

*This work is supported by NKFIH K-133046, K-147557 and MATE KKP (2023) grants.*

*Preprints are available: [arXiv:2311.18678](https://arxiv.org/abs/2311.18678) [arXiv:2311.03568](https://arxiv.org/abs/2311.03568)*

# Describing hadronic and thermal photon distributions with an analytic solution of hydrodynamics

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GÁBOR KASZA, TAMÁS CSÖRGŐ

**ZIMÁNYI SCHOOL 2023**

BUDAPEST, 12/07/2023



# Various application of hydrodynamics

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- Why is hydrodynamics so effective?

- Works well on ...

- ... *microscopic scales*
  - ... *macroscopic scales*
  - ... *cosmic scales*

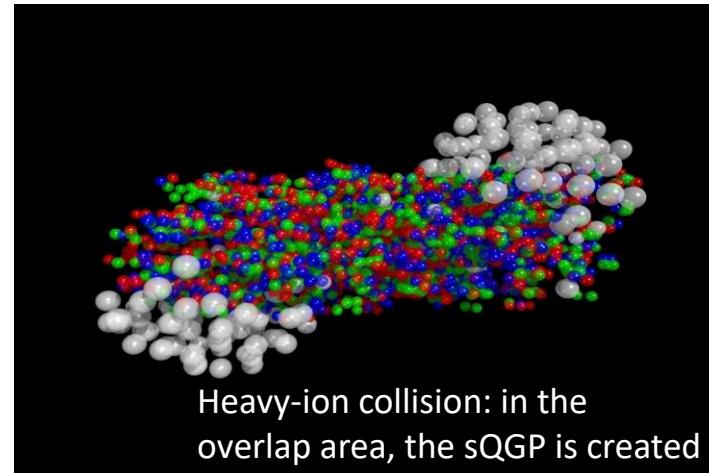
- ***Hydrodynamics has no internal scale***

- Averages the microscopic degrees of freedom

- Analytic hydro: with relatively simple equations we can ...

- ... describe the scaling behaviour of experimental data

- ... provide a tool for checking more complex calculations



Heavy-ion collision: in the overlap area, the sQGP is created

*Today's presentation:*

→ ***scaling of  $dN/dy$***

→ ***describe the thermal radiation***

# Relativistic perfect fluid solution with accelerating velocity field

- Equation of state:  $\varepsilon = \kappa_0 p$  ( $\mu=0$ )

- Rindler coordinates:  $\tau = \sqrt{t^2 - r_z^2}$

$$\eta_z = \frac{1}{2} \ln \left( \frac{t + r_z}{t - r_z} \right)$$

- Velocity field:  $u^\mu = \begin{pmatrix} \cosh(\Omega) \\ \sinh(\Omega) \end{pmatrix}$

$$\Omega \equiv \Omega(\eta_z)$$

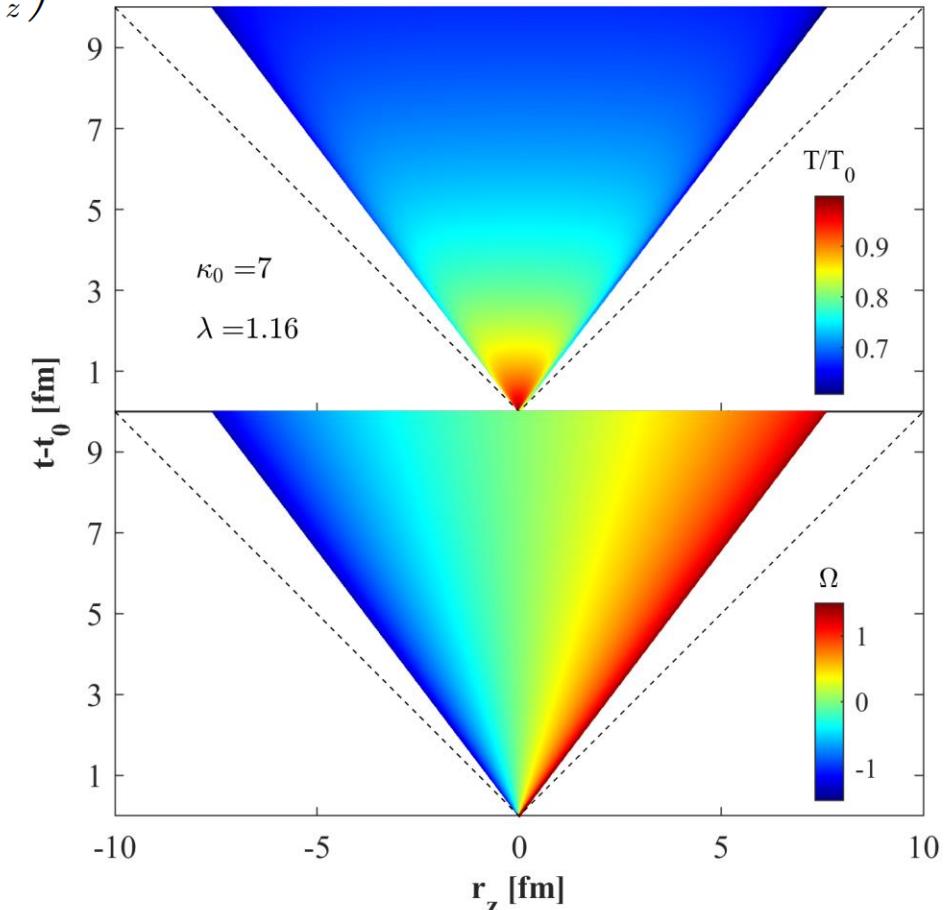
- 1+1 dimensional, parametric, almost self-similar, finite solution

$\lambda$ : rate of acceleration  
(Hwa-Bjorken:  $\lambda=1$ )

accelerating expansion

realistic  $dN/dn_p$

T. Csörgő, G. Kasza, M. Csanád and Z-F. Jiang:  
Universe 4 (2018) 69



# *Scaling of $dN/dy$*

[arXiv:2311.18678](https://arxiv.org/abs/2311.18678)

# Pseudorapidity distribution

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- Starting from the rapidity distribution, we calculated the pseudorapidity distribution
- Parametric curve:

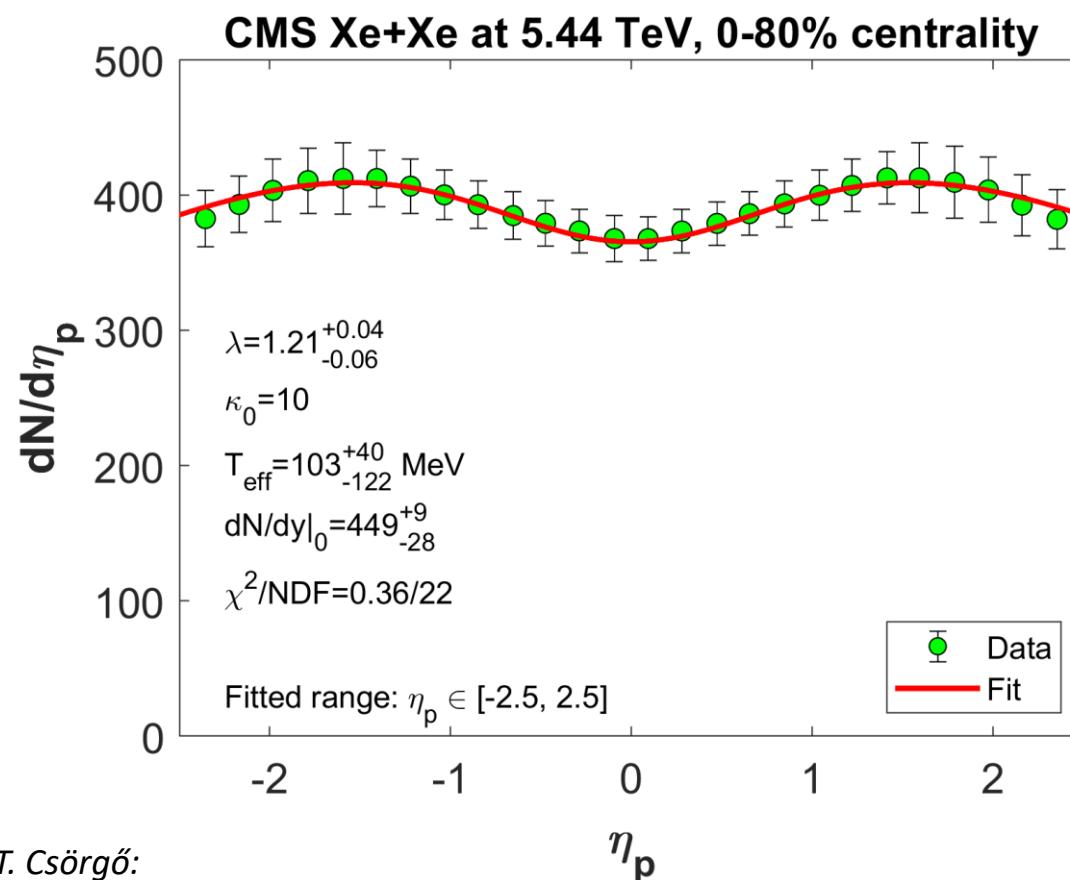
$$\left( \eta_p(y), \frac{dN}{d\eta_p}(y) \right) = \left( \frac{1}{2} \ln \left[ \frac{\langle |p(y)| \rangle + \langle p_z(y) \rangle}{\langle |p(y)| \rangle - \langle p_z(y) \rangle} \right], \frac{\langle |p(y)| \rangle}{\langle E(y) \rangle} \frac{dN}{dy} \right)$$

- We compared this curve with experimental data :
  - PHOBOS Au+Au 130 GeV, 200 GeV
  - ALICE Pb+Pb 5.02 TeV
  - CMS p+p 7 TeV, 8 TeV, 13 TeV → **described very well, collectivity in p+p collisions?**
  - CMS Xe+Xe 5.44 TeV
- Different sizes (p+p, A+A), different centralities, different collision energies

G. Kasza, T. Csörgő:  
*Int.J.Mod.Phys. A34 no.26, 1950147 (2019)*

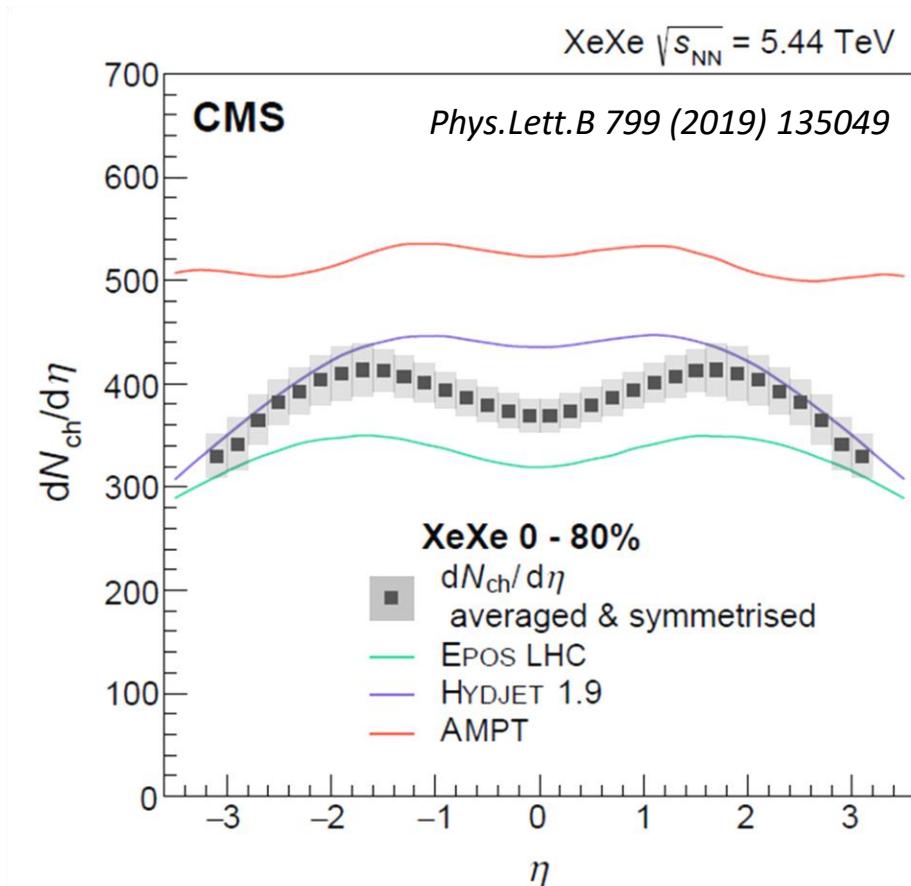
# Fits to CMS data (Xe+Xe)

- The parametric formula works well in such cases where other models fail:



G. Kasza, T. Csörgő:

Int.J.Mod.Phys. A34 no.26, 1950147 (2019)



# A reasonable approximation of the rapidity distribution

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- If  $|y| \ll 2+1/(\lambda-1)$ , then the rapidity distribution becomes Gaussian:

$$\frac{dN}{dy} \approx \frac{\langle N \rangle}{(2\pi\Delta y^2)^{1/2}} \exp\left(-\frac{y^2}{2\Delta y^2}\right)$$

- Manifest of the hydrodynamic scaling behaviour:

$$\frac{1}{\Delta y^2} = (\lambda - 1)^2 \left[ 1 + \left( 1 + \frac{1}{\kappa_0} \right) \left( \frac{1}{2} + \frac{m}{T_{\text{eff}}} \right) \right]$$

$$\langle N \rangle = (2\pi\Delta y^2)^{1/2} \left. \frac{dN}{dy} \right|_{y=0}$$

