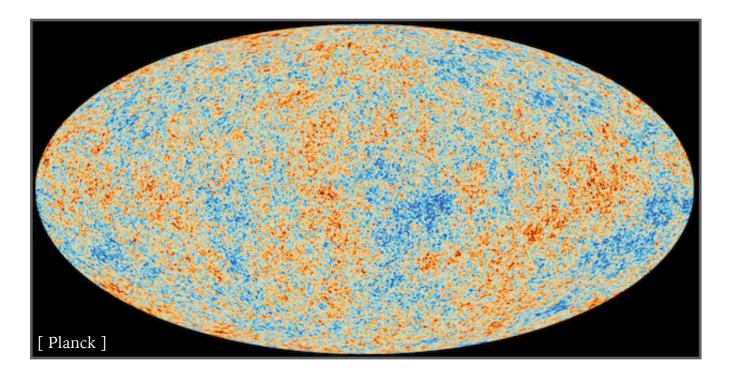
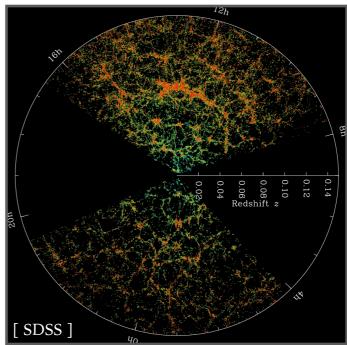


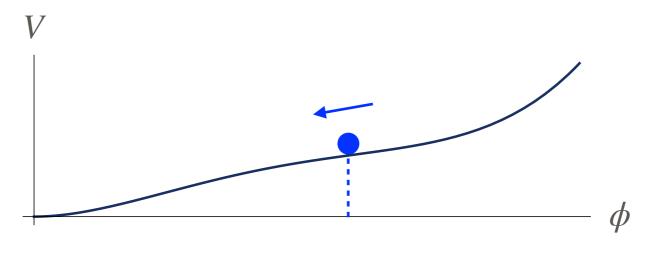
- ► Inflation [Guth '81, Linde '82, '83, Albrecht & Steinhardt '82, Starobinsky '80, Sato '81]
 - A period of time when the Universe undergoes exponential expansion driven by a hypothetical field called "inflaton"
 - Solves the flatness and horizon problems, as well as dilutes unwanted relics
 - Generates primordial seeds for late-time structures [Mukhanov & Chibisov '81]





- ► Inflation and primordial seeds
 - Time evolution of the (coarse-grained value of) inflaton is important

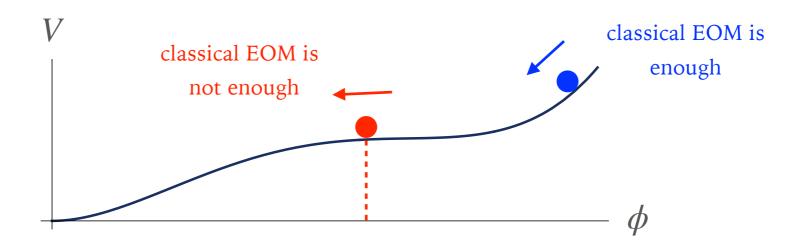
in estimating the amount of the primordial seeds



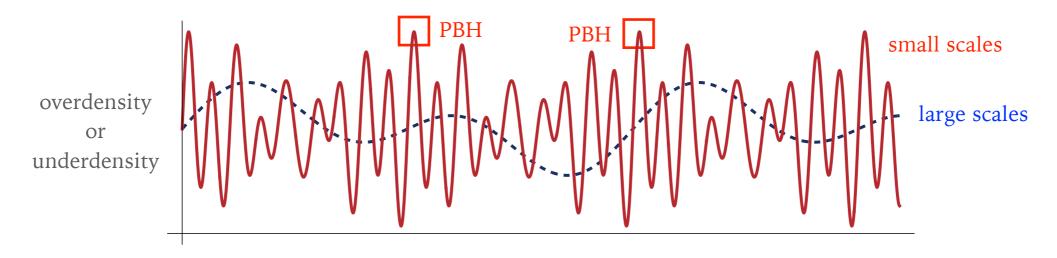
(averaged within the de Sitter Horizon)

