# Can relativistic effects explain galactic dynamics without dark matter?

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### Missing mass problem

#### Galactic flat rotation curves



#### Gravitational lensing



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- these effects cannot be accounted for based only on the visible baryonic matter
- ▶ for stars in galaxy,  $v \leq 10^{-3} \Rightarrow \frac{d^2 \vec{x}}{dt^2} = \vec{G}_N + O(|\vec{G}_N| 10^{-6})$ ⇒ relativistic corrections one million times smaller than needed to impact rotation curves
- Can full general relativity explain them, without dark matter?

#### Quasi-Maxwell formalism

Stationary spacetime: 
$$ds^2 = -e^{2\Phi}(dt - A_i dx^i)^2 + h_{ij} dx^i dx^j$$

Space part of time-like geodesic equation:

$$\frac{\tilde{D}\vec{U}}{d\tau} = \gamma \left[ \gamma \vec{G} + \vec{U} \times \vec{H} \right] \qquad \stackrel{(\tilde{L})}{\overset{(\tilde{L})}{det}}$$

 $( ilde{D}/d au \equiv ext{covariant})$  derivative wrt to  $h_{ij}$ 

Space part of null geodesic equation:

$$\frac{\tilde{D}\vec{k}}{d\lambda} = \nu \left[ \nu \vec{G} + \vec{U} \times \vec{k} \right]$$

- analogous to Lorentz force  $D\vec{U}/d\tau = (q/m)[\gamma\vec{E} + \vec{U} \times \vec{B}]$
- $G_i = -\Phi_{,i} \equiv$  "gravitoelectric" field
- $\blacktriangleright H^{i} = e^{\Phi} \epsilon^{ijk} \mathcal{A}_{k,j} \equiv \text{``gravitomagnetic'' field}$
- $h_{ij} \equiv$  space (or radar) metric



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  - $\blacktriangleright$   $ec{G} = ec{G}_{
    m N} +$  non-linear terms
  - ▶ if GR was to explain the missing mass problem, would have to be either through  $\vec{H}$ , or the non-linear terms in  $\vec{G}$



#### Gravitational lensing



- Nearly spherical lens: when the light source, lens, and observer are aligned, an Einstein ring forms in the observer's sky.
- Nearly perfect Einstein rings have been detected (e.g. "Cosmic Horseshoe", B1938+666);
- impossible to explain based only on the visible baryonic matter.
- Consistent with dark matter halos roughly spherical or moderately deformed



- The gravitomagnetic field  $\vec{H}$  cannot mimic dark matter
  - ▶ in the equatorial plane, GM "force"  $\vec{v} \times \vec{H}$  deflects rays on both sides of the body in the same direction;
  - creates no convergence of rays along axis connecting source and lens

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Gauss-Bonnet theorem applied to 2-surface S on the space manifold (of metric  $h_{ij}$ ), bounded by  $C_+$  and  $C_-$ :

$$\theta_{\rm R} = \iint_{\mathcal{S}} K d\mathcal{S} + \int_{C_+} \kappa_{\rm g} d\lambda - \int_{C_-} \kappa_{\rm g} d\lambda - \theta_{\rm S}$$

 $\blacktriangleright \ \kappa_{\rm g} = G^2 + (\vec{v} \times \vec{H})^2 \Rightarrow \text{gravitomagnetic contributions to } \theta_{\rm R} \text{ cancel out}$ 



- Kerr: rays starting at equal (in magnitude) angles will not cross along the lens-source axis (x-axis)
- those that do cross along the axis, arrive at different angles.



Dipole-like  $\vec{H}$ :

- rays with impact parameter  $ec{b}$  orthogonal to the equatorial plane are deflected orthogonally to  $ec{b}$ 
  - creates no convergence
  - deflection direction the same for  $\pm \vec{b}$ ; but opposite to equatorial plane



(Images generated with the GYOTO ray tracing code) Kerr:

for aligned setting, ring is weakened at the poles or splits into pair of arcs.

