

Supersymmetric EFTs and the Swampland

SUSY 2024 - Madrid
13th Jun 2024

Based on [arXiv:2310.07708]

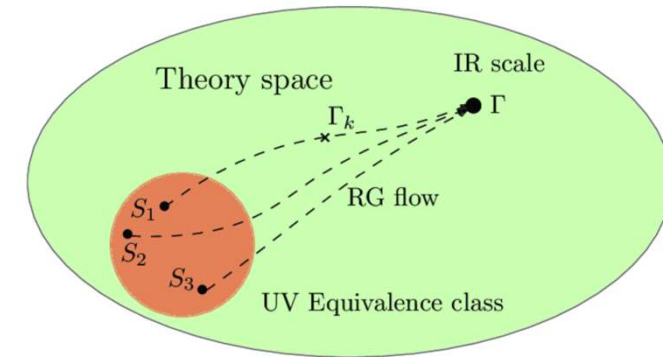
Work with L. Ibáñez and A. Herráez



The Scale of QG

☛ Low energy physics is unavoidably described by Effective Field Theory (EFT) [Weinberg '95]

$$\mathcal{L}_{\text{EFT}, d} = \mathcal{L}_{\text{renorm}} + \sum_{n \geq d} \frac{\mathcal{O}_n}{\Lambda_{\text{UV}}^{n-d}}$$



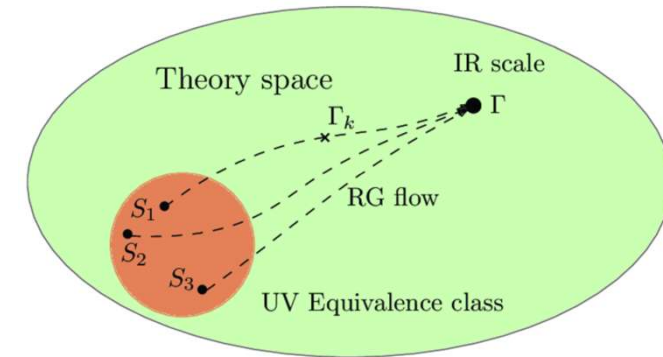
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• For gravitational interactions, the same logic should apply [Donoghue '94]

$$\mathcal{L}_{\text{EFT},d} = \frac{1}{2\kappa_d^2} \left(\mathcal{R} + \sum_n \frac{\mathcal{O}_n(\mathcal{R})}{\Lambda_{\text{QG}}^{n-2}} \right) + (\text{matter}) + \dots$$



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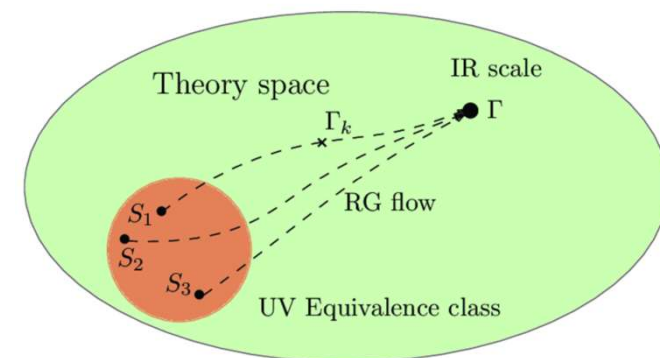
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- Q: Which energy scale should control the EFT expansion in (Quantum) Gravity?

$$\kappa_d^2 = 8\pi G_N \implies [G_N] = E^{2-d}$$

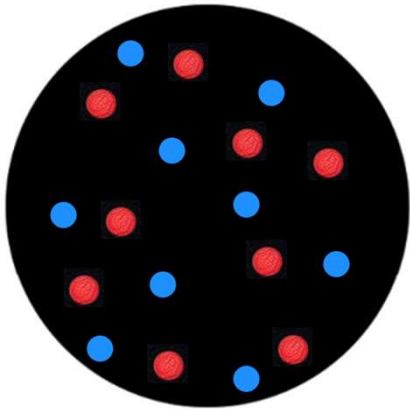
$$M_{\text{pl},d} \sim G_N^{\frac{1}{2-d}} \stackrel{?}{=} \Lambda_{\text{QG}}$$



The Scale of QG (cont'd)

• This naïve answer is essentially correct in the presence of a few number of dofs [AC, Herráez, Ibáñez '21-22]

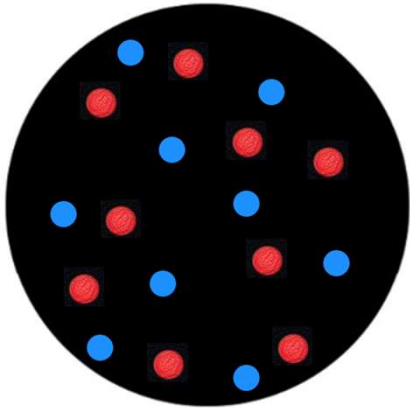
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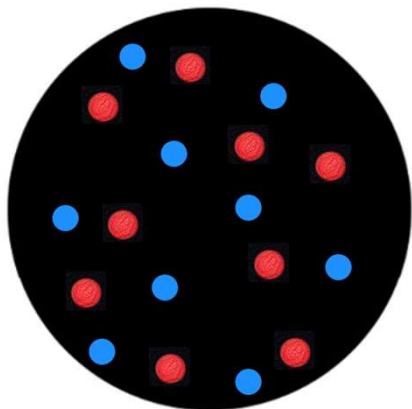


$$\begin{cases} E_{\text{gas}} \sim R^{d-1} T^d \\ S_{\text{gas}} \sim (RT)^{d-1} \end{cases}$$

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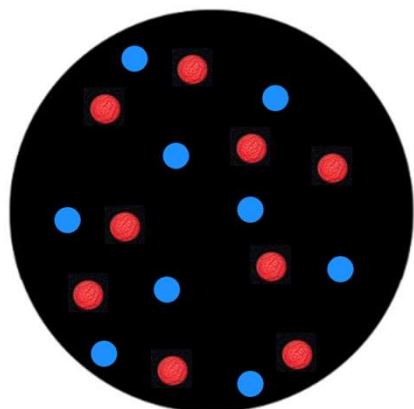


$$\left[\begin{array}{l} E_{\text{gas}} \sim R^{d-1} T^d \\ S_{\text{gas}} \sim (RT)^{d-1} \end{array} \right. \xrightarrow{R \gtrsim \left(\frac{E}{M_{\text{pl},d}}\right)^{\frac{1}{d-3}}} \boxed{S_{\text{gas}} \sim A^{\frac{d-1}{d}} < A}$$

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$$S_{\text{gas}} \sim A^{\frac{d-1}{d}} < A$$

- Everything is consistent with the holographic principle

[Bekenstein, Hawking, Susskind, Bousso, etc.]

