



Diving deep into the (N)MSSM potential: ML-enhanced analysis of vacuum stability

Work in collaboration with Fabio Campello and Georg Weiglein

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Vacuum stability: SM vs. Susy





Supersymmetry



[1504.07217]



- **Idea:** In models with extended scalar sectors the EW vacuum might not correspond to the global minimum of the potential.
 - \rightarrow The EW vacuum is not stable
 - \rightarrow If the EW vacuum is $\textbf{short-lived}^1$ a parameter point is unphysical
 - \rightarrow Constraints on the parameter space of the model

Vacuum-stability constraints in the MSSM and the NMSSM We improve the computation of vacuum decay rates using neural networks

¹in comparison to the age of the Universe



Step 1: Is the EW vacuum meta-stable?

 \rightarrow Determination of all stationary points of the scalar potential below the EW minimum.





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Step 2: What is the lifetime of the EW vacuum? \rightarrow Compute the vacuum decay rates for each possible transition

$$\frac{\Gamma}{V} = K \mathrm{e}^{-S_E}$$

[Coleman, 1977]

 S_E : Euclidean bounce action



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Step 2: What is the lifetime of the EW vacuum?

 \rightarrow Compute the vacuum decay rates for each possible transition

Step 3: Is the EW vacuum sufficiently long-lived?

 \rightarrow Compare the smallest inverse decay rate (lifetime) to the age of the universe

 $S_E < 390$: Short-lived EW vacuum, unphysical $390 < S_E < 440$: Uncertain fate of the EW vacuum $S_E > 440$: Long-lived (meta-stable) EW vacuum, physical

[W. Hollik, G. Weiglein, J. Wittbrodt: 1812.04644]

Homotopy continuation method



Solutions of a system of polynomial equations can be found using homotopy continuation methods. \rightarrow Can be used to determine all stationary points $\vec{F}(\vec{z}) = 0$ of tree-level scalar potentials

 $\vec{H}(\vec{z},t) = (1-t)\cdot\vec{F}(\vec{z}) + t\cdot\vec{g}(\vec{z}) = 0 \ , \quad \text{solutions of } \vec{g}(\vec{z}) = 0 \ \text{are known}.$



We use HOM4PS2 v.2: Polyhedral homotopy continuation [Lee, Li, Tsai, 2008]

The bounce solution: 1 field

The bounce solution is the key object describing the vacuum transition.

[C.L. Wainwright, 1109.4189]

In one dimensions this can be solved using the overshoot-undershoot method.

Shooting methods: [e.g. Stoer, Bulirsch, 1972]

Bounce action:
$$S_E = \frac{2\pi^2}{\Gamma(2)} \int \rho^3 d\rho \left[\frac{1}{2} (\dot{\vec{\phi}}(\rho))^2 + \Delta V(\vec{\phi}(\rho)) \right]$$

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Boundary conds.:

$$\left. \vec{\phi}(\infty) = \vec{\phi}_F \right., \quad \left. \frac{d\vec{\phi}}{d\rho} \right|_{\rho=0} = 0$$

Bounce solution: Multi-field case



Important obstacle in many dimensions is finding the correct tunneling path in field space.

Method 1: Straight path approximation

Compute the bounce action assuming that the true tunneling path is well approximated by a straight path connecting the true and false minima.

Method 2: Path deformation algorithm

Iterative procedure that deforms the straight path into the correct path by minimizing *perpendicular forces* along the path (cosmoTransitions).

Method 3: Using neural network (NN) to solve differential equations

Transform the problem of solving a system of differential equations into a minimization problem. Then use NN to minimize the loss function.

Methods 2 and 3 now implemented into a new version of EVADE.

Many public codes: cosmoTransitions, AnyBubble, FindBounce, SimpleBounce, BSMPT-v.3, ...

Neural-network method

Bounce equations are a system of coupled differential equations:

$$\frac{d^2\vec{\phi}}{d\rho^2} + \frac{3}{\rho}\frac{d\vec{\phi}}{d\rho} - \vec{\nabla}_{\phi}V(\vec{\phi}) = 0 \qquad \text{BC: } \vec{\phi}_B(\infty) = \vec{\phi}_{\text{false}} \quad \text{and} \quad \frac{d\vec{\phi}_B}{d\rho} \bigg|_{\rho=0} = 0$$

