# An Improved Search for Unstable Sterile Neutrinos at IceCube

#### Philip Weigel, MIT APS DPF-PHENO 2024 May 15th, 2024





#### **Neutrino Oscillation Anomalies**

• Several unresolved anomalies in various neutrino experiments



**MiniBooNE** 

Events/MeV 5.2 Phys. Rev. Lett. 121, 221801

Data (stat err.)

 $v_e$  from  $\mu^{+/-}$ 

 $v_e$  from K<sup>+/-</sup>  $v_e$  from K<sup>0</sup>  $\pi^0$  misid

 $\Delta \rightarrow N\gamma$ dirt

# 3+1 Sterile Neutrino Model

- Anomalies in short baseline neutrino experiments could be explained by a sterile neutrino
  - Introduce a new flavor and mass state, append a row and column to the PMNS matrix
  - $\circ$  These states do not interact weakly  $\rightarrow$  only accessible through oscillations

$$m_{4}^{2} \xrightarrow{m_{4}^{2}} \Delta m_{4}^{2}$$

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$$m_{3}^{2} \xrightarrow{m_{2}^{2}} \xrightarrow{m_{1}^{2}} v_{3}$$

$$v_{2} \xrightarrow{v_{2}} v_{1}$$

$$v_{2} \xrightarrow{v_{1}} v_{3}$$

$$\begin{pmatrix} \nu_{e} \\ \nu_{\mu} \\ \nu_{\tau} \\ \nu_{s} \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix} \begin{pmatrix} \nu_{1} \\ \nu_{2} \\ \nu_{3} \\ \nu_{4} \end{pmatrix} \quad \frac{|U_{e4}|^{2} = \sin^{2}(\theta_{14})}{|U_{\mu 4}|^{2} = \sin^{2}(\theta_{24})\cos^{2}(\theta_{14})} \\ |U_{\tau 4}|^{2} = \sin^{2}(\theta_{34})\cos^{2}(\theta_{24})\cos^{2}(\theta_{34})$$

#### Atmospheric Neutrinos and IceCube

The conventional atmospheric muon neutrino flux originates from  $\pi$  and K decay-in-flight



IceCube Lab

# **Sterile Neutrino Oscillations in Matter**

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  - The zenith angle is a proxy for the baseline L, the distance that the neutrinos have traveled
- There are two main features in these oscillograms:



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- There are two main features in these oscillograms:
  - Vacuum (L/E) oscillations
  - A matter-enhanced resonant disappearance for antineutrinos that travel through most of the Earth

$$E_{res} = \mp \frac{\Delta m^2 \cos(2\theta)}{\sqrt{2}G_F N_n}$$



#### Improvements to Event Selection and Analysis

- To search for sterile neutrinos, we leverage the large flux of atmospheric muon (anti)neutrinos observed in IceCube
- Major improvements since the previous search:
  - BDT for removing atmospheric muon backgrounds
  - DNN-based energy reconstruction and classifier
  - More detailed ice systematic treatment
  - Improved atmospheric flux systematics
  - Broken power law for astrophysical flux





Philip Weigel / DPF-PHENO 2024 / 05-15-2024

# Latest 3+1 Sterile Neutrino Results

• Best Fit:

$$\Delta m_{41}^2 = 3.5 \text{ eV}^2$$
$$\sin^2(2\theta_{24}) = 0.16$$
p-value =  $3.1\% \longrightarrow 2.2\sigma$ 

- Consistent with the previous sterile analyses in IceCube
- These results appeared on the arXiv today! See: <u>arXiv:2405.08070</u>



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# Invisible Sterile Neutrino Decay



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  - Previous best fit is excluded, but the LLH space was quite shallow so there is overlap
- Regions from the global fits to SBL experiments are excluded
  - Partially excluded at 99% CL
  - Mostly excluded at 95% CL



#### Systematic pulls



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#### Conclusion

- IceCube has performed an improved search for unstable eV-scale sterile neutrinos using 10.7 years of muon neutrino data
  - Probed the invisible decay case, which is preferred over the ordinary 3+1 in the global fits
- No preference for decay was found, best fit point at  $g^2 = 0$ 
  - Results are consistent with previous analysis from 2021, but the previous best fit point is now excluded at >99% CL
  - This result puts constraints on the preferred region of the SBL experiments

#### Thank you for listening!