

A modular muon telescope for tomography and radiography applications

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Introduction

• Muography

Muography

- When muons traverse through a given material
	- Deflect from their incident direction due elastic scattering from target nuclei
	- Loose their energy due to inelastic collisions with the atomic electrons in the target
- Muons have very high penetration capabilities through the material
- **Muography** is an imaging technique exploits the scattering and/or absorption characteristics of the atmospheric muons

Muography

Muon Tomography [1,2]

- When cosmic-ray muons crossing a given material, they are affected with elastic multiple coulomb scattering.
- After traversing a macroscopic amount of material, the net angular distribution features an approximately Gaussian core and the standard deviation can be expressed as:

$$
\sigma(\theta) = \frac{13.6 \text{ MeV}}{\beta cp} \sqrt{\frac{x}{X_0}} \left[1 + 0.038 \ln \frac{x}{X_0 \beta^2} \right]
$$

1 X_{0} can be expressed as:

•

$$
\frac{1}{X_0} \propto \frac{Z(Z+1)}{A} \ln\left(\frac{287}{\sqrt{Z}}\right)
$$

- $\sigma(\theta)$ is directly related to *Z*, the basis for scattering based muography.
- **Applications:** Controls of the transport for the detection of the illicit or dangerous materials, nuclear safety, etc.

Muon Radiography [1,2]

This technique is based on measurement of the attenuation of the muon flux intensity by large massive object.

• Expected muon transmission:

$$
T_E(\theta, \varphi, \overline{\rho}) = \frac{\Phi_{ET}(\theta, \varphi, \overline{\rho})}{\Phi_{EFS}(\theta, \varphi)}
$$

• Measured muon transmission:

$$
T_M(\theta,\phi) = \frac{\Phi_{MT}(\theta,\phi)}{\Phi_{MFS}(\theta,\phi)}
$$

• Relative transmission:

$$
R(\theta, \varphi, \overline{\rho}) = \frac{T_M(\theta, \varphi, \overline{\rho})}{T_E(\theta, \varphi)}
$$

Here, θ is zenith angle, ϕ is azimuthal angle, Φ is differential flux. • **Applications:** Railway tunnel investigations

Motivation

• To use the glass RPCs for muography technology

Glass RPC for muongraphy applications

- Resistive plate chambers (RPC) are simpler to construct
- Cost-effective
- Large area ($>$ 10 m²) RPCs can be built $\boxed{3}$
- Efficiency: ≥ 95% at the operating voltage **Schematic of a single gap Resistive Plate Chamber**

- The achieved position resolutions in muography experiments: a few mm [4]
- Time resolution: $<< 1$ ns $[5]$

These characteristics motivate to use the RPCs for muon tomography/radiography applications

[3] *Rev. in Phys. 5 (2020) 100038* [4] *J. Bonnard et al, High performance electronics for a muon telescope operating in an isolated site, DIRE collaboration, Muographers2022 workshop* [5] *IEEE Trans. Nucl. Sci.*, 49 (2002) 881-887

Muon tomography and radiography setups at NISER

θ µ

Layout of the muon telescope for the tomography activities

- 400 mm²
- RPC dimensions = 200×200 mm²
- Readout strips = 9 mm width, 1 mm gap between consecutive strips

We are developing the muon telescopes in modular form for both tomography and radiography applications

This telescope has an upper and lower trackers, each comprise of two RPCs

> This telescope has four layers of RPCs to track the cosmic-ray muons

Layout of the muon telescope

Muon tomography simulation studies

Goals of the studies:

- To image the high-Z materials using the cosmic-ray muon tomography technique
- Test the technique for an application

Pipeline of the simulations work

PoCA algorithm to image the materials^[8]

- For a muon event, reconstruct the muon tracks from the upper and lower trackers
- Find the minimum distance between the two tracks
- Point of Closest Approach (PoCA) point is the mid point of line joining P_A and P_B

