SUSY

Theory meets Experiment

Madrid, 10 - 14 June 2024 Pre-SUSY school: 3 – 7 June 2024 https://indico.cern.ch/e/susy2024

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Recent Dark Matter searches at BaBar and implications on some SUSY models



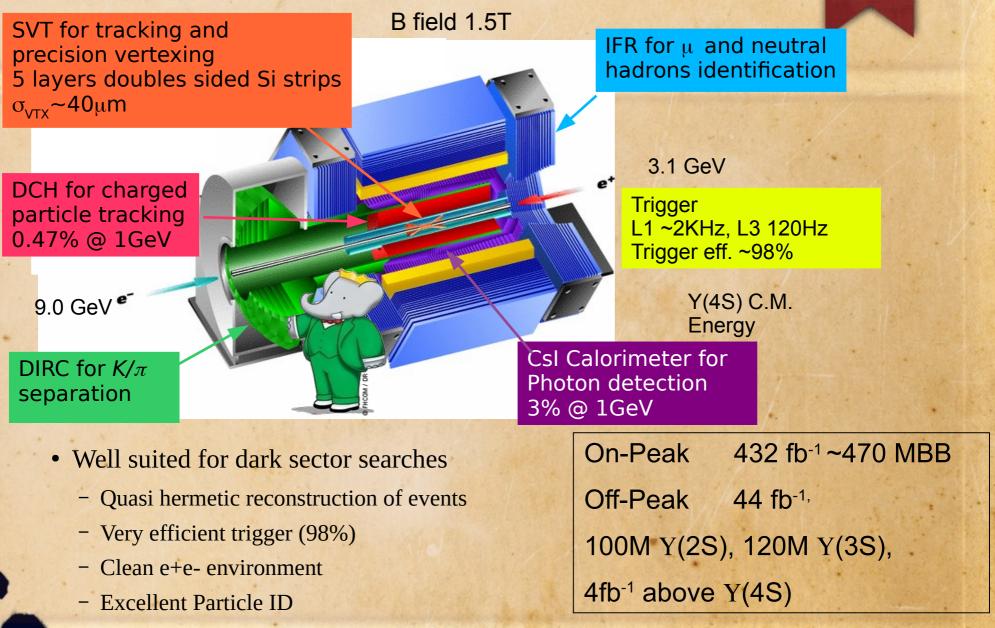
Gabriele Simi University of Padova and INFN On behalf of the BaBar collaboration



Overview

- BaBar experiment
- Introduction
- Search for Heavy Neutral Leptons in tau decays
 - Phys.Rev.D 107 (2023) 5, 052009 (BABAR)
- Search for Baryogenesis and Dark Matter
 - Phys.Rev.Lett. 131 (2023) 20, 201801 (BABAR)
 - Phys.Rev.D 107 (2023) 9, 092001 (BABAR)
- Implications for some SUSY models
 - JHEP 2023, 224 (2023)
 - Phys. Rev.Lett. 131 (2023) 20, 201801 (BABAR)

BABAR Experiment at PEPII



- Accurate Missing Energy reconstruction

Nucl. Instrum. Meth. A 729, 615 (2013)

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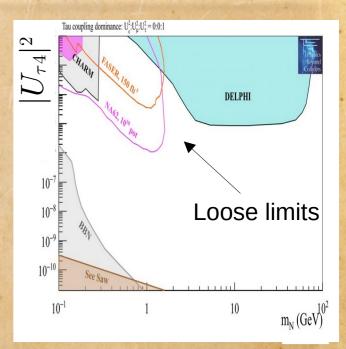
Dark Matter (DM) Portals

- To understand the nature of dark matter an effective theory approach provides several portals
- At B-Factories DM can be produced in e⁺e⁻ interactions or in B meson decays
- Searches can access different operators, constrain masses and couplings



