

# DUNE Systematic Flux Uncertainties

Ian D. Kotler FRAS on behalf of the DUNE Collaboration

[APS DPH-PHENO 2024](#)

May 14<sup>th</sup>, 2024

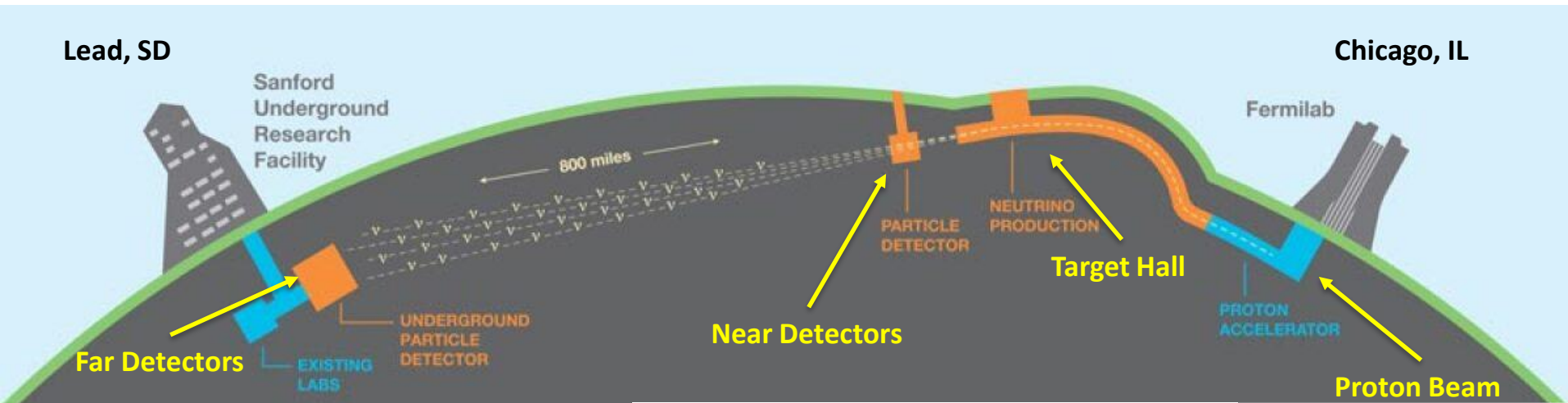
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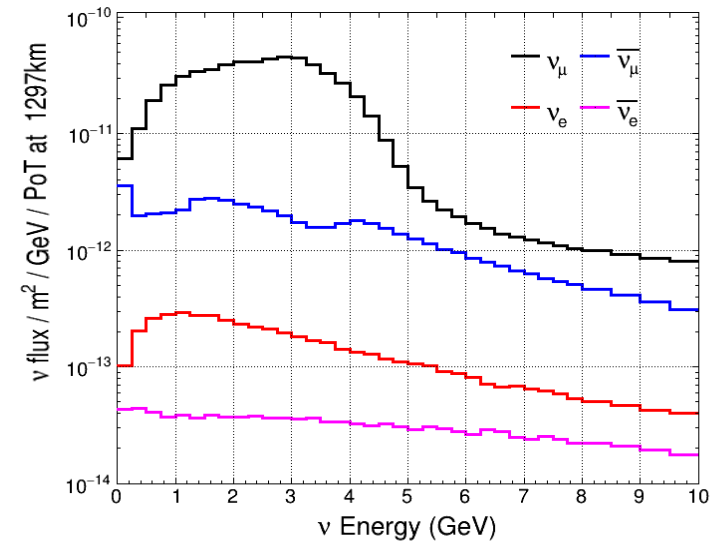
# What is DUNE?

- The **D**eep **U**nderground **N**eutrino **E**xperiment **hosted at** Fermilab
- **Comprised of** 1400+ collaborators **across** 35+ countries.
- **Physics goals include** (but not limited to):
  - **Address Baryon Asymmetry of the Universe (BAU)**
    - ◆ Measure  $\delta_{CP}$  in lepton sector.
  - **Determine the neutrino mass ordering**
    - ◆ Sign of  $|\Delta m_{32}^2|$  ?
  - **Determine the octant of  $\theta_{23}$ .**
    - ◆ Is  $\theta_{23}$  greater or less than  $\frac{\pi}{4}$  ?
  - **Near Detector Complex hosts a suite of rich physics programs .**
    - ◆ Suite of detectors {LAR, GAR, SAND, TMS, **PRISM** ...}
  - **And so much more!**
    - ◆ Interested in joining **DUNE**? Get started [here](#).

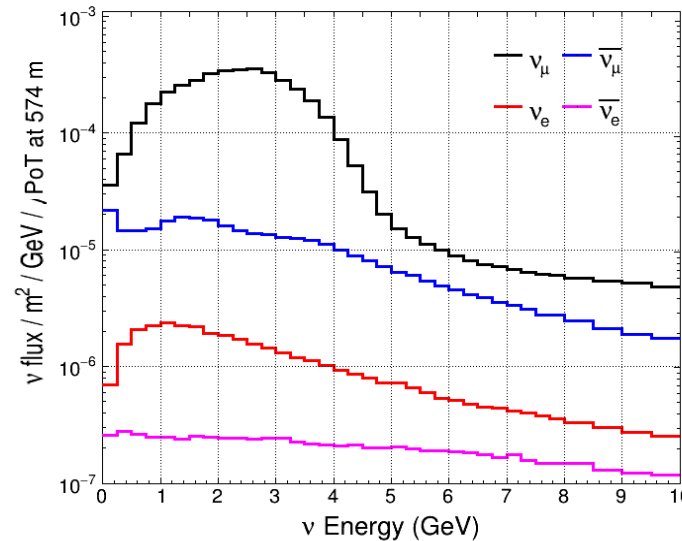
# How does DUNE work?



$\nu$  Flux Prediction at Far Det.

