

Li 12/11/2023

# Al in Rare Event Search

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# Fast & Slow: Al in Rare Event Search



## Slow

• What is rare event search?

## Fast

- Radiation detectors
- Al algorithms

## **Fast for Slow**

• Fast ML for rare event

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# High Energy Physics Experiment

### High-energy Particle Beam

600 million collisions per second



### **PIP-II Neutrino Beam**

Trillions of neutrinos per second

**Deep Underground Neutrino Experiment** 



Probability of detecting electron, muon and tau neutrinos



# **Naturally Occurring Neutrinos**









A.





# **Rare Event Search in 1950s**



## The Cowan-Reine Neutrino Experiment

First detection of neutrino (via inverse beta decay):

$$\bar{\nu_e} + p \to n + e^+$$

Extremely low cross section, but unique signature:

• 
$$e^+ + e^- \rightarrow 2\gamma$$

• Neutron capture  $\gamma$ 



Nobel Prize of 1995



# **Rare Event Search in 2023**

## Double Beta Decay $(2v\beta\beta)$

First proposed by Maria Goeppert Mayer in 1935 First detection by Elliott, Hahn, Moe, in 1987 Decay half-life  $T_{\frac{1}{2}} \sim 10^{14} - 10^{24} yrs$ 

Much longer than the age of universe!





## Neutrinoless Double-Beta Decay ( $0v\beta\beta$ )

- **ΔL = 2 lepton number violation** process
- Explain the matter-antimatter asymmetry in our universe
- Changes our fundamental understanding of particle physics
- Has not been observed at  $T_{\frac{1}{2}} > 10^{26} yrs$



## **Rare Event Search in 2023 Dark Matter**

#### The evidence for the existence of dark matter has been plenty





#### Large Scale Structure Formation

Gravitational Lens



#### Cosmic Microwave Background





## **Rare Event Search in 2023 Dark Matter**

None has been observed.

## WIMP: Weakly Interacting Massive Particle



- The evidence for the existence of dark matter has been plenty
- Many DM candidates have been proposed (WIMP, Axion, etc.)

# What Makes Rare Event Search Hard?

## It is extremely rare! Using $0v\beta\beta$ as an example ...

- We have not seen  $0v\beta\beta$  at half life of  $T_{\frac{1}{2}} > 10^{26}yrs$
- Next-generation experiments typically aims at  $T_{\frac{1}{2}} > 10^{28} yrs$  (×100 improvement)
- Correspond to 3-4 event after 10 years of data taking



# What Makes Rare Event Search Hard?



•1 event every 2.5-3.3 year, we need ultra-sensitive detector to capture every event • As our detector gets more sensitive, we also collect lots of events that are not 0vββ/WIMP DM Search for needle in a haystack







## What Makes Rare Event Search Hard? Ουββ **WIMP Dark Matter** The Cowan-Reine Exp.



Naturally radioactive and cosmic ray background

## **Control background is of unparalleled** importance in rare event search experiment!



# The Rare Event Search Pipeline



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## KamLAND-Zen Monolithic Liquid Scintillator Detector for 0vßß Search





#### From Left to Right:

- **Dr. Christopher Grant (BU Co-PI)**
- Hasung Song (BU)
- **Dr. Lindley Winslow (MIT, Co-PI)**
- Dr. Spencer Axani (MIT/UDelaware)
- Dr. Zhenghao Fu (MIT/Jump Trading)
- Dr. Joseph Smolsky (MIT/CSU)
- Dr. Aobo Li (BU/UNC/UCSD) Not on this photo:
- Dr. Sumita Ghosh (MIT)
- So Young Jeon (BU)

## The MIT-BU Analysis Group











UNIVERSITY of HAWAI'I°









## KamLAND-Zen Monolithic Liquid Scintillator Detector for 0vßß Search



## **Inner Detector PMTs**

## **Background Source**

 XeLS Background Film Background

Canon

## **Liquid Scintillator**

25-µm-thick transparent nylon film

## **Xenon Loading**

Load double beta decay isotope <sup>136</sup>Xe in LS inside inner balloon (XeLS)





# KamLAND-Zen Data



#### → 23% Quantum Efficiency ... 500 photons will produce a signal

... 500 photons will produce a signal (photoelectron).

## **Triggered PMT**

## 22% Photocoverage

... 2,200 photons will reach PMT ...





## KamLAND-Zen Data **Triggered PMT** θ-φ Sphere Map (-14.0 ns, -12.5 ns) $(R, \theta, \phi, t, q) \rightarrow E = \Sigma q$ 0.08 0.07 0.06 Normalized Amplitude 0.05 0.04 0.03 0.02 0.01 0.00 -10



## **Spatiotemporal Data**

A time series of 2D images, projected onto sphere (A spherical video)









## A Time Series of 2D Images ... **Attention Mechanism ConvLSTM**

Convolutional Long-Short Term Memory (LSTM) Network



Produce context images & provide interpretability





# ... Project onto A Sphere

Cohen, Taco et al. "Spherical CNNs." ICLR 2018









## **Spherical CNN**

SO(3) symmetry & rotational invariance



## KamNet: An Integrated Spatiotemporal Neural Network

## **Spatiotemporal Data**

A time series of images projected onto Sphere



## **AttentionConvLSTM**

for Spatiotemporal symmetry





Context Images  $(c, \theta, \phi)$ 

## KamLAND-Zen



# KamNet-enabled Background Rejection

Monolithic LS detector has been at the heart of many great discoveries in neutrino physics ...