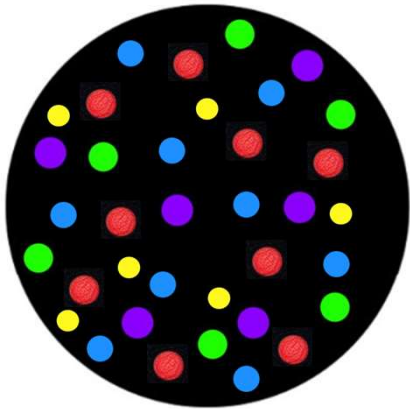


- The minimal size of the box is reached when $A \sim 1$ (in Planck units)

The Scale of QG (cont'd)

👤 But what if N is very large? [AC, Herráez, Ibáñez '21-22]

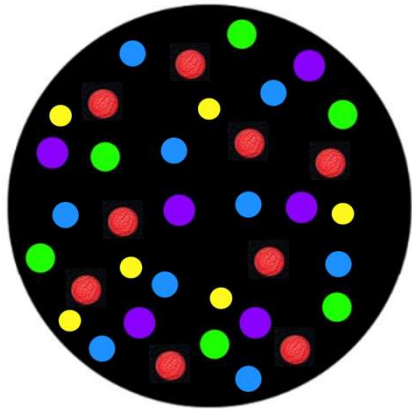
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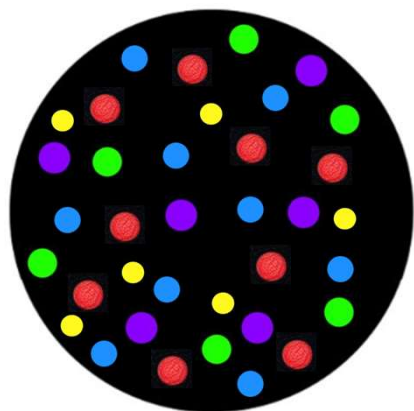


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The Scale of QG (cont'd)

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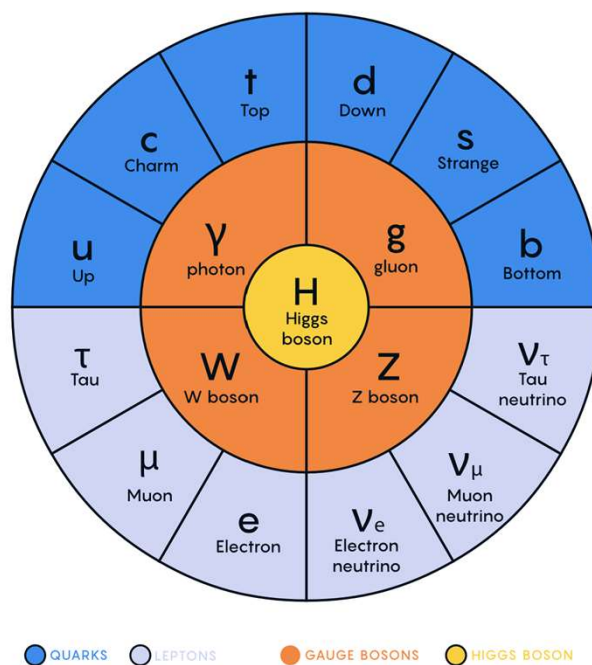
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The holographic entropy seems to be violated if $N \gtrsim A$!!



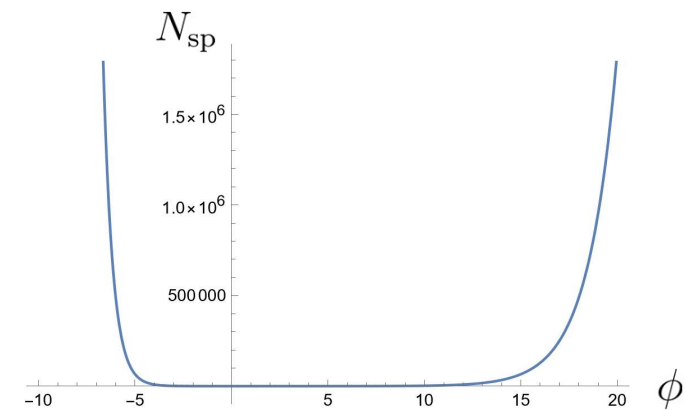
The Scale of QG (cont'd)

- This is essentially the **species problem** [Sorkin, Wald, Zhang, Unruh, etc.]
- Is this **really a problem**? After all, in our Universe and at energies probed by LHC, we only see $\mathcal{O}(10^2)$ particle species



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- Yes!! In **quantum gravity** we know that **the number of species can change** (e.g. decompactifications limits, tensionless string points, conifold loci, etc.)