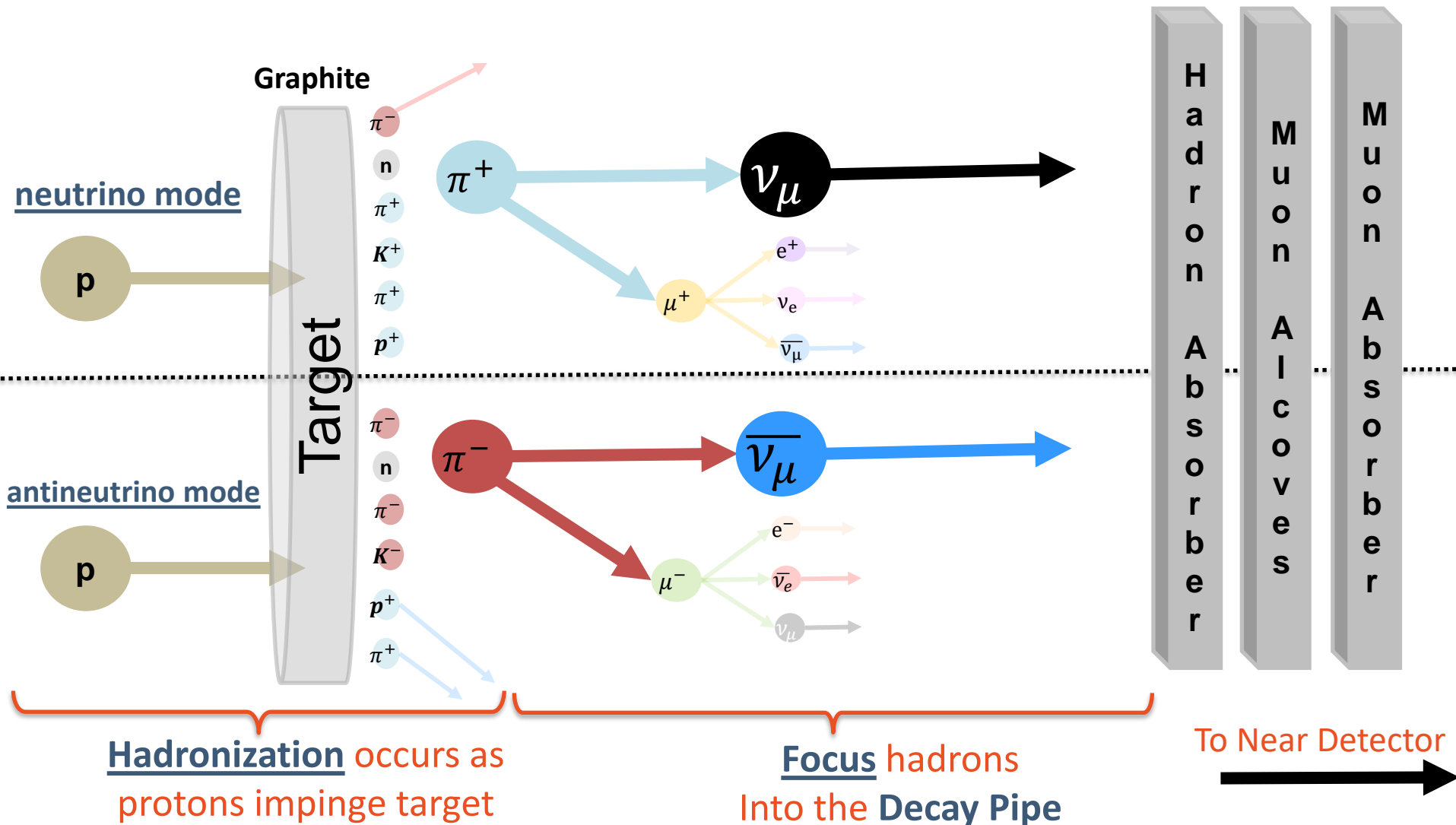


$\nu$  Flux Prediction at Near Det.



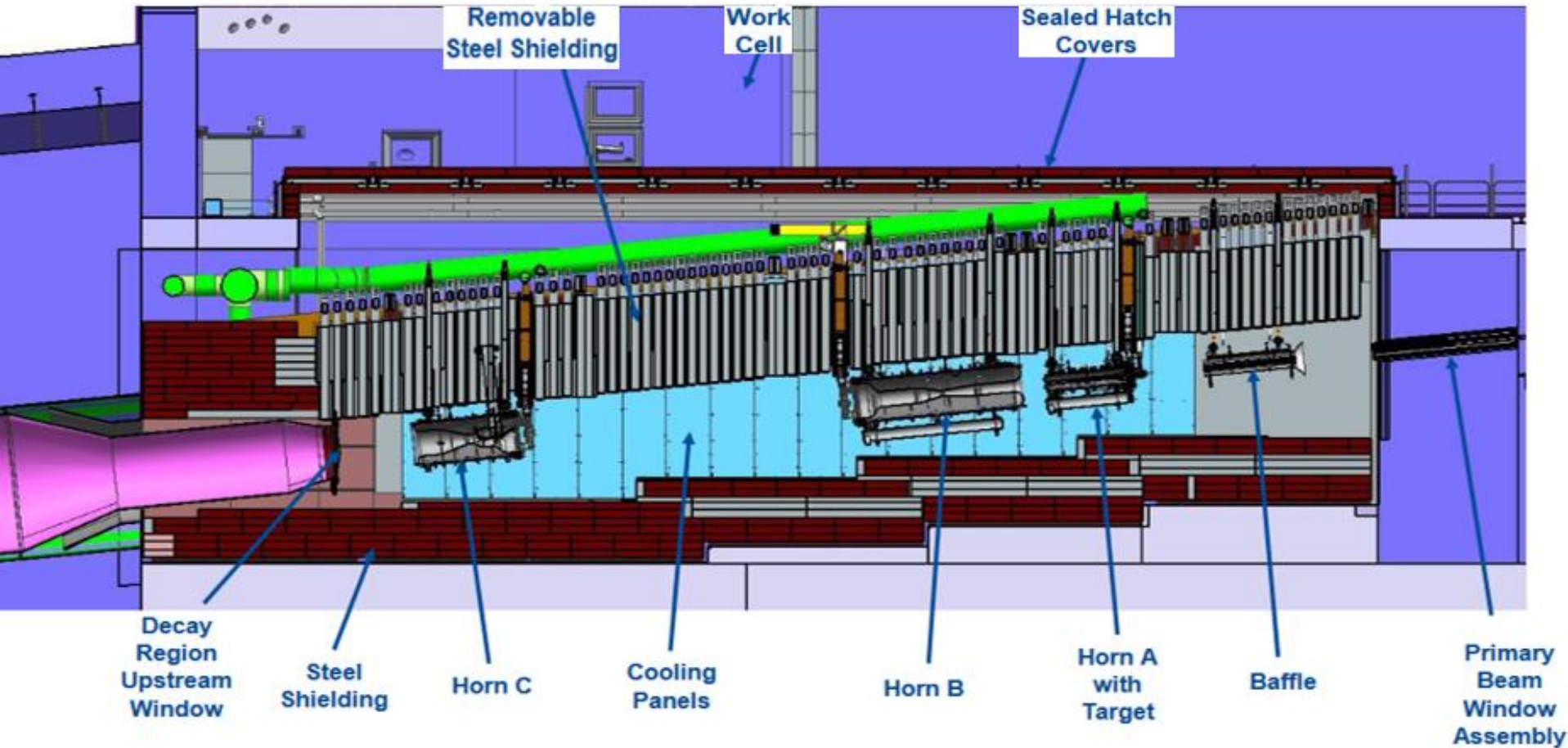
- Protons impinge target.
- Secondary beam of pions is created and focused.
- Pions decay to neutrinos and muons.
- Near Det. samples the neutrino beam.
- Neutrinos oscillate and experience matter effect.

# Hadron Production for DUNE



# Focusing Effects

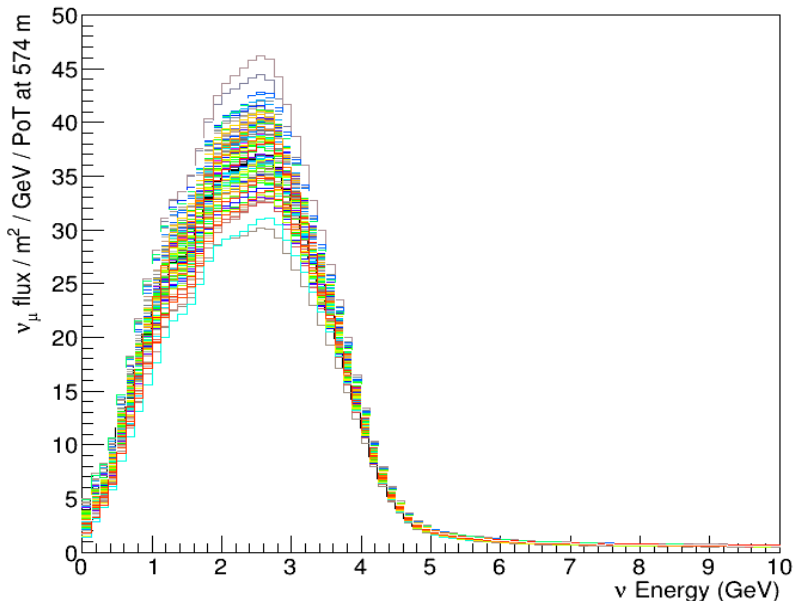
- 3 Horn Focusing System
- 2 modes of operation
  - ◆ Current:  $\pm 300\text{kA}$
- 60+ sources of beam focusing uncertainties
- Target Hall components contribute to Hadronization