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(Images generated with the GYOTO ray tracing code) Kerr:

- for aligned setting, ring is weakened at the poles or splits into pair of arcs.
- lmage shifted orthogonally to  $\vec{S}$ .
- for source at the primary caustic (off the optical axis), covering the whole caustic section: nearly perfect, shifted ring forms
- similar (*for source wider than caustic*) to non-aligned Schwarzschild lens

• same angular diameter  $\Rightarrow \vec{H}$  does not contribute to lens power

smaller sources: rings to do not form anywhere



(Images generated with the GYOTO ray tracing code)

For  $S/M^2 > 1$  (possible only for extended bodies, like stars) the ring's deformation is unavoidable

in general, the ring does not even form

▶ still H < G (typically  $H \ll G$ ) along the ray trajectory:  $H/G \sim v_{\rm rot} R/r < 1$ .



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- ▶ still H < G (typically  $H \ll G$ ) along the ray trajectory:  $H/G \sim v_{\rm rot} R/r < 1.$
- But a much larger  $\vec{H}$  would be needed in order to have an impact on galactic rotation.

Space part of time-like geodesic equation:

$$rac{ ilde{D}ec{U}}{d au} = \gamma^2 \left[ec{G} + ec{v} imes ec{H}
ight] \qquad v \lesssim 10^{-3} ext{ for stars in galaxy}$$

Space part of null geodesic equation:

$$\frac{\tilde{D}\vec{k}}{d\lambda} = \nu^2 \left[\vec{G} + \vec{v} \times \vec{H}\right] \qquad v = 1 \text{ for light}$$

▶ In order for gravitomagnetic force  $\vec{v} \times \vec{H}$  to have impact on rotation curves, one needs  $|\vec{H}| \sim 10^3 |\vec{G}|$ 

 $\blacktriangleright$  impossible for rotating body  $H/G \sim v_{
m rot} R/r < 1$ 

$$\blacktriangleright \quad \vec{v}_{\rm f} - \vec{v}_{\rm in} \approx 2 \int_{-\infty}^{\infty} \vec{G} dt + \int_{-\infty}^{\infty} \vec{v} \times \vec{H} dt$$

 $\Rightarrow$  bending angles orders of magnitude larger than observed!

 $\blacktriangleright$   $\vec{H}$  cannot be the driver of galactic dynamics

#### Non-linear GR effects work against attraction

Geodesic equation for a star in a galaxy, constrained by observed lensing to be:

$$rac{ ilde{D}ec{U}}{d au}pproxec{G}~~v_{
m circ}=\sqrt{rG_r}+O(10^{-6})$$

Remains only to clarify whether non-linear effects can amplify  $\vec{G}$  in order to sustain the rotation curves without dark matter

Field equations for $\vec{G}$ and $\vec{H}$	
$\begin{split} \tilde{\nabla} \cdot \vec{G} &= -4\pi (2\rho + T^{\alpha}_{\ \alpha}) + \vec{G}^2 + \frac{1}{2}\vec{H}^2 \\ \tilde{\nabla} \times \vec{H} &= -16\pi \vec{J} + 2\vec{G} \times \vec{H} \end{split}$	• time-time and time-space projections of $R_{\mu\nu} = 8\pi \left(T_{\mu\nu} - \frac{1}{2}g_{\mu\nu}T^{\alpha}_{\ \alpha}\right)$
$egin{array}{ll} ar{ abla}  imes egin{array}{ll} ar{ abla} &  otimes eta &  otimes &  otimes eta &  otimes eta &  otim$	► Identities

Non-linear terms  $\vec{G}^2$  and  $\vec{H}^2/2$  act as effective negative "energy" sources for  $\vec{G}$ 

- counter the attractive effect of  $2
  ho+{\cal T}^lpha_{\ lpha}$ 
  - aggravate the missing mass problem

