Proper Treatment of Flux Uncertainties in Neutrino Cross Section Measurement

London Cooper-Troendle lcoopert@proton.me May 14th 2024

Issue first raised to the neutrino cross section community in [PhysRevD.102.113012](https://journals.aps.org/prd/abstract/10.1103/PhysRevD.102.113012)

DPF-PHENO 2024

Challenges in Neutrino Interaction Modeling

- Wide range of energies
	- Spans QE, RES, DIS
- Range of nuclear targets across experiments ○ Hydrogen, Deuterium, Carbon, Argon, Iron, Lead
- Complex QCD physics inside nucleus
	- Nuclear initial state
	- Nucleon-nucleon correlations
	- Final state interactions

Credit: T. Golan

Flux-Averaged Cross Sections

- Accelerator neutrino experiments do not directly observe the incoming neutrino
	- Reconstructing Enu introduces model dependence that we want to avoid
- Cross-section measurements are flux-averaged over the beam flux they are exposed to
	- Wide energy-range beams means cross section varies significantly across measured phase space

How to Measure Flux-Averaged Cross Sections

- Directly measure N events
	- Subtract background **B**
	- \circ Correct for efficiency ϵ and smearing D
	- Scale by number of nuclei T
	- \circ Scale by total flux prediction Φ
- Flux uncertainties present in Φ , B, ϵ
	- \circ B can vary with E_{n}
	- \circ ecan vary with E_{v} within each bin
- Potential for low model dependence

$$
S \triangleq \frac{\int F(E_v) \cdot \sigma(E_v) \cdot dE_v}{\Phi}
$$

N_i = B_i + T $\cdot \Sigma_j \int_j F(E_v) \cdot \sigma(E_v) \cdot D_{ij} \cdot \epsilon_{ij} \cdot dE_v$

Signal S Measured event count N Estimated background B Estimated efficiency ϵ Detector smearing D True neutrino flux distribution F

Estimated flux distribution F

Number of target nuclei T Total estimated flux Φ Neutrino energy E Cross section σ Reco bin i Truth bin j

How to Measure Flux-Averaged Cross Sections

- Directly measure N events
	- \circ Subtract background B
	- \circ Correct for efficiency ϵ and smearing D
	- Scale by number of nuclei T
	- \circ Scale by total flux prediction Φ
- Potential for low model dependence
- Flux uncertainties present in Φ , B, ϵ
	- \circ B can vary with E_{α}
	- \circ ecan vary with E_{v} within each bin

$$
S \stackrel{\text{def}}{=} \frac{\int F(E_v) \cdot \sigma(E_v) \cdot dE_v}{\Phi}
$$
\n
$$
N_i = B_i + T \cdot \Sigma_j \int_j F(E_v) \cdot \sigma(E_v) \cdot D_{ij} \cdot \epsilon_{ij} \cdot dE_v
$$
\n
$$
\frac{\sum_i (N_i - B_i) \cdot (\epsilon \cdot D)^{-1} \cdot j}{T \cdot \Phi} = \frac{\int_j F(E_v) \cdot \sigma(E_v) \cdot dE_v}{\Phi}
$$
\n
$$
= S_j
$$

Signal S Measured event count N Estimated background B Estimated efficiency ϵ Detector smearing D True neutrino flux distribution F

Estimated flux distribution F

Number of target nuclei T Total estimated flux Φ Neutrino energy E Cross section σ Reco bin i Truth bin j

Comparing Measurements to Predictions

- Prediction uses estimated flux \overline{F}
	- Note: F contains uncertainties as well as central value
- Now both measurement and prediction contain flux uncertainties!
	- Prediction contains full flux uncertainties
	- \circ Measurement has norm from Φ and some shape effects in B, ϵ
- Exact correlation cannot be easily determined
	- No measurement to date provides sufficient info (as far as I know)

- Signal S Measured event count N Estimated background B Estimated efficiency ϵ Detector smearing D
- True neutrino flux distribution F
- Estimated flux distribution F
- Number of target nuclei T Total estimated flux Φ Neutrino energy E Cross section σ Reco bin i Truth bin j