# Modeling the DUNE Flux

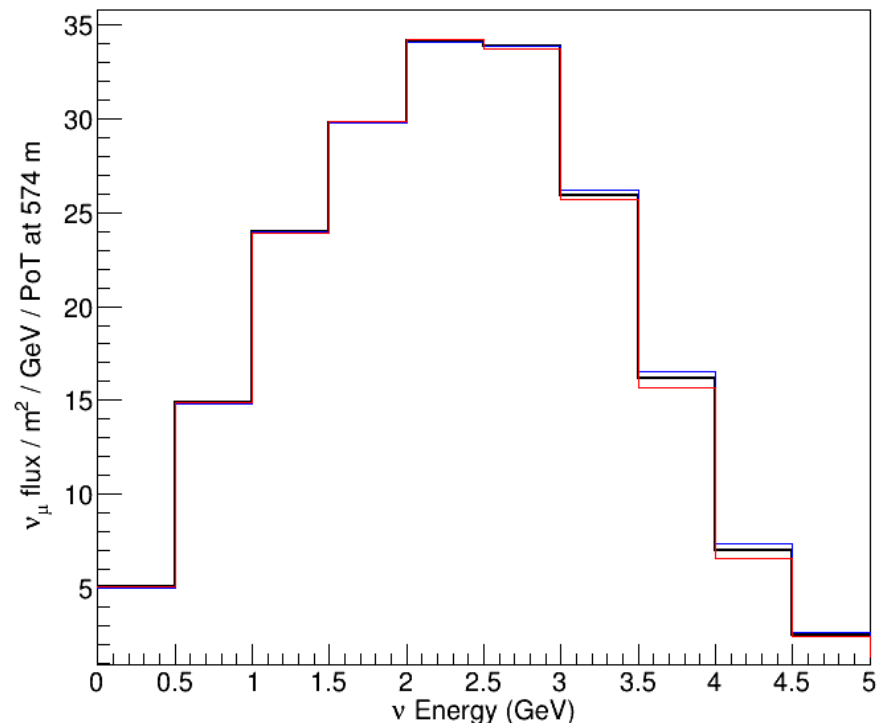
- Nominal flux is input into PPFX.
- Varies the flux parameters across 100 universes.
- Specialized reweighters and external inputs account for Hadron Production processes.

PPFX Multi-Universe FHC  $\nu_\mu$  Flux



- Nominal flux is generated in g4lbne
- Varies nominal by engineering tolerance.
  - Results in 2 universes,  $\pm 1\sigma$ .

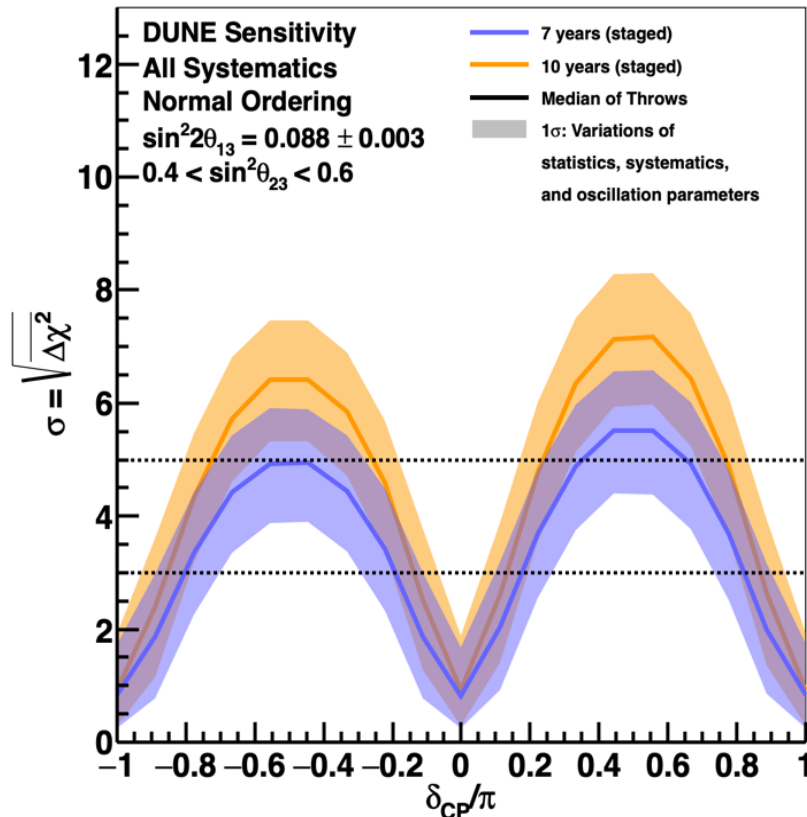
BFU Multi-Universe FHC  $\nu_\mu$  Flux



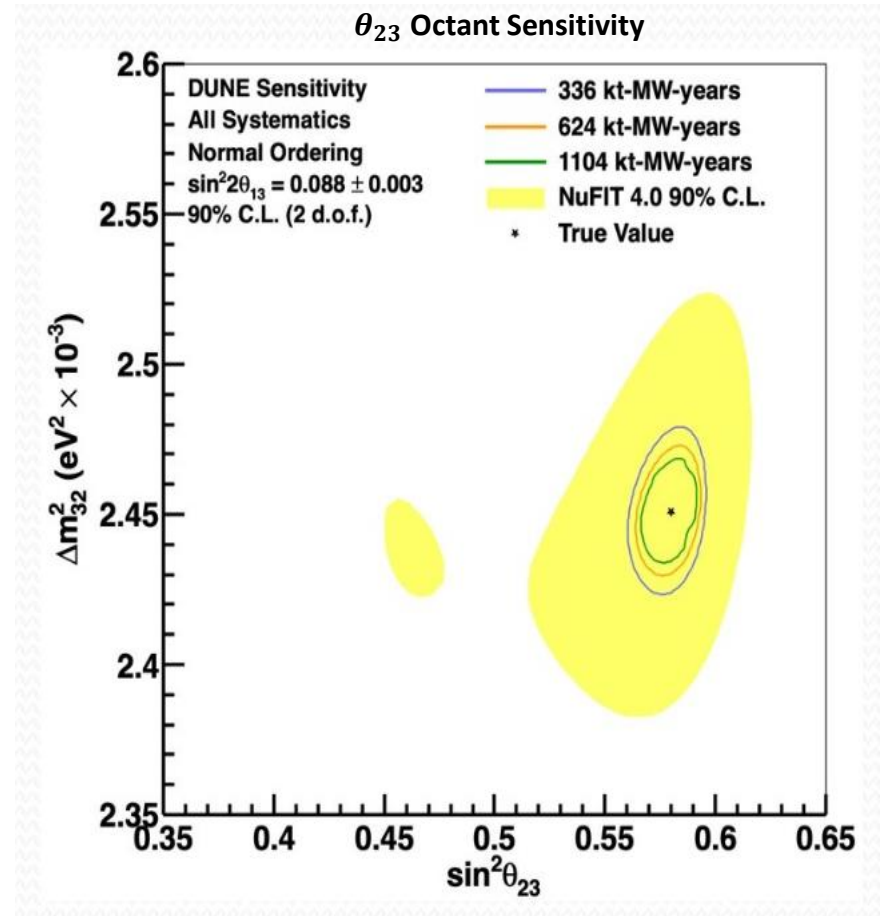
# The Importance of Systematics

- See upcoming publication, “The DUNE Neutrino Flux Simulation” details on covariance.