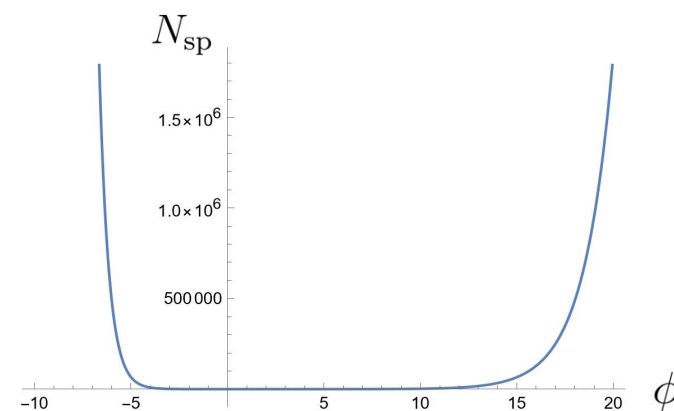
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- Yes!! In **quantum gravity** we know that **the number of species can change** (e.g. decompactifications limits, tensionless string points, conifold loci, etc.)
- In fact, this **must** be the case (By No Global Symmetry Conjecture)
- Turning this logic around:

$$A \geq N \implies \ell_{\min} = \ell_{\text{sp}} := \ell_{\text{pl}} N^{\frac{1}{d-2}}$$

[see also Vafa '24]



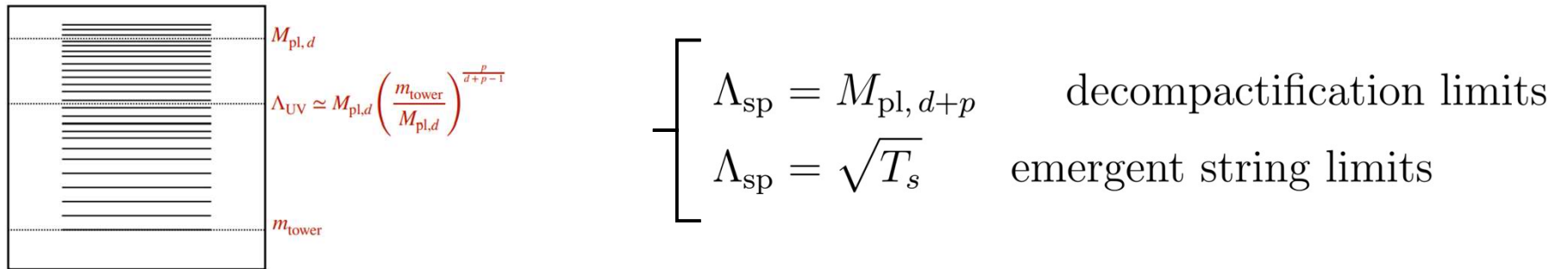
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The Species Scale

- The **Species Scale** is therefore defined (at least asymptotically) as [Dvali, Redi '07; Dvali, Gómez '10]

$$\Lambda_{\text{sp}} \approx \frac{M_{\text{pl},d}}{N^{\frac{1}{d-2}}} \leq M_{\text{pl},d}$$

- In the presence of infinite **towers** of states there can be **parametric decoupling!** [AC, Herráez, Ibáñez '21-22]

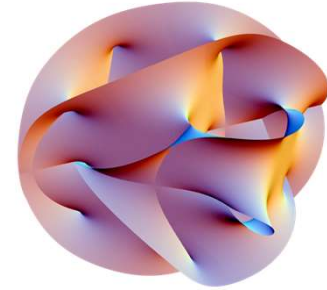


- This can be nicely **encoded** into the formula shown in the first slide [v. d. Heisteeg, Vafa, Wiesner, Wu '22-23] [AC, Herráez, Ibáñez '23]

$$\Lambda_{\text{QG}} = \Lambda_{\text{sp}} \implies \mathcal{L}_{\text{EFT},d} \supset \frac{1}{2\kappa_d^2} \left(\mathcal{R} + \sum_{n>2} \frac{\mathcal{O}_n(\mathcal{R})}{\Lambda_{\text{sp}}^{n-2}} \right)$$

String Theory Checks

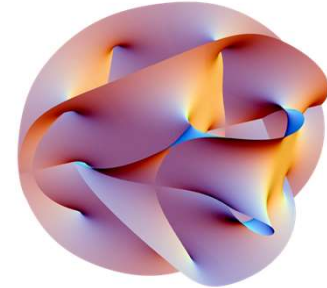
- How can we **check** this in explicit QG/string theory constructions?
- We need **full control** over certain higher-dimensional corrections.



String Theory Checks

- How can we **check** this in explicit QG/string theory constructions?
- We need **full control** over certain higher-dimensional corrections.
- This is where **supersymmetry** enters to our rescue!
- Basic strategy: Look into F-like terms or **BPS** operators in the theory

SUSY 2024



[AC, Herráez, Ibáñez '23, v. d. Heisteeg, Vafa, Wiesner, Wu '23]

$$\text{e.g., } \mathcal{L}_{4d} \supset \int d^2\theta W(\Phi) + \text{h.c.}$$

String Theory Checks

• Multiple (asymptotic) checks of this picture in string theory!

[AC, Herráez, Ibáñez '23, v. d. Heisteeg, Vafa, Wiesner, Wu '23]

10d Type IIB string theory

○ Bosonic action

$$S_{\text{IIB}}^{10\text{d}} = \frac{1}{2\kappa_{10}^2} \int d^{10}x \sqrt{-g} \left(\mathcal{R} - \frac{1}{2}(\partial\phi)^2 \right) - \frac{1}{4\kappa_{10}^2} \int e^{-\phi} H_3 \wedge \star H_3 \\ - \frac{1}{4\kappa_{10}^2} \int \left[e^{2\phi} F_1 \wedge \star F_1 + e^{\phi} \tilde{F}_3 \wedge \star \tilde{F}_3 + \frac{1}{2} \tilde{F}_5 \wedge \star \tilde{F}_5 + C_4 \wedge H_3 \wedge F_3 \right],$$

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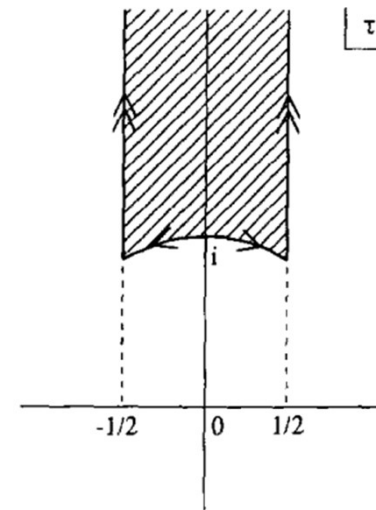
[AC, Herráez, Ibáñez '23, v. d. Heisteeg, Vafa, Wiesner, Wu '23]

