# The Light Dark Matter eXperiment (LDMX)

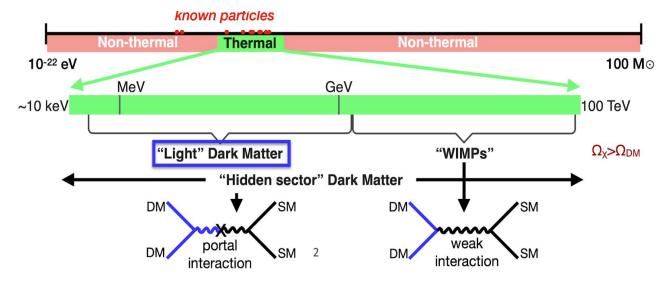
#### Jessica Pascadlo, University of Virginia On Behalf of the LDMX Collaboration May 13, 2024





#### Light Dark Matter (LDM)

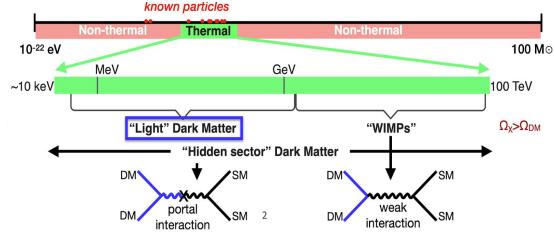
- Thermal relic DM model works for keV to TeV range
- Two categories for thermal dark matter: WIMPs & LDM
  - Still viable parameter space for WIMPs is dwindling, leading to the rise of "light" dark matter research



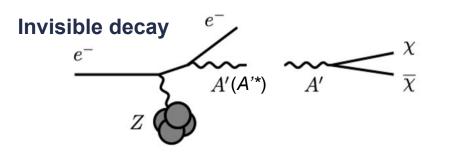


#### Light Dark Matter (LDM)

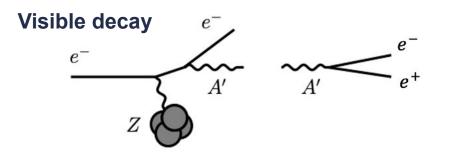
- Thermal relic DM model works for keV to TeV range
- $\begin{array}{c} \chi \\ \chi \\ \chi \\ \chi \\ \end{array} \\ \begin{array}{c} A' \\ \alpha_D \\ \epsilon^2 \\ \epsilon^2 \\ \end{array} \\ \begin{array}{c} \gamma \\ \alpha \\ SM \\ \end{array} \\ \begin{array}{c} SM \\ SM \\ SM \\ \end{array} \\ \begin{array}{c} SM \\ SM \\ SM \\ \end{array} \\ \begin{array}{c} SM \\ SM \\ SM \\ \end{array} \\ \begin{array}{c} SM \\ SM \\ SM \\ \end{array} \\ \begin{array}{c} SM \\ SM \\ SM \\ \end{array} \\ \begin{array}{c} SM \\ SM \\ SM \\ SM \\ SM \\ \end{array}$
- Two categories for thermal dark matter: WIMPs & LDM
  - Still viable parameter space for WIMPs is dwindling, leading to the rise of "light" dark matter research
- Important benchmark is vector portal with dark photon (A') kinetically mixing with SM photon
- Accelerator experiments provide generic probe for LDM independent of portal and type of DM particle



#### **Dark Photon Signatures**



- $m_{A'} > 2m_{\chi}$  or off-shell production
  - Decays into DM particles that don't interact with detector
- For LDMX, characterized by some missing energy/momentum in the detector as a whole



- $2m_e < m_{A'} < 2m_{\chi}$ 
  - Decays into SM particles
  - Long-lived
- For LDMX, characterized by a displaced, sudden appearance of energy deposited in some downstream part of the detector

