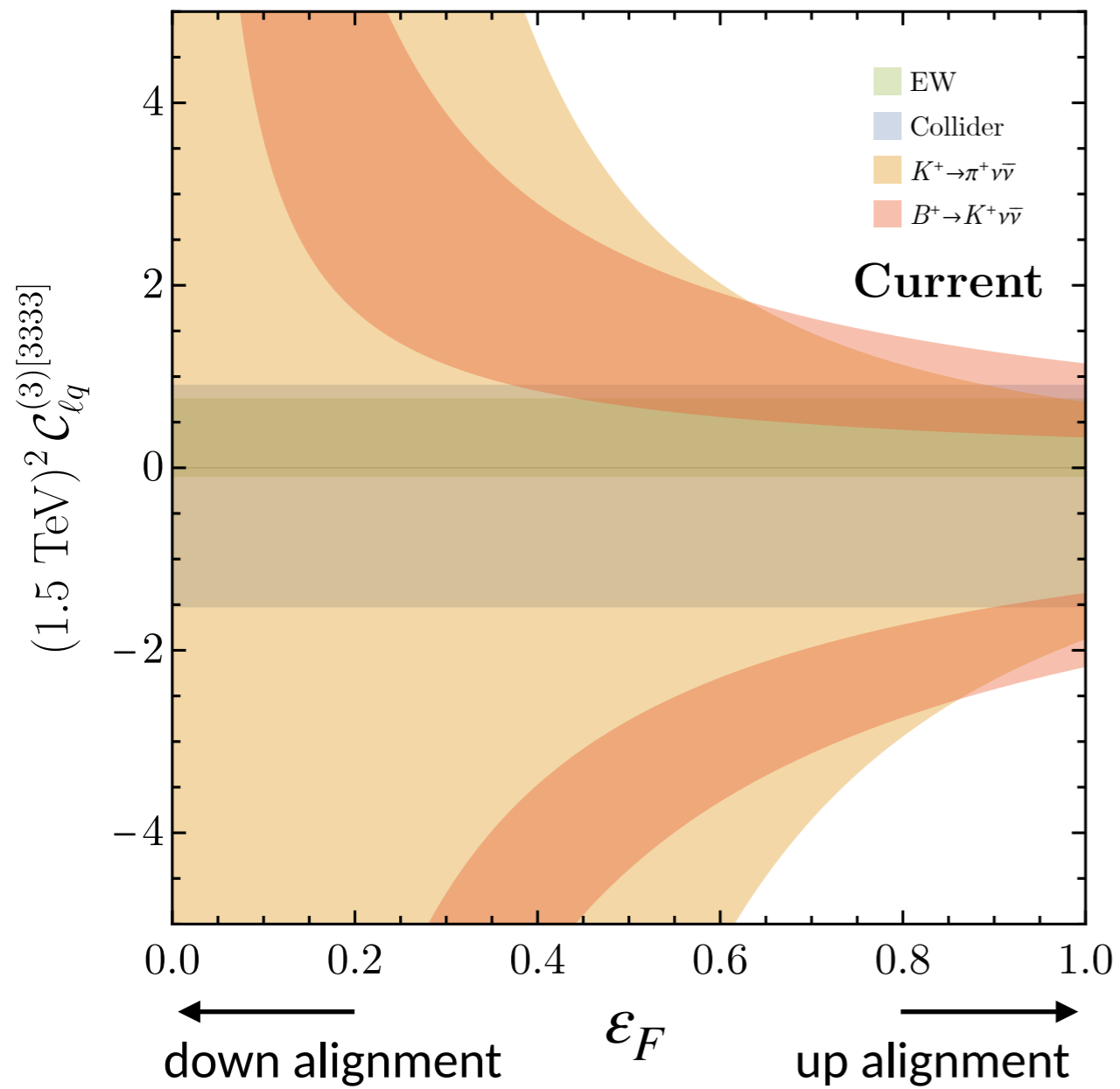
[Exp: NA62 2021; SM: Buras et al. 2015]