10d Type IIB string theory

- Scalar-gravity sector

$$S_{\text{IIB}}^{10\text{d}} \supset \frac{1}{2\kappa_{10}^2} \int d^{10}x \sqrt{-g} \left(\mathcal{R} - \frac{\partial\tau \cdot \partial\bar{\tau}}{2(\text{Im } \tau)^2} \right),$$

$$\text{under } \text{SL}(2, \mathbb{Z}) : \tau \rightarrow \frac{a\tau + b}{c\tau + d}, \quad g_{\mu\nu} \rightarrow g_{\mu\nu}$$



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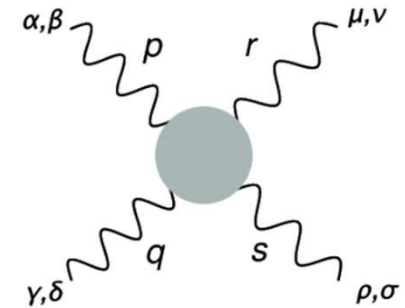
10d Type IIB string theory

- The \mathcal{R}^4 operator

$$S_{\text{IIB}}^{10\text{d}} \supset \frac{1}{\ell_{10}^2} \int d^{10}x \sqrt{-g} E_{3/2}^{sl_2}(\tau, \bar{\tau}) t_8 t_8 \mathcal{R}^4$$

where $t_8 t_8 \mathcal{R}^4 \equiv t^{\mu_1 \dots \mu_8} t_{\nu_1 \dots \nu_8} \mathcal{R}_{\mu_1 \mu_2}^{\nu_1 \nu_2} \dots \mathcal{R}_{\mu_7 \mu_8}^{\nu_7 \nu_8}$

and $E_{3/2}^{sl_2} = 2\zeta(3)\tau_2^{3/2} + 4\zeta(2)\tau_2^{-1/2} + \mathcal{O}(e^{-2\pi\tau_2})$



String Theory Checks

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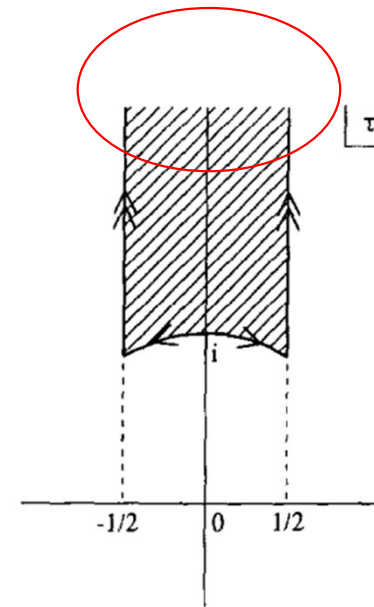
10d Type IIB string theory

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$$\text{therefore } E_{3/2}^{sl_2} \sim \left(\frac{\Lambda_{\text{sp}}}{M_{\text{Pl},10}} \right)^{-6}$$

$$\text{since } \Lambda_{\text{sp}} \sim m_s = \frac{M_{\text{Pl},10}}{(4\pi\tau_2^2)^{1/8}}$$



String Theory Checks

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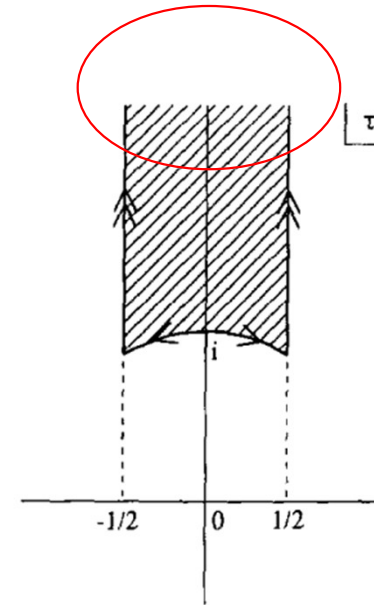
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same for dual copies :
$$E_{3/2}^{sl_2}(\tau, \bar{\tau}) = \left(4\pi^{\frac{3}{4}}\right)^3 \sum'_{(p,q) \in \mathbb{Z}^2} \left(\frac{M_{\text{Pl}, 10}}{\sqrt{T_{p,q}}}\right)^6$$

with
$$T_{p,q} = \frac{2\pi |p + q\tau|}{\ell_{10}^2 \sqrt{\tau_2}}$$



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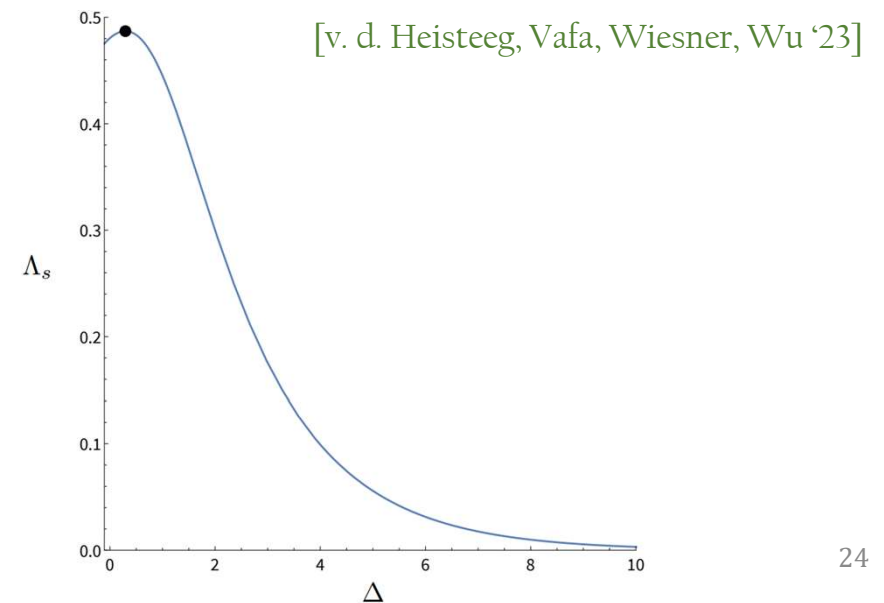
○ Theories with 32 supercharges in $7 < d < 11$

○ Theories with 16 supercharges in 9d

○ Theories with 8 supercharges in 6d, 5d and 4d
[see also AC, Herráez, Ibáñez '23]

