

Schwinger pair production in de Sitter Regularizing negative conductivities

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Ongoing work with
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Soon on 2409.XXXXX



Spontaneous **particle creation** by **time-varying backgrounds**

- **Schwinger Effect**

W. Heisenberg, H. Euler (1936); J. Schwinger (1951)

- Production of charged particles from vacuum under **strong electric fields**
- Time dependent vector potential

- **Curved backgrounds**

L. Parker (1966); S. W. Hawking (1975)

- Particle production from vacuum under **time dependent gravitational field**

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Combining the two examples: **Schwinger effect in de Sitter**

Inflationary **Magnetogenesis**

- Generate the **observed magnetic fields** present in voids our universe

Generation of Dark Sectors

- Candidates for non-thermal **dark matter**

Scalar QED in de Sitter

$$S = \int d^4x \sqrt{-g} \left\{ -g^{\mu\nu} (\partial_\mu - ieA_\mu) \phi^* (\partial_\nu + ieA_\nu) \phi - (m_\phi^2 + \xi R) \phi^* \phi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} \right\},$$

- Set a constant a electric field

$$A_\mu = \frac{E}{H^2 \tau} \delta_\mu^z, \quad F_{\mu\nu} F^{\mu\nu} = -2E^2$$

- After canonically normalizing the scalar field ϕ e.o.m. for $q \equiv a\phi$

$$q_k'' + \omega_k^2 q_k = 0,$$

- Analytical solution with Whittaker functions

$$q_k = \frac{e^{-\pi\lambda r/2}}{\sqrt{2k}} W_{i\lambda r, \mu}(2ik\tau)$$

Scalar QED in de Sitter

- A_ν e.o.m.

$$\nabla^\nu F_{\mu\nu} = J_\mu^\phi \quad \text{with} \quad J_\mu^\phi = \frac{ie}{2} \left\{ \phi^\dagger (\partial_\mu + ieA_\mu) \phi - \phi (\partial_\mu - ieA_\mu) \phi^\dagger \right\} + \text{h.c.} .$$

- With an electric field in the z-direction,

$$\langle 0 | J_z^\phi | 0 \rangle = \frac{2e}{a^2} \int \frac{d^3k}{(2\pi)^3} (k_z + eA_z) |q_k|^2 .$$

- **Divergent expectation value.** With a cut off momentum ζ

T. Kobayashi, N. Afshordi 2014

$$\langle J_z^\phi \rangle = aH \frac{e^2 E}{4\pi^2} \lim_{\zeta \rightarrow \infty} \left[\frac{2}{3} \left(\frac{\zeta}{aH} \right)^2 + \frac{1}{3} \ln \frac{2\zeta}{aH} - \frac{25}{36} + \frac{\mu^2}{3} + \frac{\lambda^2}{15} + F_\phi(\lambda, \mu, r) \right] .$$

Dimensionless quantities $\lambda = \frac{eE}{H^2}$, $r = \frac{k_z}{k}$, $\mu^2 = \frac{9}{4} - \frac{m^2}{H^2} - \lambda^2$ and $m_\xi^2 = m_\phi^2 + 12\xi H^2$.

State of the art on renormalization

- For **Scalars**: 3 different regularization/renormalization procedures

- Adiabatic Subtraction (AS)
- Point Splitting (PS)
- Pauli Villars (PV)

T. Kobayashi, N. Afshordi 2014

T. Hayashinaka, J. Yokoyama 2016

M. Banyeres, G. Domenèch, J. Garriga 2018

- All **agree** for $m > H$

- When $m < H$:

- AS and PS the result leads to **negative conductivities** coming from $\log(m/H)$ term
- In PV authors argue $\log(m/H)$ should be **reabsorbed in the running of electric charge**

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- **Fermions**

- Only AS
- Again, when $m < H$ **negative conductivities** coming from $\log(m/H)$ term

T. Hayashinaka, T. Fujita, J. Yokoyama 2016

Renormalized currents with PV

- An arbitrary number of additional **auxiliary fields are introduced to cancel divergences**

The **regularized** current $\langle J_z \rangle_{\text{reg}} = \lim_{\Lambda \rightarrow \infty} \sum_{i=0}^3 (-1)^i \langle J_z \rangle_i = aH \frac{e^2 E}{4\pi^2} \lim_{\Lambda \rightarrow \infty} \left[\frac{1}{6} \ln \frac{\Lambda^2}{H^2} - \frac{2\lambda^2}{15} + F_\phi(\lambda, \mu, r) \right].$

- In Λ/H **divergence** to be reabsorbed with **renormalization** of the charge

$$\begin{aligned}\langle J_\mu \rangle_{\text{reg}} &= (\delta_3 + 1) \nabla^\nu F_{\mu\nu} \\ \langle J_\mu \rangle_{\text{ren}} &= \nabla^\nu F_{\mu\nu} = \langle J_\mu \rangle_{\text{reg}} - (-2aHE\delta_\nu^z)\delta_3\end{aligned}$$