(From simulation studies) [8] Nucl. Instrum. Methods Phys. Res. 1014 (2021) 165732

PoCA image of the Lead block (Z=82) with dimensions 100 × 100 × 100 mm³ placed between the upper and lower trackers

High and Medium-Z material discrimination

Scattering angle distribution for Lead (Pb) block of size 100 × 100 × 100 mm³

Scattering angle distribution for Iron (Fe) block of size 100 × 100 × 100 mm³

Larger σ (width of the distribution) observed for larger-Z material

Binned Clustering Algorithm (BCA) [9]

- The 3D space between the two inner detectors D1 and D2 is divided into the voxels of the size $10 \times 10 \times 10$ mm³
- The PoCA points and the corresponding scattering angles are stored in each voxel
- For every pair of muon tracks with scatter vertices $v_i(x_i, y_i, z_i)$ and $v_j(x_j, y_j, z_j)$ located within the same voxel, calculate the metric distance $m_{ij} = ||v_i - v_j||$
- The metric distance is weighted with the scattering angles as $\widetilde{m}_{ij} = \frac{m_{ij}}{A \cdot B}$ $θ_i$. $θ_j$
- Score = Median($\ln(\widetilde{m}_{ij})$)

[9] *JINST 8 (2013) P10013*

RPC2024, Santiago de Compostela, Spain 13 anno 13 ann an t-ann an t-ann an t-ann an 13 an 13 an 13 an 13 an 14 an 14 an 15 an 16 an 17 an

Imaging with Binned Clustering Algorithm (BCA)

- With PoCA algorithm, false PoCA points are recorded outside the Lead block. These false points are from the muon scattering in the acrylic material.
- This issue is resolved using Binned Clustering Algorithm (BCA).

• **The edges of the Lead block are reconstructed better with BCA compared to PoCA**

Imaging spent nuclear fuel dry cask with BCA [10]

- Spent nuclear fuel (SNF) is leftover radioactive material (U^{238}) from nuclear reactors
- It is housed in a dry cask constructed from concrete and steel
- Muon scattering tomography technique can be used for monitoring these casks.

SNF storage dry cask like structure is built using Geant4 with the following components:

Imaging Spent Nuclear Fuel dry cask with BCA

• The dry cask is placed between the upper and lower trackers and BCA is used to reconstruct its image. Outer laver made of concrete

Spent nuclear fuel dry cask placed in between the upper and lower trackers

The XZ projection of the dry cask using BCA reconstruction

The top view of the dry cask containing the Uranium and Copper rods

RPC based muon telescope developments

- Four glass RPCs are built and four acrylic chambers are developed
- Efficiency and time resolutions are studied for one RPC
- Relative Humidity (RH) and temperature studies are performed to an acrylic chamber

Resistive plate chambers and readouts

- Glass RPCs are built with 3 mm thick and 200×200 mm² size electrodes
- PCB readout boards are developed to study the cosmic-ray muon signals detected by the RPCs
- The efficiency and time resolution of one RPC are studied using a muon telescope developed with plastic scintillator detectors

A glass RPC of 160 × 160 mm² active area

A PCB based readout panel A muon telescope developed using plastic scintillator detectors

RPC Efficiency

• The efficiency of RPC is measured using a muon coincidence unit made of plastic scintillator detectors

• The efficiency of RPC is 95% on the plateau

RPC time resolution

- Δt = The time difference between arrival time of RPC pulse and trigger pulse
- ∆t s are plotted for the collected cosmic-ray muon events
- Gaussian function is used to fit the distribution
- Time resolution is the standard deviation (sigma) from the fit

• The time resolution of RPC is 1.4 ns

DAQ and power supply

- ASIC (PETIROC) and FPGA based DAQ system is being tested to study the RPC signals
- Single channel portable power supply module will be used to power all four RPCs

DT5550W DAQ System

PETIROC ASIC:

- Input channels: 32 inputs **(Four ASICs have 128 inputs)**
- Signal polarity: positive or negative
- Dynamic range: 160 fC up to 480 pC
- 32 trigger outputs
- 40 ps bin TDC

CAEN A7512DN single channel power supply module:

- Single output channel
- The maximum output voltage $= 12$ kV
	- 200 mV monitor resolution
- The maximum current = $20 \mu A$
	- 500 pA monitor resolution

CAEN A7512DN

RH&T tests to the acrylic chamber

- RPC performance depends on the ambient Relative Humidity (RH) [11].
- Acrylic chambers are developed to house the RPCs and readout panels
- RH and Temperature (T) are monitored for 30 days inside and outside of an acrylic chamber

Acrylic chamber developed to house the RPC and signal readout panels

- DHT22 sensors are used to measure the temperature and RH.
- One sensor is placed inside the chamber and one outside.
- Arduino UNO R3 board is used to read the data from the sensors.
- RH variation inside the chamber is small whereas it is large outside the chamber
- Same variation of temperature is observed inside and outside the chamber

[11] *JINST* **5** (2010) P02007

Summary and conclusions

- Simulated a glass RPC based telescope for cosmic-ray muon tomography applications using Geant4 toolkit
- Image of a Lead block is reconstructed using PoCA and BCA algorithms. BCA is found be work better
- Simulation studies demonstrate that the designed muon telescope can distinguish medium and high-Z materials
- BCA algorithm is used to image the Spent Nuclear Fuel Dry Cask like structure with the simulated muon telescope
- Efficiency of the glass RPC is 95% to detect the cosmic-ray muons and the time resolution is 1.4 ns.
- Four acrylic chambers are developed to house the RPCs and readout panels. RH and T studies are made to one chamber. RH is well controlled inside the chamber. Temperature variation is same inside and outside the chamber. These tests are being conducted for other three RPCs and acrylic chambers.

Future plans

- Construct the telescopes for both tomography and radiography applications
- Implement the ASIC (PETIROC) and FPGA based DAQ system for cosmic-ray muon data analysis from the telescope
- Test the telescope for imaging the tomography and radiography of the targets

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Thank you

Backup slides

EcoMug Generator

Cosmic Muons Energy Distribution Cosmic Muons Theta(Zenith) Distribution \overline{z} z histGenE 3500 histGenTheta Entries 10000 100000 Entries 10 Mean 0.5727 5235 Mean Std Dev 0.2796 3000F Std Dev 1.51e+04 The zenith angle 2500 distribution of the 10 dN/de atmospheric muons dN/d₀ 2000 approximately Peak Energy follows the function: **500** 5.235 GeV $cos^2(\theta)$ 1000 500 10 Ω -0.5 0.5 1.5 $10²$ $\text{Energy}[\text{MeV}^{\text{D}}]$ $\overline{1}$ 10 $10³$ $10⁴$ $10⁵$ $10⁶$ $10⁷$ Radian Cosmic Muons Phi(Azimuthal) Distribution PDGid \geq $\mathsf{z}% _{T}\!\left(\mathsf{z}\right)$ histGenPhi histGenPDG **Emmies** 10000 ๎<mark>ไม้</mark>ให้ใ_{ในกา}รแห่ง^มหว่^าไม้รางกายในในใน -0.000747 -1.515 Mean 1000 \mathbf{F} Positive 50000 Std Dev 12.91 Muons(PDG id Negative $=-13$ 40000 Muons(PDG 800 dN/de dN/d₀ $id = +13$ The azimuthal angle distribution of the atmospheric muons are uniformly 30000 distributed. charge ratio of muons, is experimentally 20000 400 found to be constant and its value (at sea level) is ~1.28 10000 200 20 -15 -10 -5 10 15 20 -3 3 5 Radian PDGid

Telescope acceptance

Detector developments

Current-Voltage (IV) characteristics of RPC