Non-linear GR effects work *against* attraction — Post-Newtonian approximation

static point mass

$$ec{G} = -rac{M}{r^3}\left(1-rac{2M}{r}
ight)ec{r} < -rac{M}{r^3}ec{r} \equiv ec{G}_{
m N}$$

angular velocity of circular orbit:

$$\Omega_{\rm circ} = \left[ \sqrt{\frac{M}{r^3}} - \frac{3}{2} \sqrt{\frac{M^3}{r^5}} \right] \ < \ \sqrt{\frac{M}{r^3}} \ \equiv \ \Omega_{\rm N}$$

 $\Rightarrow$  non-linear term *slows down* rotation

► self gravitating disks (Mach-Malec, 2015)  

$$\Omega_{\rm circ} = \Omega_{\rm N} \left[ 1 - \frac{2}{1-\delta} \Omega_{\rm N}^2 r^2 - \frac{4h_{\rm N}}{1-\delta} \right] - \frac{\mathcal{A}_{\phi}}{r^2(1-\delta)}$$

$$\delta \in [-\infty, 0] \setminus \{-1\}$$

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#### Balasin-Grumiller "galactic" model

$$ds^{2} = -(dt - \mathcal{A}_{\phi}(r, z)d\phi)^{2} + h_{ij}(r, z)dx^{i}dx^{j}$$
$$\mathcal{A}_{\phi}(r, z) = V_{0}(R - r_{0}) + \frac{V_{0}}{2}[d_{r_{0}} + d_{-r_{0}} - d_{R} - d_{-R}]$$
$$h_{ij}dx^{i}dx^{j} = r^{2}d\phi^{2} + e^{\nu(r, z)}(dr^{2} + dz^{2})$$

$$d_{R} \equiv \sqrt{r^{2} + (z - R)^{2}}$$
$$d_{-R} \equiv \sqrt{r^{2} + (z + R)^{2}}$$
$$d_{r_{0}} \equiv \sqrt{r^{2} + (z - r_{0})^{2}}$$
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 $r_0 \equiv radius \text{ of galactic bulge}; \quad R \equiv radius \text{ of galactic disk}$  $V_0 = \text{const.} \equiv \text{dust velocity, wrt ZAMOS, in the "flat regime"}$ 

Claimed to describe, in *comoving coordinates*, a rotating dust with a flat velocity profile matching the Milky Way's. But:

- ►  $g_{\alpha\beta}$  time-independent  $\Rightarrow$  dust at rest in rigid frame  $\Rightarrow$  incompatible with flat rotation curve (demands non-constant  $\vec{\Omega}$ )
- $\vec{G} = 0$ ,  $\lim_{r \to \infty} \vec{H} = 0 \Rightarrow$  asymptotically inertial rigid frame
  - dust static with respect the asymptotic inertial frame (Costa et al, 2023)

 $\Rightarrow$  non-rotating with respect to the distant quasars

Cannot describe any galaxy.

Balasin-Grumiller "galactic" model — non-linearity

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► 
$$\vec{G} = 0$$
 and  $\vec{J} = 0$  (comoving coordinates):  
 $\tilde{\nabla} \cdot \vec{G} = -4\pi\rho + \frac{1}{2}\vec{H}^2 = 0$ 

 $\blacktriangleright$  Linearizing yields empty space equation ho= 0

 purely non-linear solution (no linear, or Newtonian limit)

• extreme repulsive action of  $\vec{H}^2/2$  cancels out exactly the attractive effect of the dust's energy density  $\rho$  ("freezes" the dust!)

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- extreme repulsive action of  $\vec{H}^2/2$  cancels out exactly the attractive effect of the dust's energy density  $\rho$  ("freezes" the dust!)
- $\blacktriangleright$   $\vec{H}$  generated by singularities along the axis, not by motion of matter.

