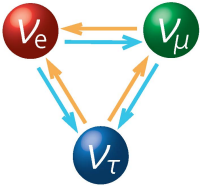
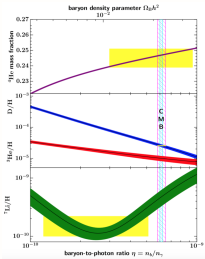


Right-handed neutrinos

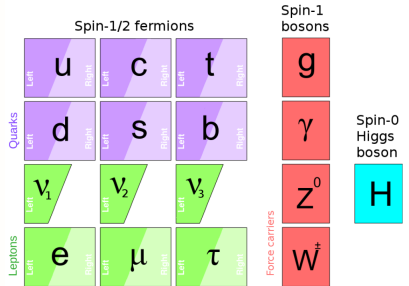


Neutrino oscillations/masses

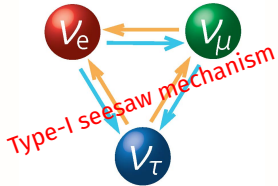


[Particle Data Group]

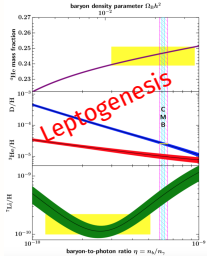
Baryon asymmetry



Right-handed neutrinos



Neutrino oscillations/masses



[Particle Data Group]

Baryon asymmetry

Spin-1/2 fermions						Spin-1 bosons			
Quarks	Left	Right	Left	Right	Left	Right	Force carriers	Spin-0 Higgs boson	
	u	c	t	g					
	d	s	b	γ					
	ν_1	N_1	ν_2	N_2	ν_3	N_3			Z^0
	e	μ	τ	W^\pm					
	H								

Type-I seesaw mechanism

Type-I seesaw Lagrangian

$$\mathcal{L} \supset Y_{\alpha i} (\bar{\ell}_\alpha \tilde{\phi}) \nu_{Ri} + \frac{1}{2} \bar{\nu}_{Ri}^c (M_M)_{ij} \nu_{Rj} + \text{h.c.}$$

Yukawa

Majorana

Seesaw relation

$$m_\nu = -v^2 (Y \cdot M_M^{-1} \cdot Y^t)$$

$$\nu \simeq U_\nu^\dagger (\nu_L - \theta \nu_R^c) + \text{h.c.}$$

Light neutrinos

$$N \simeq U_N^\dagger (\nu_R + \theta^t \nu_L^c) + \text{h.c.}$$

Heavy neutrinos (HNL)

- $n \geq 2$ HNL generations needed to explain light neutrino masses



Type-I seesaw mechanism

Type-I seesaw Lagrangian (below EWSB)

$$\mathcal{L} \supset \underbrace{v Y_{\alpha i}}_{\text{Dirac}} \bar{\ell}_{\alpha} \nu_{Ri} + \frac{1}{2} \bar{\nu}_{Ri}^c \underbrace{(M_M)_{ij}}_{\text{Majorana}} \nu_{Rj} + \text{h.c.}$$

Dirac

Majorana

Seesaw relation

$$m_{\nu} = -v^2 (Y \cdot M_M^{-1} \cdot Y^t)$$



$$\nu \simeq U_{\nu}^{\dagger} (\nu_L - \theta \nu_R^c) + \text{h.c.}$$

Light neutrinos

$$N \simeq U_N^{\dagger} (\nu_R + \theta^t \nu_L^c) + \text{h.c.}$$

Heavy neutrinos (HNL)

- $n \geq 2$ HNL generations needed to explain light neutrino masses
 - What is our prior on n ?
 - $n = 2$: Minimality (ν MSM)
 - $n = 3$: Flavour symmetries, gauge extensions (LRSM,...)

Type-I seesaw mechanism

Type-I seesaw Lagrangian (below EWSB)

$$\mathcal{L} \supset \underbrace{v Y_{\alpha i}}_{\text{Dirac}} \bar{\ell}_{\alpha} \nu_{Ri} + \frac{1}{2} \bar{\nu}_{Ri}^c \underbrace{(M_M)_{ij}}_{\text{Majorana}} \nu_{Rj} + \text{h.c.}$$

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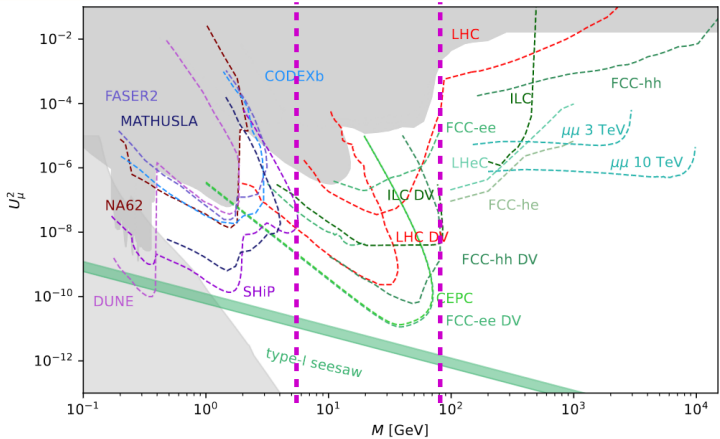
Heavy neutrinos (HNL)

- $n \geq 2$ HNL generations needed to explain light neutrino masses
- Experimental sensitivity expressed in terms of mixing

$$U_{\alpha}^2 = \sum_i |\theta_{\alpha i}|^2 = \sum_i |v(Y \cdot M_M^{-1})_{\alpha i}|^2$$

Testing the type-I seesaw

Many different ways to probe HNLs:



[Bose et al; 2209.13128]

Meson decays

W/Z decays

Virtual W/Z exchange

Leptogenesis

Sakharov conditions:

- ★ C- and CP-violation
- ★ Deviation from thermal equilibrium
- ★ Baryon number violation

Leptogenesis

Sakharov conditions:

- ★ C- and CP-violation
 - RHN oscillations and decay

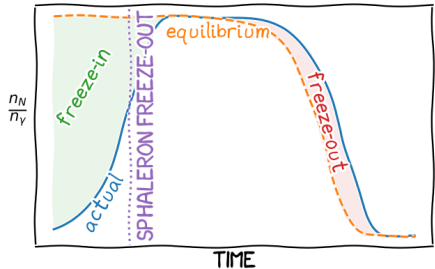
- ★ Deviation from thermal equilibrium
 - Freeze-in and freeze-out of the RHN

- ★ Baryon number violation

Leptogenesis

Sakharov conditions:

- ★ C- and CP-violation
 - RHN **oscillations** and decay
- ★ Deviation from thermal equilibrium
 - **Freeze-in** and freeze-out of the RHN
- ★ Baryon number violation

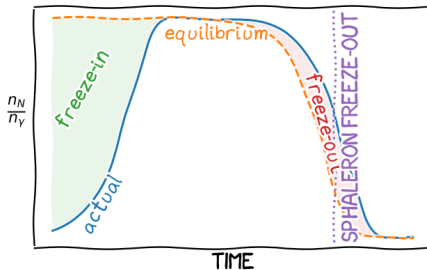


[Klarić/Shaposhnikov/Timiryasov, 2103.16545]

Leptogenesis

Sakharov conditions:

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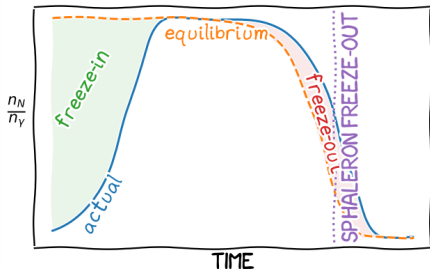


[Klarić/Shaposhnikov/Timiryasov, 2103.16545]

Leptogenesis

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 - Freeze-in and **freeze-out** of the RHN
- ★ Baryon number violation
 - Sphaleron process



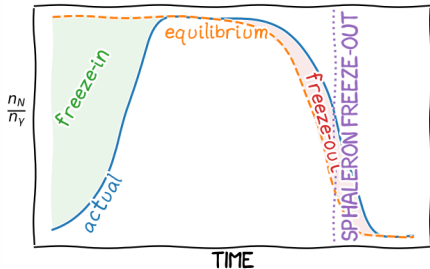
[Klarić/Shaposhnikov/Timiryasov, 2103.16545]



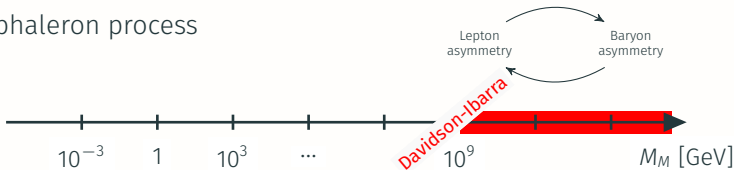
Leptogenesis

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[Klarić/Shaposhnikov/Timiryasov, 2103.16545]



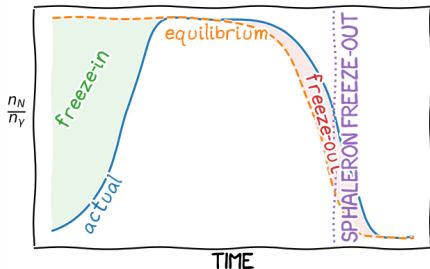
Thermal leptogenesis

[Fukugita/Yanagida '86]

Leptogenesis

Sakharov conditions:

- ★ C- and CP-violation
 - RHN oscillations and **decay**
- ★ Deviation from thermal equilibrium
 - Freeze-in and **freeze-out** of the RHN
- ★ Baryon number violation
 - Sphaleron process



[Klarić/Shaposhnikov/Timiryasov, 2103.16545]



Low-scale leptogenesis

[Akhmedov/Rubakov/Smirnov '98, Pilaftsis/Underwood '03, Asaka/Shaposhnikov '05, ...]

Thermal leptogenesis

[Fukugita/Yanagida '86]

Quantum kinetic equations

$$i \frac{d}{dt} \rho = [H, \delta \rho] - \frac{i}{2} \{ \Gamma, \delta \rho \} - i \sum_{a \in \{e, \mu, \tau\}} \tilde{\Gamma}_a \frac{\mu_a}{T} f_F (1 - f_F),$$

$$i \frac{d}{dt} \bar{\rho} = -[H, \delta \bar{\rho}] - \frac{i}{2} \{ \Gamma, \delta \bar{\rho} \} + i \sum_{a \in \{e, \mu, \tau\}} \tilde{\Gamma}_a \frac{\mu_a}{T} f_F (1 - f_F),$$

$$\frac{d}{dt} n_{\Delta_a} = - \frac{2i \mu_a}{T} \int \frac{d^3 \vec{k}}{(2\pi)^3} \text{Tr}[\Gamma_a] f_F (1 - f_F) + i \int \frac{d^3 \vec{k}}{(2\pi)^3} \text{Tr}[\tilde{\Gamma}_a (\delta \bar{\rho} - \delta \rho)].$$

Density matrix

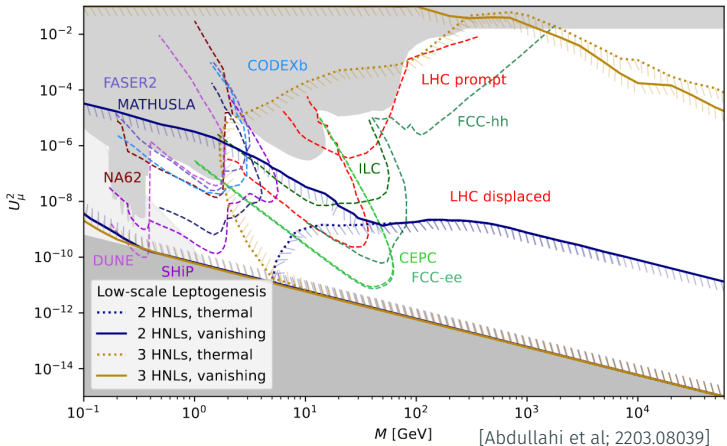
Effective Hamiltonian

Lepton asymmetry

Interaction rates

- **Interaction rates** can be
 - ★ Fermion number **conserving** $\sim (Y^\dagger Y) T$
 - ★ Fermion number **violating** $\sim (Y^\dagger Y^*) \frac{M^2}{T}$
- Refined calculation subject to intensive studies over the last years, e.g. Anisimov/Bedak/Bödeker '10, Garny/Kartavtsev/Hohenegger '11, Drewes/Garbrecht/Gueter/Klarić '16, Hernandez/Kekic/Lopez-Pavon/Racker/Salvado '16, Laine/Ghiglieri '16 '18, Klarić/Shaposhnikov/Timiryasov '21, ...