KamNet-enabled E	3;	ac
e <sup>-</sup>		0.14
		0.12
WEWE	olitude	0.10
<ul> <li>Signal are strictly single-vertex events</li> <li>All energy deposited almost immediately</li> </ul>	ad Amp	0.08
An energy deposited annost miniculatory	malize	0.06
$e^{-}$ Less than a few $\gamma$	Nor	0.04
ns later		0.02
WEW		

- Most backgrounds are closely-spaced multi-vertex events
  - part of event energy is deposited by cascading  $\gamma$ s that slightly alter event topology

KamNet captures this tiny alteration in event topology to efficiently reject most backgrounds in KamLAND-Zen!



![](_page_22_Figure_5.jpeg)

![](_page_22_Picture_6.jpeg)

# **KamNet-enabled Background Rejection**

While accepting 90% of  $0v\beta\beta$  events, KamNet rejects ~27% of

XeLS backgrounds and ~59% of film backgrounds

KamNet is **independent** and **multiplicative** to all existing background rejection methods in KamLAND-Zen

**Long-Lived Spallation** 

![](_page_23_Figure_4.jpeg)

![](_page_23_Figure_5.jpeg)

backgrounds allows for the expansion of the fiducial volume from 157cm to 165.8cm, resulting in 17.7% gain on exposure

![](_page_23_Figure_7.jpeg)

# **KamNet-enabled New Search**

**Exposure Before KamNet:** 970 kg·yr

<u>Apply KamNet to High-Background</u> Period Only:

• Conservative use of KamNet

 Veto critical backgrounds that passes all traditional methods

**Official KamLAND-Zen 800 Limit:** 

 $T_{1/2}^{0\nu\beta\beta} > 2.0 \times 10^{26} \text{yr} (90 \% \text{ C}.\text{L}.)$ 

#### **American Physical Society 2023 Dissertation Awards In Nuclear Physics**

![](_page_24_Figure_9.jpeg)

## Exposure After KamNet: 1142 kg·yr

+17.7%

![](_page_24_Picture_12.jpeg)

Worth \$2.5 million!!! (Based on 2010 Xe price)

**Official KamLAND-Zen 800 Limit:** 

 $T_{1/2}^{0\nu\beta\beta} > 2.0 \times 10^{26} \text{yr} (90 \% \text{ C}.\text{L}.)$ 

**KLZ Combined Official Limit:** 

 $T_{1/2}^{0\nu\beta\beta} > 2.3 \times 10^{26} \text{yr} (90 \% \text{ C}.\text{L}.)$ 

This Xe  $0\nu\beta\beta$  search represents the **worlds most stringent limit** on the effective Majorana mass

#### Apply KamNet to All Data:

 $T_{1/2}^{0\nu\beta\beta} > 2.7 \times 10^{26} \text{yr} (90 \% \text{ C. L.}) +35\%$ 

![](_page_24_Picture_21.jpeg)

![](_page_24_Picture_22.jpeg)

![](_page_24_Picture_23.jpeg)

![](_page_24_Picture_24.jpeg)

## Large Enriched Germanium Experiment for Neutrinoless ββ Decay – LEGEND 56 Institutions, about 270 scientists

![](_page_25_Picture_1.jpeg)

LEGEND mission: "The collaboration aims to develop a phased, <sup>76</sup>Ge based double-beta decay experimental program with discovery potential at a half-life beyond 10<sup>28</sup> years, using existing resources as appropriate to expedite physics results."

Univ. New Mexico L'Aquila Univ. and INFN Univ. Texas, Austin Lawrence Berkeley Natl. Lab. Univ. of Warwick Univ. California, Berkeley Leibniz Inst. Crystal Growth Univ. of Indiana Comenius Univ. Simon Fraser Univ.

