

# The rise and fall of light stops in the LHC top quark sample

---

Emanuele A. Bagnaschi (INFN LNF)



Istituto Nazionale di Fisica Nucleare  
Laboratori Nazionali di Frascati

10 June 2024

*SUSY2024*

**IFT Madrid**

# Introduction

---

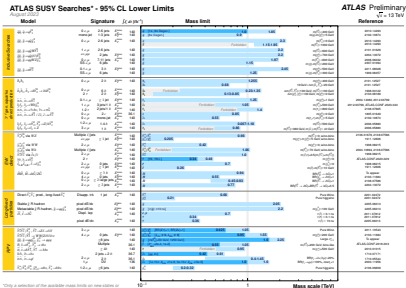
# BSM physics at the LHC: status and prospects

## Status

- No clear signs of new physics from searches, although some anomalies are present here and there in the data

## Prospects

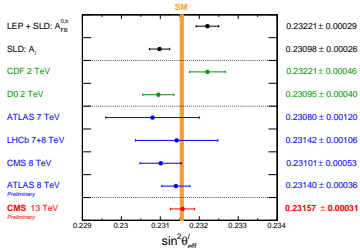
- Continue with direct searches, and move towards more complex final states target/more exclusive phase space (e.g. from long decay chains)



[ATL-PHYS-PUB-2023-025]

# BSM physics at the LHC: status and prospects

[CMS-PAS-SMP-22-010]



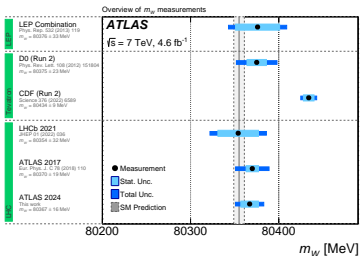
## Status

- Precision tests of observables with SM final states are as important as direct searches, however the general picture here is still of consistency with the SM as well

## Prospects

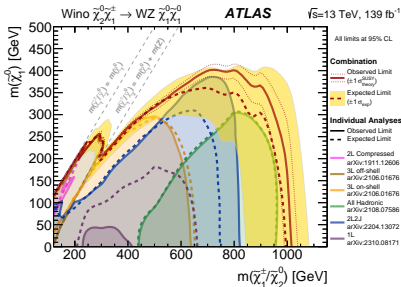
- Continue with the SM physics program at the LHC (e.g.  $M_W$ ,  $\sin^2 \theta_{\text{eff}}^{\text{lep}}$ ), identify new “precisely measured” observables as a target for BSM physics

[ATLAS, 2403.15085]

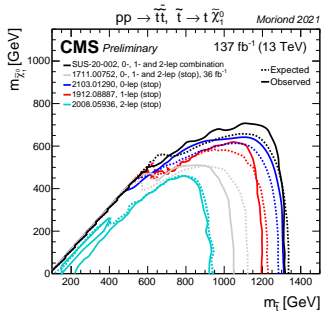


[See the talk of J. Ferrando this morning]

# Exp. searches for SUSY states



[ATLAS, 2403.02455]



[CMS,

summary plots]

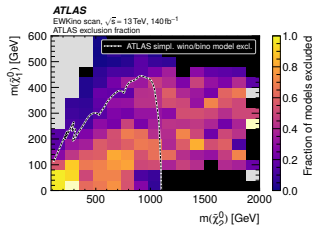
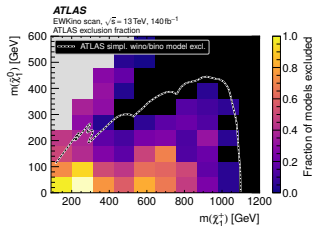
## Experimental persp.

