

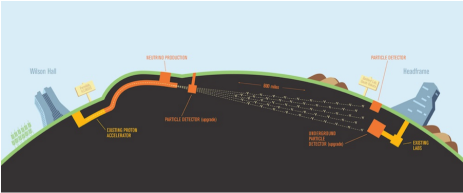
Neutrino Physics: Now and at Future Accelerators

Vedran Brdar

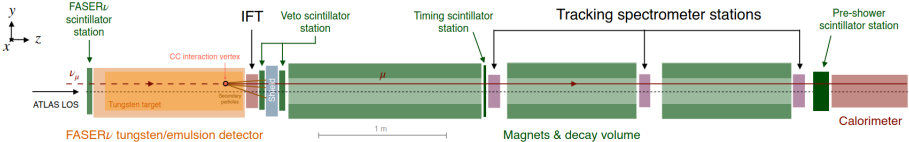


Neutrinos & Accelerators

▶ Accelerator Neutrinos (MiniBooNE, NOvA, DUNE...)



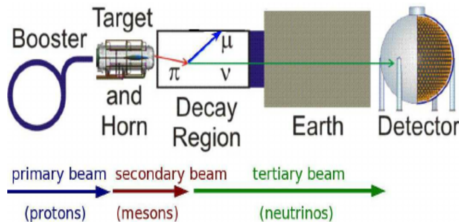
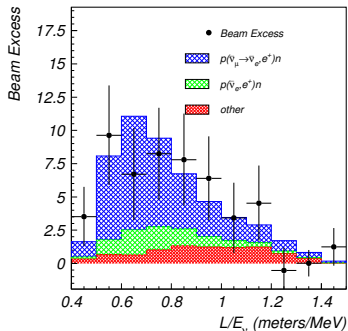
- ▶ LHC Neutrinos
- ▶ Neutrino-philic BSM searches at Colliders



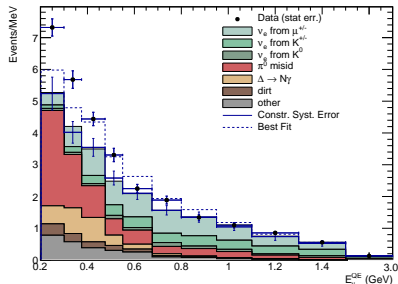
Outline

- ▶ MiniBooNE Anomaly
- ▶ New Physics Searches at DUNE
- ▶ Resonances in $\bar{\nu}_e - e^-$ scattering at FPF
- ▶ Neutrino Magnetic Moment and Future Colliders

Anomalies: LSND and MiniBooNE



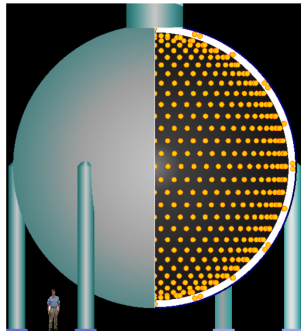
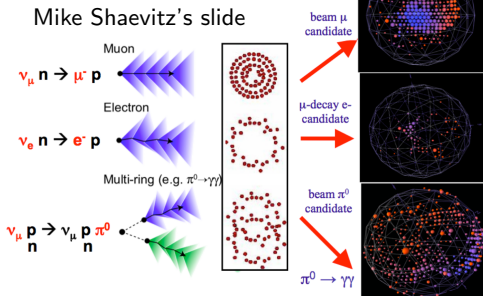
- ▶ **LSND**: $\bar{\nu}_e$ in $\bar{\nu}_\mu$ beam from stopped pion source ($> 3\sigma$) at $L/E \sim 1\text{km GeV}^{-1}$
- ▶ **MiniBooNE**: reports electron-like event excess (4.8σ)
- ▶ in combination with LSND 6.1σ



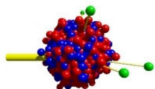
Main Channels for MiniBooNE

- ▶ shower events can be produced by e , γ , collimated e^+e^- and $\gamma\gamma$
- ▶ in SM, majority of events relevant for MiniBooNE excess arise from
 - 1) $CC \nu_e$ events,
 - 2) $NC \pi^0$ production,
 - 3) $NC \gamma$ from e.g. $\Delta(1232)$

Cerenkov rings provide the primary means of identifying the types of ν interactions in the detector



Employed MC Generators



GiBUU

The Giessen Boltzmann-Uehling-Uhlenbeck Project



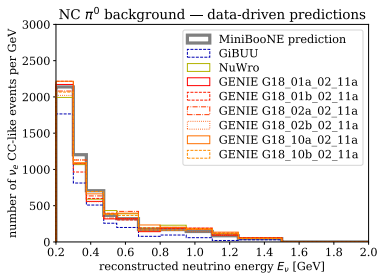
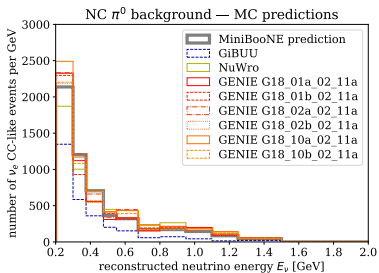
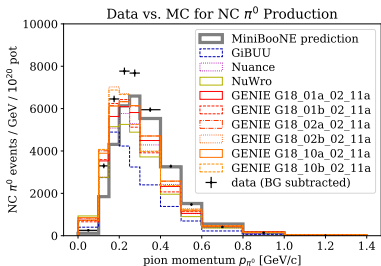
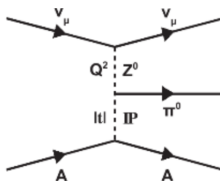
Generator	Tune	Ref.	Comments
NUANCE	-	[40]	the generator used by MiniBooNE
GiBUU	-	[42]	theory-driven generator
NuWro	-	[41]	
GENIE	G18_01a_02_11a	[39, 44]	GENIE baseline tune; see [44] for naming conventions
	G18_01b_02_11a		different FSI implementation compared to G18_01a_02_11a
	G18_02a_02_11a		updated res./coh. scattering models compared to G18_01a_02_11a
	G18_02b_02_11a		updated res./coh. scattering models and different FSI
	G18_10a_02_11a		theory-driven configuration; similar to G18_02a
	G18_10b_02_11a		theory-driven configuration; similar to G18_02b