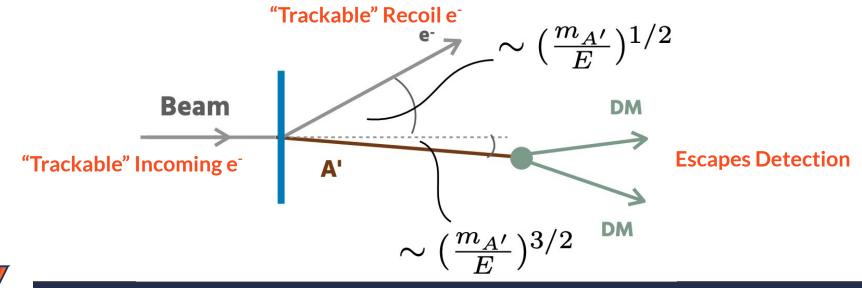


#### **Dark Photon Kinematics at a Fixed Target Experiment**

Fixed target signal characteristics:

 $\rightarrow$  Dark bremsstrahlung A' production (invisible decay)

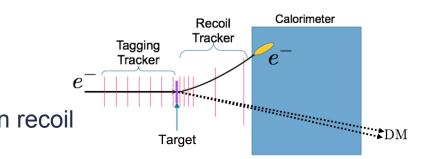
 $\rightarrow$  A's take most of the beam energy; only visible final state particle is a soft recoil electron



#### LDMX Concept

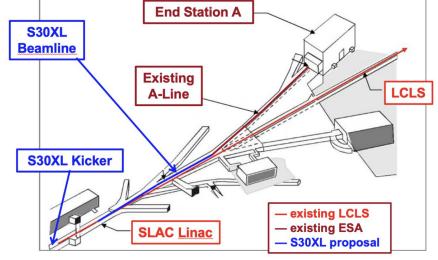
- Look for missing momentum (and energy) in recoil electron
  - Use detectors to identify the missing momentum & energy  $\rightarrow$  DM production!
  - Particle ID
  - Transverse momentum of recoil e<sup>-</sup> used as discriminator/identifier

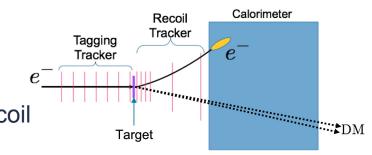




#### LDMX Concept

- Look for missing momentum (and energy) in recoil electron
  - Use detectors to identify the missing momentum & energy  $\rightarrow$  DM production! Ο
  - Particle ID  $\bigcirc$
  - Transverse momentum of recoil e<sup>-</sup> used as discriminator/identifier 0
- e<sup>-</sup> beam provided by SLAC
  - Planning on 4 GeV and 8 GeV runs Ο
- Must be able to tag and reconstruct every electron
  - Do this for up to 1e16 electrons  $\bigcirc$
  - Use low current, high repetition rate of 37 MHz, Ο
    - $\mu = 1$  to a few

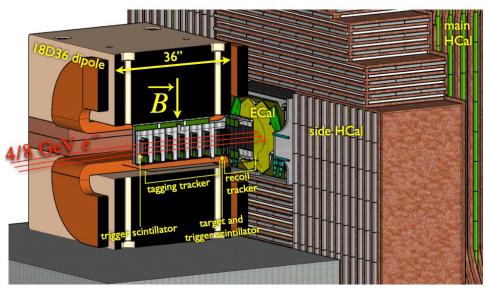




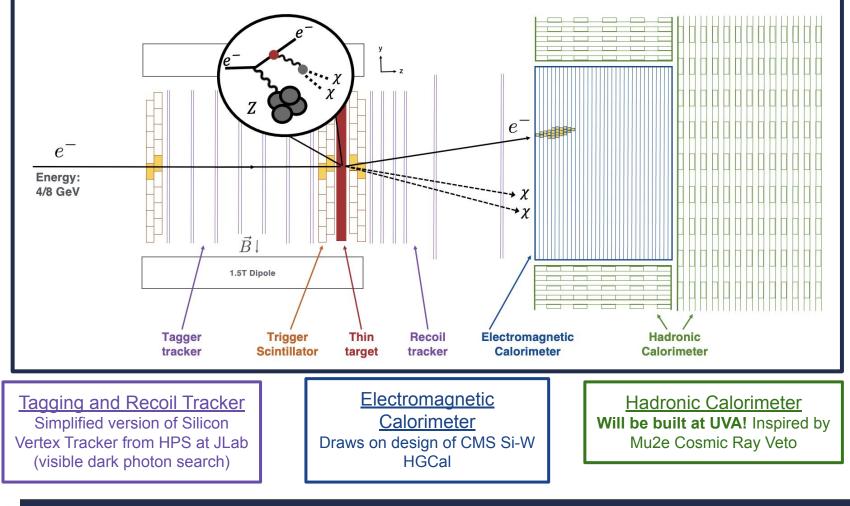
#### **LDMX Design**

Basically dumping the beam onto our detector!