- For slow-roll, classical EOM is enough to track the inflaton value: the seeds can be calculated around the classical inflaton trajectory

- ► Inflation and primordial seeds
 - However, the classical EOM is not enough when the potential is extremely flat:



- Such scenarios are often studied in the context of primordial black holes



03 / 14 Ryusuke Jinno (Kobe Univ.) "Borel resummation of secular divergences in stochastic inflation"

- Stochastic approach [Starobinsky '86]
 - Effective description of the inflaton/spectator long-wave modes

$$\phi_{\text{coarse-grained}}(t, \mathbf{x}) = \int \frac{d^3k}{(2\pi)^3} \underbrace{\Theta(\sigma a H - k)}_{\phi} \tilde{\phi}(t, \mathbf{k}) e^{i\mathbf{k}\cdot\mathbf{x}}$$

takes only small k (= coarse-grained modes) into account

- During inflation, $H \simeq \text{const}$ and $a \sim e^{Ht}$, so the Heaviside Θ behaves as

 $\Theta(\underbrace{\sigma a H}_{k} - k) \rightarrow \text{more and more } k \text{ modes contribute at late times}$ exponentially increasing

- The newly contributing modes work as Gaussian white noise to $\phi_{\text{coarse-grained}}$

$$\partial_t \phi_{\text{coarse-grained}} = -\frac{V'}{3H} + \xi$$

From Langevin to Fokker-Planck equation

- Langevin eq.:
$$\partial_t \phi = -\frac{V'}{3H} + \xi \quad \text{with } \langle \xi(t, \mathbf{x})\xi(t', \mathbf{x}') \rangle = \frac{H^3}{4\pi^2} \,\delta(t - t') \,\delta_{\mathbf{x}, \mathbf{x}'}$$
classical force stochastic noise
$$\downarrow$$
- Fokker-Planck eq.:
$$\partial_N P = \frac{1}{3H^2} \partial_\phi (V'P) + \frac{H^2}{8\pi^2} \partial_\phi^2 P$$
classical force stochastic noise

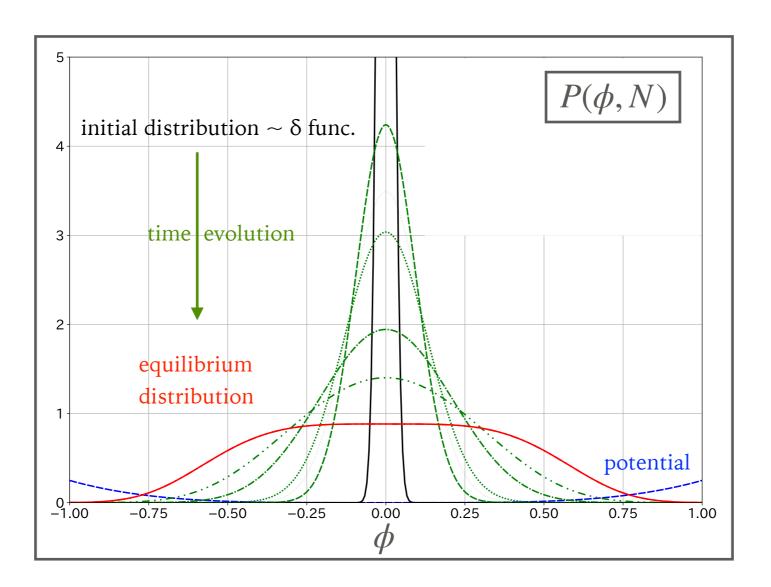
N : rescaled time variable N = Ht (e-folding) $P(\phi, N)$: probability distribution of the coarse-grained inflaton ϕ