Search For HNL: Motivation

- Several Beyond Standard Model theories introduce Heavy Neutral Leptons (HNL)
- HNL are massive additional neutrino states which have no other charges, and can explain several phenomena
 - Neutrino small mass (seesaw mechanism) and oscillations
 - Baryon asymmetry in the universe
 - The origin of dark matter
- The ν -MSM has three HNL in the keV-GeV range
- HNL can be produced in lepton decays and the τ sector is the less constrained
- BaBar produces copious amounts of τ in a clean environment
 - _ $\mathcal{L}{=}424\,fb^{-1}$, $N_{\tau\tau}{=}400M$
 - $\sigma (e^+e^-) \rightarrow \tau^+\tau^- = 0.92 \, nb, \ \sigma_{tot} \simeq 5.4 \, nb + bhabha$
- BaBar can search for a HNL that is capable of mixing with *τ* with strength |U_{τ4}|² in the mass range 100 MeV/c² < m₄ < 1300 MeV/c²



$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \\ \nu_L \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ U_{L1} & U_{L2} & U_{L3} & U_{L4} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ \nu_4 \end{pmatrix}$$

Serach for HNL: strategy

Selection

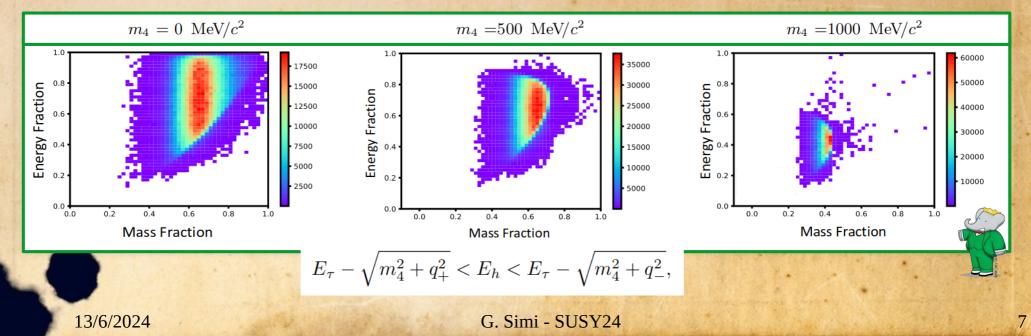
- Signal side: three prong pionic decay ($\tau \rightarrow \pi \pi \pi \nu_{\tau}$)
- Tag side one prong τ decay to reduce bkg
- Signal and tag identified in different hemispheres (3+1 topology), thrust>0.85
- P_T , $P^{miss}_{CM} > 0.9\% \sqrt{S}$ to suppress non tau bkg
- DIRC acceptance and PID requirements for e and μ
- Neutral and conversion veto for tag side
- Signal extraction: same strategy used in previous analysis in BaBar using only variables sensitive to the kinematics
 - if the decay products of the τ have recoiled against a heavy neutrino, the phase space and the kinematics of the visible particles would be modified with respect to SM τ decay with a massless neutrino: no model assumptions

HNL Signal

- Consider the process as a two body decay $au
 ightarrow h^-(E_h, m_h)
 u_{ au}$
- The signal is searched in the 2D plane of the hadron mass and energy (m_h, E_h)
- Tau decay gets separate contributions from SM and BSM decay

$$\frac{\Gamma(\tau^- \to \nu h^-)}{\mathrm{d}m_h dE_h} \bigg|_{\mathrm{Total}} = \left| U_{\tau 4} \right|^2 \frac{\mathrm{d}\Gamma(\tau^- \to \nu h^-)}{\mathrm{d}m_h \mathrm{d}E_h} \bigg|_{\mathrm{HNL}} + (1 - |U_{\tau 4}|^2) \frac{\mathrm{d}\Gamma(\tau^- \to \nu h^-)}{\mathrm{d}m_h \mathrm{d}E_h} \bigg|_{\mathrm{SM}}.$$

- 2D templates are built for each different m_4 mass. The hadronic system available phase space reduces as m_4 increases
- Signal samples from modified KK2F+TAUOLA+GEANT4 in the range $100 MeV/c^2 < m_4 < 1300 Mev/c^2$



HNL: Signal extraction

- Background from MC simulation
 - Misidentified SM τ decays
 - SM non τ bkg: B decays, light quarks, $e^+e^- \rightarrow \mu^+\mu^-(\gamma)$
- Signal (ν_{HNL}) extracted from binned ML fit assuming each bin has Poisson distribution with mean $\nu_{HNL} + \nu_{\tau-SM} + \nu_{BKG}$