CP Violation Sensitivity



$\theta_{23}$  Octant Sensitivity

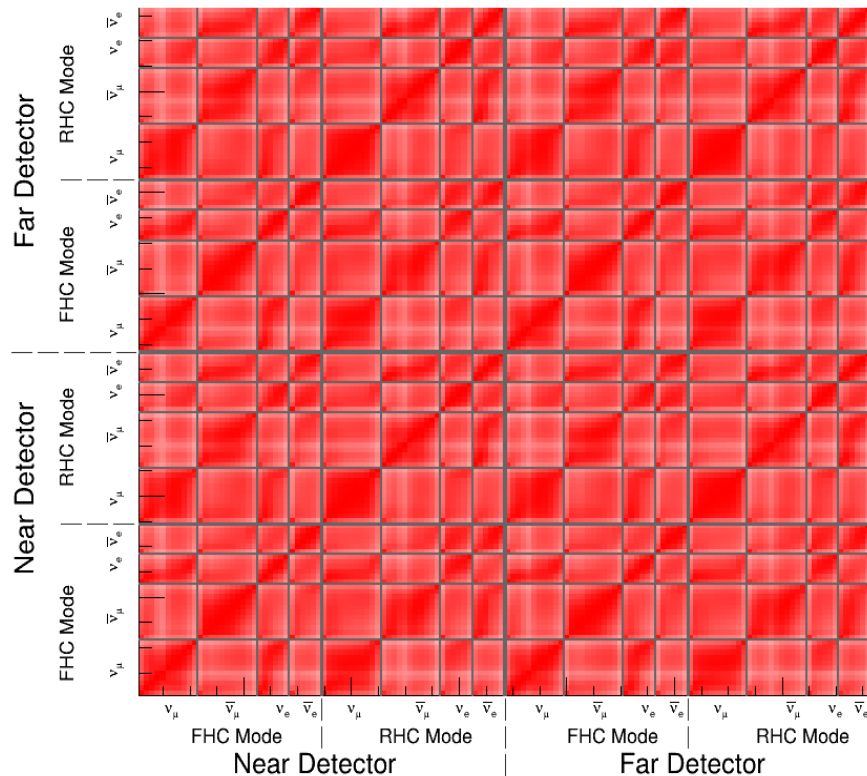




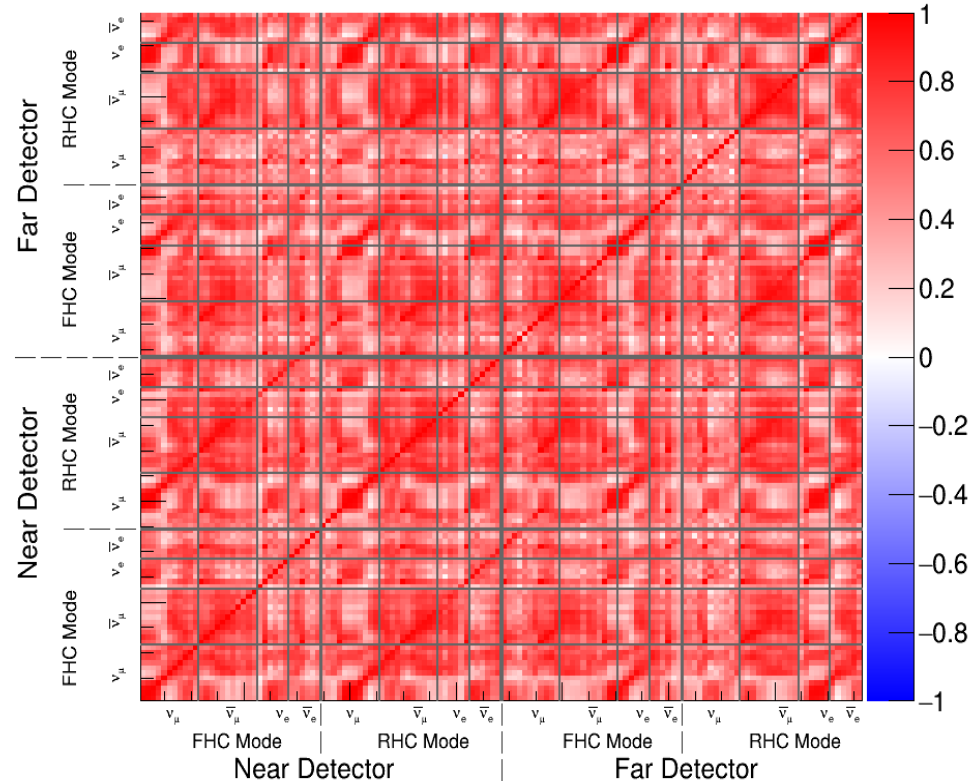
# Determining the Correlations

- The **Correlation Matrices** reveal the magnitude of the relations amongst the various sources of uncertainty across all modes, detector locations and neutrino species.

## Hadron Production Correlations

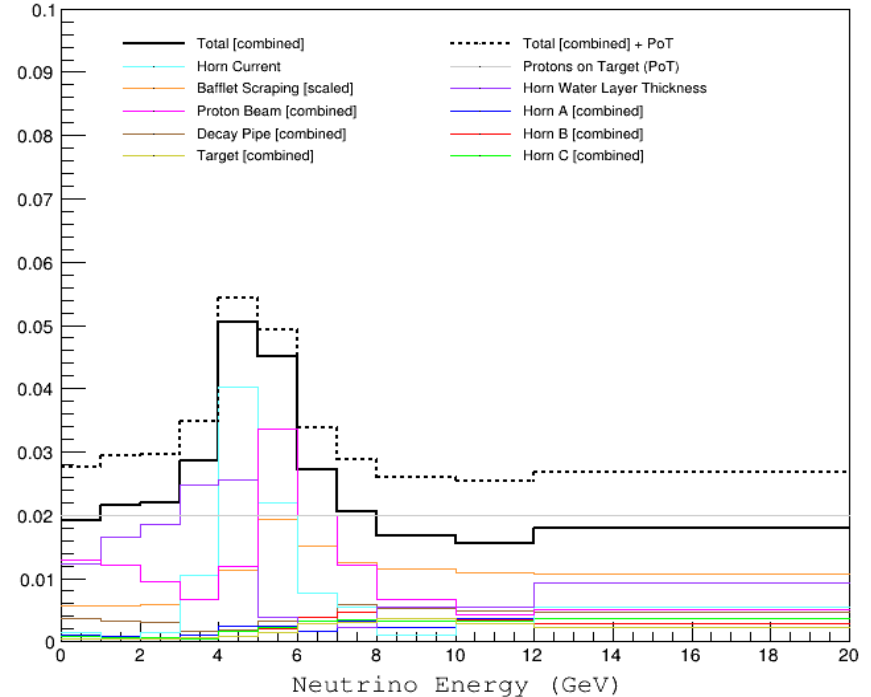
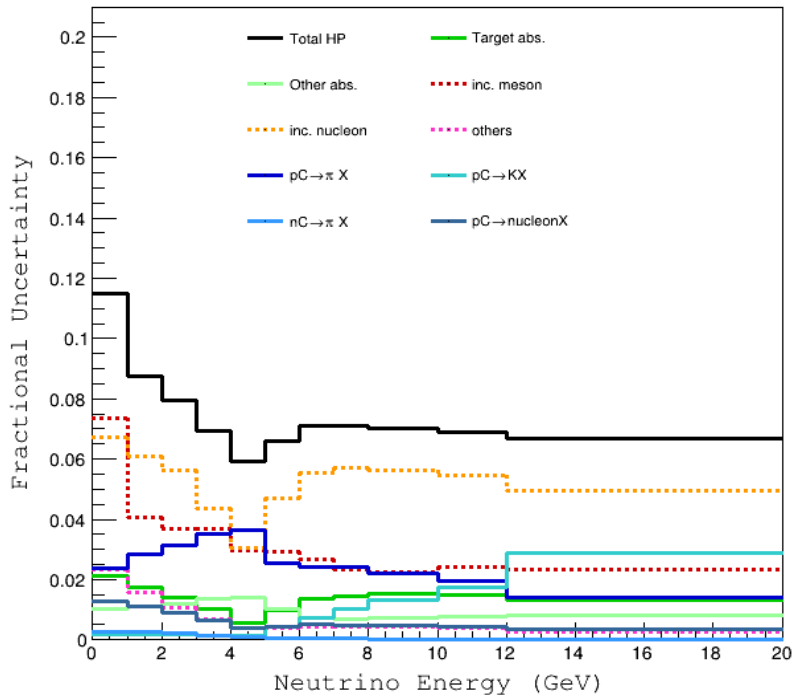