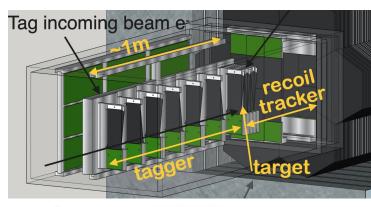
- $\rightarrow$  Need hermetic, radiation tolerant detector designed for high beam rates
- **Tagging tracker**: low acceptance and high resolution at beam energy
- **Recoil tracker**: large acceptance and high resolution at low particle momenta
- **Trigger Scintillator**: scintillator bars provide fast, accurate count of incoming electrons
- Electromagnetic calorimeter: fast, good energy resolution, and high granularity
- Hadronic calorimeter: high veto efficiency
  of neutral hadrons

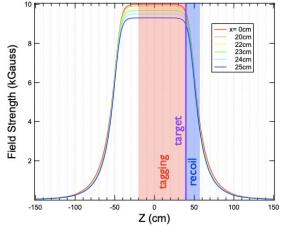






## **Trackers and Trigger Scintillator**





- Tagging tracker
  - Measures incoming beam electron

#### Recoil tracker

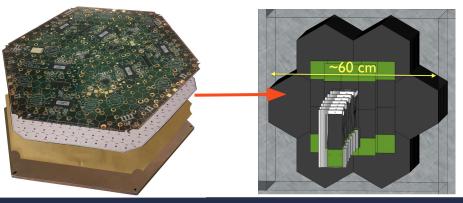
- Measures recoil electron and vetoes extra particles
- Trigger Scintillator
  - Arrays of scintillator bars provide fast count of incoming electrons
  - Used as an input to the missing energy trigger

Scintillator pads around target



#### **Electromagnetic Calorimeter (Ecal)**

- Si-W sampling calorimeter (based on CMS HGCal upgrade)
  - ~ 40  $X_0$  depth (34 Si layers) for extraordinary shower containment
- Provides fast missing energy trigger (E < 1.5 GeV for 4 GeV beam)
- <u>High granularity</u> transverse and longitudinal shower shapes can be exploited to reject backgrounds
- Capable of MIP tracking to further improve background rejection



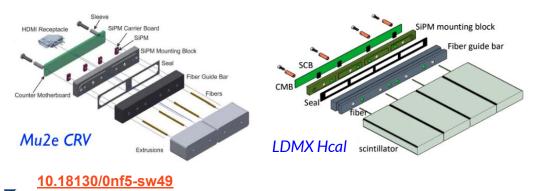
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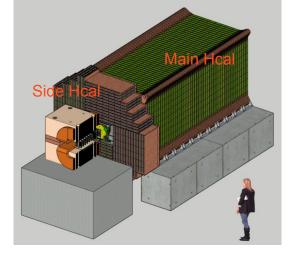


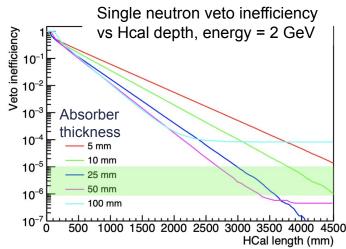
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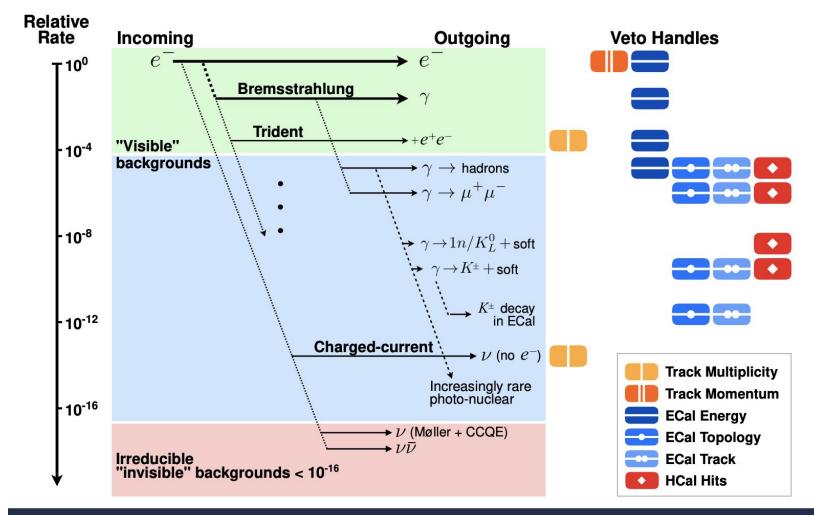
## Hadronic Calorimeter (Hcal)