Loss function L for neural network with set of weights and biases $\{w,b\}$, simplest way: Loss = "squared sum" of discretized diff. eqs. + squared BC

$$\mathcal{L}(\{w,b\}) = \sum_{i=1}^{n_{\rho}} \left(\left. \frac{d^2 \phi}{d\rho^2} \right|_{\rho_i} + \frac{3}{\rho_i} \left. \frac{d\phi}{d\rho} \right|_{\rho_i} - \nabla V(\phi(\rho_i)) \right)^2 + n_{\rho} \left[\left(\left. \frac{d\phi}{d\rho} \right|_{\rho=0} \right)^2 + \left(\phi(\rho_{\max}) - \phi_{\text{false}} \right)^2 \right]$$
[Piscopo, Spannowsky, Waite, 1902.05563]

Neural network: Trained to minimze \mathcal{L} (adam optimizer)

One neuron in input layer (ρ) and $N = \dim(\vec{\phi})$ neurons in output layer Depending on model 3 to 5 inner layers with 20 to 50 neurons each CPU and GPU implementation (tensorflow, CUDA, XLA)



[Callan, Coleman, Curtis, 1977]

Toy model: 2 singlet fields



$$V_{ ext{toy}} = \sum_{i=1}^{N=2} \lambda_i \phi_i^4 + A_i \phi_i^3 - m_i^2 \phi_i^2$$

Straight-path approximation \rightarrow Path-deformation algorithm \rightarrow Neural-network approach \rightarrow

Only the neural-network approach is able to determine the correct tunneling path

	B_1	B_2	B_3	$\mathcal{L}_{ m bc}$
Neural network	107.24	107.26	107.28	0.0915
Straight path	47.78	47.57	47.37	0.8215
Path deformation	-15332	3060	21452	73570

\mathcal{L}_{bc} : quality measure (smaller is better)



[TB, F. Campello, G. Weiglein, tbp]

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If $\mathcal{L}_i = \mathcal{L}(\rho_i) > 0$, the predicted solution $\vec{\phi}(\rho)$ does not satisfy the bounce equations at $\rho = \rho_i$

 \rightarrow Neural network found correct solution



[TB, F. Campello, G. Weiglein, tbp]

Toy model: many singlet fields



Runtime (NN on GPU) Loss Bounce action Straight Cosmo m 1000 10^{1} NN 10^{6} Straight Straight Cormo Number of fields Number of fields Number of fields [TB, F, Campello, G, Weiglein, tbp]

For sizeable number of fields only the neural network is able to determine the bounce solution \rightarrow suitable method to analyze vacuum stability in SUSY models

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The NMSSM



NMSSM: Next-to Minimal Supersymmetric Standard Model: MSSM + singlet





Susy gives a recipe for constructing the scalar potential:

$$V = F + D + V_{\text{soft}}$$



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F-term contributions from the Superpotential:

$$W = \frac{1}{3}\kappa S^{3} + \lambda SH_{u} \cdot H_{d} + y_{t}Q_{L} \cdot H_{u}\bar{t}_{R} + y_{b}H_{d} \cdot Q_{L}\bar{b}_{R} + y_{\tau}H_{d} \cdot L_{L}\bar{\tau}_{R}$$
$$F = \sum_{\phi} |\partial_{x}W|^{2} , \quad \phi \in \{h_{u}^{0}, h_{u}^{+}, h_{d}^{0}, h_{d}^{-}, \tilde{t}_{L}, \tilde{b}_{L}, \tilde{\tau}_{L}, \tilde{\nu}_{L}, \tilde{t}_{R}^{*}, \tilde{b}_{R}^{*}, \tilde{\tau}_{R}^{*}\}$$



Susy gives a recipe for constructing the scalar potential:

$$V = F + D + V_{\text{soft}}$$

D-term contributions from the gauge structure:

$$D = D_{U(1)_Y} + D_{SU(2)_L} + D_{SU(3)_c}$$
$$D_{U(1)_Y} = \frac{g_1^2}{8} \left(\sum_{\phi} Y_{\phi} |\phi|^2\right)^2$$
$$D_{SU(2)_L} = \frac{g_2^2}{8} \sum_{\Phi_i} \sum_{\Phi_j} 2(\Phi_i^{\dagger} \Phi_j) (\Phi_j^{\dagger} \Phi_i) - (\Phi_i^{\dagger} \Phi_i) (\Phi_j^{\dagger} \Phi_j)$$
$$D_{SU(3)_c} = \frac{g_3^2}{6} \left(|\tilde{t}_L|^2 - |\tilde{t}_R|^2 + |\tilde{b}_L|^2 - |\tilde{b}_R|^2\right)^2$$



