Exact Schwinger mechanism for scalar fields

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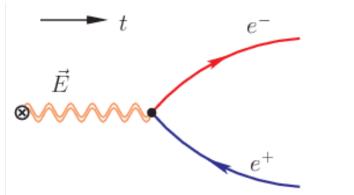
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Introduction

- Schwinger mechanism is a nonperturbative effect where electron-positron pairs are created from the vacuum in strong electric field
- We will examine a similar model with scalar fields coupled to a time independent classical source, which is exactly solveable
- Historically believed to not have scattering in this model, but we find scattering due to the vacuum decay



We model a free real scalar field coupled to a static classical source:

$$\mathcal{L} = rac{1}{2} \partial_\mu \phi \; \partial^\mu \phi - rac{1}{2} m^2 \phi^2 + g
ho(\mathbf{x}) \; \phi.$$

Solutions to e.o.m. can be computed exactly

$$\phi(x) = \phi_0(x) + ig \int_{t_0}^t dy^0 \int d^3 y \ \rho(\mathbf{y}) \int \frac{d^3 p}{(2\pi)^3 2E_{\mathbf{p}}} \left(e^{-ip \cdot (x-y)} - e^{ip \cdot (x-y)} \right)$$

It has the form "free part + Green's function".

Field and Hamiltonian

$$egin{aligned} h(\mathbf{p}) &= g \, rac{-1}{\sqrt{2E_\mathbf{p}^3}} \int d^3 y \, e^{-i\mathbf{p}\cdot\mathbf{y}}
ho(\mathbf{y}), \ f(t,\mathbf{p}) &= g \, rac{e^{iE_\mathbf{p}t}-1}{\sqrt{2E_\mathbf{p}^3}} \int d^3 y \, e^{-i\mathbf{p}\cdot\mathbf{y}}
ho(\mathbf{y}), \end{aligned}$$

The field can be written in the compact way

$$\phi(\mathbf{x}) = \int \frac{d^3p}{(2\pi)^3 \sqrt{2E_{\mathbf{p}}}} \left[\left(a_{\mathbf{p}} + f(t, \mathbf{p}) \right) e^{-ip \cdot \mathbf{x}} + \left(a_{\mathbf{p}}^{\dagger} + f^*(t, \mathbf{p}) \right) e^{ip \cdot \mathbf{x}} \right]$$

We can see the time dependent ladder operators $a_{\mathbf{p}}(t) = a_{\mathbf{p}} + f(t, \mathbf{p})$

$${\cal H} = \int rac{d^3 p}{(2\pi)^3} \, {\cal E}_{f p} \, \left[\left(a^\dagger_{f p}(0) + h^*({f p})
ight) \left(a_{f p}(0) + h({f p})
ight) - |h({f p})|^2
ight] \; .$$

Different ladder operators in the Hamiltonian and field. No mutual eigenstate of particle number and Hamiltonian exists (for all times).

Vacuum decay probability

Initial vacuum $|0\rangle$ gains a nonzero particle number expectation value after time *t*

$$\mathcal{N}_0(t) = \int rac{d^3 p}{(2\pi)^3} \langle 0 | a^\dagger_{f p}(t) a_{f p}(t) | 0
angle = 2 \int rac{d^3 p}{(2\pi)^3} \left(1 - \cos(E_{f p} t)
ight) | h(f p) |^2 \, .$$

Eigenstates of the Hamiltonian form a basis so we can insert them with the identity in the expression

$$P_{0\rightarrow n}(t) = \int \frac{d^3 p_1 \cdots d^3 p_n}{(2\pi)^{3n} 2E_{\mathbf{p}_1} \cdots 2E_{\mathbf{p}_n}} |\langle \mathbf{p}_1, \cdots, \mathbf{p}_n| U(t,0) |0\rangle|^2$$

Result is a Poisson distribution with time dependent particle number expectation value

$$P_{0\to n}(t) = \frac{e^{-N_0(t)}}{n!} \left[N_0(t)\right]^n$$

Small and large time behaviour

Small time behaviour slower than exponential decay.

$$\mathcal{P}_{0
ightarrow 0}(t) = \exp\left\{-2\int rac{d^3q}{(2\pi)^3}\left(1-\cos(E_{\mathbf{q}}t)
ight)\left|h(\mathbf{q})
ight|^2
ight\}\sim e^{-\lambda t^2}$$