$n = 3$ parameter space, NH

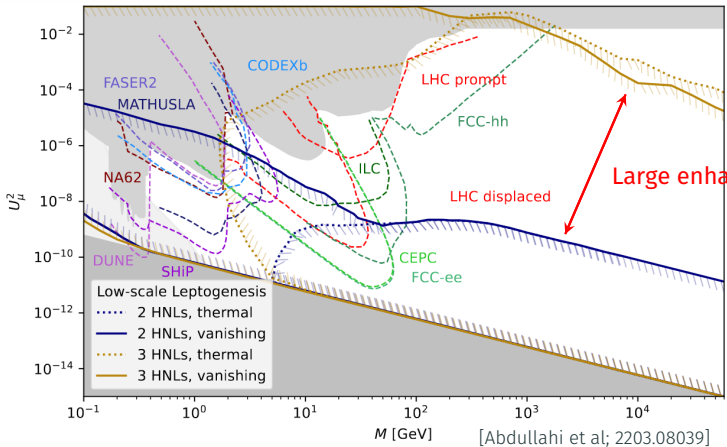


$n = 2$ lines from [Klarić/Shaposhnikov/Timiryasov, 2103.16545]

$n = 3$ lines from [Drewes/YG/Klarić, 2106.16226]

- Can potentially produce enough HNLs to **test leptogenesis** !

$n = 3$ parameter space, NH



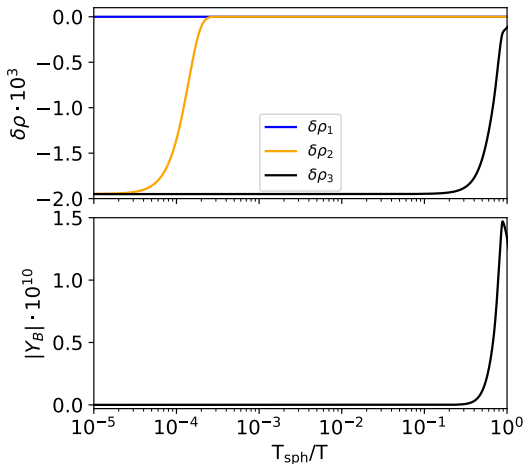
$n = 2$ lines from [Klarić/Shaposhnikov/Timiryasov, 2103.16545]

$n = 3$ lines from [Drewes/YG/Klarić; 2106.16226]

- Can potentially produce enough HNLs to **test leptogenesis** !

Why such large mixings ?

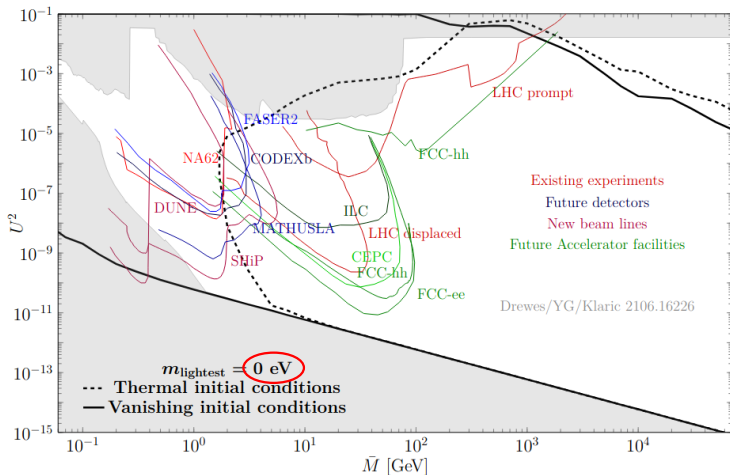
$$U^2 = 0.0248, \bar{M} = 100 \text{ GeV and } m_{\text{lightest}} = 0 \text{ eV}$$



[YG; 2305.06663]

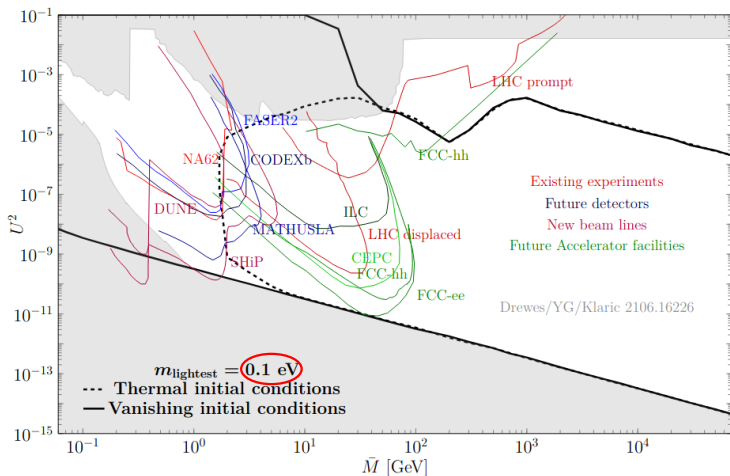
- Large mixing angles allow **late equilibration** of one HNL $U_i^2 \ll 1$
↳ Late BAU production, less time for washout

$n = 3$ parameter space, NH



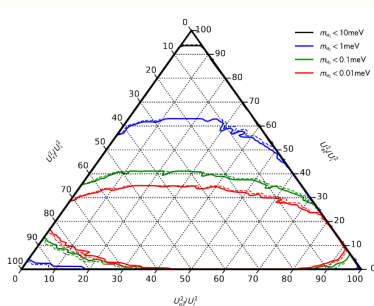
- Can potentially produce enough HNLs to test leptogenesis !

