

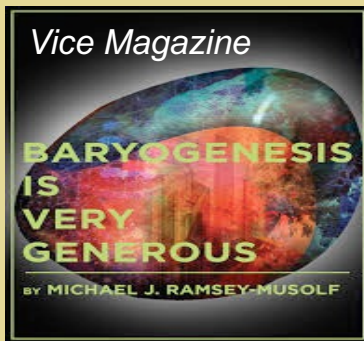
Was There an Electroweak Phase Transition?

M.J. Ramsey-Musolf

- *T.D. Lee Institute/Shanghai Jiao Tong Univ.*
- *UMass Amherst*
- *Caltech*

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- mjrm@umass.edu
- 微信 : mjrm-china
- <https://michaelramseymusolf.com/>

About MJRM:



Science



Family



Friends

My pronouns: he/him/his
MeToo

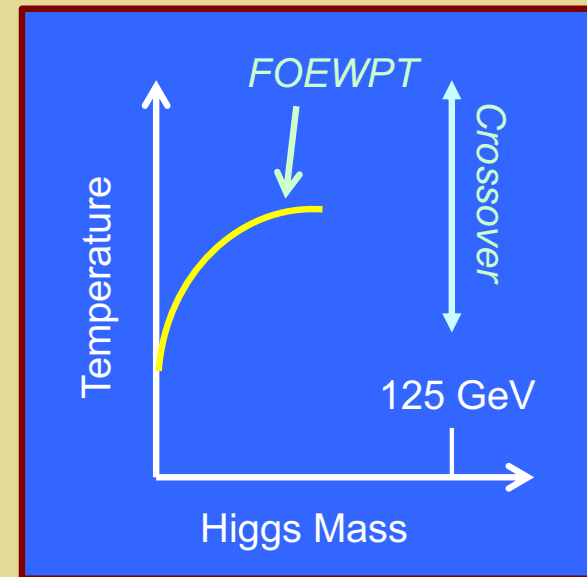
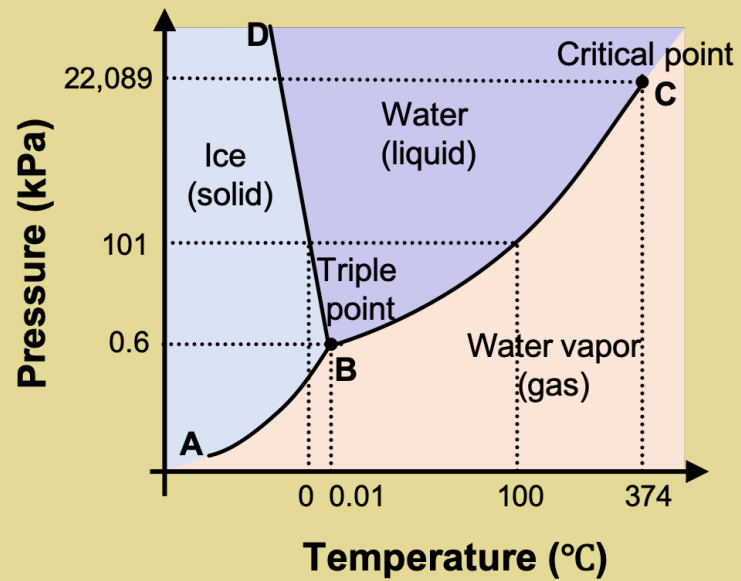
SUSY 2024 Madrid
June 13, 2024

I. Context & Questions

Was There an Electroweak Phase Transition ?

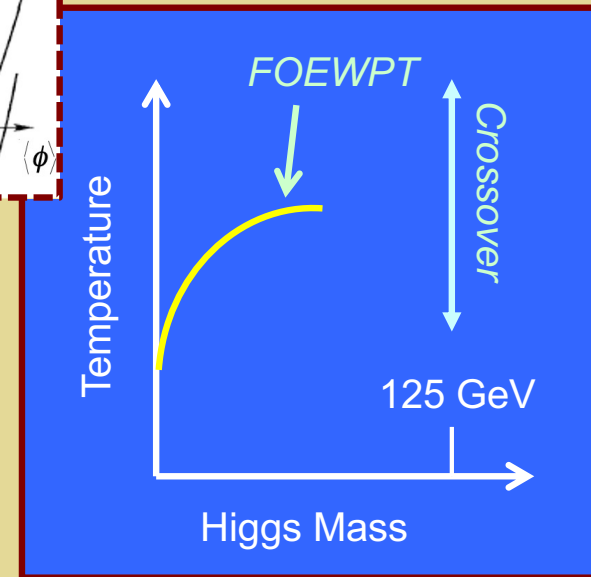
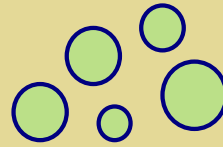
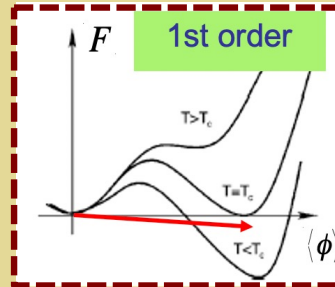
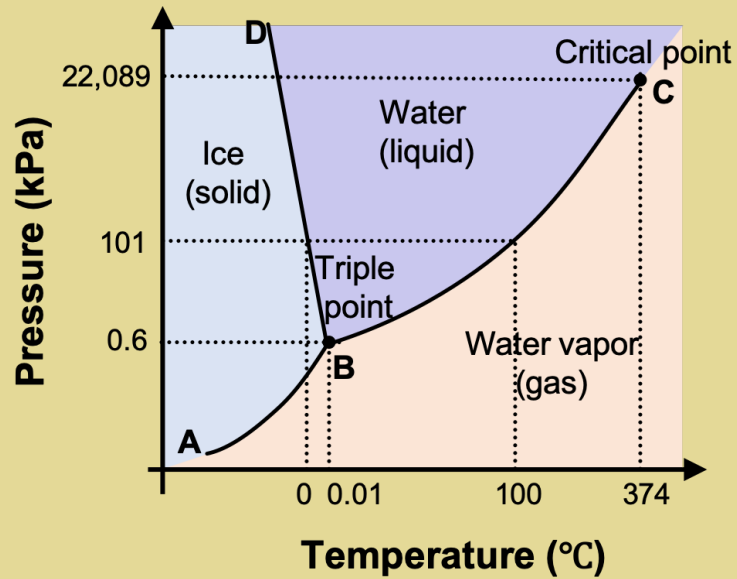
- ***Interesting in its own right***
- ***Key ingredient for EW baryogenesis***
- ***Source of gravitational radiation***

Was There an Electroweak Phase Transition ?



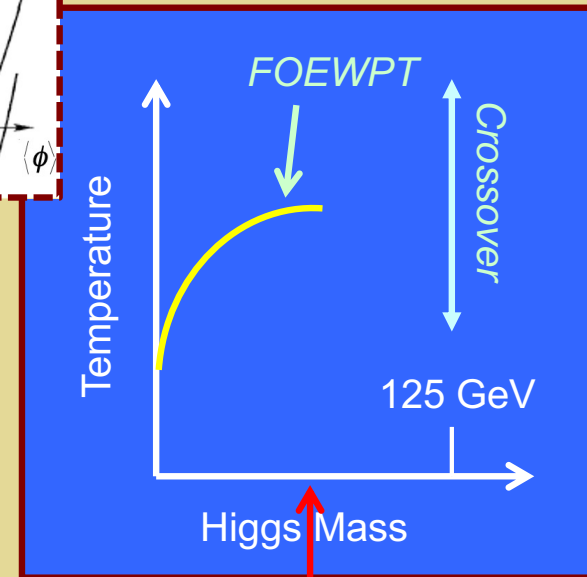
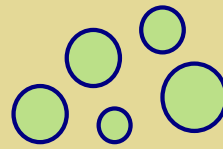
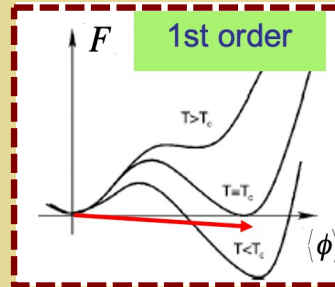
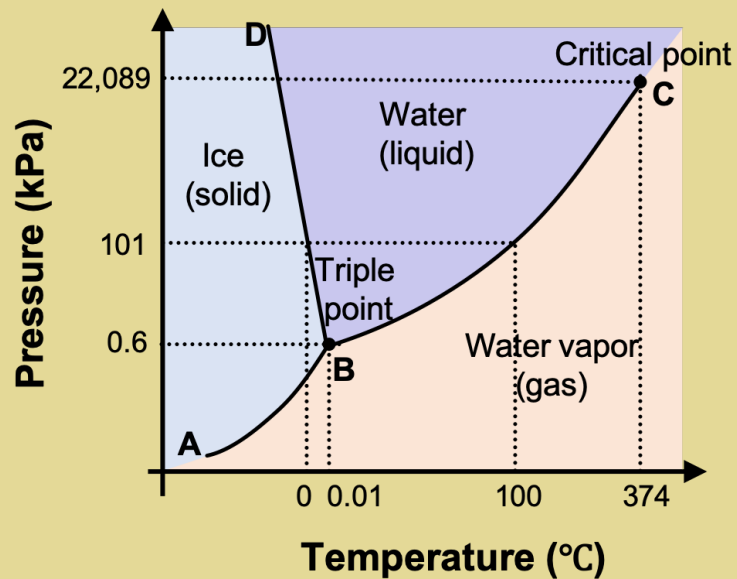
EW Phase Diagram

Was There an Electroweak Phase Transition ?



EW Phase Diagram

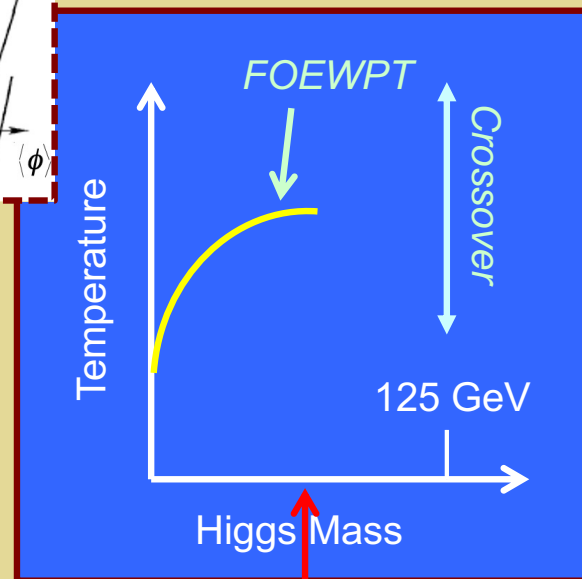
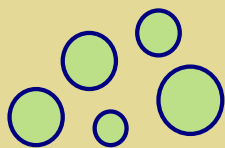
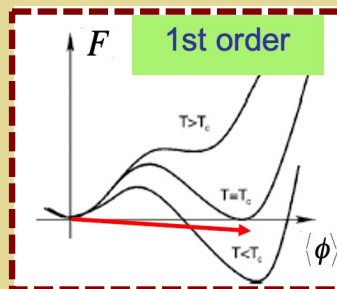
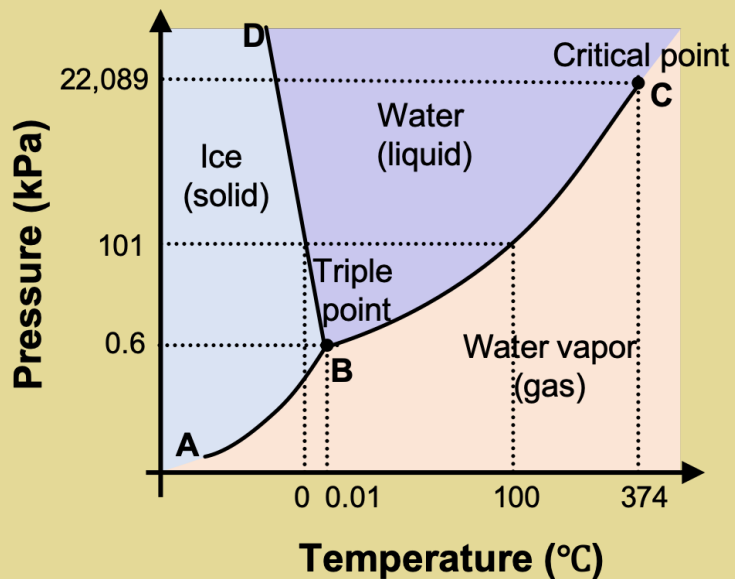
Was There an Electroweak Phase Transition ?



EW Phase Diagram

| Lattice | Authors | M_h^C (GeV) |
|----------------|---------|----------------|
| 4D Isotropic | [76] | 80 ± 7 |
| 4D Anisotropic | [74] | 72.4 ± 1.7 |
| 3D Isotropic | [72] | 72.3 ± 0.7 |
| 3D Isotropic | [70] | 72.4 ± 0.9 |

Was There an Electroweak Phase Transition ?

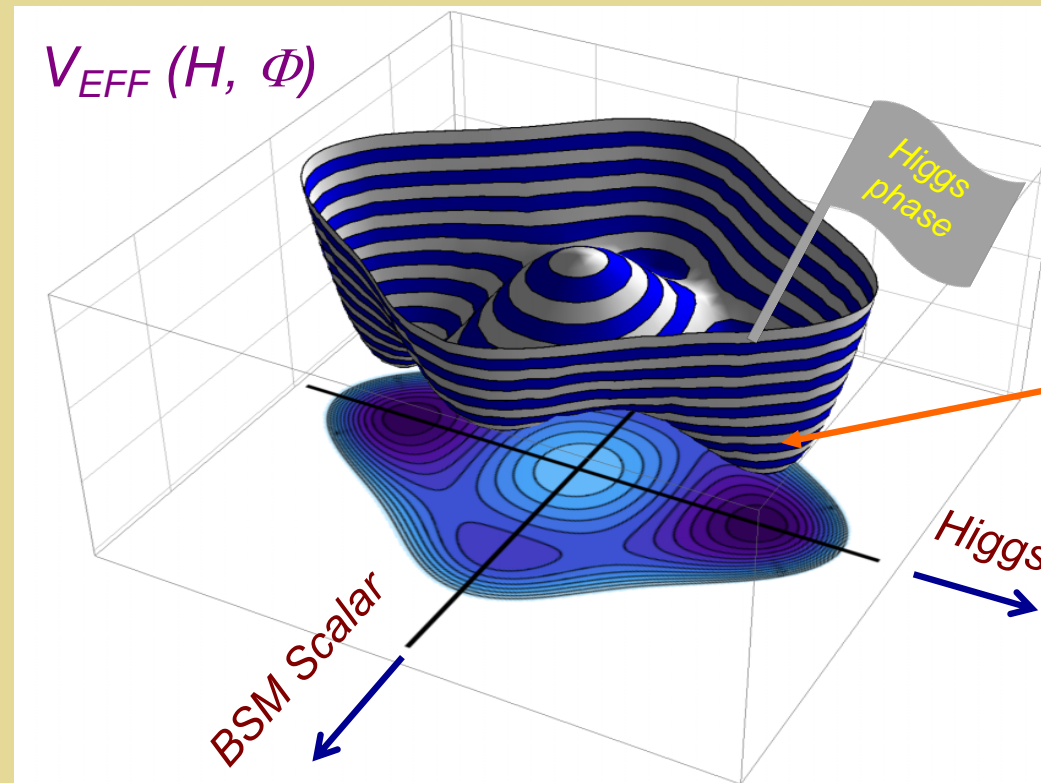


EW Phase Diagram

How does this picture change in presence of new TeV scale physics ? What is the phase diagram ? SFOEWPT ?

| Lattice | Authors | M_h^C (GeV) |
|----------------|---------|----------------|
| 4D Isotropic | [76] | 80 ± 7 |
| 4D Anisotropic | [74] | 72.4 ± 1.7 |
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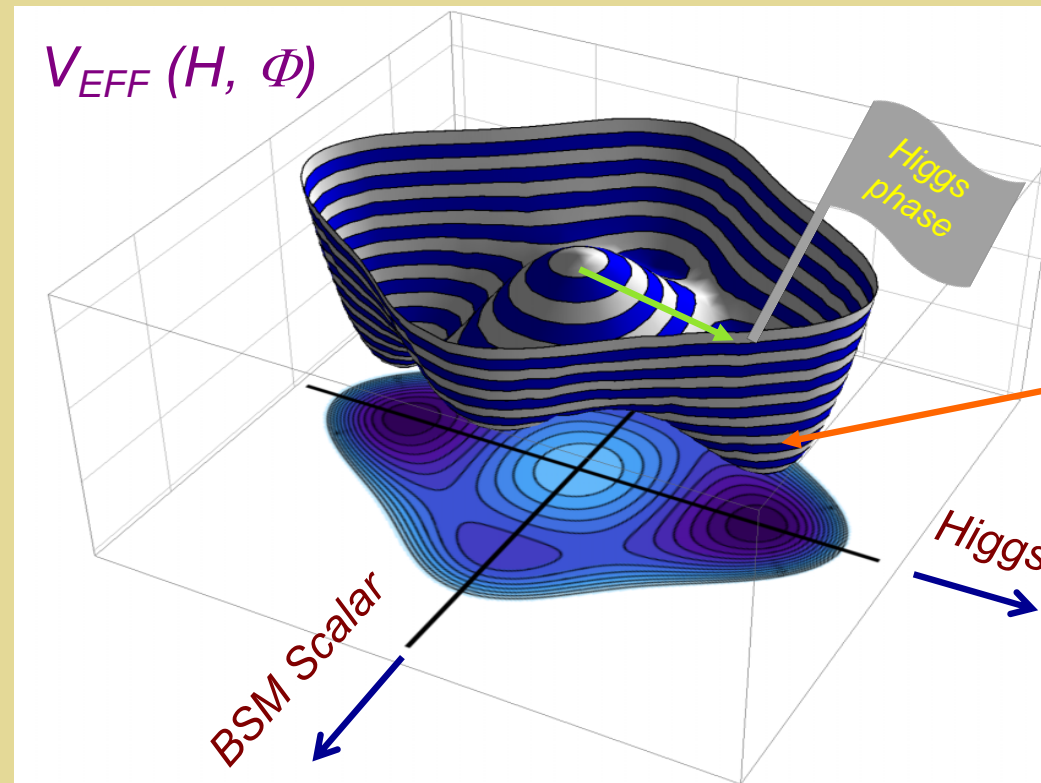
What Was the EWSB Thermal History ?



How did we end up here ?

Extrema can evolve differently as T evolves \rightarrow rich possibilities for symmetry breaking

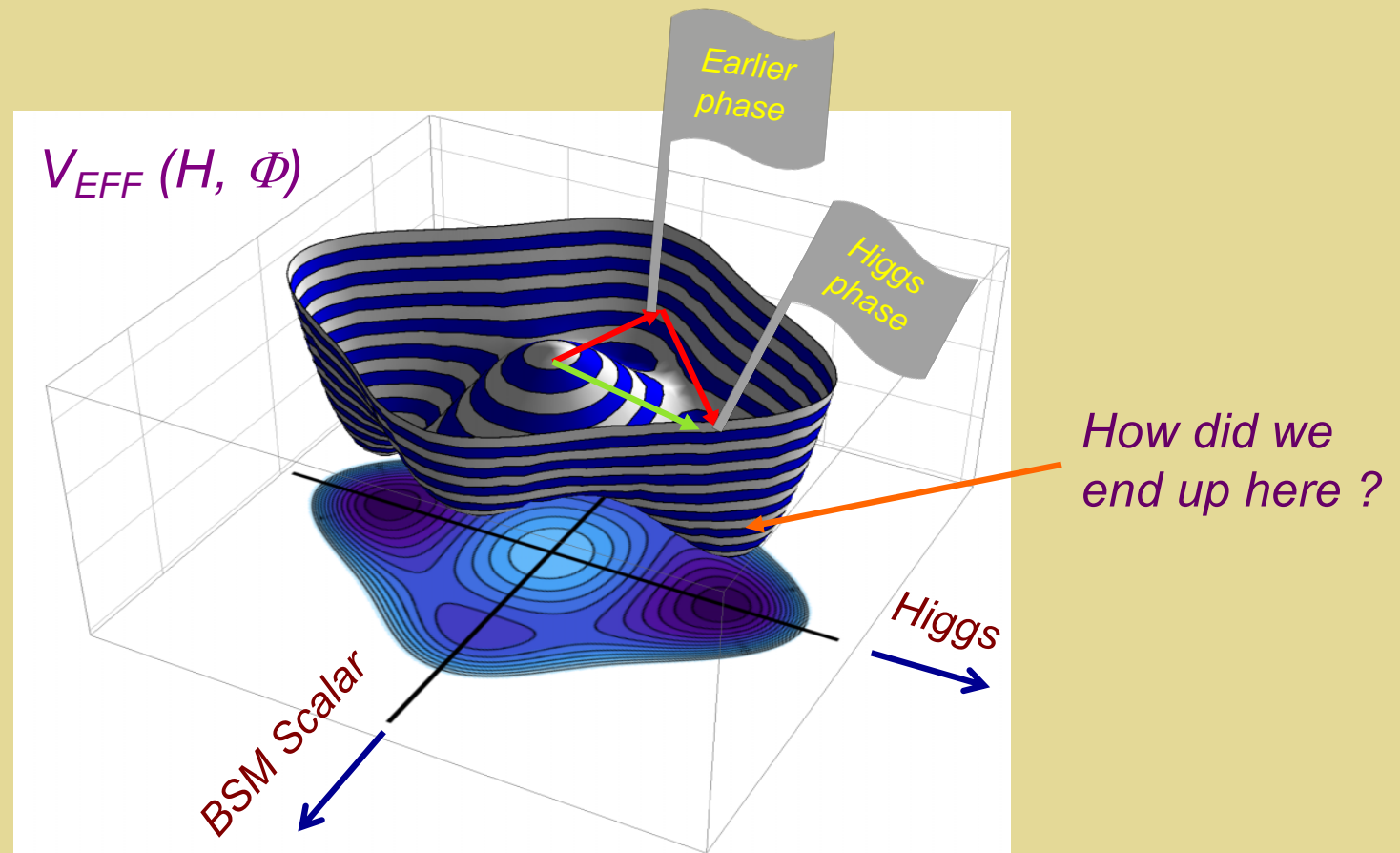
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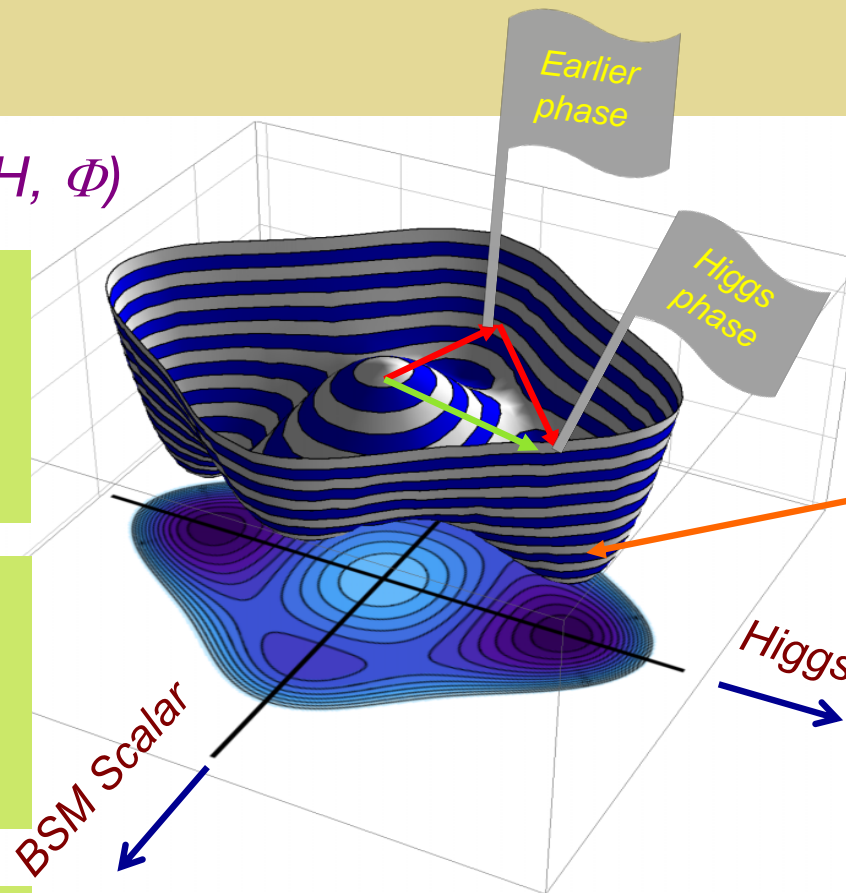
What Was the EWSB Thermal History ?