 $u_{HNL} \propto |U_{ au 4}|^2, \
u_{ au - SM} \propto (1 - |U_{ au 4}|^2)$

- Limits include nuisance parameters
- Dominant systematic from shape uncertainty

Mass [MeV/ c^2]	No Sys.	With Sys.
100	1.58×10^{-2}	2.31×10^{-2}
200	1.33×10^{-2}	1.95×10^{-2}
300	6.91×10^{-3}	9.67×10^{-3}
400	1.57×10^{-3}	2.14×10^{-3}
500	4.65×10^{-4}	5.85×10^{-4}
600	5.06×10^{-4}	6.22×10^{-4}
700	3.82×10^{-4}	4.85×10^{-4}
800	3.12×10^{-4}	3.85×10^{-4}
900	4.70×10^{-5}	5.38×10^{-5}
1000	8.34×10^{-5}	9.11×10^{-5}
1100	4.49×10^{-5}	4.78×10^{-5}
1200	4.70×10^{-6}	5.04×10^{-6}
1300	3.85×10^{-5}	4.09×10^{-5}

$$\mathcal{L} = \prod_{ij} f(n_{ij}; n_{\text{obs}}, \vec{\theta}) = \prod_{ij} \frac{(\nu_{\text{HNL}} + \nu_{\tau-\text{SM}} + \nu_{\text{BKG}})_{ij}^{(n_{\text{obs}})_{ij}} e^{-(\nu_{\text{HNL}} + \nu_{\text{BKG}} + \nu_{\tau-\text{SM}})_{ij}}}{(n_{\text{obs}})_{ij}!} \times \prod_{k} f(\theta_k, \tilde{\theta}_k),$$

• For each mass re-weighted 2D templates for signal are used

Nuisance parameters

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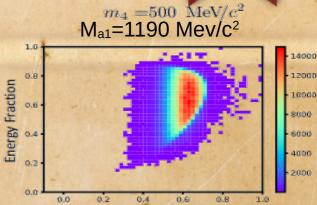
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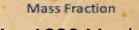
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HNL: systematic uncertainties

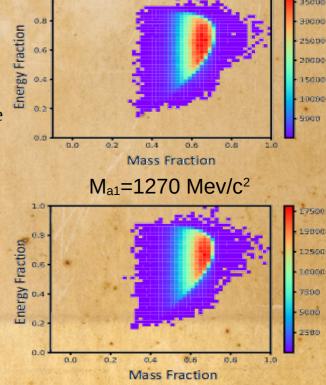
- Normalization uncertainties
 - Affect all bins uniformly
 - Have small effect on the yield (<3%)
 - Biggest contribution from PID efficiency
 - Characterized by Gaussian nuisance parameters
- Shape uncertainties
 - − τ → 3-prong BF through a_1 resonance = 97%
 - Large uncertainties on a_1 parameters mass (+/-3%) and Γ_{a1} (250 MeV/c² 600 MeV/c²)
 - Modeling of signal and bkg shape in TAUOLA affected by these uncertainties
 - To account for this looked at templates with mass and width varied to these extremes (re-weighted MC) and re-calculated the likelihood
 - Γ_{a1} has the largest effect, especially on the RMS of m_h

	E _h	m _h
RMS shift	1%-3%	6%-7%
Mean shift	1%-2%	1%-2%







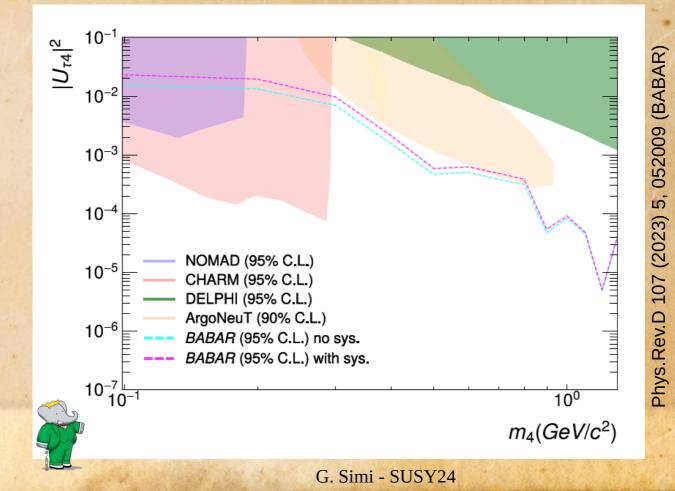


^ohys.Rev.D 107 (2023) 5, 052009 (BABAR)