$n = 3$ parameter space, NH

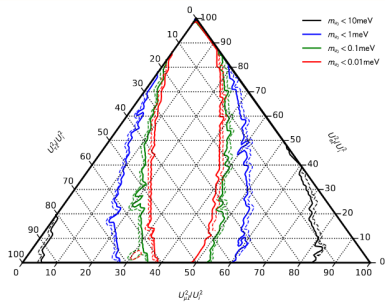


- Can potentially produce enough HNLs to test leptogenesis !

Flavour ratios



Normal ordering

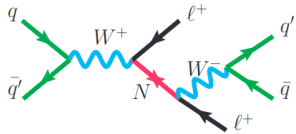


Inverted ordering

[Chrzaszcz et al; 1908.02302]

- For optimistic scenarios ($\sim 10^5$ events at FCC-ee), can measure flavour ratios $\frac{U_{\alpha}^2}{U^2}$ at percent level !
- Can give a hint on value of m_0 !

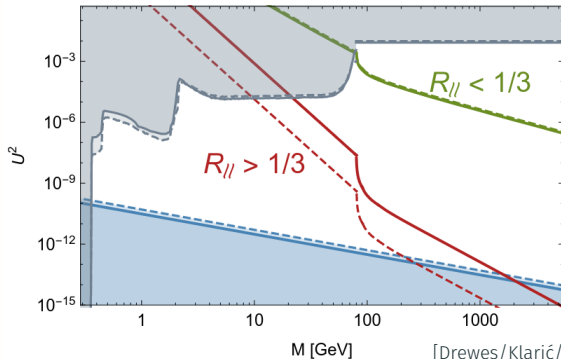
Lepton number violation at colliders



[CMS collaboration; 1806.10905]

- Ratio of lepton number **violating** to **conserving** decays parametrised by

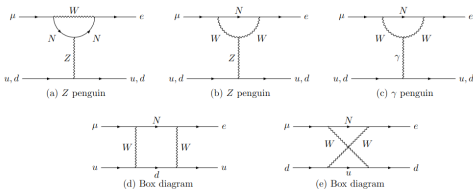
$$R_{ll} = \frac{\Delta M_{\text{phys}}^2}{2\Gamma_N^2 + \Delta M_{\text{phys}}^2}$$



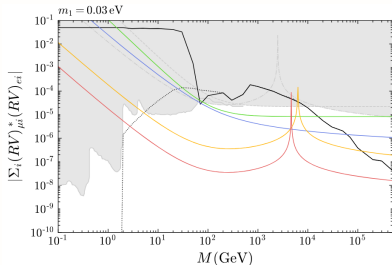
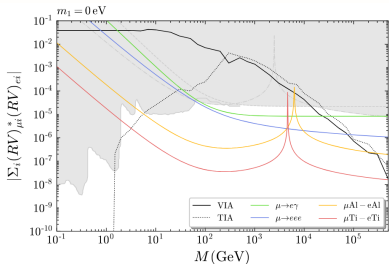
[Drewes/Klarić/Klose; 1907.13034]

Testing leptogenesis through CLFV experiments

- HNLs also lead to charge lepton flavour violation.



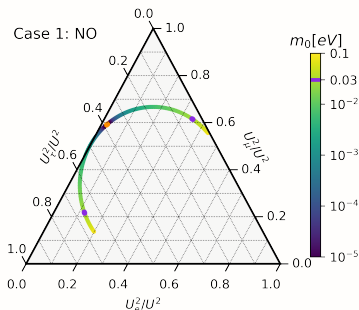
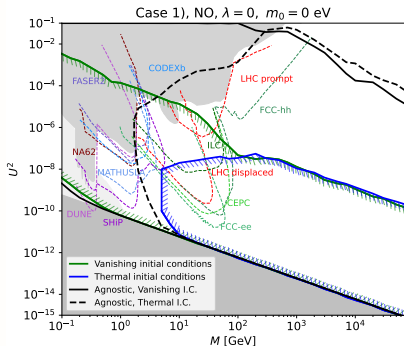
[Urquia-Calderon/Timiryasov/Ruchayskiy; 2206.04540]



[Granelli/Klarić/Petcov; 2206.04342]

Impact of flavour symmetries

- Discrete flavour symmetries ($\Delta(6n^2)$) provide explanation for PMNS matrix structure
- Scenarios with reduced parameter space but highly predictive !



[Drewes/YG/Hagedorn/Klarič 24xx.xxxxx; YG 2401.04840]

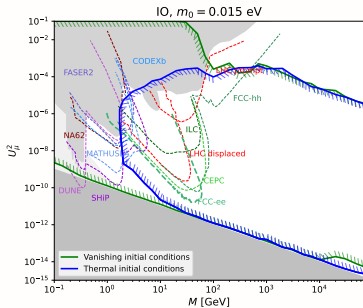
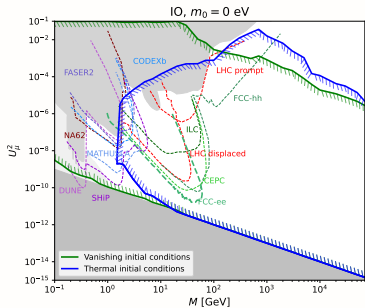
What should I take home ?

- Right-handed neutrinos provide minimal solution for ν masses + baryon asymmetry
- Leptogenesis parameter space largely enhanced for $n = 3$
- Large mixing angle opens up the possibility of testing leptogenesis by combining information from colliders, CLFV, ν oscillations, ...
- Combined with flavour symmetric explanation of PMNS: very predictive !
- Collider testability of $n = 3$ scenario to be further explored

Thanks for your attention!

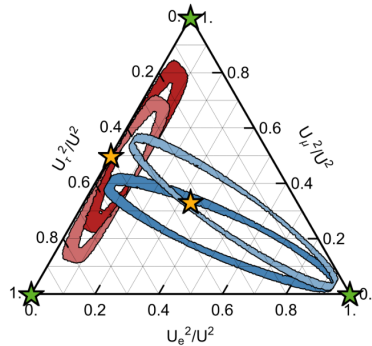
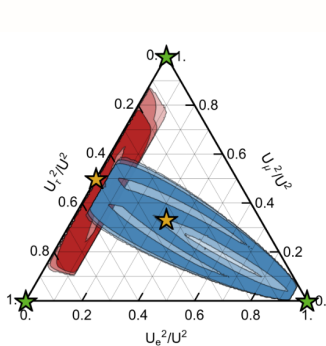
Appendix

$n = 3$ parameter space, IO



- Similar enhancement of the parameter space for IO.

