

Interplay of inert doublet and vector-like lepton triplet with displaced vertices at the LHC/FCC and MATHUSLA

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Plan

- Hypercharged zero $SU(2)$ triplet fermion: Type-III seesaw
- Unconventional searches: Displaced decays. Double recoil
- A special case of non-zero hypercharged $SU(2)$ triplet
- Interplay of Vector like inert triplet lepton and inert doublet
- MATHUSLA detector
- Displaced decay signature of the heavy leptons

Tiny neutrino mass

- Simple extension of SM with a right-handed neutrino (N_R) would solve the problem as

$$\mathcal{L}_{m_\nu} = -\frac{y_\nu v}{\sqrt{2}} \bar{\nu}_L N_R + h.c.$$

- This is a Dirac type mass $m_D = \frac{Y_N v}{\sqrt{2}}$

- To have a neutrino mass of ~ 0.1 eV, one needs $Y_N \sim 10^{-12} \implies$ Leads to fine-tuning

- Being charged neutral, a Majorana mass term $M_N \bar{N}_R N_R^c$ is possible

- A popular mechanism called Seesaw explains the small neutrino mass by having heavy Majorana neutrinos mass term

\implies existence of heavy Majorana neutrino

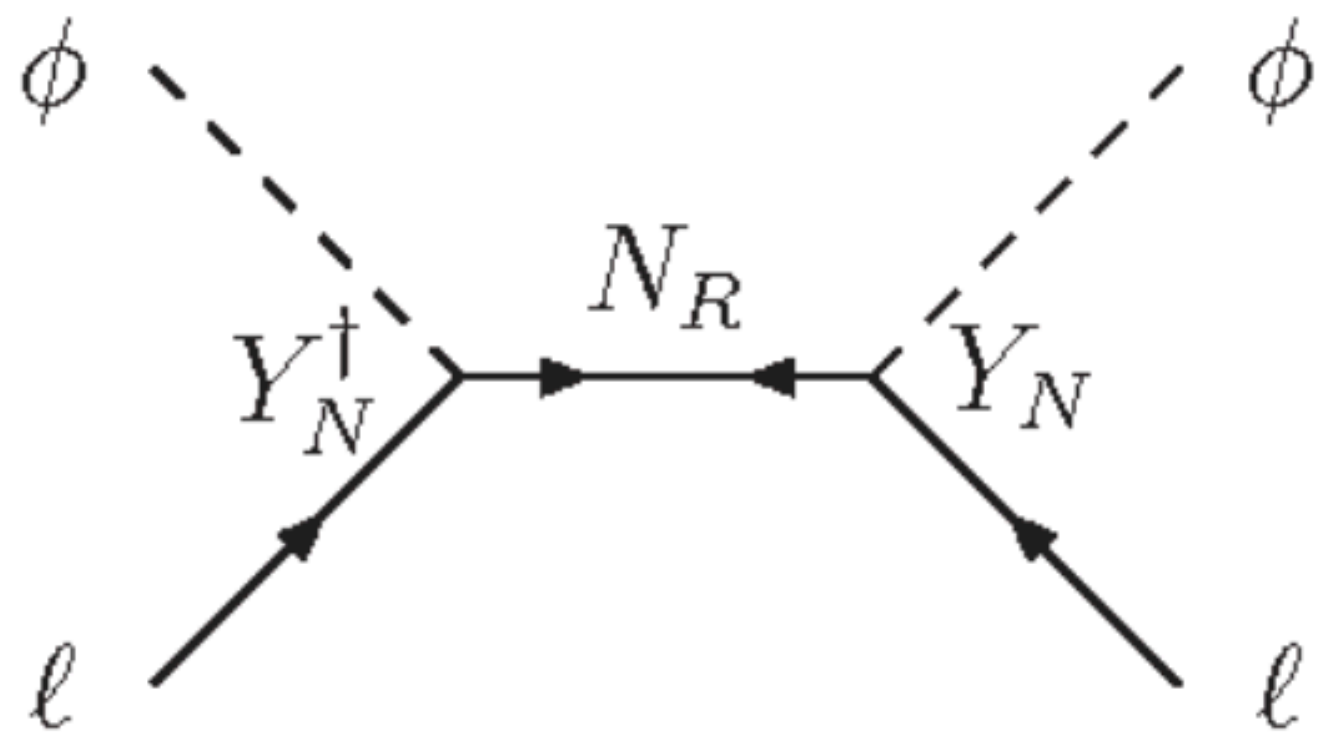


$$m_\nu = \begin{pmatrix} 0 & m_D \\ m_D^T & M_N \end{pmatrix}$$

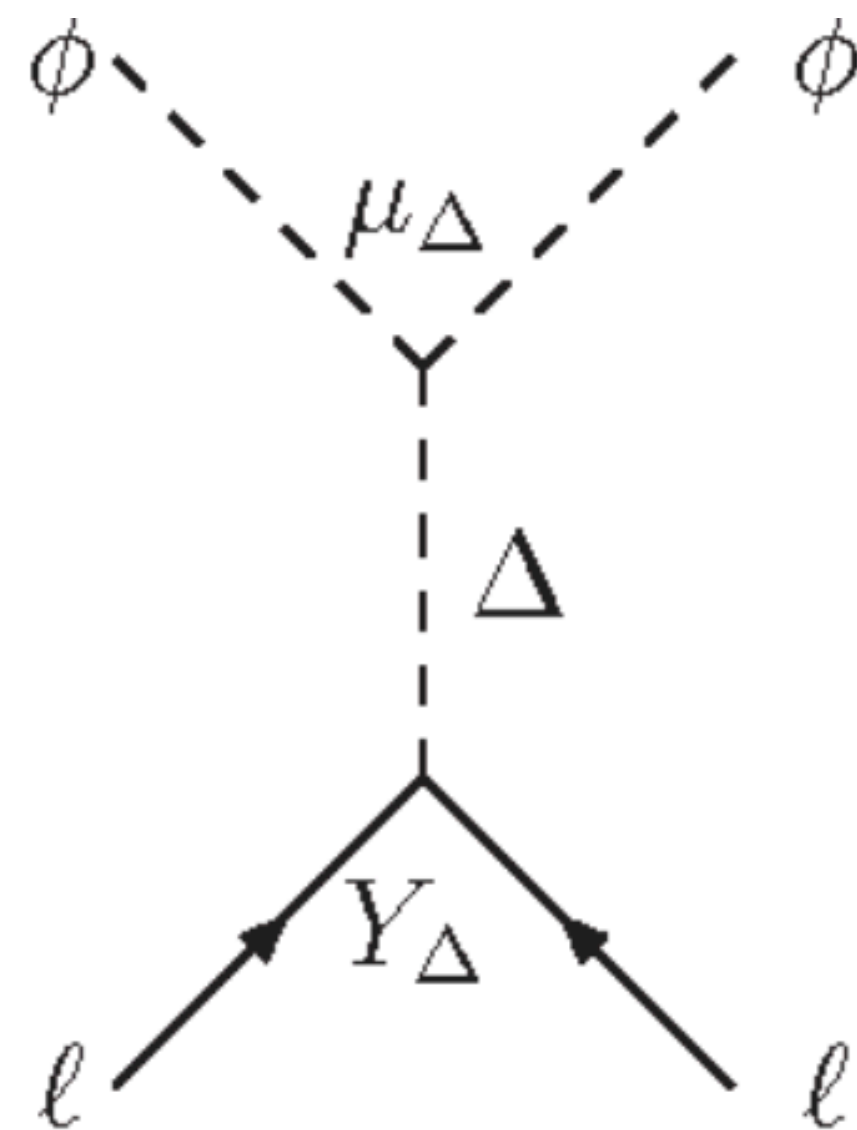
More

Seesaw mechanisms: Most popular scenarios

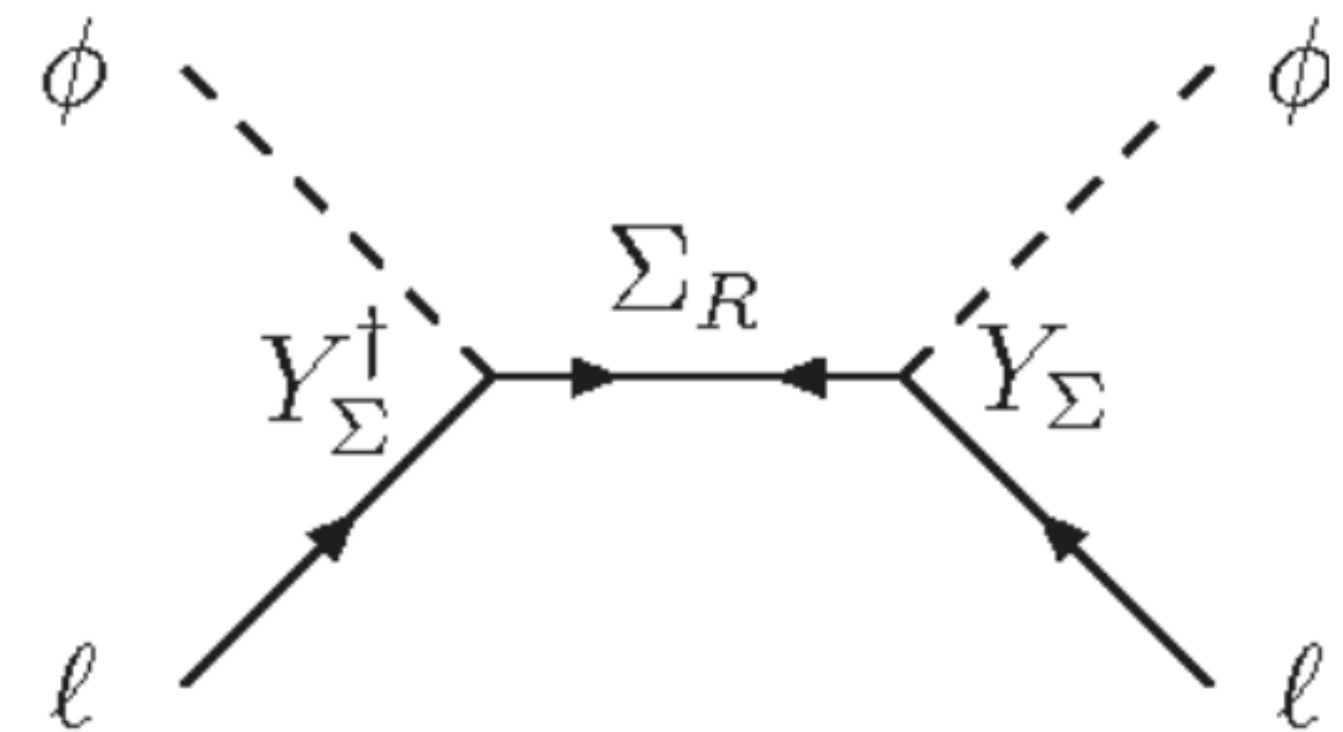
- The most popular variants are Type-I, Type-II and Type-III seesaw
- Type-I and Type-III involve heavy fermions
- Type-II involves heavy scalar