Compatible with the SM at 1 σ

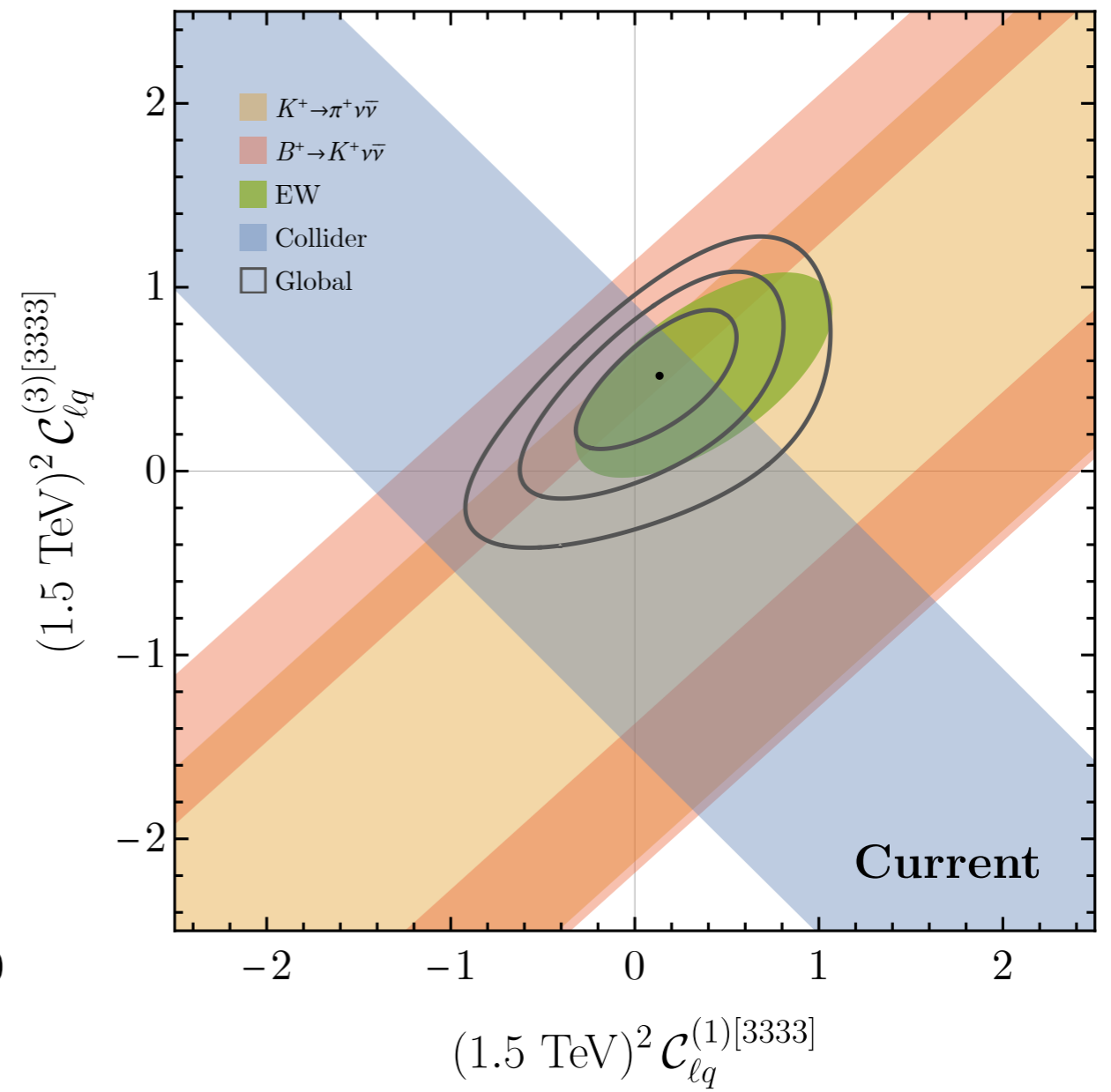
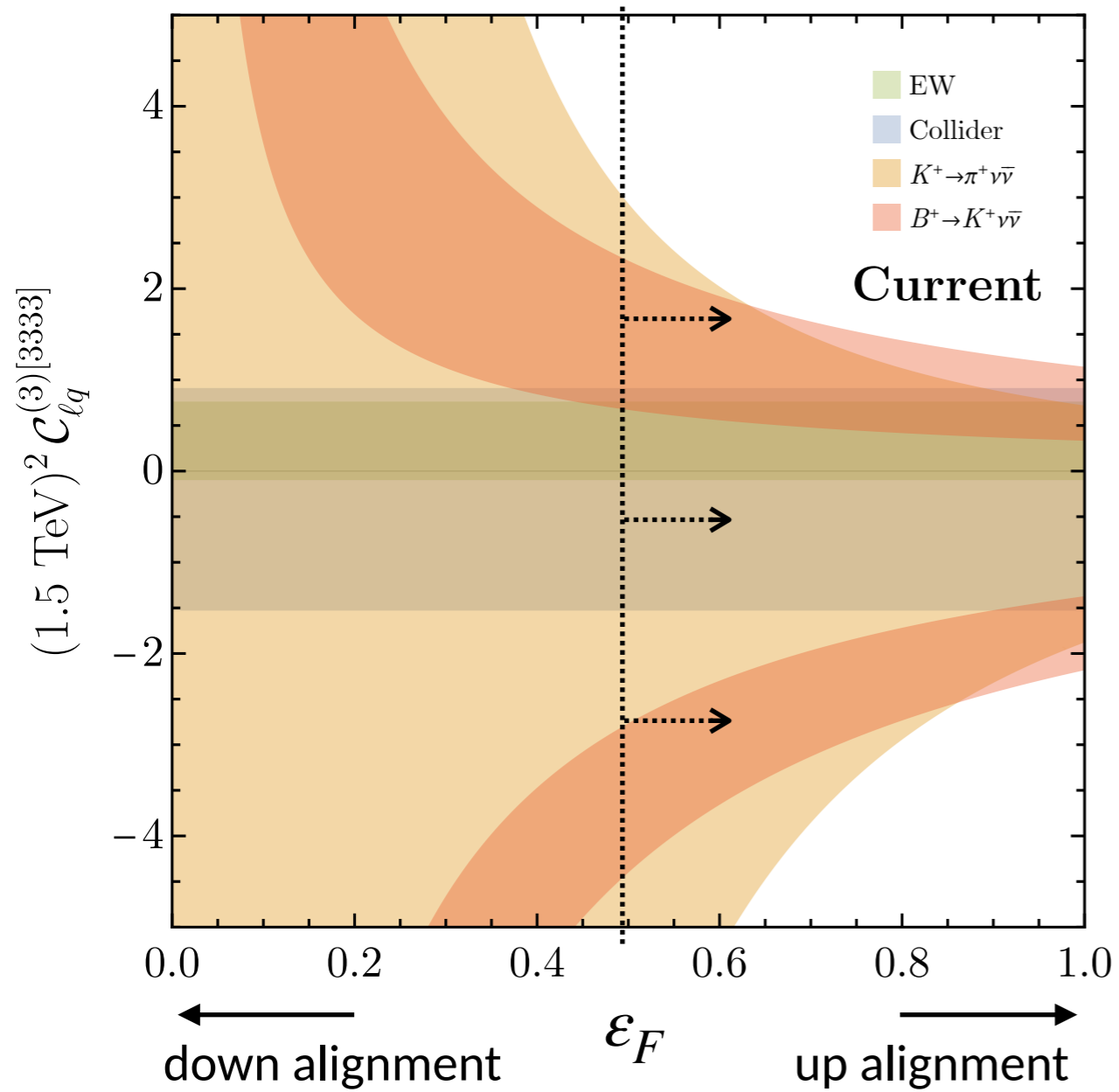


- theoretically clean
- significant improvements expected in the next years:
Belle II will measure $B \rightarrow K\nu\bar{\nu}$ @ 1%, and NA62(HIKE) $K \rightarrow \pi\nu\bar{\nu}$ @ 15%(5%)
- sensitive to a limited number of EFT operators: $C_{\ell q}^{(3)[3333]}$, $C_{\ell q}^{(1)[3333]}$
- scale differently with the alignment parameter ε_F