- Extensive research program, focusing (mostly) on SUSY simplified topologies
- Move towards non-minimal simplified models, new ideas etc.
- [See J. Hoya and M. Kazana talks this morning]

# Simplified models vs pMSSM

## Experimental persp.

- Well known in the theory community of the important differences between SSM results and realistic pMSSM scenarios
- For instance see the work of the global fitting collaborations (MasterCode, Gambit, ...) but also pheno papers
- Recent work from ATLAS in that direction, [2402.01392]
- Two similar studies from ATLAS and CMS for Run 1
- *Important effort to share exp. data in such a way that can be easily recasted*



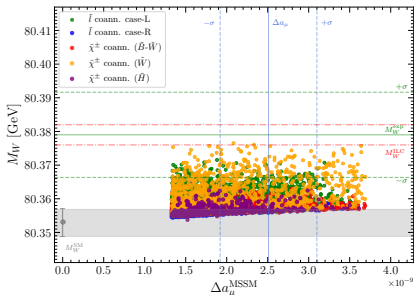
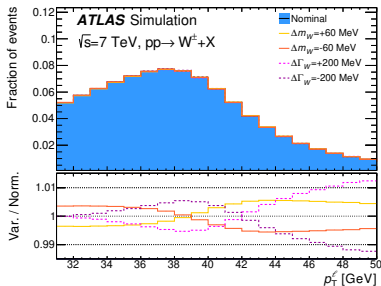
[ATLAS, 2402.01392]

# Indirect probe via precision measurements

## Indirect probes

For example

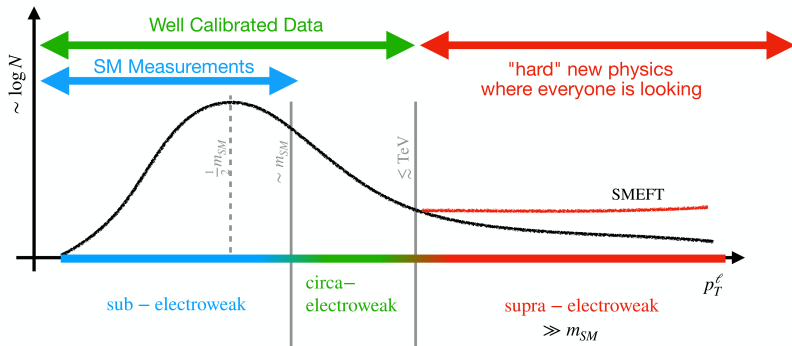
- Higgs mass
- anomalous magnetic moment of the muon
- EWPOs (e.g.  $M_W$ )
- Flavor observables



[EB et al., EPJC82 (2022)5,474, 2203.15710]

# Are the two approaches so well disconnected?

- “SM” parameter measurements are characterized by an extremely good understanding of the experimental phase space



[R. Franceschini's talk at LHCP2024]



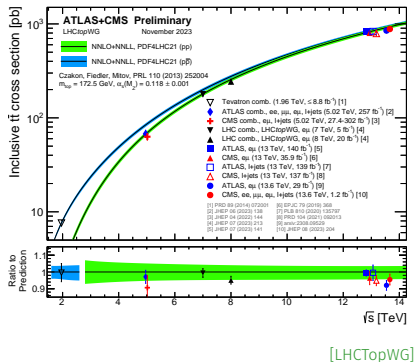
## Our study

---

# Why the top?

## Top@LHC

- Large production cross section at the LHC  $\rightarrow$  LHC is a top factory
- Allow for a precise characterization of its properties (mass, BRs, etc.)
- Precise measurement of differential observables (e.g. spin correlations)
- Top is the heaviest quark  $\rightarrow$  BSM physics portal



# Our idea

## The focus

- Try to understand if top quark observables currently measured with good precision (i.e. understanding of systematics, good statistics) and up to now only used for SM measurements can be used to probe new physics
- We focused ourselves on the invariant mass of the b-jet and the lepton,  $m_{bl}$  that has been used to extract the top quark mass

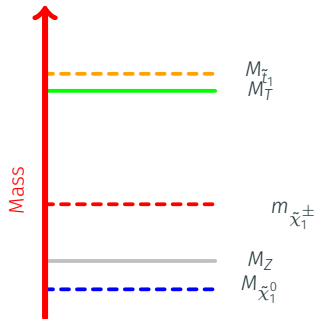
## BSM in $m_{bl}$ ?