**Extrema can evolve differently as T evolves \rightarrow
rich possibilities for symmetry breaking**

What Was the EWSB Thermal History ?

$$V_{\text{EFF}}(H, \Phi)$$



- What is the landscape of potentials and their thermal histories?

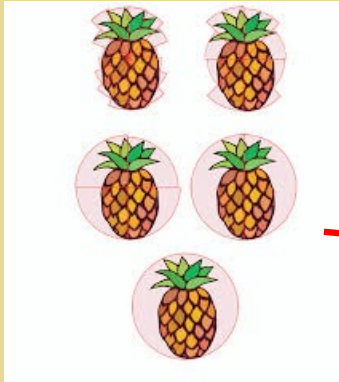
- How can we probe this $T > 0$ landscape experimentally?

- How reliably can we compute the thermodynamics?

**n evolve differently as T evolves \rightarrow
abilities for symmetry breaking**

Was There an EW Phase Transition?

Bubble Collisions

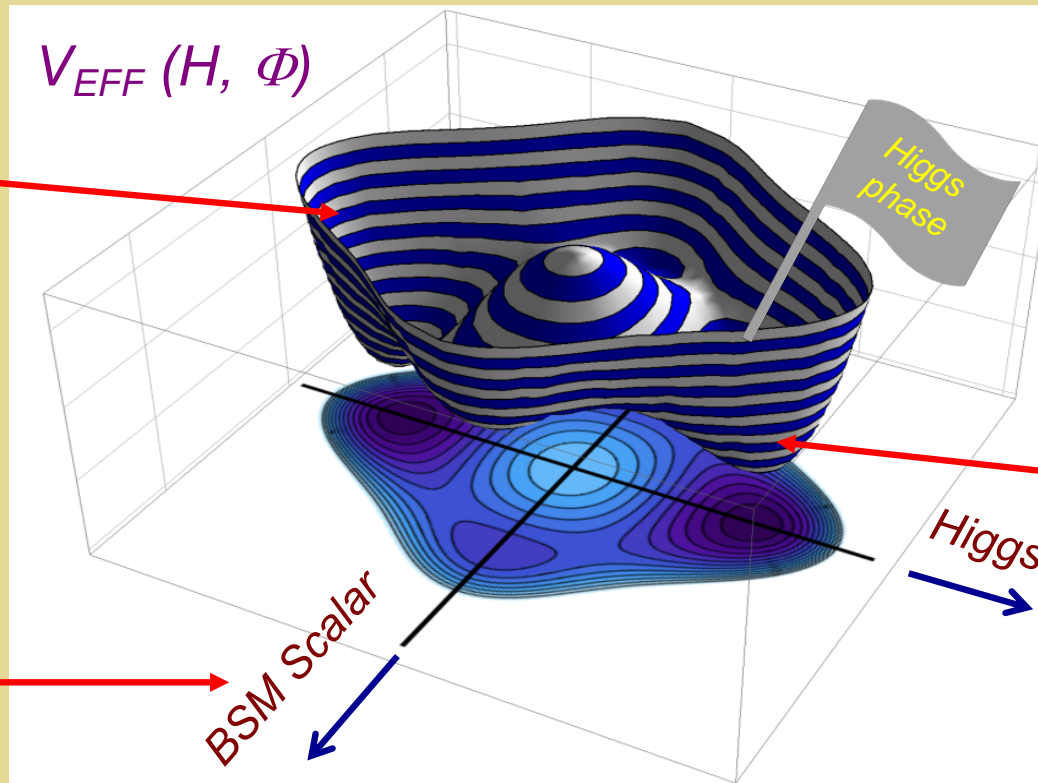


Grav Radiation

Direct Production



BSM Higgs



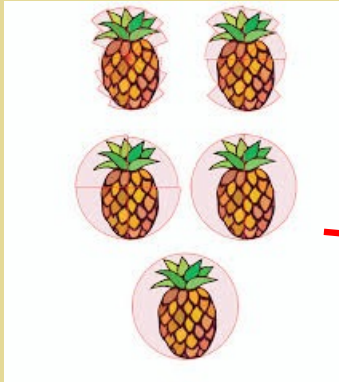
Higgs precision tests



Extrema can evolve differently as T evolves \rightarrow rich possibilities for symmetry breaking

Was There an EW Phase Transition?

Bubble Collisions

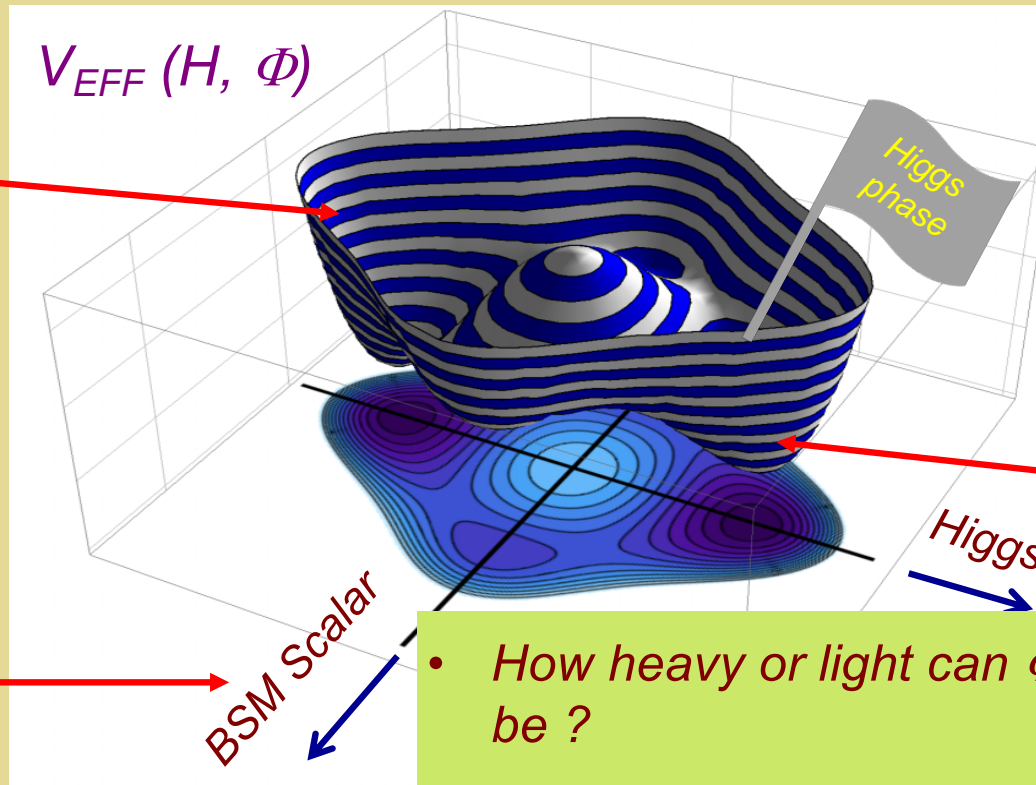


Grav Radiation

Direct Production



BSM Higgs



Higgs precision tests



SM Higgs BSM Higgs

- How heavy or light can Φ be ?
- How coupled to H ?
- Can it be discovered with colliders & GW probes ?

Extrema can evolve
rich possibilities for



$T_{EW} \rightarrow$ Scale for Colliders & GW probes

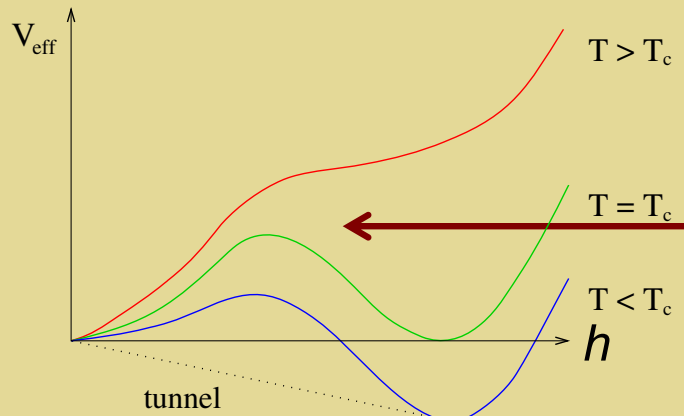
High- T SM Effective Potential

$$V(h, T)_{\text{SM}} = D(T^2 - T_0^2) h^2 + \lambda h^4 + \dots$$

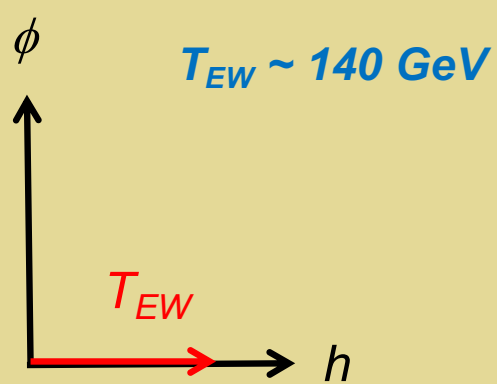
$$T_0 \sim 140 \text{ GeV}$$

$$\equiv T_{EW}$$

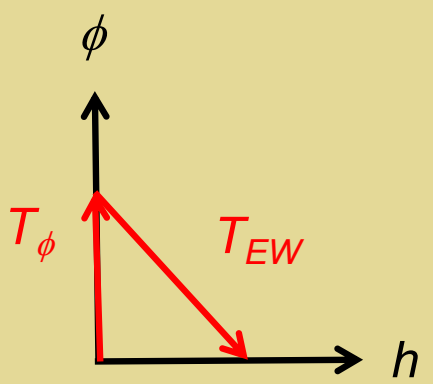
First Order EWPT from BSM Physics



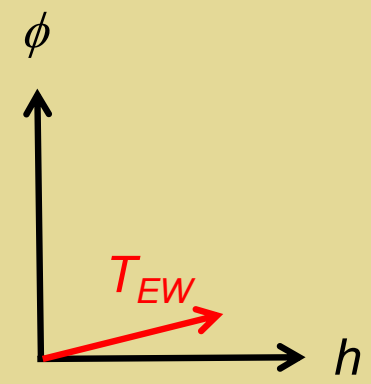
Representative thermal histories \rightarrow barrier for SFOEWPT



$a_2 H^2 \phi^2 : T > 0$
loop effect

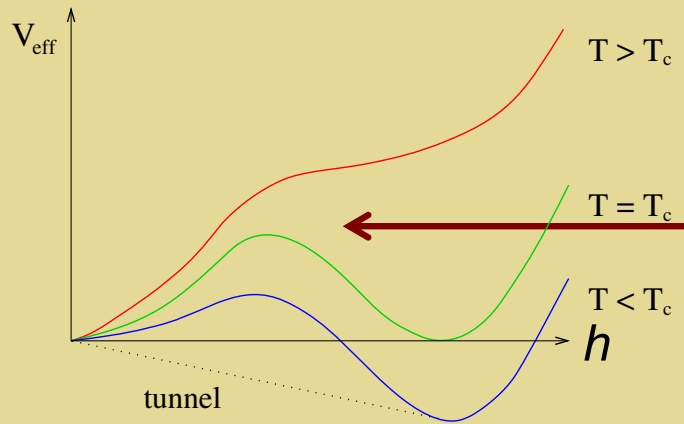


$a_2 H^2 \phi^2 : T = 0$
tree-level effect

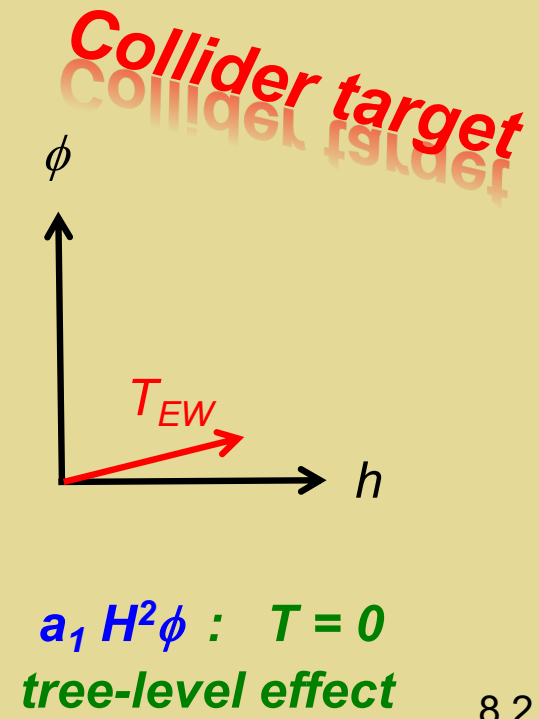
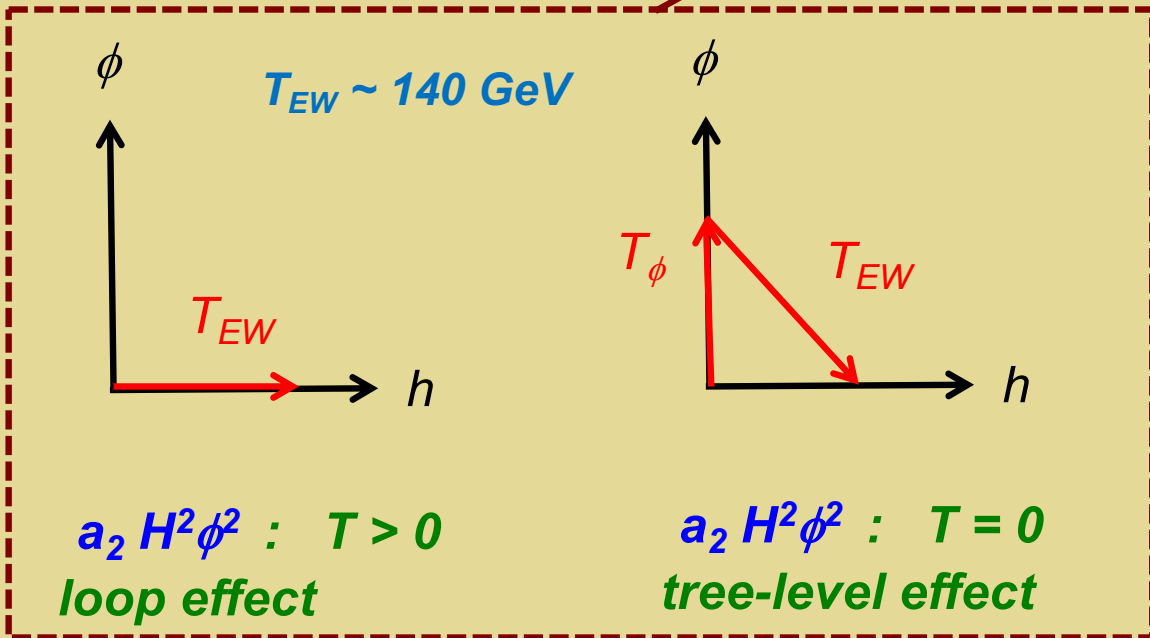


$a_1 H^2 \phi : T = 0$
tree-level effect

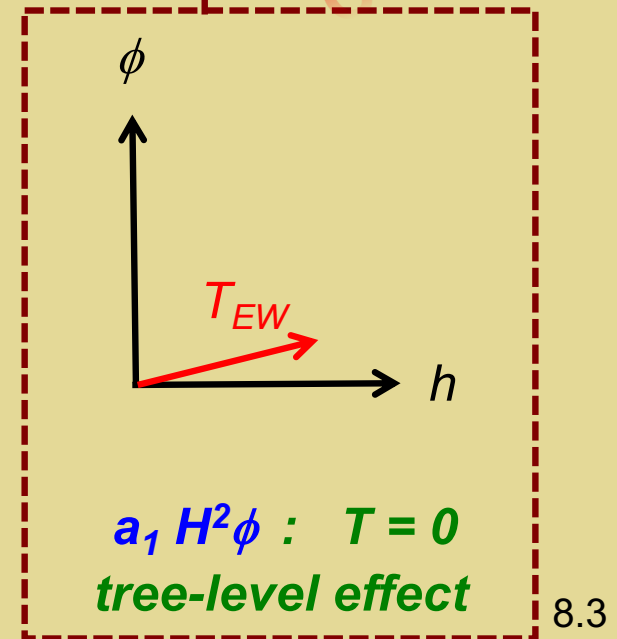
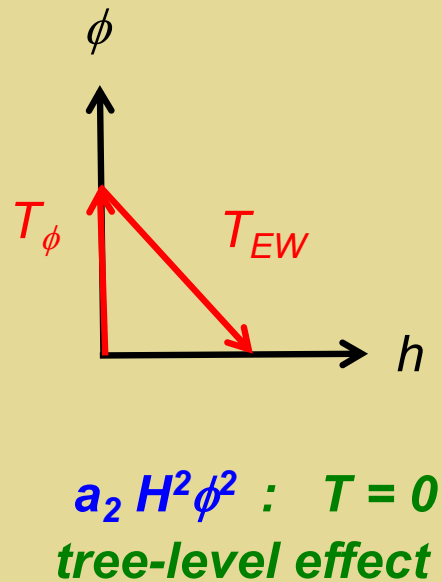
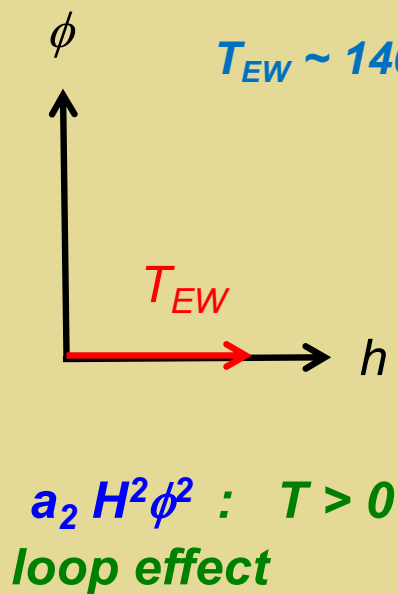
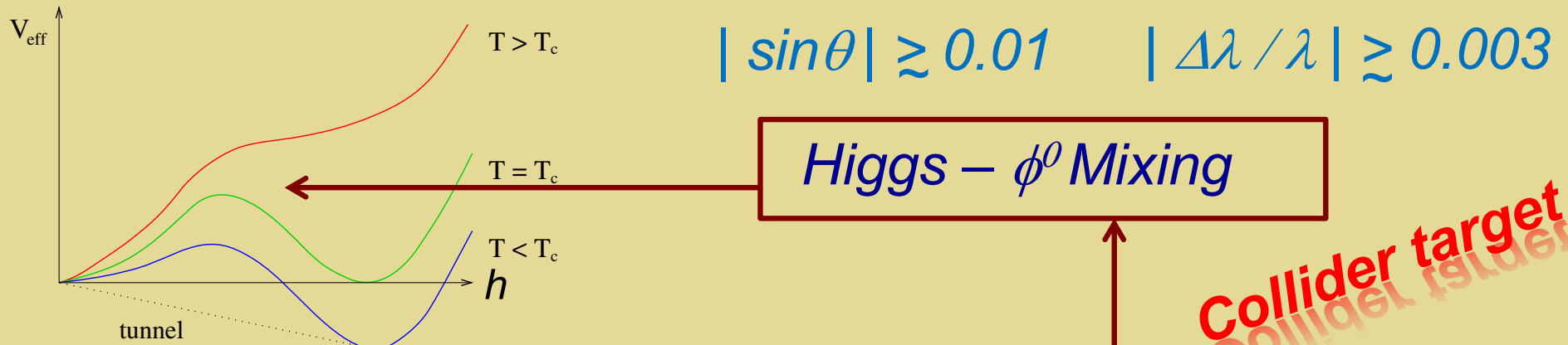
First Order EWPT from BSM Physics



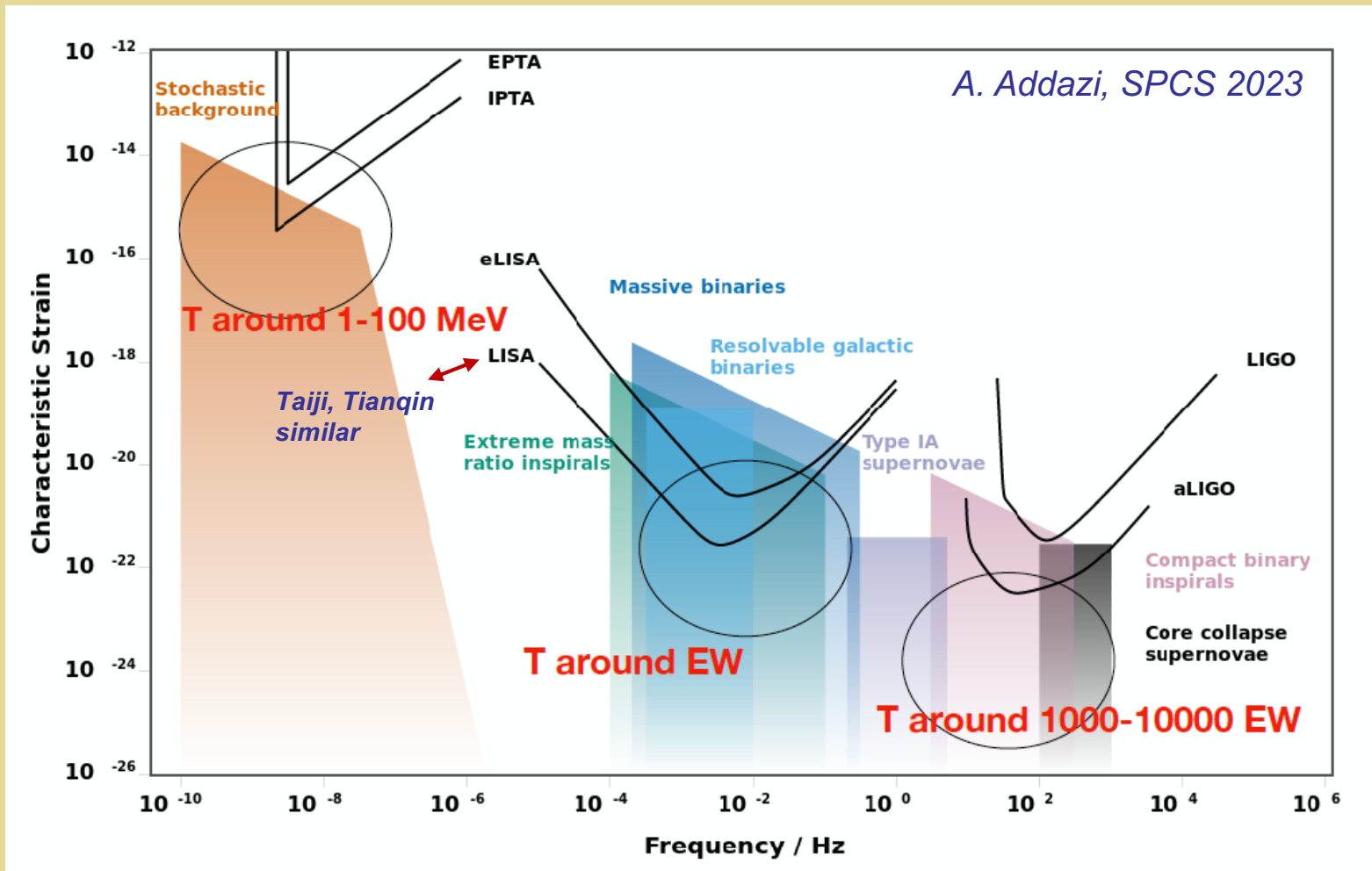
Simple arguments: $T_{EW} +$
 first order EWPT \rightarrow
 $M_\phi \lesssim 700 \text{ GeV}$