> We consider spectator ϕ in the following for simplicity



BEHAVIOR OF THE PROB. DISTRIBUTION FUNCTION

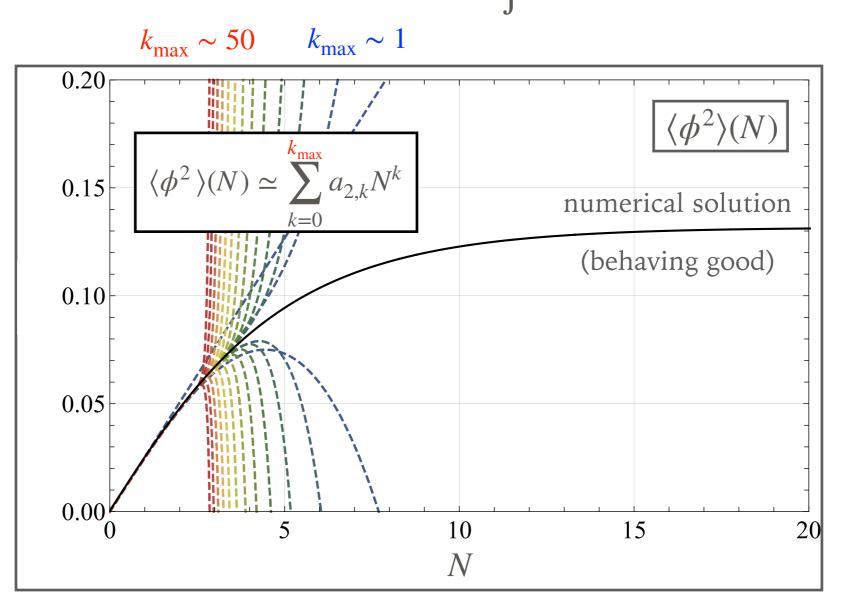
- ► At first sight, Fokker-Planck eq. seems to be well behaved
 - Example : numerical solution for $\partial_N P = \frac{1}{3H^2} \partial_\phi (V'P) + \frac{H^2}{8\pi^2} \partial_\phi^2 P$ with $V = \frac{\lambda}{4} \phi^4$



06 / 14 Ryusuke Jinno (Kobe Univ.) "Borel resummation of secular divergences in stochastic inflation"

BEHAVIOR OF THE CORRELATOR

- However, once analytically expanded, correlators behave badly
 - Example : two-point correlator $\langle \phi^2 \rangle (N) = \int d\phi \ \phi^2 P(\phi, N)$ with $V = \frac{\lambda}{4} \phi^4$



07 / 14 Ryusuke Jinno (Kobe Univ.) "Borel resummation of secular divergences in stochastic inflation"

DIVERGENT BEHAVIOR OF THE CORRELATOR

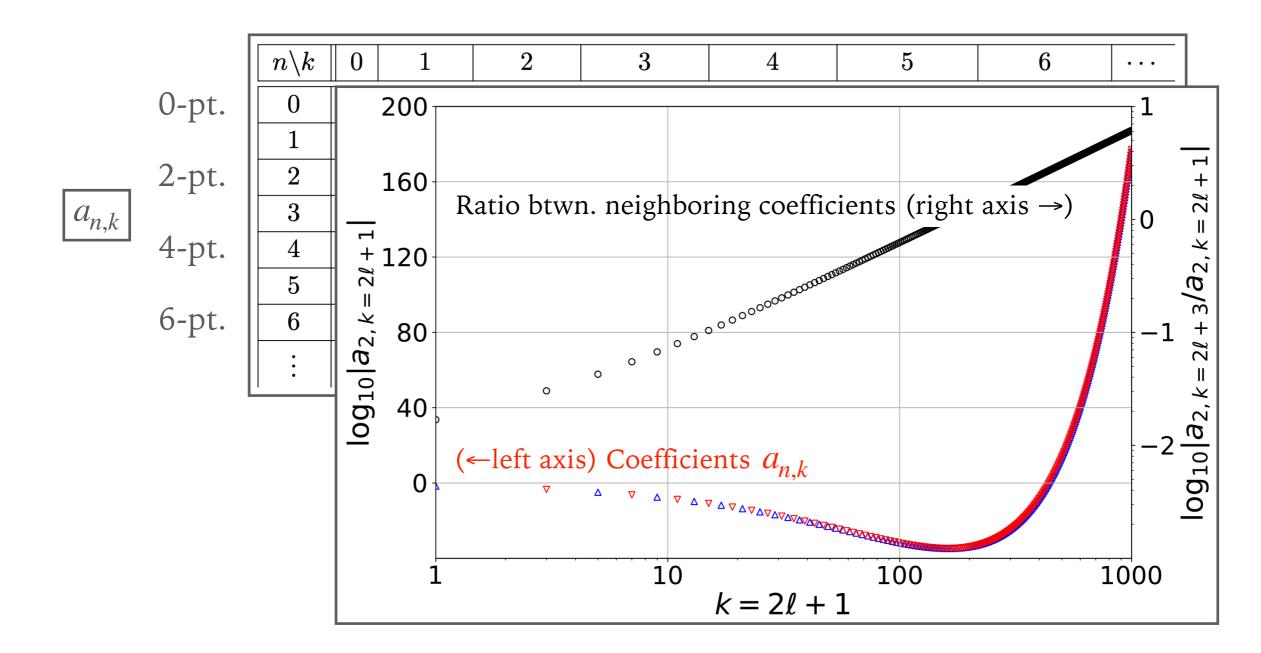
► Coefficients in the 2-point correlator $\langle \phi^n \rangle (N) = \sum_{0}^{\infty} a_{n,k} N^k$ with $V = \frac{\lambda}{4} \phi^4$