Example 1: Under-Estimate χ^2

- Consider a total cross section measurement in 1 bin
	- Suppose there is a 10% data excess
- \bullet Assume a cross section prediction $\sigma \propto \mathsf{E}_{v}$
- Assume a background prediction B ∞ $1/E$ _v
- Ignoring correlations under-estimates χ^2 :
	- Meas and pred uncertainties can address tension, but require pulling flux in **opposite directions**
	- Correct treatment requires **larger** deviation from nominal flux of the second control of the seco

Example 2: Over-Estimate χ^2

- Consider a total cross section measurement in 1 bin
	- Suppose there is a 10% data excess
- \bullet Assume a cross section prediction $\sigma \propto \mathsf{E}_{v}$
- \bullet Assume a background prediction B ∞ E_v
- Ignoring correlations over-estimates χ^2 :
	- Uncertainties are under-counted when added in quadrature: arise from the same flux deviation
	- Correct treatment requires **smaller** deviation from nominal flux

Solutions to the Flux Treatment Problem

- Provide more info to allow correlations be determined
	- Publish full set of flux universes and extracted cross section for each
	- Theorist could compute predicted cross section for each flux universe, construct joint covariance between meas and pred cross section
- Extremely messy and difficult
	- Asks a lot of work from theorists
	- Perhaps a standardized framework could be written to allow plug-in and compute
- Alternative approach: measure **nominal**-flux-averaged cross section

Include estimated flux entirely in measurement

$$
\widetilde{S} = \frac{\int \overline{F}(E_v) \sigma(E_v) dE_v}{\Phi}
$$

- Measurement contains all flux uncertainties
- Prediction only requires nominal flux estimate
- Much easier to make comparison to theory

Nominal-flux-averaged signal S Nominal-flux-averaged signal $\widetilde{\mathbf{S}}$ Measured event count N Estimated background B Estimated efficiency ϵ Detector smearing D True neutrino flux distribution F Estimated flux distribution F

Include estimated flux entirely in measurement

- Measurement contains all flux uncertainties
- Prediction only requires nominal flux estimate
- Much easier to make comparison to theory

Nominal-flux-averaged signal S Nominal-flux-averaged signal $\widetilde{\mathbf{S}}$ Measured event count N Estimated background B Estimated efficiency ϵ Detector smearing D True neutrino flux distribution F Estimated flux distribution F

$$
\widetilde{S} \stackrel{\text{def}}{=} \frac{\int \overline{F}(E_{\nu}) \cdot \sigma(E_{\nu}) \cdot dE_{\nu}}{\Phi}
$$

N_i - B_i = T· Σ_j $\int_j F(E_{\nu}) \cdot \sigma(E_{\nu}) \cdot D_{ij} \cdot \epsilon_{ij} dE_{\nu}$

Include estimated flux entirely in measurement

- Measurement contains all flux uncertainties
- Prediction only requires nominal flux estimate
- Much easier to make comparison to theory

Nominal-flux-averaged signal S Nominal-flux-averaged signal $\widetilde{\mathbf{S}}$ Measured event count N Estimated background B Estimated efficiency ϵ Detector smearing D True neutrino flux distribution F Estimated flux distribution F

$$
\widetilde{S} \stackrel{\text{def}}{=} \frac{\int \overline{F(E_v) \cdot \sigma(E_v)} \cdot dE_v}{\Phi}
$$
\n
$$
N_i - B_i = T \cdot \sum_j \int_j F(E_v) \cdot \sigma(E_v) \cdot D_{ij} \cdot \varepsilon_{ij} \cdot dE_v
$$
\n
$$
= T \cdot \sum_j \int_j F(E_v) \cdot \sigma(E_v) \cdot D_{ij} \cdot \varepsilon_{ij} \cdot dE_v \cdot T \cdot \Phi \cdot \int_j \overline{F(E_v) \cdot \sigma(E_v)} \cdot dE_v
$$
\n
$$
T \cdot \int_j \overline{F(E_v) \cdot \sigma(E_v)} \cdot dE_v
$$

Include estimated flux entirely in measurement

- Measurement contains all flux uncertainties
- Prediction only requires nominal flux estimate
- Much easier to make comparison to theory