Type-I involves SM gauge singlet heavy neutrino $\nu_R(N_R/N)$



Type-II involves SU(2) triplet scalar Δ



Type-III involves SU(2) triplet fermion $\Sigma(N)$

Type-III Seesaw: Extension with zero a hyper ($Y=0$) charged SU(2) triplet

Type-III Seesaw

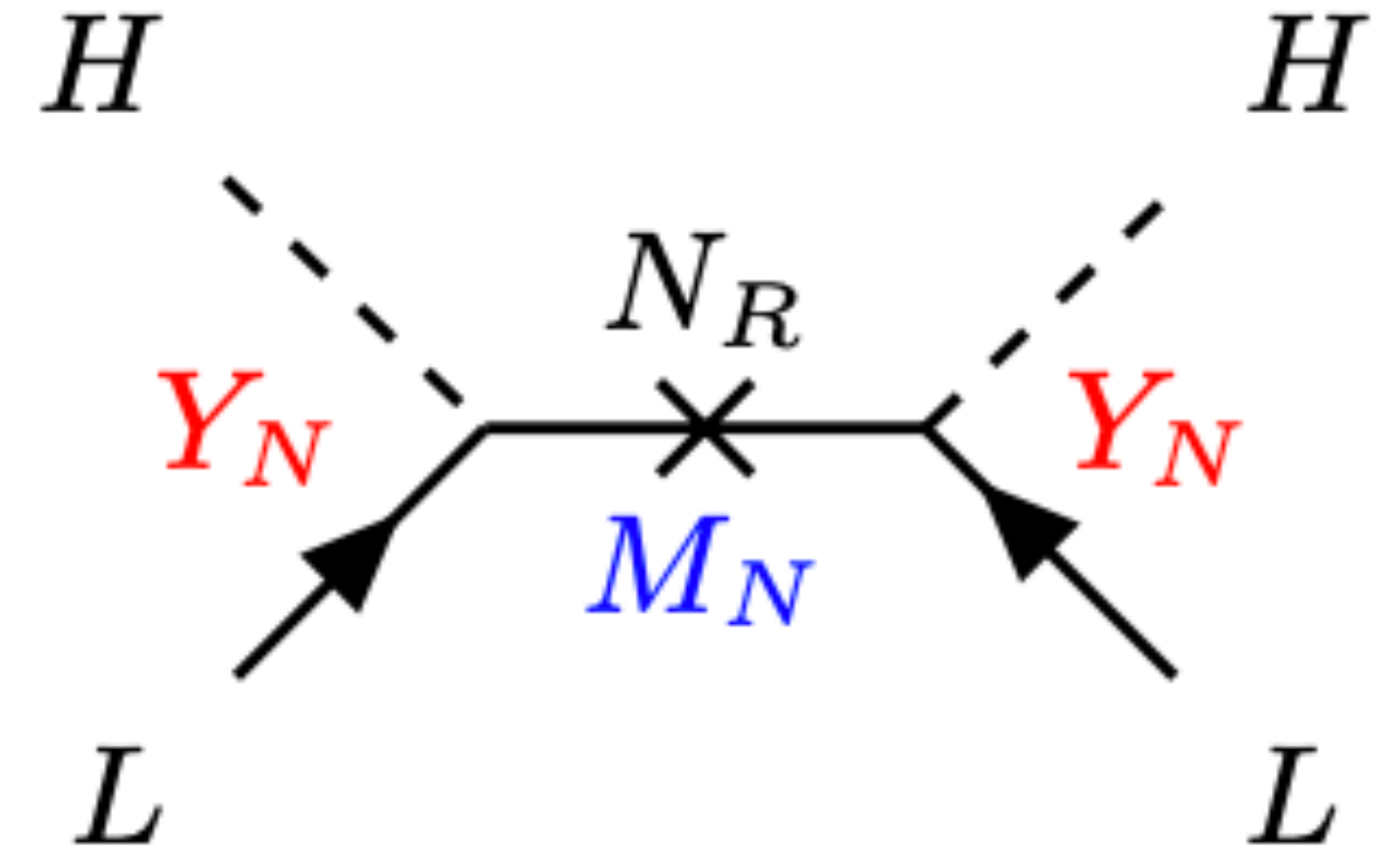
- A $Y=0$, $SU(2)$ triplet fermion can be added which generates tiny neutrino mass

$$\mathcal{L}_N = -\frac{1}{2} \text{Tr}[(\bar{N} M_N N) + \bar{N}^c M^* N] - Y_N \bar{L} N \tilde{\phi} + h.c$$

- A Majorana mass term is possible

The heavy triplet fermions have charged (N^\pm) and neutral (N^0) part as given by

$$N = \begin{pmatrix} N^0 & \sqrt{2}N^+ \\ \sqrt{2}N^- & -N^0 \end{pmatrix}$$



Type-III Seesaw

- For Heavy N the decay widths follows: $\Gamma_{N^\pm}^{Z\ell} \simeq \Gamma_{N^\pm}^{Z\ell} \simeq \frac{1}{2} \Gamma_{N^\pm}^{W^\pm \nu} \simeq \frac{Y_N^2 M_N}{32\pi^2}$
- There is another mode modinant for lower Yukawa , i.e. $N^\pm \rightarrow N^0 \pi^\pm$

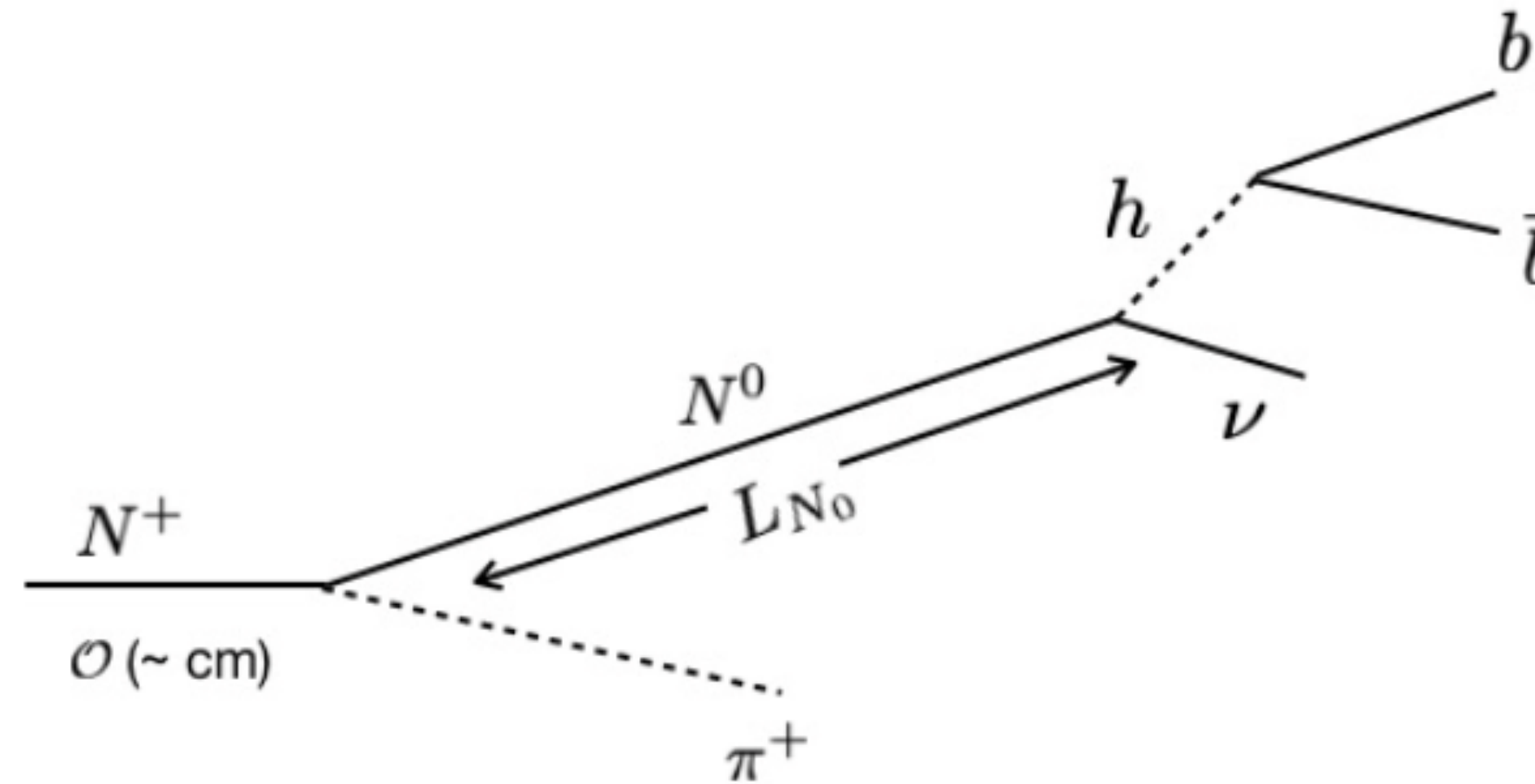
$$\text{with } \Gamma(N^\pm \rightarrow N^0 \pi^\pm) = \frac{2G_F^2 V_{ud}^2 (\Delta M)^3 f_\pi^2}{\pi} \sqrt{1 - \frac{m_\pi^2}{\Delta M^2}}$$

Cirelli et al. :Nucl.Phys.B753:178–194,2006

Where, ΔM is the mass difference between N^\pm and N^0 arising from the quantum corrections which is $\mathcal{O}(166) \text{ MeV}$

- For $Y_N \sim 5 \times 10^{-7}$, the branching ratio in this mode is less than 1 %
- However, for $Y_N \sim 5 \times 10^{-10}$ the branching fraction is $\sim 97.5 \%$

Displaced Double recoil



- For lower Yukawa, i.e. $Y_N \lesssim 5 \times 10^{-8}$, $N^\pm \rightarrow N^0 \pi^\pm$ dominates with a decay length of $\mathcal{O}(cm)$
- For similar low values of Yukawa the decay of $N^0 \rightarrow h\nu$ can be a few meeter to 100s of meter
- A displaced double recoil is predicted.

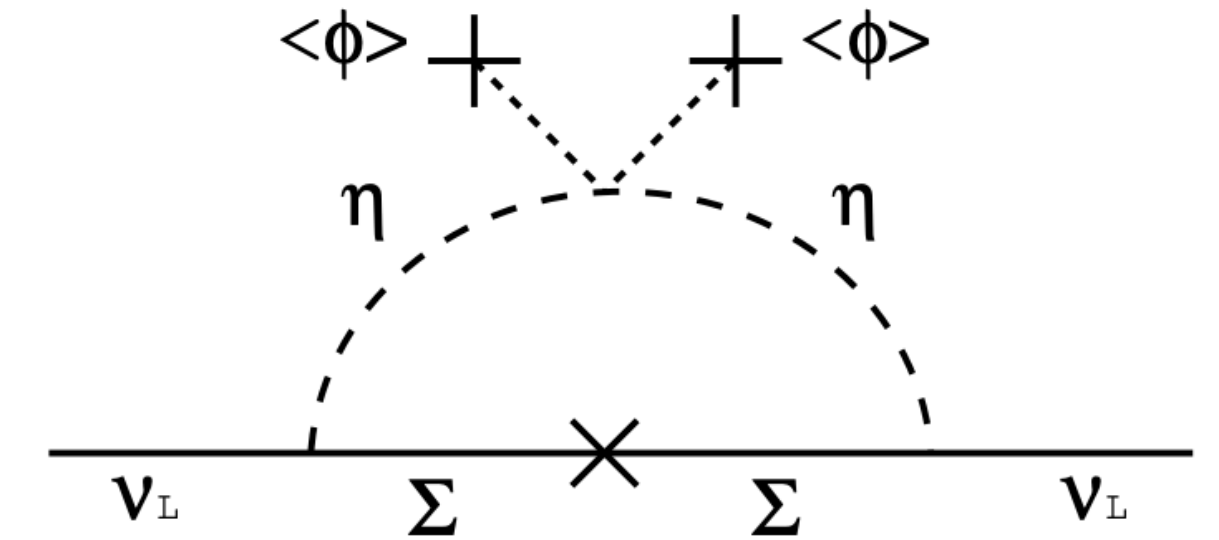
**Scotogenic Type-III: Extension with a zero hyper ($Y=0$) charged Z_2 -odd
SU(2) triplet**

Scotogenic Type-III

- $Y=0$ triplet fermion (Σ) and inert doublet (η) both are Z_2 -odd

$$-\mathcal{L}_\Sigma = \sum_{\alpha=1}^{n_\Sigma} \left(\sum_{i=e,\mu,\tau} h_{i\alpha} \bar{\ell}_{L_i} \Sigma_\alpha \eta + \frac{1}{2} M_\alpha \text{tr}(\bar{\Sigma}_\alpha \Sigma_\alpha^c) + \text{h.c.} \right)$$