		$n \setminus k$	0	1	2	3	4	5	6	•••
$a_{n,k}$	0-pt.	0	1	0	0	0	0	0	0	•••
	_	1	0	0	0	0	0	0	0	•••
	2-pt.	2	0	$1/4\pi^{2}$	0	$-1/24\pi^{4}$	0	$1/80\pi^{6}$	0	•••
		3	0	0	0	0	0	0	0	•••
	4-pt.	4	0	0	$3/16\pi^4$	0	$-3/32\pi^6$	0	$53/960\pi^{8}$	•••
		5	0	0	0	0	0	0	0	•••
	6-pt.	6	0	0	0	$15/64\pi^{6}$	0	$-15/64\pi^{8}$	0	•••
			:	:		•	•	•	•	•.

converging? NO!

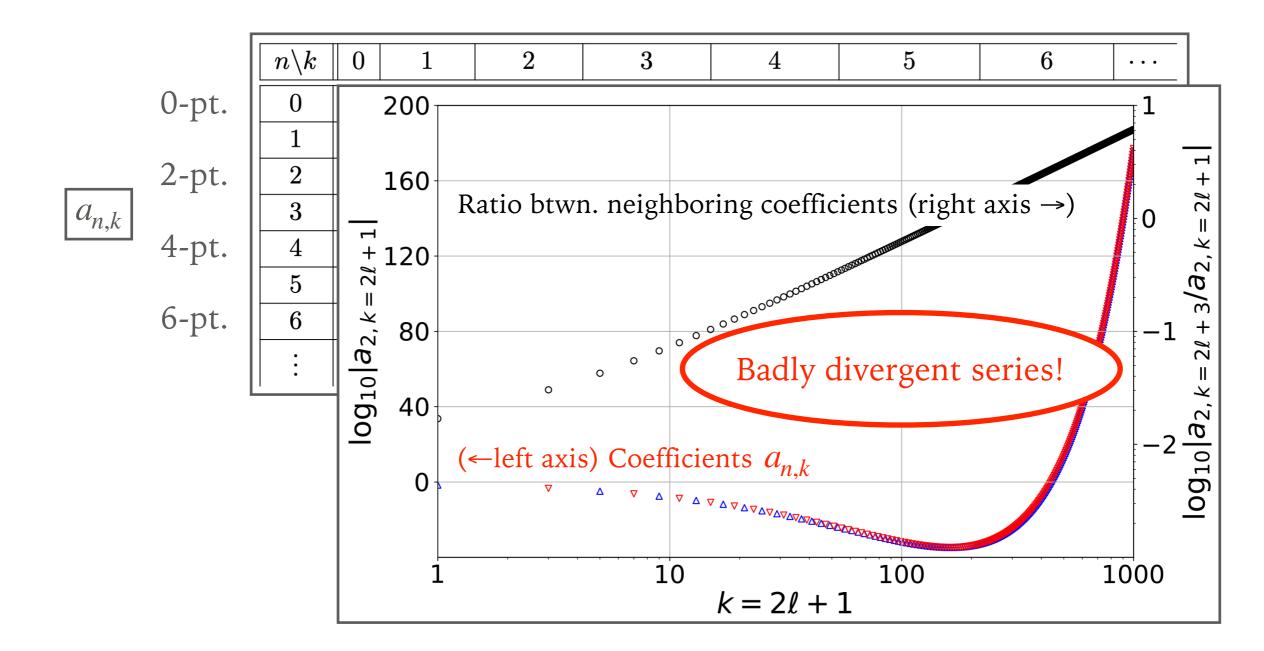
DIVERGENT BEHAVIOR OF THE CORRELATOR

► Coefficients in the 2-point correlator $\langle \phi^n \rangle (N) = \sum_{0}^{\infty} a_{n,k} N^k$ with $V = \frac{\lambda}{4} \phi^4$



DIVERGENT BEHAVIOR OF THE CORRELATOR

► Coefficients in the 2-point correlator $\langle \phi^n \rangle (N) = \sum_{0}^{\infty} a_{n,k} N^k$ with $V = \frac{\lambda}{4} \phi^4$



QUESTIONS

► Truncated correlators $\langle \phi^2 \rangle(N) \simeq \sum_{k=0}^{k_{\text{max}}} a_{2,k} N^k$ are badly divergent

► Q1: Is it because of our poor way of looking at the series?

Or is it truly badly divergent?

► Q2: Any nonperturbative effects in the correlators?



BOREL RESUMMATION

- ► Borel resummation [Borel 1899]
 - Method to construct, from a formally divergent series,

an analytic function that gives the same expansion as the original series

► Example in 3 steps

- Step1: Divergent series
$$\mathcal{O}(N) = \sum_{k=0}^{\infty} \underbrace{(-1)^k k! N^{k+1}}_{a_k} = N - N^2 + 2N^3 - 6N^4 + \cdots$$