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- In Banyeres et al $\delta_3 = -\frac{e^2}{48\pi^2} \ln \frac{\Lambda^2}{m^2}$ as $\mathbf{p}^2 = \mathbf{0}$ in the **vacuum polarization** diagram

$$\langle J_z^\phi \rangle_{\text{ren}} = aH \frac{e^2 E}{4\pi^2} \left[\frac{1}{6} \ln \frac{m^2}{H^2} - \frac{2\lambda^2}{15} + F_\phi(\lambda, \mu, r) \right]$$

- It is argued that when problematic $\ln m/H$ is reabsorbed in **running** of \mathbf{e} to scale H

Renormalized currents with PV

- However, taking $\mathbf{p}^2 = -2\mathbf{H}^2$, as required to ensure a **constant electric field** in de Sitter in e.o.m.

$$\delta_3 = \left(\frac{e}{12\pi}\right)^2 \left(3 \ln \left(\frac{m^2}{\Lambda^2} \right) - 12 \left(\frac{m}{H} \right)^2 + 6 \left(2 \left(\frac{m}{H} \right)^2 + 1 \right)^{3/2} \coth^{-1} \left(\sqrt{2 \left(\frac{m}{H} \right)^2 + 1} \right) - 8 \right)$$

- And we find the **renormalized** current to be

$$\langle J_z^\phi \rangle_{\text{ren}}^{PV} = aH \frac{e^2 E}{4\pi^2} \left[\frac{1}{3} \ln \frac{m}{H} - \frac{4}{9} - \frac{2}{3} \left(\frac{m}{H} \right)^2 - \frac{2\lambda^2}{15} + \frac{\left(1 + 2 \left(\frac{m}{H} \right)^2 \right)^{3/2}}{3} \coth^{-1} \left(\sqrt{2 \left(\frac{m}{H} \right)^2 + 1} \right) + F_\phi \right]$$

Renormalizing currents with AS

- The subtraction is done **mode by mode** removing the expectation evaluated in the adiabatic approx

WKB expansion
$$q_{\mathbf{k}}(\tau) = \frac{1}{\sqrt{2W_{\mathbf{k}}(\tau)}} \exp \left\{ -i \int^{\tau} d\tilde{\tau} W_{\mathbf{k}}(\tilde{\tau}) \right\}$$

- Running / Physical Scale AS** with an **arbitrary adiabatic expansion scale** \bar{m}

A. Ferreira, S. Monin, J. Navarro Salas, F. Torrenti 2018, 2022, 2023

- Value of \bar{m} has to be set to obtain the **appropriate adiabatic vacuum** evolution

$$\langle \mathbf{J}_z^\phi \rangle_{\text{reg}}^{\text{AS}} = \langle \mathbf{J}_z^\phi \rangle - \langle \mathbf{J}_z^\phi \rangle^{(2)} = aH \frac{e^2 E}{4\pi^2} \left[\frac{1}{3} \ln \frac{\bar{m}}{H} - \frac{2\lambda^2}{15} + F_\phi(\lambda, \mu, r) \right]$$

- If $m > H$ $\bar{m} = m$;
- If $m < H$ $\bar{m} = H$

Renormalizing currents with DR

- Applying DR, in the Whitaker function we have a **scaleless argument** and integral gives zero
- **Expanding** the argument for a **large energy-like quantity**,

$$e_k = \sqrt{k^2 + a^2 x^2}$$

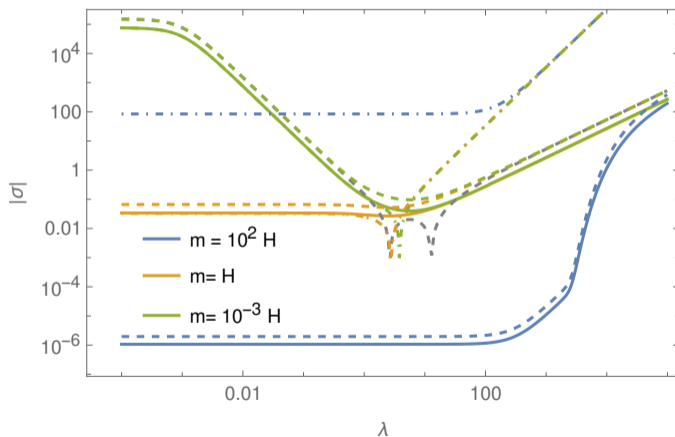
Isolates the divergent pieces and **introduce an artificial IR regulator**.

