

# MODified Characteristics of Hadronic Interactions

## in ultra-high-energy cosmic-ray showers



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# Modified hadronic interactions

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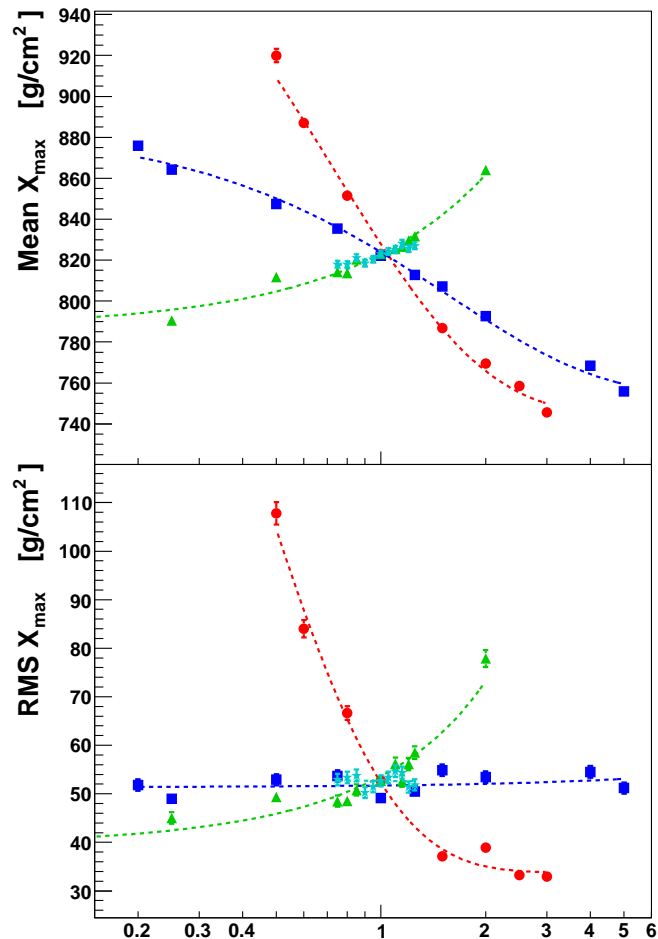
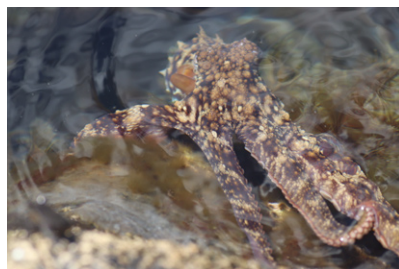
- individual changes of multiplicity, elasticity and cross-section in CONEX - 1D simulations
- 215 citations

$$f(E, f_{19}) = 1 + (f_{19} - 1) \cdot \frac{\log_{10}(E/E_{\text{thr}})}{\log_{10}(10 \text{ EeV}/E_{\text{thr}})}$$

CONEX in Corsika: 3D information

MOCHI: CORSIKA 7.741 with CONEX option, Sibyll 2.3d

- nuclear projectiles treated as a set of p-Air interactions
- POS(ICRC2023)245
- POS(ICRC2021)441
- EPJ WoC 283:05005



# “Allowed” modifications and thresholds

Cross-section ( $E_{\text{thr}} = 10^{16}$  eV)

- well constrained for p-p at LHC to a few %
- unc. in conversion to p-A limited by CMS p-Pb measurement

Multiplicity ( $E_{\text{thr}} = 10^{15}$  eV)

- no p-A data, limited rapidity coverage

Elasticity ( $E_{\text{thr}} = 10^{14}$  eV)

- difficult at accelerators, limits from nuclear emulsion chambers
- recent LHCf neutron elasticity measurement?
- range of modifications limited by internal consistency



0.8

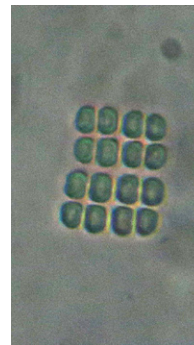
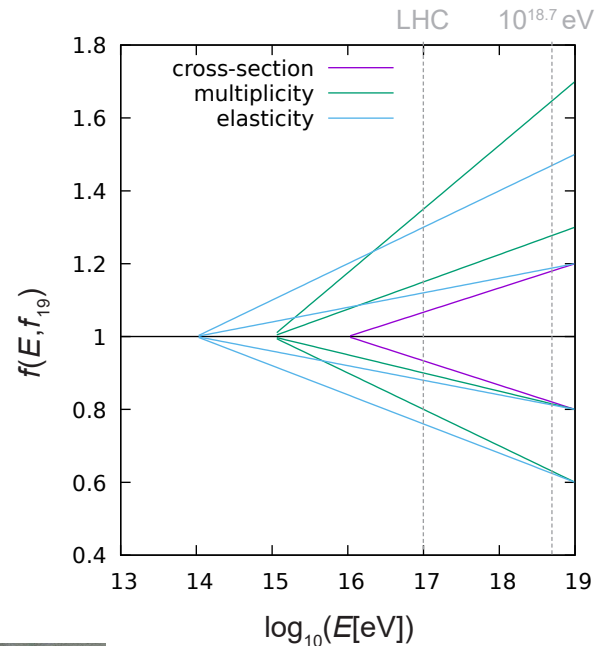
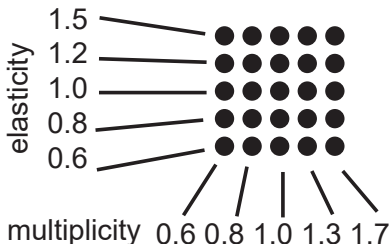