- Neutrino mass cannot be generated at the tree-level
- Neutrino mass is generated at one-loop
- A $Y=0$, $SU(2)$ triplet fermion with Z_2 -odd can provide Majorana dark matter
- Due to $Y=0$, the coupling with Z boson is zero
- It can evade the direct dark matter constraint mediated by Z boson
- For $Y=0$, a scotogenic Type-III can provide the neutral component of Σ can be a Majorana dark matter



Extension with a non-zero hyper charged SU(2) triplet

Non-zero Hypercharged fermion

• For $Y \neq 0$, we cannot write a Majorana mass term for $N = \begin{pmatrix} \frac{N^-}{\sqrt{2}} & N^0 \\ N^{--} & \frac{-N^-}{\sqrt{2}} \end{pmatrix}$

• So usual Seesaw will not work

• The gauge invariant Lagrangian also needs a $SU(2)$ doublet scalar $\Phi_2 = \begin{pmatrix} \phi_2^+ \\ \phi_2^0 \end{pmatrix}$

$$\mathcal{L}_{VLL} \supset \left[-\frac{M_N}{2} \overline{N}_L N_R + \mathcal{Y}_N \overline{L}_L^e N_R \Phi_2 \right] + h.c.$$

• We make both N, Φ_2 Z_2 -odd

• Being Z_2 -odd, in principle, neutral component of both (N^0, ϕ_2^0) can provide the dark matter

• Due to $Y \neq 0$, N^0 couples to Z boson

• For $Y \neq 0$, the possible mass term is Dirac Type, which makes it vector like

SU(2) VLL and IDM

Description	Field definition	Gauge charges			
		$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	Z_2
Vectorlike lepton (VLL)	$N = \begin{pmatrix} \frac{N^-}{\sqrt{2}} & N^0 \\ N^{--} & -\frac{N^-}{\sqrt{2}} \end{pmatrix}$	1	3	-1	-
Scalars	$\Phi_1 = (\phi_1^+ \ \phi_1^0)^T$	1	2	1/2	+
	$\Phi_2 = (\phi_2^+ \ \phi_2^0)^T$	1	2	1/2	-

- The scalar potential can be written as

$$V_{\text{scalar}} = -m_{\Phi_1}^2 \Phi_1^\dagger \Phi_1 - m_{\Phi_2}^2 \Phi_2^\dagger \Phi_2 + \lambda_1 (\Phi_1^\dagger \Phi_1)^2 + \lambda_2 (\Phi_2^\dagger \Phi_2)^2 + \lambda_3 (\Phi_1^\dagger \Phi_1) (\Phi_2^\dagger \Phi_2) + \lambda_4 (\Phi_1^\dagger \Phi_2) (\Phi_2^\dagger \Phi_1) + [\lambda_5 (\Phi_1^\dagger \Phi_2)^2 + h.c.]$$

- Being Z_2 odd, Φ_2 does not get vev

SU(2) VLL and IDM

- The physical masses are

$$M_h^2 = 2\lambda_1 v^2$$

$$M_{H^0/A^0}^2 = m_{\Phi_2}^2 + \frac{1}{2}v^2 \lambda_{L/S}$$

$$M_{H^\pm}^2 = m_{\Phi_2}^2 + \frac{1}{2}v^2 \lambda_3, \quad \text{where, } \lambda_{L/S} = \lambda_3 + \lambda_4 \pm 2\lambda_5.$$

- Being Z_2 odd, Φ_2 does not get vev
- Masses of VLLs are all degenerate to M_N at the tree-level
- However, at loop-level there is a mass splitting between the different charged and the neutral components

$$\Delta M_{N^\pm N^0} = \frac{\alpha_2 M_N}{4\pi} \left[(s_W^2 + 1) \mathcal{G} \left(\frac{M_Z}{M_N} \right) - \mathcal{G} \left(\frac{M_W}{M_N} \right) \right],$$

$$\Delta M_{N^{\pm\pm} N^0} = \frac{\alpha_2 M_N}{4\pi} \left[4 s_W^2 \mathcal{G} \left(\frac{M_Z}{M_N} \right) \right]. \quad \text{where, } \mathcal{G}(x) = \frac{x}{2} \left[2x^3 \ln x - 2x + (x^2 + 2)\sqrt{x^2 - 4} \ln \left(\frac{x^2 - 2 - x\sqrt{x^2 - 4}}{2} \right) \right]$$

- For $M_N \geq 400$ GeV, $\Delta M_{N^\pm N^0} \sim 500$ MeV, $\Delta M_{N^{\pm\pm} N^0} \sim 1.5$ GeV

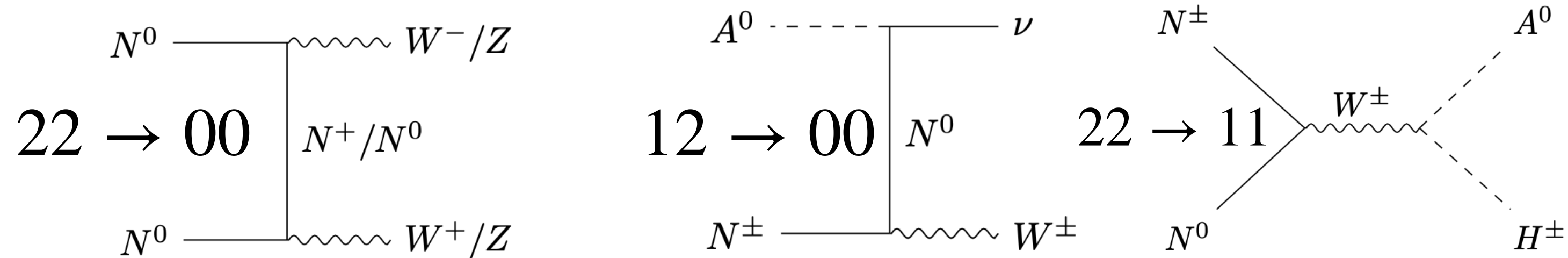
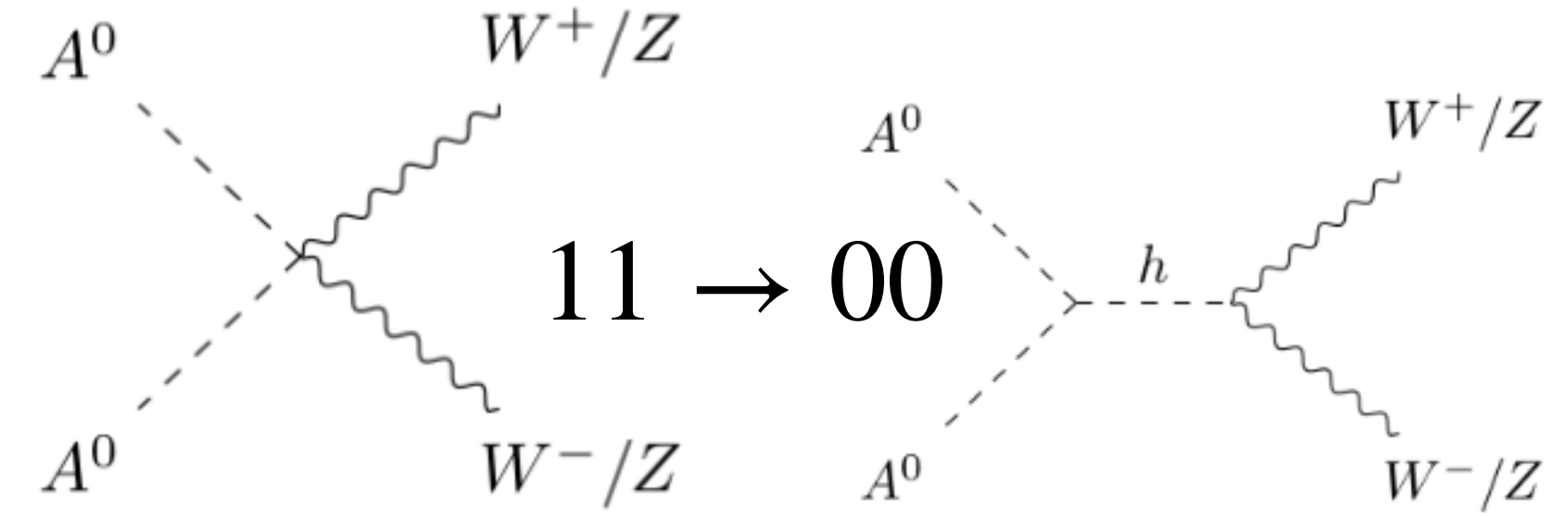
SU(2) VLL and IDM

- Unlike $Y=0$ Schotogenic Type-III, N^0 can not be dark matter here
- Due to non-Majorana nature, its coupling with Z boson forbids N^0 to be the dark matter, owing to large direct dark matter cross-section
- Whereas, A^0/H^0 due to CP-conserving lagrangian do not couple to Z boson. The lighter of these two can be a dark matter candidate, which we choose as A^0
- The direct dark matter constraint even prohibits a multi-component dark matter scenario as well
- So for all the practical purpose, this is IDM scenario
- However, the interplay of N^0 and A^0/H^0 plays an important role in the freeze-out of the dark matter and attaining the correct relic

Interplay between IDM and VLL

Convention: Sector 0: SM particles; Sector 1: IDM: A^0, H^0, H^\pm ; Sector 2: VLL: $N^0, N^\pm, N^{\pm\pm}$

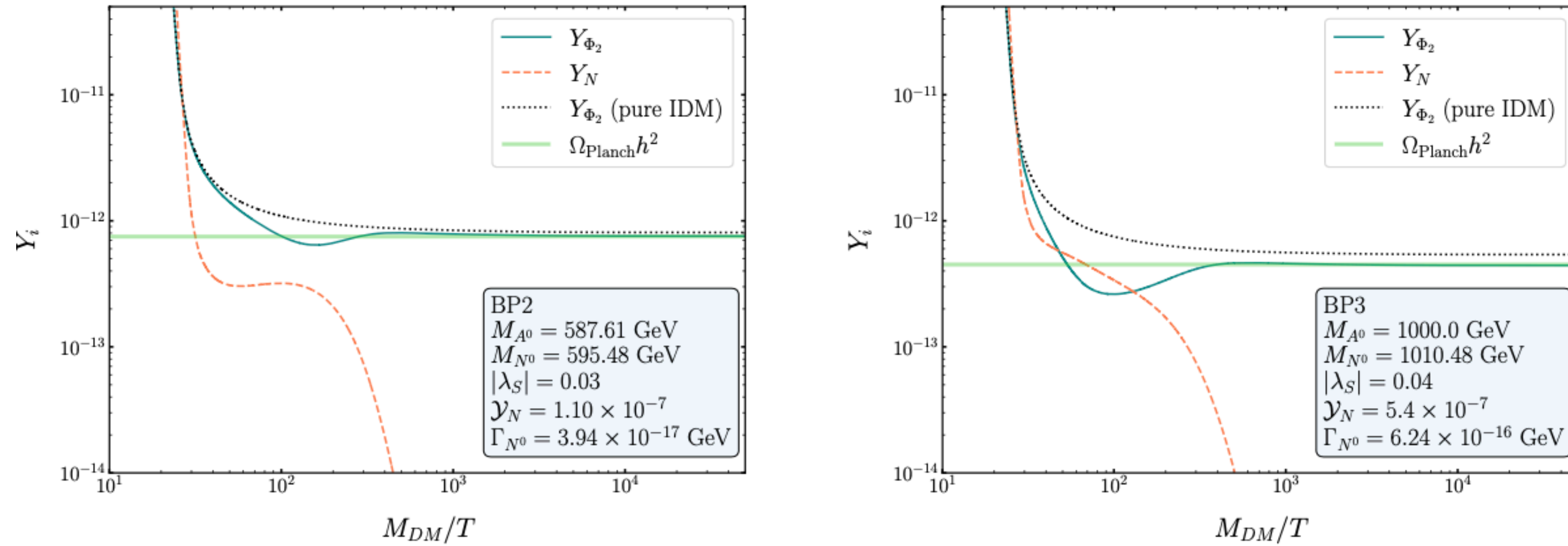
- For $m_{\Phi_2} < M_N$, IDM is the real dark matter
- Pair-annihilation of Z_2 odd particles
 $\Phi_2\Phi_2 \rightarrow \text{SM SM}(11 \rightarrow 00)$ and
 $NN \rightarrow \text{SM SM}(22 \rightarrow 00)$
- Co-annihilations of Z_2 odd particles
 $N\Phi_2 \rightarrow \text{SM SM}(21 \rightarrow 00)$
- Co-scattering of Z_2 odd particles
 $\Phi_2\Phi_2 \leftrightarrow NN$
- Late decay effect: $N \rightarrow \Phi_2 \text{ SM}$