A. V. Lysenko, O. O. Sobol, E. V. Gorbar, A. I. Momot, and S. I. Vilchinskii 2020, 2023

$$\begin{aligned} \langle J_Z^\phi \rangle_{\text{ren}}^{DR} &= \langle J_Z^\phi \rangle_{\text{reg}}^{DR} - (-2aHE\delta_\nu^z)\delta_3 \quad \text{with} \quad \langle J_Z^\phi \rangle_{\text{reg}}^{DR} = \langle J_Z^\phi \rangle - \langle J_Z^\phi \rangle^{e_k} + \langle J_Z^\phi \rangle_{\text{reg}}^{e_k} \\ &= aH \frac{e^2 E}{4\pi^2} \left[\frac{1}{3} \ln \frac{2m}{H} - \frac{7}{18} - \left(\frac{m}{H}\right)^2 - \frac{4\lambda^2}{15} + \frac{\left(1 + 2\left(\frac{m}{H}\right)^2\right)^{3/2}}{3} \coth^{-1} \left(\sqrt{2\left(\frac{m}{H}\right)^2 + 1} \right) + F_\phi \right] \end{aligned}$$

Renormalized Currents and conductivities

- Successfully **removed the infrared divergences** ($\ln m/H$) that lead to negative conductivities



$$\sigma_z \equiv \frac{1}{aH} \frac{\langle J_z \rangle}{e^2 H}$$

DR **disagrees** when:

- $\lambda \gg 1$
- $m \gg H$

Non physical results

Work in progress (soon on 24XX.XXXXX)

Conclusion & Outlook

- We have revised PV, AS and DR renormalization in the literature
- We were able to address and clarify literature's negative conductivities in $H > m$ case
 - **Unphysical** result comes from **wrong physical** conditions
- With both PV and AS we have always recovered physically sensible results
 - Currents show **small deviations**
 - In PV we seem to have a **better knowledge on the physical system.**
 - With the the physical scale AS criteria to determine the scale \bar{m} seems more unsatisfactory.

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Next steps

- Apply this into generation of Dark Sectors during inflation
- Check Gravitational wave spectrum in Dark matter compatible scenarios
- Extend analysis to fermions

Backup

- An arbitrary number of additional **auxiliary fields are introduced to cancel divergences**
- The mass of these extra fields will then be sent to infinity, making them **non-dynamical**

Introduce **3** fields $\sum_{i=0}^3 (-1)^i = 0$ and $\sum_{i=0}^3 (-1)^i m_i^2 = 0,$

$$m_0 = m, \quad m_2^2 = 4\Lambda^2 - m^2 \quad \text{and} \quad m_1^2 = m_3^2 = 2\Lambda^2, \quad \Lambda \rightarrow \infty$$

The **regularized** current $\langle J_z \rangle_{\text{reg}} = \lim_{\Lambda \rightarrow \infty} \sum_{i=0}^3 (-1)^i \langle J_z \rangle_i.$

$$\langle J_z^\phi \rangle_{\text{reg}} = aH \frac{e^2 E}{4\pi^2} \lim_{\Lambda \rightarrow \infty} \left[\frac{1}{6} \ln \frac{\Lambda^2}{H^2} - \frac{2\lambda^2}{15} + F_\phi(\lambda, \mu, r) \right]$$

- In Λ/H **divergence** to be reabsorbed with **renormalization** of the charge

$$\begin{aligned} (\delta_3 + 1) \nabla^\nu F_{\mu\nu} &= \langle J_\mu \rangle_{\text{reg}} \\ \langle J_\mu \rangle_{\text{ren}} = \nabla^\nu F_{\mu\nu} &= \langle J_\mu \rangle_{\text{reg}} - (-2aHE\delta_\nu^z)\delta_3 \end{aligned}$$

Constant electric field in de Sitter

$$S = - \int d^4x \sqrt{-g} \frac{1}{4} F^{\mu\nu} F_{\mu\nu}.$$

From Euler Lagrange equations we would expect

$$g^{\alpha\nu} g^{\beta\sigma} \partial_\alpha F_{\nu\sigma} = 0.$$

However, including the details on the physical setting of our problem, a de Sitter metric and $A_\mu = \frac{E}{H^2\tau} \delta_\mu^z$, we find

$$\begin{aligned} g^{\alpha\nu} g^{\beta\sigma} \partial_\alpha F_{\nu\sigma} &= g^{00} g^{jj} \partial_0 F_{0j} + g^{jj} g^{00} \partial_j F_{j0} \\ &= (-a^{-2}) a^{-2} \partial_\tau \left(-\frac{E}{\tau^2 H^2} \delta_i^z \right) \\ &= -2a^{-4} \frac{E}{\tau^3 H^2} \delta_i^z \neq 0. \end{aligned}$$

We see that an abelian gauge theory with only a kinetic term is not consistent with a constant electric field in a de Sitter background.