Susy gives a recipe for constructing the scalar potential:

$$V = F + D + V_{\text{soft}}$$

Contributions from soft Susy breaking:

$$\begin{split} V_{\text{soft}} &= m_S^2 s^{\dagger} s + m_{H_u}^2 h_u^{\dagger} h_u + m_{H_d}^2 h_d^{\dagger} h_d + \left(A_{\lambda} s h_u \cdot h_d + \frac{1}{3} A_{\kappa} s^3 + \text{h.c.} \right) \\ &+ m_{Q_3}^2 \tilde{Q}_L^{\dagger} \tilde{Q}_L + m_{L_3}^2 \tilde{L}_L^{\dagger} \tilde{L}_L + m_{U_3}^2 |\tilde{t}_R|^2 + m_{D_3}^2 |\tilde{b}_R|^2 + m_{E_3}^2 |\tilde{\tau}_R|^2 \\ &+ \left(y_t A_t \tilde{t}_R^* \tilde{Q}_L \cdot h_u + y_b A_b \tilde{b}_R^* h_d \cdot \tilde{Q}_L + y_\tau A_\tau \tilde{\tau}_R^* h_d \cdot \tilde{L}_L + \text{h.c.} \right) \,, \end{split}$$



Susy gives a recipe for constructing the scalar potential:

$$V = F + D + V_{\text{soft}}$$

Physical EW vacuum (the one we want to be in):

$$H_d = \begin{pmatrix} (v_d + h_d + ia_d)/\sqrt{2} \\ h_d^+ \end{pmatrix}, \quad H_d = \begin{pmatrix} h_u^+ \\ (v_u + h_u + ia_u)/\sqrt{2} \end{pmatrix}, \quad s = (v_s + h_s + ia_s)/\sqrt{2}$$

with $v_d, v_u, v_s \in \mathbb{R}$ and $v^2 = v_d^2 + v_u^2 = 246^2 \text{ GeV}^2$

BUT: V is a function in many (field) dimensions with many parameters

 \rightarrow In general there can be several local (dangerous) minima below the EW minimum

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 $M_h^{125}(\tilde{\tau})$ benchmark scenario



Comparison: straight-path approximation \leftrightarrow true tunneling path



[W. Hollik, G. Weiglein, J. Wittbrodt, 1812.04644]



[TB, F. Campello, G. Weiglein, tbp]



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[TB. F. Campello, G. Weiglein, tbp]

EVADE v.2: attempts to find bounce with path-deformation algorithm (CosmoTransitions) if this fails switch to neural-network approach

Conclusions



Analysis of vacuum is difficult but worth the effort

 \rightarrow constrains the way in which BSM theories might manifest themselves at colliders

Meta-stable EW vacua are in most cases sufficiently long-lived

ightarrow demanding a global EW minimum is too restrictive, especially in the NMSSM

New method for computing bounce solutions using neural networks

- \rightarrow Reliable determination of tunneling path for $\mathcal{O}(10-100)$ fields
- ightarrow Can find the bounce solution in cases where the path-deformation algorithm fails
- \rightarrow Drawback: longer runtime of about 10 s for 10 fields and about 100 s for 50 fields

CPU and GPU implementation of neural-network method in new version of EVADE

A logo for EVADE



ChatGPT-4: "graphic design is my passion"



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Benchmark scenarios



BP	$\tan \beta$	μ	M_A	m_{L_3,e_3}	X_t	$A_{ au}$	A
${ m M}_h^{125}$	[0, 60]	1000	[0, 2000]	2000	2800	$A_{\tau} = A$	$A(X_t, \mu, \tan \beta)$
$\mathbf{M}_{h}^{125}(\widetilde{\tau})$	[0, 60]	1000	[0, 2000]	350	2800	800	$A(X_t, \mu, \tan \beta)$
$\mathcal{M}_{h}^{125}(A)$	20	[-5000, 5000]	1500	2000	$X_t(A,\mu, aneta)$	$A_{\tau} = A$	[-6000, 6000]

Table 2: Parameter values for the M_h^{125} and $M_h^{125}(\tilde{\tau})$ scenarios as defined in Ref. [65] and the $M_h^{125}(A)$ scenario defined in Ref. [33]. The remaining MSSM parameters are set equally in all scenarios: $m_{Q_3,u_3,d_3} = 1500$, $M_{1,2} = 1000$, $M_3 = 2500$. In the first row, we also set $A \equiv A_t = A_b = A_\tau$ and $A = X_t + \mu/\tan\beta$. In the second row we use the same relations with the exception of $A_\tau = 800$. In the third row, A is varied and X_t is derived from the relation given above. In the NMSSM we additionally set $A_\kappa = -100$, $\mu_{\text{eff}} = \mu$ and $\lambda = \kappa = 0.1$, such that the singlet vev v_S varies according the relation shown in Eq. (35). All dimensionful parameters are given in GeV.

Ref. [65]: [E. Bagnaschi et al., 1808.07542]

Ref. [33]: [W. Hollik, G. Weiglein, J. Wittbrodt, 1812.04644]

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