- Can we use this measurement to probe BSM physics?
- Note: here by BSM physics we mean the production and decay of BSM states
- Modification of the  $t\bar{t}$  production process (i.e. loop corrections) is also a BSM effect that we have not considered

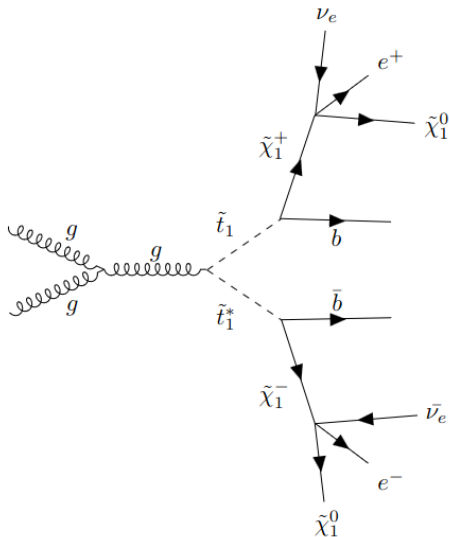
# The scenario

## The spectrum

- We consider a scenario with a light  $\tilde{t}_1 \simeq \tilde{t}_R$
- Stop mass close to the top quark mass:  $m_{\tilde{t}_1} \simeq M_T$
- Bino-like/Bino-higgsino  $\tilde{\chi}_1^0$  and higgsino-like  $\tilde{\chi}_1^\pm$ .
- The MSSM parameters are adjusted to keep the stop mass roughly constant as we vary the chargino ( $\mu$  parameter) and the bino masses



# The signal



## Stop-top contamination

- The final state is similar to the top one
- $\tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^0 \rightarrow l + MET$
- W OS or not depending on the  $\tilde{\chi}_1^\pm, \tilde{\chi}_1^0$  mass difference
- A similar signature could be obtained in the decay channel  $\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$ , which however is much strongly constrained

# The scenario

## The Higgs-stop connection

- In the MSSM, the stop sector closely connected to the Higgs sector because of the Higgs mass
- All the benchmark points are at  $\pm 2$  GeV from the measured value of  $M_H$
- However **SARAH-SPHeno** does not include the best calculation for these spectra
- In any case, one can move to NMSSM or adjust a bit more the parameters

## Input parameters

- $m_{\tilde{u}}(1, 1)^2 = m_{\tilde{u}}(2, 2)^2 = m_{\tilde{q}}(i, i)^2 = m_{\tilde{l}}(i, i)^2 = m_{\tilde{e}}(i, i)^2 = m_{\tilde{d}}(i, i)^2 = 1.2 \cdot 10^7 \text{ GeV}^2$  for  $i = 1, 2, 3$
- All the off-diagonal squark and slepton mass terms are set to zero;
- $m_{\tilde{u}}(3, 3)^2 = 1.7 \cdot 10^5 \text{ GeV}^2 \rightarrow \tilde{t}_1$  almost pure SU(2) singlet state;
- $M_1 \in [5, 1000] \text{ GeV}$ ,  $M_2 = 1 \text{ TeV}$ ,  $M_3 = 3.5 \text{ TeV}$ ;
- $M_A^2 = 2 \cdot 10^6 \text{ GeV}^2$ ,  $\tan \beta = 10$ ;
- $\mu \in [100 \text{ GeV}, m_{\tilde{t}_1}]$ ;
- $A_t \in X_t + \mu \cdot \cot \beta + [-100, 100] \text{ GeV}$ ; the exact value is obtained by trial-and-error to get the desired  $m_{\tilde{t}_1}$ .
- All other trilinear couplings are set to zero.