## Beam Focusing Correlations



# Individual Uncertainties

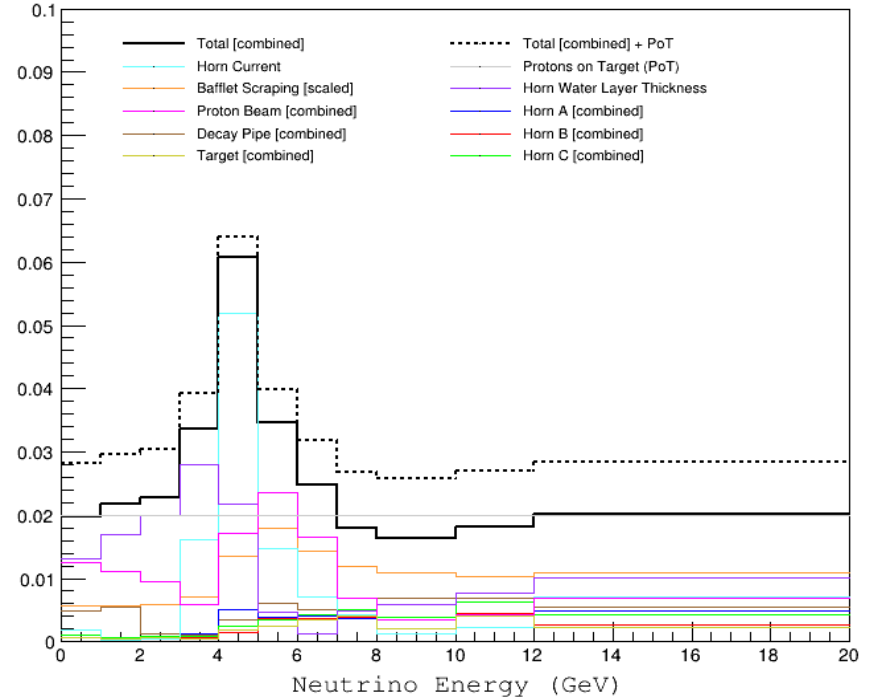
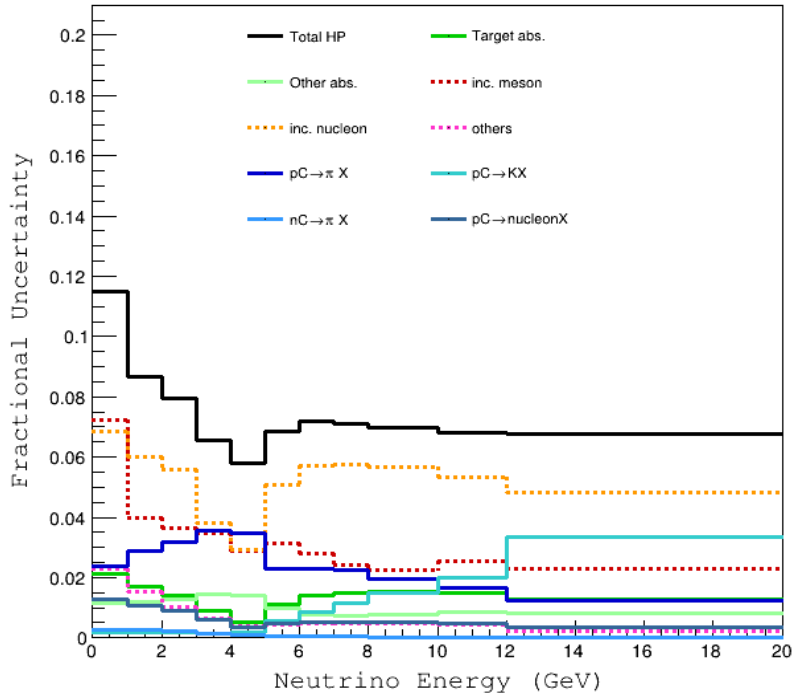
- Taking the square root of the diagonals of each matrix yields the individual uncertainties.



Mode	Location	Species	Mode	Location	Species
neutrino	Far Det.	$\nu_\mu$	neutrino	Far Det.	$\nu_\mu$
<b>Hadron Production</b>			<b>Beam Focusing</b>		

# Individual Uncertainties

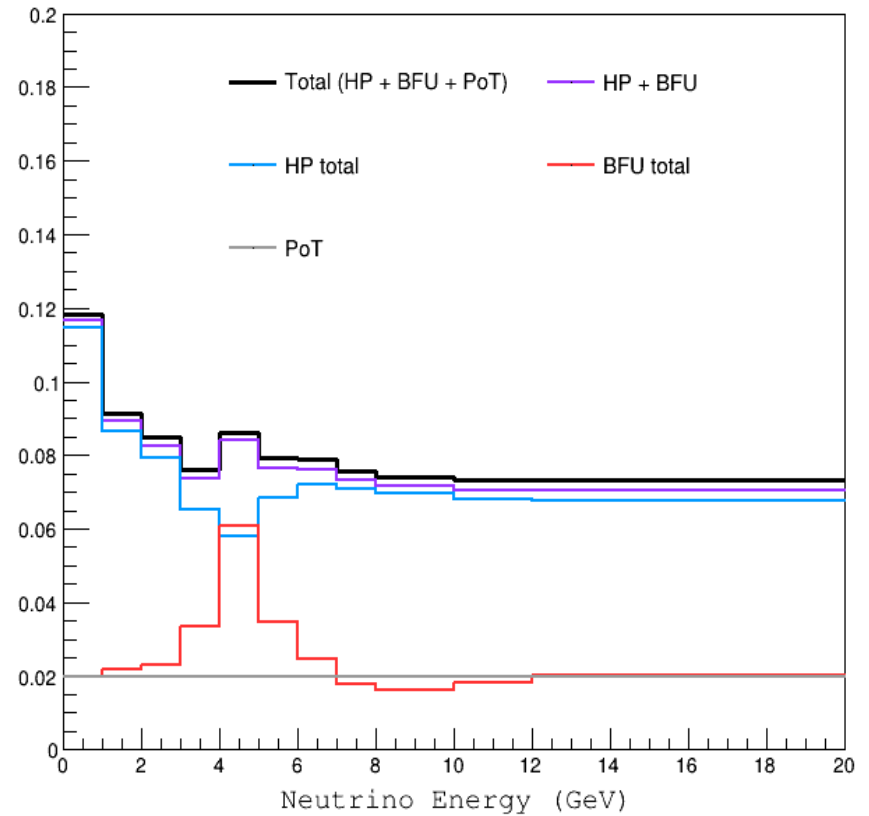
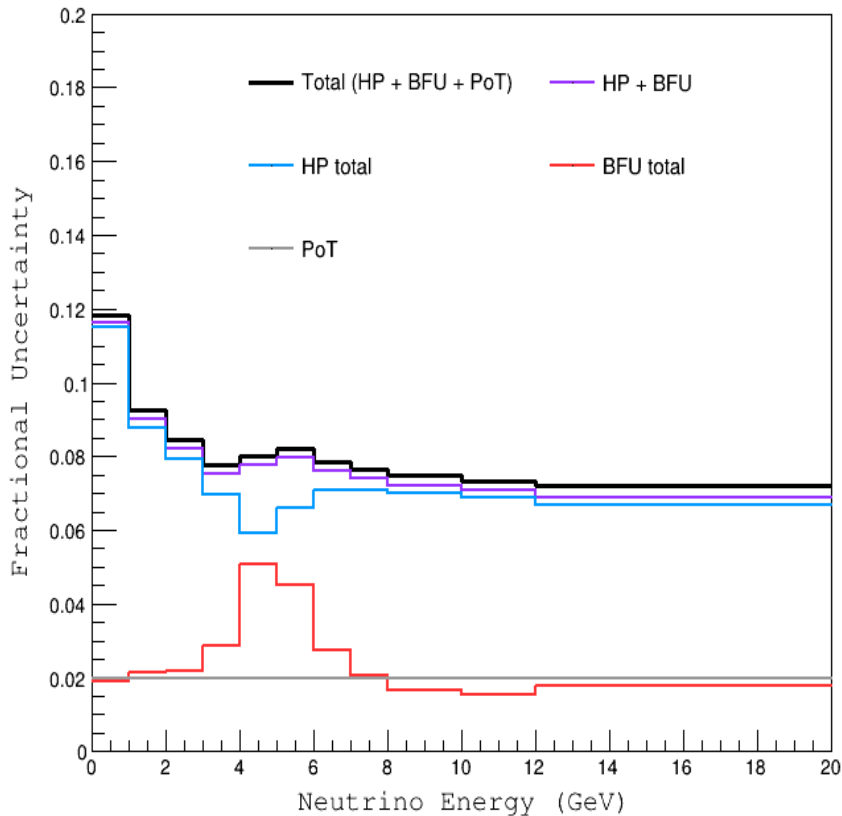
- Taking the square root of the diagonals of each matrix yields the individual uncertainties.