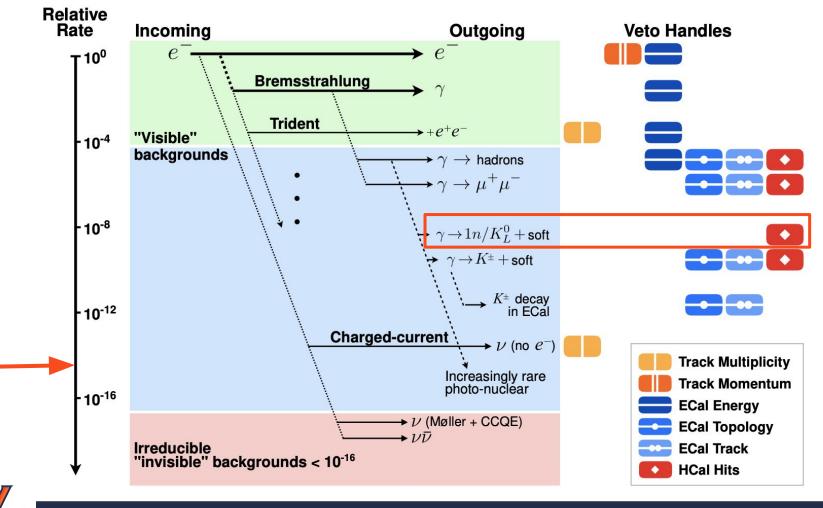
- Segmented scintillator/steel sampling calorimeter
  - $_{\circ}$  96 layers of 20 (25) mm of polystyrene (Fe) → ~17λ
- Main Hcal detects neutral hadrons (mostly K<sub>L</sub>, n) produced in photonuclear reactions, and MIPs
- Side Hcal rejects wide angle brem and  $\gamma \rightarrow \mu + \mu$ -
- Extruded scintillator bars with inserted wavelength-shifting fibers, read out with SiPMs