1.0

cross-section



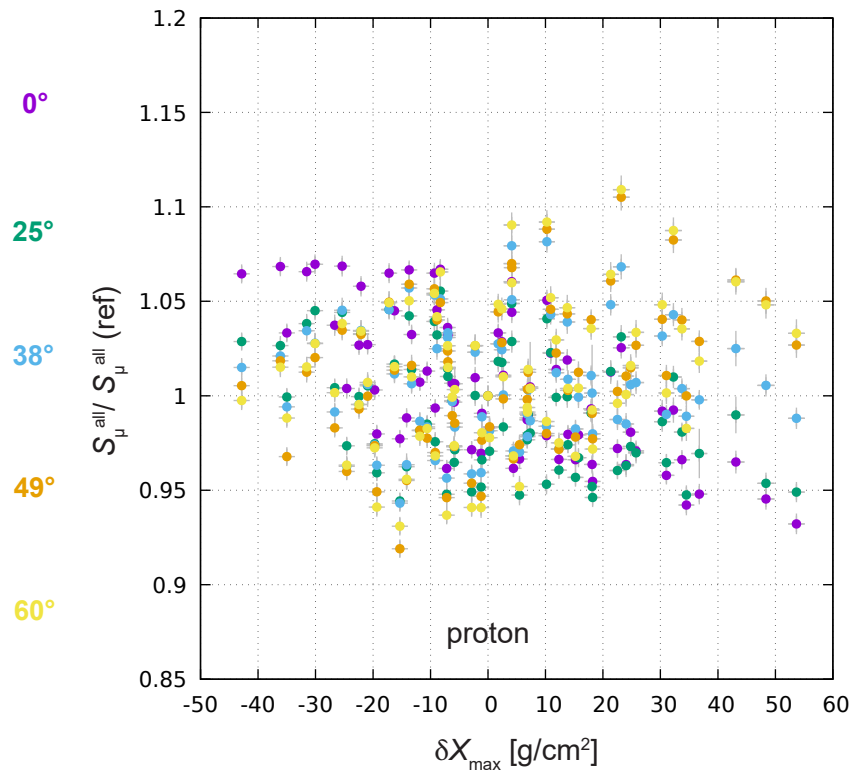
1.2



- energy  $10^{18.7}$  eV
- proton and iron
- 5 zenith angles
- 1000 showers per „bin“
- 750 000 showers

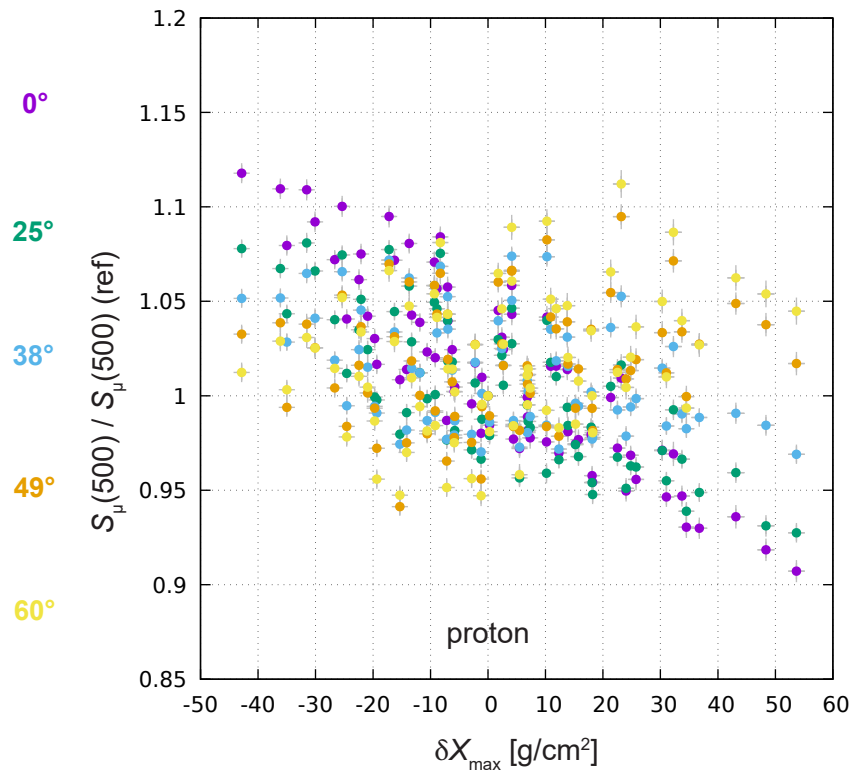
# Number of muons vs. $X_{\max}$ for all muons