# BG "galactic" model — gravitomagnetic field $\vec{H}$



This is the gravitomagnetic field of a pair of oppositely charged NUT rods along z-axis, of gravitomagnetic charges

$$Q_{\text{NUT}} = \frac{1}{4\pi} \int_{\mathcal{S}} d\boldsymbol{\mathcal{A}} = \frac{1}{4\pi} \int_{\mathcal{S}} \vec{H} \cdot \vec{dS} = \mp V_0 (R - r_0)/2$$

► matches the magnetic field  $\vec{B}_{rods}$  of a pair of magnetically charged rods, identifying  $V_0/2$  with charge density  $\lambda_M$ :  $(B_{rods})_i \stackrel{\lambda_M \to V_0/2}{=} H_i$ . (length of the rods approximately equal to galactic diameter...)

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- Plus a curl-free term in  $\mathcal{A} \Rightarrow$  potential of an infinite spinning cosmic string, of angular momentum per unit mass  $j = -V_0(R_p r_0)/4$ .





(Images generated with the GYOTO ray tracing code)

- Rays do not cross along optical axis for aligned setting
- Multiple images at equator for y > 0, where light rays cross
- No Einstein rings





(Images generated with the GYOTO ray tracing code)

▶ Deflection angles much larger than observed (spherical lens with Milky Way's mass  $M = 10^{12} M_{\odot}$ , yields Einstein ring 18arcsec wide).

#### BG "galactic" model — origin of claimed rotation curves

BG velocity is measured wrt zero angular momentum observers (ZAMOs) (azimuthal angular momentum:  $u_{\phi}$ )

- **ZAMOs**:  $(u_Z)_{\phi} = 0$
- have angular velocity  $\Omega_{Z} \equiv \frac{u_{Z}^{\phi}}{u_{Z}^{0}} = -\frac{g_{0i}}{g_{00}} = \frac{e^{2\Phi} \mathcal{A}_{\phi}}{g_{\phi\phi}}$ relative to asympt. inertial frame
- $\blacktriangleright$  are *dragged* by  ${\cal A}$

Kerr spacetime

 at the horizon, ZAMO angular velocity coincides with that of the horizon (ZAMO comoves with the horizon)

$$\Omega_{\mathrm{Z}}(r_{+}) = rac{a}{r_{+}^{2}+a^{2}} = \Omega_{\mathrm{H}}$$

by confusing the ZAMOs with observers at rest relative to distant stars, one would conclude that Kerr black holes do not rotate!

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- $\blacktriangleright$  are "dragged" by  $oldsymbol{\mathcal{A}}$



- $\blacktriangleright$  artificially large gm potential  ${\cal A}$  created by the singularities
- ZAMOs misunderstood as at rest relative to the axis' asymptotic rest frame
- ► the velocity curve obtained: $v_{\rm rZ}^{\phi} = -\sqrt{-g^{00}}\Omega_Z$ is but minus the velocity of the ZAMOs with respect to the rigid asymptotic inertial frame
- Is the ZAMOs, not the dust (static in such frame), that rotates

#### Conclusions

We have demonstrated that, in light of the experimentally measured galactic rotation curves and gravitational lensing, relativistic effects cannot resolve (or even be relevant) to the missing mass problem

- gravitational lensing rules out the gravitomagnetic field as a player;
- non-linear effects only aggravate the need for dark matter (besides negligible in realistic models)
- general relativistic "galactic" models in the literature originate from pathologies:
  - unphysical singularities, generating artificially large gravitomagnetic fields (ruled out by the observed gravitational lensing);
  - in "exact" models, rotation curves moreover computed relative to unsuitable reference observers — the ZAMOs, being dragged by the singularities

Thank you! 🔊 🖉

(e.g. BG model is actually static, does not even rotate!)

References:

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Costa-Natário arXiv:2312.12302 (PRD to appear),
Costa et al PRD 108 (2023) 4, 044056 [arXiv:2303.17516]
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