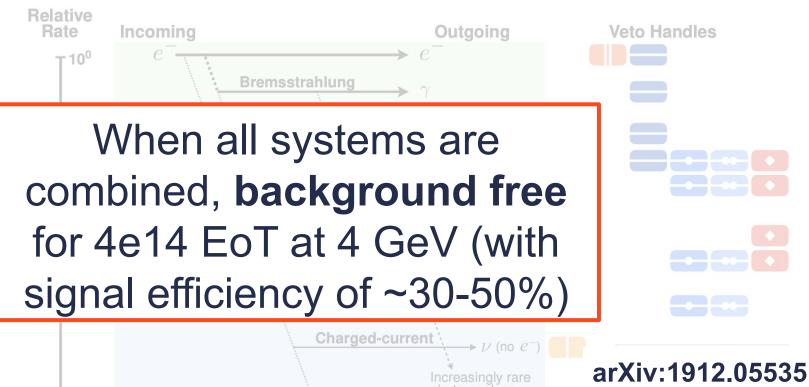


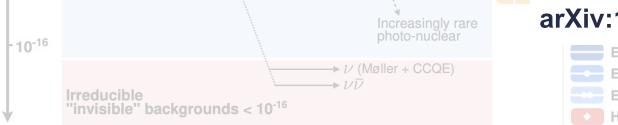








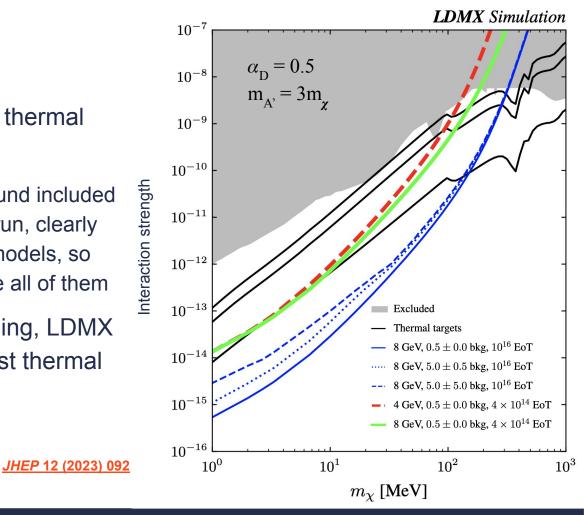






## **Projected Sensitivity**

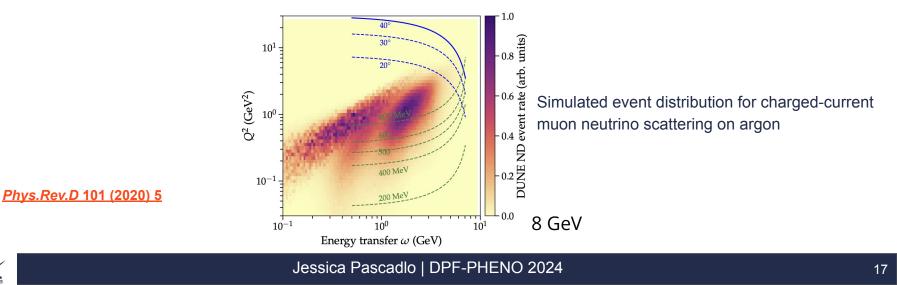
- LDMX is able to reach ALL thermal targets!
  - Even with some background included for the 8 GeV/1e16 EoT run, clearly pass all the benchmark models, so we will be able to exclude all of them
- Within a few weeks of running, LDMX could begin to reach the first thermal targets





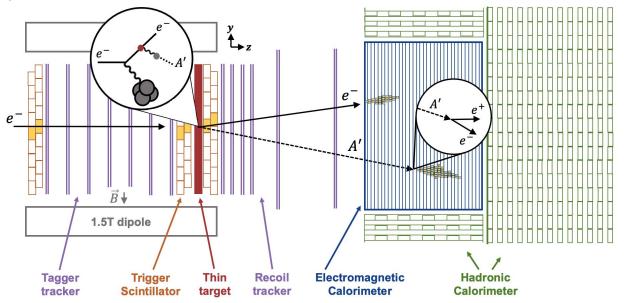
#### **Additional Physics Program: Neutrino Cross-Sections**

- LDMX offers a broader physics program beyond a missing momentum search for LDM
- Small angle acceptance (nearly hermetic) and fully reconstructing final/initial states allows for several unique measurements
  - Electro-nuclear scattering measurements of interest to neutrino experiments such as DUNE



#### **Additional Physics Program: Visible Signatures**

- LDMX offers a broader physics program beyond a missing momentum search for LDM
- Many models could be tested (minimal dark photon, ALPs, SIMPS, etc.)



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#### Conclusions

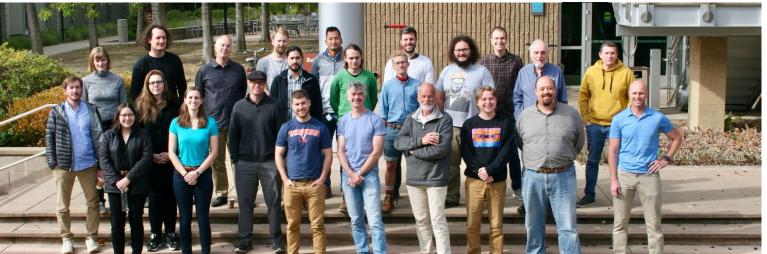
- Thermal DM is a simple and compelling scenario, and MeV-GeV scale is a logical place to look extension of WIMPs!
- LDMX will be able to provide world-leading sensitivity to sub-GeV DM and is able to test many LDM scenarios along the way
  - Impressive physics discovery potential and guaranteed deliverables
- The experiment is ready to move forward with the construction phase
- LDMX could be taking data in 2-3 years after establishing the funding profile and **make a major discovery shortly thereafter**