More dependent on  $r$  for vertical showers, less for inclined



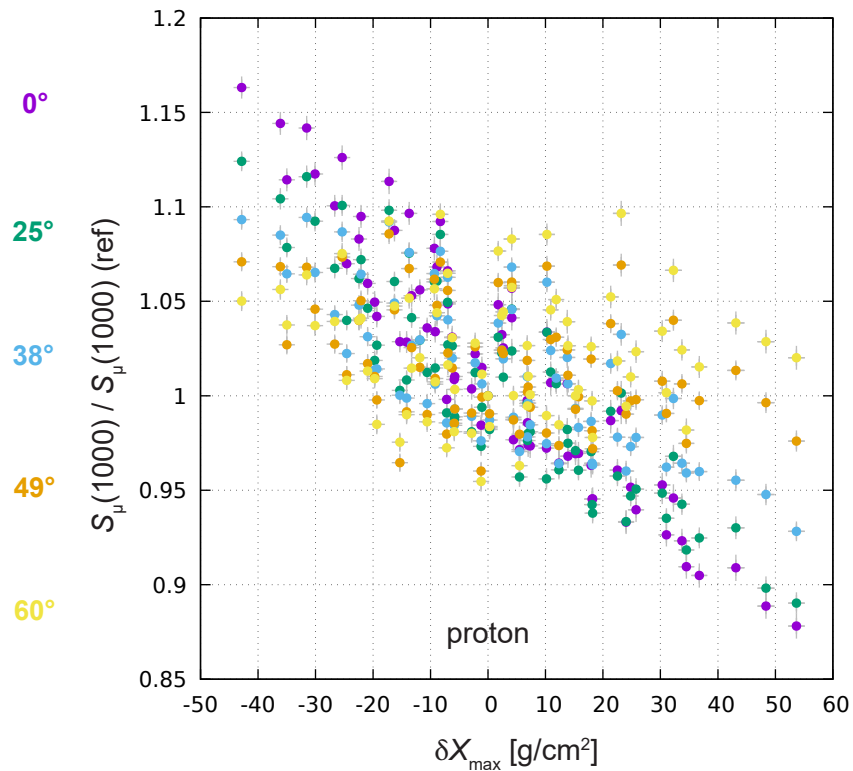
# Number of muons vs. $X_{\max}$ @ 500 meters

More dependent on  $r$  for vertical showers, less for inclined



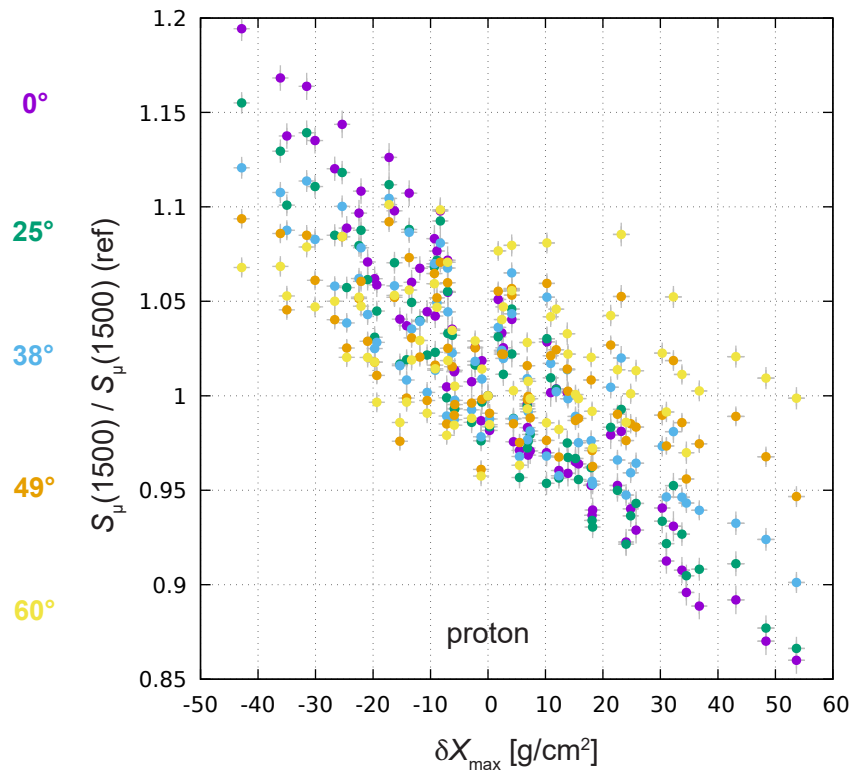
# Number of muons vs. $X_{\max}$ @ 1000 meters