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HNL: Results

- Result is expressed as an upper limit on $|U_{ au 4}|^2 vs m_4$
- Covers nicely region around 1 GeV/c²
- World leading constraint at time of publication



Baryogenesys and dark matter: Introduction

- A new dark sector anti-baryon has been proposed to simultaneously explain both the dark matter (DM) abundance and the Baryon Asymmetry in the Universe (BAU)
- Baryon asymmetry generated via
 - 1) Production of bb pairs

- 2) CPV in B oscillations generates more B⁰ than anti-B⁰
- 3) B meson decay into a baryon \mathcal{B} (= Λ or

p in this search) and a dark sector antibaryon (Ψ_D) + additional mesons (\mathcal{M})

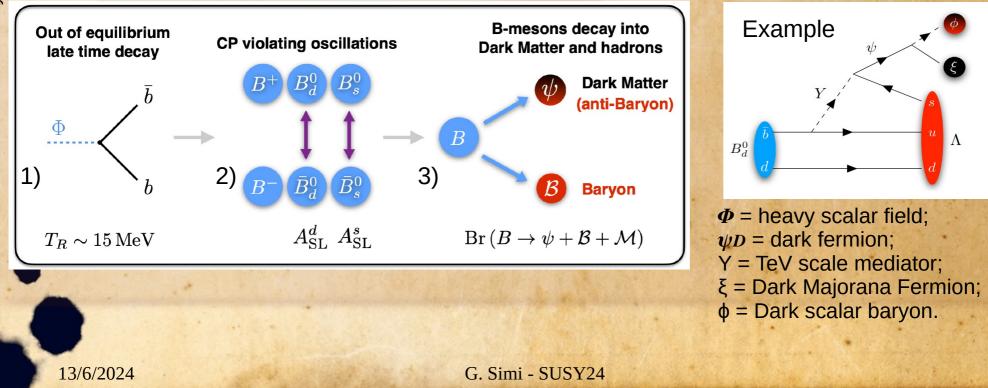
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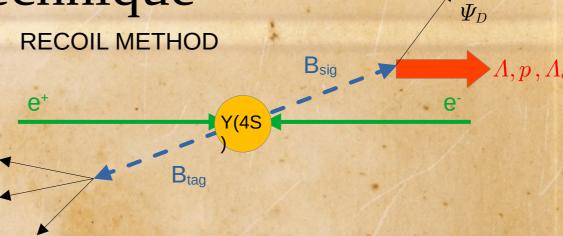
generated visible asymmetry



Operators sensitivity and Reconstruction Technique

- Decay described by effective operators O_{ij} corresponding to transitions $b \rightarrow q_i q_j$ where i=u,cj=d,s
- Different decay channels are sensitive to different effective operators

In this talk		In preparation	
Meson	Decay	Operato r	M _{max} [MeV]
B^0	$\varPsi_{D} + \varLambda$	O_{us}	4163.95
B^+	$\varPsi_{D} + p$	O_{ud}	4341.05
B^+	$\varPsi_D{+}\Lambda_c$	O_{cd}	2992.86



- B_{tag}: fully reconstructed hadronic B-decays (mostly charmed)
- B_{sig}: the rest of the event = a single reconstructed baryon
- $\Psi_{D:}$ Missing momentum since dark-sector particles escape detection
- Kinematic constraints limit the search for $0.94 \, GeV < m_{\Psi_D} < 4.34 \, GeV$

Selection

• B_{Tag}

- Beam Energy substituted mass
- $-\Delta E = E_{beam} E_{BTag}$
- Spherical event topology
- Signal Baryon
 - BDT built from bkg rejection variables, baryon (and B_{Tag} and) reconstruction quality variables with small correlation with missing mass

-data

 $\times 10^3$

 $^{27}_{beam}$

#Events / (1.0 MeV/c²)

20

15

10

 $B^0\overline{B}^0$

cc

5.25 5.26 5.27 5.28 5.29 5.3

m_{ES} [GeV/c²]