$$\frac{dY_{\Phi_2}}{dx} = -\frac{1}{x^2} \frac{s(M_{A^0})}{H(M_{A^0})} \left[\langle \sigma v \rangle_{1100} \left(Y_{\Phi_2}^2 - (Y_{\Phi_2}^{eq})^2 \right) - \langle \sigma v \rangle_{2211} \left(Y_{\Phi_2}^2 - Y_N^2 \frac{(Y_{\Phi_2}^{eq})^2}{(Y_N^{eq})^2} \right) \right. \\ \left. + \langle \sigma v \rangle_{1200} \left(Y_{\Phi_2} Y_N - Y_{\Phi_2}^{eq} Y_N^{eq} \right) \right] + \frac{x \Gamma_{N \rightarrow \Phi_2 X}}{H(M_{A^0})} \left(Y_N - Y_{\Phi_2} \frac{Y_N^{eq}}{Y_{\Phi_2}^{eq}} \right),$$

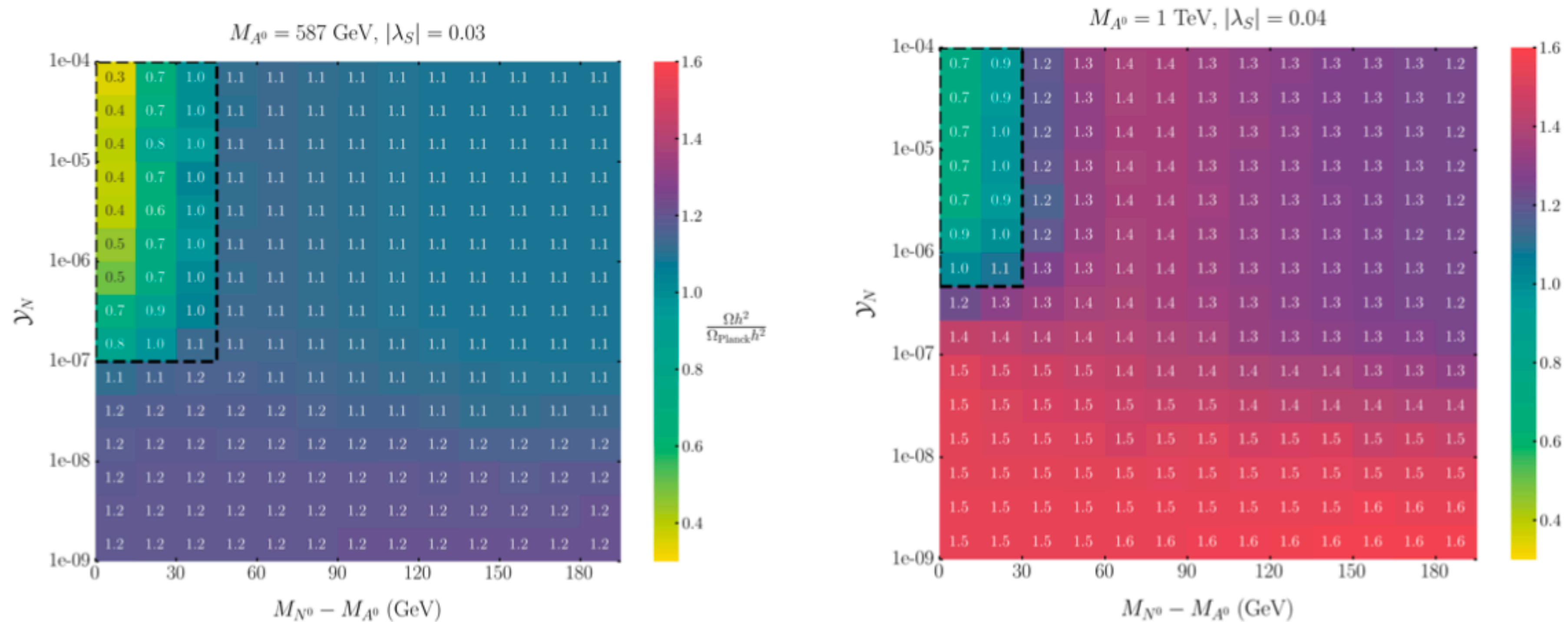
$$\frac{dY_N}{dx} = -\frac{1}{x^2} \frac{s(M_{A^0})}{H(M_{A^0})} \left[\langle \sigma v \rangle_{2200} \left(Y_N^2 - (Y_N^{eq})^2 \right) + \langle \sigma v \rangle_{2211} \left(Y_{\Phi_2}^2 - Y_N^2 \frac{(Y_{\Phi_2}^{eq})^2}{(Y_N^{eq})^2} \right) \right. \\ \left. + \langle \sigma v \rangle_{1200} \left(Y_{\Phi_2} Y_N - Y_{\Phi_2}^{eq} Y_N^{eq} \right) \right] - \frac{x \Gamma_{N \rightarrow \Phi_2 X}}{H(M_{A^0})} \left(Y_N - Y_{\Phi_2} \frac{Y_N^{eq}}{Y_{\Phi_2}^{eq}} \right).$$

Interplay between IDM and VLL



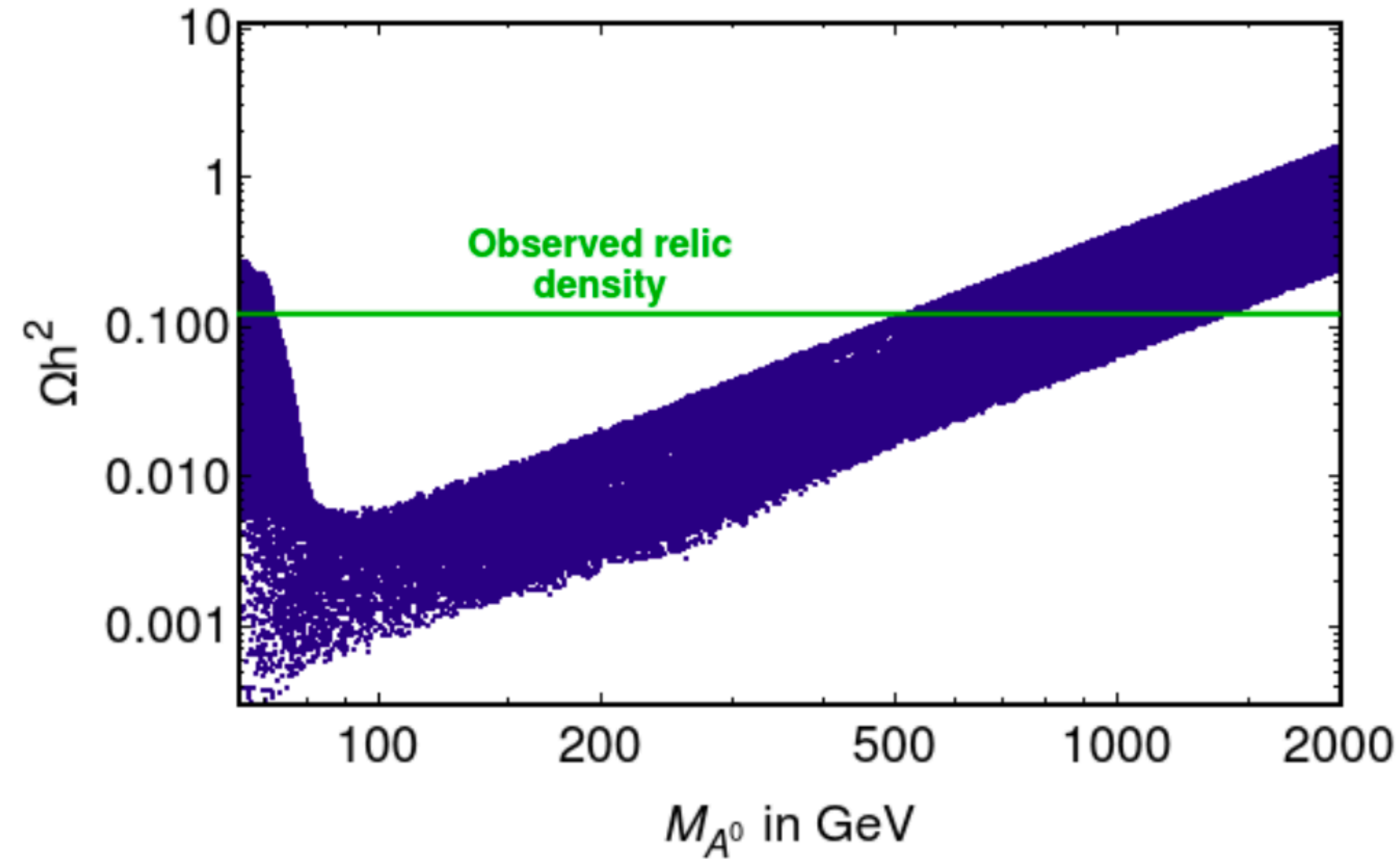
- The A^0 suffers a dip due to co-annihilation due to compressed spectra
- The number density of A^0 increases as N^0 decays completely
- Pure IDM shows over abundant, but due to the interplay with the VLL, the DM relic is back to the allowed
- The situation corresponds to relatively late decay of N^0 , $\Gamma_{N^0} \sim 10^{-16} - 10^{-17} \text{ GeV}$, giving rise to displaced decays N