- Step2: Borel transf. $\mathcal{BO}(t) = \sum_{k=0}^{\infty} \frac{a_k}{\Gamma(k+1)} t^k = 1 - t + t^2 - t^3 + \cdots = \frac{1}{1+t}$
- Step3: Borel resum. $\mathcal{SO}(N) = \int_0^{\infty} dt \ e^{-\frac{t}{N}} \mathcal{BO}(t) = \int_0^{\infty} dt \ \frac{e^{-\frac{t}{N}}}{1+t} = -e^{\frac{1}{N}} \operatorname{Ei}\left(-\frac{1}{N}\right)^{\frac{1}{N}}$

10 / 14 Ryusuke Jinno (Kobe Univ.) "Borel resummation of secular divergences in stochastic inflation"

BOREL RESUMMATION

- ► Borel resummation [Borel 1899]
 - Method to construct, from a formally divergent series,

an analytic function that gives the same expansion as the original series

- Moreover, the ambiguity of Borel resummation $(\mathcal{SO}(N) = \int_0^{e^{i\theta}\infty} dt \ e^{-\frac{t}{N}} \mathcal{BO}(t)),$

if exists, tells us about nonperturbative properties of the system ("resurgence")

No nonperturbative effects for real N

Nonperturbative effects for real N



10 / 14 Ryusuke Jinno (Kobe Univ.) "Borel resummation of secular divergences in stochastic inflation"

PADÉ APPROXIMATION

- ► Padé approximation to Borel transformation [Padé 1892]
 - Practically, it is often difficult to obtain the coefficients \boldsymbol{a}_k

or their asymptotic form $a_k \simeq a_k^{(\text{asympt.})}$

- Even in such cases, the system can be analyzed with Padé approximation

$$\mathscr{BO}(t) = \sum_{k=0}^{k_{\max}} \frac{a_k}{\Gamma(k+1)} t^k \qquad \simeq \qquad \frac{\sum_{n=0}^m c_n t^n}{1 + \sum_{n=1}^l d_n t^n} \equiv P_{l,m}(t)$$