With N repeated measurements, survival prob. after T = Nt

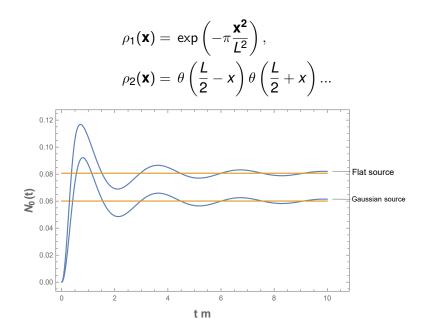
$$P_{0 \to 0}(T) = (P_{0 \to 0}(t))^N \approx \exp\left[-N\lambda t^2\right] = \exp\left[-\frac{\lambda}{N}T^2\right] \xrightarrow{N \to \infty} 1$$

Zeno effect: initial state is preserved by repeated measurement for as long as we want

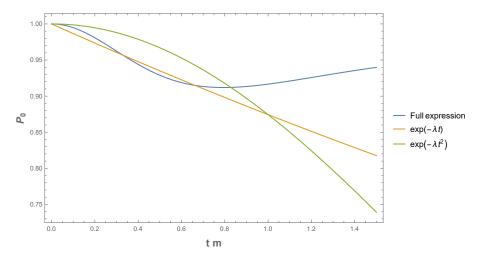
For large time, the oscillation term dies out (Riemann-Lebesgue):

$$\lim_{t\to\infty}N_0(t)=2\int\frac{d^3p}{(2\pi)^3}\left|h(\mathbf{p})\right|^2<\infty.$$

Numerical examples



Small time behaviour II



 Vacuum survival probability and some approximations for the Gaussian shaped source The particle density distribution is well-defined for large volume

$$\lim_{L \to \infty} \lim_{t \to \infty} \frac{1}{L^3} N_0(t) = 2 \frac{g^2}{m^3} \qquad \text{(flat source)},$$
$$\lim_{L \to \infty} \lim_{t \to \infty} \frac{1}{L^3} N_0(t) = \frac{1}{\sqrt{2}} \frac{g^2}{m^3} \qquad \text{(Gaussian source)}.$$

Different prefactor depending on shape; the tails matter (somewhat)

One particle Scattering

With initial condition

$$|\psi_1
angle = \int rac{{\mathcal O}^3 k}{(2\pi)^3 2 E_{f k}} \; \psi_1({f k}) \; |{f k}
angle,$$

and the n-particle partial wave of the time evolving vacuum

$$\psi_{0}(t;\mathbf{p}_{1},\cdots,\mathbf{p}_{n})=\langle\mathbf{p}_{1},\cdots,\mathbf{p}_{n}|U(t,0)|0\rangle$$

Vacuum decay causes dynamics in scattering

$$\begin{aligned} \langle \mathbf{p}_{1}\cdots\mathbf{p}_{n}| \ U(t,0) |\psi_{1}\rangle &= -\psi_{0}(t;\mathbf{p}_{1}\cdots\mathbf{p}_{n}) \int \frac{d^{3}k}{(2\pi)^{3}\sqrt{2E_{\mathbf{k}}}} \ f^{*}(t,\mathbf{k}) \ \psi_{1}(\mathbf{k}) \\ &+ \frac{1}{\sqrt{n}} \sum_{j=1}^{n} \left[\psi_{1}(\mathbf{p}_{j}) \ e^{-iE_{\mathbf{p}_{j}}t} \ \psi_{0}(t;\mathbf{p}_{1},\cdots\mathbf{p}_{j-1},\mathbf{p}_{j+1},\cdots,\mathbf{p}_{n}) \right]. \end{aligned}$$

First term: absorption and vacuum decay into *n* particles, Second term: free propagation and vacuum decay into n - 1 particles

Conclusion

- Contrary to what was thought historically, we find scattering of which a key component is the vacuum decay
- Past arguments rely on the adiabatic switching off. This changes the algebra of the operators, even if done at infinity, which gives different results than the exact calculation
- The vacuum decay is conceptually similar to the Schwinger mechanism. The exact calculations show qualitative differences to the expected result, such as a saturation of particle number, non-exponential decay and Zeno effect
- Vacuum decay might be a relevant non-perturbative production mechanism of the dilaton, whose production is suppressed at the tree level

Thank you for your attention!