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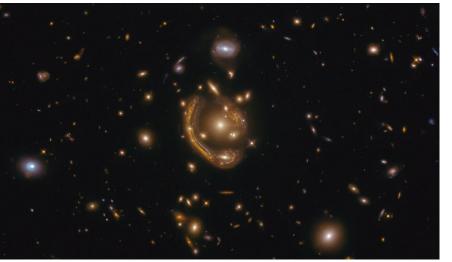


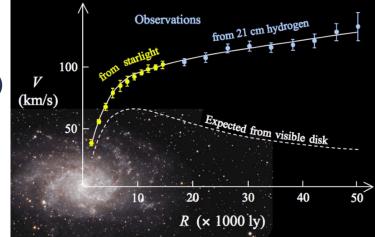




#### **Evidence for Dark Matter**

- Strong case for the existence of dark matter (DM)
  - Galaxy rotation curves
  - Gravitational lensing
  - Cosmic Microwave Background anisotropy
  - Cluster collisions



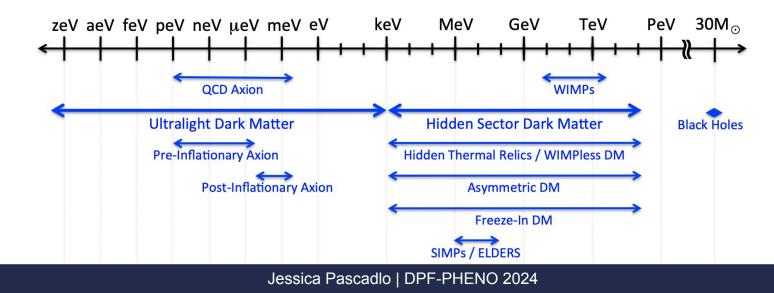


- No detection (yet!) the origin and nature of DM is a key puzzle for particle physics
  - Standard model does not include dark matter
  - How do we narrow down a search region to determine what DM is?



#### **Trying to Understand Dark Matter**

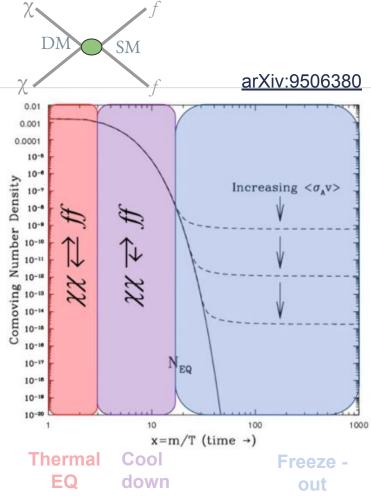
- What <u>do</u> we know?
  - Interacts gravitationally
  - Cosmological abundance
  - Limited interactions with known (SM) matter
- We don't know the mass of the DM



#### **Thermal Dark Matter**

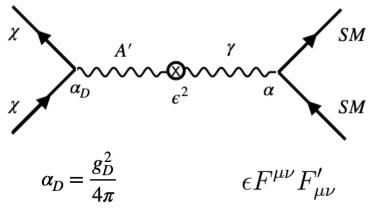
- Assume we are dealing with *particle-like DM*
- DM and SM particles in thermal equilibrium in the very early universe
- As universe cools and expands, DM pairs are no longer in equilibrium, resulting in decreasing amount of interactions
- Universe expands and cools enough such that DM is too dilute to interact → freeze-out
- The current relic density  $\Omega_X$  is related to the annihilation cross section  $\langle \sigma v \rangle$

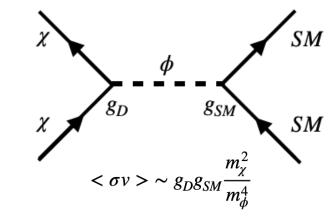
$$\Omega_{\chi} \propto \frac{1}{\langle \sigma v \rangle} \longrightarrow \langle \sigma v \rangle = 3 \times 10^{-26} \ \frac{\mathrm{cm}^3}{\mathrm{s}}$$