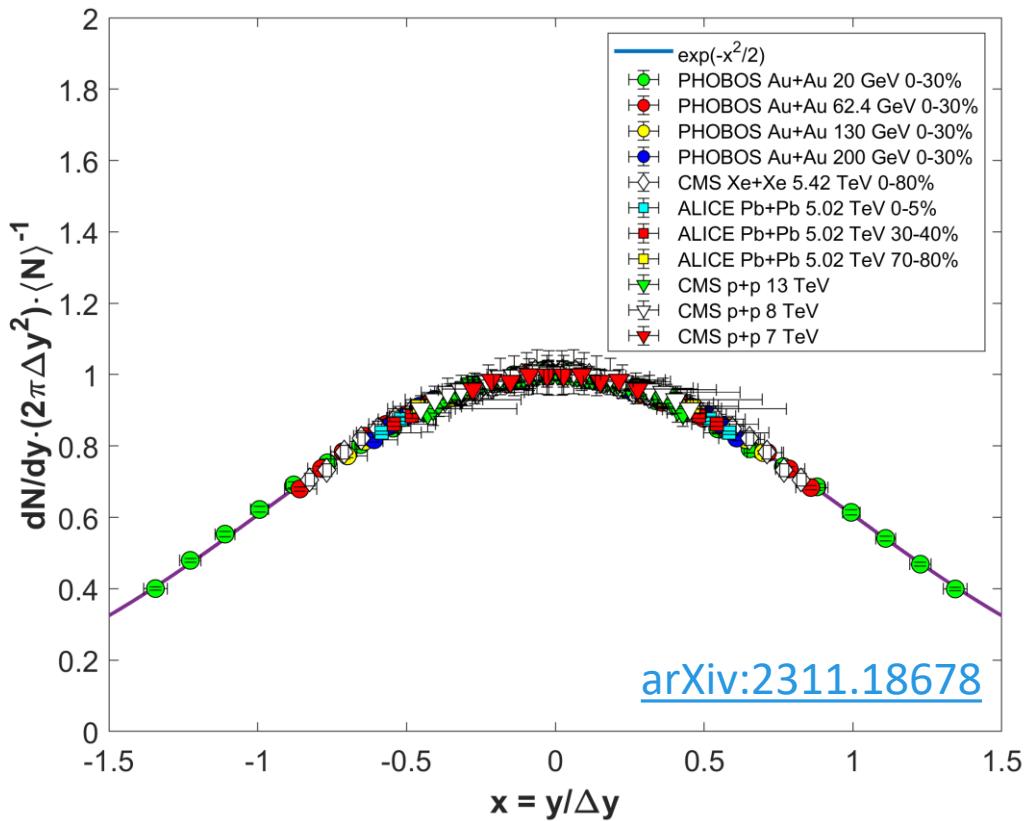
- The physical properties of different collisions (vs, centrality, size) are scaled out:

$$\frac{dN}{dy} = \left. \frac{dN}{dy} \right|_{y=0} \exp\left(-\frac{y^2}{2\Delta y^2}\right) \rightarrow f(x) = \exp\left(-\frac{x^2}{2}\right)$$

[arXiv:2311.18678](https://arxiv.org/abs/2311.18678)

# Data collapsing

- Pseudorapidity distributions were transformed into rapidity distributions → fits to the  $dN/dy$  data series



- Data collapsing on the  $f(x)=\exp(-x^2/2)$  curve of the scale function

For each dataset, the fit range satisfies:  
 $|y| \ll 2 + 1/(\lambda - 1)$

This condition is the strictest in the case of  
*PHOBOS Au+Au@20 GeV* →  $|y| \ll 3$ ,  $|x| \ll 1.8$

Another prediction for Gaussian  $dN/dy$  from  
Landau hydrodynamics:

C-Y. Wong: *Phys. Rev. C* 78, 054902

C-Y. Wong et al.: *Phys. Rev. C* 90 (2014) 6, 064907

A. Sen et al.: *J. Phys. Conf. Ser.* 630 (2015) 1, 012042

# *Thermal photon radiation*

[arXiv:2311.03568](https://arxiv.org/abs/2311.03568)

# Derivation of the thermal radiation

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Source function:

$$S(x^\mu, p^\mu) d^4x = \frac{g}{(2\pi\hbar)^3} \frac{H(\tau)}{\tau_R} \frac{p_\mu d\Sigma^\mu}{\exp\left(\frac{p^\mu u_\mu}{T}\right) - 1}$$

Using the **1+1 dimensional CKCJ solution**:



$$d\Sigma^\mu = \frac{u^\mu \tau d\tau d\eta_z dr_x dr_y}{\cosh(\Omega(\eta_z) - \eta_z)}$$

Assuming **homogeneous transverse distribution** of temperature

**Window function:** sQGP is transparent for photons

$$H(\tau) = \Theta(\tau - \tau_f) - \Theta(\tau - \tau_0)$$

[arXiv:2311.03568](https://arxiv.org/abs/2311.03568)

# The new analytic formula for the thermal radiation

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*The source function is integrated over space and time*

- Motivated by earlier results:  $\lambda$  was assumed to be close to 1
- The integral was performed by saddle point approximation
- The result is evaluated at midrapidity ( $y \approx 0$ )

$$\frac{d^2N}{2\pi p_T dp_T dy} \Big|_{y=0} = N_0 \frac{2\alpha}{3\pi^{3/2}} \left[ \frac{1}{T_f^\alpha} - \frac{1}{T_0^\alpha} \right]^{-1} p_T^{-\alpha-2} \left[ \Gamma \left( \alpha + \frac{5}{2}, \frac{p_T}{T_0} \right) - \Gamma \left( \alpha + \frac{5}{2}, \frac{p_T}{T_f} \right) \right]$$

$\lambda$  and  $\kappa$  are collapsed into  $\alpha$  (typical behaviour of hydro):

$T_f$ : freeze-out temperature

$$\alpha = \frac{2\kappa}{\lambda} - 3$$

$T_0$ : initial temperature

$N_0$ : multiplicity at midrapidity

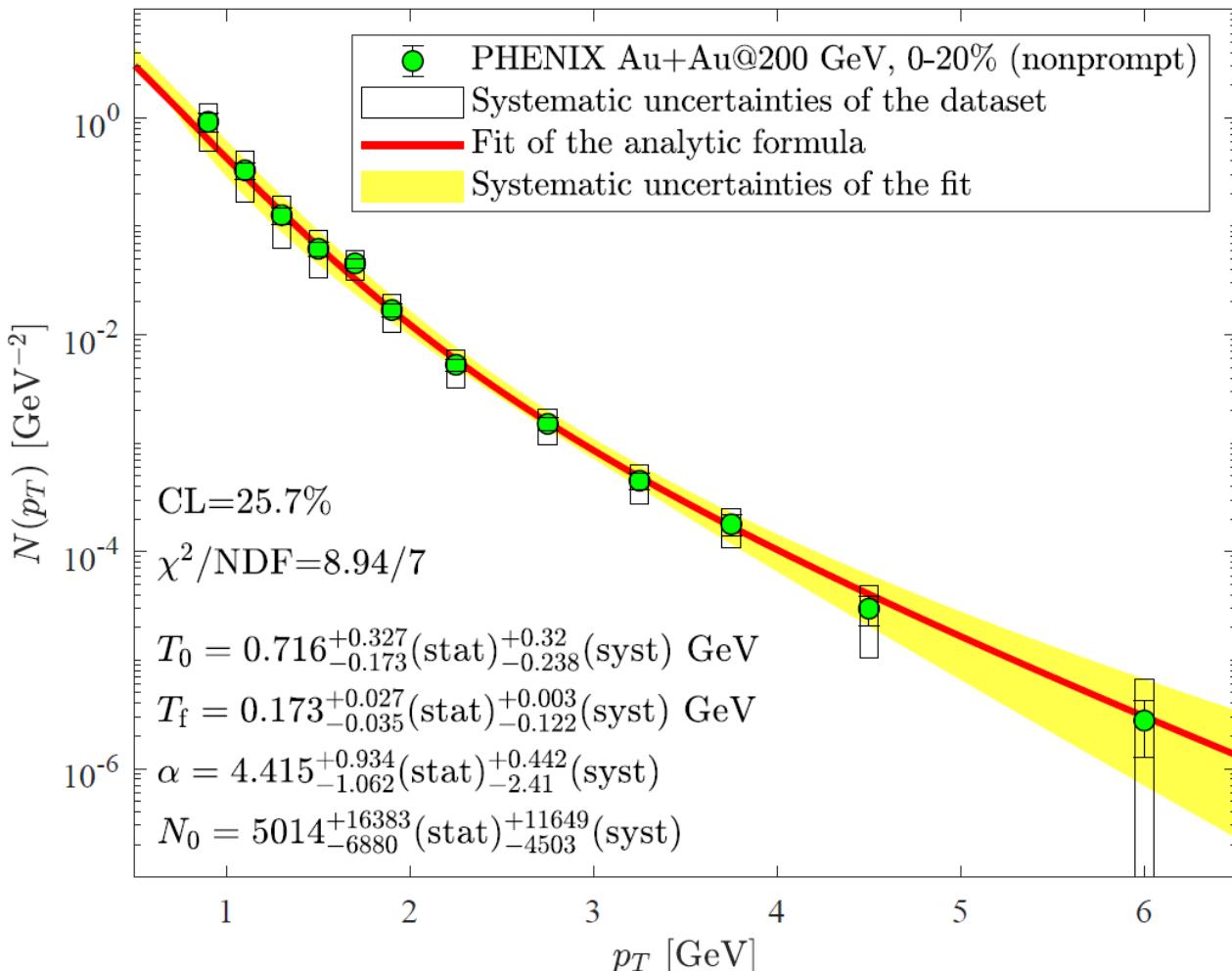
[arXiv:2311.03568](https://arxiv.org/abs/2311.03568)