Correlation of mass gap and Yukawa coupling



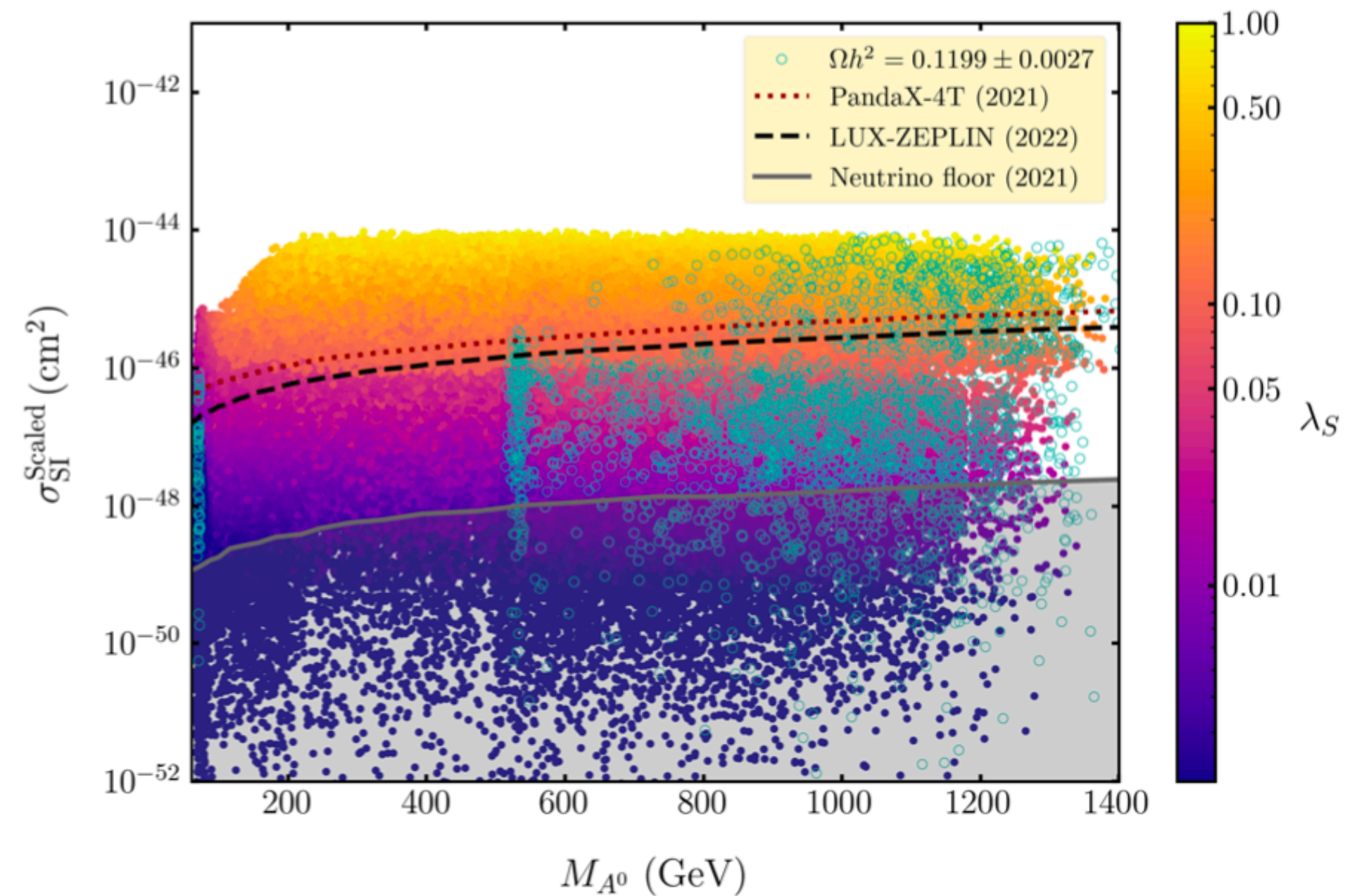
- **Lower Yukawa couplings:** less co-annihilation and corresponds to very late decay of N
- **Higher mass splitting:** Less phase space for co-annihilation, and leads to overabundance of dark matter number density
- **Higher dark matter mass:** Lesser annihilation of DM, needs more compressed spectrum and larger Yukawa for the correct DM relic

Dark matter relic and the interplay



- The observed relic is $\Omega h^2 = 0.1199 \pm 0.0027$
- The lower region around $M_{A^0} \sim 70$ GeV due to the s-channel annihilation via Higgs boson
- $M_{A^0} > 1.4$ TeV is ruled out due to over abundance

Direct Dark matter Detection

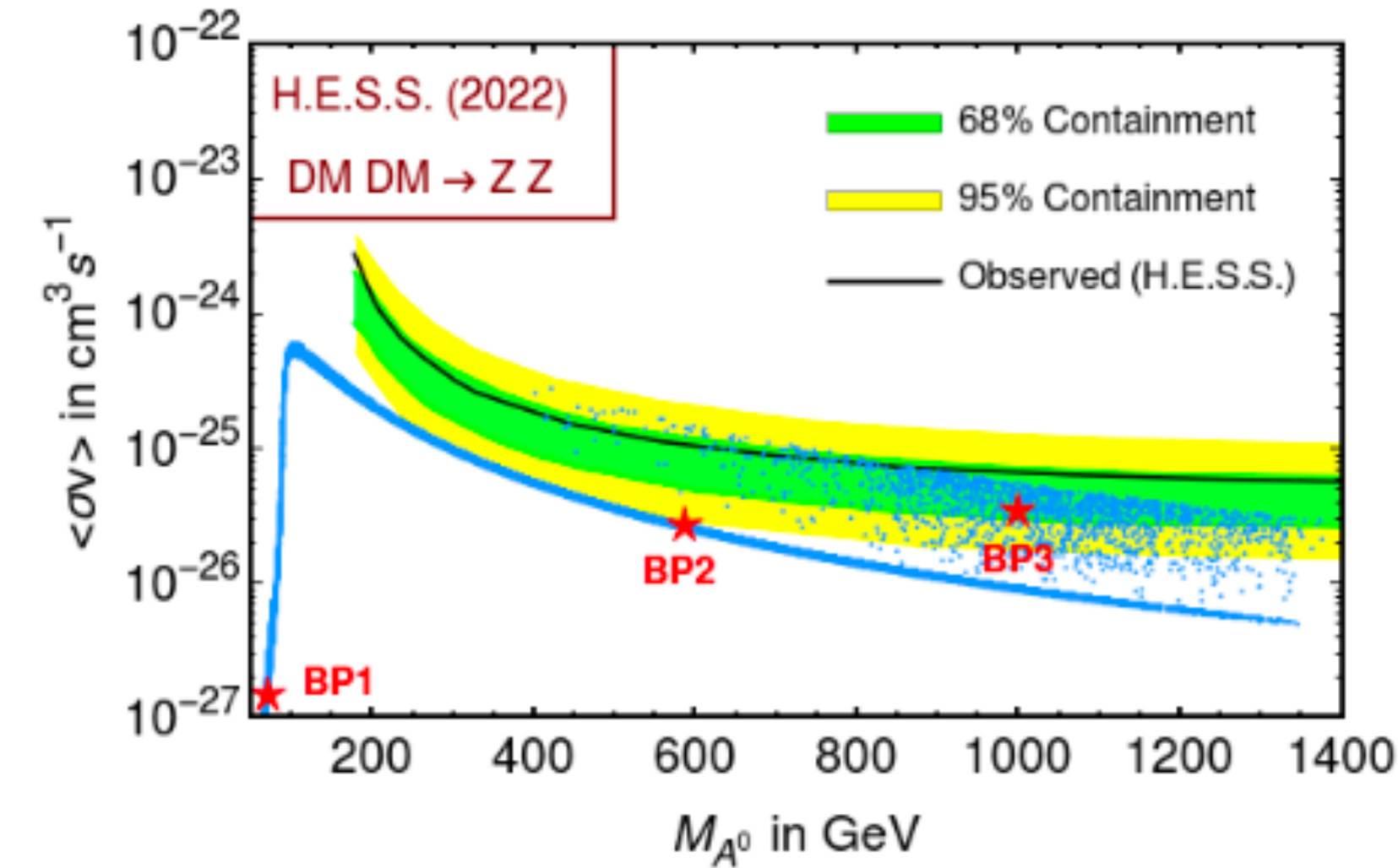
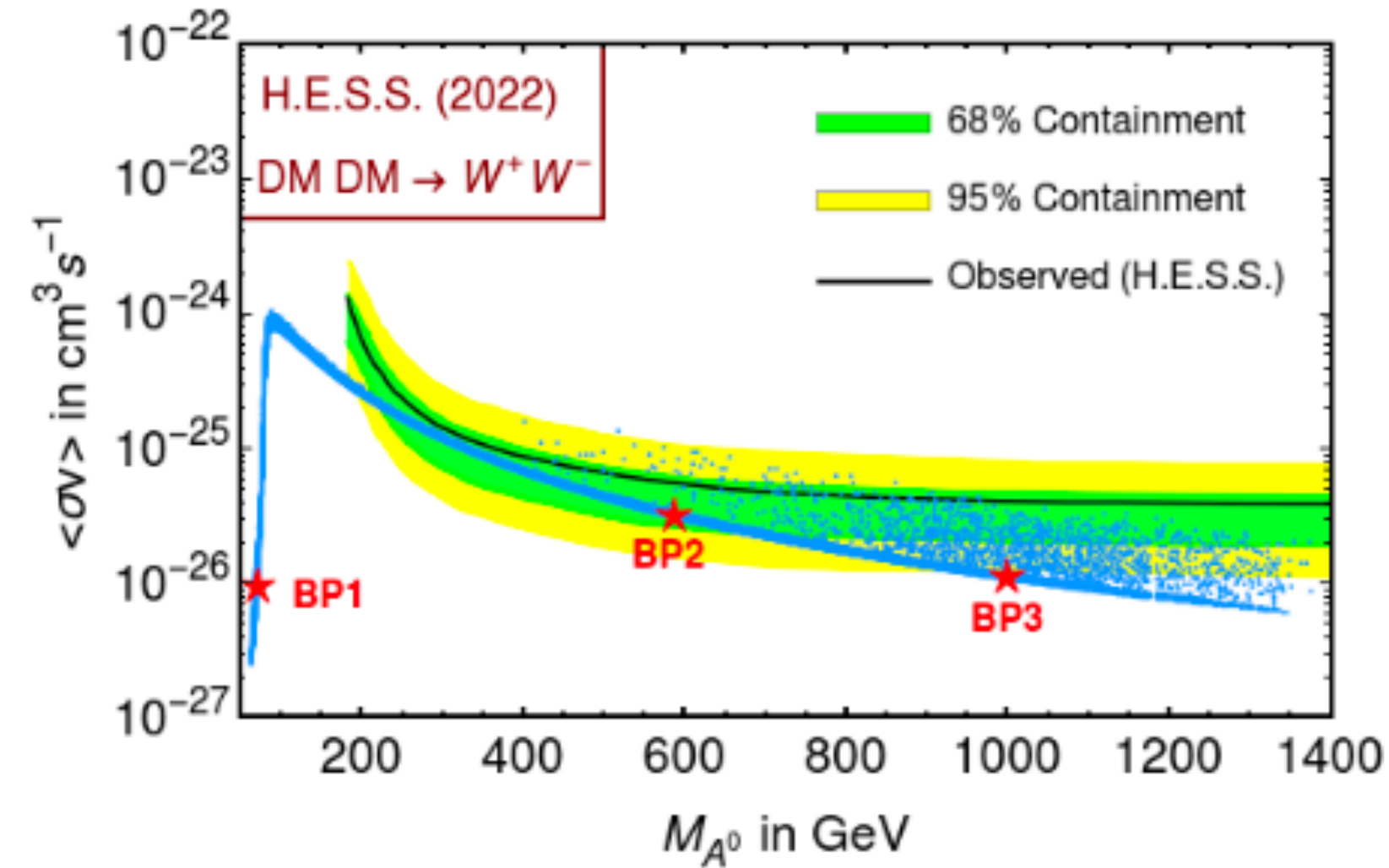
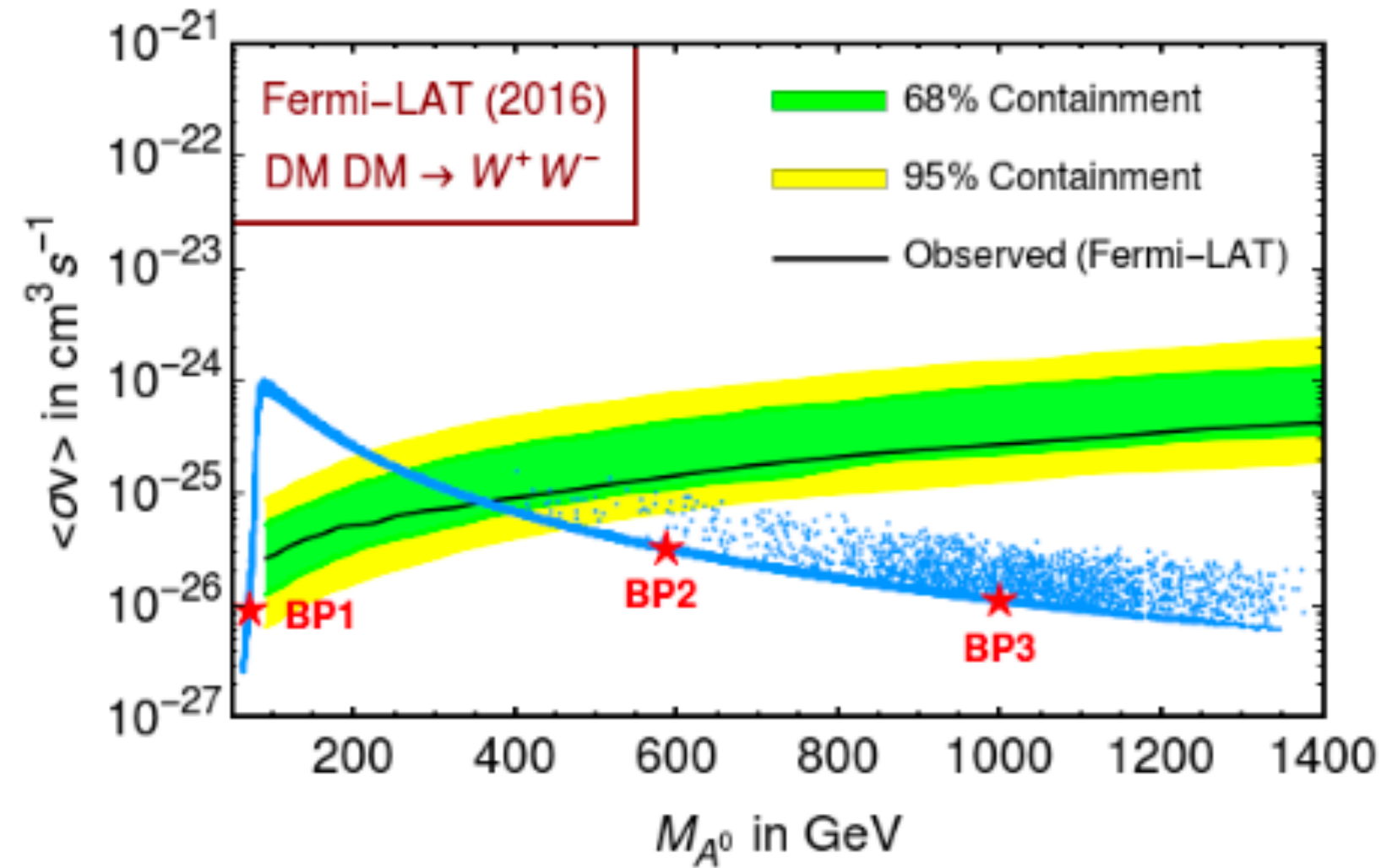