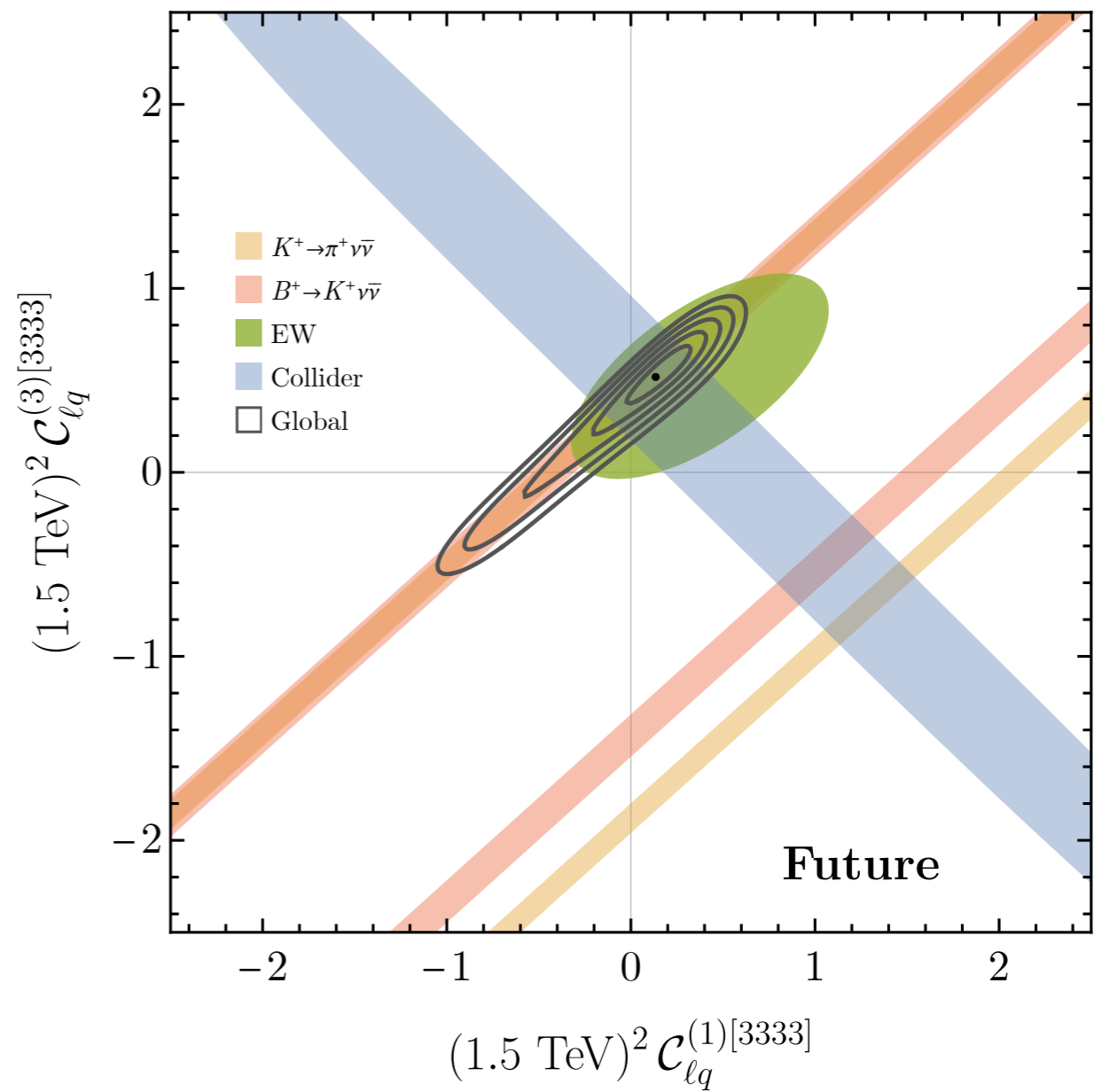
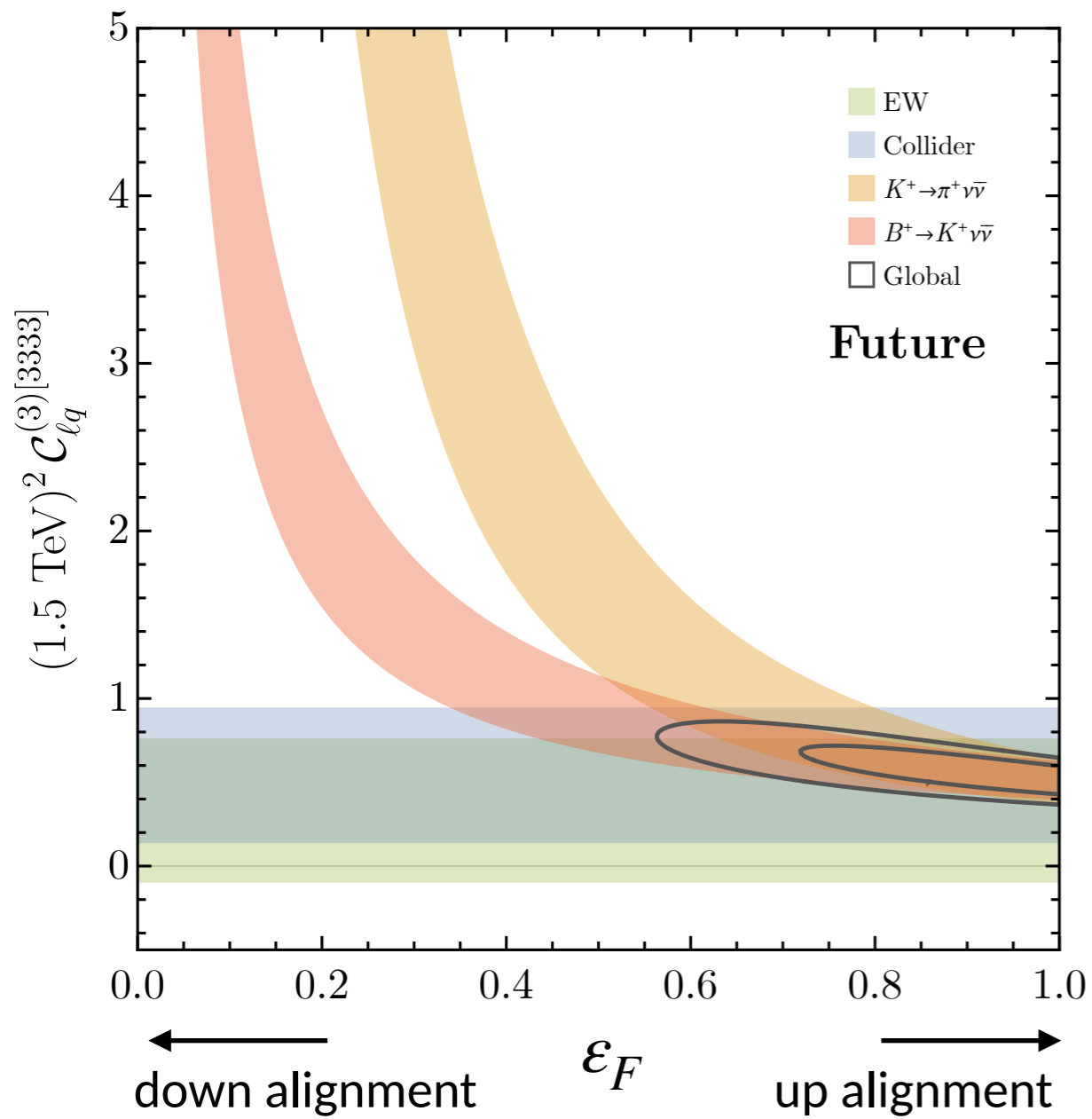
Rare decays and 3rd generation NP: current data