More dependent on  $r$  for vertical showers, less for inclined



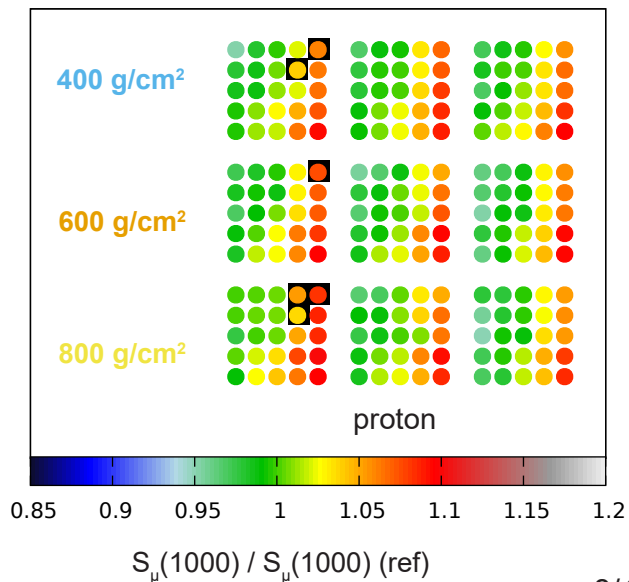
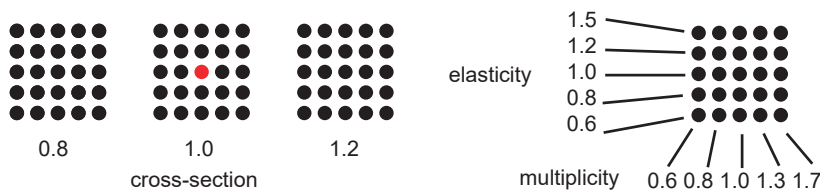
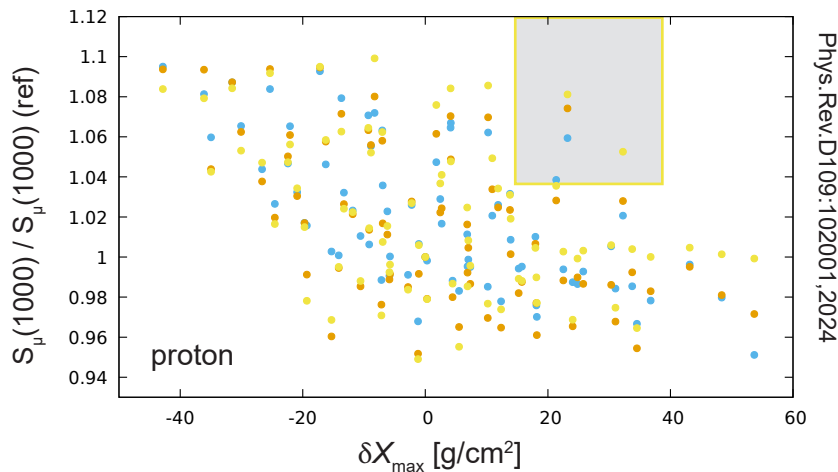
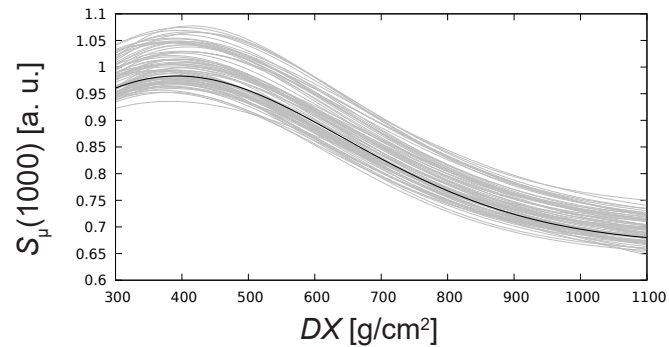
# Number of muons vs. $X_{\max}$ @ 1500 meters

More dependent on  $r$  for vertical showers, less for inclined



# Muons at 1000 m at fixed $DX$

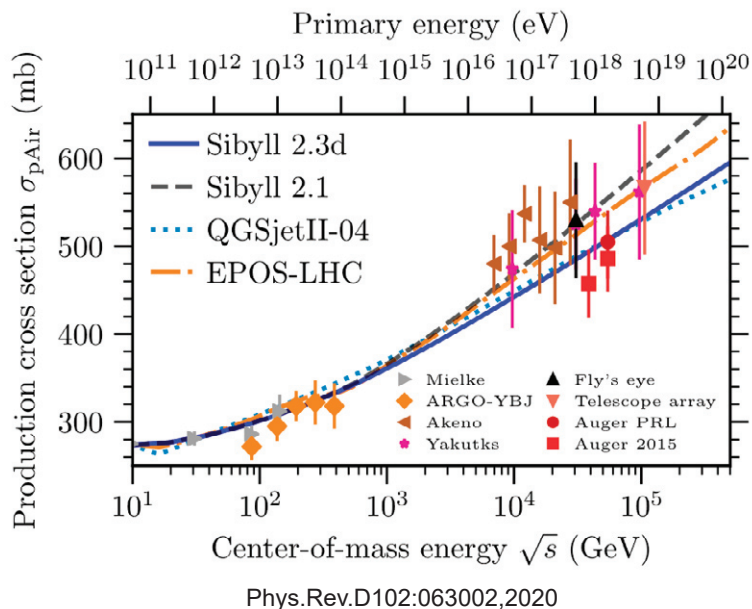
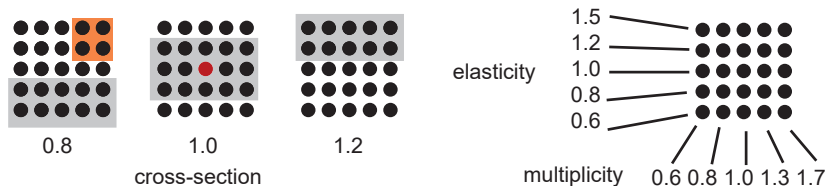
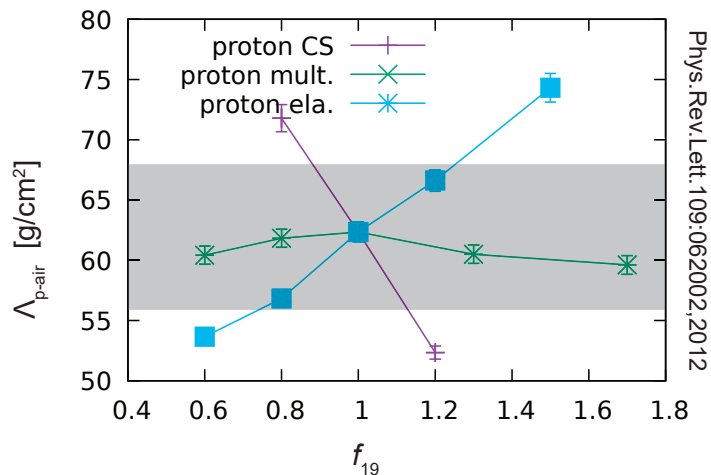
- remove effects of shifting  $X_{\max}$  on  $S_{\mu}$  by fitting a dependence on  $DX$  - allows comparison recent with Auger analysis
- note: Auger has universal correction to ground signal