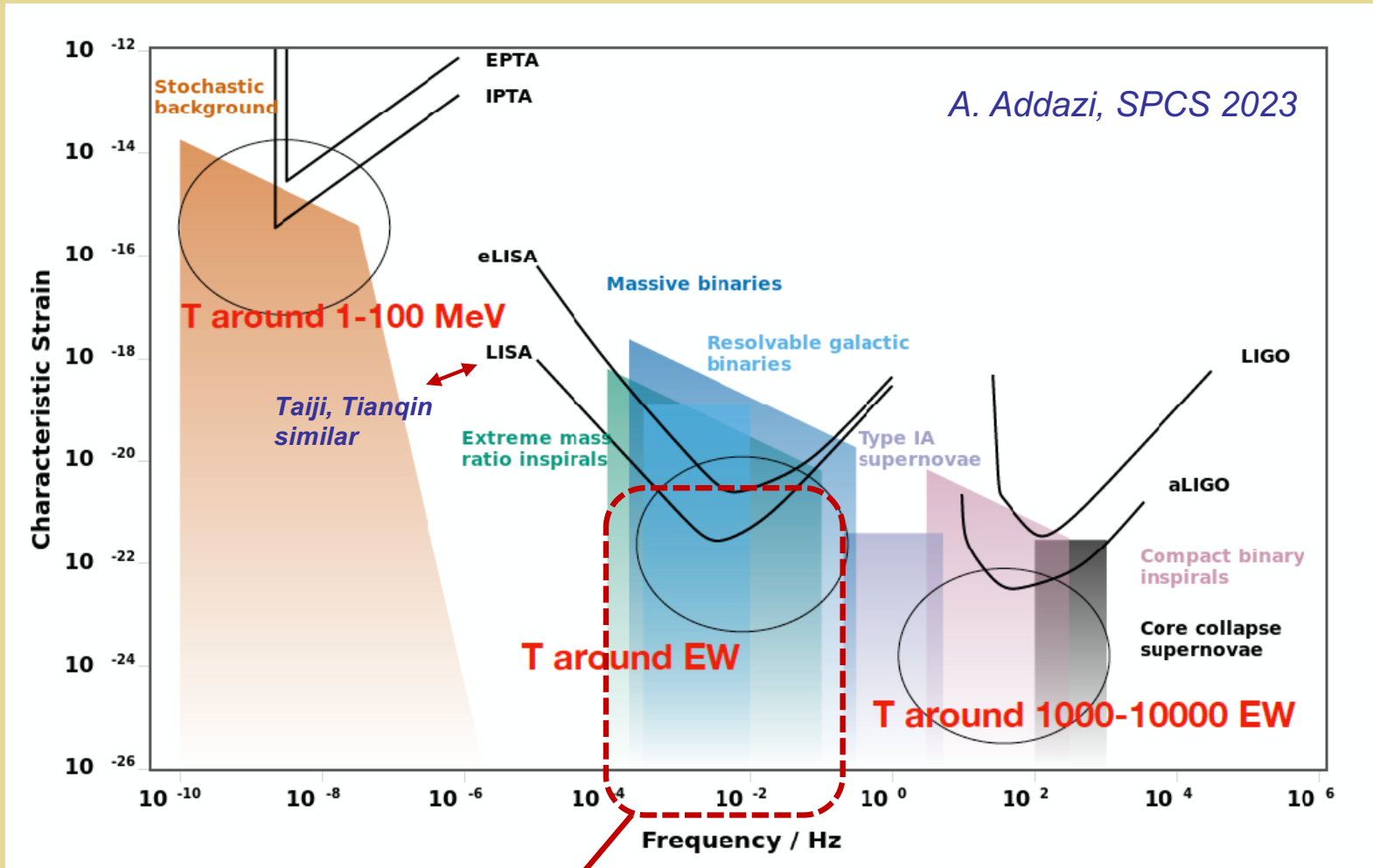
First Order EWPT from BSM Physics



Gravitational Waves



Gravitational Waves



EWPT laboratory for GW micro-physics: colliders can probe particle physics responsible for non-astro GW sources → test our framework for GW microphysics at other scales

II. Theory-Pheno Interface

***Theoretical developments →
phenomenological implications***

Models & Phenomenology

What BSM Scenarios?

SM + Scalar Singlet

Espinosa, Quiros 93, Benson 93, Choi, Volkas 93, Vergara 96, Branco, Delepine, Emmanuel-Costa, Gonzalez 98, Ham, Jeong, Oh 04, Ahriche 07, Espinosa, Quiros 07, Profumo, Ramsey-Musolf, Shaughnessy 07, Noble, Perelstein 07, Espinosa, Konstandin, No, Quiros 08, Barger, Langacker, McCaskey, Ramsey-Musolf, Shaughnessy 09, Ashoorioon, Konstandin 09, Das, Fox, Kumar, Weiner 09, Espinosa, Konstandin, Riva 11, Chung, Long 11, Barger, Chung, Long, Wang 12, Huang, Shu, Zhang 12, Fairbairn, Hogan 13, Katz, Perelstein 14, Profumo, Ramsey-Musolf, Wainwright, Winslow 14, Jiang, Bian, Huang, Shu 15, Kozaczuk 15, Cline, Kainulainen, Tucker-Smith 17, Kurup, Perelstein 17, Chen, Kozaczuk, Lewis 17, Gould, Kozaczuk, Niemi, Ramsey-Musolf, Tenkanen, Weir 19...

SM + Scalar Doublet
(2HDM)

Turok, Zadrozny 92, Davies, Froggatt, Jenkins, Moorhouse 94, Cline, Lemieux 97, Huber 06, Froome, Huber, Seniuch 06, Cline, Kainulainen, Trott 11, Dorsch, Huber, No 13, Dorsch, Huber, Mimasu, No 14, Basler, Krause, Muhlleitner, Wittbrodt, Wlotzka 16, Dorsch, Huber, Mimasu, No 17, Bernon, Bian, Jiang 17, Andersen, Gorda, Helset, Niemi, Tenkanen, Tranberg, Vuorinen, Weir 18...

SM + Scalar Triplet

Patel, Ramsey-Musolf 12, Niemi, Patel, Ramsey-Musolf, Tenkanen, Weir 18 ...

MSSM

Carena, Quiros, Wagner 96, Delepine, Gerard, Gonzalez Felipe, Weyers 96, Cline, Kainulainen 96, Laine, Rummukainen 98, Carena, Nardini, Quiros, Wagner 09, Cohen, Morrissey, Pierce 12, Curtin, Jaiswal, Meade 12, Carena, Nardini, Quiros, Wagner 13, Katz, Perelstein, Ramsey-Musolf, Winslow 14...

NMSSM...

Pietroni 93, Davies, Froggatt, Moorhouse 95, Huber, Schmidt 01, Ham, Oh, Kim, Yoo, Son 04, Menon, Morrissey, Wagner 04, Funakubo, Tao, Yokoda 05, Huber, Konstandin, Prokopec, Schmidt 07, Chung, Long 10, Kozaczuk, Profumo, Stephenson Haskins, Wainwright 15...

Models & Phenomenology

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Theory Meets Phenomenology

Perturbation theory

- *I.R. problem: poor convergence*
- *Thermal resummations*
- *Gauge Invariance (radiative barriers)*
- *RG invariance at $T > 0$*

BSM proposals



Non-perturbative (I.R.)

- *Computationally and labor intensive*

Theory Meets Phenomenology

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BSM proposals

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Challenges for Theory

Perturbation theory

Non-perturbative (I.R.)

- *I.R. problem: poor convergence*

- *Computationally and labor intensive*

- *Thermal resummations*
- *Gauge Invariance (radiative barriers)*
- *RG invariance at $T > 0$*

BSM proposals

Benchmark pert theory

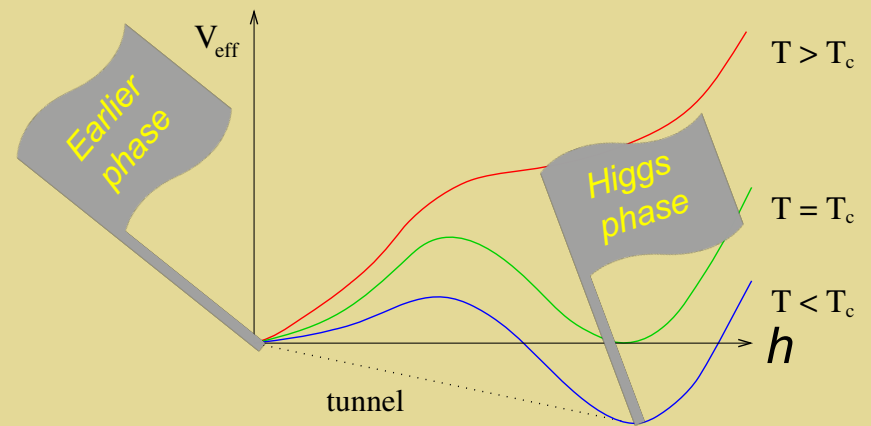
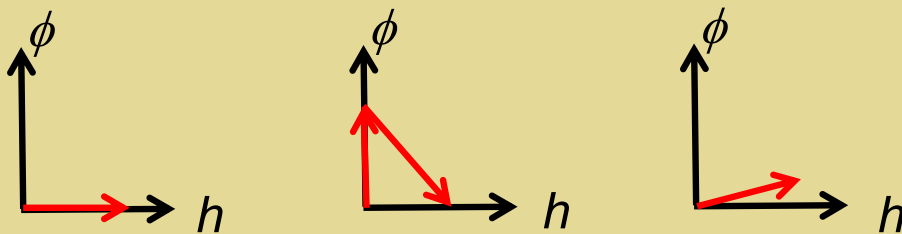
Theory-Pheno Interface



Simple Higgs portal models:

- *Real gauge singlet (SM + 1)*
- *Real EW triplet (SM + 3)*

$$V \subset a_1 H^2 \phi + a_2 H^2 \phi^2$$



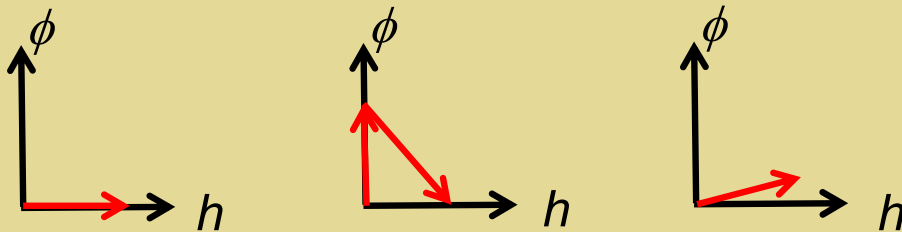
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Phenomenology

$$h_1 = \sin \theta s + \cos \theta h$$

$$h_2 = \cos \theta s - \sin \theta h$$

$m_{1,2}; \theta; h_i h_j h_k$ couplings

Collider Probes

- *Resonant di-Higgs ($h_1 h_1$) production **
- *Heavy h_2 production **
- *Associated production ($Z h_1$) and non-resonant di-Higgs production **
- *Exotic Higgs decays ***

** Heavy h_2*

*** Light h_2*

Collider Probes

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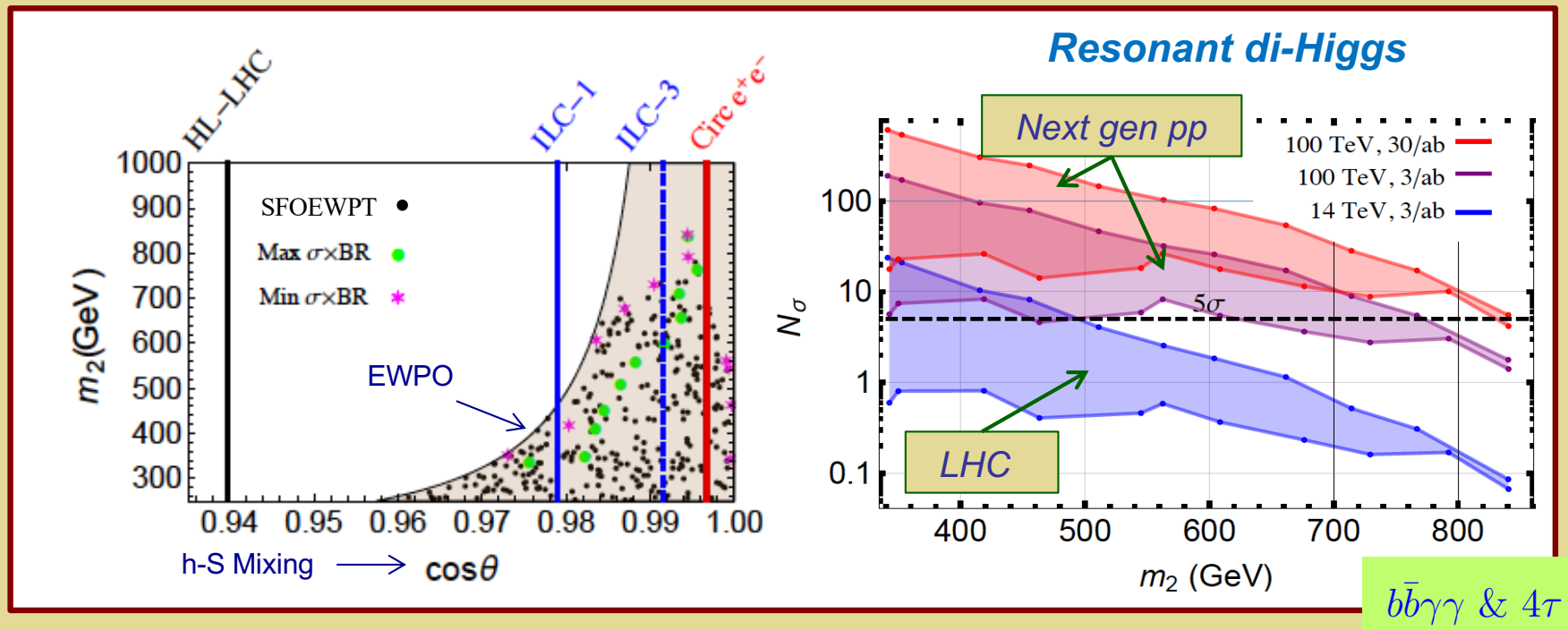
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** Heavy h_2*

*** Light h_2*

Singlets: Precision & Res Di-Higgs Prod

SFOEWPT Benchmarks: Resonant di-Higgs & precision Higgs studies

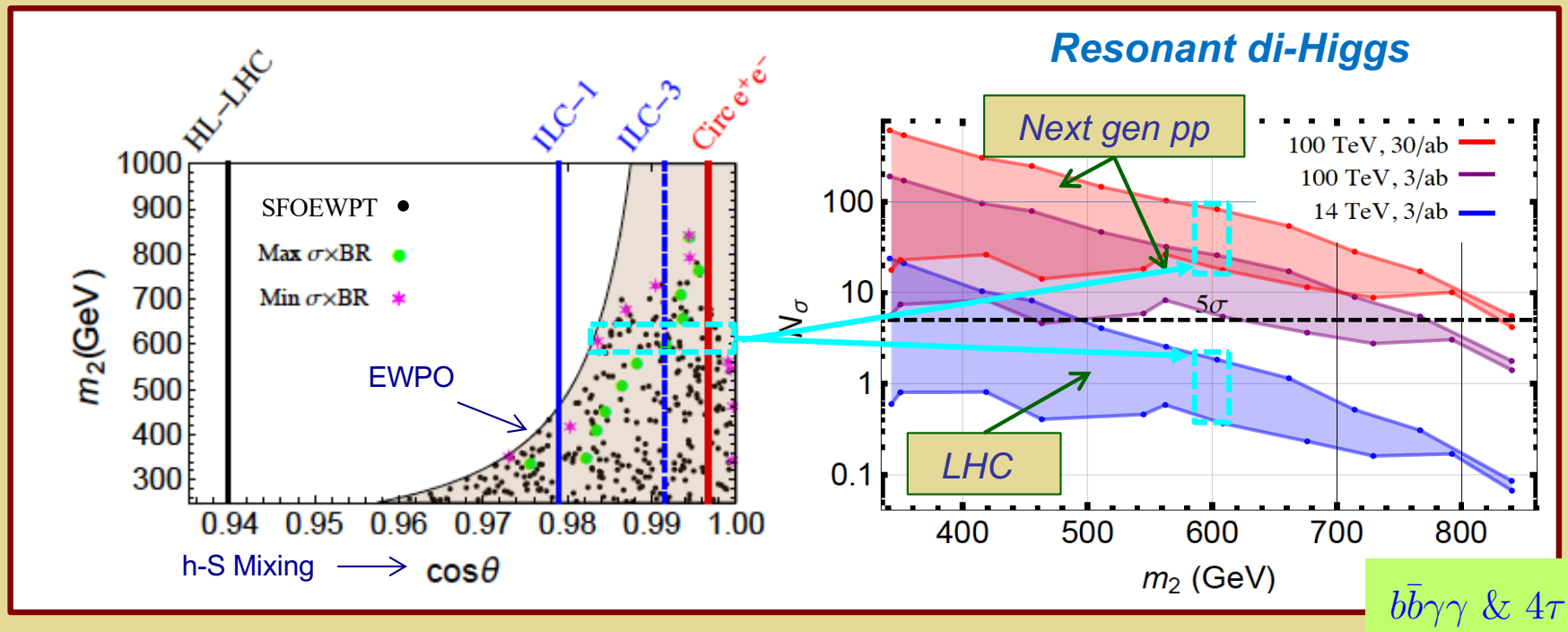


Kotwal, No, R-M, Winslow 1605.06123

See also: Huang et al, 1701.04442;
Li et al, 1906.05289

Singlets: Precision & Res Di-Higgs Prod

SFOEWPT Benchmarks: Resonant di-Higgs & precision Higgs studies

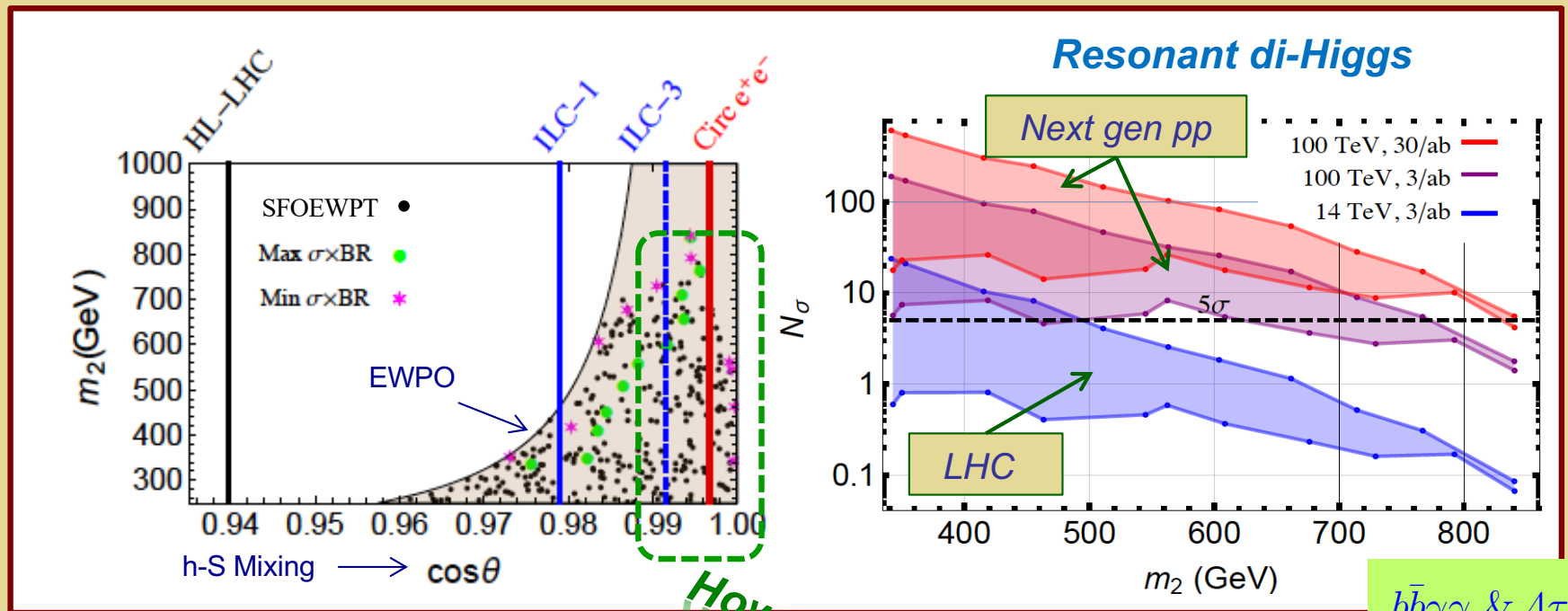


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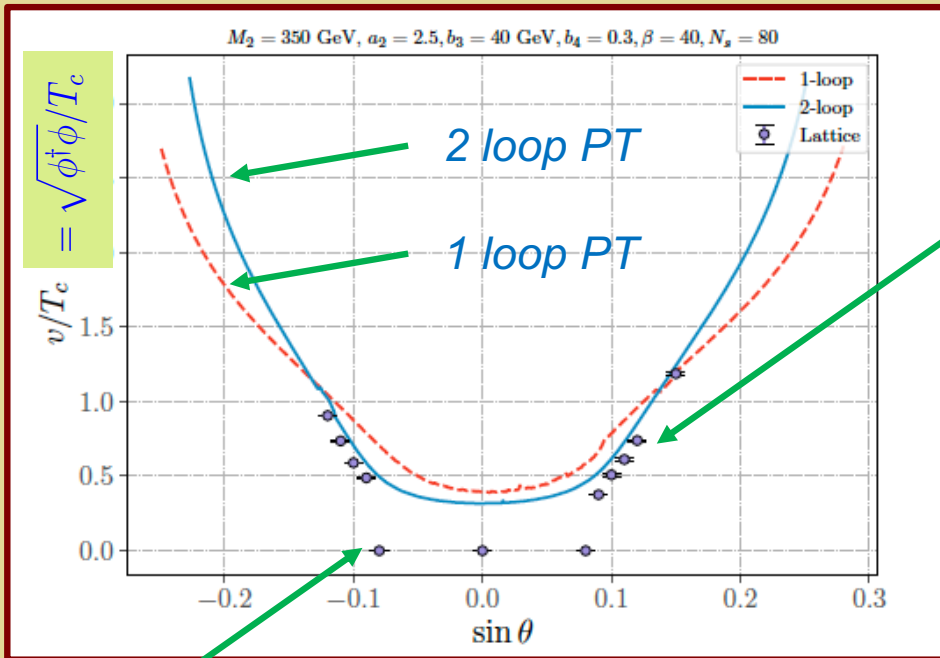


How reliable?
How reliable?