Constant electric field in de Sitter

A possible solution is the inclusion of an effective mass term in the action

$$S = - \int d^4x \sqrt{-g} \left(\frac{1}{4} F^{\mu\nu} F_{\mu\nu} + \frac{1}{2} m_A^2 A_\mu A^\mu \right).$$

Then, the Euler Lagrange equations give

$$\begin{aligned} g^{\alpha\nu} g^{\beta\sigma} \partial_\alpha F_{\nu\sigma} - m_A^2 g^{\beta\mu} A_\mu &= 0 \\ -a^{-4} 2 \frac{E}{\tau^3 H^2} \delta_i^z - m_A^2 a^{-2} \frac{E}{\tau H^2} \delta_i^z &= 0, \end{aligned}$$

and the system becomes consistent for $m_A^2 = -2H^2$.

In order to have a consistent electric field in de Sitter, the gauge boson must have something like an effective tachyonic mass, or a source term, breaking the conformal invariance.

Revising AS

- In a **time-dependent background** the **vacuum** of the theory is generally **evolving** making the concept of “vacuum contribution” **ambiguous**
- The subtraction is done **mode by mode** removing the expectation evaluated in the adiabatic approx

WKB expansion $q_{\mathbf{k}}(\tau) = \frac{1}{\sqrt{2W_{\mathbf{k}}(\tau)}} \exp \left\{ -i \int^{\tau} d\tilde{\tau} W_{\mathbf{k}}(\tilde{\tau}) \right\}$

$$\langle J_z^{\phi} \rangle = -\frac{2e}{(2\pi)^3 a^2} \int d^3k (k_z + eA_z) \frac{1}{2W_{\mathbf{k}}}$$

Inserting the mode function q in the e.o.m.

$$W_{\mathbf{k}}^2 = \omega_{\mathbf{k}}^2 + \frac{3}{4} \left(\frac{W'_{\mathbf{k}}}{W_{\mathbf{k}}} \right)^2 - \frac{1}{2} \frac{W''_{\mathbf{k}}}{W_{\mathbf{k}}}$$

Expanded at the n^{th} order

$$W_{\mathbf{k}} = W_{\mathbf{k}}^{(0)} + W_{\mathbf{k}}^{(1)} + W_{\mathbf{k}}^{(2)} + \dots$$

Running / Physical Scale AS

- Take $\Omega_{\mathbf{k}}^{\bar{m}}$ with **arbitrary adiabatic expansion scale** \bar{m} (opposed to automatically set $\bar{m} = m$)
A. Ferreiro, S. Monin, J. Navarro Salas, F. Torrenti 2018, 2022, 2023

$$\Omega_{\mathbf{k}}^{\bar{m}^2} = (k_z + eA_z)^2 + k_x^2 + k_y^2 + a^2 \bar{m}^2 = \omega_{\mathbf{k}}^2 + a^2(\bar{m}^2 - m^2) + \frac{a''}{a}$$

And set $W_{\mathbf{k}}^2{}^{(0)} = \Omega_{\mathbf{k}}^{\bar{m}^2}$

Find second order $W_{\mathbf{k}}^2$ with e.o.m. $W_{\mathbf{k}}^2{}^{(2)} = \Omega_{\mathbf{k}}^{\bar{m}^2} - a^2(\bar{m}^2 - m^2) - \frac{a''}{a} + \frac{3}{4} \left(\frac{\Omega_{\mathbf{k}}^{\bar{m}'}}{\Omega_{\mathbf{k}}^{\bar{m}}} \right)^2 - \frac{1}{2} \frac{\Omega_{\mathbf{k}}^{\bar{m}''}}{\Omega_{\mathbf{k}}^{\bar{m}}}$

$$\langle J_z^\phi \rangle^{(2)} = \lim_{\zeta \rightarrow \infty} \frac{eaH^3}{(2\pi)^2} \left[\frac{2\lambda}{3} \left(\frac{\zeta}{aH} \right)^2 - \frac{2\lambda^3}{15} - \frac{\lambda}{3} \left(\frac{m}{H} \right)^2 + \frac{\lambda}{3} \ln \left(\frac{2\zeta}{a\bar{m}} \right) + \frac{\lambda}{18} \right]$$

- And the **renormalized** current is given by

$$\langle J_z^\phi \rangle_{\text{ren}}^{\text{AS}} = \langle J_z^\phi \rangle - \langle J_z^\phi \rangle^{(2)} = aH \frac{e^2 E}{4\pi^2} \left[\frac{1}{3} \ln \frac{\bar{m}}{H} - \frac{2\lambda^2}{15} + F_\phi(\lambda, \mu, r) \right] \quad (\text{Similar to Banyeres et al})$$

- Value of \bar{m} has to be set to obtain the **appropriate adiabatic vacuum** evolution