Nominal-flux-averaged signal S Nominal-flux-averaged signal $\widetilde{\mathbf{S}}$ Measured event count N Estimated background B Estimated efficiency ϵ Detector smearing D True neutrino flux distribution F Estimated flux distribution F

$$
\widetilde{S} \stackrel{\text{def}}{=} \frac{\int \overline{F}(E_{v}) \cdot \sigma(E_{v}) dE_{v}}{\Phi}
$$
\n
$$
N_{i} - B_{i} = T \cdot \sum_{j} \int_{j} F(E_{v}) \cdot \sigma(E_{v}) \cdot D_{ij} \cdot \varepsilon_{ij} dE_{v}
$$
\n
$$
= T \cdot \sum_{i} \int_{j} F(E_{v}) \cdot \sigma(E_{v}) \cdot D_{ij} \cdot \varepsilon_{ij} dE_{v} \cdot T \cdot \Phi \cdot \int_{j} \overline{F(E_{v}) \cdot \sigma(E_{v}) dE_{v}}
$$
\n
$$
T \cdot \int_{j} \overline{F}(E_{v}) \cdot \sigma(E_{v}) dE_{v}
$$
\n
$$
\frac{\Phi}{\widetilde{S}_{j}}
$$

Include estimated flux entirely in measurement

- Measurement contains all flux uncertainties
- Prediction only requires nominal flux estimate
- Much easier to make comparison to theory

Nominal-flux-averaged signal S Nominal-flux-averaged signal $\widetilde{\mathbf{S}}$ Measured event count N Estimated background B Estimated efficiency ϵ Detector smearing D True neutrino flux distribution F Estimated flux distribution F

$$
\widetilde{S} \stackrel{\text{def}}{=} \frac{\int \overline{F(E_{v}) \cdot \sigma(E_{v})} \, dE_{v}}{\Phi}
$$
\n
$$
N_{i} - B_{i} = T \cdot \sum_{j} \int_{j} F(E_{v}) \cdot \sigma(E_{v}) \cdot D_{ij} \cdot \varepsilon_{ij} \, dE_{v}
$$
\n
$$
= T \cdot \sum_{i} \int_{j} F(E_{v}) \cdot \sigma(E_{v}) \cdot D_{ij} \cdot \varepsilon_{ij} \, dE_{v} \cdot T \cdot \Phi \cdot \int_{j} \overline{F(E_{v}) \cdot \sigma(E_{v})} \cdot dE_{v}
$$
\n
$$
T \cdot \int_{j} \overline{F(E_{v}) \cdot \sigma(E_{v})} \cdot dE_{v} \qquad \qquad \underbrace{\rho}_{\overline{F}_{j}} \cdot \frac{\Phi}{S_{j}}
$$

Total estimated flux Φ Neutrino energy E Cross section σ

> Flux constant F Reco bin i Truth bin j

Monte-Carlo smearing matrix M

Include estimated flux entirely in measurement

- Measurement contains all flux uncertainties
- Prediction only requires nominal flux estimate
- Much easier to make comparison to theory

Nominal-flux-averaged signal S Nominal-flux-averaged signal $\widetilde{\mathbf{S}}$ Measured event count N Estimated background B Estimated efficiency ϵ Detector smearing D True neutrino flux distribution F Estimated flux distribution F

Include estimated flux entirely in measurement

- Measurement contains all flux uncertainties
- Prediction only requires nominal flux estimate
- Much easier to make comparison to theory

Nominal-flux-averaged signal S Nominal-flux-averaged signal $\widetilde{\mathbf{S}}$ Measured event count N Estimated background B Estimated efficiency ϵ Detector smearing D True neutrino flux distribution F Estimated flux distribution F

Summary

- Cross section measurements are vital to improving our neutrino interaction modeling
	- We need to be able to accurately compare measurements to predictions
- Industry standard real-flux-averaged cross section contains complicated correlations between meas and pred
	- Existing measurements contain insufficient info for accurate comparison
	- In theory info release is possible flux universes, each cross section extracted, let theorists construct covariance across joint distribution
	- However, this is messy and asks a lot of work on theorists
- Nominal-flux-averaged cross section allows for direct comparison to prediction