Kotwal, No, R-M, Winslow 1605.06123

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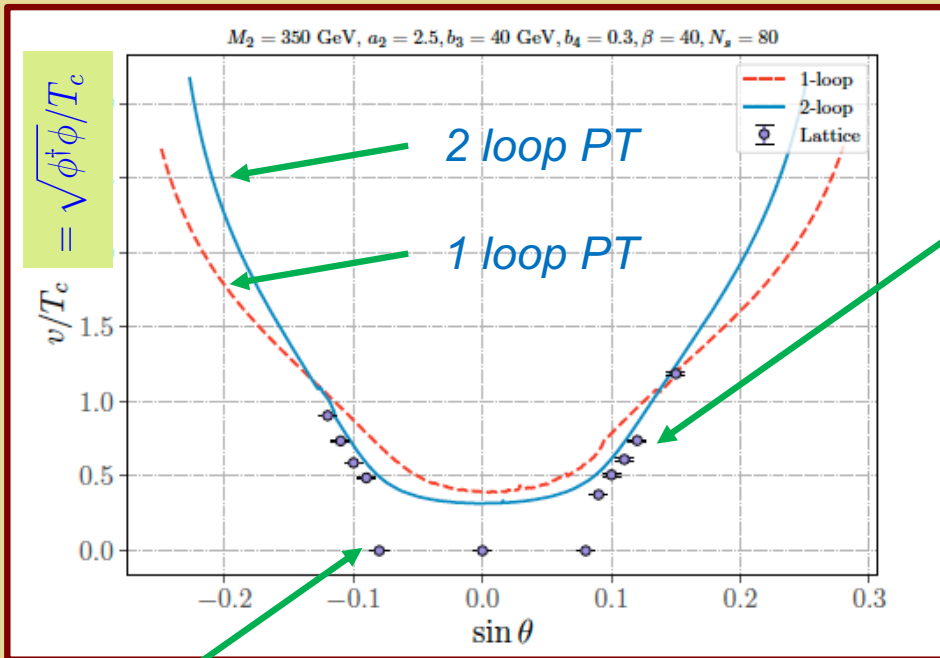
Singlets: Lattice vs. Pert Theory



Lattice:
FOEWPT

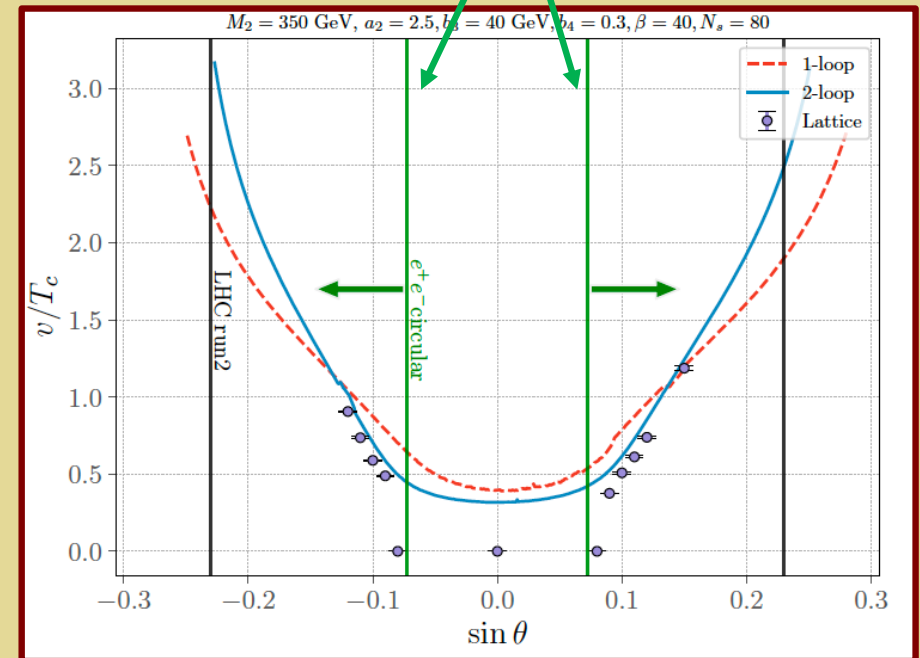
Lattice:
Crossover

Singlets: Lattice vs. Pert Theory



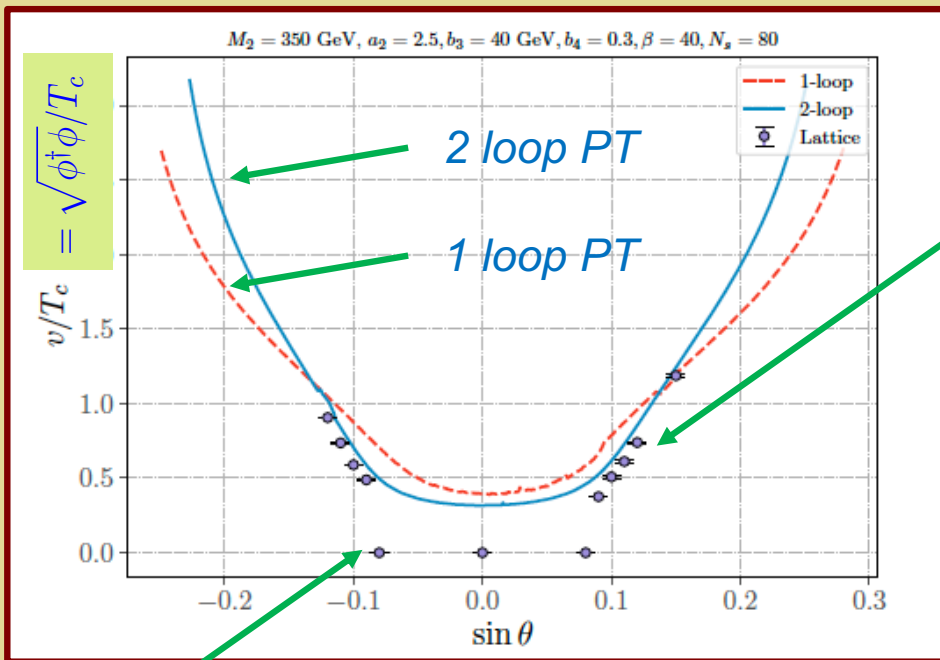
Lattice:
FOEWPT

Future e^+e^-



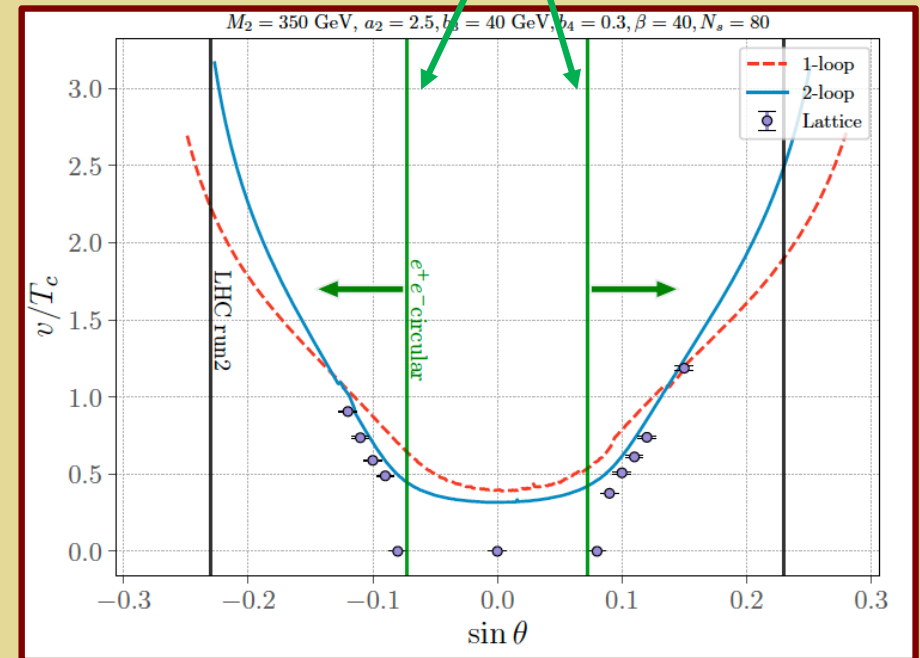
Lattice:
Crossover

Singlets: Lattice vs. Pert Theory



Lattice:
FOEWPT

Future e^+e^-



Lattice:
Crossover

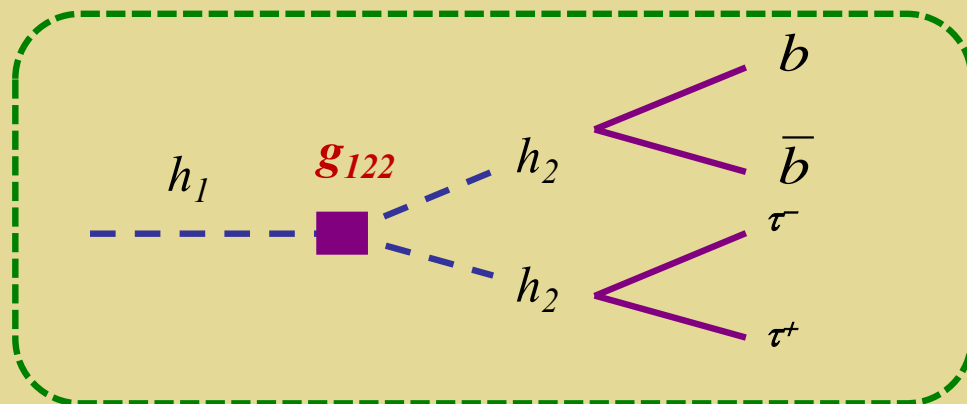
- Lattice: crossover-FOEWPT boundary
- FOEWPT region: PT-lattice agreement
- Pheno: precision Higgs studies may be sensitive to a greater portion of FOEWPT-viable param space than earlier realized

Collider Probes

- Resonant di-Higgs ($h_1 h_1$) production *
- Heavy h_2 production *
- Associated production ($Z h_1$) and non-resonant di-Higgs production *
- Exotic Higgs decays **

* Heavy h_2

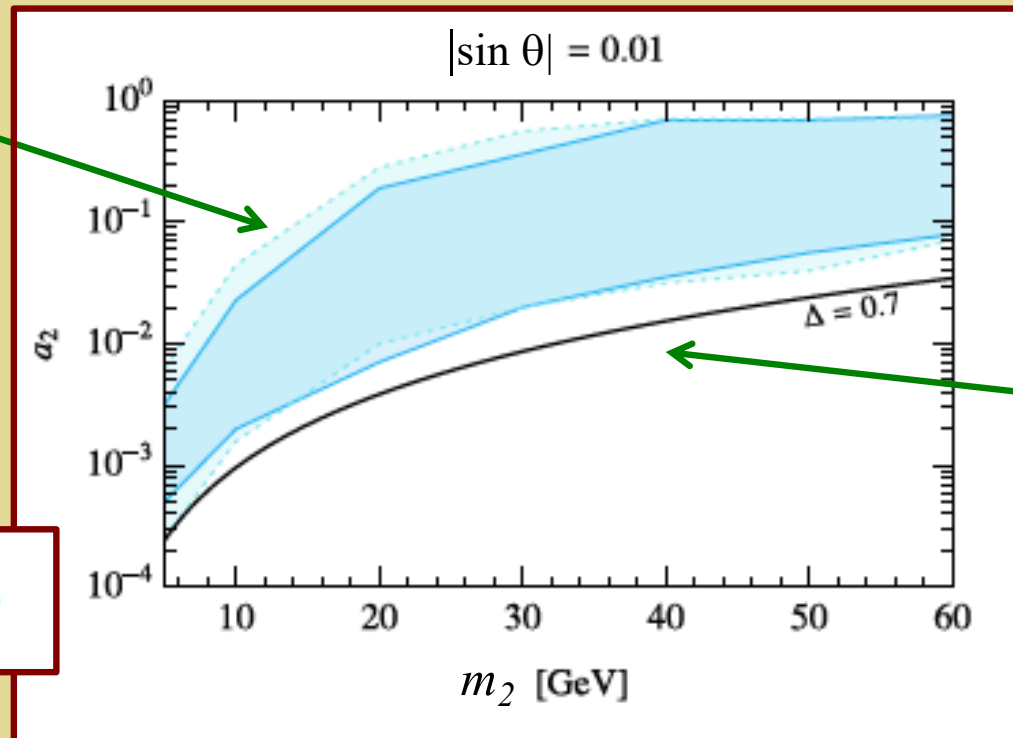
** Light h_2



Light Singlets: Exotic Higgs Decays

One loop perturbation theory

EWPT viable:
numerical

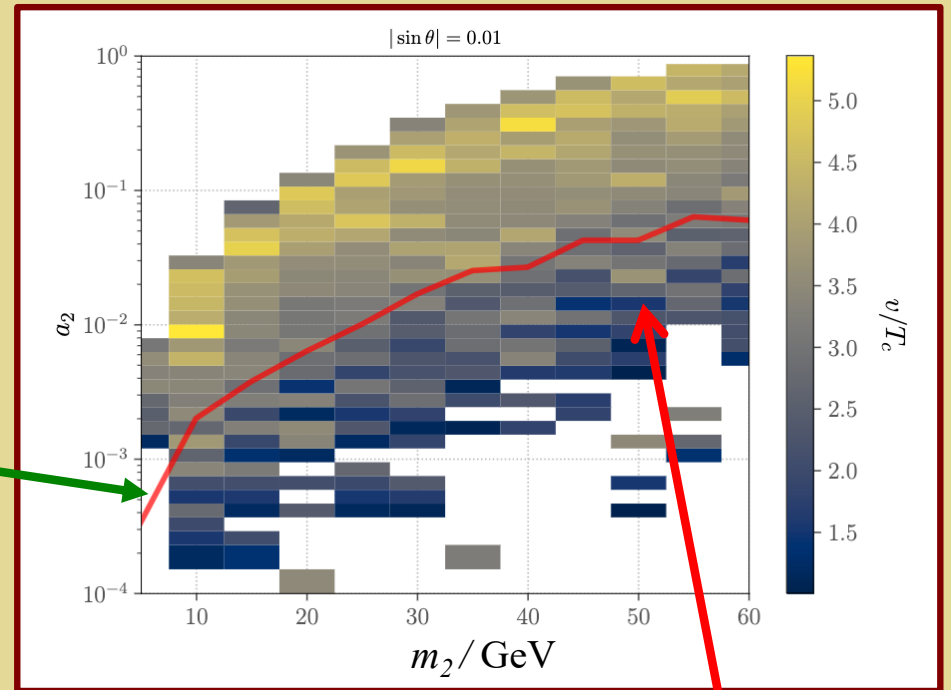
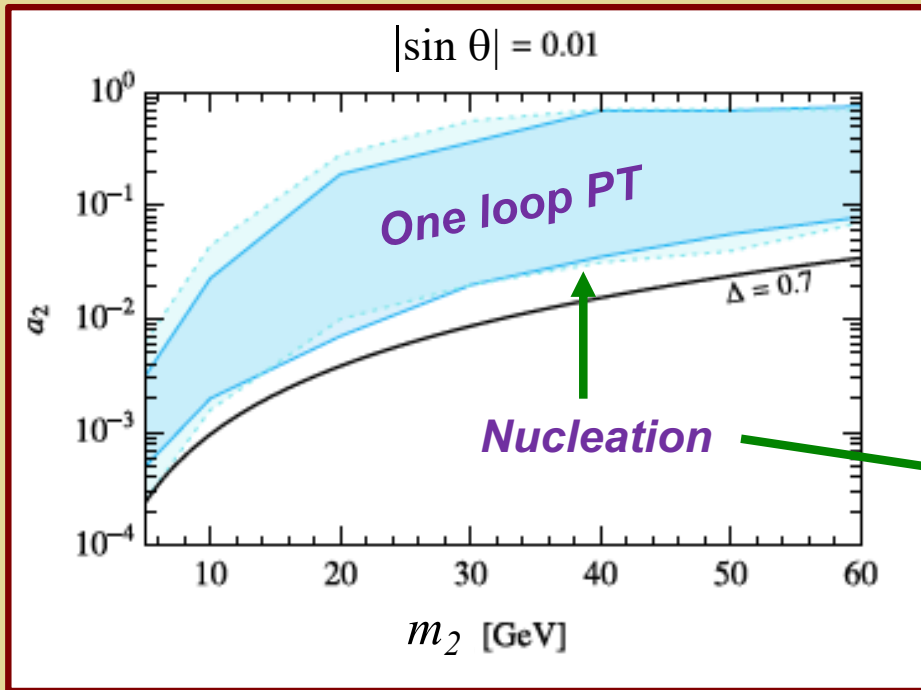


EWPT viable:
Semi analytic
→ nucleation
decisive

$$g_{122} = \frac{1}{2}va_2 + \mathcal{O}(\theta^2)$$

J. Kozaczuk, MR-M, J. Shelton 1911.10210
See also: Carena et al 1911.10206, Carena
et al 2203.08206, Wang et al 2203.10184,

New: Lattice + EFT @ $T > 0$

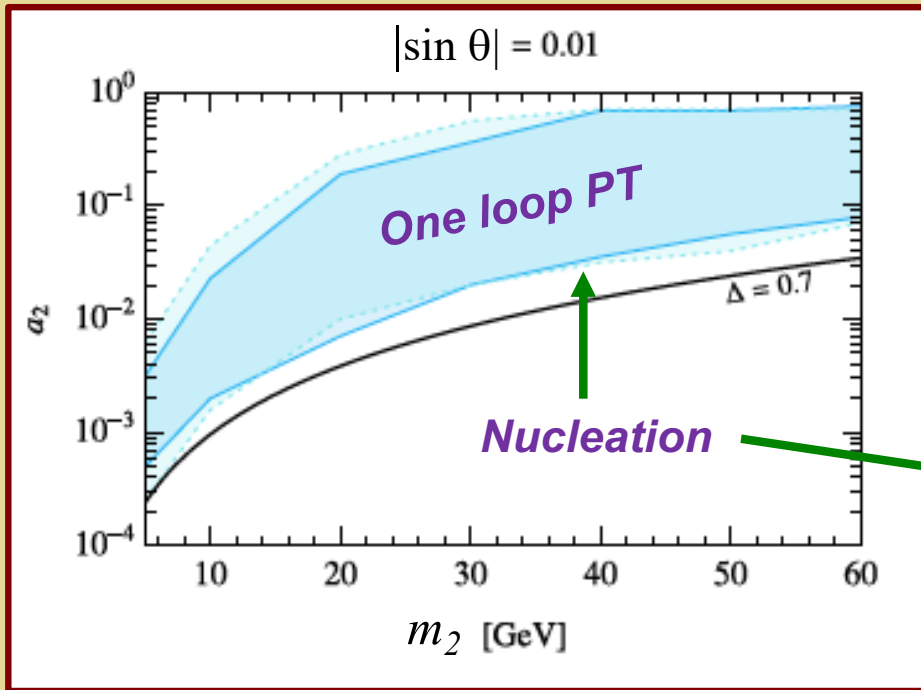


J. Kozaczuk, MR-M, J. Shelton 1911.10210

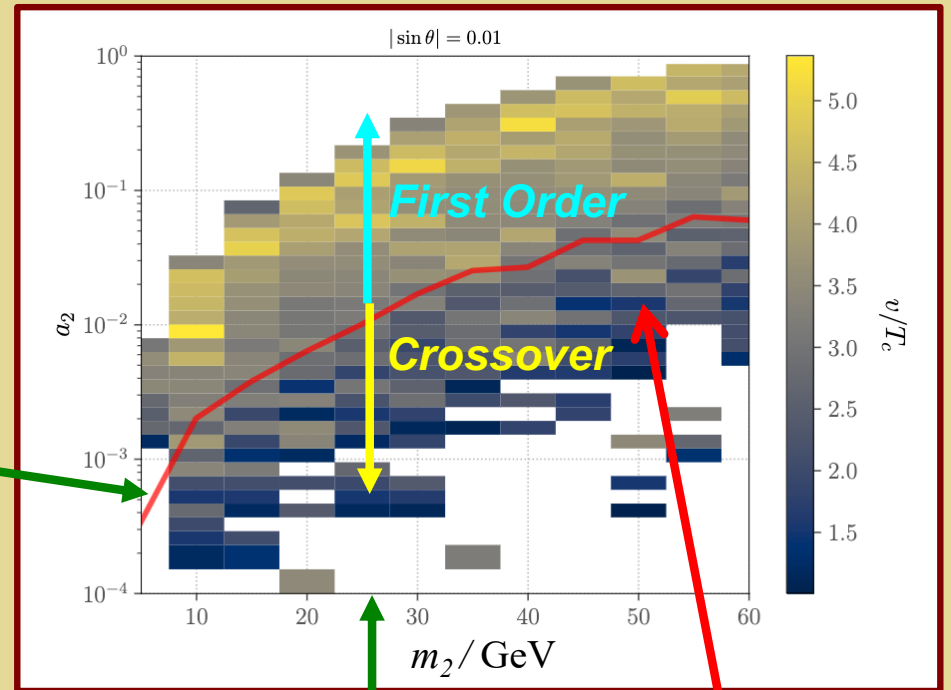
**Two-loop PT:
3d EFT**

L. Niemi, MJRM, G. Xia 2405.01191

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J. Kozaczuk, MR-M, J. Shelton 1911.10210

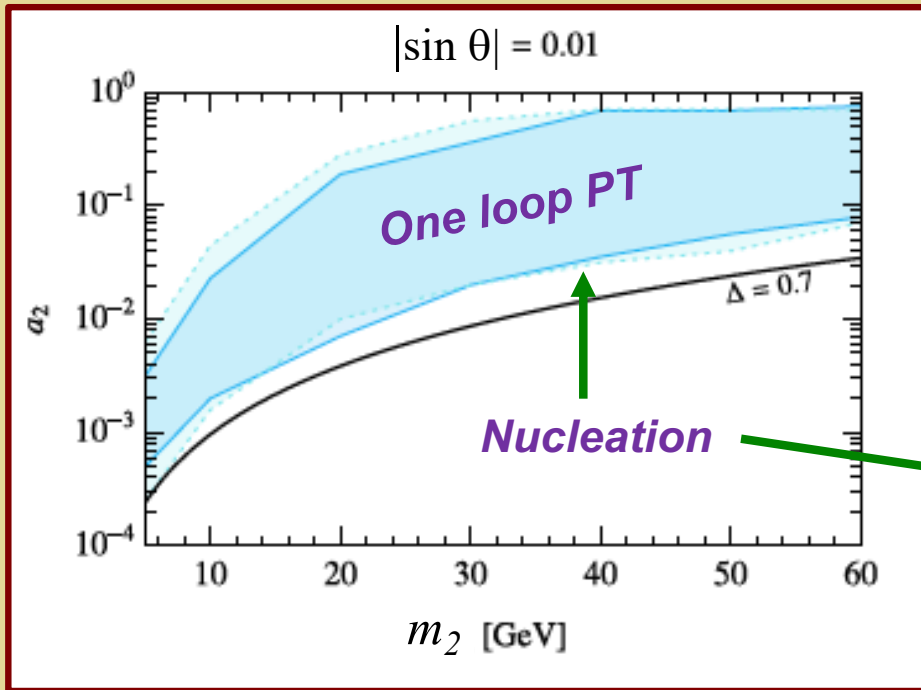


Lattice study

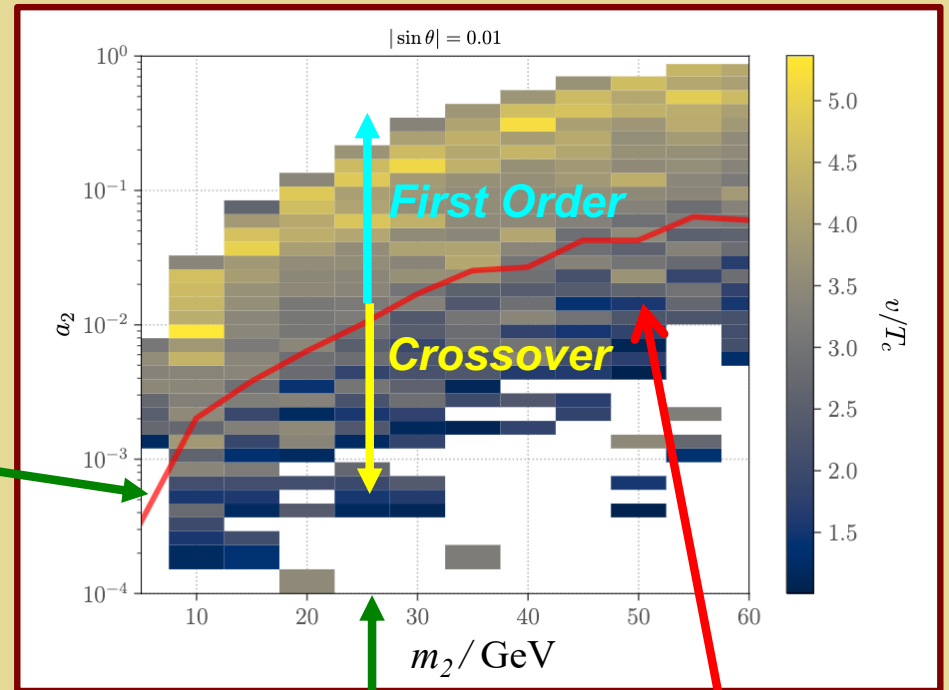
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Lattice study