- Poles of the Padé approximant $P_{m,n}(t)$ tells us about nonpert. effects in the system

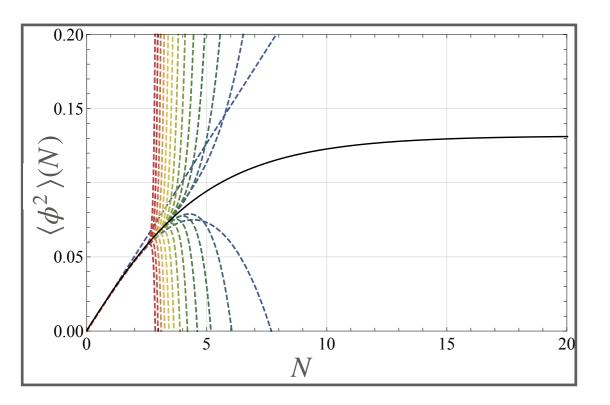
BOREL-PADÉ APPLIED TO CORRELATORS IN STOCHASTIC INFLATION

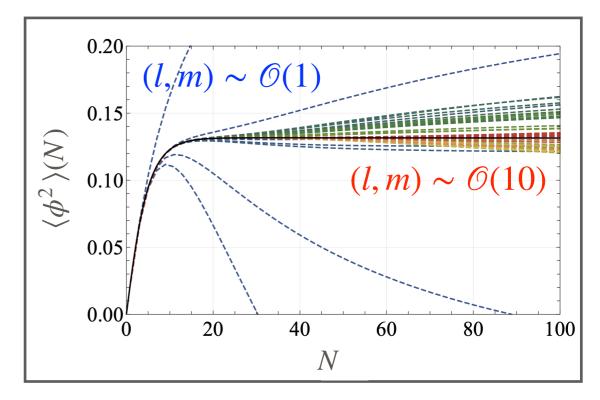
► Naive truncated series

$$\langle \phi^2 \rangle(N) \simeq \sum_{k=0}^{k_{\text{max}}} a_k N^{k+1}$$

- ► Borel-Padé approx.
 - Step1: Divergent series
 - Step2: Borel transf. & Padé approx. (with $(l, m) \sim O(10)$)

Step3: Borel resum.