Mode	Location	Species	Mode	Location	Species
neutrino	Near Det.	$\nu_\mu$	neutrino	Near Det.	$\nu_\mu$
<b>Hadron Production</b>			<b>Beam Focusing</b>		

# Total Systematic Uncertainties

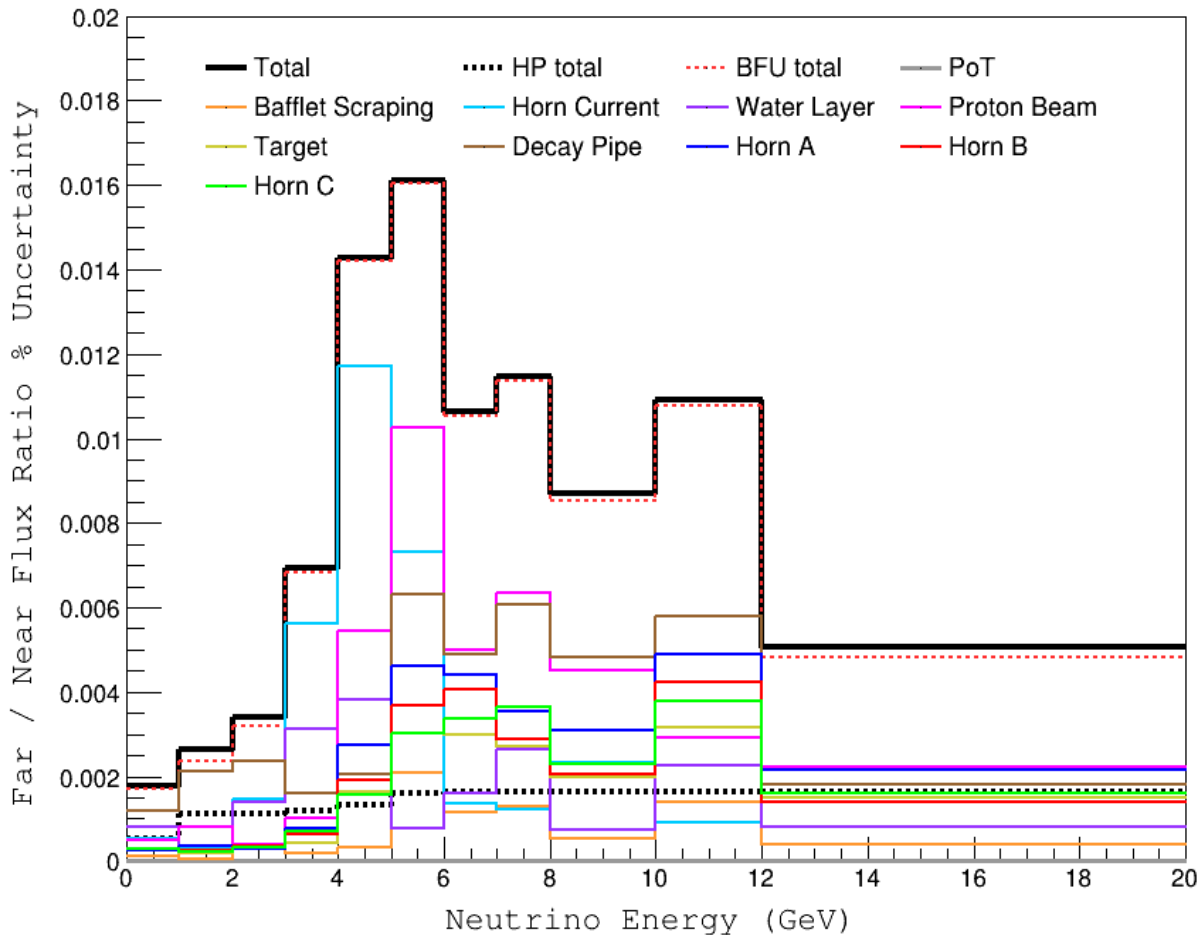
- Adding Hadron Production and Beam Focusing Covariances to obtain Total Beam Covariance.



Mode	Location	Species	Mode	Location	Species
neutrino	Far Det.	$\nu_\mu$	neutrino	Near Det.	$\nu_\mu$

# The Far to Near Flux Ratio

FHC  $\nu_\mu$  Far / Near Ratio % Uncertainties



- **Beam Focusing dominates.**
- **Hadron Production now < 0.2%**
- **Previous max of 6% at 4.2 GeV ~ 1.5%**
- **Flux Peak Unc. at 2.5 GeV ~ 0.4%**

# Conclusions

- **DUNE** is an accelerator-based neutrino experiment hosted at Fermilab
- Among **DUNE's** many goals includes determining:
  - ◆  $\delta_{CP}$
  - ◆ neutrino mass hierarchy
  - ◆ octant of  $\theta_{23}$
- To achieve the high sensitivity required to measure parameters requires covariance matrices for all Systematic Uncertainties
- The covariance matrix encapsulates the all information regarding uncertainties and correlations.
- Hadron Production and Beam Focusing are the largest contributors to beam systematics uncertainties.

# Back Up Slides

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(2:30 - 2:45) pm

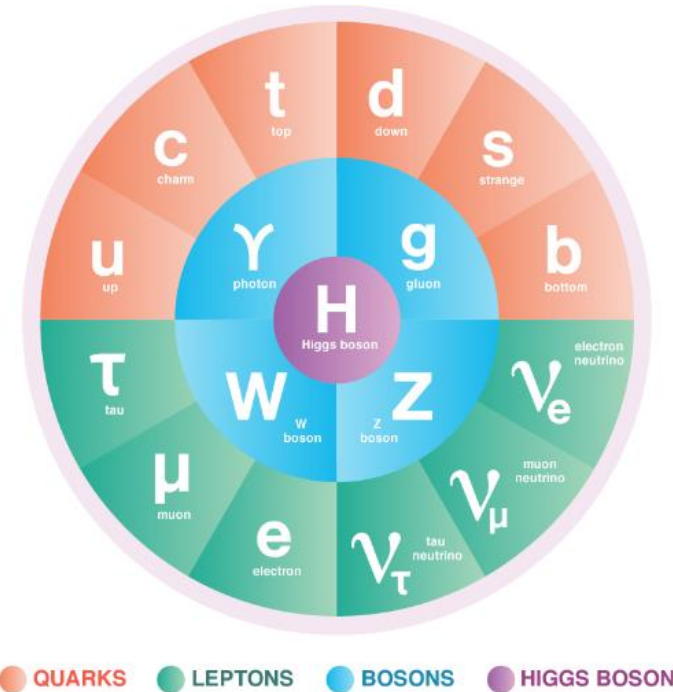
David Lawrence Hall, 107

University of Pittsburgh



# What are neutrinos?

- Fundamental particles of the SM.
- Colorless, neutral leptons
- 3 distinct flavors:  $\nu_e, \nu_\mu, \nu_\tau$
- 3 distinct masses:  $\nu_1, \nu_2, \nu_3$
- Can oscillate between flavors, governed by the PMNS matrix.