# Methodology

---

# Methodology

## Point generation

- Keeping the stop mass fixed, scan the parameter space
- Spectrum generation performed with **SARAH+SPHeno**
- Check the Higgs mass constraint
- Check the collider constraints with **SModelS** (e.g. from searches for direct production of the electroweakinos)

## Collider simulation

- The parameter space points that pass the selection are simulated using **Pythia8**
- Cuts applied inspired by exp. analyses

$$p_T(l) \geq 25 \text{ GeV}, \quad |\eta(l)| < 2.5$$

$$p_T(j) \geq 25 \text{ GeV}, \quad |\eta(j)| < 2.5$$

$$\text{anti-kT jets}, R = 0.4$$

$$\Delta R(l, j) > 0.2, \quad \Delta R(j, j) > 0.5$$

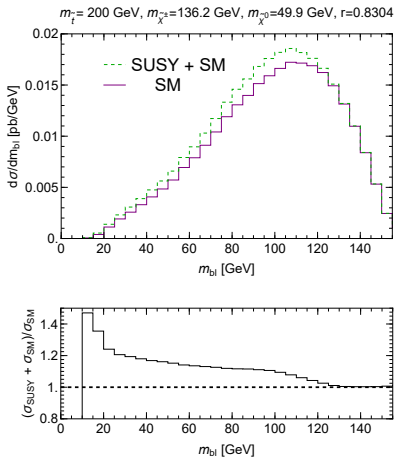
$$\Delta R(l, l) > 0.1$$



# Results

---

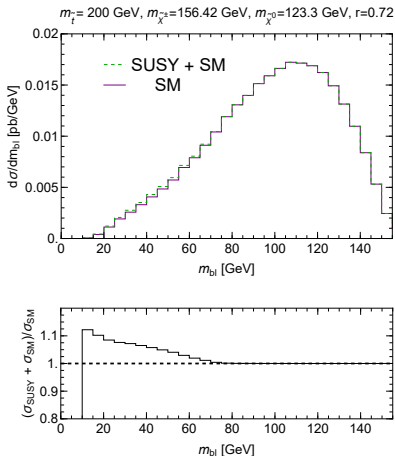
# The $m_{bl}$ distribution



## Benchmark 1

- Point allowed by SModelS
- $m_{\tilde{t}_1} \simeq 200 \text{ GeV}$ ,  
 $m_{\tilde{\chi}_1^\pm} \simeq 136 \text{ GeV}$ ,  
 $m_{\tilde{\chi}_1^0} \simeq 50 \text{ GeV}$
- Significant effect that grows monotonically as  $m_{bl}$  decreases

# The $m_{bl}$ distribution



## Benchmark 1

- Point allowed by SModelS
- $m_{\tilde{t}_1} \simeq 200 \text{ GeV}$ ,  
 $m_{\tilde{\chi}_1^\pm} \simeq 156 \text{ GeV}$ ,  
 $m_{\tilde{\chi}_1^0} \simeq 132 \text{ GeV}$ ,
- Milder effect that grows monotonically as  $m_{bl}$  decreases

# Are we sensitive to these effects?

Of course, it depends on the precision on the measurement (stat+syst uncert.) We try to estimate the significance by:

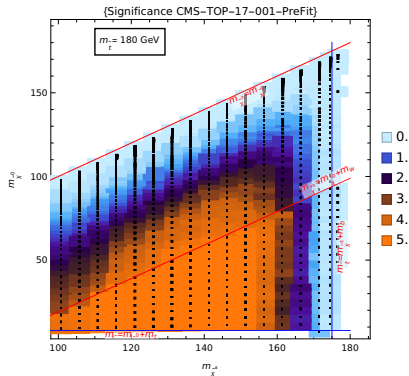
## Significance

- We calculate the signal yield for a given luminosity from the value computed above with `Pythia8`:  $S_i$ , for the  $i$ -th bin
- We compute the background uncertainty from our estimation of the background multiplied by relative uncertainty obtained from the exp. papers:  $\delta B_i = B_i \times u_i$ , where  $u_i$  is the relative exp. uncertainty for the  $i$ -th bin
- For a given point we defined the significance as