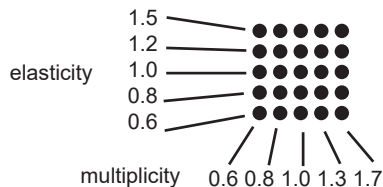
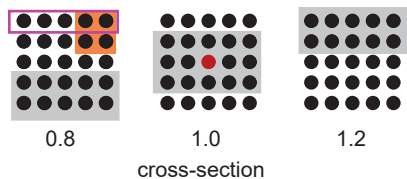
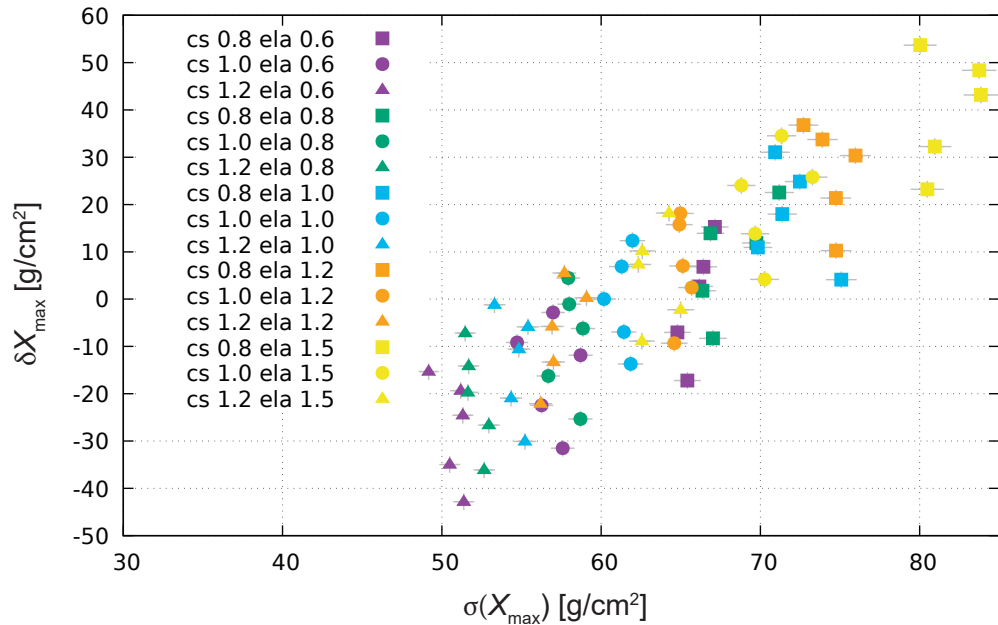
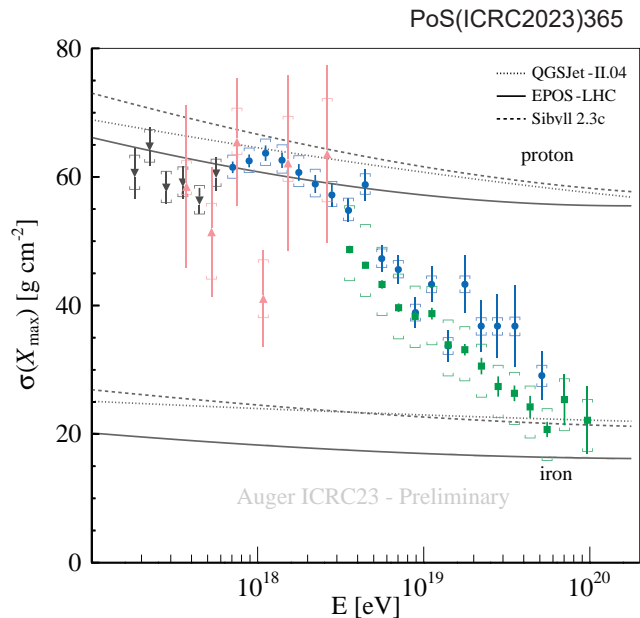


# Auger cross-section measurement

- modifications of elasticity change  $\Lambda \rightarrow \sigma$  conversion: Auger CS = constraint in  $\sigma$ -elasticity space
- unmodified Sibyll 2.3d smack on data; uncertainty extrapolation with  $f(E, f_{19})$