Impact of low energy measurements on $\frac{U_{\alpha}^2}{U^2}$



Current ν oscillation data

DUNE projections

- New (more realistic) benchmarks proposed beyond the 1-flavour approximation
- DUNE measurement of δ could constrain the mixing to each SM flavour, hence leptogenesis

How to reach large coupling ? B-L approximate symmetry

Naive seesaw bound

$$m_\nu = -v^2(Y \cdot M_M^{-1} \cdot Y^t) \Leftrightarrow U_i^2 \sim \frac{m_\nu}{M_i} \sim 10^{-10} \frac{\text{GeV}}{M_i}$$

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B-L approximate symmetry

Majorana mass

$$\bar{M} \cdot \begin{pmatrix} 1 - \mu & 0 & 0 \\ 0 & 1 + \mu & 0 \\ 0 & 0 & \mu' \end{pmatrix}$$

Yukawa coupling

$$\begin{pmatrix} f_e(1 + \epsilon_e) & if_e(1 - \epsilon_e) & f_e \epsilon'_e \\ f_\mu(1 + \epsilon_\mu) & if_\mu(1 - \epsilon_\mu) & f_\mu \epsilon'_\mu \\ f_\tau(1 + \epsilon_\tau) & if_\tau(1 - \epsilon_\tau) & f_\tau \epsilon'_\tau \end{pmatrix}$$

Technically natural: Small m_ν from small symmetry breaking parameters $\mu, \epsilon, \epsilon' \ll 1$
Consistent with large U^2 .

How to reach large coupling ? B-L approximate symmetry

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$$m_\nu = -v^2(Y \cdot M_M^{-1} \cdot Y^t) \Leftrightarrow U_i^2 \sim \frac{m_\nu}{M_i} \sim 10^{-10} \frac{\text{GeV}}{M_i}$$

B-L approximate symmetry

Majorana mass

$$\bar{M} \cdot \begin{pmatrix} 1 - \mu & 0 & 0 \\ 0 & 1 + \mu & 0 \\ 0 & 0 & \mu' \end{pmatrix}$$

Yukawa coupling

Pseudo-Dirac pair

Decoupled

$$\begin{pmatrix} f_e(1 + \epsilon_e) & if_e(1 - \epsilon_e) \\ f_\mu(1 + \epsilon_\mu) & if_\mu(1 - \epsilon_\mu) \\ f_\tau(1 + \epsilon_\tau) & if_\tau(1 - \epsilon_\tau) \end{pmatrix} \begin{pmatrix} f_e \epsilon'_e \\ f_\mu \epsilon'_\mu \\ f_\tau \epsilon'_\tau \end{pmatrix}$$

Technically natural: Small m_ν from small symmetry breaking parameters $\mu, \epsilon, \epsilon' \ll 1$
Consistent with large U^2 .

Seesaw parameter space

Consistency with ν -oscillation data induced by Casas-Ibarra parametrisation

$$F = \frac{i}{V} U_\nu \sqrt{m_\nu^{diag}} R \sqrt{M_M}$$

n=2

- 2 CP-violating phases
- 3 PMNS angles (3/3 fixed)
- 2 light neutrino masses (2/2 fixed)
- 1 complex Euler angle
- 2 Majorana masses

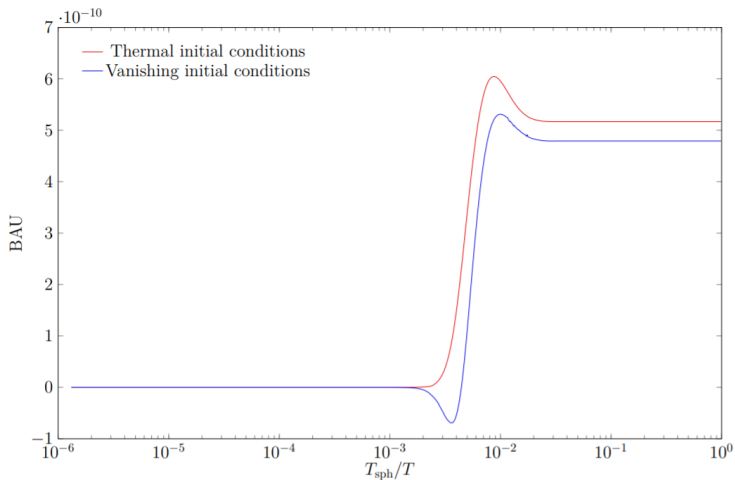
6 free parameters

n=3

- 3 CP-violating phases
- 3 PMNS angles (3/3 fixed)
- 3 light neutrino masses (2/3 fixed)
- 3 complex Euler angles
- 3 Majorana masses

13 free parameters

Thermal vs vanishing initial conditions



At large \bar{M} , parameter space for thermal I.C. is larger because **asymmetry** produced during **freeze-in** and **freeze-out** have **opposite** signs.

Vanilla thermal leptogenesis

Assumptions:

- * Asymmetry generated by heavy neutrino decays
- * Hierarchical mass spectrum $M_1 \ll M_i$
- * Unflavoured

Vanilla thermal leptogenesis

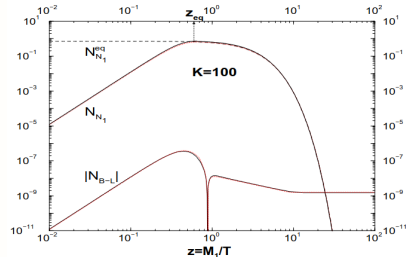
Assumptions:

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Boltzmann equations

$$\frac{d}{dz} n_1 = -\frac{\Gamma_D}{HZ} (n_1 - n_1^{eq})$$

$$\frac{d}{dz} n_{B-L} = \epsilon_1 \frac{\Gamma_D}{HZ} (n_1 - n_1^{eq}) - \frac{\Gamma_W}{HZ} n_{B-L}$$



[Buchmüller/Di Bari/Plümacher, '04]

Vanilla thermal leptogenesis

Assumptions:

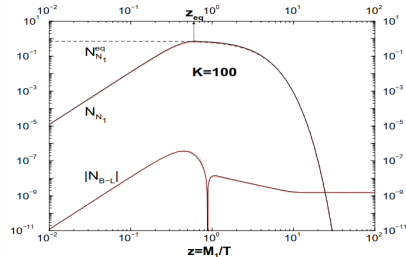
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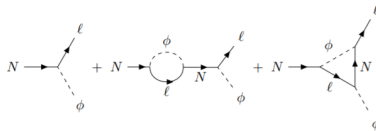
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• Decay asymmetry $\epsilon_1 \equiv \frac{\Gamma_{N_1 \rightarrow \ell + \phi} - \Gamma_{N_1 \rightarrow \bar{\ell} + \phi^*}}{\Gamma_{N_1 \rightarrow \ell + \phi} + \Gamma_{N_1 \rightarrow \bar{\ell} + \phi^*}}$



[Buchmüller/Di Bari/Plümacher, '04]



Vanilla thermal leptogenesis

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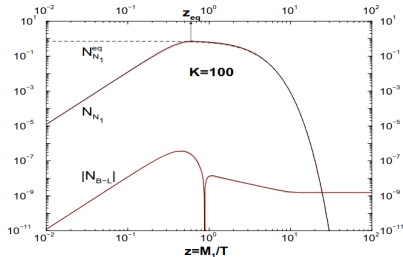
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- Decay asymmetry $\epsilon_1 \equiv \frac{\Gamma_{N_1 \rightarrow \ell + \phi} - \Gamma_{N_1 \rightarrow \bar{\ell} + \phi^*}}{\Gamma_{N_1 \rightarrow \ell + \phi} + \Gamma_{N_1 \rightarrow \bar{\ell} + \phi^*}}$
- For large mass splittings $|\epsilon_1| \lesssim \frac{3}{8\pi} \frac{M_1}{v^2} \sqrt{\Delta m_{23}^2}$ leading to the Davidson-Ibarra bound

$$M_1 \gtrsim 4 \cdot 10^8 \text{ GeV}$$

↪ Direct detection ☹️



[Buchmüller/Di Bari/Plümacher, '04]

Low-scale models

- Traditionally, 2 main mechanisms:

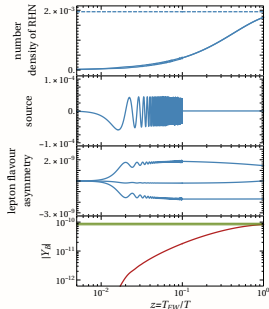
ARS Leptogenesis

Asymmetry produced during
freeze-in from CP-violating
HNL oscillations

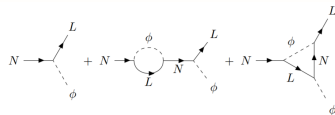


Resonant leptogenesis

Resonant enhancement of
CP-violation from small mass
splittings



[Drewes/Garbrecht/Gueter/Klarić; 1606.06690]



Decay asymmetry:

$$\epsilon_i \simeq \frac{\text{Im}(\gamma^\dagger \gamma)_{ij}^2 (M_{N_i}^2 - M_{N_j}^2) \cdot M_{N_i} \Gamma_N}{(\gamma^\dagger \gamma)_{ii} (\gamma^\dagger \gamma)_{jj} (M_{N_i}^2 - M_{N_j}^2)^2 + M_{N_i}^2 \Gamma_N^2}$$

Low-scale models

- Traditionally, 2 main mechanisms:

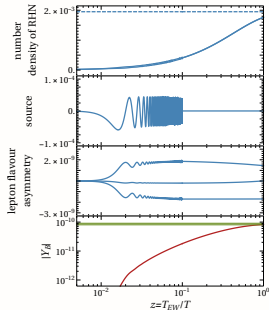
ARS Leptogenesis

Asymmetry produced during
freeze-in from CP-violating
HNL oscillations

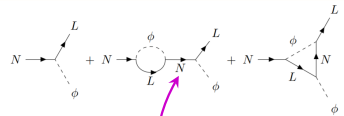


Resonant leptogenesis

Resonant enhancement of
CP-violation from small mass
splittings



[Drewes/Garbrecht/Gueter/Klarić; 1606.06690]



Decay asymmetry:

$$\epsilon_i \simeq \frac{\text{Im}(\gamma^\dagger \gamma)_{ij}^2 (M_{N_i}^2 - M_{N_j}^2) \cdot M_{N_i} \Gamma_N}{(\gamma^\dagger \gamma)_{ii} (\gamma^\dagger \gamma)_{jj} (M_{N_i}^2 - M_{N_j}^2)^2 + M_{N_i}^2 \Gamma_N^2}$$

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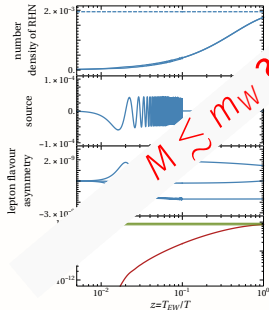
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$M > 130 \text{ GeV} ?$

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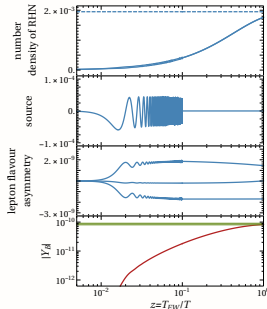
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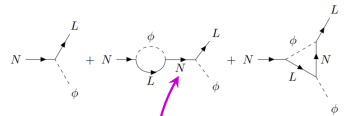
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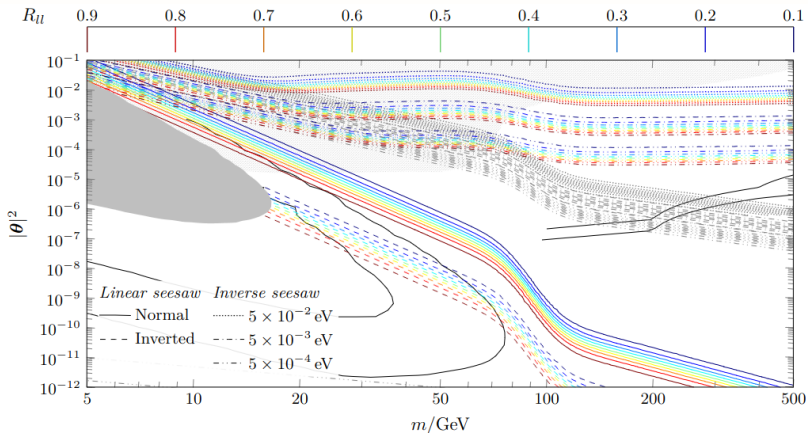
→ Two regimes of the same mechanism ! Represented by the same set of kinetic equations (cfr. [Garbrecht; 1812.02651] for a review)