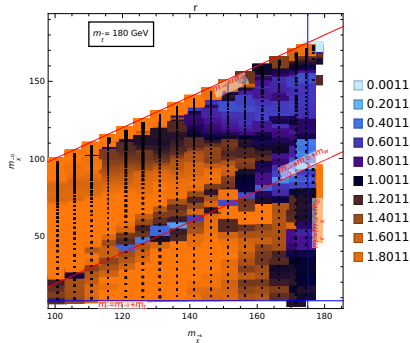
$$z = \sqrt{\sum_i \left( \frac{S_i}{\delta B_i} \right)^2}$$

- We decided to use the “pre-fit” value for the exp. uncertainties

# Exclusion planes: $m_{\tilde{t}_1} = 180$ GeV

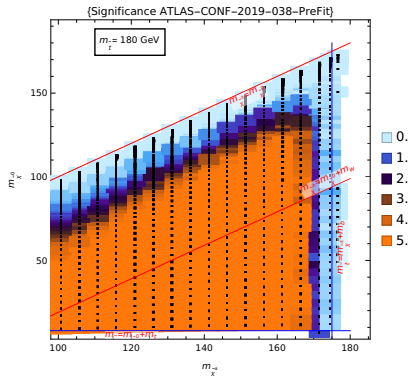


Uncertainties from [CMS, 1812.10505]

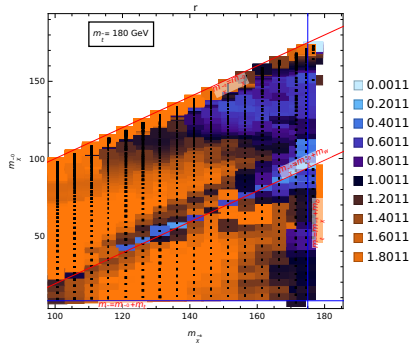


Exclusions from SModelS,  $r = \frac{\sigma(\text{signal})}{\sigma(\text{UL})}$

# Exclusion planes: $m_{\tilde{t}_1} = 180$ GeV

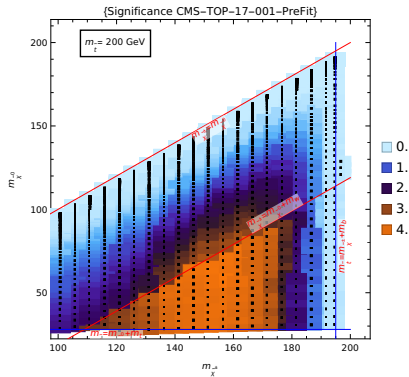


Uncertainties from [ATLAS-CONF-2019-03]

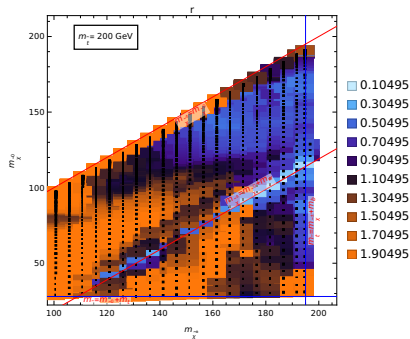


Exclusions from SModelS,  $r = \frac{\sigma(\text{signal})}{\sigma(\text{UL})}$

# Exclusion planes: $m_{\tilde{t}_1} = 200$ GeV

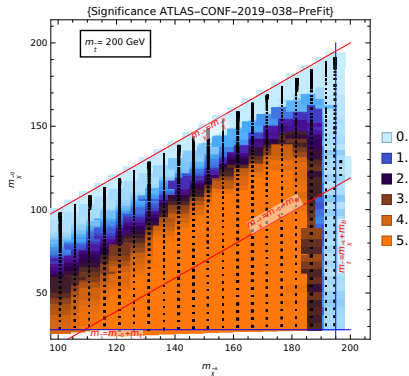


Uncertainties from [CMS, 1812.10505]