Strategy

1. Make a prediction using **NUANCE**
2. Compare with the prediction obtained by the MiniBooNE collaboration; the differences are compensated by bin-by-bin tuning
3. Predict the event sample using **GiBUU**, **NuWro** and **GENIE** using the same cuts as for **NUANCE**; apply the tuning factors

Neutral Current π^0 Production

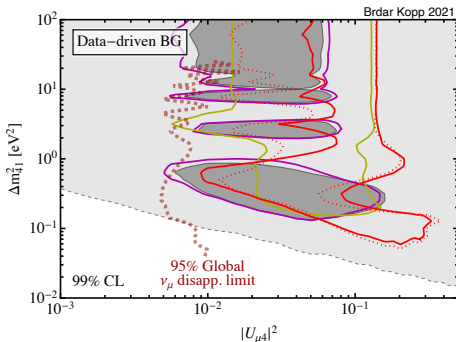
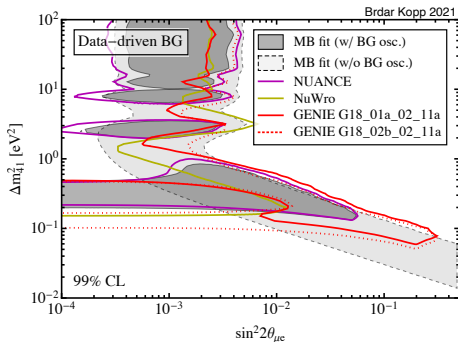
$$\nu + N \rightarrow \nu + N + \pi^0(\gamma\gamma)$$



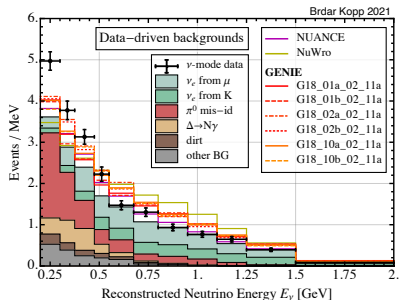
3+1 Model with eV-scale Sterile Neutrino

$$U^{4\text{flavor}} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} & U_{\mu4} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} & U_{\tau4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix}$$

$$\sin^2 2\theta_{\mu e} = 4|U_{e4}|^2|U_{\mu4}|^2$$



Results: Data-driven backgrounds



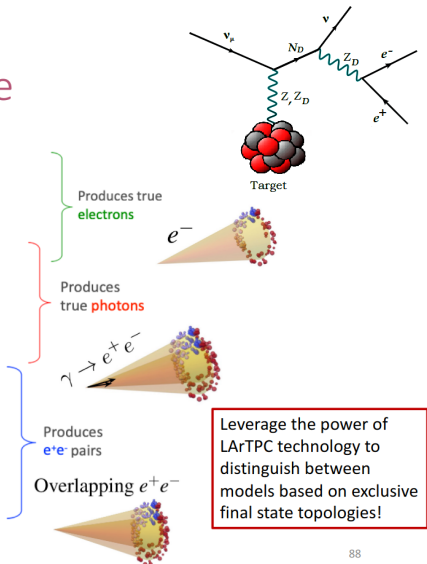
Generator	Tune	Δm_{41}^2 [eV ²]	$\sin^2 2\theta_{\mu e}$	Significance
MB official		0.25	0.01	4.0 σ
NUANCE	–	0.32	0.0079	4.0 σ
NuWro	–	3.2	0.0016	3.1 σ
GENIE	G18_01a_02_11a	0.79	0.00020	4.4 σ
	G18_01b_02_11a	0.79	0.0001	3.5 σ
	G18_02a_02_11a	0.13	0.063	4.0 σ
	G18_02b_02_11a	0.13	0.050	3.7 σ
	G18_10a_02_11a	0.25	0.016	3.5 σ
	G18_10b_02_11a	0.40	0.013	4.0 σ

Non-oscillatory Explanations of MiniBooNE Anomaly

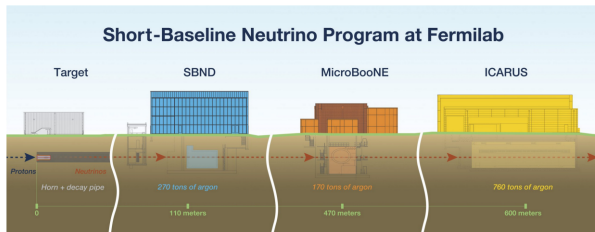
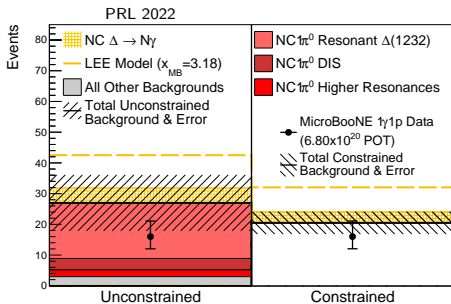
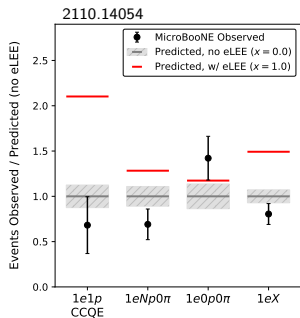
slide from MicroBooNE presentations