## **Light Thermal Dark Matter - Hidden Sector**

- DM could belong to some "hidden sector" that is secluded from the SM
- Sub-GeV DM requires an additional non-SM interaction to maintain the correct relic abundance
  - Mediated by new massive gauge boson



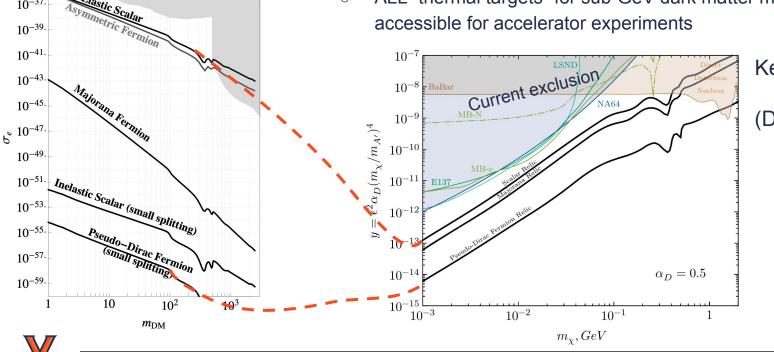


- Additional spin-one gauge boson (dark photon or A')
  - neutral under SM
  - Hidden, broken symmetry **U(1)**
- Kinetically mixing with SM U(1), with factor ε
- Visible and invisible final states

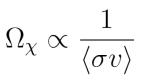


#### **Advantage of DM Production at Accelerators**

- LDM production at accelerators is fairly independent of the DM model, especially when compared to direct detection
  - ALL "thermal targets" for sub-GeV dark matter models are much more Ο accessible for accelerator experiments



Key difference is the non-relativistic (DM-e) vs relativistic (accel.) DM scattering





Thermal and Asymmetric Targets for DM-e Scattering

Elastic Scalar

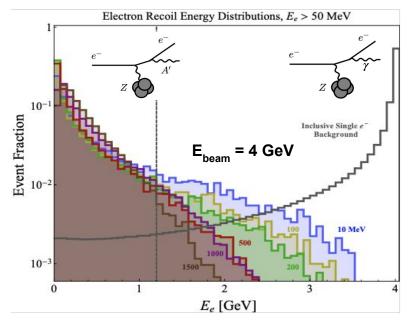
Current exclusion

 $10^{-35}$ 

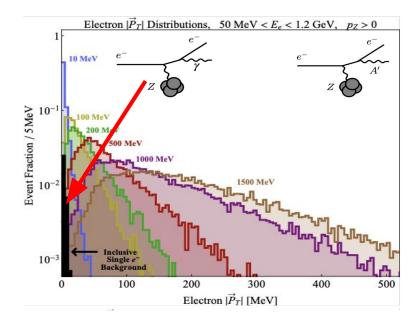
 $10^{-37}$ 

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#### **Dark Photon Kinematics at a Fixed Target Experiment**



- A'→χχ carry away most of the beam energy and escape undetected
  - Opposite behavior for the bremsstrahlung emission

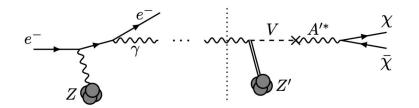


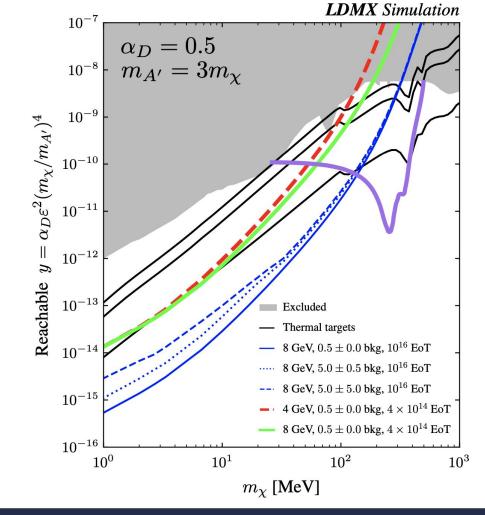
- Recoil electron p<sub>T</sub> spectrum depends strongly on m<sub>A</sub>, for signal
  - Signal identification or extra handle for background rejection