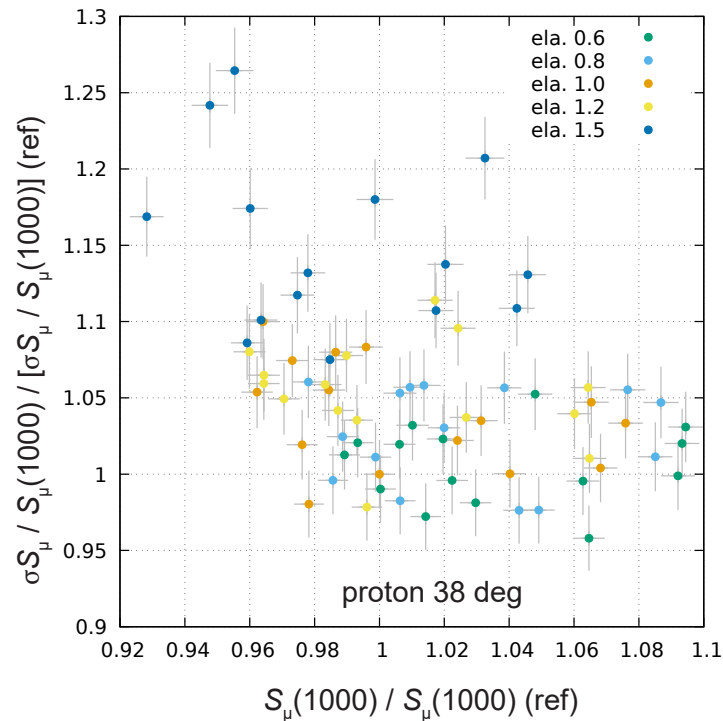
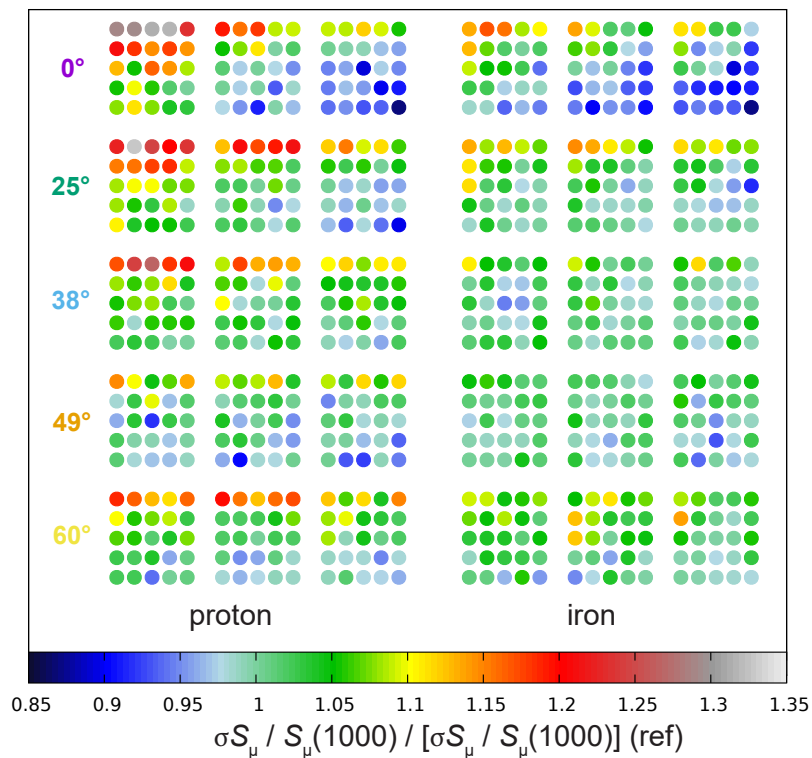
# $X_{\max}$ fluctuations



Low CS and high elasticity leads to very high  $X_{\max}$  fluctuations that may be difficult to reconcile with Auger data  
 n.b. no change of fluctuations for Fe