Evolving Theory Landscape

- Decay of O(keV) Sterile Neutrinos to active neutrinos
 - [13] Dentler, Esteban, Kopp, Machado Phys. Rev. D 101, 115013 (2020)
 - [14] de Gouvêa, Peres, Prakash, Stenico JHEP 07 (2020) 141
- New resonance matter effects
 - [5] Asaadi, Church, Guenette, Jones, Szeic, PRD 97, 075021 (2018)
- Mixed O(1eV) sterile oscillations and O(100 MeV) sterile decay
 - [7] Vergani, Kamp, Diaz, Arguelles, Conrad, Shaevitz, Uchida, arXiv:2105.06470
- Decay of heavy sterile neutrinos produced in beam
 - [4] Gninenko, Phys.Rev.D83:015015,2011
 - [12] Alvarez-Ruso, Saul-Sala, Phys. Rev. D 101, 075045 (2020)
 - [15] Magill, Plestid, Pospelov, Tsai Phys. Rev. D 98, 115015 (2018)
 - [11] Fischer, Hernandez-Cabezudo, Schwetz, PRD 101, 075045 (2020)
- Decay of upscattered heavy sterile neutrinos or new scalars mediated by Z' or more complex higgs sectors
 - [1] Bertuzzo, Jana, Machado, Zukanovich Funchal, PRL 121, 241801 (2018)
 - [2] Abdullahi, Hostert, Pascoli, Phys.Lett.B 820 (2021) 136531
 - [3] Ballett, Pascoli, Ross-Lonegan, PRD 99, 071701 (2019)
 - [10] Dutta, Ghosh, Li, PRD 102, 055017 (2020)
 - [6] Abdallah, Gandhi, Roy, Phys. Rev. D 104, 055028 (2021)
- Decay of axion-like particles
 - [8] Chang, Chen, Ho, Tseng, Phys. Rev. D 104, 015030 (2021)
- A model-independent approach to any new particle
 - [9] Brdar, Fischer, Smirnov, PRD 103, 075008 (2021)



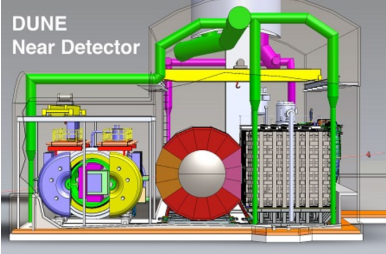
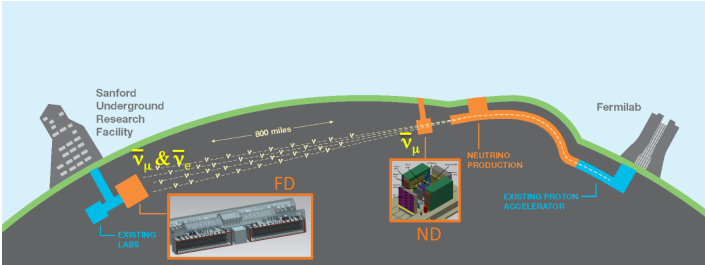
First Results from Short-Baseline Neutrino Program



Outline

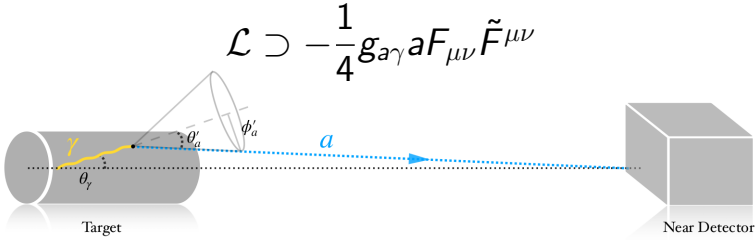
- ▶ MiniBooNE Anomaly
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- ▶ Resonances in $\bar{\nu}_e - e^-$ scattering at FPF
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DUNE

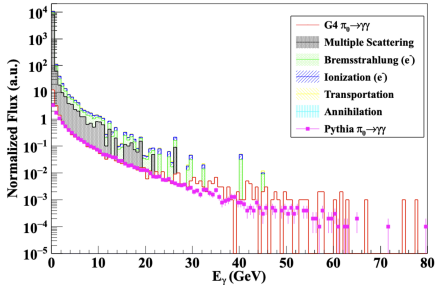


DUNE
DEEP UNDERGROUND
NEUTRINO EXPERIMENT

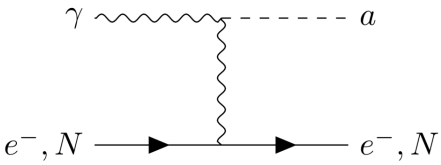
BSM at DUNE: Axion-like Particles (ALPs)