- The direct detection can only happen via the Higgs portal coupling $\lambda_S = \lambda_3 + \lambda_4 - 2\lambda_5$

$$\sigma_{\text{SI}} \simeq \frac{\lambda_S^2 f_n^2}{4\pi M_h^4} \frac{M_n^4}{(M_n + M_{A^0})^2}$$

- The most stringent upper bound comes from the LUX-ZEPLIN experiment, which excludes $|\lambda_S| \geq 0.5$ for $M_{A^0} > 500$ GeV
- $|\lambda_S| \leq 0.01$ is excluded from the neutrino floor bound

Dark matter Indirect Detection



- The dominant annihilation mode of A^0 is $W^\pm W^\mp, ZZ$
- The Fermi-Lat and HESS detects energetic photons which put bounds on the annihilation cross-section

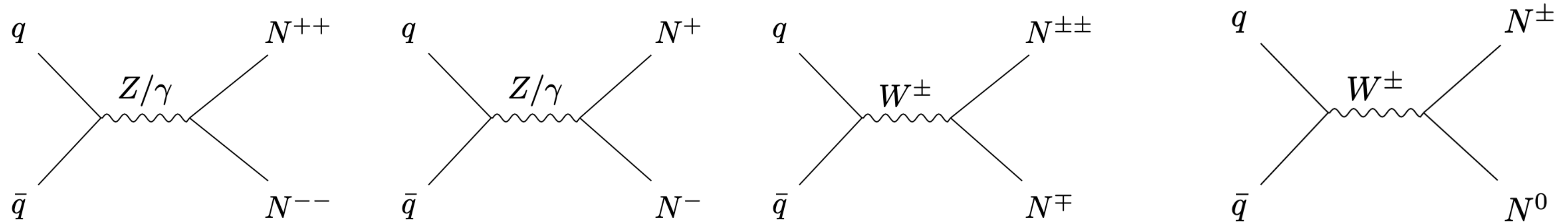
Benchmark Points

BP	M_{A^0} (GeV)	M_{H^0} (GeV)	M_{H^\pm} (GeV)	M_{N^0} (GeV)	M_{N^-} (GeV)	$M_{N^{--}}$ (GeV)	y_N
BP1	71.57	117.16	84.76	98.25	98.61	99.28	4.2×10^{-9}
BP2	587.6	589.4	588.2	595.5	595.9	596.8	1.1×10^{-7}
BP3	1000.0	1010.5	1001.0	1010.5	1011.0	1011.9	5.4×10^{-7}

- All the benchmark points satisfy the dark matter relic, direct and indirect bounds
- All the benchmark points have displaced vertex signature.

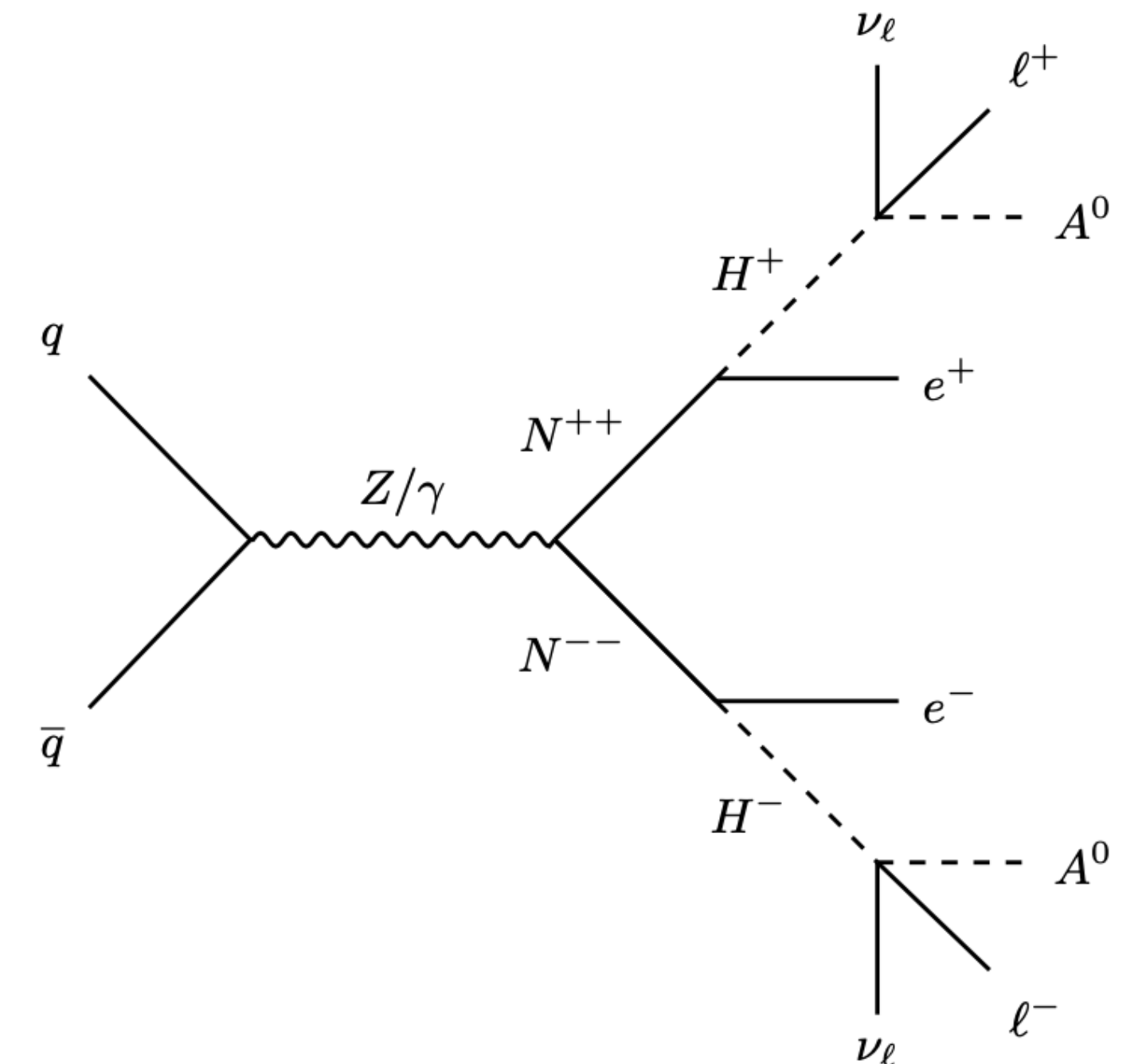
Collider signature

- $pp \rightarrow N^{\pm\pm}N^{\mp\mp}, N^{\pm}N^{\mp}, N^{\pm\pm}N^{\mp}, N^{\pm}N^0$ are looked into at the LHC and FCC-hh