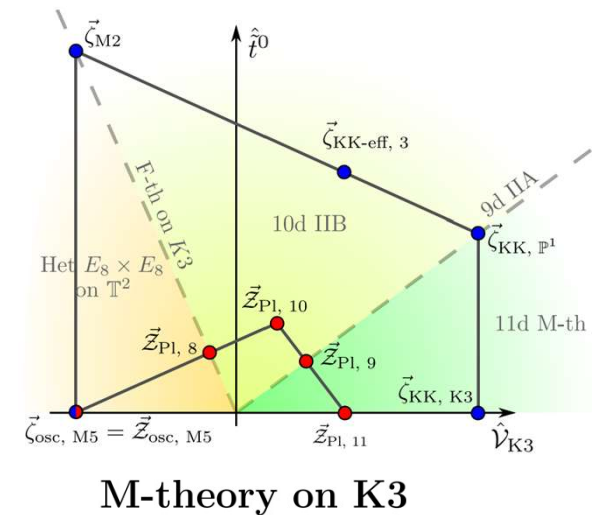
M-theory on Calabi–Yau

$$\Lambda_{\text{sp}} \sim \left(\frac{1}{12} \int c_2 \wedge J \right)^{-1/2}$$



Summary & Outlook

- **Supersymmetry** allow us to study certain relevant quantities for QG in a controlled manner
- The underlying physics is not tied to SUSY, since there are **numerous** motivations
- It also allow us **uncover** new relations/dualities [AC, Ruiz, Valenzuela, '23]
- **Many** things yet to learn!



Thank you for your attention!

QUESTIONS?

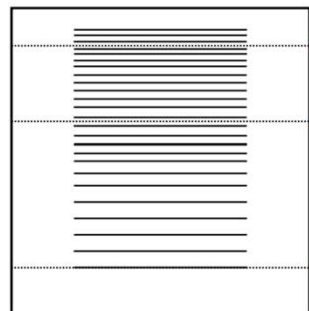
 Contact: alberto.castellano@csic.es

The Species Scale: Definition

• The **Species Scale** is therefore defined (at least asymptotically) as [Dvali, Redi '07; Dvali, Gómez '10]

$$\Lambda_{\text{sp}} \approx \frac{M_{\text{pl},d}}{N^{\frac{1}{d-2}}} \leq M_{\text{pl},d}$$

• In the presence of infinite **towers** of states there can be **parametric decoupling!** [AC, Herráez, Ibáñez '21-22]



The diagram shows a vertical rectangle representing a tower of states. The top horizontal line is labeled $M_{\text{pl},d}$. The bottom horizontal line is labeled m_{tower} . Inside the rectangle, there are many horizontal lines representing states. A dashed horizontal line is labeled $\Lambda_{\text{UV}} \simeq M_{\text{pl},d} \left(\frac{m_{\text{tower}}}{M_{\text{pl},d}} \right)^{\frac{p}{d+p-1}}$.

$$\left[\begin{array}{l} \Lambda_{\text{sp}} = M_{\text{pl},d+p} \quad \text{decompactification limits} \\ \Lambda_{\text{sp}} = \sqrt{T_s} \quad \text{emergent string limits} \end{array} \right.$$

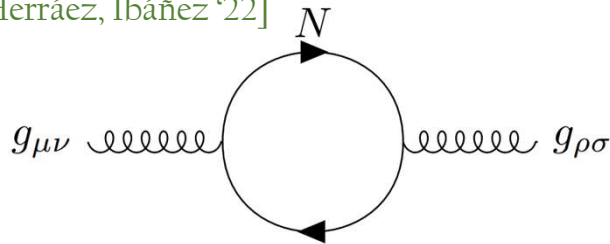
• One can arrive at the same answer via **two** different kind of **arguments**

- i) Perturbative
- ii) Non-perturbative

A (species) cut-off for gravity

- The Species Scale is the scale at which the perturbative series of the graviton breaks down

[see e.g. AC, Herráez, Ibáñez '22]



$$i \Pi^{\mu\nu\rho\sigma} = i (P^{\mu\rho} P^{\nu\sigma} + P^{\mu\sigma} P^{\nu\rho} - P^{\mu\nu} P^{\rho\sigma}) \pi(p^2)$$

$$\pi^{-1}(p^2) = 2p^2 \left(1 - \frac{Np^2}{120\pi M_{\text{pl},4}^2} \log(-p^2/\mu^2) \right)$$

- This can be nicely encoded into the formula shown in the first slide [v. d. Heisteeg, Vafa, Wiesner, Wu '22-23]
[AC, Herráez, Ibáñez '23]

$$\Lambda_{\text{QG}} = \Lambda_{\text{sp}} \implies \mathcal{L}_{\text{EFT},d} \supset \frac{1}{2\kappa_d^2} \left(\mathcal{R} + \sum_{n>2} \frac{\mathcal{O}_n(\mathcal{R})}{\Lambda_{\text{sp}}^{n-2}} \right)$$

- It also explains heuristically why the BH argument fails when $R \sim \ell_{\text{sp}}$

- Important corrections to BH entropy!

[v. d. Heisteeg, Vafa, Wiesner, Wu '23 ;
Calderón-Infante, Delgado, Uranga '23;
Cribiori, Lust, Staudt '23]

String Theory Checks

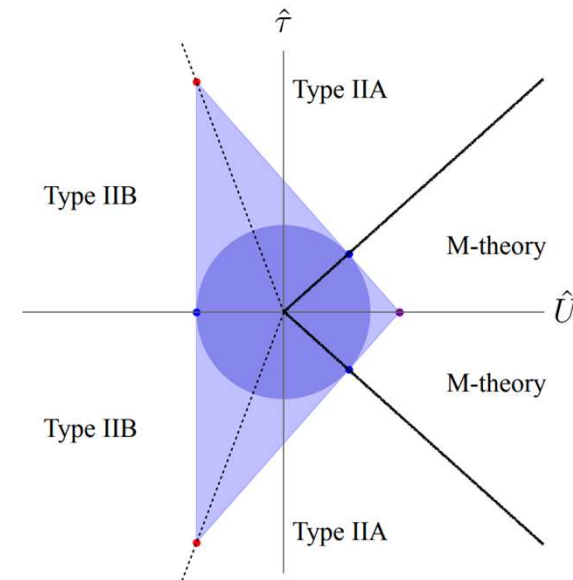
• Multiple (asymptotic) checks of this picture in string theory!