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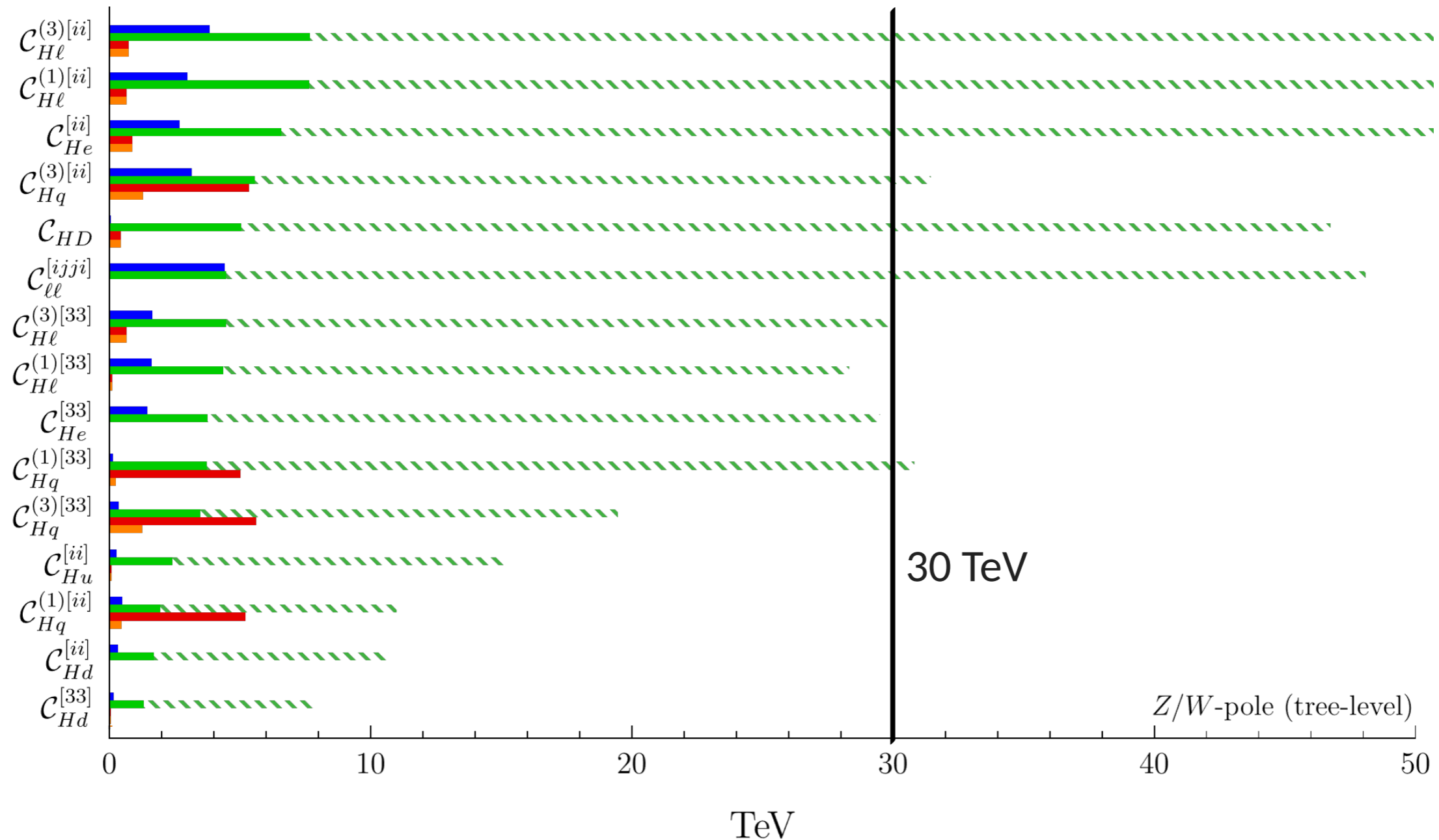