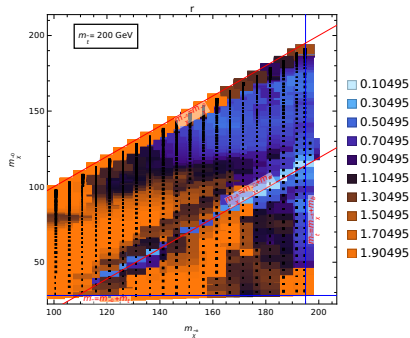


Exclusions from SModels,  $r = \frac{\sigma(\text{signal})}{\sigma(\text{UL})}$

# Exclusion planes: $m_{\tilde{t}_1} = 200$ GeV



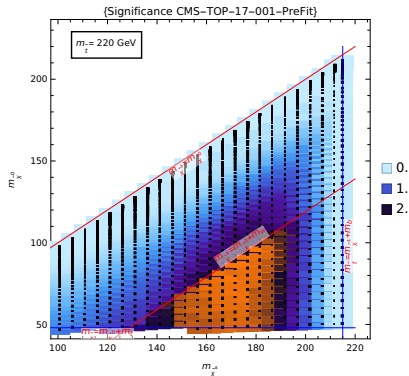
Uncertainties from [ATLAS-CONF-2019-03]



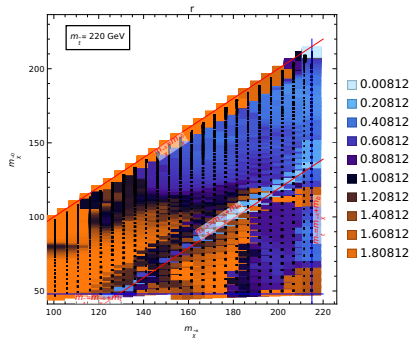
Exclusions from SModels,  $r = \frac{\sigma(\text{signal})}{\sigma(\text{UL})}$



# Exclusion planes: $m_{\tilde{t}_1} = 220$ GeV

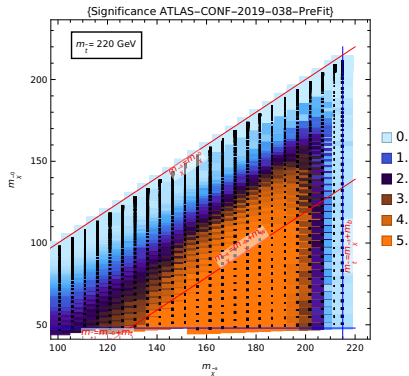


Uncertainties from [CMS, 1812.10505]

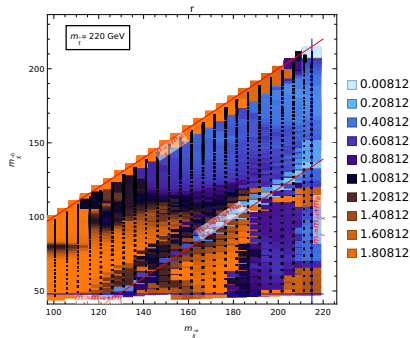


Exclusions from SModelS,  $r = \frac{\sigma(\text{signal})}{\sigma(\text{UL})}$

# Exclusion planes: $m_{\tilde{t}_1} = 220$ GeV



Uncertainties from [ATLAS-CONF-2019-03]



Exclusions from SModelS,  $r = \frac{\sigma(\text{signal})}{\sigma(\text{UL})}$

# Conclusions

## Conclusions

- Realizations of low-scale SUSY are difficult to probe → new strategies are needed
- Looking at observables that up to now were used only for SM measurement is a new avenue to pursue
- We have shown that the  $m_{bl}$  distribution seems to be sensitive to stop production, and it is therefore a new possible target for such searches.
- Similar approach taken by [Agashe et al. in 2310.13687 and 2404.17574] for the  $l + MET$  final state (connection to  $M_W$  extraction)

## Future developments

- Improve analysis with **SModelS** (SR combination enabled)
- Implement the signal and test it in **Contur**
- Release the SLHA files on **Zenodo**
- Develop a more “complete” MSSM scenario