production of photons

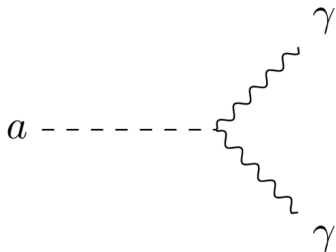


production of ALPs

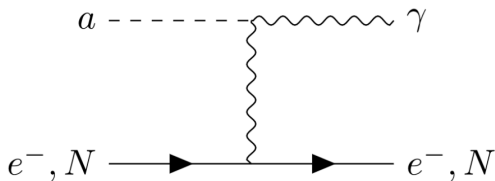


Detection of ALPs

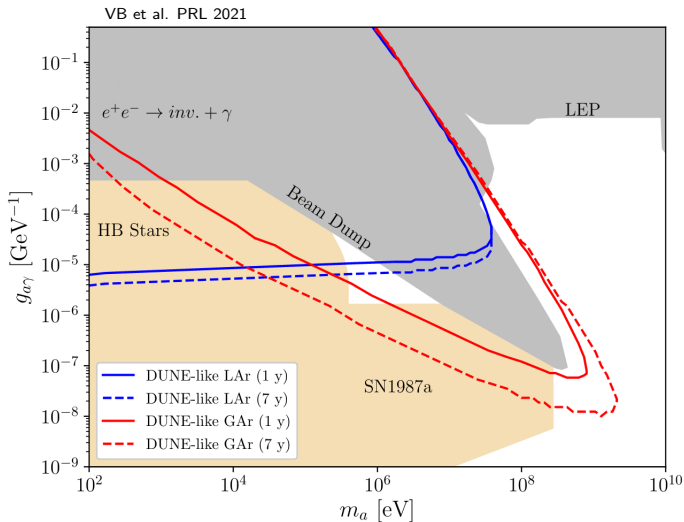
Decay



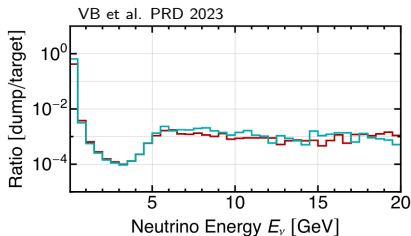
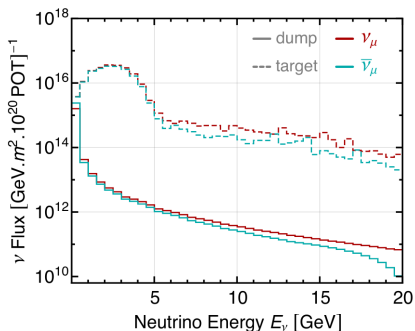
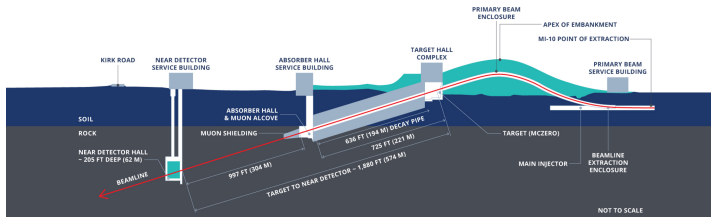
Scattering



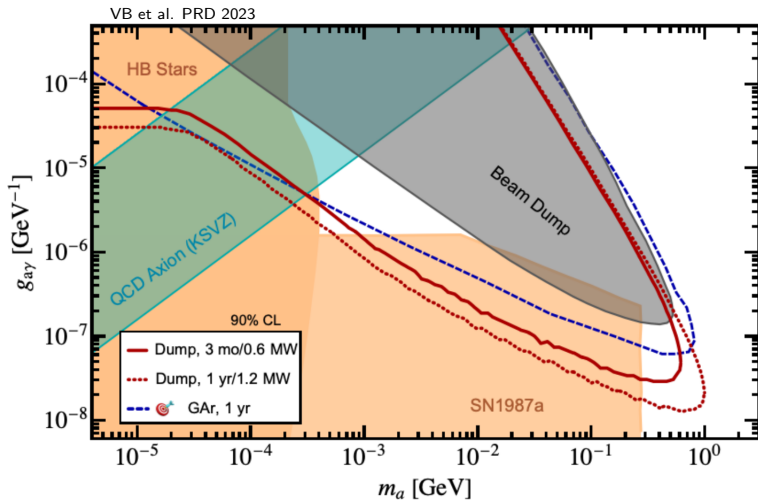
Sensitivity



Target-less DUNE



Sensitivity



Further BSM Scenarios

adopted from Jae Yu

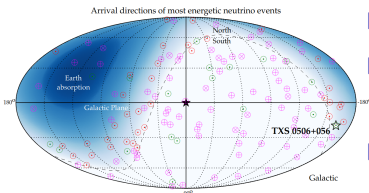
Process	Signatures	Background
ALP	Scattering: $\gamma+e/\gamma+N$ (n) Decay in flight : $\gamma\gamma$	ν coherent, NC w/ π^0 , ν_e CC w/ π^0 , etc
LDM	$\chi e^- \rightarrow \chi e^-$, $\chi N \rightarrow N' n$	NC w/ π^0 , ν_e CC, QE, RES
mCP	Multiple e^- scatterings	ν_e CC w/ π^0
Dark Photon	$A \rightarrow e^- e^+$, $\mu^- \mu^+$	ν CC + mis-ID π , Accidental overlap of CC
HNL	$N \rightarrow \nu e^- e^+$, $\nu \mu^- \mu^+$, $\nu e \mu$, $\nu \pi^0$, $e\pi$, $\mu\pi$	ν CC + mis-ID π , ν_e CC w/ π^0
ν trident	$\nu \rightarrow \nu e^- e^+$, $\nu \mu^- \mu^+$, $\nu e \mu$	$\nu_\mu N \rightarrow \nu_\mu \pi N \square$ (ν CC)
BDM/ iBDM	$\chi N \rightarrow e^- N$	ν coherent, NC w/ π^0 , ν_e CC

Outline

- ▶ MiniBooNE Anomaly
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- ▶ Neutrino Magnetic Moment and Future Colliders

First Glashow Resonance Event at IceCube

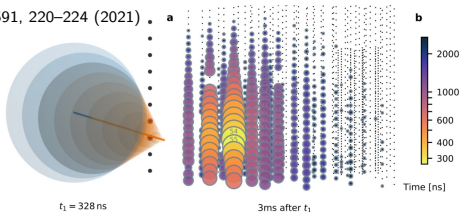
Bull. Am. Astron. Soc. 51, 185 (2019)



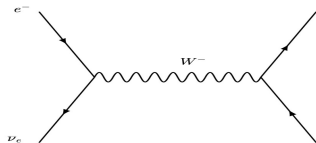
- ▶ $\mathcal{O}(100)$ upgoing tracks and HESE events
- ▶ Glashow resonance: cross section enhancement from on-shell W^- production
- ▶ $\sigma \propto \frac{1}{(E-E_0)^2 + \Gamma^2}$, $E_0 = \frac{M_W^2}{2m_e} \approx 6.3 \text{ PeV}$

- ▶ $\bar{\nu}_e$ in the astrophysical flux \implies way to distinguish ν from $\bar{\nu}$
- ▶ PeV partially contained events: shower with energy of $6.05 \pm 0.72 \text{ PeV}$
- ▶ $\mathcal{O}(\text{PeV})$ cosmic muon as BKG yields 10^{-7} events \implies rejected at 5σ