# Fit to experimental data

Good confidence level with realistic values of physical parameters

Intermediate  $p_T$  regime  $\rightarrow T_0$  can be determined more precisely

Data is from: [arXiv:2203.17187](https://arxiv.org/abs/2203.17187)



# In conclusion...

*Thank you for your attention!*

- Scaling behaviour of  $dN/dy$

- p+p collisions can be described as collective systems
- Our fits indicate low  $c_s$  value ( $\approx 0.35$ ) → indicate the presence of fluid
- p+p and A+A collisions: self-similar systems

- New formula for thermal radiation

- lacks of radial flow and viscous effects
- but describes the data very well
- compare the obtained initial temperature to the Hagedorn temperature (< 350 MeV):

} *Message?*

$$T_H \ll T_0 = 0.7_{-0.2}^{+0.3}(\text{stat})_{-0.2}^{+0.3}(\text{syst})$$

*Backup slides*

# Csörgő-Kasza-Csanád-Jiang (CKCJ) hydro solution

- Rindler coordinates, velocity field:

$$(\tau, \eta_x) = \left( \sqrt{t^2 - r_z^2}, \frac{1}{2} \ln \left[ \frac{t + r_z}{t - r_z} \right] \right)$$

$$u^\mu = (\cosh(\Omega), \sinh(\Omega))$$

- 1+1 dimensional perfect fluid solution:

$$\eta_x(H) = \Omega(H) - H,$$

$$\Omega(H) = \frac{\lambda}{\sqrt{\lambda-1}\sqrt{\kappa-\lambda}} \arctan \left( \sqrt{\frac{\kappa-\lambda}{\lambda-1}} \tanh(H) \right)$$

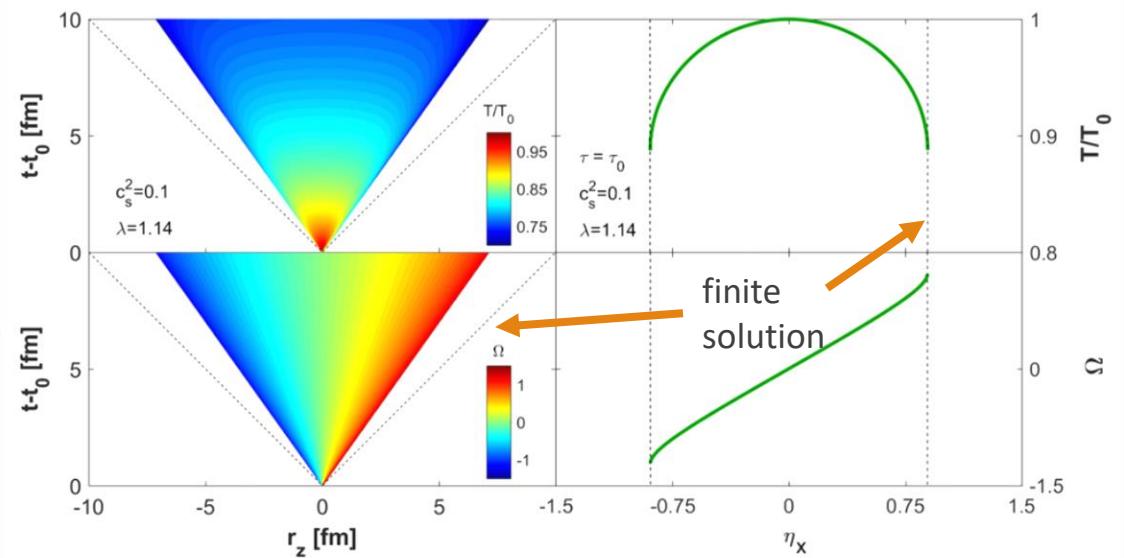
$$\sigma(\tau, H) = \sigma_0 \left( \frac{\tau_0}{\tau} \right)^\lambda \mathcal{V}_\sigma(s) \left[ 1 + \frac{\kappa-1}{\lambda-1} \sinh^2(H) \right]^{-\frac{\lambda}{2}},$$

$$T(\tau, H) = T_0 \left( \frac{\tau_0}{\tau} \right)^{\frac{\lambda}{\kappa}} \mathcal{T}(s) \left[ 1 + \frac{\kappa-1}{\lambda-1} \sinh^2(H) \right]^{-\frac{\lambda}{2\kappa}},$$

$$\mathcal{T}(s) = \frac{1}{\mathcal{V}_\sigma(s)},$$

$$s(\tau, H) = \left( \frac{\tau_0}{\tau} \right)^{\lambda-1} \sinh(H) \left[ 1 + \frac{\kappa-1}{\lambda-1} \sinh^2(H) \right]^{-\lambda/2}$$

*Universe 4 (2018) 6, 69*



Equation of State:

$$\varepsilon = \kappa p$$

(with  $\mu=0$ )

$\lambda$ : acceleration parameter



accelerating solution

realistic  $dN/d\eta_p$

# Rapidity distribution

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- We applied the Cooper-Frye formula
- Temperature is determined on the freeze-out hypersurface
- Integrals were calculated by saddle-point approximation
- Fluid rapidity could be well approximated by a linear function:  $\Omega \approx \lambda \eta_z$
- The 1+1 dimensional rapidity distribution was embedded in 1+3 dimension

$$\frac{dN}{dy} \approx \left. \frac{dN}{dy} \right|_{y=0} \cosh^{-\frac{\alpha(\kappa_0)}{2}-1} \left( \frac{y}{\alpha} \right) \exp \left( -\frac{m}{T_{\text{eff}}} \left[ \cosh^{\alpha(\kappa_0)} \left( \frac{y}{\alpha} \right) - 1 \right] \right)$$

G. Kasza, T. Csörgő:  
*Int.J.Mod.Phys. A34 no.26, 1950147 (2019)*

# Pseudorapidity distribution

- Starting from the rapidity distribution, we calculated the pseudorapidity distribution
- Parametric curve:

$$\left( \eta_p(y), \frac{dN}{d\eta_p}(y) \right) = \left( \frac{1}{2} \ln \left[ \frac{\langle |p(y)| \rangle + \langle p_z(y) \rangle}{\langle |p(y)| \rangle - \langle p_z(y) \rangle} \right], \frac{\langle |p(y)| \rangle}{\langle E(y) \rangle} \frac{dN}{dy} \right)$$



$$\eta_p(y) = \tanh^{-1} (\mathcal{J}^{-1} \tanh(y)) = \tanh^{-1} \left( \frac{\tanh(y)}{\sqrt{1 - \frac{\langle m_T(y) \rangle^2 \cosh^2(y)}{m^2}}} \right)$$