- Signal

-B⁺B

-uds

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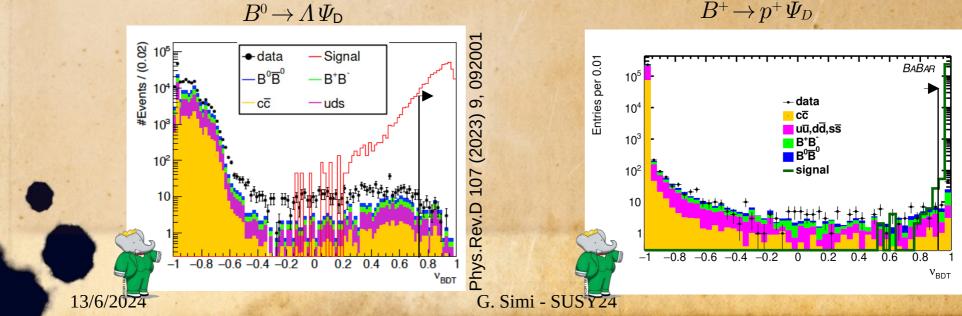
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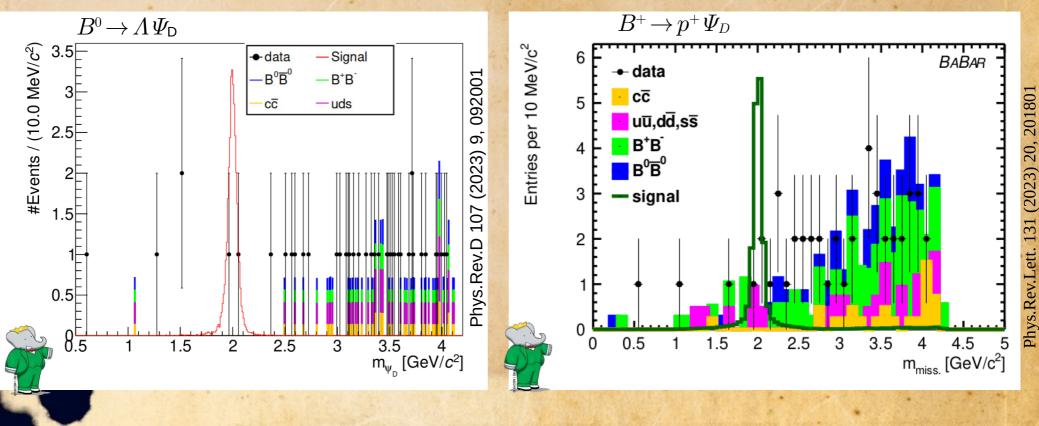
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Results

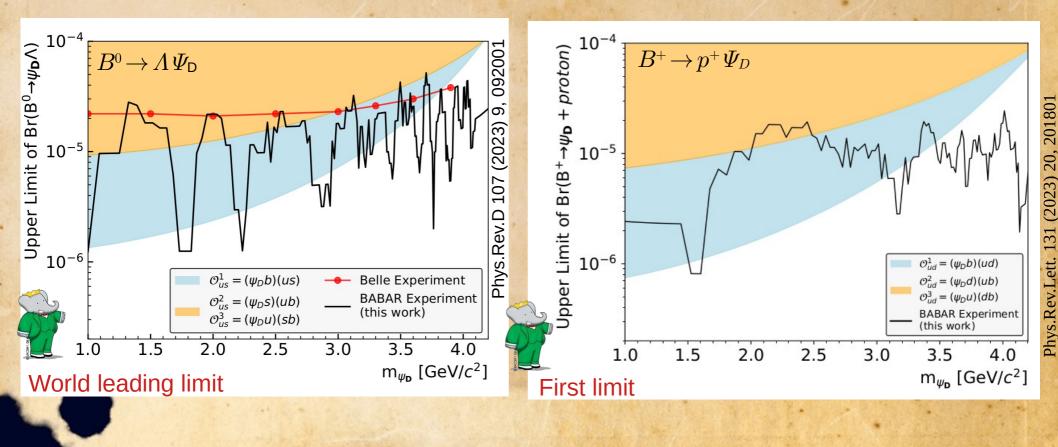
- Total data used 389ifb
- 10% of data used to optimize analysis
- Signal is extracted from the missing mass distribution fitted to a Crystal Ball

