

# Forbidden Conformal Dark Matter at a GeV

Steven Ferrante In collaboration with Ameen Ismail, Seung J. Lee, Yunha Lee



# **Gravitational Wave Signals!**

- June 2023 NANOGrav observes new GW signals
- Pulsar timing array sensitive to very low frequencies !



NANOGrav Collaboration, 2306.16217, 2306.16213, 2306.16219

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angle \sim 1/m_{\phi}^2 \gg obs. \ & ext{`secluded''} \ & (m_{\phi} > m_{\sigma}) \end{aligned}$$



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 $(m_\phi < m_\sigma)$ 

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• Provides 5d explanation for  $\, \mathcal{O}(m_\sigma) \sim \mathcal{O}(m_\phi)$  !



Ruderman et al, 1505.07107

$$\int d^5x ~~ {\cal L}_\eta ~~
ightarrow~\int d^4x ~{\cal L}_\chi$$

$$\int d^5x \, {\cal L}_\eta \, 
ightarrow \, \int d^4x \, {\cal L}_\chi \, \sim \chi^4 - \chi^{4+\epsilon}$$

$$egin{array}{rcl} \int d^5x \ \mathcal{L}_\eta &
ightarrow \int d^4x \ \mathcal{L}_\chi &
ightarrow \int d^4x \ \mathcal{L}_\chi &
ightarrow \int d^4x \ \mathcal{L}_\sigma &ec \chi = \langle \chi 
angle + \sigma \end{array}$$

-

$$\begin{split} \int d^5 x \ \mathcal{L}_{\eta} &\to \int d^4 x \ \mathcal{L}_{\chi} \to \int d^4 x \ \mathcal{L}_{\sigma} \\ \mathcal{L} = \mathcal{L}_{\rm SM} + \frac{1}{2} (\partial_{\mu} \sigma)^2 - \frac{1}{2} m_{\sigma}^2 \sigma^2 - \frac{5}{6} \frac{m_{\sigma}^2}{f} \sigma^3 - \frac{11}{24} \frac{m_{\sigma}^2}{f^2} \sigma^4 \end{split} \Big\}_{\text{Potential}}^{\text{GW}}$$

$$\int d^5 x \ \mathcal{L}_{\eta} \rightarrow \int d^4 x \ \mathcal{L}_{\chi} \rightarrow \int d^4 x \ \mathcal{L}_{\sigma}$$

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$$+ \frac{1}{2} (\partial_{\mu} \phi)^2 - \frac{1}{2} m_{\phi}^2 \phi^2 - \frac{1}{4!} \lambda_{\phi}^4 - \left(\frac{2\sigma}{f} + \frac{\sigma^2}{f^2}\right) \frac{1}{2} m_{\phi}^2 \phi^2$$

$$- \frac{\sigma}{\Lambda^2/f} \left[ \sum_{\text{fermions}} m_{\psi} \overline{\psi} \psi + m_h^2 h^2 - 2m_W^2 W_{\mu}^+ W^{-\mu} - m_Z^2 Z_{\mu} Z^{\mu} \right] \begin{cases} \text{Spurion} \\ \text{Analysis} \\ m_{\phi} \to m_{\phi} \frac{\chi}{\langle \chi \rangle} \\ \text{Spurion} \\ \text{Analysis} \\ m \to m_{\chi} \frac{\chi}{\langle \chi \rangle} (R \langle \chi \rangle)^2 \end{cases}$$

• Dilaton interactions come from Goldberger-Wise stabilization

$$\int d^5 x \ \mathcal{L}_{\eta} \rightarrow \int d^4 x \ \mathcal{L}_{\chi} \rightarrow \int d^4 x \ \mathcal{L}_{\sigma}$$

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• These interactions are used to calculate Relic Abundance!

#### Belanger et al, 1005.4133

• Deconfined ightarrow Confined ... @ critical temperature  $~T_c \sim \sqrt{m_\sigma f}$ 

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- Gravitational Waves:
  - Frequency

$$\left. f_{
m GW} \sim rac{eta_{
m GW}}{H} = T rac{dS_b}{dT} 
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- Gravitational Waves:
  - Frequency
  - Amplitude

$$lpha_{
m GW} \sim rac{
m energy released during PT}{
m energy of radiation bath} \sim rac{T_c^4}{T_n^4} \gg 1$$

# **Dilaton Phenomenology**

- Dilaton coupling to fermions = Higgs couplings, scaled by  $\kappa = \frac{vf}{\Lambda}$
- **Theoretical Bounds** 
  - Dilaton EFT  $m_{\sigma} < 4\pi f$ Ο
  - PT Completion  $\Gamma > H^4$
  - Thermalization  $\Gamma_{\sigma} > H$

# **Dilaton Phenomenology**

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  - $\circ$  Dilaton EFT  $m_\sigma < 4\pi f$
  - $\circ$  PT Completion  $\Gamma > H^4$
  - $\circ$  Thermalization  $\Gamma_{\sigma} > H$
- Experimental Bounds
  - Cosmological bounds on late decays
  - Supernovae cooling
  - CHARM beam dump experiment
- NANOGrav prediction
  - $\circ$  Frequency  $eta_{
    m GW}/H < 27$
  - $\circ$  Amplitude  $T_R \in (0.017, 3.3)~{
    m GeV}$