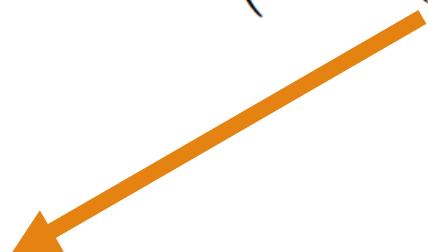
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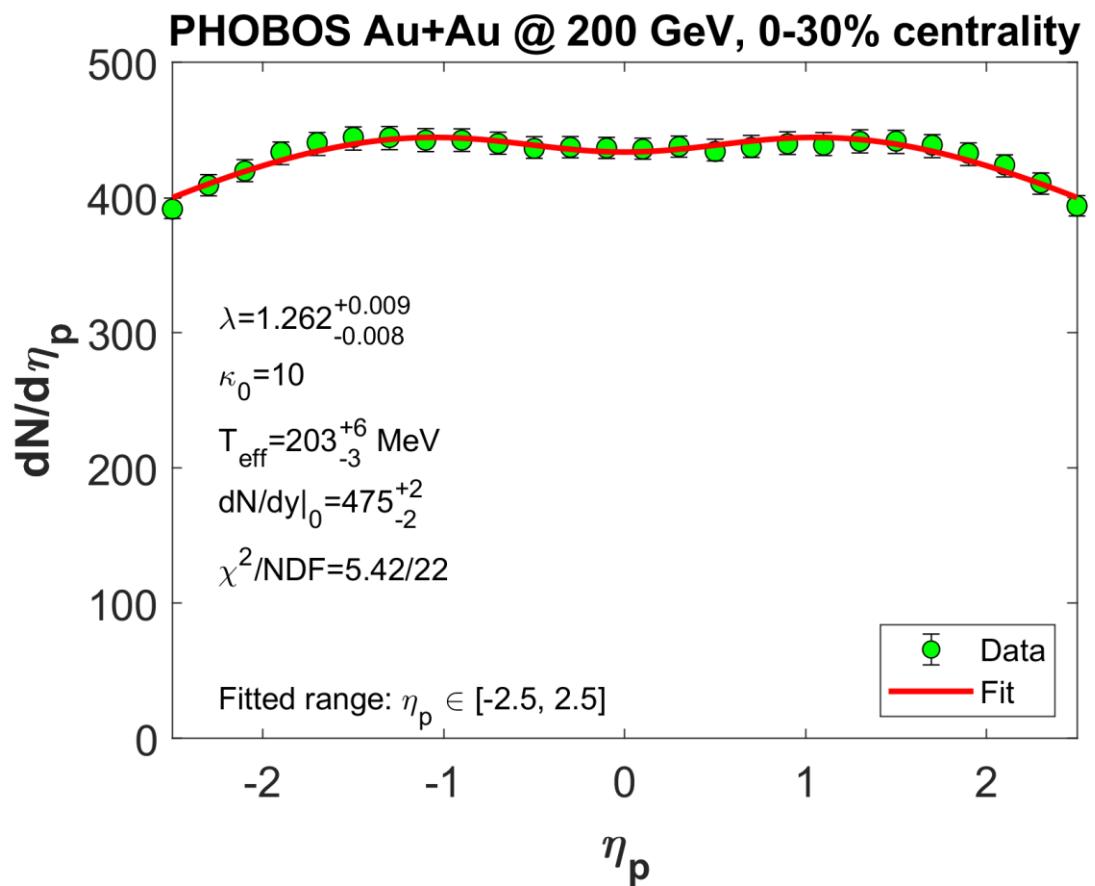
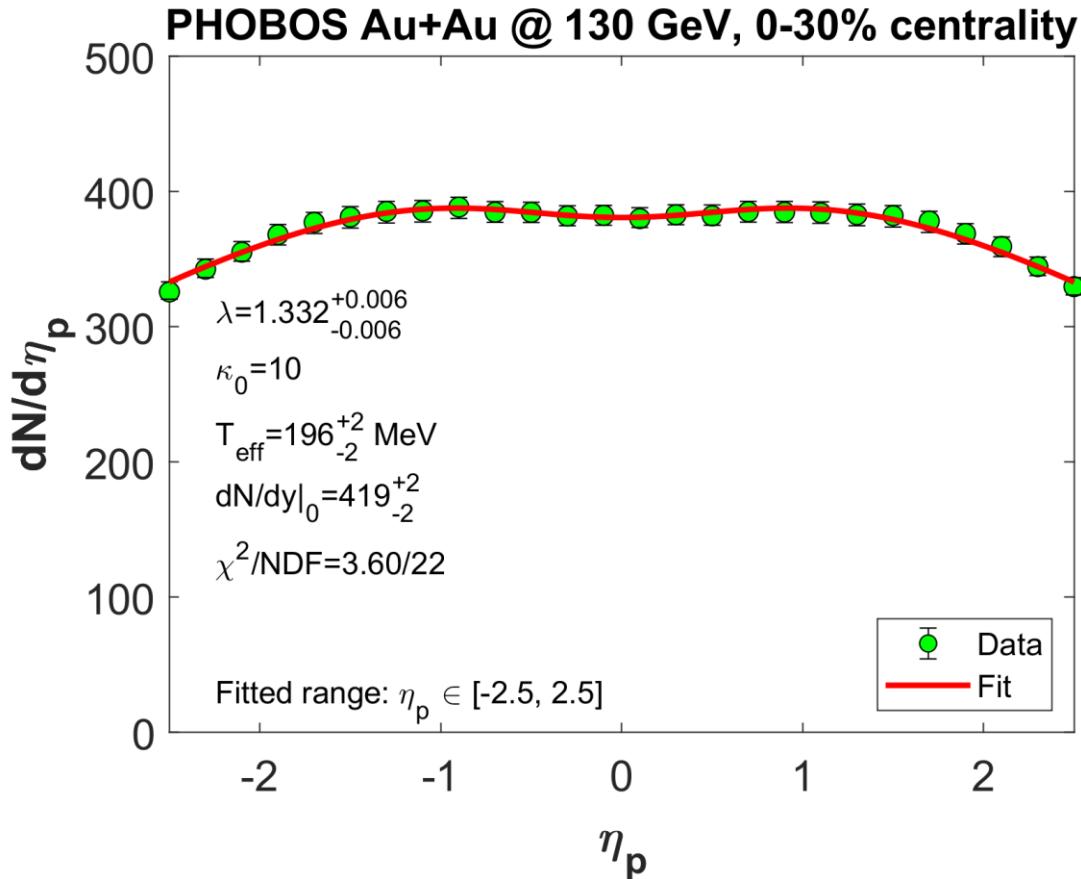
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$$\frac{dN}{d\eta_p}(y) \approx \left. \frac{dN}{dy} \right|_{y=0} \sqrt{1 - \frac{m^2}{\langle m_T(y) \rangle^2 \cosh^2(y)}} \cosh^{-\frac{\alpha(\kappa_0)}{2}-1} \left( \frac{y}{\alpha} \right) \exp \left( -\frac{m}{T_{\text{eff}}} \left[ \cosh^{\alpha(\kappa_0)} \left( \frac{y}{\alpha} \right) - 1 \right] \right)$$