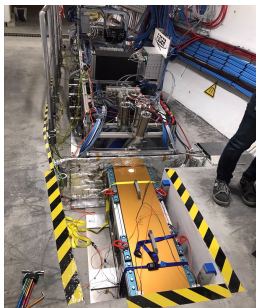
Nature 591, 220–224 (2021)



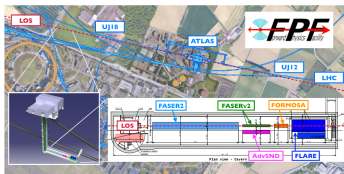
S. Glashow, Phys. Rev. 118, 316



Collider Neutrinos

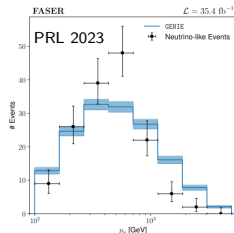
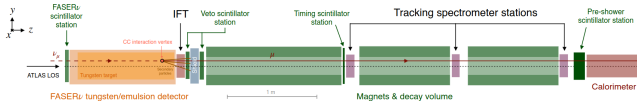


- ▶ **FASER ν** is a 1.2 tonne detector located 480 m from the ATLAS interaction point containing emulsion films and tungsten plates

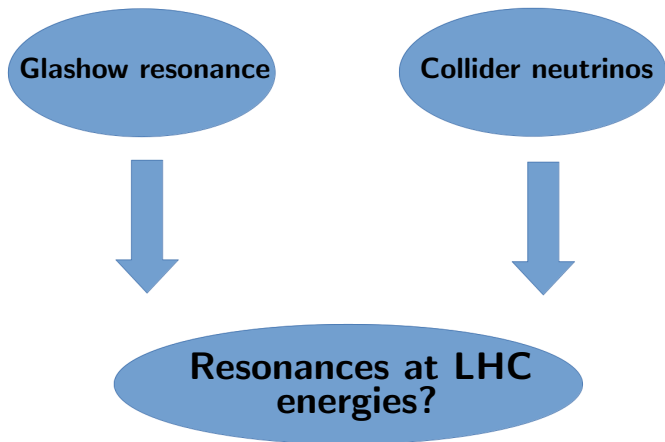


- ▶ **FASER ν 2** and **FLArE (LAR)** at FPF

- ▶ 153 ν_μ charged-current interactions at FASER
 \implies first direct observation of ν interactions at a particle collider experiment



“Glashow-like” Events at Low Energies?



“Glashow-like” Events at Low Energies?

$$\bar{\nu}_e e^- \rightarrow R^-$$

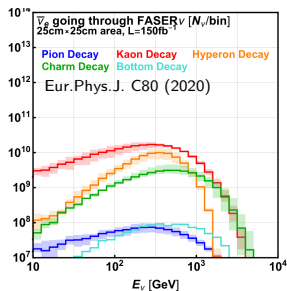
- ▶ electrically charged
- ▶ no baryon number
- ▶ no lepton number



W^- works but **mesons as well!**

$$s = 2E_\nu m_e \sim \text{GeV}^2$$

- ▶ there are many GeV-scale charged mesons



“Glashow-like” Events at Low Energies?

$$\bar{\nu}_e e^- \rightarrow \text{meson} \rightarrow \text{anything}$$

Breit-Wigner:
$$\sigma_{\text{res}} = (2J + 1)8\pi \Gamma^2 \text{Br}_{\text{in}} \text{Br}_{\text{fi}} \frac{s/M^2}{(s - M^2)^2 + M^2\Gamma^2}$$

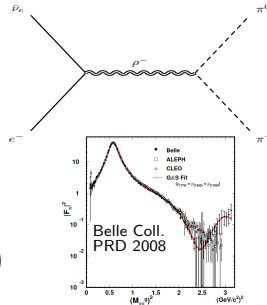
- ▶ pseudoscalar mesons: $\Gamma(\text{m} \rightarrow \bar{\nu}_e e^-) = \frac{G_F^2}{8\pi} f^2 m_{\text{lep}}^2 M \left(1 - \frac{m_{\text{lep}}^2}{M^2}\right) |V_{\text{CKM}}|^2$
- ▶ vector mesons: $\Gamma(\text{m} \rightarrow \bar{\nu}_e e^-) = \frac{G_F^2}{12\pi} f^2 M^3 |V_{\text{CKM}}|^2$

$$\bar{\nu}_e e^- \rightarrow \rho^- \rightarrow \pi^0 \pi^-$$

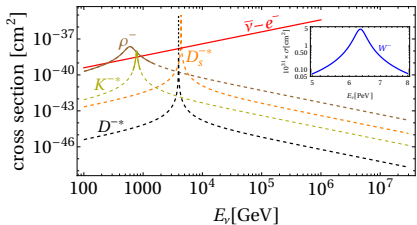
$$E_{\nu}^{\text{res}}(\rho^-) = \frac{(770\text{MeV})^2}{2m_e} \approx 580 \text{ GeV}$$

$$E_{\nu}^{\text{res}}(K^{*-}) \approx 780 \text{ GeV}$$

- ▶ alternative calculation using $\langle \pi^-(k_1) \pi^0(k_2) | V_{\mu} | 0 \rangle = (k_1 - k_2)_{\mu} F(q^2)$

