Generalized Siklos space-times Ángel Jesús Murcia Gil, INFN Padova (Italy)



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Based on arXiv:2407:01359 [gr-qc] with B. Araneda.

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Do there exist **exact solutions** of **GR** which admit a **gravitational-wave interpretation**?

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(x,y) ℓ^a

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Physical interpretation: **plane-fronted waves** with **parallel rays** propagating on **Minkowski** [*e.g.* Ehlers, Kundt '62; Roche, Aazami, Cederbaum '22].

If our space-time (M,g) is spin, **pp-waves** are **equivalent** to existence of **parallel spinor**! [*e.g.* Tod '83; Bryant '00; <u>ÁM</u>, Shahbazi '21; Araneda '22]

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Parallel spinor: Geometric technique to solve Einstein equations (2nd order) through 1st order equation!

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Inspired by **supersymmetry** conditions in **supergravity**, Siklos analyzed spacetimes (M,g) endowed with the following type of spinors [Siklos '85]:

$$abla_m \psi = rac{b}{\sqrt{2}} \gamma_m \psi, \quad b \in \mathbb{C}.$$

We call (M,g) a **Siklos space-time**. ψ is said to be a **Killing spinor** with Killing constant b.

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If (M,g) not AdS, $b \in \mathbb{R}$ and the Killing spinor $\psi = (o_A, \tilde{o}_{A'})^T$ is Majorana [Araneda, <u>ÁM</u> '24]. Now $\ell^a = o^A \tilde{o}^{A'}$ is no longer parallel. It is Killing:

 $\nabla_{(a}\ell_{b)} = 0.$

Siklos space-times are locally isometric to [Siklos '85]:

$$ds_{Siklos}^{2} = \frac{1}{b^{2}x^{2}} \left[H(v, x, y)dv^{2} + 2dudv - dx^{2} - dy^{2} \right]$$



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Einstein tensor of Siklos space-times:

$$G_{ab} = -b^4 x^4 \left(H_{xx} + H_{yy} - \frac{2}{x} H_x \right) \ell_a \ell_b + 6b^2 g_{ab} \,.$$

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Physical interpretation: Siklos space-times correspond to exact GWs propagating on top of AdS [Podolský '97], with propagation vector ℓ_a .

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A given **space-time** is said to be **generalized Siklos** if it possesses a spinor ψ satisfying [Araneda, <u>ÁM</u> '24]:

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- May be useful to construct solutions to Einstein equations with certain matter content.

If (M,g) is a generalized Siklos space-time:

• If (M,g) not conformally flat, Killing function λ is **real** and $\psi = (o_A, \tilde{o}_{A'})$ may be taken to be **Majorana**.



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(M, g**) conformal** to **pp-wave**:

$$ds_{\text{gen-Siklos}}^2 = \frac{1}{\Omega^2(x,v)} \left[2dudv - dx^2 - dy^2 + Hdv^2 \right],$$
$$\Omega(x,v) = \int \lambda(x,v) \, dx + f(v), \quad f \in C^\infty(\mathbb{R}).$$

Weyl tensor is **type N**, so **gravitational dofs** are interpreted as **waves**: u is **affine parameter** along the rays given by $\ell^a = (\partial_u)^a$ and **wave surfaces** are spanned by coordinates (x, y).

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$$\begin{split} G_{ab} &= -(T_{ab}^{(\ell)} + T_{ab}^{(s)}) \,, \\ T_{ab}^{(\ell)} &= \rho_{(\ell)} \ell_a \ell_b \qquad T_{ab}^{(s)} = (\rho_{(s)} + p_{(s)}) s_a s_b - p_{(s)} g_{ab}, \end{split}$$

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with $s_a s^a = -2\Omega_{xx}^2 < 0$ and

$$\rho_{(\ell)} = 2\Omega^3 \left[\frac{\Omega}{2} (\hat{H}_{xx} + \hat{H}_{yy}) + \Omega_{vv} - \Omega_x \hat{H}_x - \frac{\Omega_{vx}^2}{\Omega_{xx}} \right],$$
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Physically relevant case: $\Omega_{xx} = 0$, which is equivalent to standard Siklos space-times.

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Muito obrigado!