Rare decays and 3rd generation NP: projections



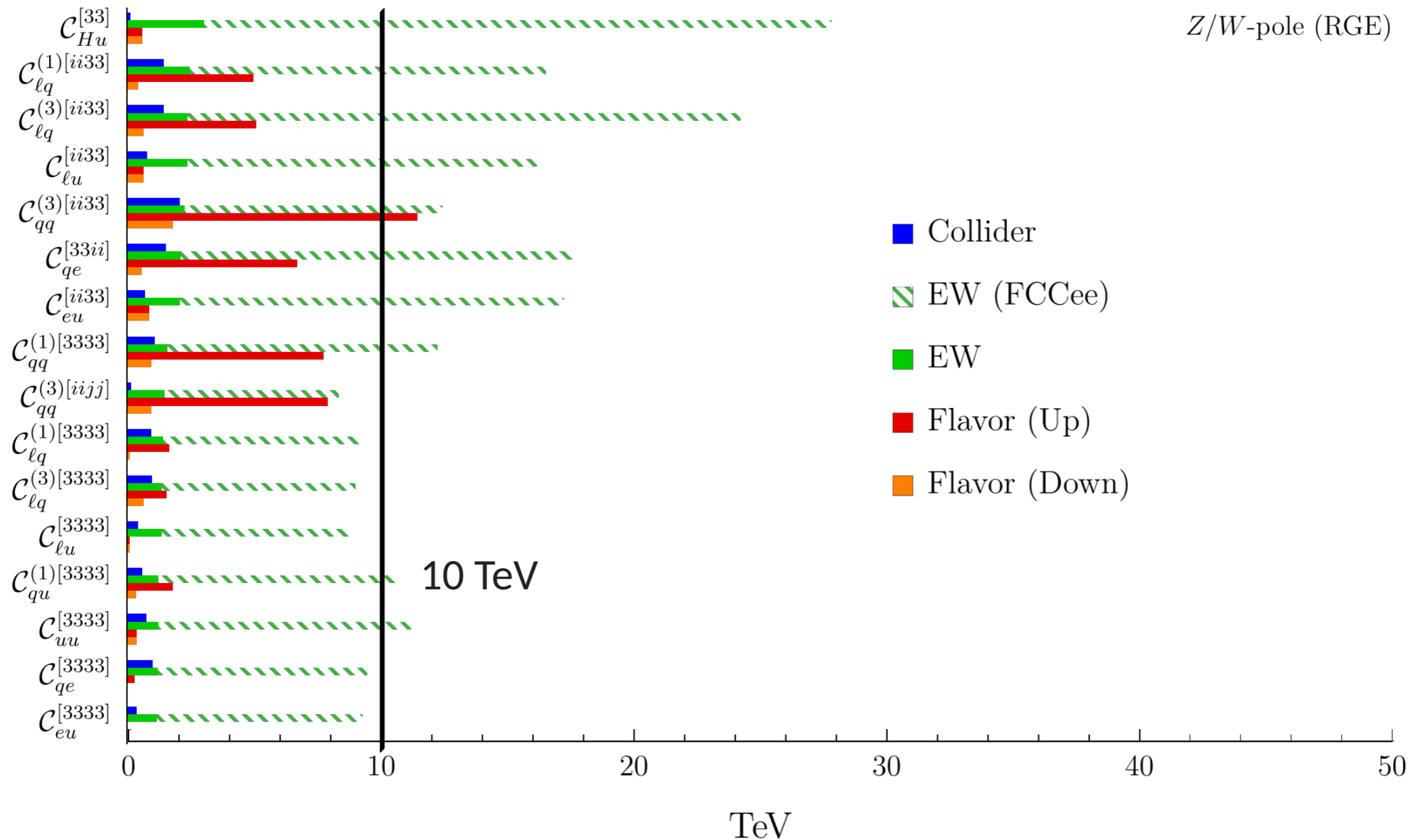
FCC-ee for “generic” U(2) NP (no suppression factors)

- Operators entering Z-pole observables at tree-level get bounds of 30-50 TeV



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Two comments:

- A future EW precision machine such as FCC-ee is the best way to probe NP with sizeable couplings to the Higgs
- NP that does not couple directly to the Higgs but does couple to the 3rd generation can be probed up to effective scales of about 10 TeV

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FCC-ee can push most of the existing bounds on NP from the EW sector by one order of magnitude!

Conclusions

We investigated NP scenarios characterized by a $U(2)^5$ symmetry acting on the light families. We included EW, flavor, and collider data, and accounted for RG effects.

Our main focus was **NP coupled mostly to the 3rd generation**, because of its strong theoretical motivation.

1. **How low** can the energy **scale of new physics** be for these class of models, and which conditions make this possible?
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We investigated NP scenarios characterized by a $U(2)^5$ symmetry acting on the light families. We included EW, flavor, and collider data, and accounted for RG effects.

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1. **How low** can the energy **scale of new physics** be for these class of models, and which conditions make this possible?
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1. **NP in the 3rd family is compatible with a scale as low as 1.5 TeV** under simple, non-tuned assumptions. Well-motivated NP models can be nearby!
2. **Precision flavor measurement** can help probe this scenario. For example, the rare decays $B \rightarrow K\nu\bar{\nu}$ and $K \rightarrow \pi\nu\bar{\nu}$ can help determine the orientation of the 3rd family in flavor space.

A future tera-Z machine like **FCC-ee** can probe third-generation NP up to **10 TeV**.

Back-up slides

Higgs bi-fermion operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	Λ_{EW}	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
$\mathcal{C}_{H\ell}^{(1)[33]}$	0.1	0.1	4.4	1.6	4.3	R_τ	4.3	R_τ
$\mathcal{C}_{H\ell}^{(1)[ii]}$	0.7	0.7	7.6	3.	7.8	σ_{had}	7.8	σ_{had}
$\mathcal{C}_{H\ell}^{(3)[33]}$	0.7	0.7	4.5	1.7	4.4	R_τ	4.4	R_τ
$\mathcal{C}_{H\ell}^{(3)[ii]}$	0.7	0.7	7.7	3.8	7.7	σ_{had}	7.7	σ_{had}
$\mathcal{C}_{He}^{[33]}$	-	-	3.8	1.5	3.7	R_τ	3.7	R_τ
$\mathcal{C}_{He}^{[ii]}$	0.9	0.9	6.6	2.7	6.7	σ_{had}	6.7	σ_{had}
$\mathcal{C}_{Hq}^{(1)[33]}$	0.3	5.	3.7	0.1	3.7	Γ_Z	5.1	$B_s \rightarrow \mu\mu$
$\mathcal{C}_{Hq}^{(1)[ii]}$	0.5	5.2	1.9	0.5	2.	R_c	5.4	$B_s \rightarrow \mu\mu$
$\mathcal{C}_{Hq}^{(3)[33]}$	1.3	5.6	3.5	0.4	3.4	R_b	5.5	$B_s \rightarrow \mu\mu$
$\mathcal{C}_{Hq}^{(3)[ii]}$	1.3	5.3	5.6	3.1	5.7	R_τ	7.7	Γ_Z
$\mathcal{C}_{Hd}^{[33]}$	-	-	1.3	0.2	1.3	R_b	1.3	R_b
$\mathcal{C}_{Hd}^{[ii]}$	-	-	1.7	0.3	1.7	R_τ	1.7	R_τ
$\mathcal{C}_{Hu}^{[33]}$	0.6	0.6	3.	0.1	3.1	A_b^{FB}	3.1	A_b^{FB}
$\mathcal{C}_{Hu}^{[ii]}$	-	-	2.4	0.3	2.4	R_τ	2.4	R_τ