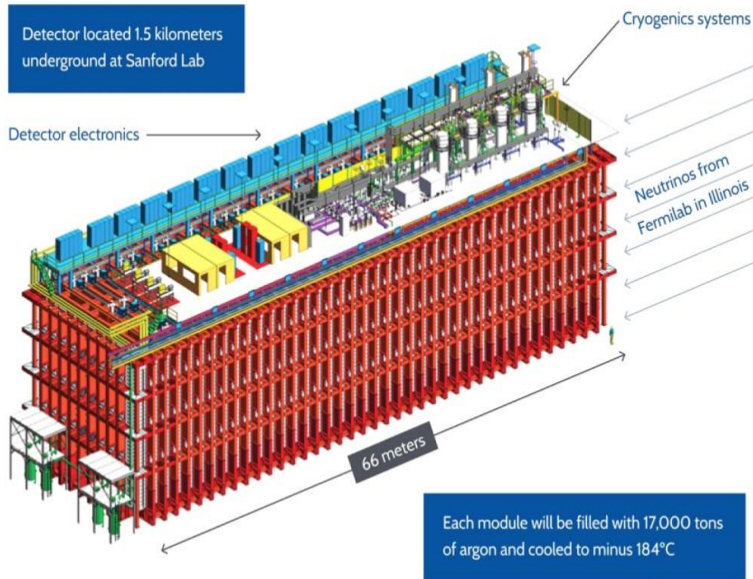
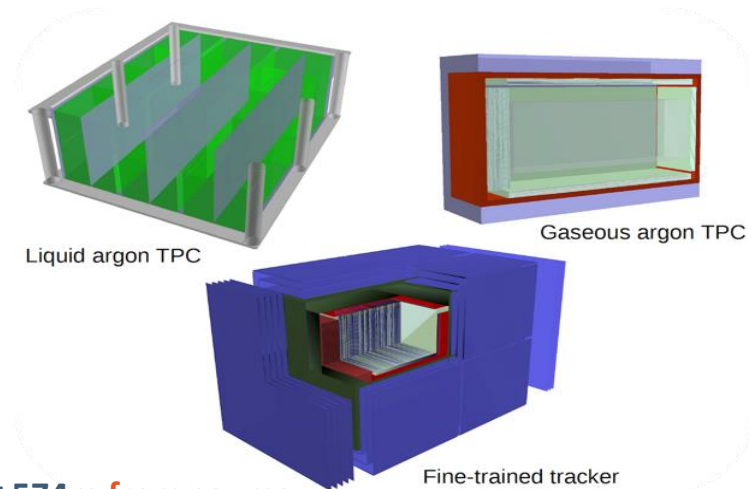


$$|U| = \begin{bmatrix} |U_{e1}| & |U_{e2}| & |U_{e3}| \\ |U_{\mu 1}| & |U_{\mu 2}| & |U_{\mu 3}| \\ |U_{\tau 1}| & |U_{\tau 2}| & |U_{\tau 3}| \end{bmatrix} = \begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta_{CP}} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta_{CP}} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta_{CP}} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta_{CP}} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta_{CP}} & c_{23}c_{13} \end{bmatrix}$$



# DUNE Detection Systems

- 2 Detector Design (Near and Far Detectors)
- Far Det. ~ 1297km from source located at SURF
- Comprised of 4, 17kt LAr TPC's
- Cooled to ~ 90 Kelvin
- Fiducial Volume of 10kt per TPC
- Primary Goal: Measure Oscillated Flux



- Near Det. ~ 574m from source
- Comprised of 3 subdetectors:
  - LAr TPC (50-ton)
  - GAr TPC (1-ton)
  - SAND (8-ton plastic scintillator)
- LAr TPC measures unoscillated flux
- GAr TPC monitors the muon flux
- SAND (Fine-trained tracker) measures:
  - On-Axis beam flux, possible Neutron detection.
- Argon TPC's can move off axis (PRISM)
  - Deconvolves flux from  $\nu$  cross-section on Argon.
- Primary Goals:
  - Measure Unoscillated Flux /
  - Characterize the Beam

# What is Hadron Production?

- Largest source of systematic uncertainty for DUNE flux prediction.
- Sources of Hadron Production in DUNE include:
  - Protons impinging on Graphite target:
$$p + {}^{12}\text{C} \rightarrow \pi^{\pm} + X$$
$$p + {}^{12}\text{C} \rightarrow K^{\pm} + X$$
$$p + {}^{12}\text{C} \rightarrow p(n) + X$$
  - Secondary Interactions of neutrons:  $n + {}^{12}\text{C} \rightarrow \pi^{\pm} + X$
  - Hadron Absorption both inside and outside the target.
  - Secondary meson and nucleon interactions
  - And many others!
- Simulating these Hadron Production uncertainties requires:
  - Input data from dedicated experiments [NA49, SHINE, NA61\*]
  - Package to Predict the Flux (PPFX), developed originally for Minerva by Leonidas Aliaga Soplín of U. Houston.

# Beam Focusing Effects

- 2<sup>nd</sup> largest source of systematic uncertainty in DUNE flux prediction.
- Over 60 sources, all arising from engineering tolerances, such as:
  - Horn Current ( $\pm 300\text{kA}$ )
  - Thickness of Water Layer cooling Horns.
  - Scraping of proton beam against the Baffle.
  - Various characteristics of:
    - ◆ Proton Beam characteristics (Radius, Position, Angle, ...)
    - ◆ Target characteristics (Density, Position, Length, ...)
    - ◆ Horns A,B,C characteristics (Position, Ellipticity, Tilt, ...)
    - ◆ Decay Pipe characteristics (Radius, Position, Cross-Section, ...)

# Calculating BFU Covariance

- Calculate individual covariances for each source of uncertainty (i) in both universes.

$$\text{Cov}_{\text{BFU},+}^{(i)}(x_j, x_k) = \frac{(x_j^{(i)} - \bar{x}_j)(x_k^{(i)} - \bar{x}_k)}{\bar{x}_j \bar{x}_k}$$

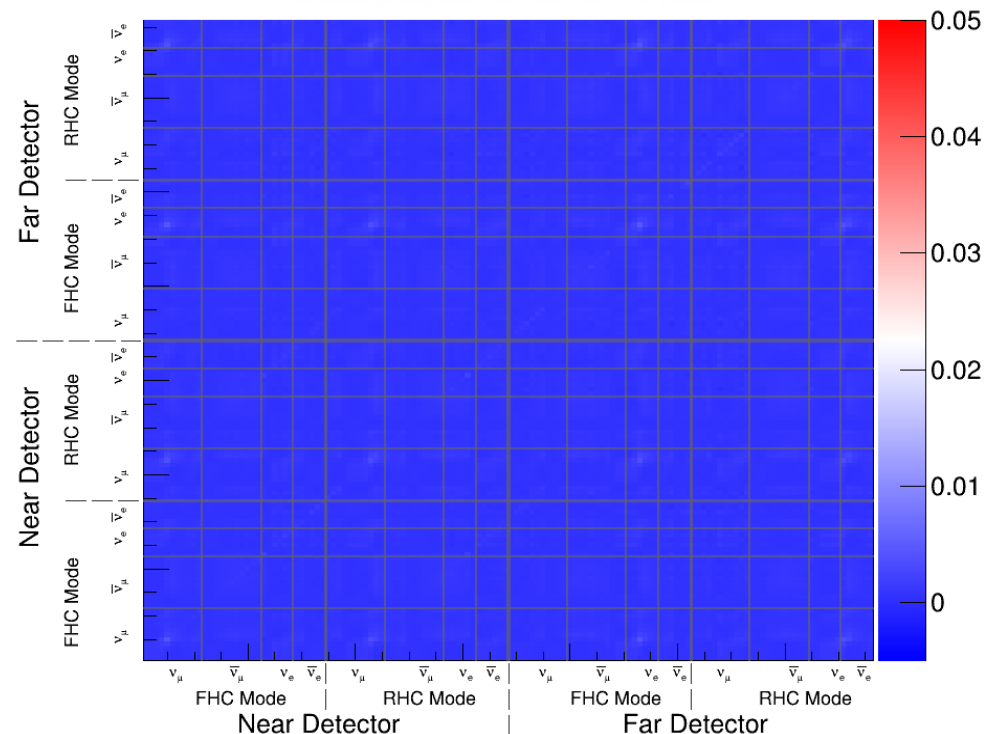
$$\text{Cov}_{\text{BFU},-}^{(i)}(y_j, y_k) = \frac{(y_j^{(i)} - \bar{y}_j)(y_k^{(i)} - \bar{y}_k)}{\bar{y}_j \bar{y}_k}$$

- Total BFU Covariance is average of universe covariances.

$$\left\langle \text{Cov}_{\text{BFU}}^{(i)}(x_j, x_k, y_j, y_k) \right\rangle = \frac{1}{2} \left[ \text{Cov}_{\text{BFU},+}^{(i)}(x_j, x_k) + \text{Cov}_{\text{BFU},-}^{(i)}(y_j, y_k) \right]$$

- Here we see the BFU Covariance is quite small indicating the magnitudes of the focusing uncertainties are likewise, small.