Decay asymmetry:

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Lepton number violation at colliders

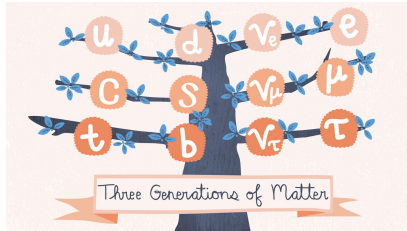


[Antusch/Hajer/Roskopp, 2307.06208]

In practice, decoherence effects can make testability prospects even more optimistic !

Motivations for flavour symmetries

- Why 3 generations in the Standard Model ?



[Sandbox Studio, Chicago]

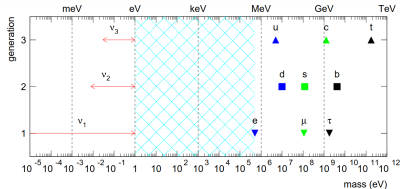
Motivations for flavour symmetries

- Why 3 generations in the Standard Model ?
- Hierarchy in the CKM matrix structure ?

$$|U_{\text{CKM}}| \approx \begin{pmatrix} 0.97 & 0.22 & 0.004 \\ 0.22 & 0.99 & 0.04 \\ 0.008 & 0.04 & 1.01 \end{pmatrix}$$

Motivations for flavour symmetries

- Why 3 generations in the Standard Model ?
- Hierarchy in the CKM matrix structure ?
- Hierarchy in the fermion masses ?



[A. de Gouvea; hep-ph/0411274]

Motivations for flavour symmetries

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- Why such neutrino mixing pattern ? In particular, why the PMNS matrix

$$|U_{\text{PMNS}}| \approx \begin{pmatrix} 0.82 & 0.55 & 0.15 \\ 0.29 & 0.59 & 0.75 \\ 0.49 & 0.59 & 0.64 \end{pmatrix}$$

is so close to a tri-bimaximal mixing

$$U_{\text{TB}} = \begin{pmatrix} \sqrt{2/3} & \sqrt{1/3} & 0 \\ -\sqrt{1/6} & \sqrt{1/3} & \sqrt{1/2} \\ -\sqrt{1/6} & \sqrt{1/3} & -\sqrt{1/2} \end{pmatrix}.$$

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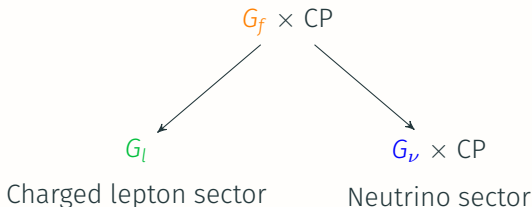
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Discrete flavour symmetries

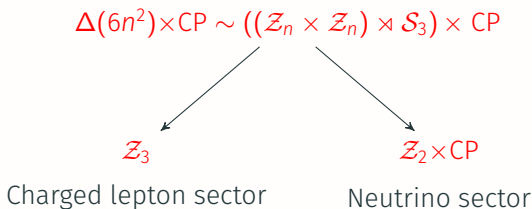
- Discrete symmetry G_f at high scale, broken at low scale into residual symmetries $G_l, G_\nu \subset G_f$.



- What group to choose ?
 - * G_f discrete subgroup of $U(3)$ (not always necessary)
 - * G_f non-abelian to avoid texture zero
 - * G_l abelian and minimal to avoid imposing too strong constraints on the charged lepton masses
 - * G_ν as minimal as possible

Discrete flavour symmetries

- Discrete symmetry G_f at high scale, broken at low scale into residual symmetries $G_l, G_\nu \subset G_f$.



In our setup

Prediction

$$U_{\text{PMNS}} = \Omega(\mathbf{3}) R_{ij}(\theta_L) K_\nu$$

$$Y = \Omega(\mathbf{3}) R_{ij}(\theta_L) \text{diag}(y_1, y_2, y_3) P_{kl}^{ij} R_{kl}(-\theta_R) \Omega(\mathbf{3}')^\dagger$$

Parametrisation of flavour symmetries

- 4 qualitatively different scenarios:

Case 1), Case 2), Case 3 a) and Case 3 b.1).

- 13 \rightarrow 6 or 7 free parameters: For Case 1),

$$\phi_S, \theta_R, M_1 \approx M_2 \approx M_3, m_0.$$

\rightarrow Better **analytical understanding** of the parameter space.

- Total coupling proportional to

$$U^2 \propto \frac{1}{|\cos(2\theta_R)|}, \frac{1}{|\sin(2\theta_R)|}.$$

$\hookrightarrow \theta_R \rightarrow k\frac{\pi}{4}, k \in \mathbb{Z}$ (but enhanced residual symmetry) leads to experimentally testable scenarios !