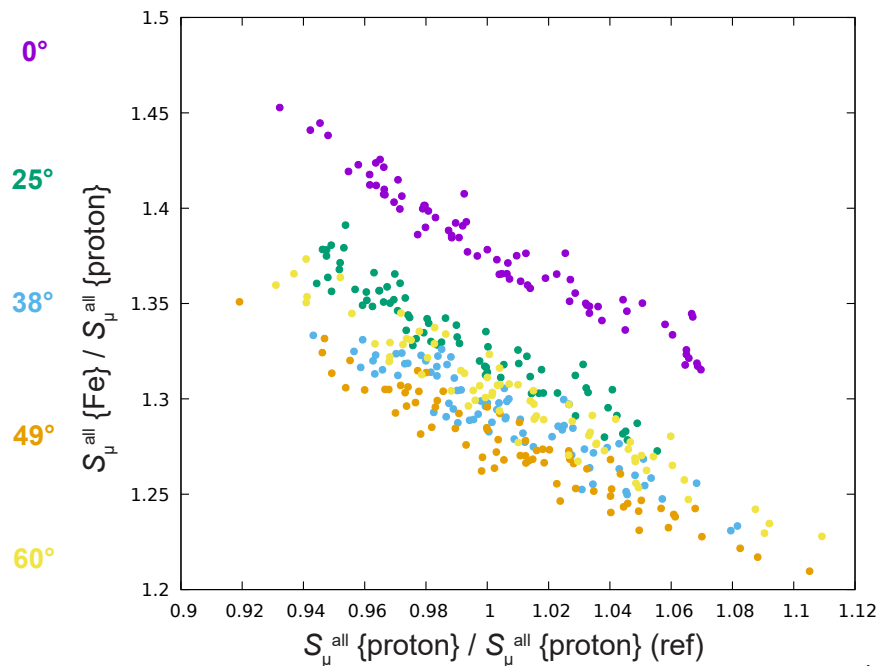
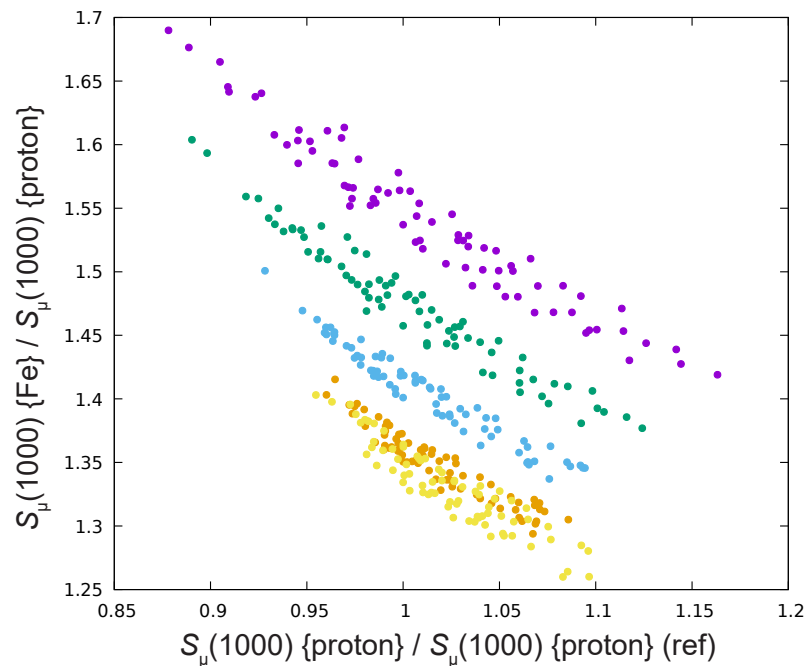
# Ground particles: relative muon number fluctuations at 1000 meters

- not correlated with absolute changes in muon number, sensitive to high elasticity changes



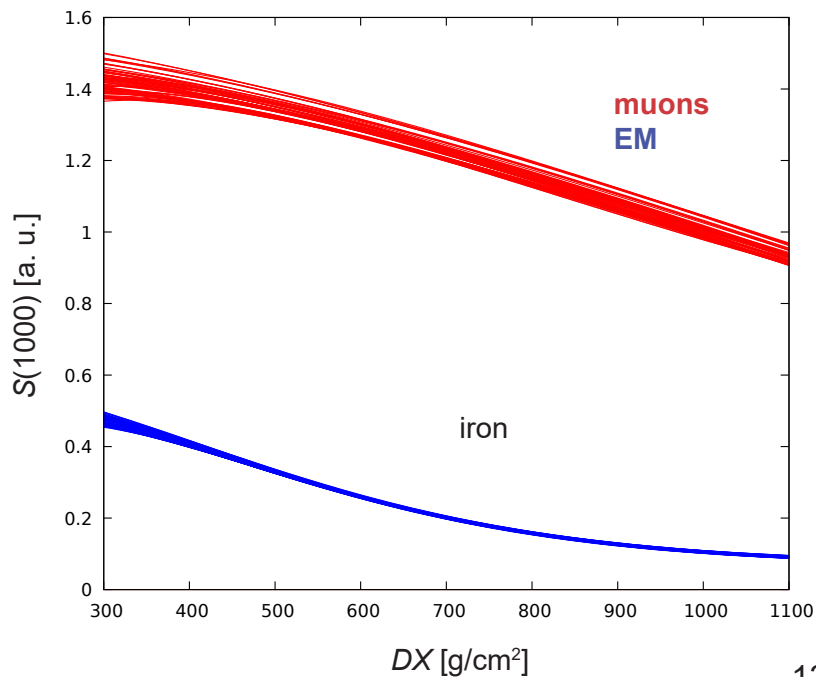
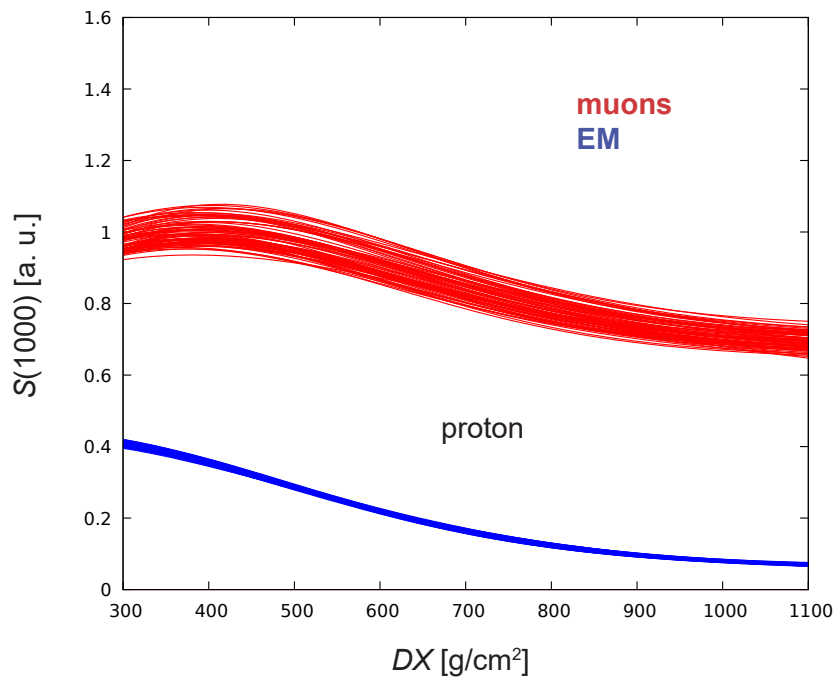
## Adding muons and proton/iron separation