When all systems are background free for 4 signal efficiency of ~3	sstrahlun	g →	outgoing $e^-$	Veto Handles					
10-3 10-4 10-5		*	ent		$+e^+e^-$ $\rightarrow \gamma \rightarrow \text{hadrons}$ $\rightarrow \gamma \rightarrow \mu^+\mu^-$				
	Photo-i Target-area		Muon con Target-area		γ γ μα μα				
EoT equivalent	$4 \times 10^{14}$	$2.1 \times 10^{14}$	$8.2 \times 10^{14}$		$\gamma \to 1n/K_L^0 + \text{soft}$				
Total events simulated	$8.8 \times 10^{11}$	$4.65\times10^{11}$	$6.27 \times 10^8$	0 1 10	$\gamma \rightarrow K^{\pm} + \text{soft}$				
Trigger, ECal total energy < 1.5 GeV		$2.63 \times 10^{8}$	$1.6 \times 10^{7}$	$1.6 \times 10^{8}$	$K^{\pm}$ decay	,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,			
Single track with $p < 1.2 \text{ GeV}$ ECal BDT (> 0.99)	$\begin{array}{c c} 2 \times 10^7 \\ 9.4 \times 10^5 \end{array}$	$2.34 \times 10^{8}$ $1.32 \times 10^{5}$	$3.1 \times 10^4$ < 1	$1.5 \times 10^{8}$ < 1	in ECal				
HCal max PE $< 5$	$  \frac{9.4 \times 10}{< 1}  $	$1.52 \times 10$ 10	< 1	<1					
ECal MIP tracks = $0$	< 1	< 1	< 1	< 1	🔪 arXiv:'	1912.05535 ack			
10-15 10-16	"visible" backgrou			'	increasingly rare photo-nuclear	Extra Tracks			
····	"invisible"	' backgrou	unds $\ll 10^{\circ}$	-16 ν νi	(M  øller + CCQE)	ECal Feature HCal Hits			
Jessica Pascadlo   DPF-PHENO 2024 2									

# Gain additional sensitivity from invisible meson decay channel

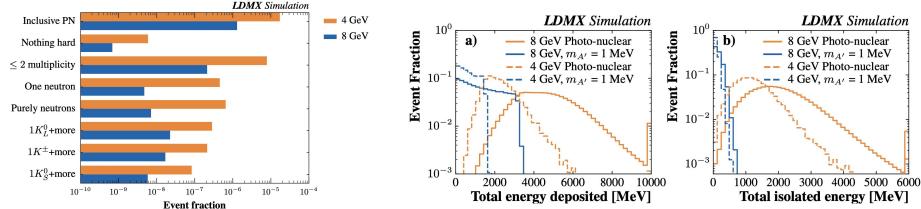








#### **Advantages of 8 GeV Beam**



Overall reduction of PN events that pass the trigger and higher multiplicity

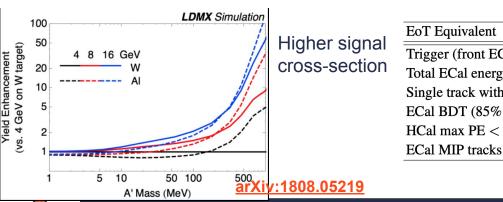


	Photo-1	nuclear	Muon conversion	
	Target-area	ECal	Target-area	ECal
EoT Equivalent	$2.00  imes 10^{14}$	$2.00\times10^{14}$	$2.00\times10^{14}$	$2.00  imes 10^{14}$
Trigger (front ECal energy < 3160 MeV)	$7.57  imes 10^7$	$4.43 \times 10^8$	$2.37  imes 10^7$	$8.12 \times 10^7$
Total ECal energy < 3160 MeV	$2.73 imes10^7$	$7.27 \times 10^7$	$1.76  imes 10^7$	$6.06  imes 10^7$
Single track with $p < 2400$ MeV/c	$3.03 imes10^6$	$6.64  imes 10^7$	$5.32  imes 10^4$	$5.69  imes 10^7$
ECal BDT (85% eff. $m_{A'} = 1 \text{ MeV}$ )	$1.50  imes 10^5$	$1.04 \times 10^5$	< 1	< 1
HCal max $PE < 8$	< 1	2.02	< 1	< 1
ECal MIP tracks = $0$	< 1	< 1	< 1	< 1