- Can relate low- and high-scale parameters. For Case 1):

$$\sin(\delta) = 0, |\sin(\alpha)| = |\sin(6\phi_S)|, \sin(\beta) = 0.$$

CP-violation combinations

- Perturbatively,

$$Y_B \propto \text{Tr} \left(\tilde{\Gamma}_\alpha (\delta\rho - \delta\bar{\rho}) \right) \propto \text{Tr} \left(\tilde{\Gamma}_\alpha [H_N, \Gamma] \right)$$

$$H_N = \frac{M_M^2}{2E} + h_+(T) Y^\dagger Y + h_-(T) Y^t Y^*, \quad \Gamma, \tilde{\Gamma} = \pm \gamma_+(T) Y^\dagger Y + \gamma_-(T) Y^t Y^*$$

- BAU production governed by

$$C_{\text{LFV},\alpha} = i \text{Tr} \left([M_M^2, Y^\dagger Y] Y^\dagger P_\alpha Y \right),$$

$$C_{\text{LNV},\alpha} = i \text{Tr} \left([M_M^2, Y^\dagger Y] Y^T P_\alpha Y^* \right),$$

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Flavour violating only $\mathcal{C}_{\text{LFV},\alpha} = i \text{Tr} \left([M_M^2, Y^\dagger Y] Y^\dagger P_\alpha Y \right),$

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Flavour violating only, can be $\neq 0$ for $\Delta M = 0!$

Violates lepton number

$$\sum_\alpha \mathcal{C}_{\text{LNV},\alpha} \neq 0$$

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$$C_{\text{LFV},\alpha} \sim \frac{8}{3} M^2 \kappa y_2 y_3 (y_2^2 - y_3^2) \sin \theta_{L,\alpha} \sin \theta_R \cos 3\phi_S,$$

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where

$$\theta_{L,\alpha} = \theta_L + \rho_\alpha \frac{4\pi}{3} \text{ with } \rho_e = 0, \rho_\mu = +1, \rho_\tau = -1.$$

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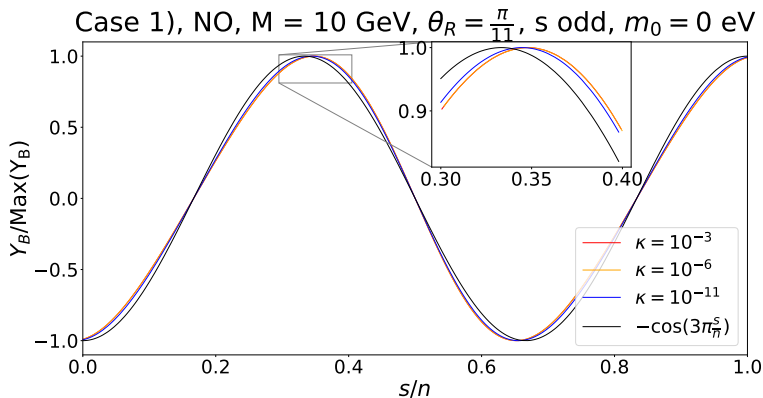


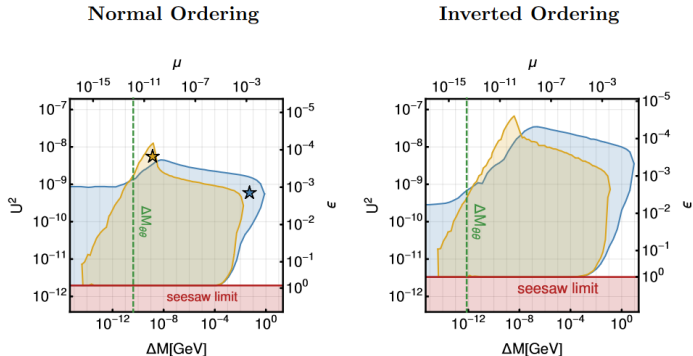
Figure 1: Vanishing initial conditions, $\lambda = 0$

[Drewes/Hagedorn/YG/Klarić; 2203.08538]

- Correlation between Y_B and low-energy observables. Here,

$$\sin(\alpha) = \sin(6\pi \frac{S}{n}).$$

Leptogenesis in the mass degenerate case



[Antusch/Cazzato/Drewes/Fischer/Garbrecht/Gueter/Klarić; 1710.03744]

See also [Sandner/Hernandez/Lopez-Pavon/Rius; 2305.14427]

- Leptogenesis possible for $\Delta M = 0$ thanks to Higgs and thermal mass splittings

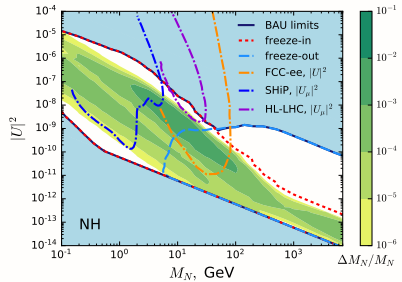
$$\Delta M_{\text{phys}} \sim h_+(T)Y^\dagger Y + h_-(T)Y^t Y^*$$

- Lepton asymmetry proportional to CP-violating combination

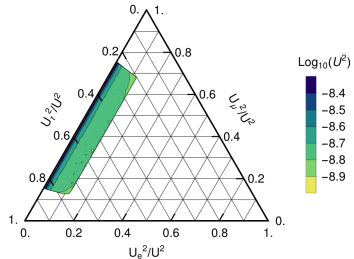
$$\text{Tr} \left(\tilde{\Gamma}_\alpha [H_N, \Gamma] \right) \sim \text{Tr} \left(\left[\hat{Y}^t \hat{Y}^*, \hat{Y}^\dagger \hat{Y} \right] \hat{Y}^t P_\alpha \hat{Y}^* \right) \neq 0!$$

$n = 2$ (ν MSM) parameter space

- Parameter space for **freeze-in** and **freeze-out** are **connected**
- Sizeable fraction of the parameter space can be tested at colliders or fixed target experiments
- Relies on **flavour hierarchies** to reach large U^2
- IH parameter space larger than for NH for $M \lesssim \mathcal{O}(100)$ GeV due to stronger washout



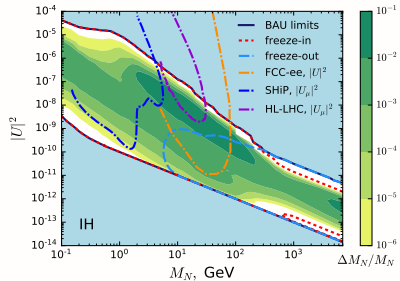
[Klarić/Shaposhnikov/Timiryasov; 2103.16545]



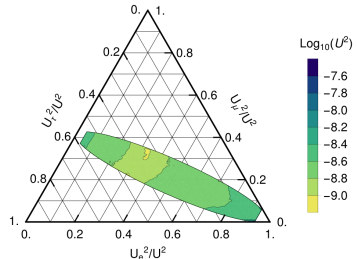
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