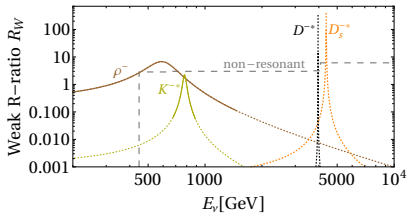


Event Rates

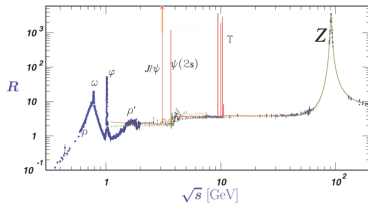


Experiment	$\rho^-, \pm\Gamma/2$	$\rho^-, \pm 2\Gamma$	$K^{*-}, \pm\Gamma/2$	$K^{*-}, \pm 2\Gamma$
FASER ν	0.3	0.5	–	–
FASER $\nu/2$	23	37	0.7	3
FLArE-10	11	19	0.3	2
FLArE-100	63	103	2	8
DeepCore	3 (1)	5 (2)	–	–
IceCube	8 (40)	17 (83)	–	–

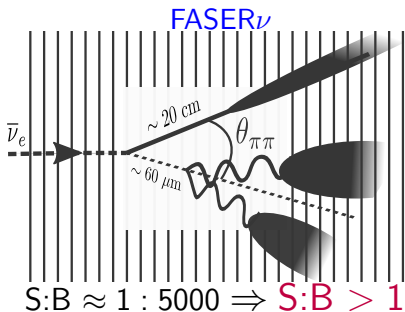
$$R_W = \frac{\sigma(\bar{\nu}_e e^- \rightarrow \text{hadrons})}{\sigma(\bar{\nu}_e e^- \rightarrow \bar{\nu}_\mu \mu^-)}$$



$$R = \frac{\sigma(e^+ e^- \rightarrow \text{hadrons})}{\sigma(e^+ e^- \rightarrow \mu^+ \mu^-)}$$



Signature of ρ^- Resonance



- ▶ cut on $E_{\pi^-} + E_{\pi^0}$ to lie near 580 GeV
- ▶ $\theta_{\nu N} \sim 1/\gamma_{\text{cm}} \sim 28 \text{ mrad} \times \sqrt{600 \text{ GeV}/E_\nu}$ for **deep inelastic scattering**
- ▶ $\theta_{\pi\pi} = 28 \text{ mrad} \sqrt{m_e/m_N} \times \sqrt{600 \text{ GeV}/E_\nu} = 0.7 \text{ mrad} \times \sqrt{600 \text{ GeV}/E_\nu}$ for $\bar{\nu} - e$ scattering
- ▶ cut on charged track and photon multiplicity
- ▶ reconstruct the invariant mass of the $\pi^0\pi^-$ pair, $m_{\pi\pi}^2 = m_{\pi^0}^2 + m_{\pi^-}^2 + E_{\pi^0}E_{\pi^-} - \theta_{\pi\pi}^2$, and require it to lie within $\Gamma_\rho \sim 150 \text{ MeV}$ of $m_\rho \approx 770 \text{ MeV}$

- ▶ Sweeper Magnet for FASER ν 2

IceCube:

- ▶ large background and difficult to identify $\pi^-\pi^0$ topology
- ▶ S : B \approx 1 : 100

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Magnetic Moment for Massive Neutrinos

$$\mathcal{L} \supset \frac{1}{2} \mu_{\nu}^{\alpha\beta} \bar{\nu}_L^{\alpha} \sigma^{\mu\nu} \nu_R^{\beta} F_{\mu\nu}$$

- ▶ neutrinos are massless in the Standard Model

$$\implies \mu_{\nu}^{\alpha\beta} = 0$$

- ▶ adding ν_R to the Standard Model generates Dirac mass $m_{\nu} \bar{\nu}_L \nu_R$ which generates nonzero μ_{ν}

$$\mu_{\nu}^{\text{diag}} = \frac{3eG_F m_{\nu}}{8\sqrt{2}\pi^2} \approx 3 \times 10^{-20} \mu_B \left(\frac{m_{\nu}}{0.1 \text{ eV}} \right)$$

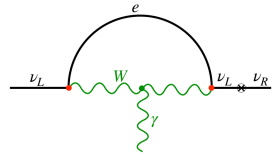
- ▶ for Majorana neutrinos, only non-diagonal elements are nonzero

$$\mu_{\nu}^{ij} = -\frac{3eG_F(m_i+m_j)}{16\sqrt{2}\pi^2} \sum_k \text{Im} [U_{ki}^* U_{kj}] \frac{m_k^2}{m_W^2} \lesssim 10^{-23} \mu_B \quad \mu_B = \frac{e}{2m_e}$$

- ▶ **not testable** \implies more BSM physics required

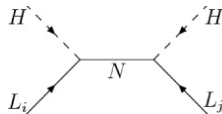
Three Generations of Matter (Feynman) spin 1/2

	I	II	III	spin 0
mass - charge	$\frac{2}{3}$	$\frac{2}{3}$	$\frac{2}{3}$	0
name	u up	c charm	t top	g gluon
Quarks	$\frac{1}{3}$ d down	$\frac{1}{3}$ s strange	$\frac{1}{3}$ b bottom	γ photon
	$\frac{2}{3}$ ν_u neutrino	$\frac{2}{3}$ ν_c neutrino	$\frac{2}{3}$ ν_t neutrino	H Higgs boson
Leptons	-1 e electron	-1 μ muon	-1 τ tau	W boson
	0 ν_e neutrino	0 ν_{μ} neutrino	0 ν_{τ} neutrino	Z boson
				W boson
				spin 0



Sterile Neutrinos

Type-I Seesaw



Minkowski, Mohapatra, Senjanović,
Gell-Mann, Ramond, Slansky, Yanagida

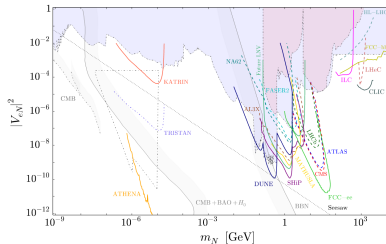
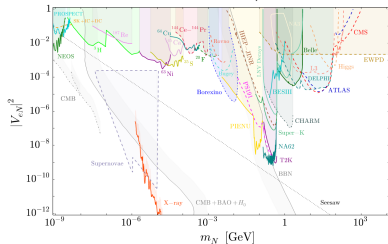
$$\mathcal{M} = \begin{pmatrix} 0 & M_D \\ M_D & M_R \end{pmatrix}$$

$$m_\nu = -M_D M_R^{-1} M_D^T$$

$$\theta = M_D / M_R$$

$$\mathcal{L} \supset \frac{1}{2} \bar{N}^c M_R N + \bar{L} Y_\nu \tilde{H} N + \text{h.c.}$$