# **Dilaton Phenomenology**



# **DM Phenomenology**

• Mass splitting defined by  $\Delta = rac{m_\sigma - m_\phi}{m_\phi}$ 

... "how 'forbidden' the DM is"

- Other experimental constraints:
  - $\circ$  Merging Galaxy Clusters (Self-Interactions  $\sigma_{SI}$  )

 $m_{\phi}$ 

 $\circ$  Direct Detection Constraints (DM-Nucleon  $\sigma_{\phi N}$  )

# **DM Phenomenology**



 $f = 4.5 m_{\sigma}$ 

# Conclusion

- Stochastic GWs can provide a new probe of DM models or other new physics scenarios.
- Forbidden conformal DM at a GeV is a natural hypothesis for the NANOGrav signal
  - Even outside of NANOGrav, provides a viable DM candidate
- Future:
  - Consider entire frequency spectrum and compare to data, SMBHBs
  - Consider secluded case -- (BHs can open channels for d-wave DM...)
  - Consider non-minimal scenarios
  - Wait eagerly for more data !!

BACKUP

#### Backup – nanograv bounds



 $T_R \sim T_*$  $eta/H = (8\pi)^{1/3}/(H_*R_*)$ 

# **Backup – dilaton couplings to trace anomaly**

$$-\frac{\sigma}{\Lambda^2/f} \left[ \frac{\beta_e(e)}{2e^3} F_{\mu\nu}^2 + \frac{\beta_3(g_3)}{2g_3^3} \left( G_{\mu\nu}^a \right)^2 + \sum_{\text{fermions}} \gamma_\psi \overline{\psi} \psi \right]$$

#### **Backup – detailed result for Relic Abundance approx**

$$\Omega_{\phi}h^{2} \sim 0.1g_{\Delta}(x_{f}) \frac{9\pi (f/m_{\phi})^{4}m_{\phi}^{2}}{(20 \text{ TeV})^{2}}e^{2\Delta x_{f}},$$
$$g_{\Delta}(x_{f}) = \frac{2(1+\Delta)^{4}}{\sqrt{\Delta(2+\Delta)}(1-4\Delta-2\Delta^{2})^{2}} \left[1-2\Delta x_{f}e^{2\Delta x_{f}}\int_{2\Delta x_{f}}^{\infty} dt \frac{e^{-t}}{t}\right]^{-1}.$$

#### **Backup – thick wall approx**



# **Backup – model details**

$$V(\chi) = \frac{3N^2}{2\pi^2} \left[ -\lambda \chi^4 + \lambda_{\rm GW} \frac{\chi^{4+\alpha}}{R^{-\alpha}} \right] + V_0$$
$$\langle \chi \rangle = R^{-1} \left( \frac{\lambda}{\lambda_{\rm GW}} \right)^{1/\alpha}$$
$$\chi = \langle \chi \rangle + \sqrt{2\pi^2/3N^2} \sigma$$
$$m_{\sigma}^2 = 4\alpha \lambda \langle \chi \rangle^2$$
$$f = \sqrt{3N^2/2\pi^2} \langle \chi \rangle$$

# **Backup – additional constraints**

- DM annihilation to SM through dilaton portal
  - Cross section below constraints on EM energy injection in early universe
  - Can get a Sommerfeld enhancement, ends up being too small
- Higgs → invisible decays into KK modes
- Can estimate

$$\Gamma(h o ext{invisible}) \sim rac{m_h}{8\pi} \Big(rac{f}{\Lambda}\Big)^4 < 0.11$$

 $(20 \text{ TeV})^{-2}$  $\Omega h^2 \sim$  $\langle \sigma v 
angle_{_{\phi\phi o \sigma\sigma}}$ 

 $(20 \text{ TeV})^{-2}$  $\Omega h^2 \sim$  $\left\langle oldsymbol{\sigma} \mathcal{V} 
ight
angle_{\sigma\sigma o \phi\phi} igg( rac{n_{\sigma}^{eq}}{n_{\phi}^{eq}} igg)^2$ 

 $\Delta = (m_\sigma - m_\phi)/m_\phi$  .  $x=m_{\phi}/T$ 

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$$\Omega h^2 \sim rac{\left(20 \ {
m TeV}
ight)^{-2}}{\left| {}^{*}_{
m s} {\scriptstyle \sim \sigma} {\scriptstyle \circ \sigma} {\scriptstyle + \sigma} {\scriptstyle \circ \sigma} {\scriptstyle \circ \sigma} {\scriptstyle + ({
m t}, {
m u})} 
ight|^2 e^{-2\Delta x}}$$

• Set by  $\phi\phi
ightarrow\sigma\sigma$ 

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# • Numerics: micrOMEGAs

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 $\Gamma > H^4 \Rightarrow T_n$  $\overset{\parallel}{\underset{b}{\mathbb{S}_b(T_n)}} \overset{\parallel}{<} 4\logig(rac{M_{pl}T_n}{T^2}ig)$ 300 250 200 150 0.001 0.010 0.100

Harling et al, 1711.11554

# **DM Candidates** "What the hell is going on" – Nima 2017

• WIMPs facing increasing experimental pressure ... exciting time for other scales !



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- Dark sector w/ spontaneously broken conformal symmetry
- GeV-scale phase transition

#### Asai and Carena DM-DS Summary

#### **DM Phenomenology**





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