3H and dipole operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	Λ_{EW}	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
$\mathcal{C}_{eH}^{[33]}$	-	-	5.1	-	5.1	$H \rightarrow \tau\tau$	5.1	$H \rightarrow \tau\tau$
$\mathcal{C}_{uH}^{[33]}$	-	-	0.2	-	0.2	$H \rightarrow \tau\tau$	0.2	$H \rightarrow \tau\tau$
$\mathcal{C}_{dH}^{[33]}$	-	-	3.7	-	3.7	$H \rightarrow bb$	3.7	$H \rightarrow bb$
$\mathcal{C}_{Hud}^{[33]}$	3.2	3.2	0.5	-	3.2	$B \rightarrow X_s\gamma$	3.2	$B \rightarrow X_s\gamma$
$\mathcal{C}_{eB}^{[33]}$	-	-	0.2	1.2	1.2	$pp \rightarrow \tau\tau$	1.2	$pp \rightarrow \tau\tau$
$\mathcal{C}_{uB}^{[33]}$	0.7	0.8	2.4	1.9	2.7	A_b^{FB}	2.7	A_b^{FB}
$\mathcal{C}_{dB}^{[33]}$	15.2	74.8	0.4	0.7	15.2	$B \rightarrow X_s\gamma$	74.8	$B \rightarrow X_s\gamma$
$\mathcal{C}_{eW}^{[33]}$	-	-	1.	1.9	1.8	$pp \rightarrow \tau\nu$	1.8	$pp \rightarrow \tau\nu$
$\mathcal{C}_{uW}^{[33]}$	0.5	0.9	2.3	3.6	3.7	QuarkDipoles	3.8	QuarkDipoles
$\mathcal{C}_{dW}^{[33]}$	15.7	53.	1.4	0.6	15.7	$B \rightarrow X_s\gamma$	53.	$B \rightarrow X_s\gamma$
$\mathcal{C}_{uG}^{[33]}$	0.1	0.3	0.5	2.7	2.7	QuarkDipoles	2.7	QuarkDipoles
$\mathcal{C}_{dG}^{[33]}$	4.	25.5	0.3	-	4.	$B \rightarrow X_s\gamma$	25.5	$B \rightarrow X_s\gamma$

Scalar and tensor operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	Λ_{EW}	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
$\mathcal{C}_{ledq}^{[3333]}$	0.6	-	0.1	1.2	1.1	$pp \rightarrow \tau\tau$	1.2	$pp \rightarrow \tau\tau$
$\mathcal{C}_{quqd}^{(1)[3333]}$	1.8	5.5	1.7	0.4	2.2	$B \rightarrow X_s\gamma$	5.5	$B \rightarrow X_s\gamma$
$\mathcal{C}_{quqd}^{(8)[3333]}$	1.	5.1	0.7	0.2	1.	$B \rightarrow X_s\gamma$	5.1	$B \rightarrow X_s\gamma$
$\mathcal{C}_{lequ}^{(1)[3333]}$	-	-	2.1	-	2.1	$H \rightarrow \tau\tau$	2.1	$H \rightarrow \tau\tau$
$\mathcal{C}_{lequ}^{(3)[3333]}$	-	-	0.8	-	0.8	$H \rightarrow \tau\tau$	0.8	$H \rightarrow \tau\tau$

LLLL vector operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	Λ_{EW}	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
$\mathcal{C}_{\ell\ell}^{[3333]}$	-	-	0.3	0.2	0.3	σ_{had}	0.3	σ_{had}
$\mathcal{C}_{\ell\ell}^{[ii33]}$	-	-	0.8	3.4	3.3	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	3.3	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$\mathcal{C}_{\ell\ell}^{[i33i]}$	-	-	3.3	3.3	4.2	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	4.2	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$\mathcal{C}_{\ell\ell}^{[iijj]}$	-	-	0.9	4.4	4.4	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	4.4	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$\mathcal{C}_{\ell\ell}^{[ijji]}$	-	-	4.5	4.4	4.9	A_b^{FB}	4.9	A_b^{FB}
$\mathcal{C}_{qq}^{(1)[3333]}$	1.	7.8	1.6	1.1	1.7	Γ_Z	7.6	$ C_{B_s} $
$\mathcal{C}_{qq}^{(1)[ii33]}$	1.3	11.2	0.9	1.5	1.7	FourQuarksTop	11.3	$ C_{B_s} $
$\mathcal{C}_{qq}^{(1)[i33i]}$	2.5	11.3	0.7	1.6	2.6	$B_s \rightarrow \mu\mu$	11.3	$ C_{B_s} $
$\mathcal{C}_{qq}^{(1)[iijj]}$	0.9	8.1	0.4	-	0.9	$\text{Im}(C_D)$	8.1	$ C_{B_s} $
$\mathcal{C}_{qq}^{(1)[ijji]}$	1.1	8.1	0.5	-	1.	$\text{Im}(C_D)$	8.1	$ C_{B_s} $
$\mathcal{C}_{qq}^{(3)[3333]}$	1.	8.2	1.2	1.1	1.5	m_W	8.2	$ C_{B_s} $
$\mathcal{C}_{qq}^{(3)[ii33]}$	1.8	11.5	2.3	2.1	3.	R_b	11.3	$ C_{B_s} $
$\mathcal{C}_{qq}^{(3)[i33i]}$	2.6	11.2	0.9	2.4	3.1	$B_s \rightarrow \mu\mu$	11.3	$ C_{B_s} $
$\mathcal{C}_{qq}^{(3)[iijj]}$	1.	7.9	1.5	0.2	1.5	R_τ	7.9	$ C_{B_s} $
$\mathcal{C}_{qq}^{(3)[ijji]}$	1.1	8.	0.9	0.1	1.2	$K^+ \rightarrow \pi^+\nu\bar{\nu}$	8.	$ C_{B_s} $
$\mathcal{C}_{\ell q}^{(1)[3333]}$	0.1	1.7	1.4	1.	1.4	R_τ	1.6	$K^+ \rightarrow \pi^+\nu\bar{\nu}$
$\mathcal{C}_{\ell q}^{(1)[ii33]}$	0.4	5.	2.5	1.5	2.5	σ_{had}	5.1	$B_s \rightarrow \mu\mu$
$\mathcal{C}_{\ell q}^{(1)[33ii]}$	-	1.6	0.3	3.4	3.4	$pp \rightarrow \tau\tau$	3.4	$pp \rightarrow \tau\tau$
$\mathcal{C}_{\ell q}^{(1)[iijj]}$	0.5	5.	0.5	5.4	5.4	$pp \rightarrow \mu\mu$	5.6	$pp \rightarrow \mu\mu$
$\mathcal{C}_{\ell q}^{(3)[3333]}$	0.7	1.5	1.4	1.	1.6	R_τ	1.6	$K^+ \rightarrow \pi^+\nu\bar{\nu}$
$\mathcal{C}_{\ell q}^{(3)[ii33]}$	0.7	5.1	2.4	1.5	2.5	A_b^{FB}	5.	$B_s \rightarrow \mu\mu$
$\mathcal{C}_{\ell q}^{(3)[33ii]}$	0.1	1.4	2.	8.6	8.8	$pp \rightarrow \tau\nu$	8.7	$pp \rightarrow \tau\nu$
$\mathcal{C}_{\ell q}^{(3)[iijj]}$	0.5	5.1	2.1	22.5	22.5	$pp \rightarrow \mu\nu$	23.7	$pp \rightarrow \mu\nu$