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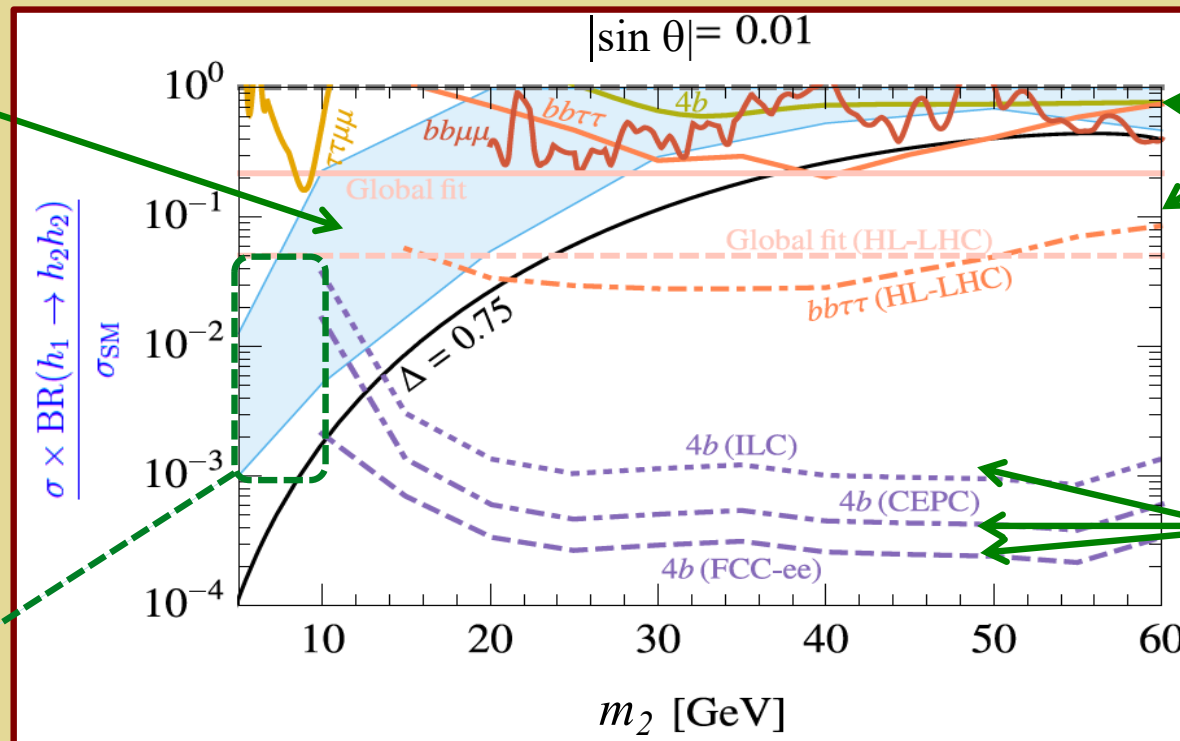
Small portal couplings
→ FO EWPT unlikely

L. Niemi, MJRM, G. Xia 2405.01191

Light Singlets: Exotic Higgs Decays

Prompt decays: $h_1 \rightarrow h_2 h_2 \rightarrow AA BB$

EWPT viable:
numerical



LHC: 2019 &
HL

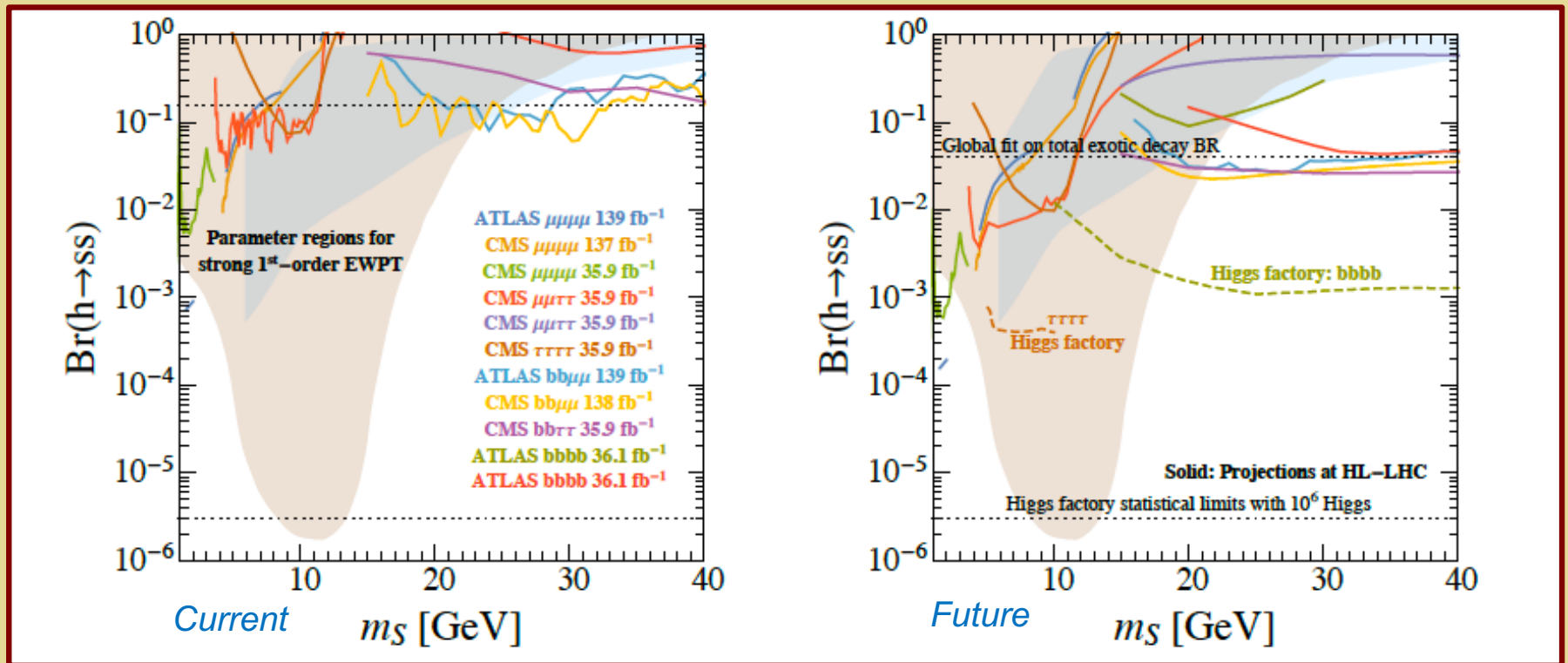
Future e^+e^-

Other
probes?

J. Kozaczuk, MR-M, J. Shelton 1911.10210
See also: Carena et al 1911.10206, Carena
et al 2203.08206, Wang et al 2203.10184,

Light Singlets: Exotic Higgs Decays

Z_2 breaking: prompt h_2 decays

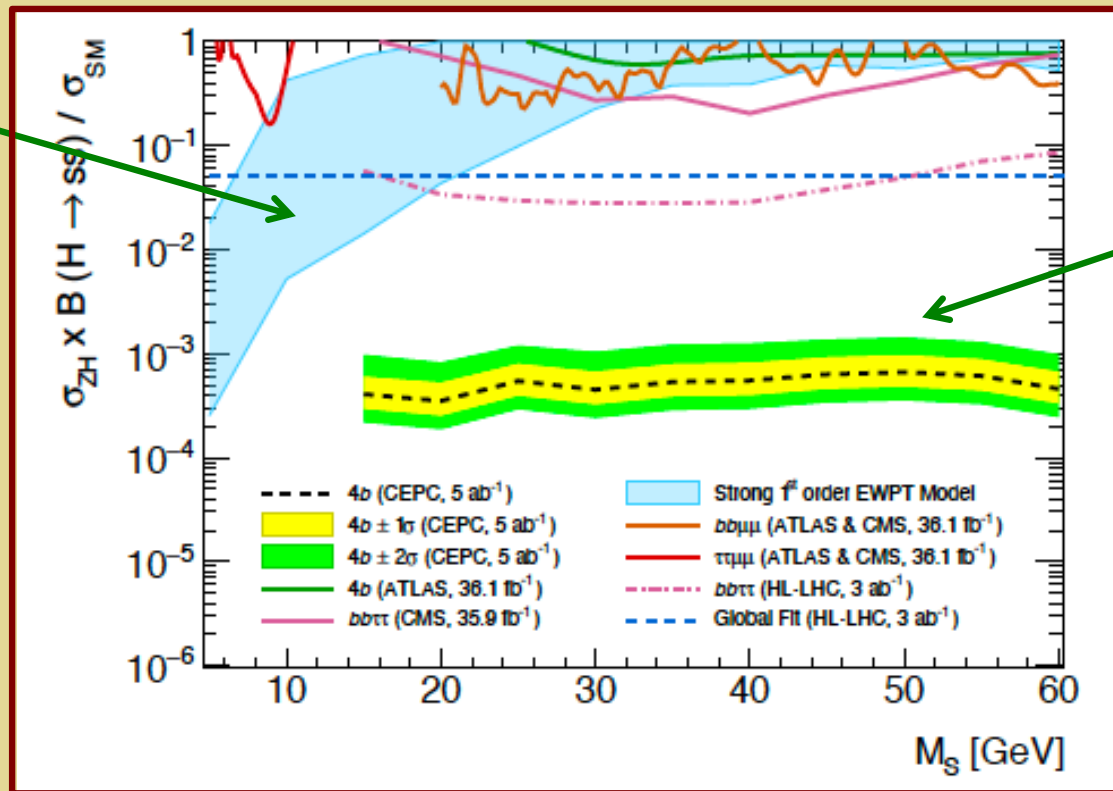


Carena et al (Snowmass) 2203.08206

Light Singlets: Exotic Higgs Decays

$$h_1 \rightarrow h_2 \quad h_2 \rightarrow 4b \text{ (prompt)}$$

EWPT viable:
numerical



CEPC 4b

J. Wang et al (Snowmass) 2203.10184

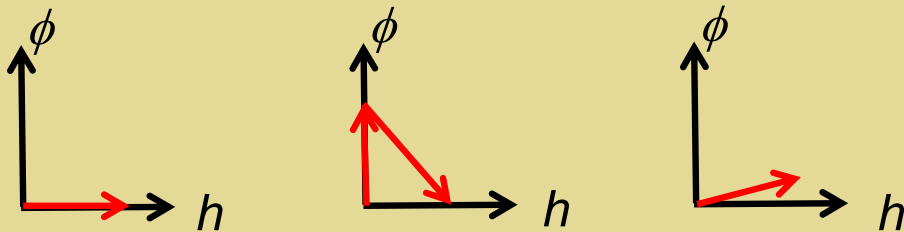
Theory-Pheno Interface



Simple Higgs portal models:

- *Real gauge singlet (SM + 1)*
- *Real EW triplet (SM + 3)*

$$V \subset a_1 H^2 \phi + a_2 H^2 \phi^2$$



Theory-Pheno Interface

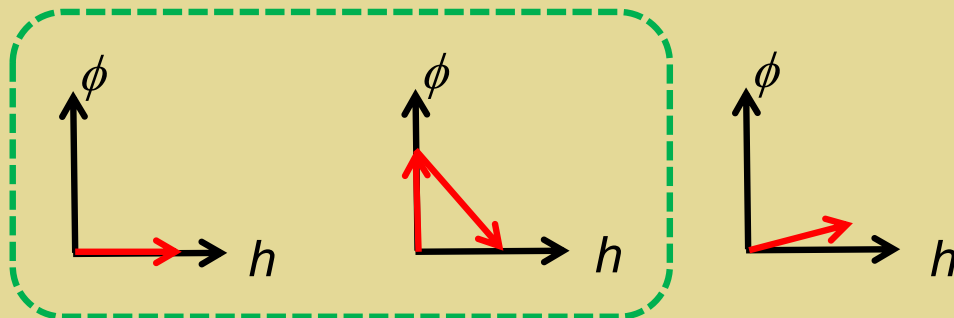


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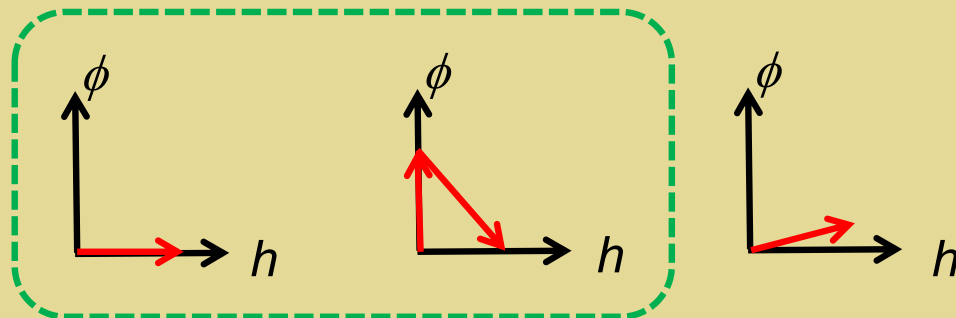


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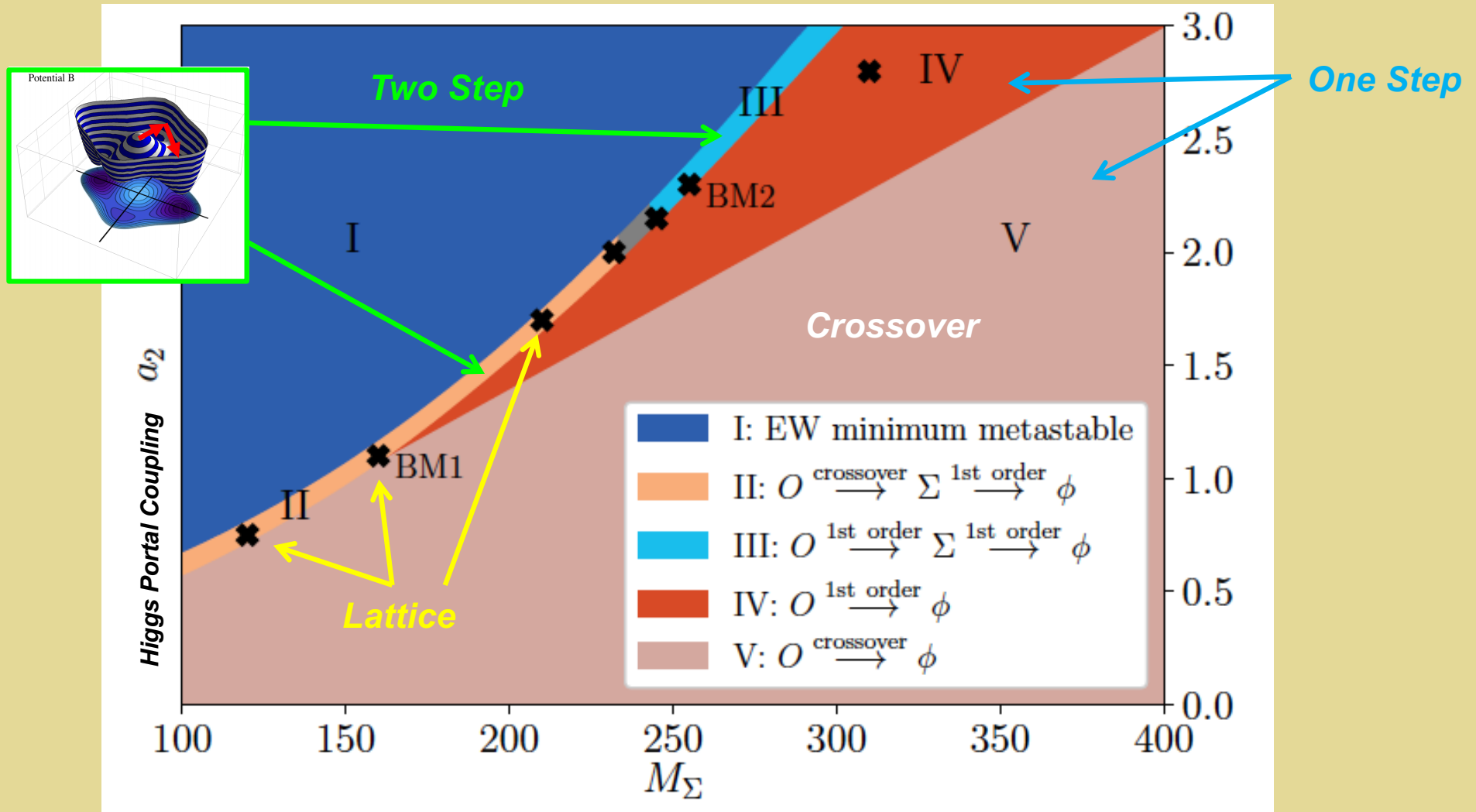
$$V \subset a_1 H^2 \phi + a_2 H^2 \phi^2$$



Phenomenology

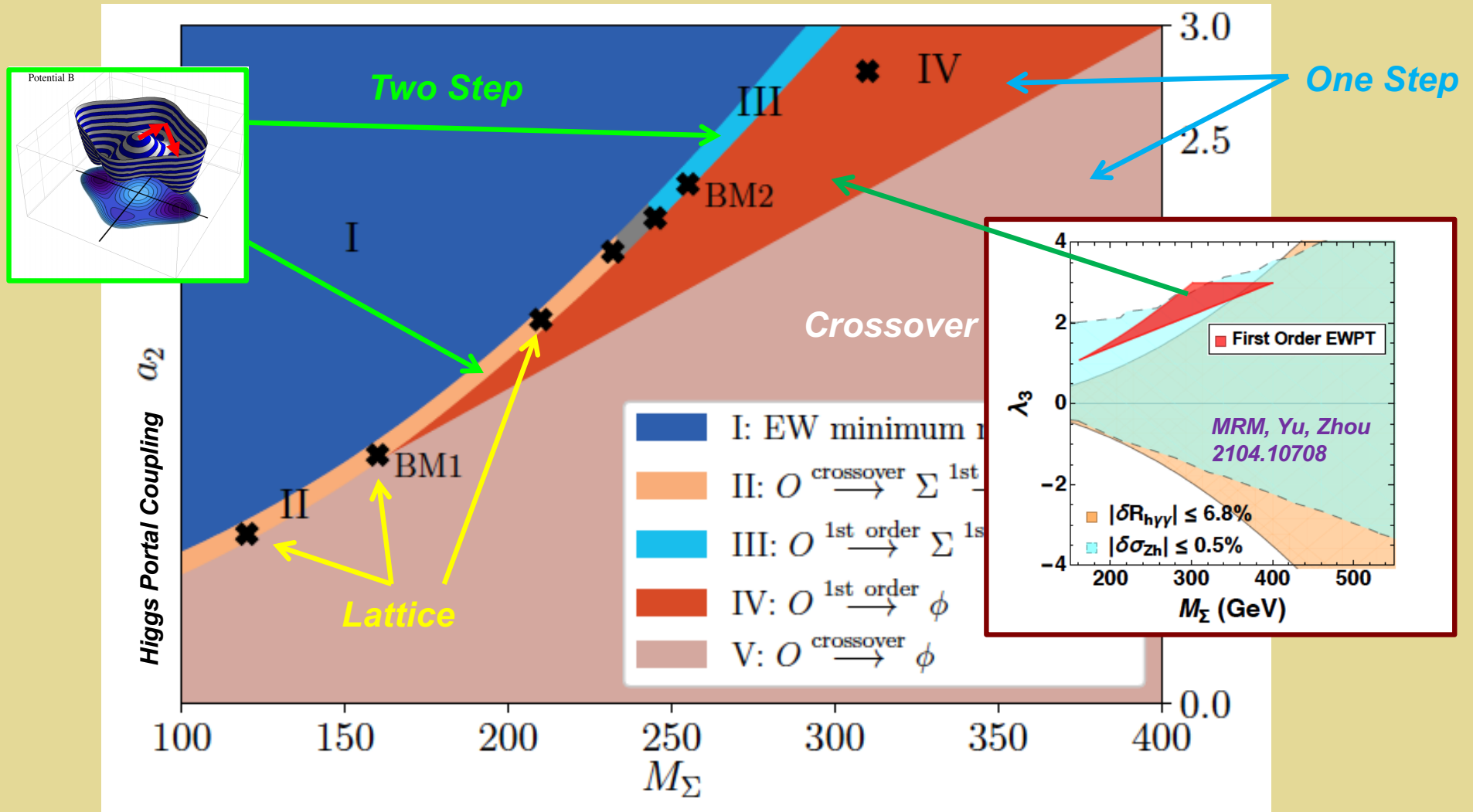
- Gravitational waves
- Collider: $h \rightarrow \gamma\gamma$, dis charged track, NLO $e^+e^- \rightarrow Zh...$