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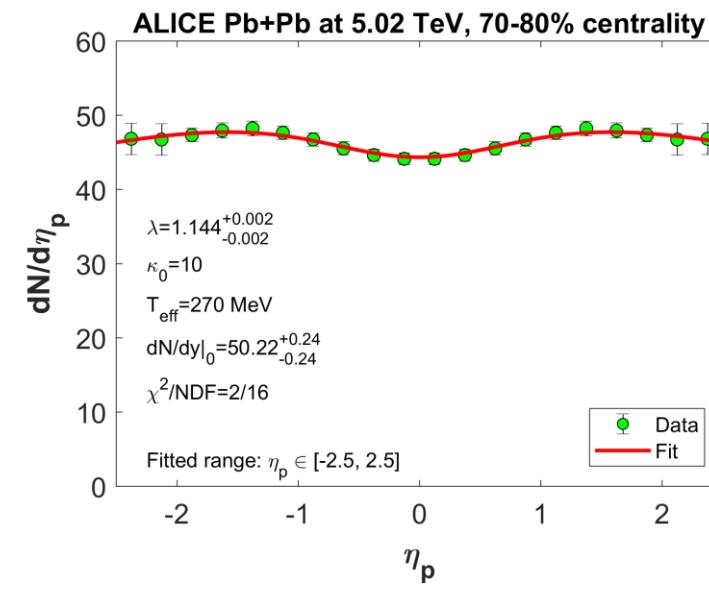
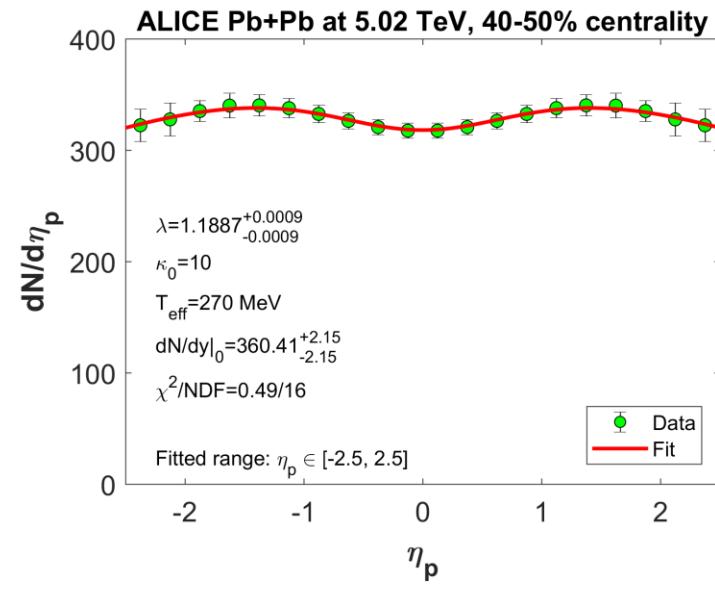
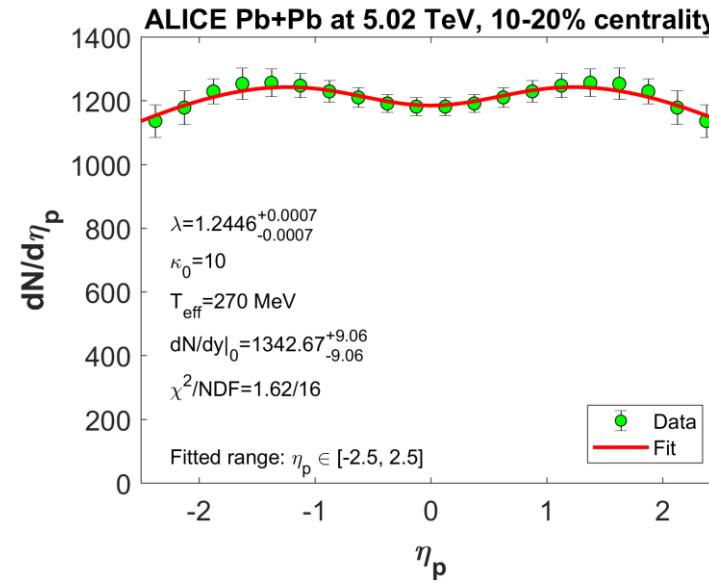
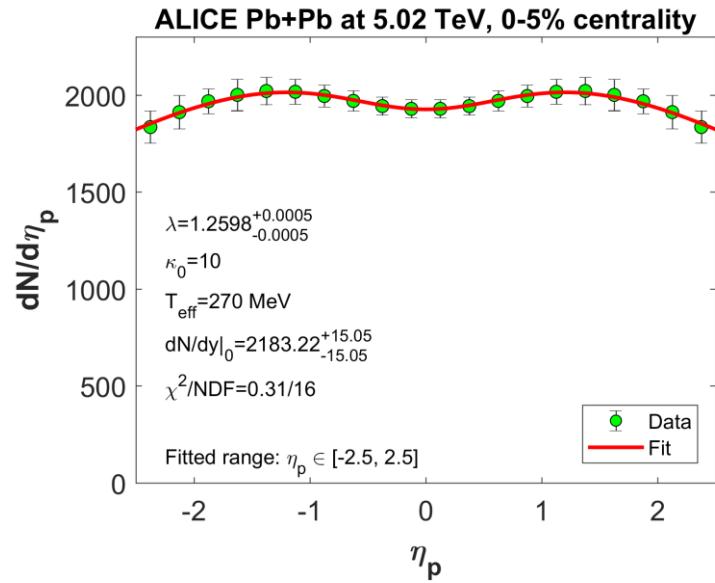
# Fits to PHOBOS data (Au+Au)