RRRR vector operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	Λ_{EW}	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
$\mathcal{C}_{ee}^{[3333]}$	-	-	0.3	0.2	0.3	R_τ	0.3	R_τ
$\mathcal{C}_{ee}^{[ii33]}$	-	-	0.7	3.2	3.2	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	3.2	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$\mathcal{C}_{ee}^{[ijjj]}$	-	-	0.8	4.2	4.2	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	4.2	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$\mathcal{C}_{uu}^{[3333]}$	0.4	0.4	1.2	0.8	1.3	A_b^{FB}	1.3	A_b^{FB}
$\mathcal{C}_{uu}^{[ii33]}$	0.1	0.1	1.1	1.3	1.4	FourQuarksTop	1.4	FourQuarksTop
$\mathcal{C}_{uu}^{[i33i]}$	-	-	0.5	1.3	1.4	FourQuarksTop	1.4	FourQuarksTop
$\mathcal{C}_{uu}^{[ijjj]}$	-	-	0.3	-	0.3	R_τ	0.3	R_τ
$\mathcal{C}_{uu}^{[ijji]}$	-	-	0.3	-	0.3	R_τ	0.3	R_τ
$\mathcal{C}_{dd}^{[3333]}$	-	-	-	-	-	R_b	-	R_b
$\mathcal{C}_{dd}^{[ii33]}$	-	-	0.1	-	0.1	R_τ	0.1	R_τ
$\mathcal{C}_{dd}^{[i33i]}$	-	-	-	-	-	Γ_Z	-	Γ_Z
$\mathcal{C}_{dd}^{[ijjj]}$	-	-	0.2	-	0.2	R_τ	0.2	R_τ
$\mathcal{C}_{dd}^{[ijji]}$	-	-	0.1	-	0.1	R_τ	0.1	R_τ
$\mathcal{C}_{eu}^{[3333]}$	-	-	1.2	0.4	1.2	R_τ	1.2	R_τ
$\mathcal{C}_{eu}^{[ii33]}$	0.9	0.9	2.1	0.7	2.2	σ_{had}	2.2	σ_{had}
$\mathcal{C}_{eu}^{[33ii]}$	-	-	0.3	2.8	2.8	$pp \rightarrow \tau\tau$	2.8	$pp \rightarrow \tau\tau$
$\mathcal{C}_{eu}^{[ijjj]}$	-	-	0.6	7.4	7.4	$pp \rightarrow ee$	7.4	$pp \rightarrow ee$
$\mathcal{C}_{ed}^{[3333]}$	-	-	0.2	1.	1.	$pp \rightarrow \tau\tau$	1.	$pp \rightarrow \tau\tau$
$\mathcal{C}_{ed}^{[ii33]}$	-	-	0.3	1.5	1.5	$pp \rightarrow \mu\mu$	1.5	$pp \rightarrow \mu\mu$
$\mathcal{C}_{ed}^{[33ii]}$	-	-	0.2	2.8	2.8	$pp \rightarrow \tau\tau$	2.8	$pp \rightarrow \tau\tau$
$\mathcal{C}_{ed}^{[ijjj]}$	-	-	0.4	4.4	4.4	$pp \rightarrow \mu\mu$	4.4	$pp \rightarrow \mu\mu$
$\mathcal{C}_{ud}^{(1)[3333]}$	0.1	0.1	0.4	0.3	0.4	R_b	0.4	R_b
$\mathcal{C}_{ud}^{(1)[ii33]}$	-	-	0.1	-	0.1	R_τ	0.1	R_τ
$\mathcal{C}_{ud}^{(1)[33ii]}$	-	-	0.5	1.2	1.2	FourQuarksTop	1.2	FourQuarksTop
$\mathcal{C}_{ud}^{(1)[ijjj]}$	-	-	0.2	-	0.2	R_τ	0.2	R_τ
$\mathcal{C}_{ud}^{(8)[3333]}$	0.1	0.1	-	0.2	0.2	FourQuarksBottom	0.2	FourQuarksBottom
$\mathcal{C}_{ud}^{(8)[ii33]}$	-	-	-	-	-	-	-	-
$\mathcal{C}_{ud}^{(8)[33ii]}$	-	-	0.1	0.7	0.7	FourQuarksTop	0.7	FourQuarksTop
$\mathcal{C}_{ud}^{(8)[ijjj]}$	-	-	-	-	-	-	-	-