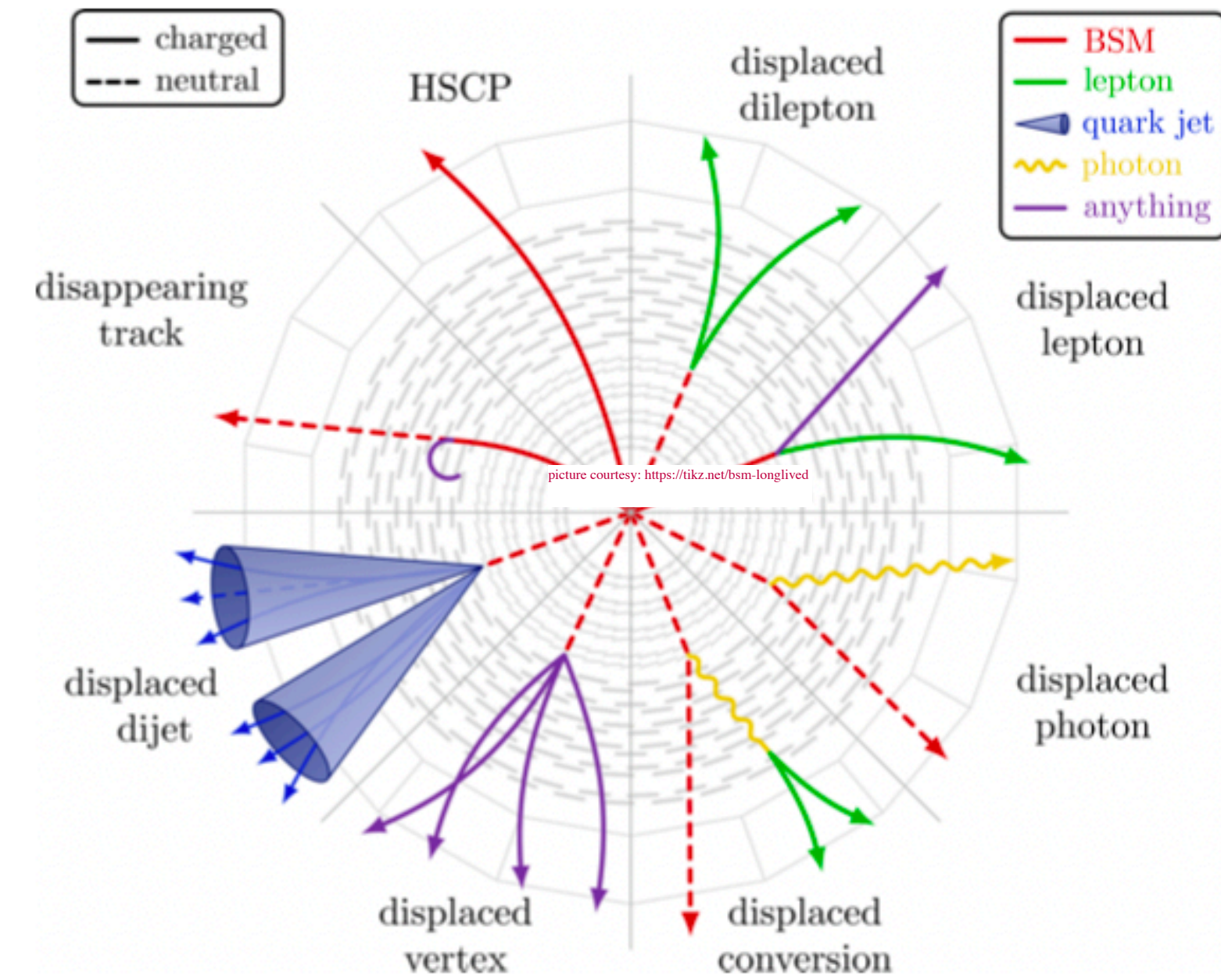
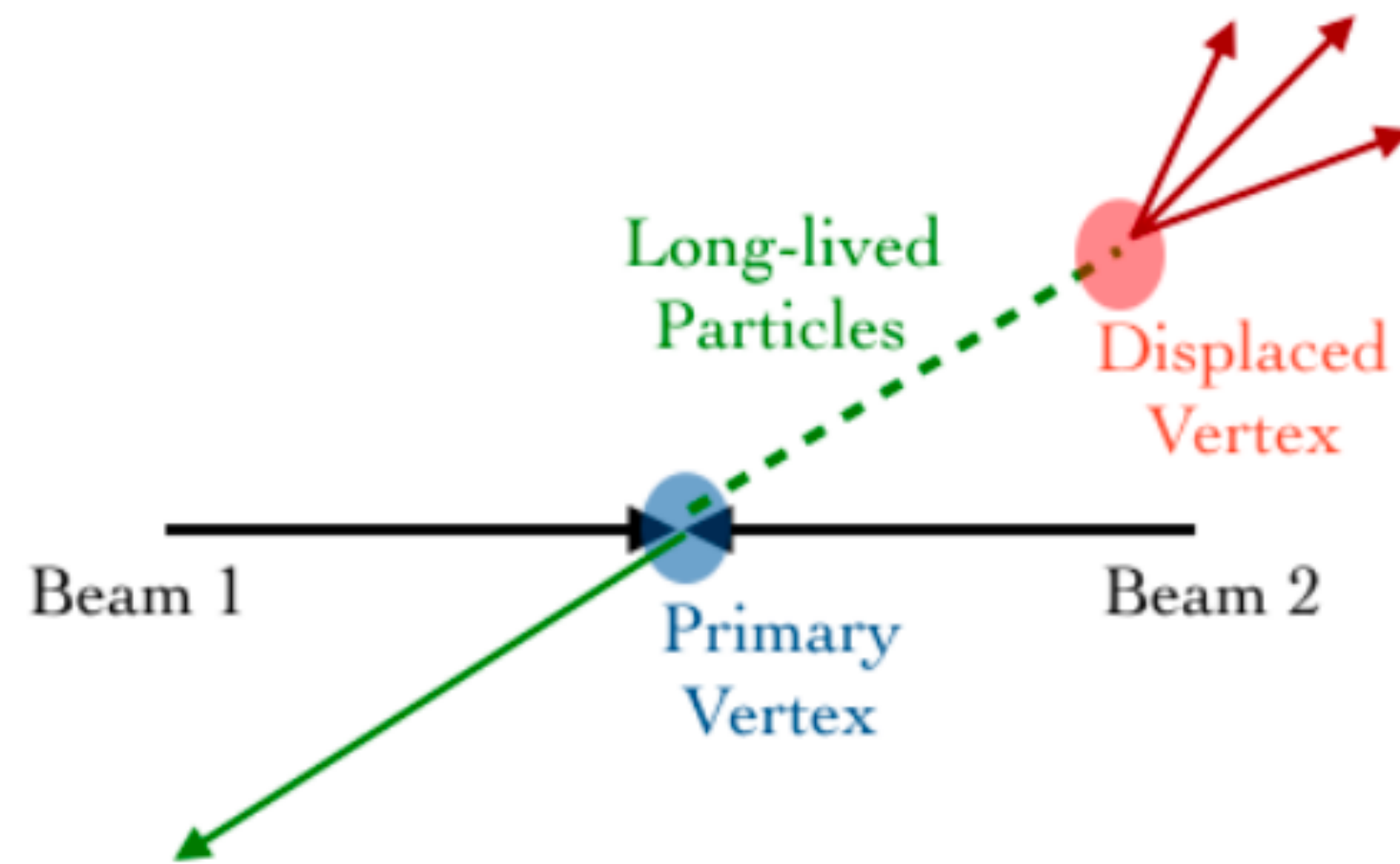


- Typically hadronically quiet displaced multi-leptonic ($1\ell - 6\ell$) signature can be achieved. However, we present the numbers only till 4ℓ

- Pair production of doubly charged VLL can lead to 4 displaced leptons for small Yukawa couplings and compressed mass spectrum



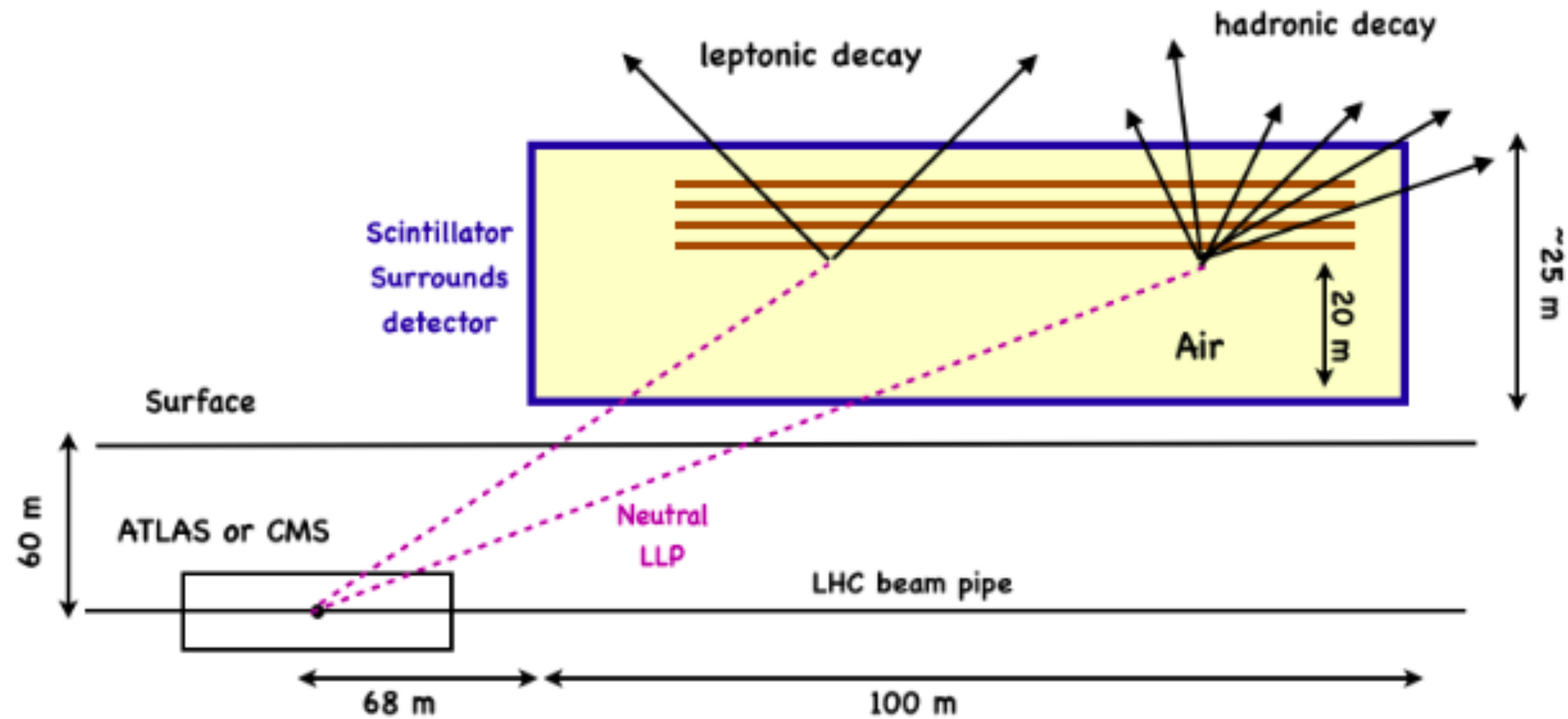
Dispalced decays



picture courtesy: <https://tikz.net/bsm-longlived>

- The decay width $\Gamma = \frac{1}{\tau} \sim g^2 |\mathcal{M}|^2 \Phi$, which can be small due to
 1. Small coupling g
 2. Heavy intermediary particle or cancellation in matrix element
 3. Low phase space due to compressed spectrum