Univ. of North Carolina Max Planck Univ. of South Carolina Tennessee Tech. Jagiellonian Univ. Tech. Univ. – Dresden Joint Inst. Nucl. Res. Inst. Duke Univ. Triangle Univ. Nuclear. Lab. Los Alamos Joint Res. Centre, Geel

Inst., Heidelberg	Milano Univ. and Milano INFN	North Carolina State
iv.	Inst. Nucl. Res. Russ. Acad. Sci.	South Dakota Mines
ssee	Natl. Res. Center Kurchatov Inst.	Univ. of Regina
Science Inst.	Lab. Exper. Nucl. Phy. MEPhl	Roma Tre
Gran Sasso	Max Planck Inst., Munich	Univ. Washington
niv.	Tech. Univ. Munich	SNOLAB
ool	Oak Ridge Natl. Lab.	Laurentian Universi
e London	Padova Univ.	Univ. Tuebingen
Natl. Lab.	Padova INFN	Univ. South Dakota
o Bicocca	Czech Tech. Univ. Prague	Univ. Zurich

![](_page_25_Picture_10.jpeg)

![](_page_25_Picture_11.jpeg)

![](_page_25_Picture_12.jpeg)

![](_page_25_Picture_13.jpeg)

![](_page_26_Figure_0.jpeg)

![](_page_26_Picture_1.jpeg)

90 x 76 mm 2.68 kg

**High Purity Ge Detector (HPGe)** 

<sup>76</sup>Ge is a double-beta decay isotope

92 x 112 mm 4.11 kg

![](_page_26_Figure_6.jpeg)

![](_page_26_Picture_7.jpeg)

# **LEGEND** data

![](_page_27_Picture_2.jpeg)

![](_page_28_Figure_0.jpeg)

# Improve LEGEND Baseline Model

**Upper turning edge: Critical** 

Waveform tail: Irrelevant

**Rising edge: Critical** 

**baseline:** Bias

## Background Rejection Hooking a fully connected network(

to the LEGEND baseline model to identify and reject background

#### Interpretable

Allow students to see where LBM pays attention to to make decision

![](_page_28_Figure_12.jpeg)

![](_page_28_Picture_13.jpeg)

## Feature Importance Supervision (FIS) Guide ML model to be Right for the Right Reason

![](_page_29_Figure_1.jpeg)

## Feature Importance Supervision (FIS) Guide ML model to be Right for the Right Reason

## **Quantifying Bias**

Selected dataset to test for energy dependency bias

- Biased classifier: nontrivial classification power

![](_page_30_Figure_5.jpeg)

## **Rejecting Background**

Identify multi-site background in HPGe detector

# **XENONNT** 2-Phase Liquid Xenon TPC for WIMP DM Search

![](_page_31_Figure_1.jpeg)

## **XENONNT** 2-Phase Liquid Xenon Time Projection Chamber for WIMP DM Search

![](_page_32_Figure_1.jpeg)

### **Spatiotemporal Data** A 2D flat video

![](_page_32_Figure_3.jpeg)

# **XENONnT Data**

### **Spatiotemporal Data** A 2D flat video

![](_page_33_Figure_2.jpeg)

Squash along time dimension

Squash along (x,y) dimension

![](_page_33_Figure_5.jpeg)

Time Index

![](_page_33_Picture_6.jpeg)

Time Index

# **Dual Classifier**

### **Hit Pattern** 2D Image

![](_page_34_Figure_2.jpeg)

![](_page_34_Figure_3.jpeg)

![](_page_34_Figure_4.jpeg)

![](_page_34_Figure_7.jpeg)

## **Combining Dual Classifier** Combine CNN & RNN to classify confused events

![](_page_35_Figure_1.jpeg)

## Min Zhong

![](_page_35_Picture_3.jpeg)

![](_page_35_Picture_4.jpeg)

![](_page_35_Figure_5.jpeg)

# Fast & Slow: Al in Rare Event Search

![](_page_36_Picture_1.jpeg)

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# **New Electronics for KamLAND-Zen**

### 16-channel prototype for KamLAND2-Zen

![](_page_37_Picture_2.jpeg)

## **Primary Goals:**

- Digitize waveform during the chaotic period after a muon passes through the detector in order to record all neutrons, allowing us to reduce the Long-Lived spallation background.
- 2. Streaming data (deadtime free system), large data throughput.
- 3. Large memory buffers.

Reduction in PCB footprint	Machine learning on FPGA	*50% cost savings	*30-40% pow consumptio savings
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\* compared to standard RF signal chain

![](_page_37_Picture_9.jpeg)

# **Al Trigger for Low Energy**

**Exponentially Increasing Event Rate** Lots of electronic noise

![](_page_38_Figure_2.jpeg)

![](_page_38_Figure_3.jpeg)

# **Real-Time Al Analysis**

![](_page_39_Picture_1.jpeg)

![](_page_39_Figure_2.jpeg)

![](_page_39_Picture_3.jpeg)

#### Data Stream

![](_page_39_Picture_5.jpeg)

# Fast & Slow: Al in Rare Event Search

## Thank you for you liaobo77@ucsd.edu if you

## Slow

- Rare event search provides a unique window to unravel the mystery of Neutrino & Dark Matter
- High sensitivity and low background is required

## Fast

- Radiation detectors: KamLAND-Zen, LEGEND, XENONnT
- Al algorithms: KamNet, FIS, Dual Classifier

## **Fast for Slow**

- Hardware: RFSoC for KamLAND-Zen and my other experiments
- Algorithms: AI Trigger and online AI analysis
- Thank you for your attention. Please email
- liaobo77@ucsd.edu if you are interested in collaborating!

![](_page_40_Figure_13.jpeg)