LLRR vector operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	Λ_{EW}	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
$\mathcal{C}_{\ell e}^{[3333]}$	-	-	0.2	0.1	0.2	A_τ	0.2	A_τ
$\mathcal{C}_{\ell e}^{[ii33]}$	-	-	0.4	2.	1.9	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	1.9	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$\mathcal{C}_{\ell e}^{[33ii]}$	-	-	0.3	1.9	2.	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	2.	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$\mathcal{C}_{\ell e}^{[iijj]}$	-	-	0.5	3.8	3.8	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$	3.8	$(e^+e^- \rightarrow \mu^+\mu^-)_{\text{FB}}$
$\mathcal{C}_{\ell u}^{[3333]}$	0.1	0.1	1.4	0.4	1.3	R_τ	1.3	R_τ
$\mathcal{C}_{\ell u}^{[ii33]}$	0.7	0.7	2.4	0.8	2.3	σ_{had}	2.3	σ_{had}
$\mathcal{C}_{\ell u}^{[33ii]}$	-	-	-	-	-	-	-	-
$\mathcal{C}_{\ell u}^{[iijj]}$	-	-	-	-	-	-	-	-
$\mathcal{C}_{\ell d}^{[3333]}$	-	-	-	-	-	-	-	-
$\mathcal{C}_{\ell d}^{[ii33]}$	-	-	-	-	-	-	-	-
$\mathcal{C}_{\ell d}^{[33ii]}$	-	-	-	-	-	-	-	-
$\mathcal{C}_{\ell d}^{[iijj]}$	-	-	-	-	-	-	-	-
$\mathcal{C}_{qe}^{[3333]}$	-	0	-	-	-	-	-	-
$\mathcal{C}_{qe}^{[33ii]}$	0.6	6	-	-	-	-	-	-
$\mathcal{C}_{qe}^{[ii33]}$	-	0	-	-	-	-	-	-
$\mathcal{C}_{qe}^{[iijj]}$	-	-	-	-	-	-	-	-
$\mathcal{C}_{qu}^{(1)[3333]}$	0.3	1	-	-	-	-	-	-
$\mathcal{C}_{qu}^{(1)[ii33]}$	0.3	1	-	-	-	-	-	-
$\mathcal{C}_{qu}^{(1)[33ii]}$	-	0	-	-	-	-	-	-
$\mathcal{C}_{qu}^{(1)[iijj]}$	-	0	-	-	-	-	-	-
$\mathcal{C}_{qu}^{(8)[3333]}$	0.2	0	-	-	-	-	-	-
$\mathcal{C}_{qu}^{(8)[ii33]}$	0.3	0	-	-	-	-	-	-
$\mathcal{C}_{qu}^{(8)[33ii]}$	-	0	-	-	-	-	-	-
$\mathcal{C}_{qu}^{(8)[iijj]}$	-	0	-	-	-	-	-	-
$\mathcal{C}_{qd}^{(1)[3333]}$	0.2	0	-	-	-	-	-	-
$\mathcal{C}_{qd}^{(1)[ii33]}$	-	0	-	-	-	-	-	-
$\mathcal{C}_{qd}^{(1)[33ii]}$	-	0	-	-	-	-	-	-
$\mathcal{C}_{qd}^{(1)[iijj]}$	-	0	-	-	-	-	-	-
$\mathcal{C}_{qd}^{(8)[3333]}$	-	-	-	-	-	-	-	-
$\mathcal{C}_{qd}^{(8)[ii33]}$	0.1	-	-	-	0.1	$B \rightarrow X_s \gamma$	-	$B \rightarrow X_s \gamma$
$\mathcal{C}_{qd}^{(8)[33ii]}$	-	-	0.1	0.7	0.7	FourQuarksTop	0.7	FourQuarksTop
$\mathcal{C}_{qd}^{(8)[iijj]}$	-	-	-	-	-	R_τ	-	$ C_{B_s} $

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	Λ_{EW}	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.
\mathcal{C}_H	-	-	-	-	-	-
$\mathcal{C}_{H\Box}$	0.2	0.2	0.6	0.1	0.6	A_b^{FB}
\mathcal{C}_{HD}	0.5	0.5	5.1	-	5.	A_b^{FB}
\mathcal{C}_{HG}	0.8	0.8	0.4	-	0.9	$B \rightarrow X_s \gamma$
\mathcal{C}_{HB}	0.5	0.5	0.9	-	0.9	A_b^{FB}
\mathcal{C}_{HW}	0.7	0.7	0.9	-	1.	A_b^{FB}
\mathcal{C}_{HWB}	1.	1.	9.	-	9.	A_b^{FB}
\mathcal{C}_G	1.1	1.1	0.1	-	1.1	$B \rightarrow X_s \gamma$
\mathcal{C}_W	0.3	0.3	0.9	-	0.9	A_b^{FB}

Bosonic operators

coeff.	$\Lambda_{\text{flav.}}^{\text{down}}$	$\Lambda_{\text{flav.}}^{\text{up}}$	Λ_{EW}	$\Lambda_{\text{coll.}}$	$\Lambda_{\text{all}}^{\text{down}}$	Obs.	$\Lambda_{\text{all}}^{\text{up}}$	Obs.
\mathcal{C}_H	-	-	-	-	-	-	-	-
$\mathcal{C}_{H\Box}$	0.2	0.2	0.6	0.1	0.6	A_b^{FB}	0.6	A_b^{FB}
\mathcal{C}_{HD}	0.5	0.5	5.1	-	5.	A_b^{FB}	5.	A_b^{FB}
\mathcal{C}_{HG}	0.8	0.8	0.4	-	0.9	$B \rightarrow X_s \gamma$	0.9	$B \rightarrow X_s \gamma$
\mathcal{C}_{HB}	0.5	0.5	0.9	-	0.9	A_b^{FB}	0.9	A_b^{FB}
\mathcal{C}_{HW}	0.7	0.7	0.9	-	1.	A_b^{FB}	1.	A_b^{FB}
\mathcal{C}_{HWB}	1.	1.	9.	-	9.	A_b^{FB}	9.	A_b^{FB}
\mathcal{C}_G	1.1	1.1	0.1	-	1.1	$B \rightarrow X_s \gamma$	1.1	$B \rightarrow X_s \gamma$
\mathcal{C}_W	0.3	0.3	0.9	-	0.9	A_b^{FB}	0.9	A_b^{FB}