Real Triplet & EWPT: Novel EWSB



- 1 or 2 step
- Non-perturbative

Real Triplet & EWPT: Novel EWSB



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- Non-perturbative

BSM EWPT: Inter-frontier Connections

***Robust theory:
EFT + lattice***



***Observables:
model specific***



***Hydro:
 $\alpha, \beta / H_*$***

BSM EWPT: Inter-frontier Connections

GW – Collider “inverse problem” **

**Robust theory:
EFT + lattice**

**Observables:
model specific**

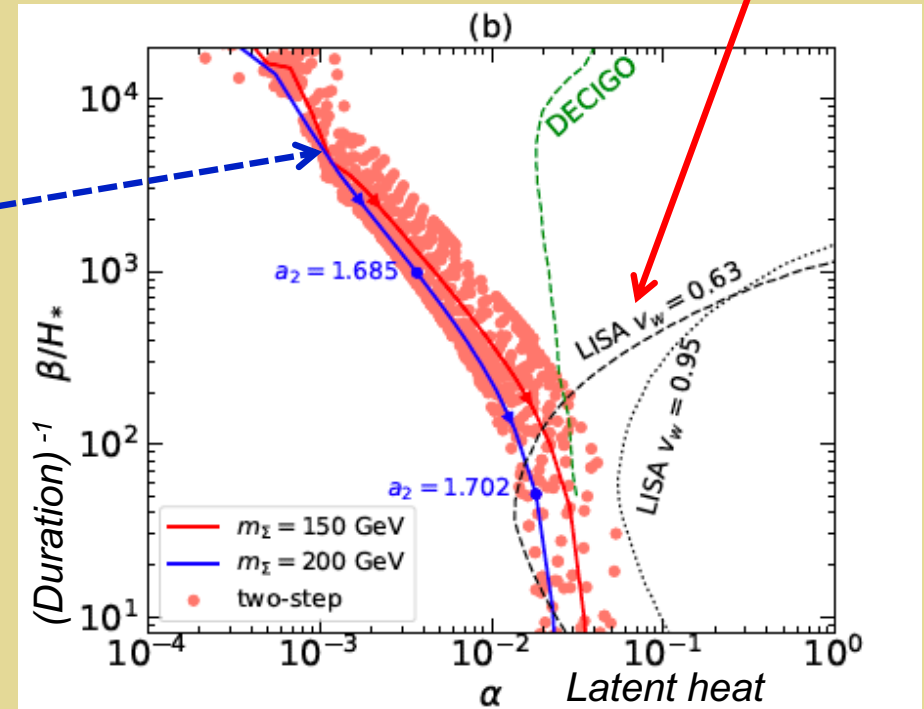
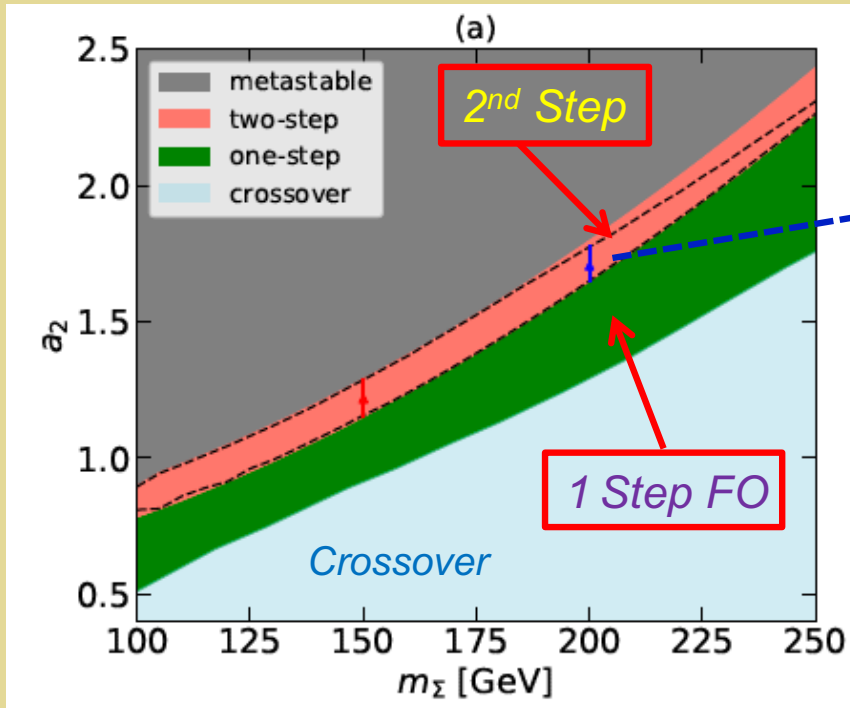


**Combined
reach**

**Hydro:
 $\alpha, \beta / H^*$**

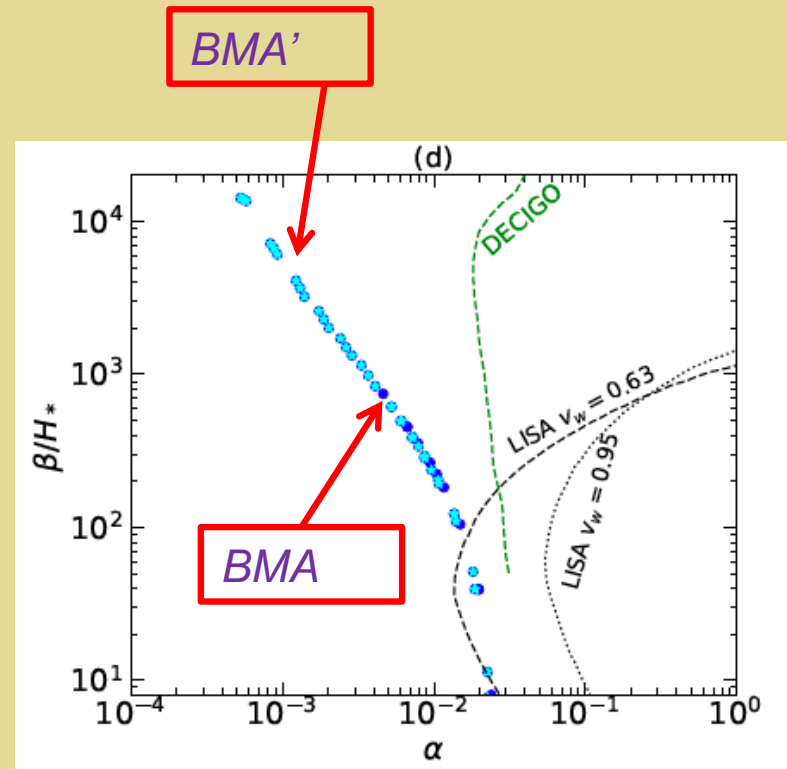
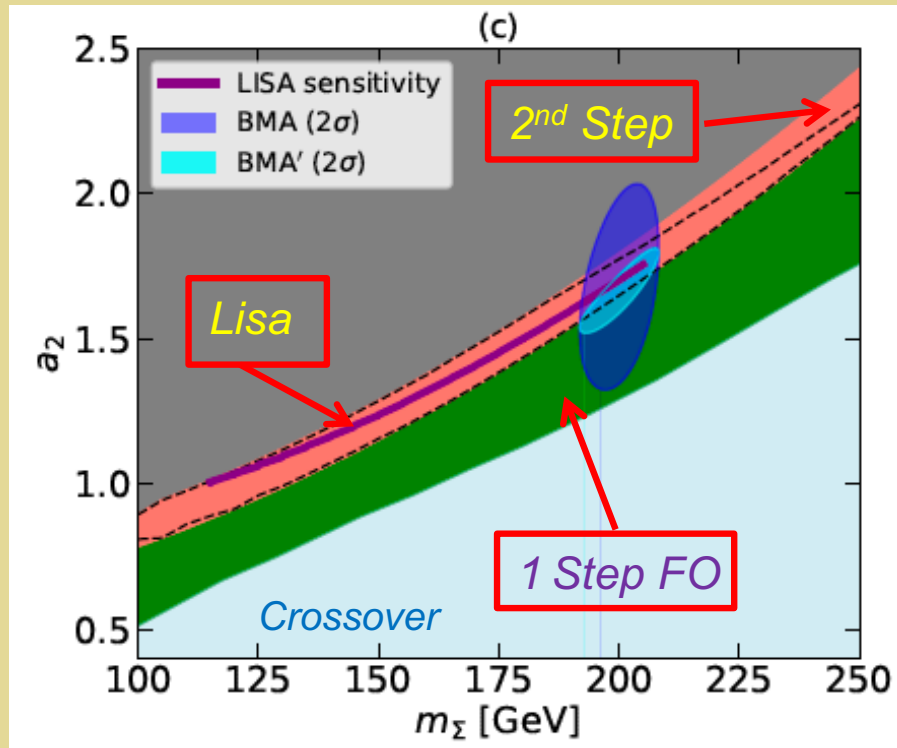
**** How can we exploit experiment to identify EWPT-viable models & parameters ?**

GW & EWPT Phase Diagram



- *Single step transition: GW well outside LISA sensitivity*
- *Second step of 2-step transition can be observable*
- *Significant GW sensitivity to portal coupling*

GW & EWPT Phase Diagram



$$BMA: m_\Sigma + h \rightarrow \gamma\gamma$$

$$BMA': BMA + \Sigma^0 \rightarrow ZZ$$

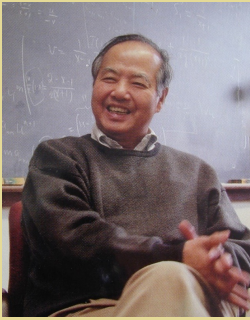
- *Two-step*
- *EFT+ Non-perturbative*

III. Outlook

Was There an Electroweak Phase Transition ?

- *Answering this question is an important and exciting challenge for Higgs Physics @ LHC/CEPC/FCC-ee/ILC...*
- *The relevant scale T_{EW} makes this physics a prime target for collider and gravitational wave probes*
- *The EWPT question entails a rich interplay of model building, thermal QFT, phenomenology & experiment → **robust thermal field theory is vital***
- *The **collider – gravitational wave “inverse problem”** has emerged as a particularly compelling arena for further exploration and opportunity HEP community and beyond*

T. D. Lee Institute / Shanghai Jiao Tong U.



Director



Prof Jie Zhang

A point of convergence of the world's top scientists

A launch pad for the early-career scientists

A world famous source of original innovation



Founded 2016

100+

faculty members from 17 countries and regions, with over 40% of them foreign (non-Chinese) citizens

Theory & Experiment

Particle & Nuclear Physics

Astronomy & Astrophysics

Quantum Science

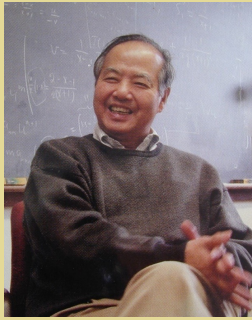
Dark Matter & Neutrino

Laboratory Astrophysics

Topological Quantum Computation

<https://tdli.sjtu.edu.cn/EN/>
<https://www.youtube.com/watch?v=z0awD6q8FTI>

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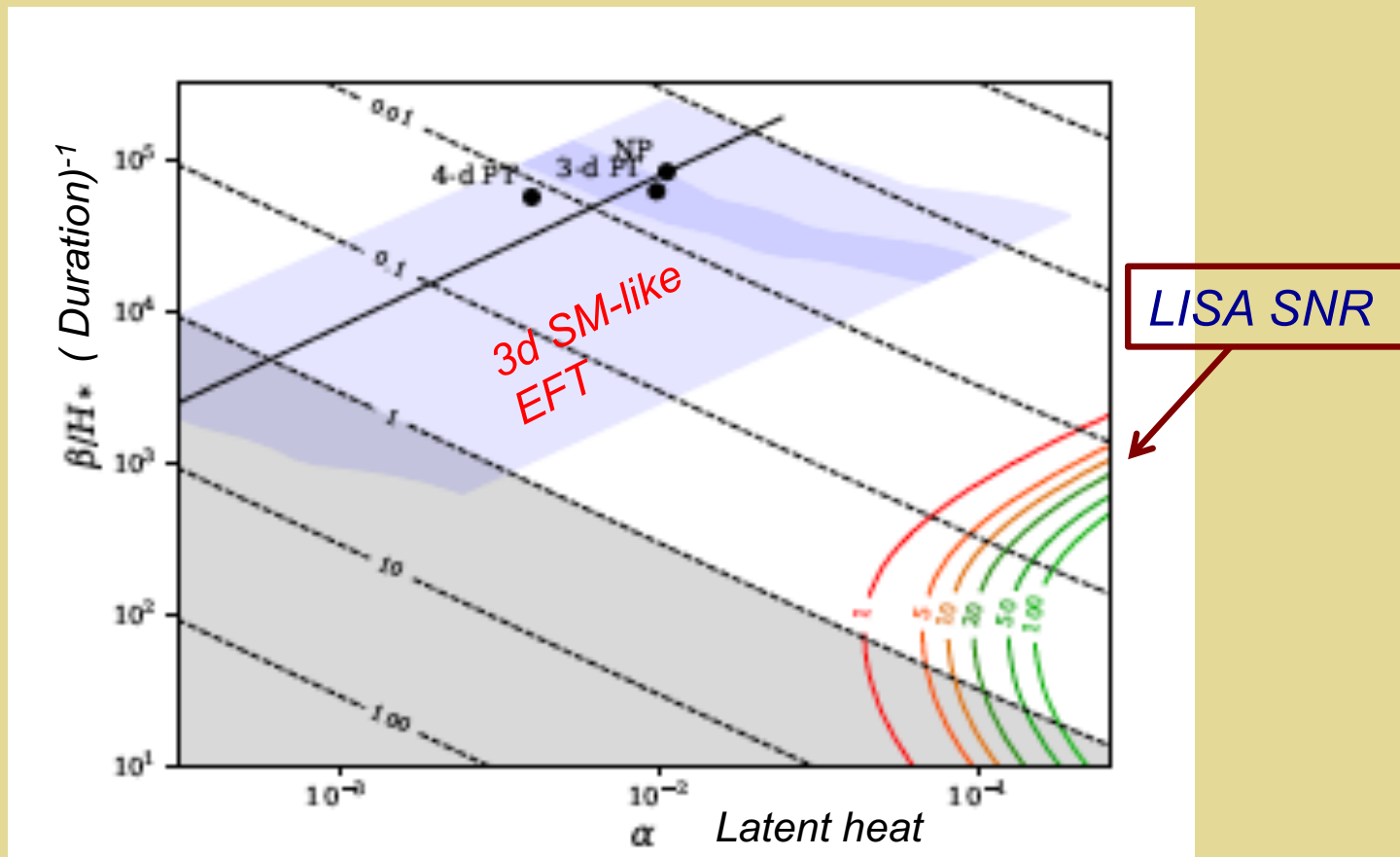
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谢谢！

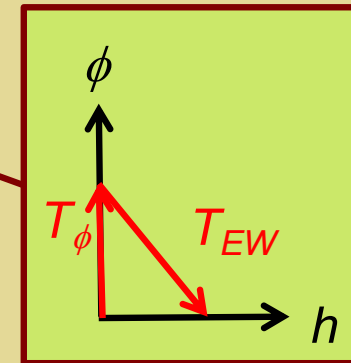
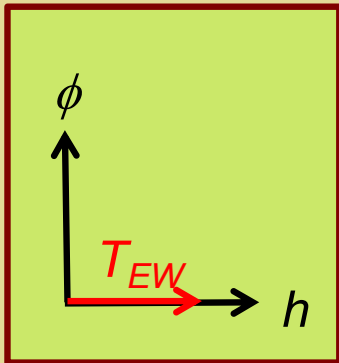
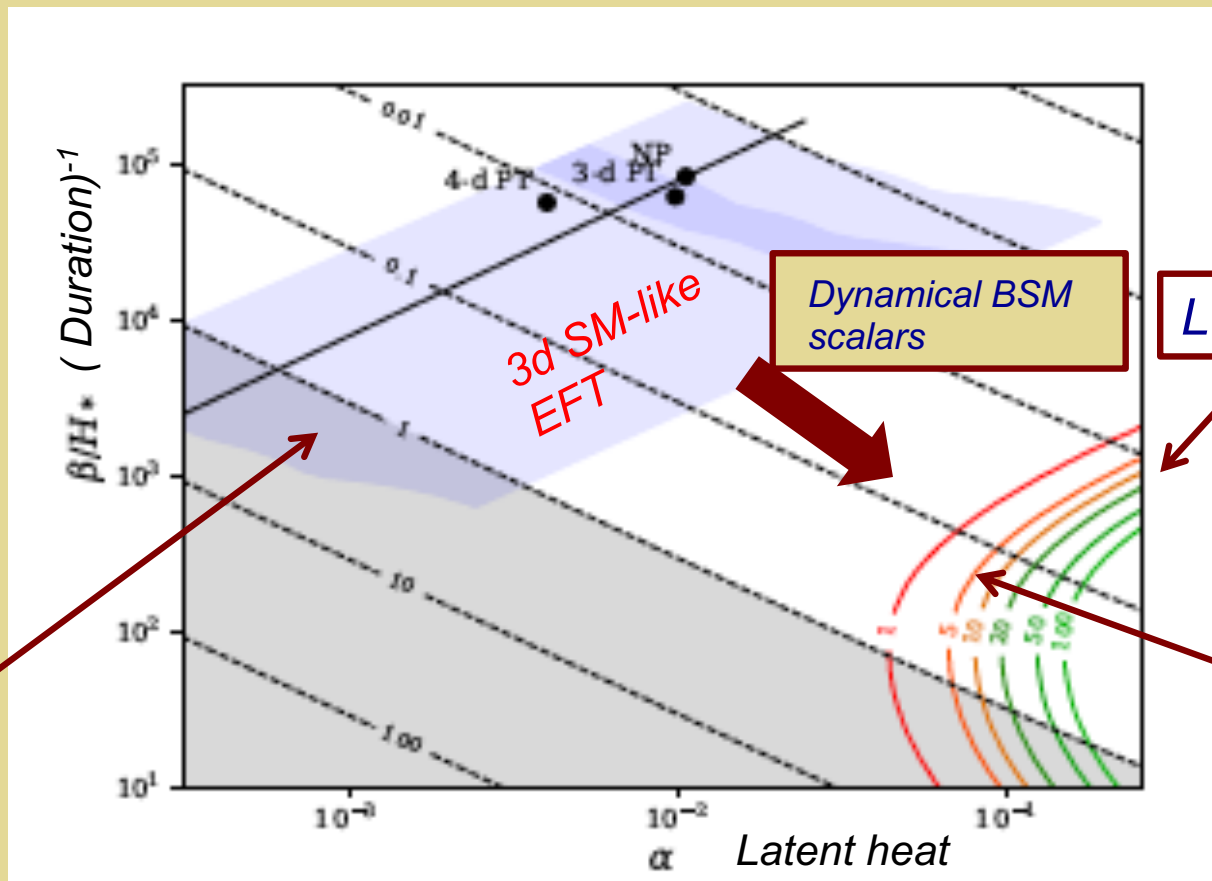
Back Up Slides

GW-Collider

BSM Scalar: EWPT & GW

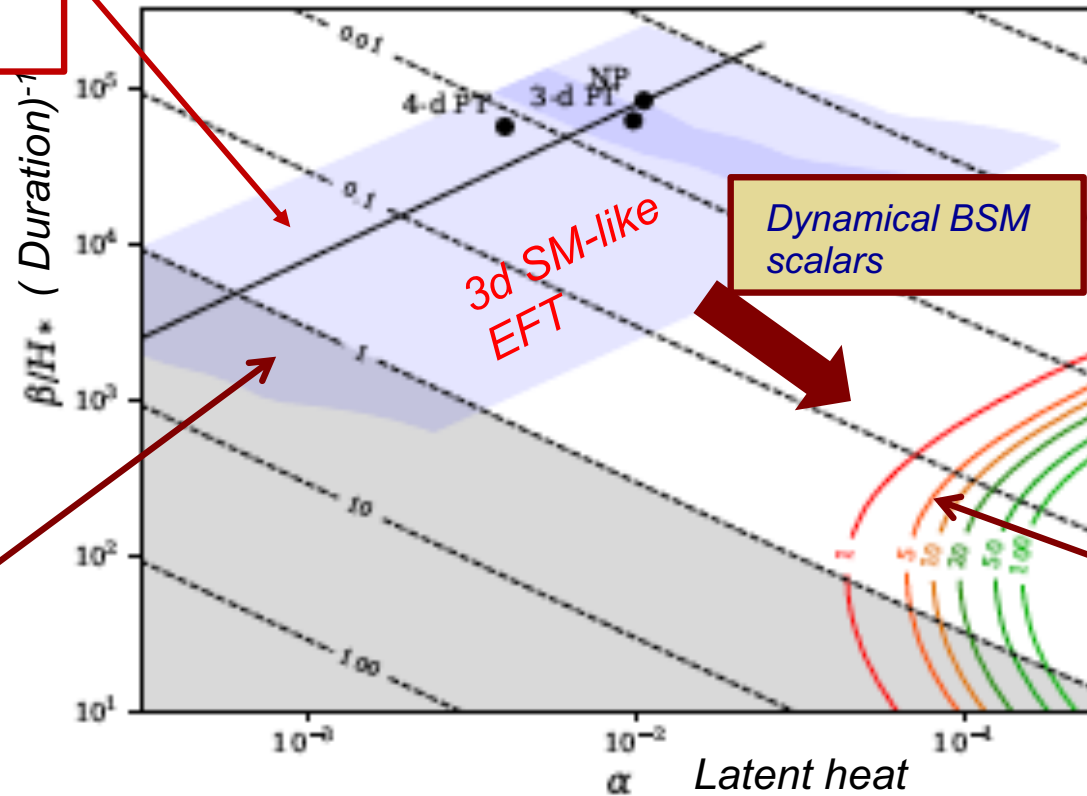


BSM Scalar: EWPT & GW

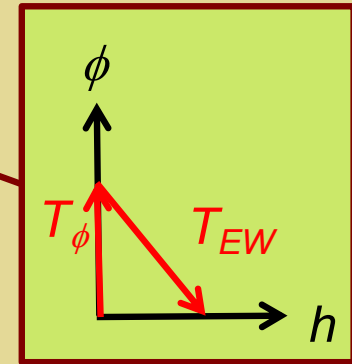
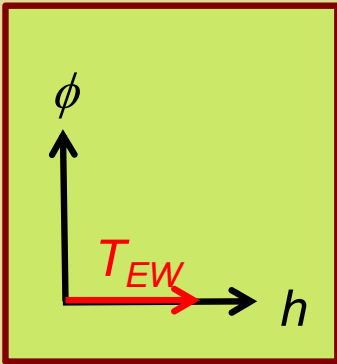


BSM Scalar: EWPT & GW

Collider probes of BSM parameters in \mathcal{L}_{full}



LISA SNR



Nucleation

Tunneling @ $T > 0$: Gravitational Waves

Amplitude & frequency: latent heat & intrinsic time scale

Normalized latent heat

$$\Delta Q = \Delta F + T \Delta S$$

$$S = -\partial F / \partial T$$

$$F \approx V$$

Time scale

$$\frac{\beta}{H_*} = T \frac{d}{dT} \frac{S_3}{T}$$

$$\Delta Q \approx \Delta V - T \partial \Delta V / \partial T$$

$$\alpha = \frac{30 \Delta q}{\pi^2 g_* T^4}$$

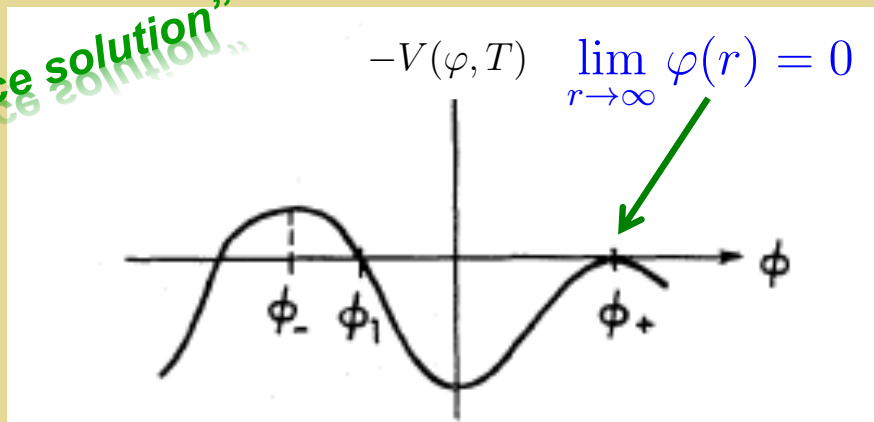
How Reliable?
How Believable?