[AC, Herráez, Ibáñez '23, v. d. Heisteeg, Vafa, Wiesner, Wu '23]

○ Theories with 32 supercharges in $7 < d < 11$

M-theory on T^2

$$\Lambda_{\text{sp}} \sim \left(\frac{2\pi^2}{3} \mathcal{V}_2^{6/7} + \mathcal{V}_2^{-9/14} E_{3/2}^{sl_2}(\tau) \right)^{-1/6}$$



[AC, Herráez, Ibáñez '23]

String Theory Checks

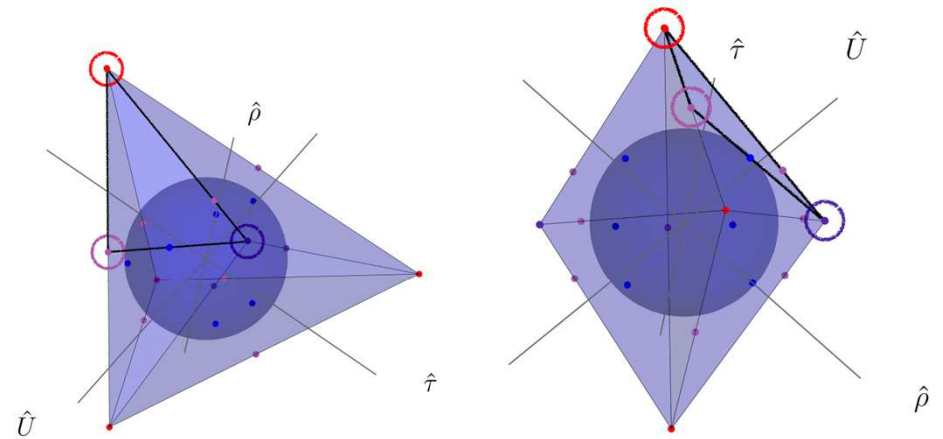
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o Theories with 32 supercharges in $7 < d < 11$

M-theory on T^3

$$\Lambda_{\text{sp}} \sim \left(\hat{E}_{3/2}^{sl_3} + 2\hat{E}_1^{sl_2} \right)^{-1/6}$$



[AC, Herráez, Ibáñez '23]

Interesting examples (Part I)

• One can go beyond the leading \mathcal{R}^4 term in e.g. 10d Type IIB [AC, Herráez, Ibáñez '23]

$$S_{\text{IIB}}^{10\text{d}} \supset \int d^{10}x \sqrt{-g} \left(\frac{1}{\ell_{10}^2} E_{3/2}^{sl_2}(\tau, \bar{\tau}) t_8 t_8 \mathcal{R}^4 + \frac{\ell_{10}^2}{2} E_{5/2}^{sl_2}(\tau, \bar{\tau}) \partial^4 \mathcal{R}^4 \right)$$

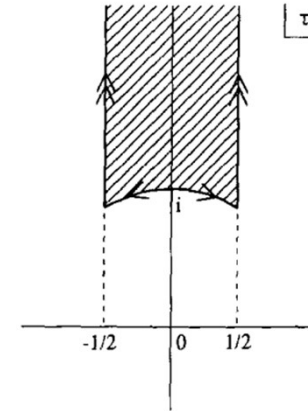
where [Green, Gutperle, Vanhove '97]

$$E_{3/2}^{sl_2} = 2\zeta(3)\tau_2^{3/2} + 4\zeta(2)\tau_2^{-1/2} + \mathcal{O}(e^{-2\pi\tau_2})$$

$$E_{5/2}^{sl_2}(\tau, \bar{\tau}) = 2\zeta(5)\tau_2^{5/2} + \frac{4\pi^4}{135}\tau_2^{-3/2} + \mathcal{O}(e^{-4\pi\tau_2})$$

and

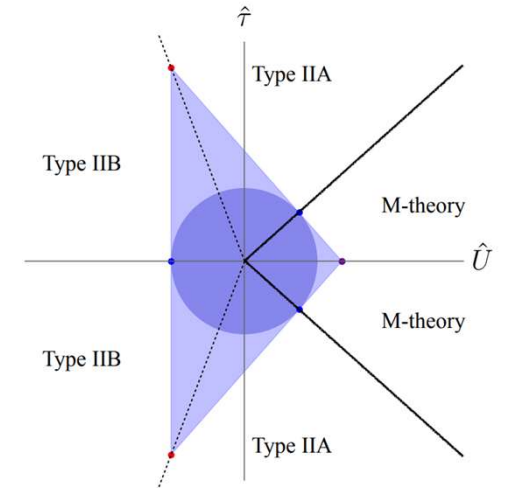
$$m_s = \frac{M_{\text{Pl}; 10}}{(4\pi\tau_2^2)^{1/8}}$$



Interesting examples (Part II)

- Certain higher-curvature corrections in M-theory compactified on tori seem to not follow the pattern

$$\mathcal{L}_{\text{EFT}, d} \supset \frac{1}{2\kappa_d^2} \left(\mathcal{R} + \sum_{n>2} \frac{\mathcal{O}_n(\mathcal{R})}{\Lambda_{\text{sp}}^{n-2}} \right)$$



- E.g. M-theory on T^2 [AC, Herráez, Ibáñez '23]

$$S_{\text{M-th}}^{9d} \supset \ell_9^3 \int d^9x \sqrt{-g} \left(\frac{1}{2} \mathcal{V}_2^{-15/14} E_{5/2}^{sl_2}(\tau) + \frac{2\zeta(2)}{15} \mathcal{V}_2^{27/14} E_{3/2}^{sl_2}(\tau) + \frac{4\zeta(2)\zeta(3)}{15} \mathcal{V}_2^{-18/7} \right) \partial^4 \mathcal{R}^4$$

but

$$\frac{M_{\text{Pl}; 11}}{M_{\text{Pl}; 9}} = (4\pi)^{-2/9} \mathcal{V}_2^{-1/7}$$

$$\frac{M_{\text{Pl}; 10}^{\text{IIA}}}{M_{\text{Pl}; 9}} = (4\pi)^{1/56} \mathcal{V}_2^{-9/112} \tau_2^{-1/16}$$

$$\frac{M_{\text{Pl}; 10}^{\text{IIB}}}{M_{\text{Pl}; 9}} = (4\pi)^{1/56} \mathcal{V}_2^{3/28}$$