- Scan is performed on the missing mass with step size equal to the resolution
- Largest systematic uncertainty from data/MC corrections (efficiency)



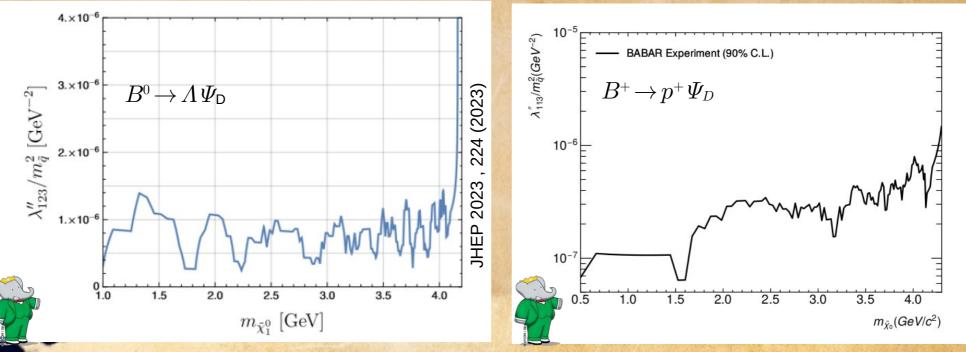
Results

- Upper limits (black line) computed assuming Poisson distribution
- 95% exclusion limits (black line) obtained scanning the missing mass distribution for many values
- Global significance at 1σ level
- Colored bands are the allowed regions consistent with observed baryon asymmetry: large regions excluded



Results: limits on RPV SUSY models

- Since all we are requiring is missing mass these results can be reinterpreted to constrain other models, including the R Parity Violating (RPV) supersymmetry process $B^+ \rightarrow \tilde{\chi}_0 + p$ where $\tilde{\chi}_0$ is the lightest neutralino. It violates R parity since it produces only one neutralino.
 - BF bounds are interpreted as limits on λ''_{113} vs neutralino mass



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Phys.Rev.Lett. 131 (2023) 20, 201801

d/s

Conclusions

- BaBar continues to produce excellent physics results and is well suited to search for dark matter candidates
- We performed a model independent search searched for Heavy Neutral Leptons (mostly sterile)
 - Could explain simultaneously BAU and DM origin
 - No signal observed but set the world leading limits for the mixing strength with the tau in the region around 1 GeV/c²
 - Largest systematic coming from hadronic tau decays
- We performed a search for dark matter candidates in exotic B decays to a baryon a dark sector anti-baryon
 - No signal is observed and we put limits on the BF
 - In the context of DM and BAU B-mesogenesys models we have put stringent limits on some operators
 - In the context of SUSY these limits can be recast as limits on RPV SUSY models





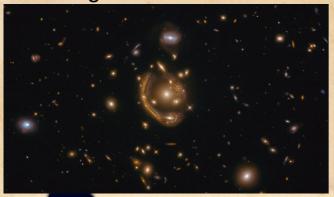
BACKUP

Dark Matter in the Universe

Cosmic microwave bkg

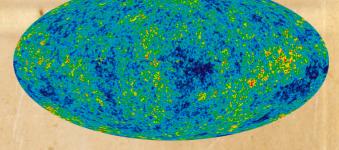
- Many experimental evidences of Dark matter
 - Lambda CDM: 5% Baryonic Matter, 27%
 Dark Matter, 68% Dark Energy
 - Rotational curve of galaxies
 - Large scale structure formation
 - Bullet galaxy cluster
 - Gravitational lensing

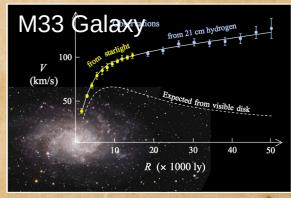
lensing



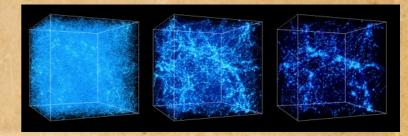
Dark matter Bullet cluster







[simulated] structure formation



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