$T=0$: S. Coleman, *PRD* 15 (1977) 2929

Tunneling @ $T > 0$

Scalar Quantum Field Theory

“Bounce solution”
“Bounce solution”



Tunneling rate / unit volume:

$$\Gamma = A e^{-\beta S_3 / \hbar} [1 + \mathcal{O}(\hbar)]$$

$$\frac{d^2 \varphi}{dr^2} + \frac{2}{r} \frac{d\varphi}{dr} = V'(\varphi, T)$$

Exponent in Γ

Path: minimize S_E

$$S_3 = \int d^3x \left\{ \frac{1}{2} (\vec{\nabla} \varphi)^2 + V(\varphi, T) \right\}$$

Friction term

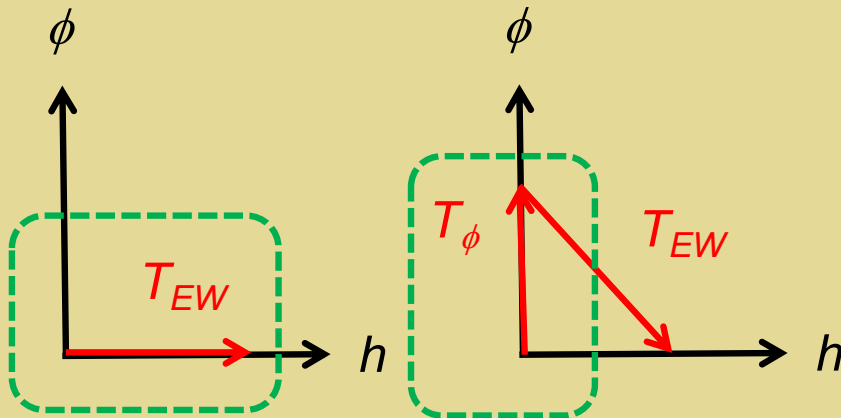
$$A \sim \mathcal{O}(1) \times T^4$$

Tunneling @ $T > 0$

Radiative barriers \rightarrow st'd method gauge-dependent

Tunneling rate / unit volume:

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Tunneling @ $T > 0$

Theoretical issues:

- *Radiatively-induced barrier (St'd Model) → gauge dependence*
 - $T = 0$ Abelian Higgs: E. Weinberg & D. Metaxas: hep-ph/9507381
 - $T=0$ St'd Model: A. Andreassen, W. Frost, M. Schwartz 1408.0287
 - $T > 0$ Gauge theories: **recently solved in 2112.07452 (→ PRL) and 2112.08912**
- *Multi-field problem (still gauge invar issue)*
 - Cosmotransitions: C. Wainwright 1109.4189
 - Espinosa method: J. R. Espinosa 1805.03680

(Re) Organize the Perturbative Expansion

Illustrate w/ Abelian Higgs

$$\mathcal{L} = \frac{1}{4} F_{\mu\nu} F_{\mu\nu} + (D_\mu \Phi)^* (D_\mu \Phi) + \mu^2 \Phi^* \Phi + \lambda (\Phi^* \Phi)^2 + \mathcal{L}_{\text{GF}} + \mathcal{L}_{\text{FP}}$$

- Lofgren, MRM, Tenkanen, Schicho 2112.0752 → PRL
- Hirvonen, Lofgren, MRM, Tenkanen, Schicho 2112.08912

Full 3D effective action

$$S_3 = \int d^3x \left[V^{\text{eff}}(\phi, T) + \frac{1}{2} Z(\phi, T) (\partial_i \phi)^2 + \dots \right]$$

Adopt appropriate power-counting in couplings

$$S_3 = a_0 g^{-\frac{3}{2}} + a_1 g^{-\frac{1}{2}} + \Delta$$

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G.I. perturbative expansion only valid up to NLO → Δ : higher order contributions only via other methods

G.I. perturbative expansion

Tunneling @ $T > 0$: G.I. & Nielsen Identities

Adopt appropriate power-counting in couplings

Lofgren, MRM, Tenkanen,
Schicho 2112.0752 → PRL

$$S_3 = a_0 g^{-\frac{3}{2}} + a_1 g^{-\frac{1}{2}} + \Delta$$

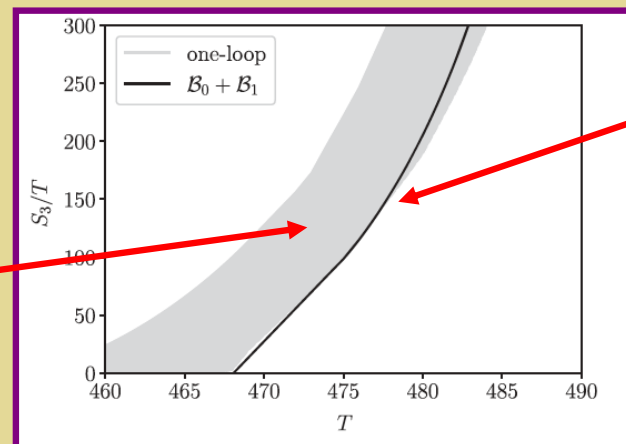
Order-by-order consistent with Nielsen Identities

$$\xi \frac{\partial S^{\text{eff}}}{\partial \xi} = - \int d^d \mathbf{x} \frac{\delta S^{\text{eff}}}{\delta \phi(x)} \mathcal{C}(x)$$

$$\mathcal{C}(x) = \frac{ig}{2} \int d^d \mathbf{y} \left\langle \chi(x) c(x) \bar{c}(y) \times \left[\partial_i B_i(y) + \sqrt{2} g \xi \phi \chi(y) \right] \right\rangle$$

Numerical comparison with conventional approach

Conventional:
 $0 < \xi < 4$



S_3 to $\mathcal{O}(g^{1/2})$:
 $0 < \xi < 4$

Tunneling @ $T > 0$: Take Aways

- *For a radiatively-induced barrier, a gauge-invariant perturbative computation of nucleation rate can be performed for S_3 to $\mathcal{O}(g^{-1/2})$ by adopting an appropriate power counting for T in the vicinity of T_{nuc}*
- *Abelian Higgs example generalizes to non-Abelian theories as well as other early universe phase transitions*
- *Remaining contributions to Γ_{nuc} beyond $\mathcal{O}(g^{-1/2})$ in S_3 and including long-distance (nucleation scale) contributions require other methods*
- *Assessing numerical reliability will require benchmarking with non-perturbative computations*

IR Problem

EWPT & Perturbation Theory: IR Problem

Bosonic loop at $T > 0$

$$I(T) = g^2 \int \frac{d^3 p}{(2\pi)^3} f_B(E, T) \frac{1}{(p^2 + m^2)^n} \xrightarrow{\text{Bose dist fn}} \boxed{\frac{g^2 T}{m}} \int_{\text{I.R.}} \frac{d^3 p}{(2\pi)^3} \frac{1}{(p^2 + m^2)^n}$$

Small p regime

$$f_B(E, T) \longrightarrow \frac{T}{m}$$

Effective expansion parameter

Field-dependent thermal mass


$$m^2(\varphi, T) \sim C_1 g^2 \varphi^2 + C_2 g^2 T^2 \equiv m_T^2(\varphi)$$

- Near phase transition: $\varphi \sim 0$
- $m_T(\varphi) < g T$

EWPT & Perturbation Theory

Expansion parameter

$$g_{\text{eff}} \equiv \frac{g^2 T}{\pi m_T(\varphi)}$$



*Infrared sensitive
near phase trans*

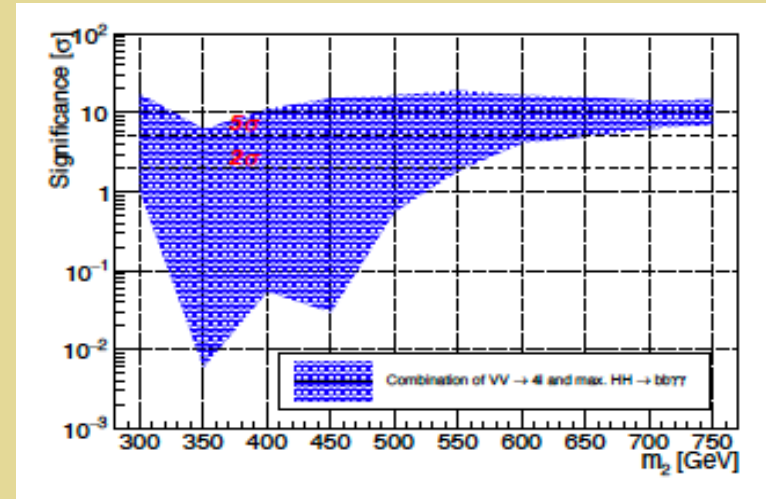
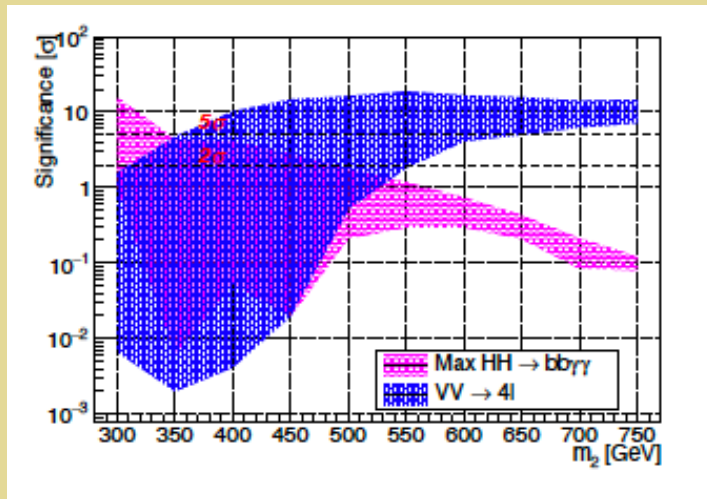
***SM lattice studies: $g_{\text{eff}} \sim 0.8$ in vicinity of EWPT for
 $m_H \sim 70 \text{ GeV}$ ****

** Kajantie et al, NPB 466 (1996) 189; hep/lat 9510020 [see sec 10.1]*

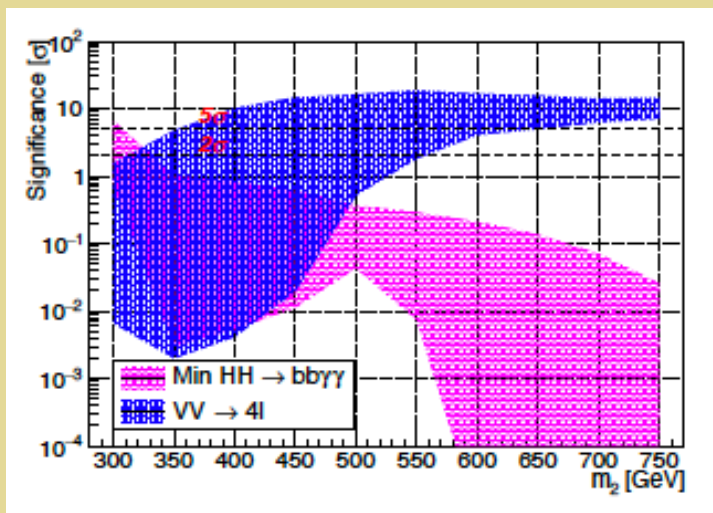
Additional Pheno

Singlets: Resonant Di-Higgs & $H_2 \rightarrow VV$

SFOEWPT Max Benchmarks: HL LHC Combination $bb\gamma\gamma$ & 4 lepton

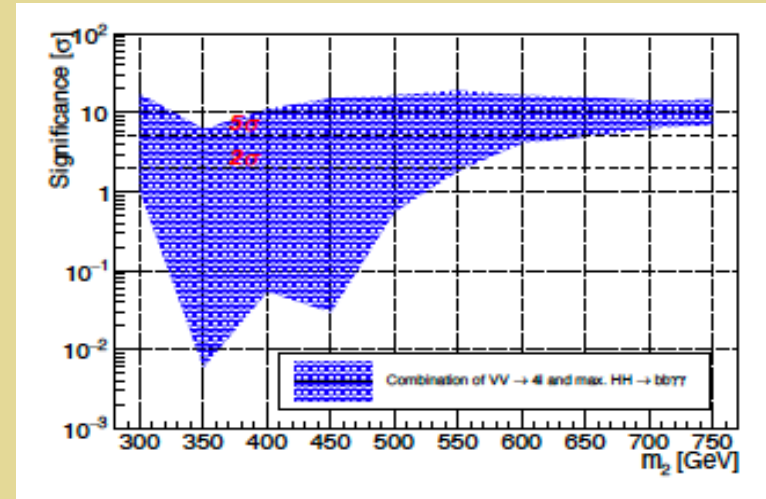
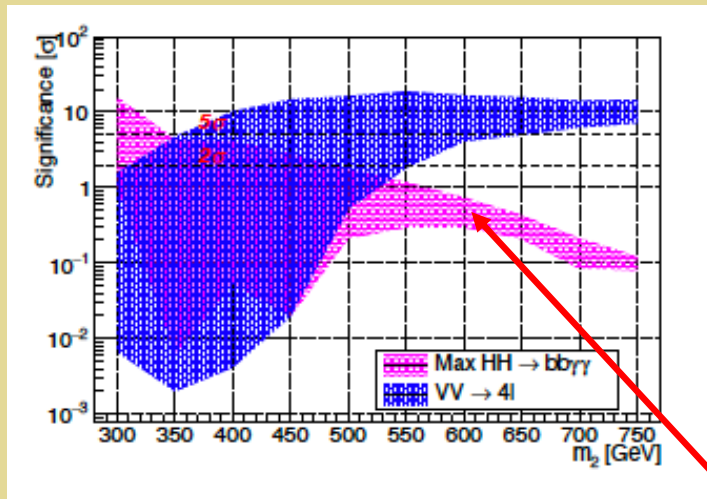


SFOEWPT Min Benchmarks:



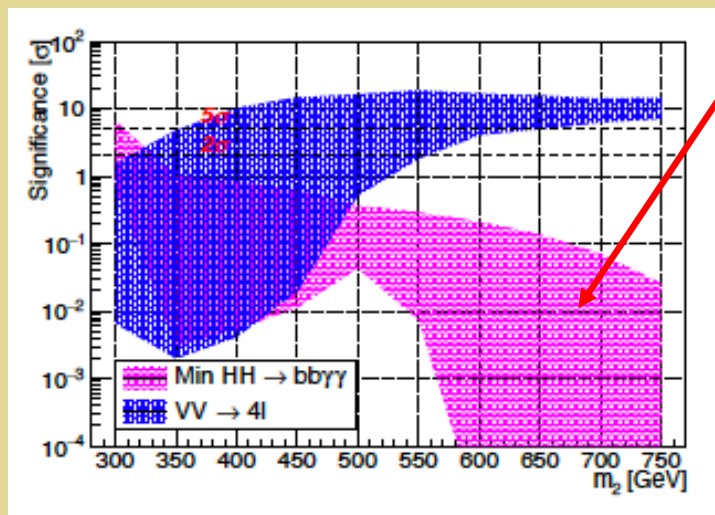
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100 TeV accessible

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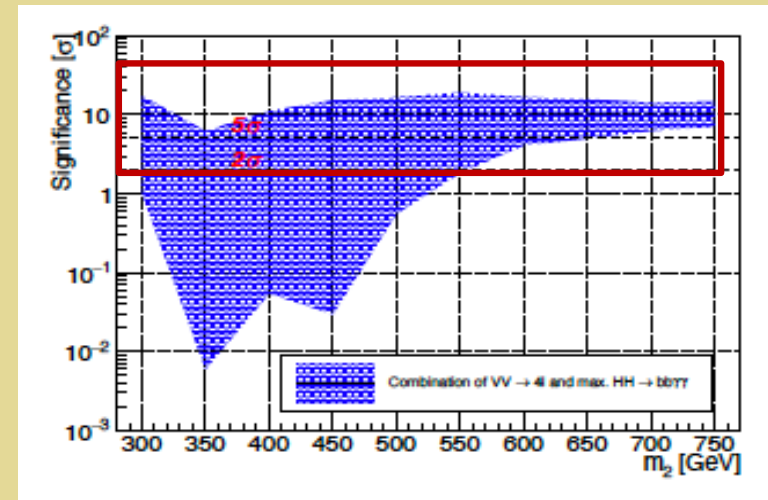
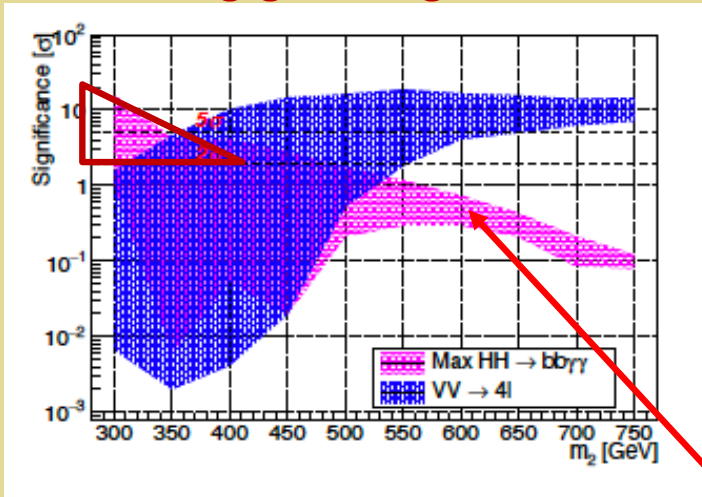


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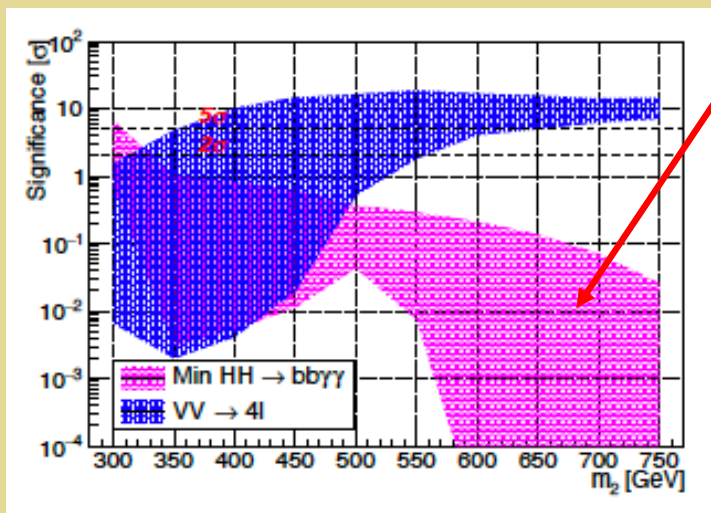
“Smoking gun” region

Parameter exclusion region



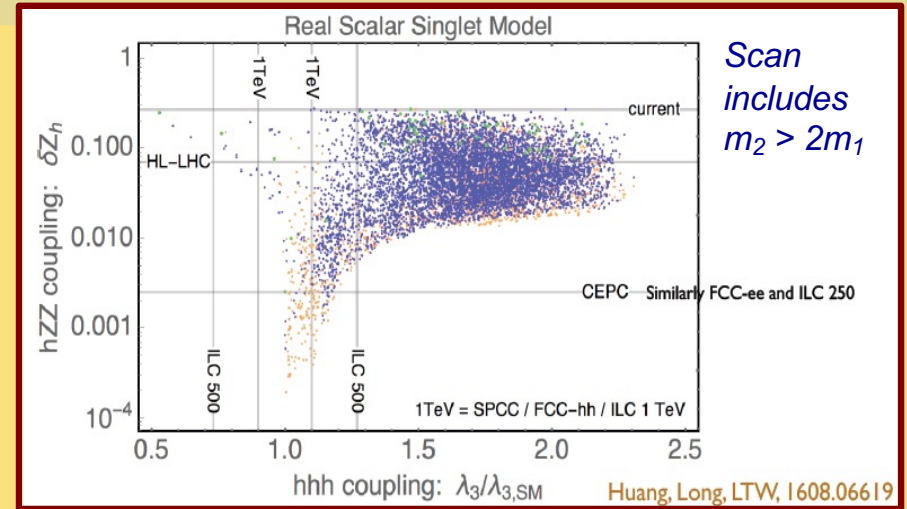
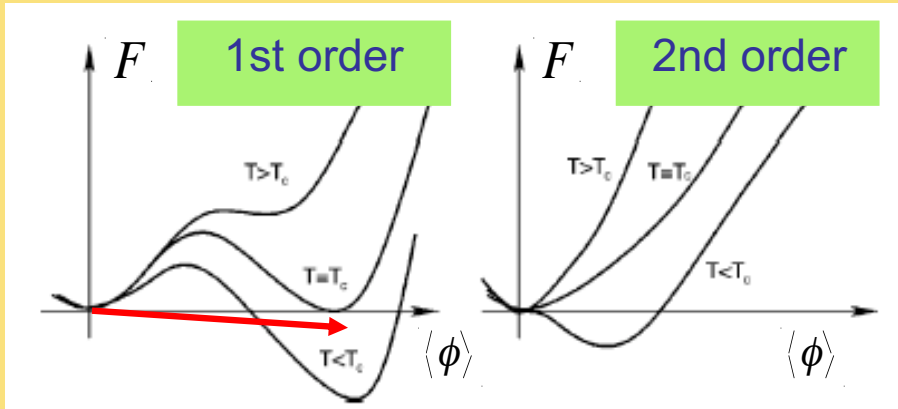
100 TeV accessible

SFOEWPT Min Benchmarks:

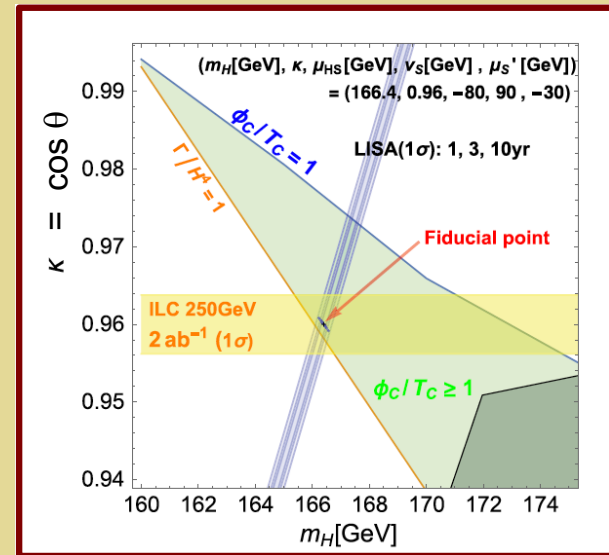
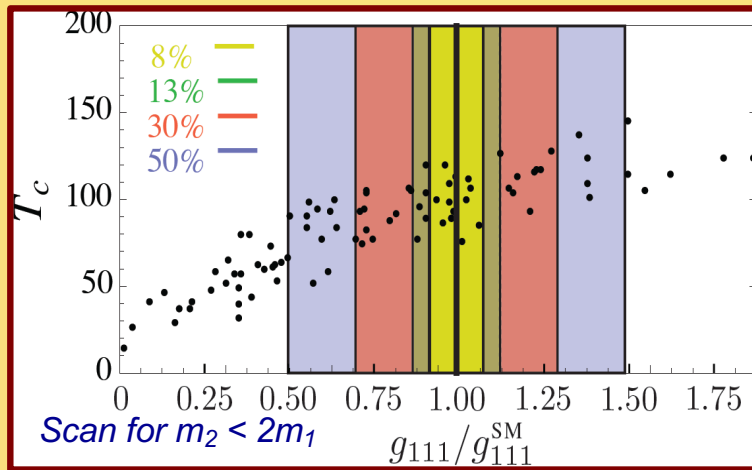


- Observation of $4l$ channel would indicate existence of heavy resonance consistent with xSM SFOEWPT
- “Smoking gun” region would provide nearly definitive evidence & narrow down model parameter space
- Exclusion would leave ample room for 100 TeV pp discovery

EW Phase Transition: Singlet Scalars



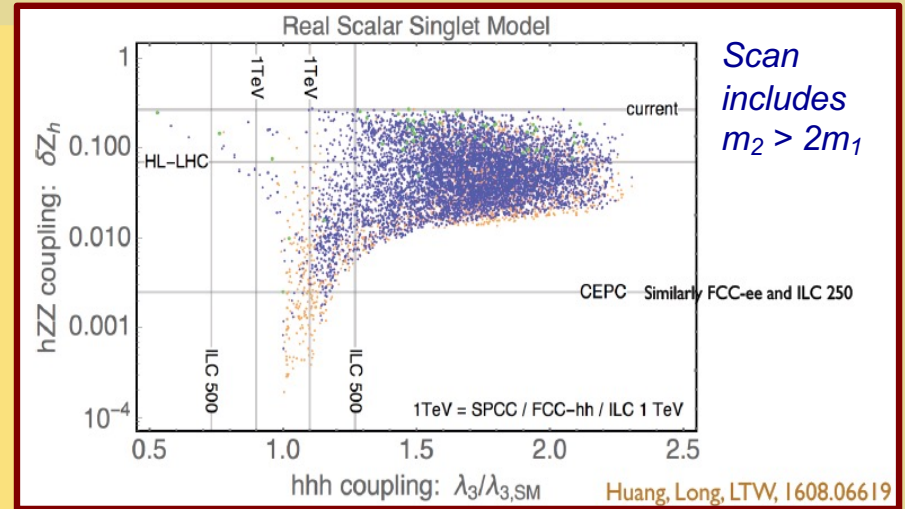
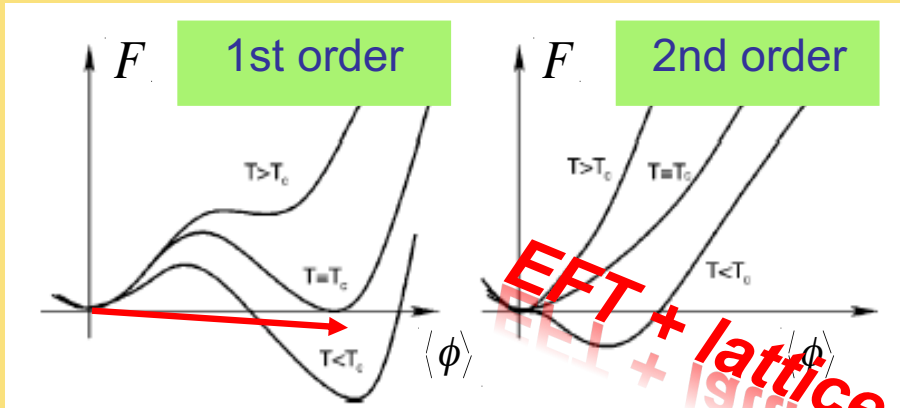
Modified Higgs Self-Coupling