Ratio between number of muons for iron and proton tends down when muons are added  
- whatever the answer to the muon problem is, it may make primary separation more difficult



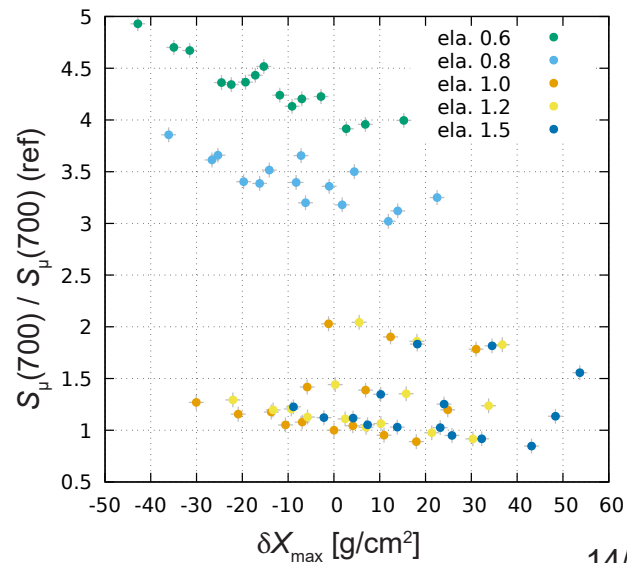
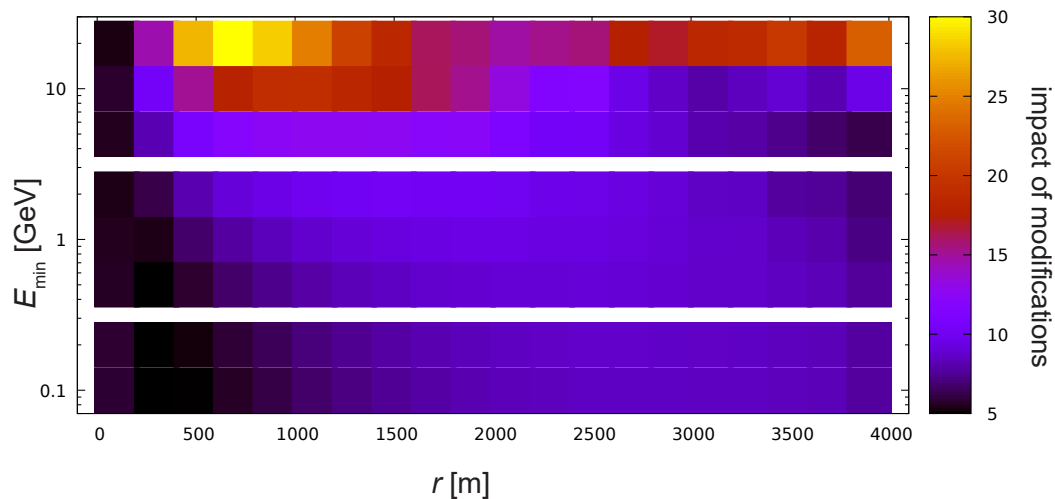
# Universality

Muon signal far more affected than EM (also lateral shape of EM well preserved)  
(notes: no complete 4-component model, arbitrary normalization between muon/EM)



# Sensitivity of muon number to modification as a function of $E_{\min}$ , $r$

- sum of absolute values of changes of muon density divided by statistical uncertainty (1000 showers)
- example: proton @ 38 degrees
- large deviations in the most significant point in  $(E_{\min}, r)$  space overwhelmingly due to low elasticity bins
  - deep underground measurements highly interesting for particle physics!



## Conclusions

- changing cross-section, elasticity and multiplicity within reasonable limits can have major impact on air-shower properties
- the impact can be quite different for quantities depending on 3D geometry as opposed to 1D sums
- the changes of hadronic interactions indicated by the Pierre Auger Observatory are just reachable
  - but only with a *combination* of modifications!
  - and already in a tension with other measurements
- a wealth of other features can be studied - see POS(ICRC2023)245 (full papers soon)
- even if some modifications are not realistic, we can learn interesting insights