Interesting examples (Part III)

- One can even find **infinite families** of grav. operators behaving this way
- E.g. **4d N=2 theories** (say Type IIA on Calabi-Yau)

$$S_{\text{IIA}}^{4\text{d}} \supset \int d^4x \sqrt{-g} \int d^4\theta \sum_{g \geq 1} \mathcal{F}_g(X^A) \mathcal{R}_+^2 F_+^{2g-2} + \text{h.c.}$$

- For **M-theory limits** one finds [Gopakumar, Vafa '98 (x2)]

$$\mathcal{F}_{g>1}^{\text{D0}} = \chi(X_3) \frac{2(2g-1)\zeta(2g)\Gamma(2g-2)}{(2\pi)^{2g}} \frac{\zeta(2g-2)}{m_{\text{D0}}^{2g-2}}$$

- And similarly for **other inf. distance degenerations** (e.g. F-theory/emergent string limits) [AC, Herráez, Ibáñez '23]

A simple explanation

- Consider **KK thresholds** for k-(super)graviton scattering in d dimensions

$$\mathcal{A}_{k,d} = \int_0^\infty d^d p \int_0^\infty \frac{d\tau}{\tau} \sum_n e^{-\tau \left(p^2 + \frac{n^2}{R^2} \right)} \text{tr} \left\langle \prod_{r=1}^k \left(\int_0^\tau dt_r V_{g_{\mu\nu}}(t_r) \right) \right\rangle$$

- If corresponding operator is **BPS** then

$$\mathcal{A}_{k,d} = \tilde{K} \sum_{n \in \mathbb{Z} \setminus \{0\}} \int d^d p \int_0^\infty \frac{d\tau}{\tau} \tau^k e^{-\tau(p^2 + m_n^2)}$$

- For **irrelevant ops** (i.e. $2k > d+1$) the amplitude **converges** [AC, Herráez, Ibáñez '23]

$$\mathcal{A}_{k,d} = 2\tilde{K} \Gamma \left(k - \frac{d}{2} \right) \sum_{n>0} \frac{1}{m_n^{2k-d}} = 2\tilde{K} \Gamma \left(k - \frac{d}{2} \right) \frac{\zeta(2k-d)}{m_{\text{KK}}^{2k-d}}$$

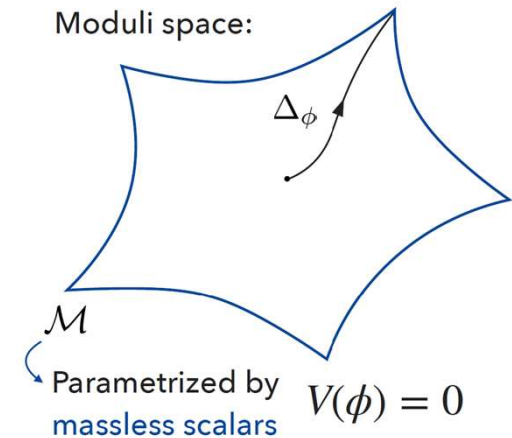
Bounds on exponential rates

It is interesting to study this scale **systematically**. Where?

Distance Conj.
$$\left[\begin{array}{l} m_{\text{tower}} \sim e^{-\lambda \kappa_d \Delta\phi}, \quad \lambda = \mathcal{O}(1) \\ N_{\text{sp}} \rightarrow \infty, \quad \Lambda_{\text{sp}} \sim e^{-\lambda_{\text{sp}} \kappa_d \Delta\phi} \rightarrow 0 \end{array} \right.$$

[Ooguri, Vafa '06]

Moduli space:



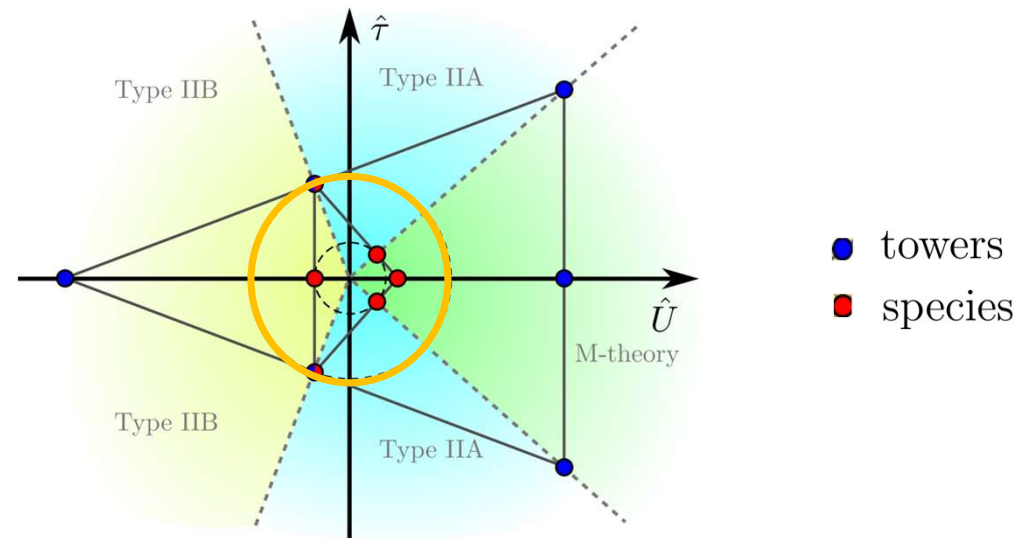
Can we say something **universal** about λ_{sp} ?