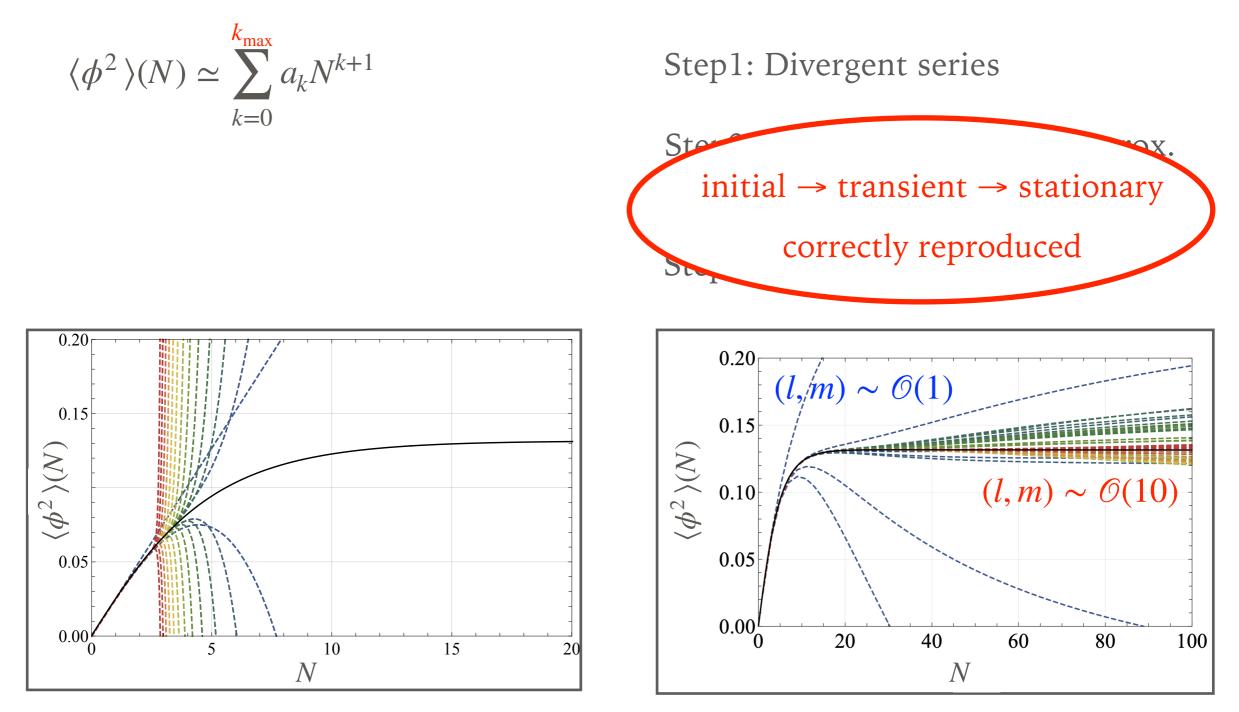


12 / 14 Ryusuke Jinno (Kobe Univ.) "Borel resummation of secular divergences in stochastic inflation"

BOREL-PADÉ APPLIED TO CORRELATORS IN STOCHASTIC INFLATION

Borel-Padé approx.

► Naive truncated series



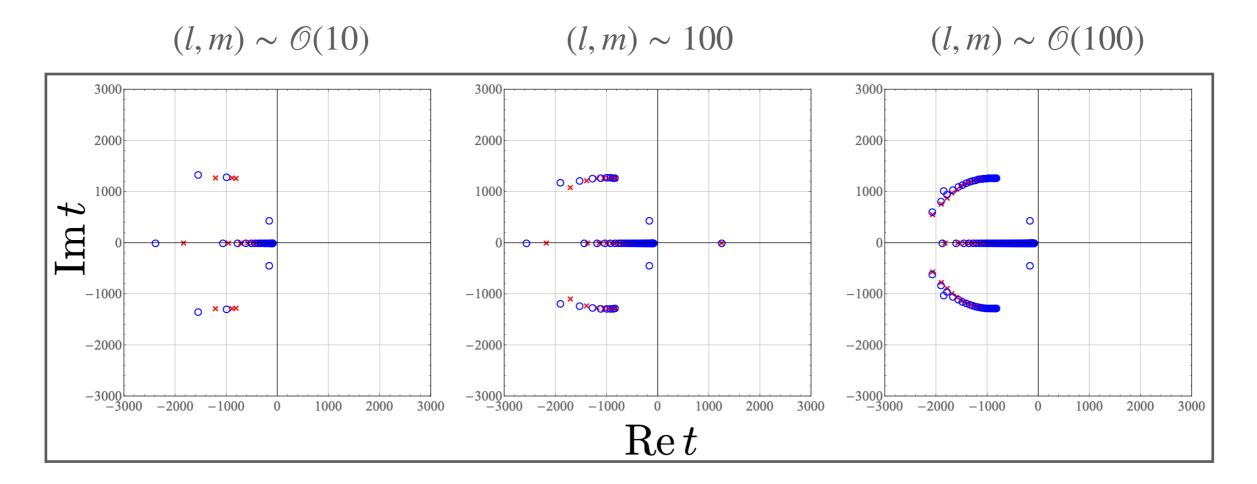
12 / 14 Ryusuke Jinno (Kobe Univ.) "Borel resummation of secular divergences in stochastic inflation"