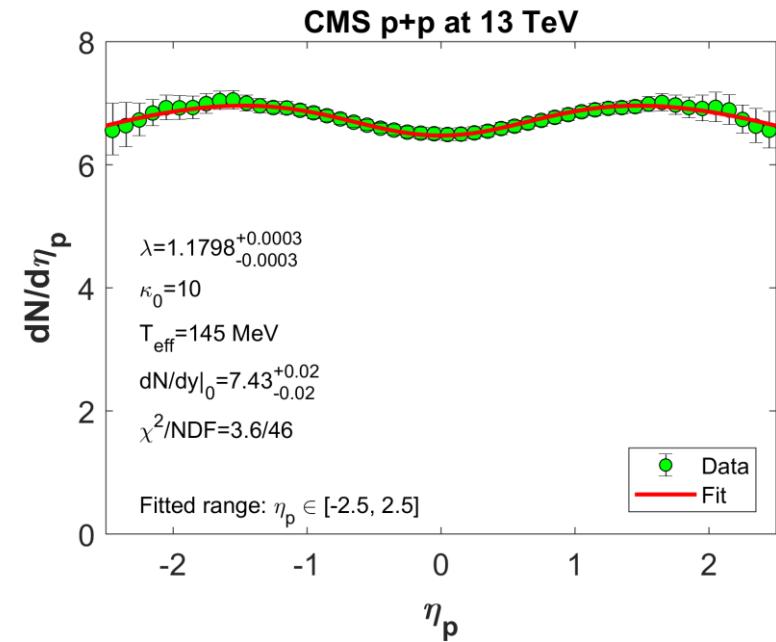
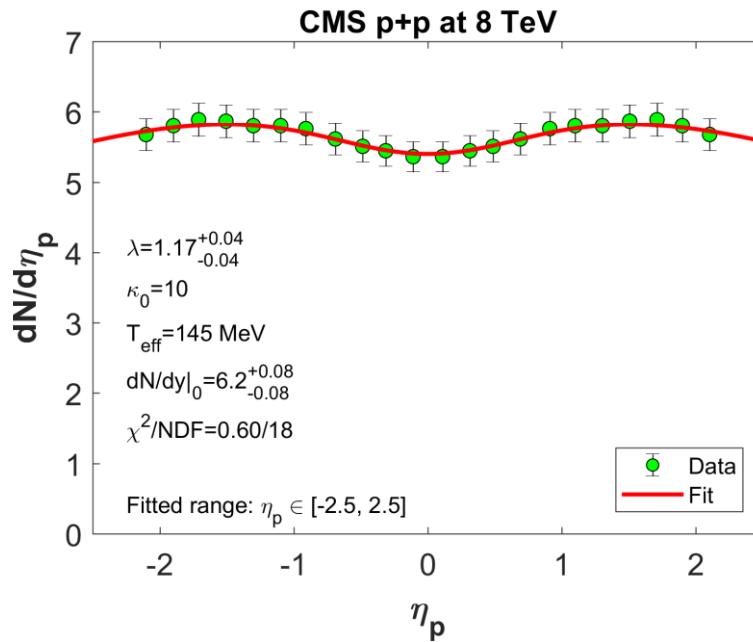
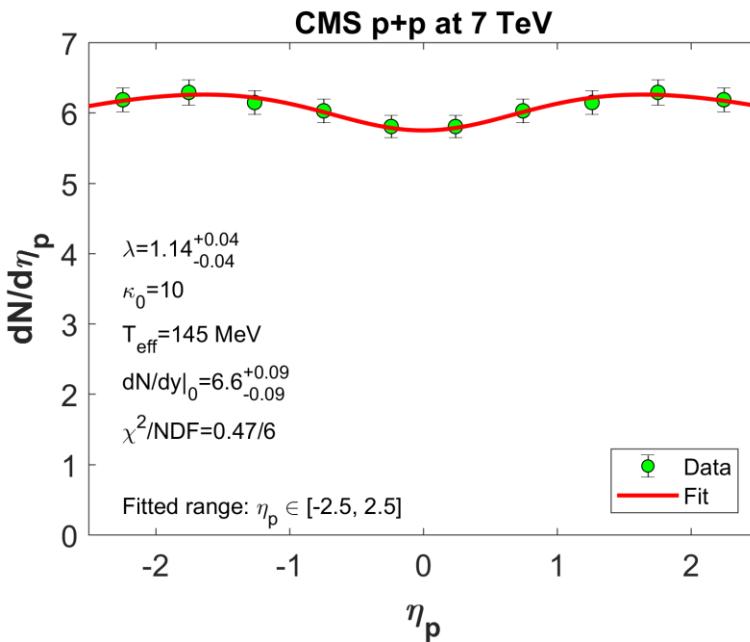
G. Kasza, T. Csörgő:

Int.J.Mod.Phys. A34 no.26, 1950147 (2019)

# Fits to ALICE data (centrality dependence)



# Fits to CMS data (p+p)



**Collectivity in p+p collisions?**

# The new analytic formula for the thermal radiation

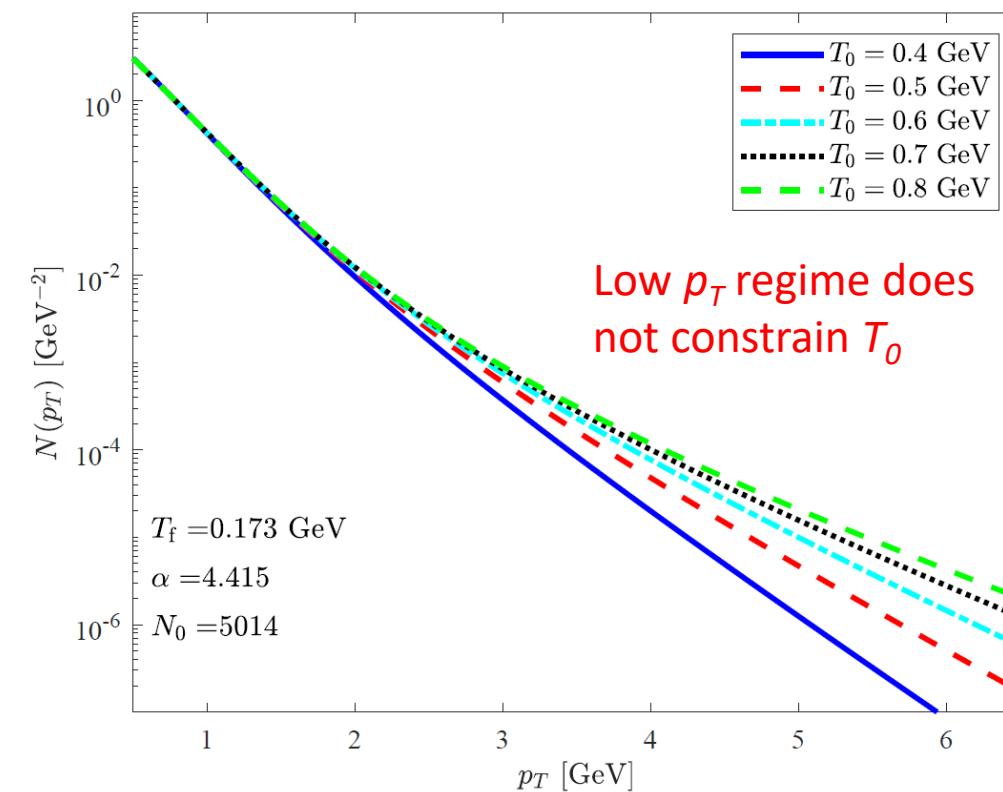
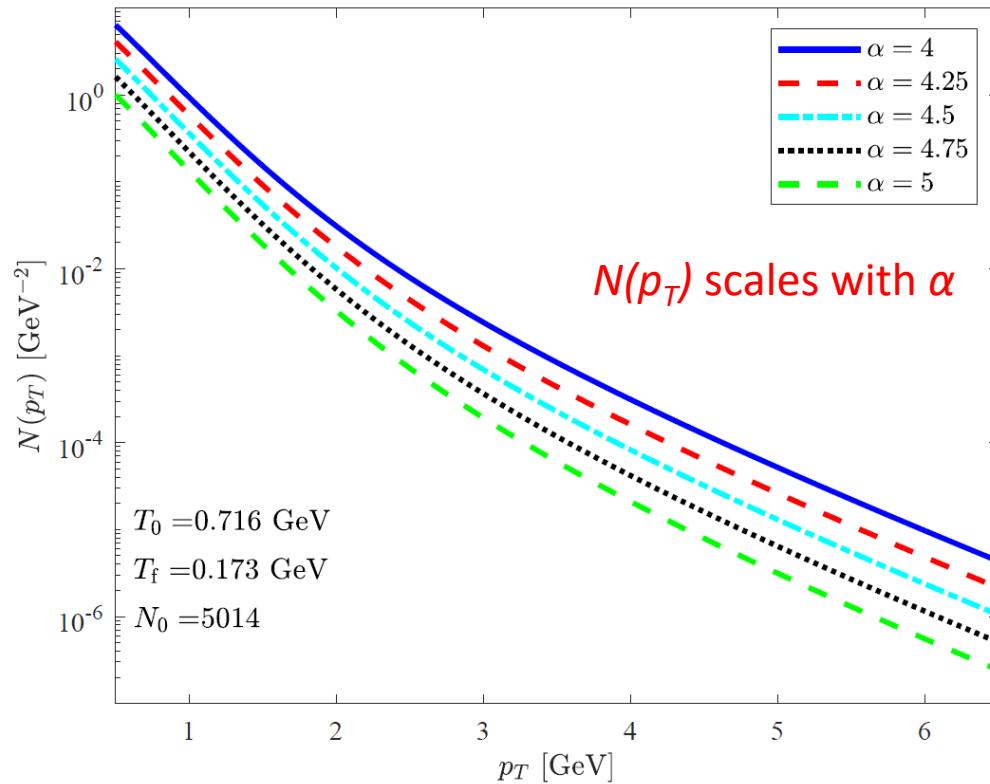
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$$\frac{d^2N}{2\pi p_T dp_T dy} \Big|_{y=0} = N_0 \frac{2\alpha}{3\pi^{3/2}} \left[ \frac{1}{T_f^\alpha} - \frac{1}{T_0^\alpha} \right]^{-1} p_T^{-\alpha-2} \left[ \Gamma \left( \alpha + \frac{5}{2}, \frac{p_T}{T_0} \right) - \Gamma \left( \alpha + \frac{5}{2}, \frac{p_T}{T_f} \right) \right]$$