... As well as **upper bounds**

$$\lambda_{\text{sp}} \leq \frac{1}{\sqrt{d-2}}$$

[v. d. Heisteeg, Vafa, Wiesner, Wu '22]

[Calderón-Infante, AC, Herráez, Ibáñez '23]



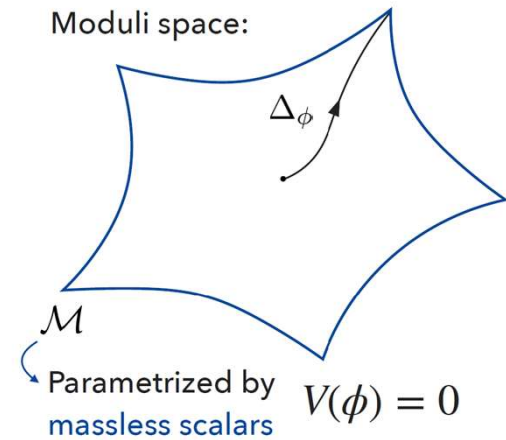
M-theory on T^2

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[Ooguri, Vafa '06]

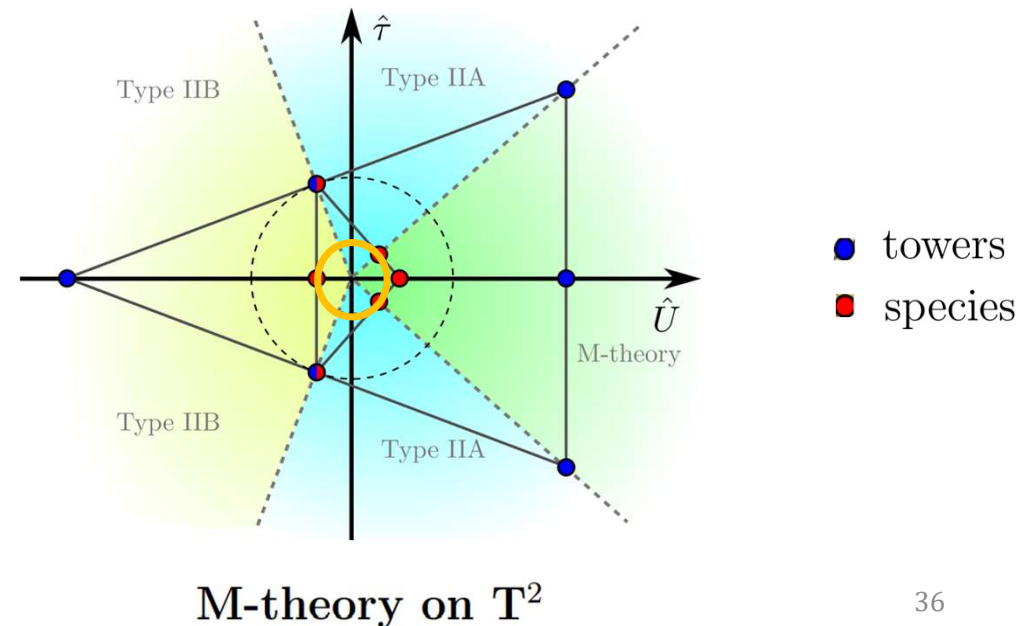


Can we say something **universal** about λ_{sp} ?

Yes!! There seems to be **lower bounds**

$$\lambda_{\text{sp}} \geq \frac{1}{\sqrt{(d-1)(d-2)}}$$

[Calderón-Infante, AC, Herráez, Ibáñez '23]

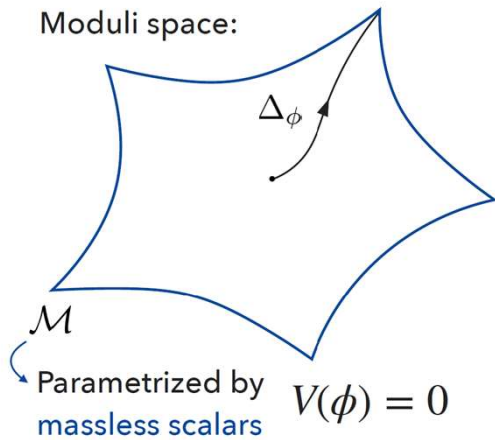


Bounds on exponential rates

- It is interesting to study this scale **systematically**. Where?

Distance Conj.
$$\begin{cases} m_{\text{tower}} \sim e^{-\lambda \kappa_d \Delta\phi}, & \lambda = \mathcal{O}(1) \\ N_{\text{sp}} \rightarrow \infty, & \Lambda_{\text{sp}} \sim e^{-\lambda_{\text{sp}} \kappa_d \Delta\phi} \rightarrow 0 \end{cases}$$

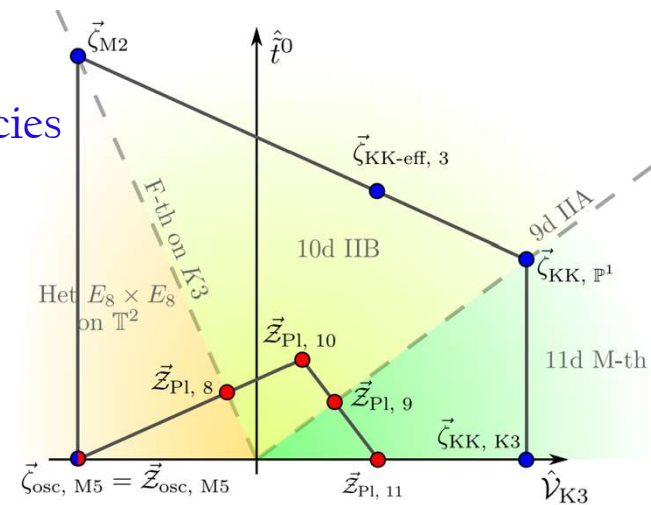
[Ooguri, Vafa '06]



- Can we say something **universal** about λ_{sp} ?
- ... And even some interesting **pattern** relating towers and species

$$\frac{\vec{\nabla} m_t}{m_t} \cdot \frac{\vec{\nabla} \Lambda_{\text{sp}}}{\Lambda_{\text{sp}}} = \frac{\kappa_d^2}{d-2}$$

[AC, Ruiz, Valenzuela, '23]



M-theory on K3