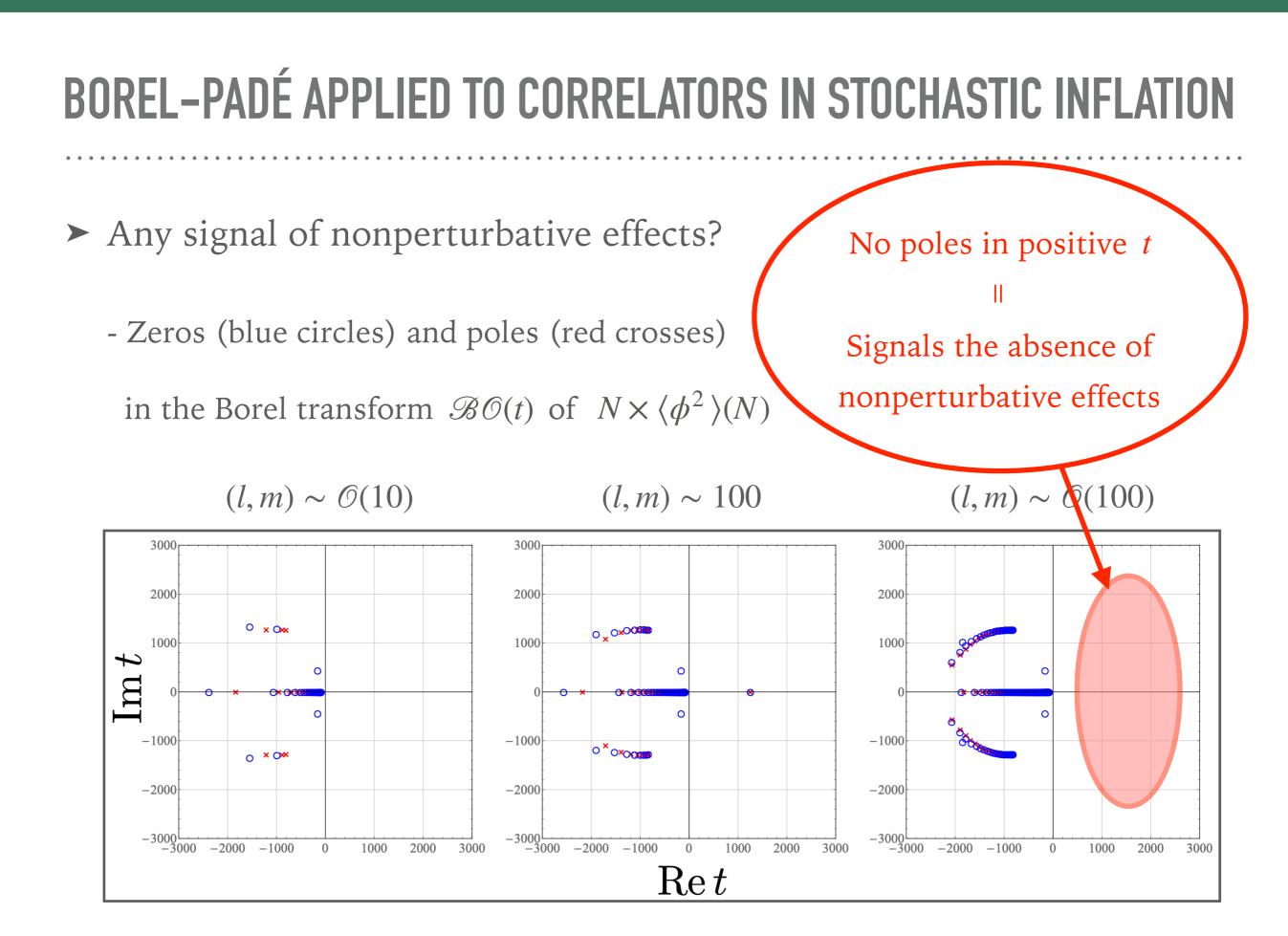
BOREL-PADÉ APPLIED TO CORRELATORS IN STOCHASTIC INFLATION

Any signal of nonperturbative effects?

- Zeros (blue circles) and poles (red crosses)

in the Borel transform $\mathscr{BO}(t)$ of $N \times \langle \phi^2 \rangle(N)$







SUMMARY

- ► In stochastic inflation, correlators are often said to be divergent series
- Using Borel resummation (& Pade approximation),
 we have shown that the series can be brought back to converging ones
- ► From the Borel-plane (= *t*-plane) analysis,

we confirmed no signs of nonperturbative effects in the system

- ► Further questions (~ random thoughts)
 - any cases where correlators have poles in positive *t*?
 - relation to resummation of IR divergence in de Sitter?

spectator \rightarrow inflaton?