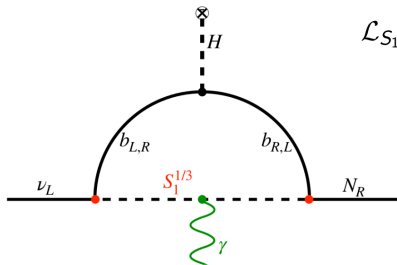
Bolton et al., JHEP 2020



Active-to-Sterile Neutrino Transition Magnetic Moment

$$\mathcal{L} \supset \frac{1}{2} \mu_N \bar{\nu}_L^\alpha \sigma^{\mu\nu} N_R F_{\mu\nu}$$

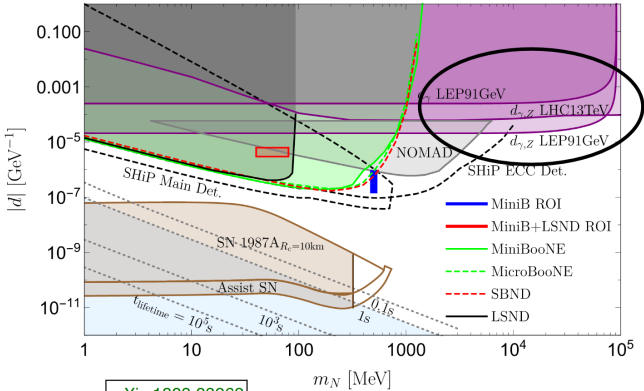
Example for a model generating transition magnetic moment:



$$\mathcal{L}_{S_1} \supset y_1 \bar{b}_R^c N_R S_1 + y_2 \bar{Q}_L^3 L_L^c S_1^\dagger + \text{h.c.}$$

$$\mu_N \approx \frac{e y_1 y_2}{8\pi^2 m_{LQ}^2} m_b \log \frac{m_b^2}{m_{LQ}^2}$$

Collider Probes



arXiv:1803.03262

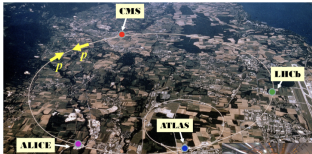
Future Colliders

High Lumi LHC

$$\sqrt{s} = 14 \text{ TeV},$$

$$\mathcal{L} = 3 \text{ ab}^{-1}$$

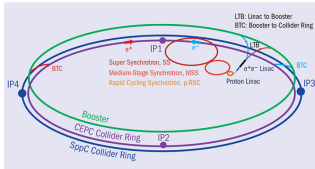
Operational by 2029



Circular e^-e^+ Collider (CEPC)

$$\sqrt{s} = 240 \text{ GeV},$$

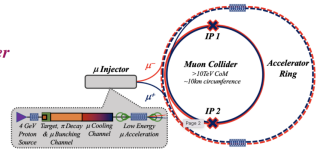
$$\mathcal{L} = 20 \text{ ab}^{-1}$$



Muon Collider

$$\sqrt{s} = 10 \text{ TeV},$$

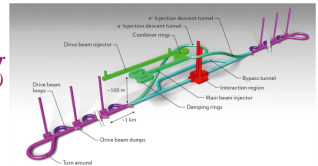
$$\mathcal{L} = 10 \text{ ab}^{-1}$$



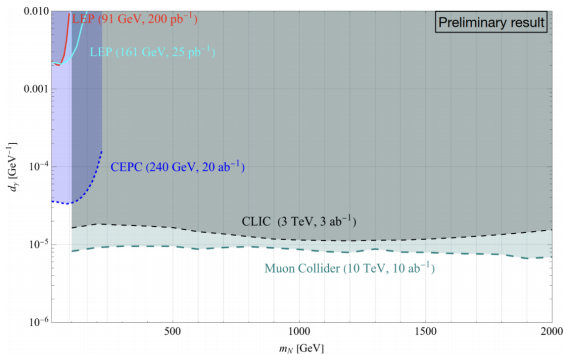
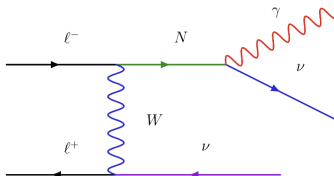
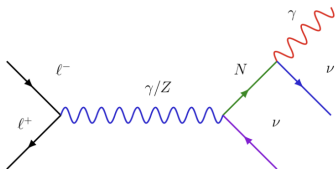
Compact Linear Collider (CLIC)

$$\sqrt{s} = 3 \text{ TeV},$$

$$\mathcal{L} = 3 \text{ ab}^{-1}$$



Collider Probes of Neutrino Transition Magnetic Moment



Summary

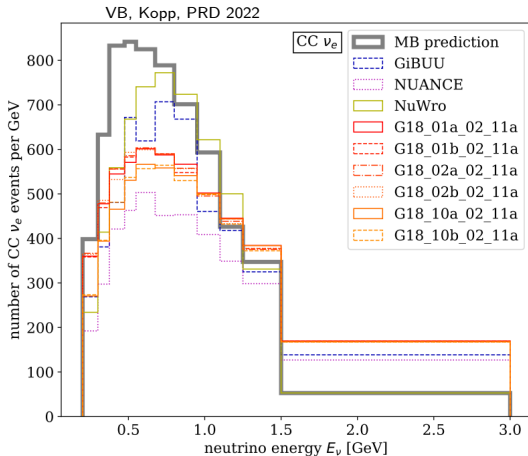
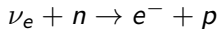
- ▶ Accelerators and colliders are invaluable for the present success and future progress of ν physics

Some examples...