$$N_0 = \frac{dN}{dy} \Big|_{y=0} = \frac{g A_T}{(2\pi\hbar)^3} \frac{\tau_0}{\tau_R} \frac{T_0^{\alpha+3}}{\alpha} \left[ \frac{1}{T_f^\alpha} - \frac{1}{T_0^\alpha} \right] \frac{3\pi^{3/2}\kappa}{2\lambda} \left( \frac{2\pi\kappa}{\lambda^2(2\kappa-1) - \lambda(\kappa-1)} \right)^{1/2}$$

# The new analytic formula for the thermal radiation

$$\frac{d^2N}{2\pi p_T dp_T dy} \Big|_{y=0} = N_0 \frac{2\alpha}{3\pi^{3/2}} \left[ \frac{1}{T_f^\alpha} - \frac{1}{T_0^\alpha} \right]^{-1} p_T^{-\alpha-2} \left[ \Gamma \left( \alpha + \frac{5}{2}, \frac{p_T}{T_0} \right) - \Gamma \left( \alpha + \frac{5}{2}, \frac{p_T}{T_f} \right) \right]$$

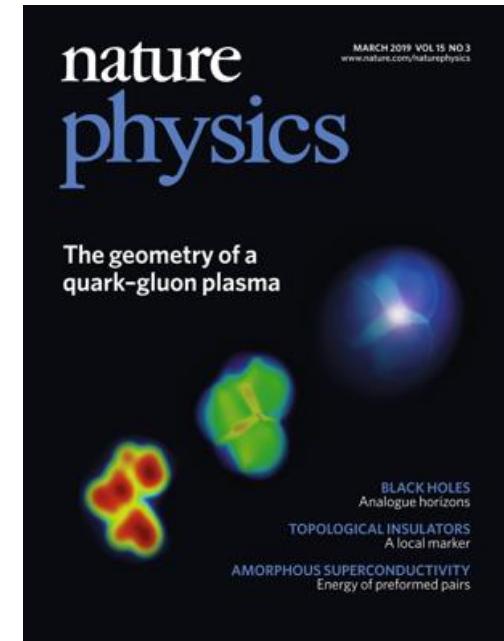


# In conclusion...

- p+p collisions can be described as collective systems
- Our fits indicate low  $c_s$  value ( $\approx 0.35$ )
- Low  $c_s$  value indicate the presence of fluid, so the presence of QGP
- p+p and A+A collisions: self-similar systems

*Is the hydrodynamic description well-accepted?*

- A+A collisions: become a major trend since 2005
- p+A, d+A and He+A collisions: accepted since 2019
- p+p collisions: not widely accepted yet



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*Is the hydrodynamic description well-accepted?*

- A+A collisions: become a major trend since 2005
- p+A, d+A and He+A collisions: accepted since 2019
- p+p collisions: not widely accepted yet
- However, describing H+H systems by hydro is not a recent idea

arXiv:hep-ex/9711009v2 19 Dec 1997



Nijmegen preprint  
HEN-405  
Dec. 97

ESTIMATION OF HYDRODYNAMICAL MODEL PARAMETERS FROM  
THE INVARIANT SPECTRUM AND THE BOSE-EINSTEIN CORRELATIONS OF  
 $\pi^-$  MESONS PRODUCED IN  $(\pi^+/K^+)p$  INTERACTIONS AT 250 GeV/c

EHS/NA22 Collaboration

N.M. Agababyan<sup>e</sup>, M.R. Atayan<sup>e</sup>, T. Csörgő<sup>b</sup>, E.A. De Wolf<sup>a,1</sup>, K. Dziunikowska<sup>b,2</sup>, A.M.F. Endler<sup>e</sup>, Z.Sh. Garutchava<sup>f</sup>, H.R. Gulkanyan<sup>e</sup>, R.Sh. Hakobyan<sup>e</sup>, J.K. Karanyan<sup>e</sup>, D. Kisielewska<sup>b,2</sup>, W. Kittel<sup>d</sup>, S.S. Mehrabyan<sup>e</sup>, Z.V. Metreveli<sup>f</sup>, K. Olkiewicz<sup>b,2</sup>, F.K. Rizatdinova<sup>e</sup>, E.K. Shabalina<sup>c</sup>, L.N. Smirnova<sup>e</sup>, M.D. Tabidze<sup>f</sup>, L.A. Tikhonova<sup>c</sup>, A.V. Tkabladze<sup>f</sup>, A.G. Tomaradze<sup>f</sup>, F. Verbeure<sup>a</sup>, S.A. Zotkin<sup>c</sup>

<sup>a</sup> Department of Physics, Universitaire Instelling Antwerpen, B-2610 Wilrijk, Belgium  
<sup>b</sup> Institute of Physics and Nuclear Techniques of Academy of Mining and Metallurgy and Institute of Nuclear Physics, PL-30055 Krakow, Poland

<sup>c</sup> Nuclear Physics Institute, Moscow State University, RU-119899 Moscow, Russia  
<sup>d</sup> High Energy Physics Institute Nijmegen (HEFIN), University of Nijmegen/NIKHEF, NL-6525 ED Nijmegen, The Netherlands

<sup>e</sup> Centro Brasileiro de Pesquisas Físicas, BR-22290 Rio de Janeiro, Brazil

<sup>f</sup> Institute for High Energy Physics of Tbilisi State University, GE-380086 Tbilisi, Georgia

<sup>g</sup> Institute of Physics, AM-375036 Yerevan, Armenia

<sup>h</sup> KFKI, Hungarian Academy of Sciences, H-1525 Budapest 114, Hungary

**Abstract:** The invariant spectra of  $\pi^-$  mesons produced in  $(\pi^+/K^+)p$  interactions at 250 GeV/c are analysed in the framework of the hydrodynamical model of three-dimensionally expanding cylindrically symmetric finite systems. A satisfactory description of experimental data is achieved. The data favour the pattern according to which the hadron matter undergoes predominantly longitudinal expansion and non-relativistic transverse expansion with mean transverse velocity  $\langle u_t \rangle = 0.20 \pm 0.07$ , and is characterized by a large temperature inhomogeneity in the transverse direction: the extracted freeze-out temperature at the center of the tube and at the transverse rms radius are  $140 \pm 3$  MeV and  $82 \pm 7$  MeV, respectively. The width of the (longitudinal) space-time rapidity distribution of the pion source is found to be  $\Delta\eta = 1.36 \pm 0.02$ . Combining this estimate with results of the Bose-Einstein correlation analysis in the same experiment, one extracts a mean freeze-out time of the source of  $\langle \tau_f \rangle = 1.4 \pm 0.1$  fm/c and its transverse geometrical rms radius,  $R_G(\text{rms}) = 1.2 \pm 0.2$  fm.