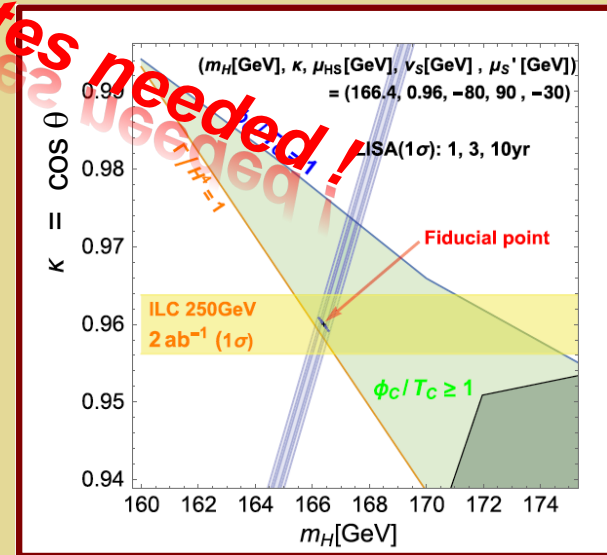
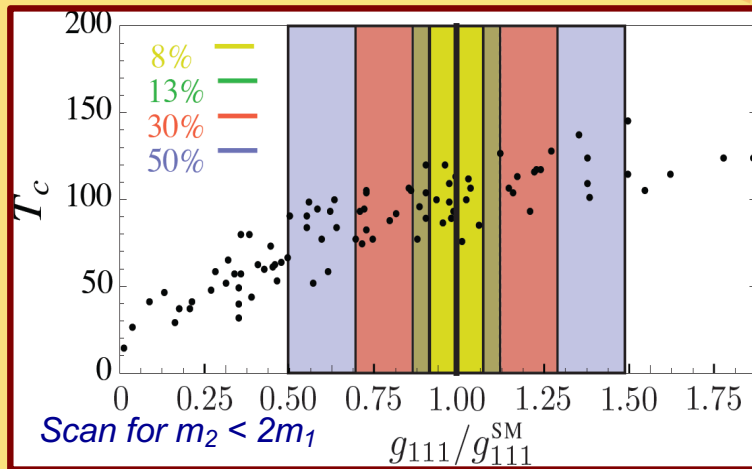
Profumo, R-M, Wainwright, Winslow: 1407.5342; see also Noble & Perelstein 0711.3018

K. Hasino et al, PRD 99 (2019) 075011

EW Phase Transition: Singlet Scalars



Modified Higgs Self-Coupling



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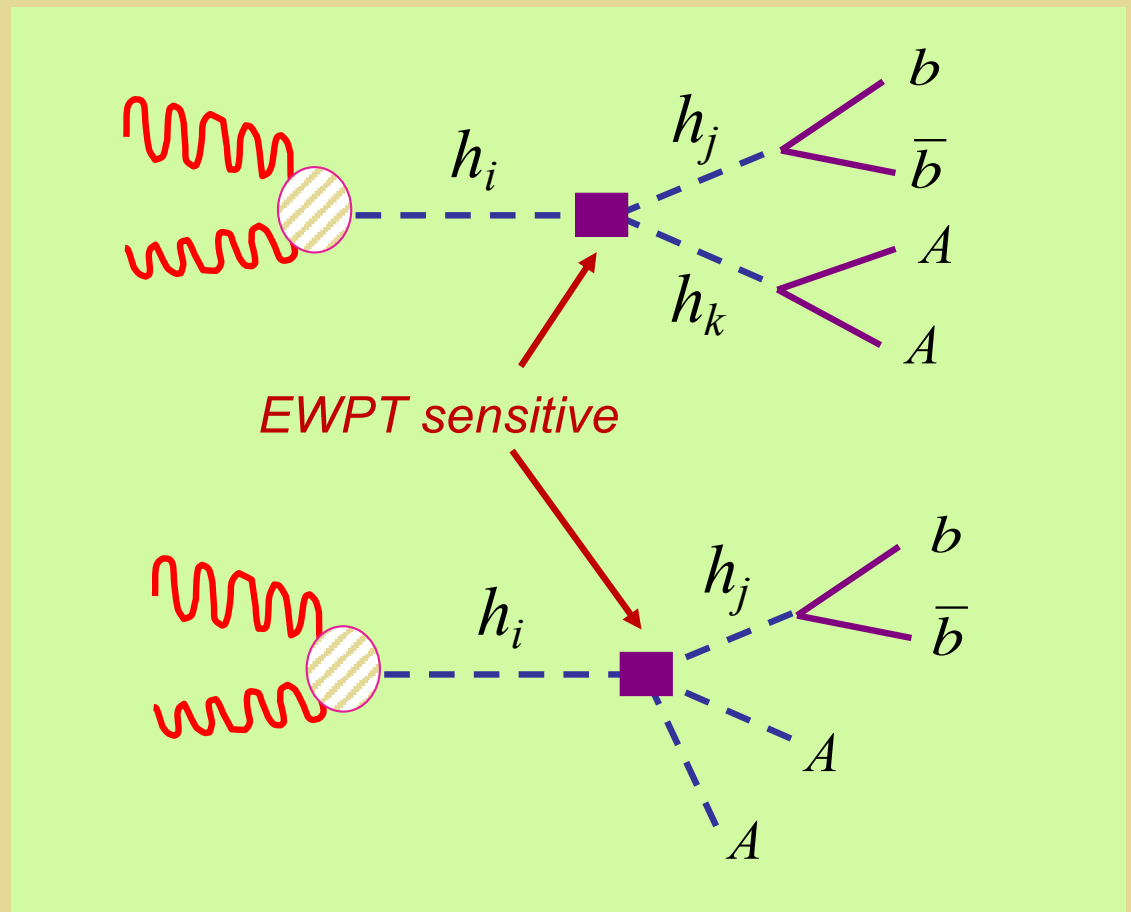
44.2

Complex Singlet: DM + EWPT

Original Model:

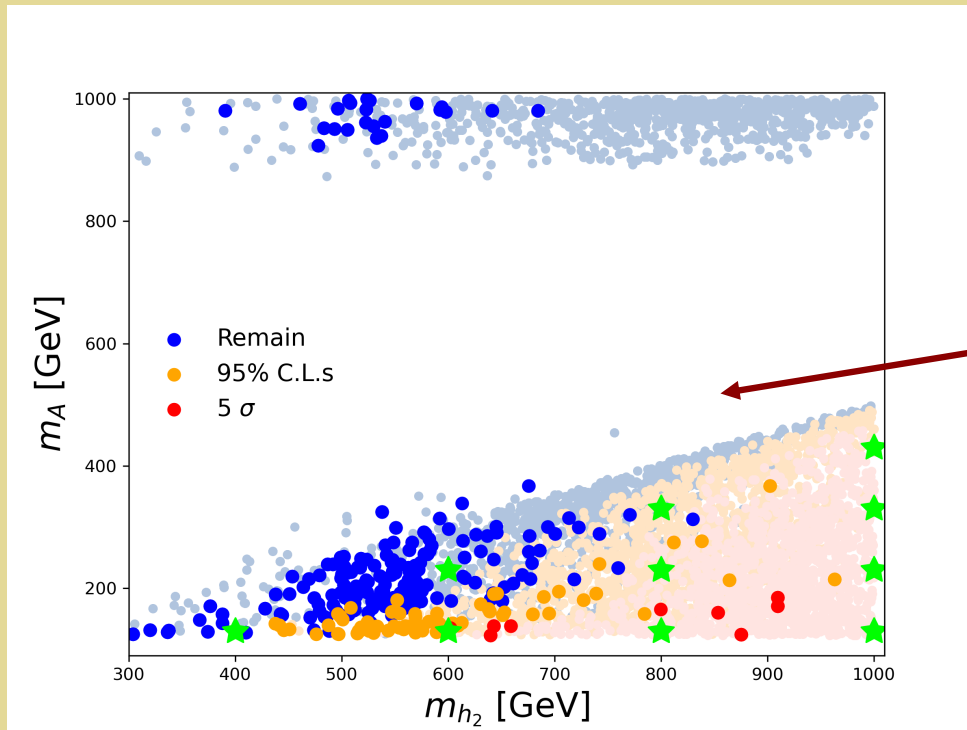
- SM + complex scalar singlet
- Global $U(1)$: broken spontaneously & softly
- Particle spectrum
 - Mixed doublet-singlet scalars $h_{1,2}$
 - Scalar dark matter A

Search for $bb + MET$: example sub-processes

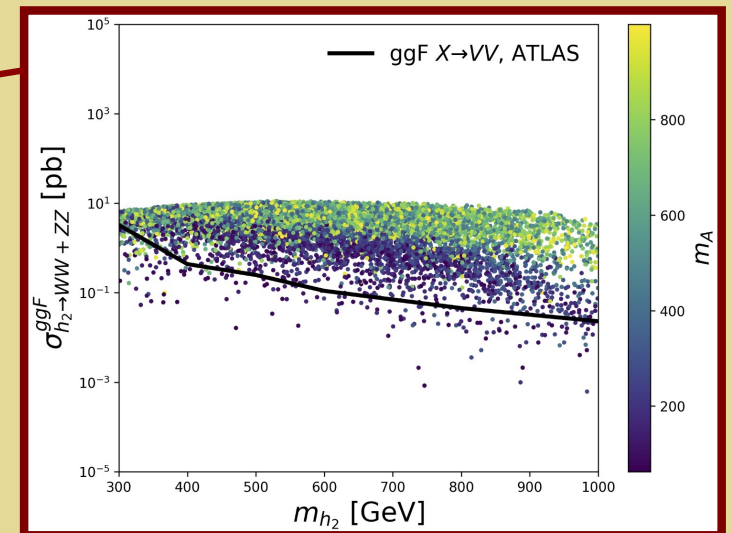


Complex Singlet: DM + EWPT

Search for $bb + MET$

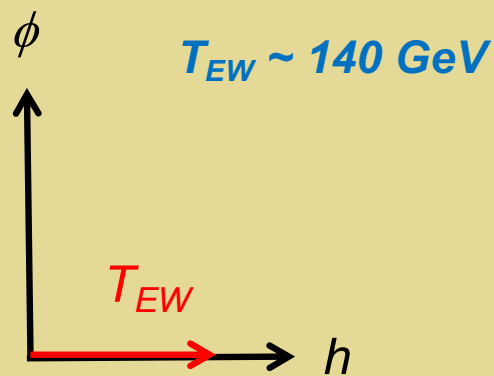


Heavy Higgs $\rightarrow VV$
exclusion: $BR(h_2 \rightarrow VV)$
larger when $m_{h_2} < 2 m_A$

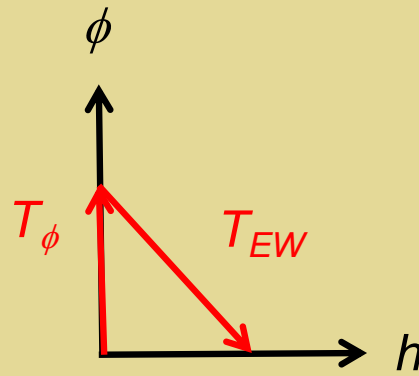


Yizhou Cai, MJRM, Lei Zhang,
Wenxing Zhang 2311.NNNNN

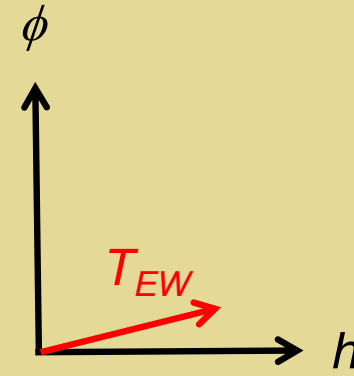
Exotic Higgs Decays & EWPT



$a_2 H^2 \phi^2 : T > 0$
loop effect



$a_2 H^2 \phi^2 : T = 0$
tree-level effect



$a_1 H^2 \phi : T = 0$
tree-level effect

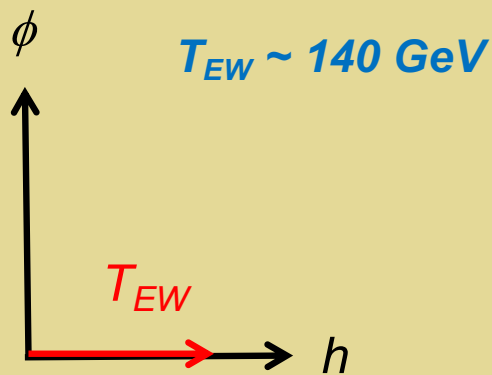
$$g_{122} = \frac{1}{2} v a_2 + \mathcal{O}(\theta^2)$$

Exotic decays

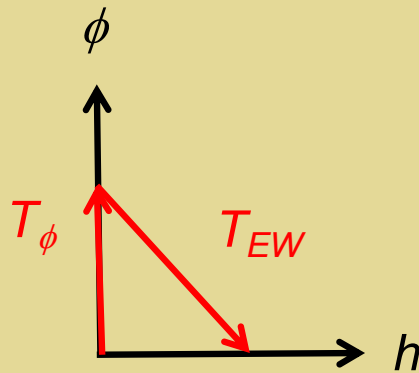
$$h_1 \rightarrow h_2 h_2$$

$$\Gamma(h_2, m_2) = \sin^2 \theta \Gamma(h_{\text{SM}}, m_2)$$

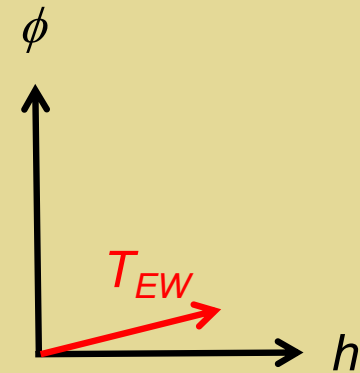
Exotic Higgs Decays & EWPT



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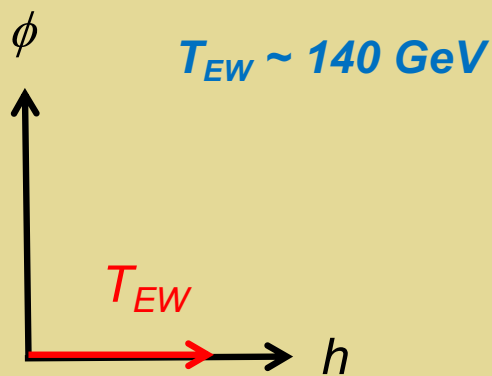
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Exotic decays

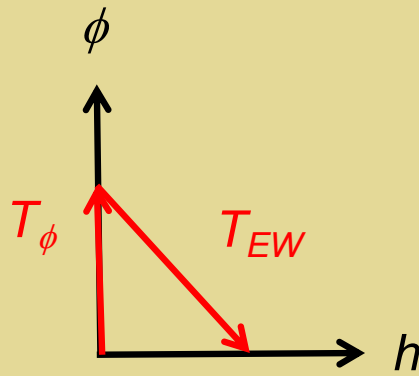
$$h_1 \rightarrow h_2 h_2$$

$$\Gamma(h_2, m_2) = \sin^2 \theta \Gamma(h_{SM}, m_2)$$

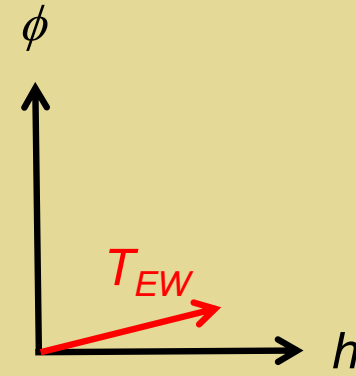
Exotic Higgs Decays & EWPT



$a_2 H^2 \phi^2 : T > 0$
loop effect



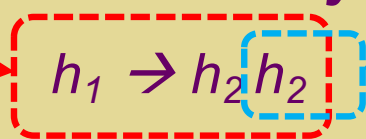
$a_2 H^2 \phi^2 : T = 0$
tree-level effect



$a_1 H^2 \phi : T = 0$
tree-level effect

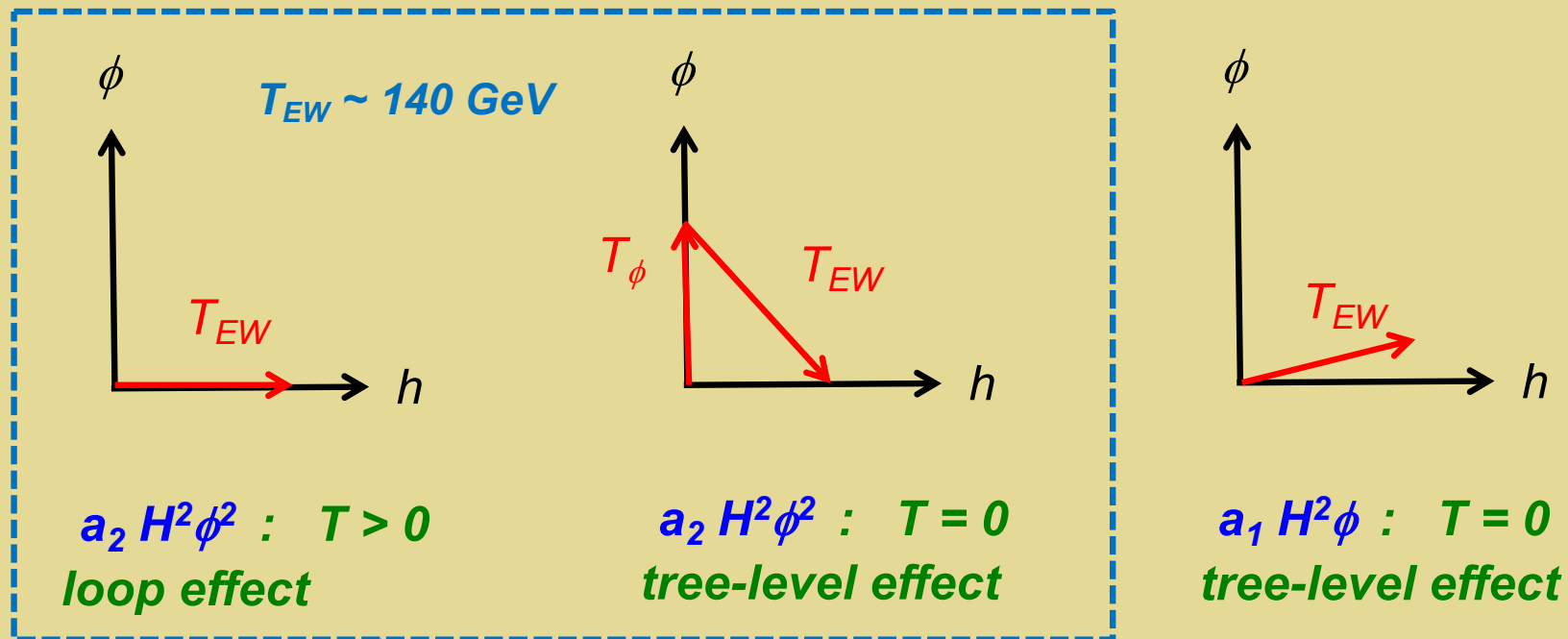
$$g_{122} = \frac{1}{2} v a_2 + \mathcal{O}(\theta^2)$$

Exotic decays



$$\Gamma(h_2, m_2) = \sin^2 \theta \Gamma(h_{\text{SM}}, m_2)$$

Exotic Higgs Decays & EWPT



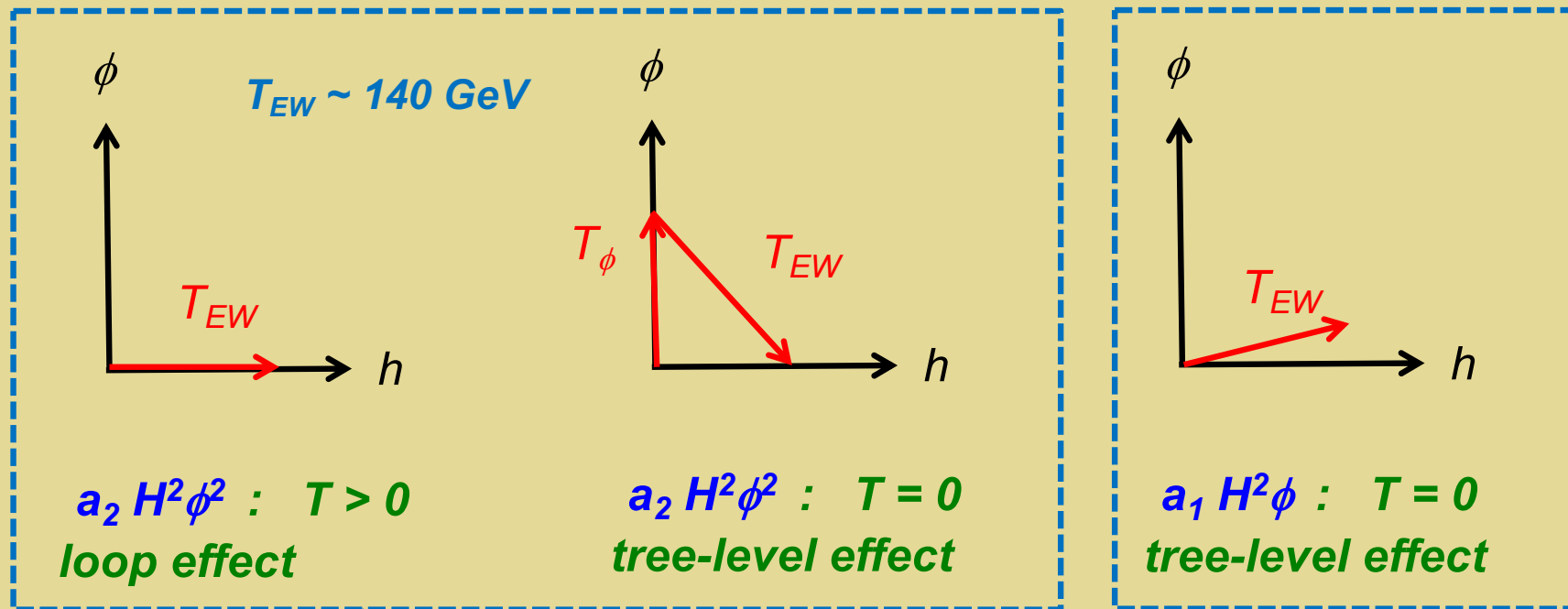
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Exotic Higgs Decays & EWPT



$$g_{122} = \frac{1}{2} v a_2 + \mathcal{O}(\theta^2)$$

Exotic decays

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Z_2 breaking 46.5

Spont Explicit