MATHUSLA Detecor



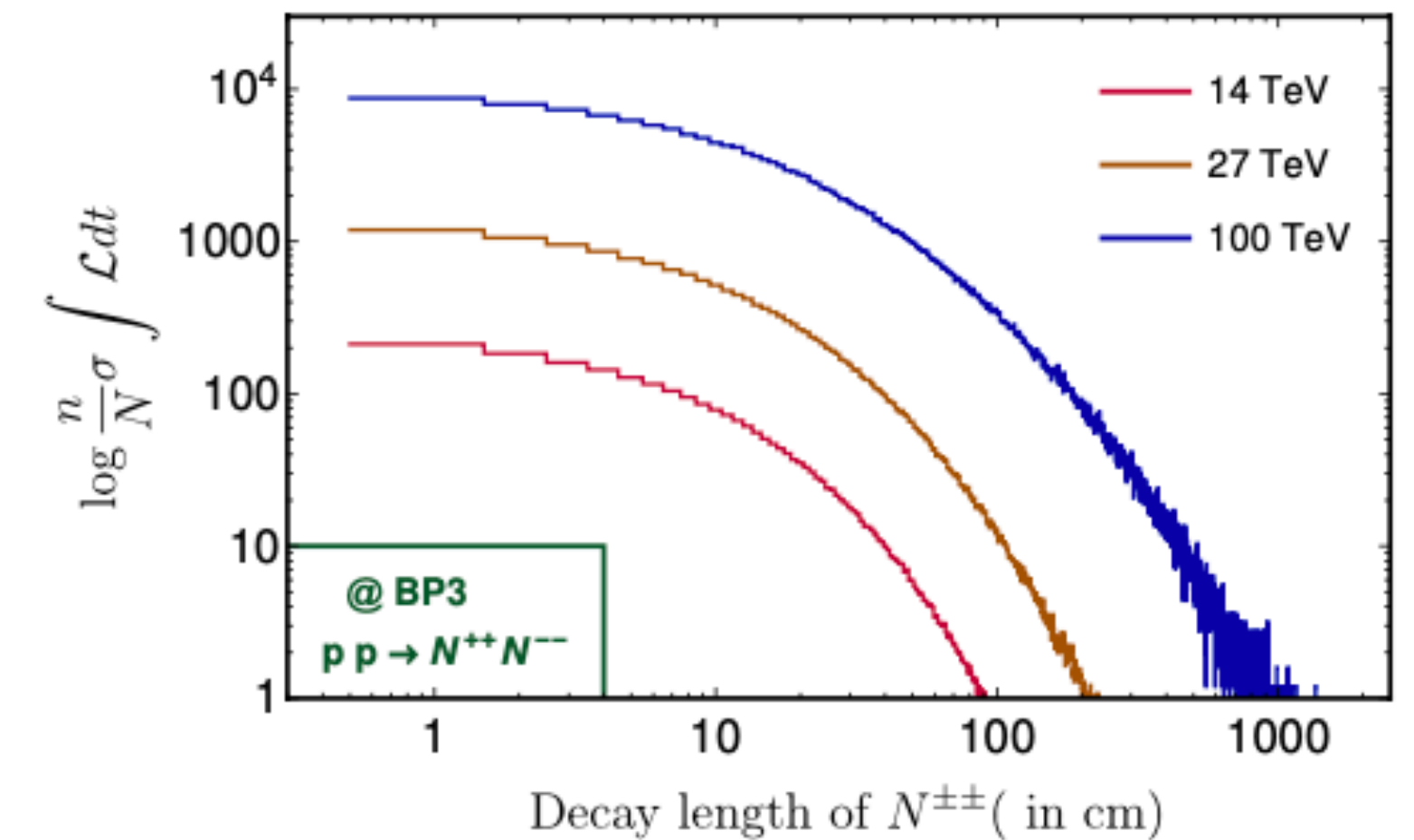
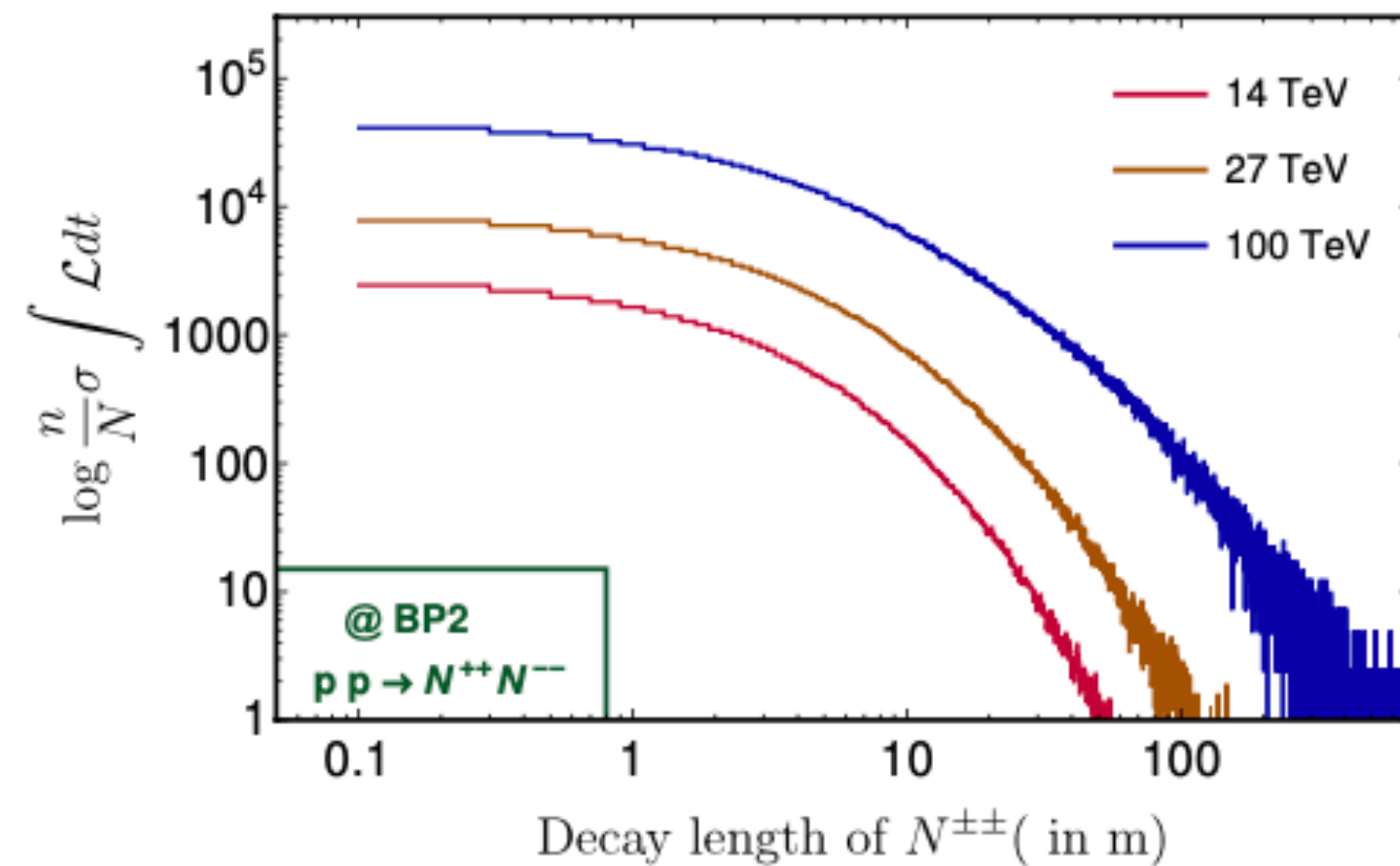
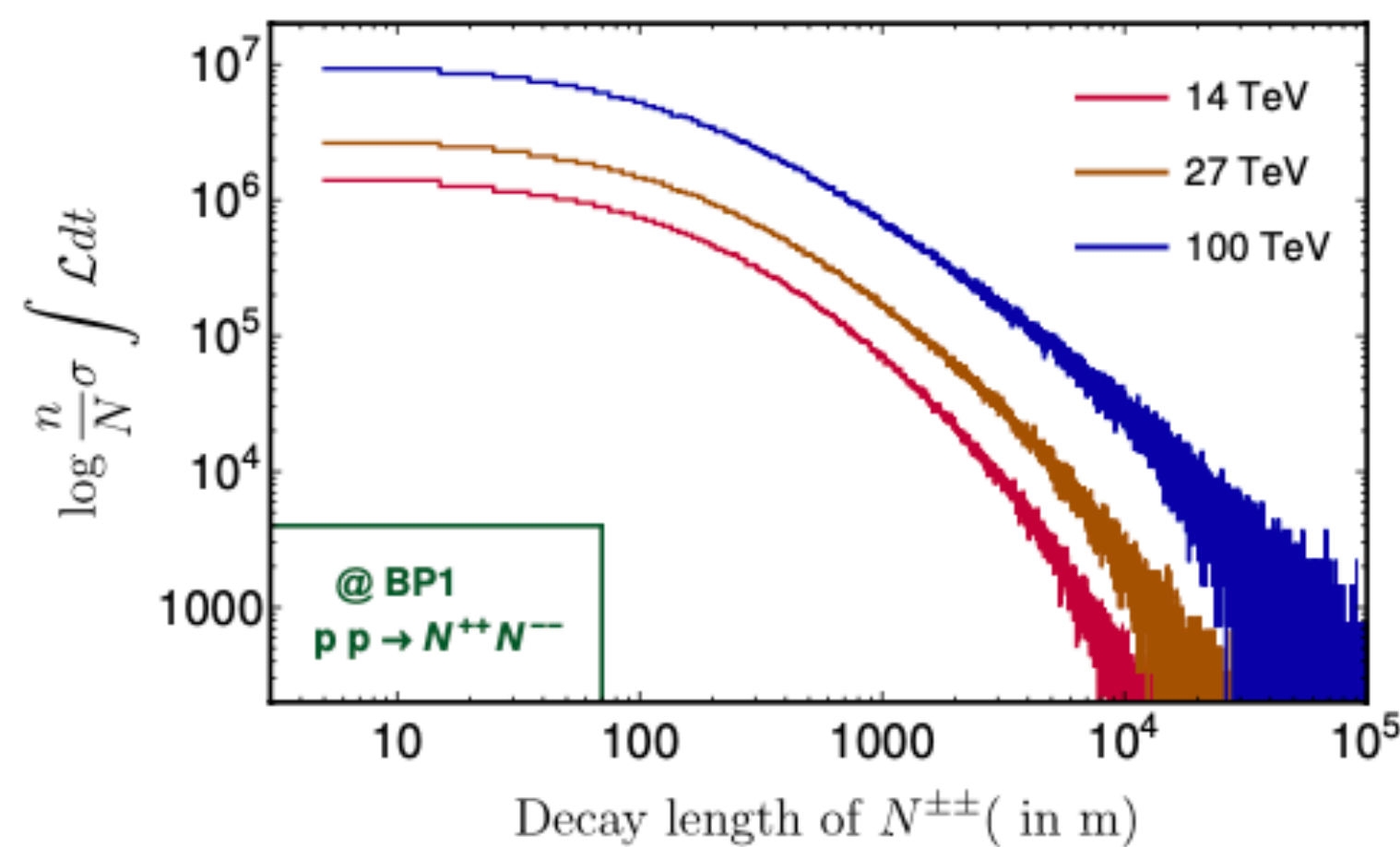
- MATHUSLA is proposed to be built 68 meters from the CMS interaction point in the longitudinal direction, 60 meters in the transverse direction.
- MATHUSLA should be able to search for the long-lived particles with near-zero backgrounds
- MATHUSLA geometry: $25 \times 100 \times 100 \text{ m}^3$

Displaced decay of $N^{\pm\pm}$

- The decay widths and the rest mass decay lengths of $N^{\pm\pm}$

$N^{\pm\pm}$	BP1		BP2		BP3	
	$\mathcal{Y}_N = 4.2 \times 10^{-9}$		$\mathcal{Y}_N = 1.1 \times 10^{-7}$		$\mathcal{Y}_N = 5.4 \times 10^{-7}$	
	Γ_{tot} (GeV)	$c\tau_0$ (m)	Γ_{tot} (GeV)	$c\tau_0$ (m)	Γ_{tot} (GeV)	$c\tau_0$ (m)
	1.27×10^{-18}	155.42	5.92×10^{-17}	3.33	1.34×10^{-15}	0.15

- The decay length distributions including the boost effect



Numbers for 2 displaced leptons

- A typical comparison can be made for the two displaced leptons inside CMS, ATLAS for different centre of mass energies at 300 fb^{-1} luminosity

Final States: 2 displaced leptons + 0 jet with	BPs	Centre of mass energies at				
		14 TeV		27 TeV		100 TeV
		CMS	ATLAS	CMS	ATLAS	FCC-hh detector
$p_{T_{\ell_1}} \geq 20 \text{ GeV}$ & $p_{T_{\ell_2}} \geq 10 \text{ GeV}$	BP1	1108.2	2112.5	2107.0	3831.5	40162.2
$p_{T_{\ell_{1,2}}} \geq 10 \text{ GeV}$	BP2	647.3	1068.5	2300.2	3927.4	28388.6
	BP3	329.7	329.7	2113.5	2113.5	11401.5

- The events inside MATHUSLA is possible only for BP1 and at 300 fb^{-1} luminosity the numbers are

Final States for BP1 at MATHUSLA	production modes	centre-of-mass energies		
		14 TeV	27 TeV	100 TeV
FS6: 2 displaced leptons with $p_{T_{\ell_{1,2}}} \geq 10 \text{ GeV}$	$N^{++}N^{--}$	249.1	507.7	1690.6
	$N^{\pm\pm}N^{\mp}$	259.8	623.8	2114.7
	N^+N^-	1.7	1.8	10.3
	$N^\pm N^0$	0.0	0.0	0.0
	Total	510.6	1133.3	3815.6

Conclusions

- Displaced decays can be used as non-standard searches to probe New physics scenarios
- Displaced decay of neutral fermions and charged fermions in the case of Type-III or VLL can probe the models for the unexplored territory
- Lower Yukawa coupling and compressed spectrum can alter the DM phenomenology
- $\mathcal{O}(100)$ m decay lengths can be probed inside MATHUSLA detector
- Double recoil: $N^\pm \rightarrow N^0 \pi (\sim \text{cm}) \rightarrow h\nu\pi (\sim 1-100 \text{ m})$ can be probed along with a displaced Higgs production
- A doubly charged displaced lepton charged track and its decay can be searched for

THANK

You!