- ▶ Exciting prospects for new physics at near-future experiments such as DUNE (e.g., probing ALPs)
- ▶ The MiniBooNE anomaly, scrutinized from both the SM and BSM sides, is being tested at the Short-Baseline Neutrino Program at Fermilab
- ▶ $\mathcal{O}(100)$ resonance events from $\bar{\nu} - e^-$ scattering are expected at proposed FPF detectors
- ▶ For neutrinophilic BSM models (e.g., neutrino magnetic moment), unconstrained regions in the parameter space can be tested at future colliders

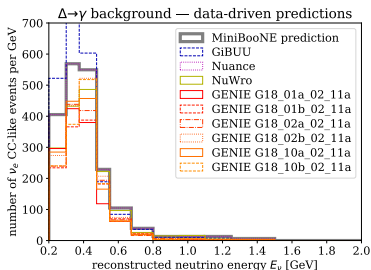
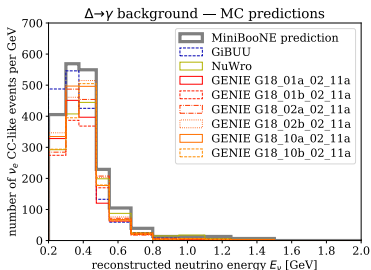
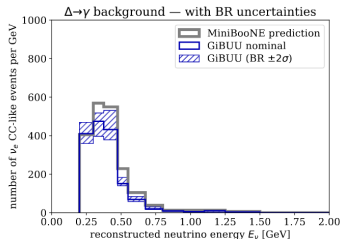
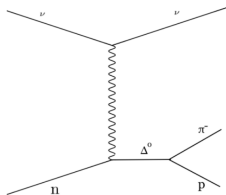
BACKUP

Charged Current Events

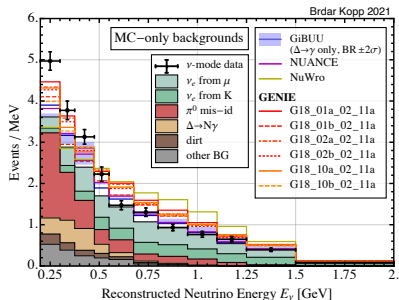


Neutral Current Single γ Production

$$\Delta \rightarrow N \gamma$$



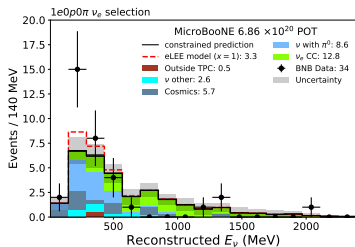
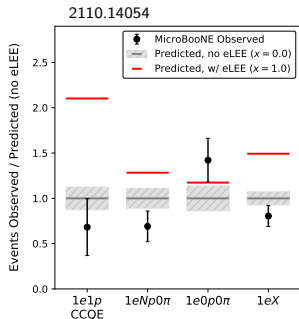
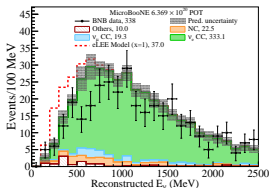
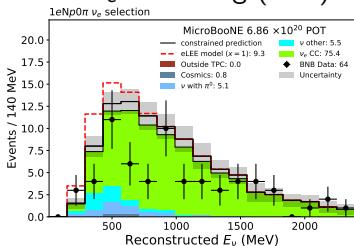
Results: Monte Carlo-only predictions



Generator	Tune	Δm_{41}^2 [eV ²]	$\sin^2 2\theta_{\mu e}$	Significance
MB official		0.25	0.01	4.0 σ
GiBUU	default	0.25	0.01	4.6 σ
	BR($\Delta \rightarrow \gamma$) - 2 σ	0.32	0.0063	4.9 σ
	BR($\Delta \rightarrow \gamma$) + 2 σ	0.32	0.0050	4.2 σ
NUANCE	-	0.32	0.0079	4.0 σ
NuWro	-	3.2	0.0020	3.5 σ
GENIE	G18_01a_02_11a	0.13	0.079	4.3 σ
	G18_01b_02_11a	0.79	0.0001	3.6 σ
	G18_02a_02_11a	0.13	0.050	3.5 σ
	G18_02b_02_11a	0.13	0.050	3.5 σ
	G18_10a_02_11a	0.25	0.016	2.9 σ
	G18_10b_02_11a	0.40	0.013	3.8 σ

First Results from MicroBooNE

two-body ν_e CCQE scattering ($1e1p$)
 pionless ν_e scattering ($1eNp0\pi$, $1e0p0\pi$)
 inclusive ν_e scattering ($1eX$)



First Results from MicroBooNE

arXiv:2110.00409

$1\gamma 1p$

$1\gamma 0p$

	$1\gamma 1p$	$1\gamma 0p$
Unconstr. bkgd.	27.0 ± 8.1	165.4 ± 31.7
Constr. bkgd.	20.5 ± 3.6	145.1 ± 13.8
NC $\Delta \rightarrow N\gamma$	4.88	6.55
LEE ($x_{MB} = 3.18$)	15.5	20.1
Data	16	153

Process	$1\gamma 1p$	$1\gamma 0p$
NC $1\pi^0$ Non-Coherent	24.0	68.1
NC $1\pi^0$ Coherent	0.0	7.6
CC $\nu_\mu 1\pi^0$	0.5	14.0
CC ν_e and $\bar{\nu}_e$	0.4	11.1
BNB Other	2.1	18.1
Dirt (outside TPC)	0.0	36.4
Cosmic Ray Data	0.0	10.0
Total Background (Unconstr.)	27.0	165.4
NC $\Delta \rightarrow N\gamma$	4.88	6.55