Total Beam Focusing Covariance



# Calculating HP Covariance

- Calculate individual covariances for each source of uncertainty (i) in both universes.

$$\left\langle \text{COV}_{\text{HP}}^{(i)}(z_j, z_k) \right\rangle = \frac{1}{N} \sum_{u=1}^N \text{COV}_{\text{HP}}^{(i,u)}(z_j, z_k) \quad \therefore \begin{cases} N = 100 & 100 \text{ universes} \\ u = [1, N] & \text{universe \#} \end{cases}$$

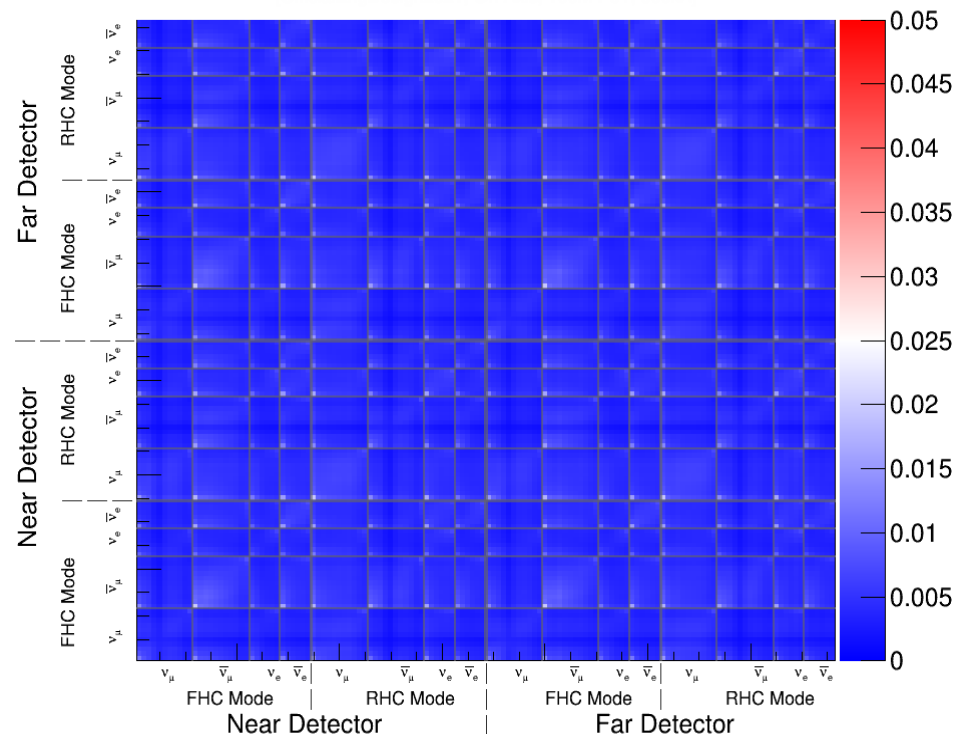
- Total BFU Covariance is average of universe covariances.

$$\text{COV}_{\text{HP}}^{(\text{total})}(z_j, z_k) = \sum_{i=0}^N \left\langle \text{COV}_{\text{HP}}^{(i)}(z_j, z_k) \right\rangle$$

$$\therefore \begin{cases} N = 9 \end{cases}$$

- Here we see the HP Covariance is likewise small indicating the magnitudes of the Hadron Production uncertainties are also, small.

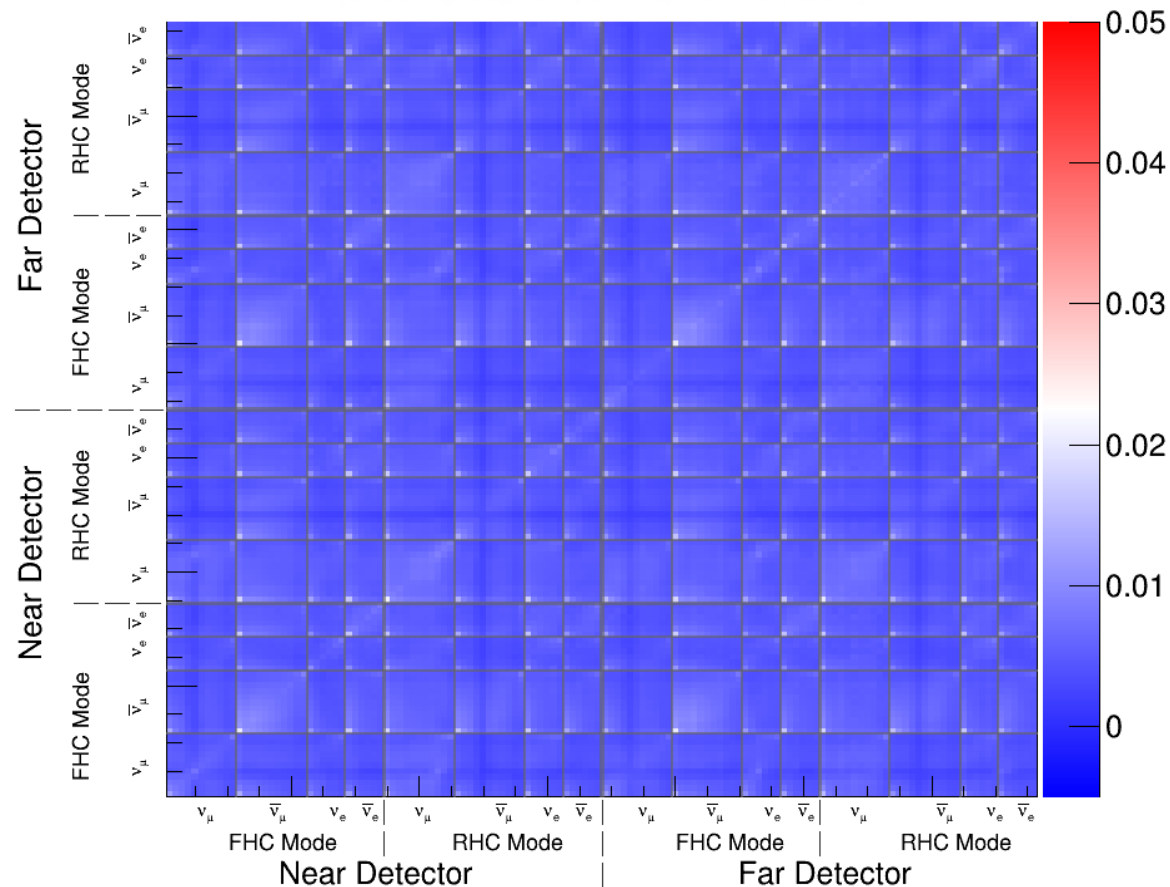
HP Covariance Matrix



# The Total Covariance

- Sum of the Hadron Production and Beam Focusing Covariance matrices.

Total DUNE Flux Covariance



# Total Flux Correlation

- The Total Correlation Matrix reveals the magnitude of the relations amongst the various sources of uncertainty across all modes, detector locations and neutrino species.

Total DUNE Flux Prediction Correlations

